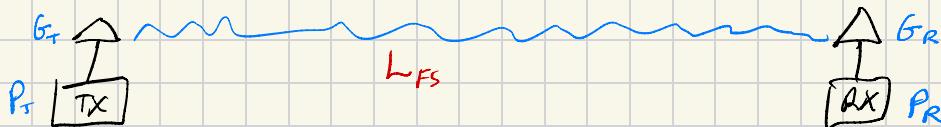


## Communication + Link Budgets

- \* Noise limits our ability to detect/decode signals
- \* The metric that matters is signal-to-noise ratio (SNR or  $S/N$ ) or closely related  $E_b/N_0$  for digital systems
- \* How do we calculate power received by a radio?



$P_T$  = Transmit Power

$G_T$  = Transmitter antenna gain

$L_{FS}$  = Free-Space Loss

$G_R$  = Receiver Antenna Gain

$R_R$  = Received power

- In a real system there are additional effects: cable losses, multi-path, atmospheric attenuation, rain fade, polarization losses.
- Can be time varying and frequency dependent
- Add everything up:

$$P_R = \frac{P_T G_T G_R}{L_{FS}}$$

- Typically done in decibels with units of dB-milliwatts (dBm)

$$P_R = P_T + G_T - L_{FS} + G_R \quad \left. \right\} \text{ "Friis Equation"}$$

(dBm)      (dBm)      (dB)      (dB)

\* Free-Space Loss

$$L_{FS} = \left( \frac{4\pi r}{\lambda} \right)^2 = 20 \log_{10}(r) - 20 \log_{10}(2) + 20 \log_{10}(4\pi) \\ = 20 \log_{10}(r) + 20 \log(f) + 20 \log_{10}\left(\frac{4\pi}{c}\right)$$

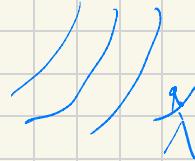
distance  
wavelength  
frequency

- Inverse-square law plus a frequency-dependent term



$$E(r) = \frac{P_T}{A(r)}$$

↑  
Influence ( $\psi_{use}$ )



- Diffraction limit  $\Rightarrow \theta \approx \frac{\lambda}{D}$

$\approx \frac{\lambda}{\sqrt{4\pi A_r}}$   
wavelength  
dish diameter  
Area of dish

$$\Rightarrow A(r) \approx \pi (\sigma r)^2 = \frac{\pi \lambda^2 r^2}{4\pi A_r} = \frac{\lambda^2 r^2}{A_r}$$

$$P_R = E(r) A_r = P_T \left( \frac{A_r A_e}{\lambda^2 r^2} \right)$$

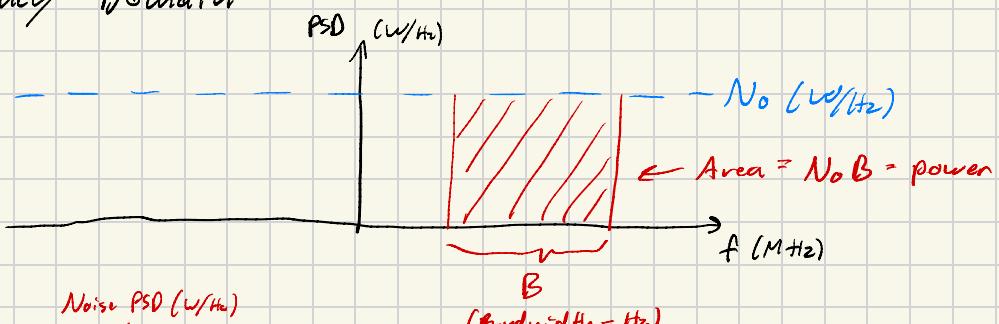
Another version of the Friis equation in terms of dish areas instead of gains (actually original)

## \* Antenna Gain

- A unitless quantity that measures how directed an antenna is
- Directly proportional to collecting area or "aperture"
- In general, larger antenna  $\Rightarrow$  higher gain
- Higher gain  $\Rightarrow$  tighter pointing requirements

\* How do we calculate noise power received by a radio?

- Noise is everywhere due to thermal excitation of (mostly) electrons
- This ambient thermal noise is "White" (equal power at all frequencies) and it scales with temperature
- There is more noise if the background is hot vs cold (e.g. deep space vs. planet).
- Frequency Domain



$$N = N_0 B \leftarrow \text{Bandwidth (Hz)}$$

$$\begin{matrix} \nearrow \\ \text{Noise power (W)} \end{matrix} \quad \begin{matrix} \searrow \\ \Rightarrow N = K_B T B \end{matrix}$$

$$N_0 = K_B T \leftarrow \text{temperature (Kelvin)}$$

R

Boltzmann Constant (J/K)

- For analog signals  $P_A/N = S/N$  is used to analyze link loss

- For digital signals, we use something like  $S/N$  integrated over a bit:

$$\frac{E_B}{N_0} = \frac{S}{N} \cdot \frac{B}{R_B} \times \text{Bandwidth (Hz)} \leftarrow \text{Bit rate (bits/sec)}$$

## Shannon - Hartley Theorem

- The maximum bit rate achievable over a channel link with white noise is:

$$C = B \log_2 \left( 1 + \frac{S}{N} \right)$$

<sup>↑</sup>  
"Channel Capacity" =  $\max R_B$  (bit/sec)

- With a little math, you can show:

$$\frac{E_b}{N_0} > 1/\ln(2) \approx -1.59 \text{ dB}$$

- It is actually possible to decode signals below the noise floor
- Modern systems can get very close to the Shannon limit in practice.

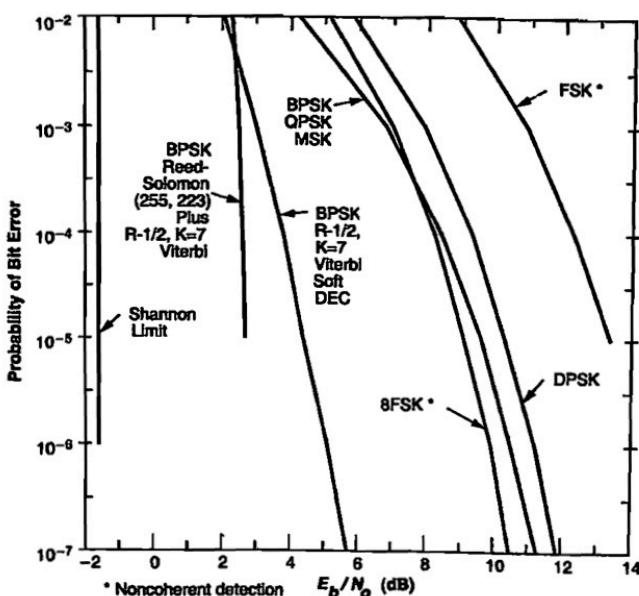


Fig. 13-9. Bit Error Probability as a Function of  $E_b/N_0$ . The theoretical performance limit can be approached by use of error correction coding.

\* Typically add several dB of margin in a practical system