

热力学与统计物理-第八次作业

吴远清-2018300001031

2020 年 8 月 29 日

Problem 6.1

Answer:

The probability that the system is in a state with energy E is proportional to the Boltzmann factor $e^{-E/kT}$.

Then the ratio of the probability of being in the first excited state to the probability of being in the ground state is:

$$\frac{e^{-3\hbar\omega/2kT}}{e^{-\hbar\omega/2kT}} = e^{-\hbar\omega/kT} \quad (1.1)$$

(b)

By the definition of mean value:

$$\bar{E} = \frac{\hbar\omega}{2} \frac{1 + 3e^{-\hbar\omega/kT}}{1 + e^{-\hbar\omega/kT}} \quad (1.2)$$

Problem 6.2

Answer:

The mean energy per particle is:

$$\bar{\epsilon} = \frac{\mu H e^{-\mu H/kT} - \mu H e^{\mu H/kT}}{e^{-\mu H/kT} + e^{\mu H/kT}} = -\mu H \tanh \frac{\mu H}{kT} \quad (2.1)$$

So:

$$E = -N\mu H \tanh \frac{\mu H}{kT} \quad (2.2)$$

Problem 6.4

Answer:

The power absorbed is proportional to the difference in the number of nuclei

in the two levels.

This is:

$$n_+ - n_- = \frac{Ne^{\mu H/KT}}{e^{\mu H/KT} + e^{-\mu H/KT}} - \frac{Ne^{-\mu H/KT}}{e^{\mu H/KT} + e^{-\mu H/KT}} \quad (3.1)$$

Since $\mu H \ll kT$:

$$n_+ - n_- \approx N \frac{\left(1 + \frac{\mu H}{kT} - 1 + \frac{\mu H}{kT}\right)}{1 + \frac{\mu H}{kT} + 1 - \frac{\mu H}{kT}} = \frac{N\mu H}{kT} \quad (3.2)$$

Problem 6.5

Answer:

The volume element of atmosphere shown must be in equilibrium under the forces of the pressure and gravity. Then if m is the mass per particle, A the area, and g the acceleration of gravity, we have:

$$P(z + dz)A - p(z)A = \frac{dp}{dz}dzA = -mn(z)Adzg \quad (4.1)$$

With $p = nKT$:

$$\frac{dn(z)}{dz} = -\frac{mg}{kT}n(z) \quad (4.2)$$

So:

$$n(z) = n(0)e^{-mgz/kT} \quad (4.3)$$

Problem 6.6

Answer:

(a)

As T approaches 0 the system tends to the low energy state, while in the

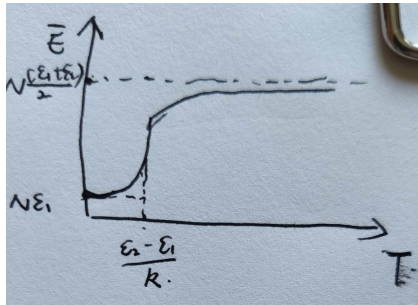


图 1: (a)

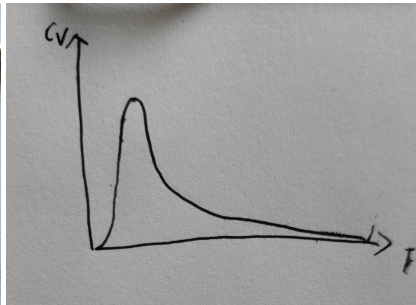


图 2: (b)

limit of high temperatures all states become equally probable. The energy goes from the low to the high temperature limit when $\epsilon_2 - \epsilon_1 \approx KT$. The specific heat is, $C_V = \left(\frac{\partial \bar{E}}{\partial T} \right)_V$, i.e., the slope of \bar{E}

(c)

$$\bar{E} = N \left[\frac{\epsilon_1 e^{-\epsilon_1/kT} + \epsilon_2 e^{-\epsilon_2/kT}}{e^{-\epsilon_1/kT} + e^{-\epsilon_2/kT}} \right] = N \left[\frac{\epsilon_1 + \epsilon_2 e^{-(\epsilon_2 - \epsilon_1)/KT}}{1 + e^{-(\epsilon_2 - \epsilon_1)/KT}} \right] \quad (5.1)$$

$$T \rightarrow 0, \bar{E} \rightarrow N\epsilon_1 \quad (5.2)$$

$$T \rightarrow \infty, \bar{E} \rightarrow N \left(\frac{\epsilon_1 + \epsilon_2}{2} \right) \quad (5.3)$$

So:

$$\frac{\epsilon_2 - \epsilon_1}{kT} = \ln 3 \approx 1 \quad \text{or} \quad (\epsilon_2 - \epsilon_1) \approx kT \quad (5.4)$$

Then the heat capacity is:

$$c_V = \frac{\partial \bar{E}}{\partial T} = \frac{N (\epsilon_2 - \epsilon_1)^2 e^{-(\epsilon_2 - \epsilon_1)/KT}}{KT^2 [1 + e^{-(\epsilon_2 - \epsilon_1)/KT}]^2} \quad (5.5)$$

$C_V \rightarrow 0$ as $T \rightarrow 0, T \rightarrow \infty$.

Problem 6.10

Answer:

(a)

The centrifugal force $m\omega^2 r$ yields the potential energy $-\frac{m\omega^2 r^2}{2}$. Since the probability that a molecule is at r is proportional to the Boltzmann factor, the density is:

$$p(r) = p(0) e^{m\omega^2 r^2 / 2kT} \quad (6.1)$$

(b)

Substituting $\mu = N_A m$, the molecular weight, into (6.1) and evaluating this expression at r_1 and r_2 we have:

$$\mu = \frac{2N_A kT}{\omega^2 (r_1^2 - r_2^2)} \ln \frac{\rho(r_1)}{\rho(r_2)} \quad (6.2)$$