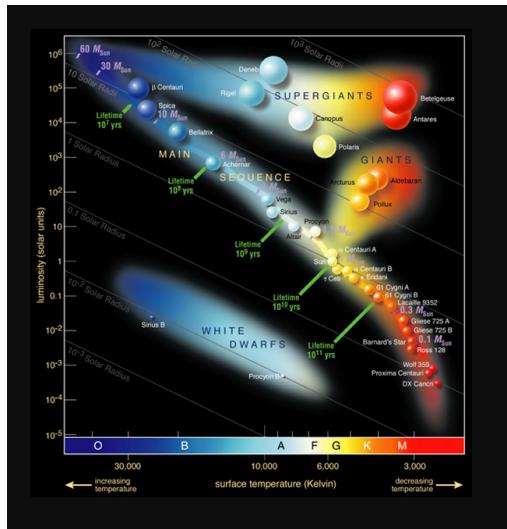
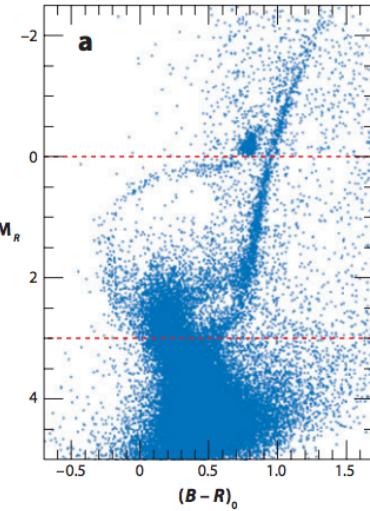


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So what is this?



2

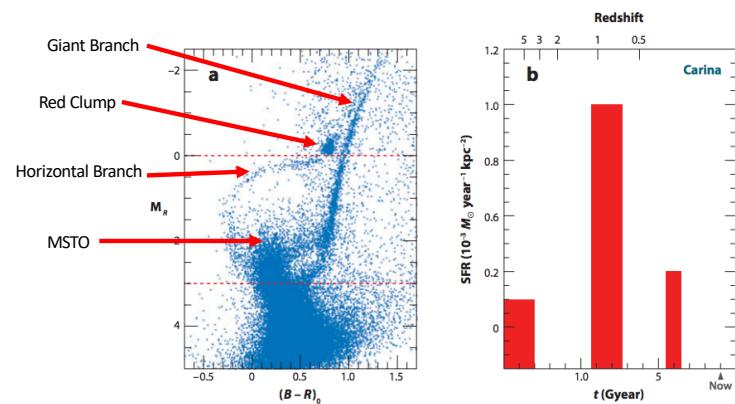
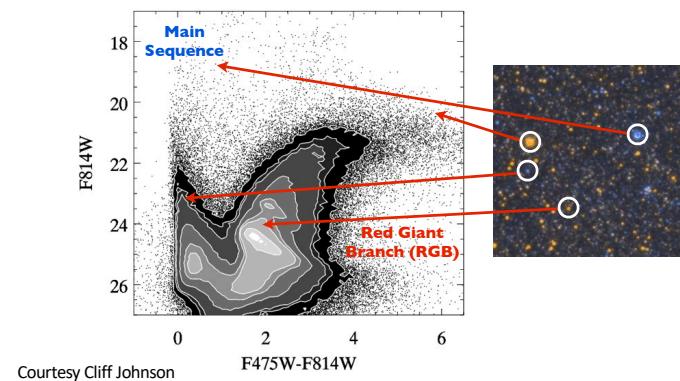


Figure 4

(a) A color-magnitude diagram of the Carina dwarf spheroidal (obtained by M. Mateo with the CTIO 4-m and MOSAIC camera, private communication) in the central 30' of the galaxy. This clearly shows the presence of at least three distinct MSTOs. (b) The star-formation history of the central region of Carina determined by Hurley-Keller, Mateo & Nemec (1998), showing the relative strength of the different bursts. The ages are also shown in terms of redshift.

## The Color-Magnitude Diagram (CMD)

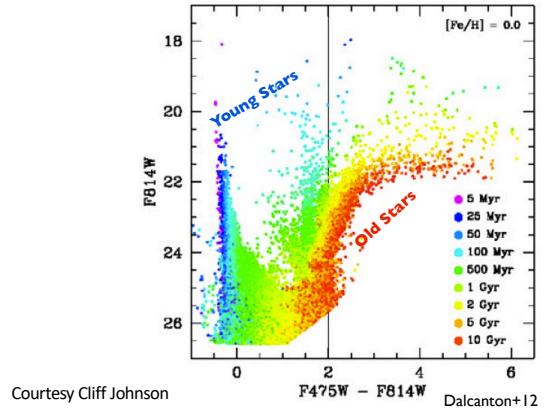
Real galaxies ....



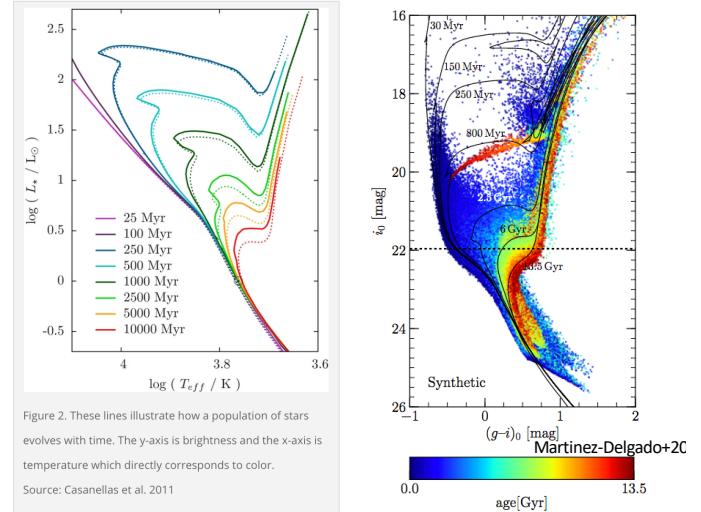
Courtesy Cliff Johnson

4

# Age Dating with CMDs



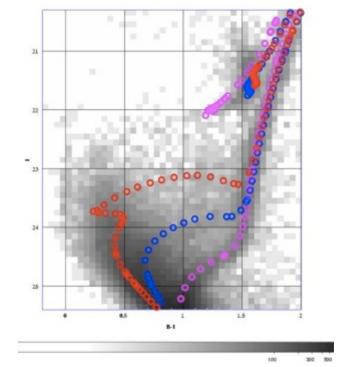
## Isochrones



6

Deriving SFH from CMDs: SFH of a quenched dwarf: Leo I

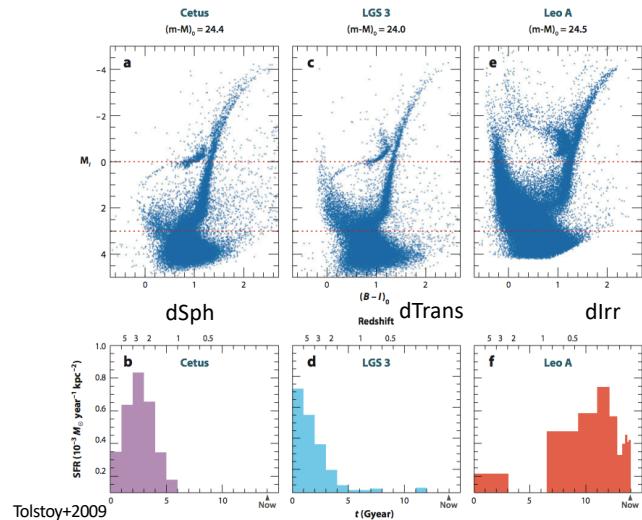
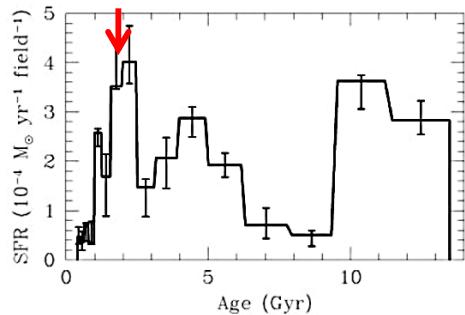
$M_* \sim 5 \times 10^6 M_\odot$



### Orbits: HST Proper motions

Accretion by Milky Way  
2 Gyr ago (Sohn, GB+2013)

Rapid quenching (Wetzel+2015)

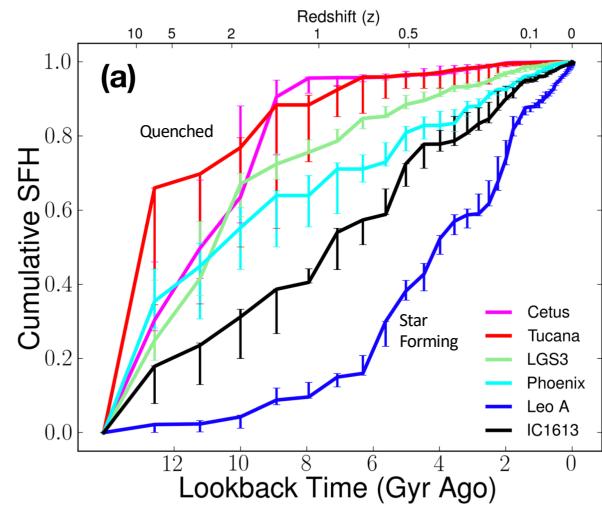


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## Cumulative SFHs

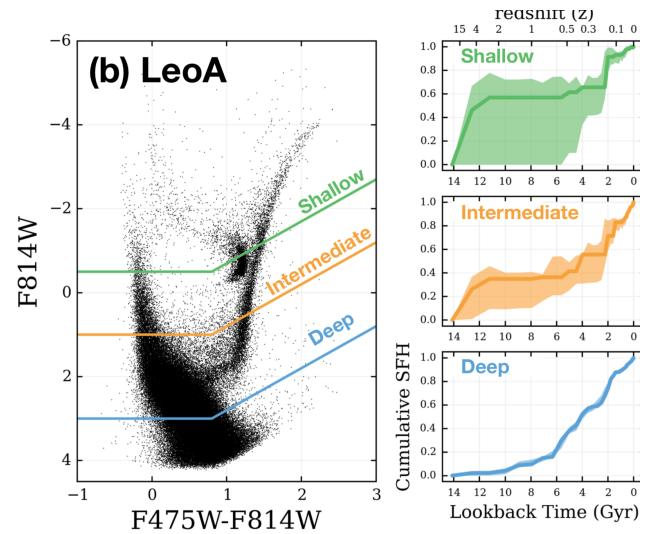
The SFHs of isolated ( $>300$  cpc) from the MW/M31) dwarf galaxies measured from deep CMDs (Skillman et al. 2014). The varied SFHs trace their diverse assembly history. *HST* can only provide deep CMDs for 9 isolated galaxies, while *JWST* can access hundreds.

WST ERS Proposal  
J Weisz



## The Need for JWST

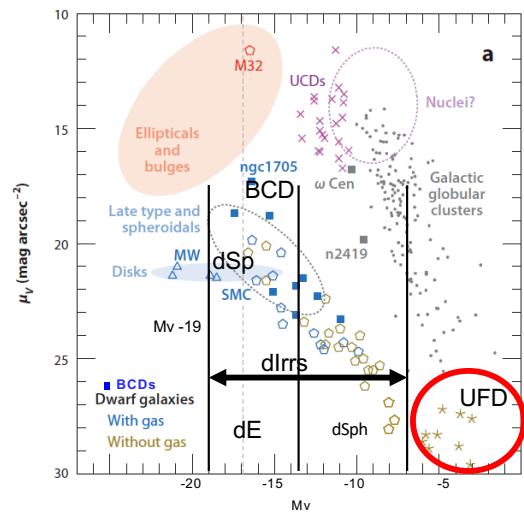
JWST ERS Proposal  
PI Weisz



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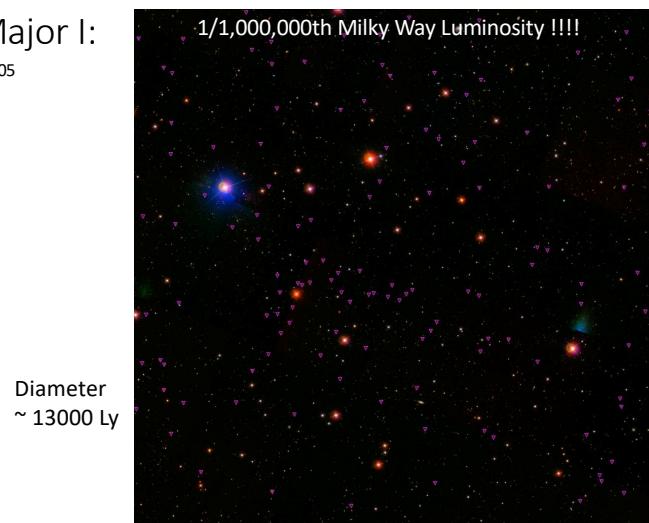
## Ultra Faint Dwarfs $M^* < 10^5 M_\odot$

Tolstoy+ 2009

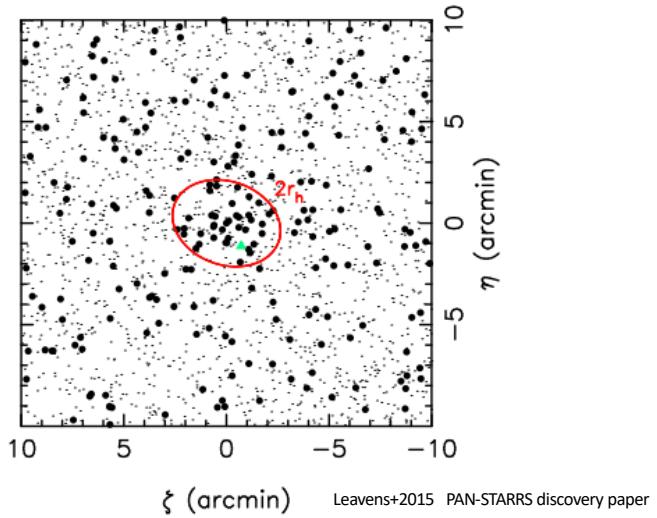


## Ursa Major I:

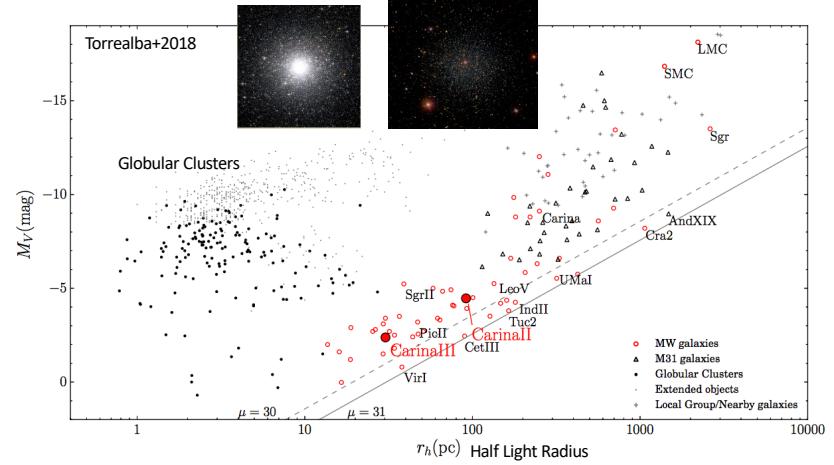
Willman + 2005



12

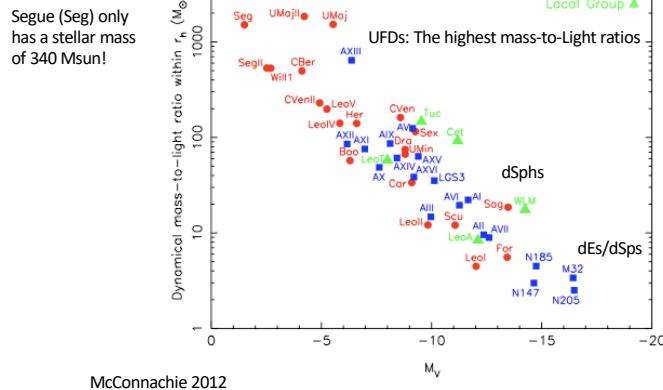


## Size vs Luminosity (abs. magnitude)



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## $M_{\text{Dynamical}} / L$ Ratios



McConnachie 2012

## Isochrone fitting to CMDs of UFDs

Green Isochrone:  
Old: 13.2 Gyr,  
Metal Poor [Fe/H] = -2.3

Single age population.

Brown+2014 ApJ 796

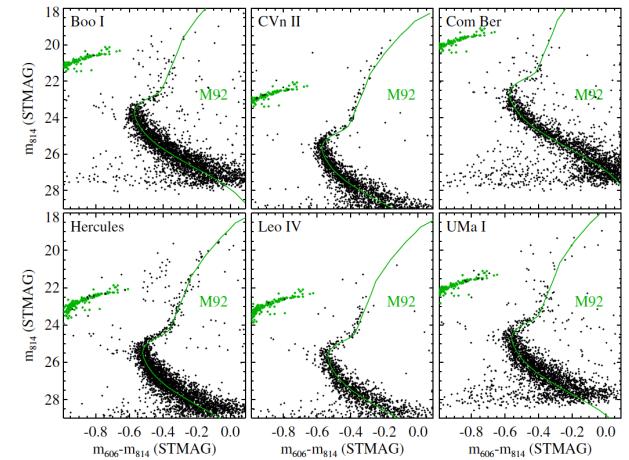


Figure 1. CMD of each UFD in our sample (black points). For reference, we show the empirical ridge line for the MS, SGB, and RGB M92 (green curve), along with the HII locus in M92 (green points). The M92 fiducial has been placed at the distance and reddening for each galaxy (Table 1), matching the luminosity of H stars and the color of the lower MS stars. Because the CMD of each galaxy looks, to first order, like that of a ancient metal-poor globular cluster, the stellar populations of each galaxy is dominated by ancient metal-poor stars. The CMDs of these galaxies are all extremely similar to one another, implying they have similar stellar populations and star formation histories.

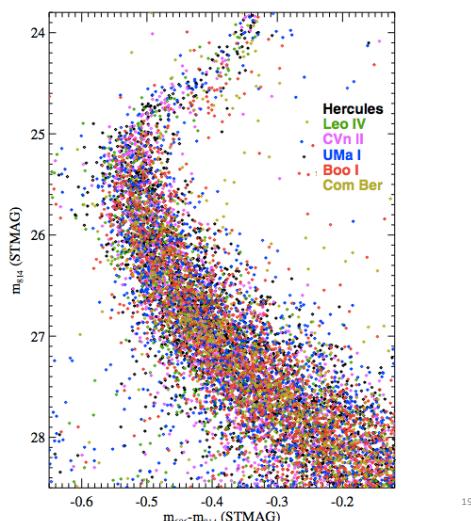
16

Brown +2014

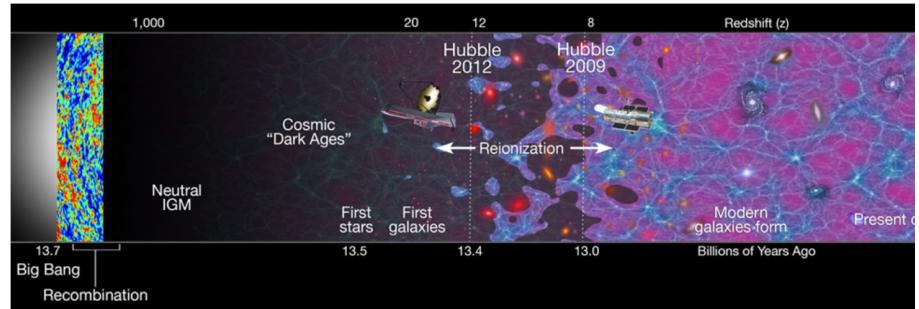
CMD of each UFD (colored points) shifted to the distance and reddening of Hercules dwarf and zoomed into the CMD region most sensitive to age.

The similarities of the six CMDs imply that the UFD populations are extremely similar in age and metallicity.

A global event caused quenching

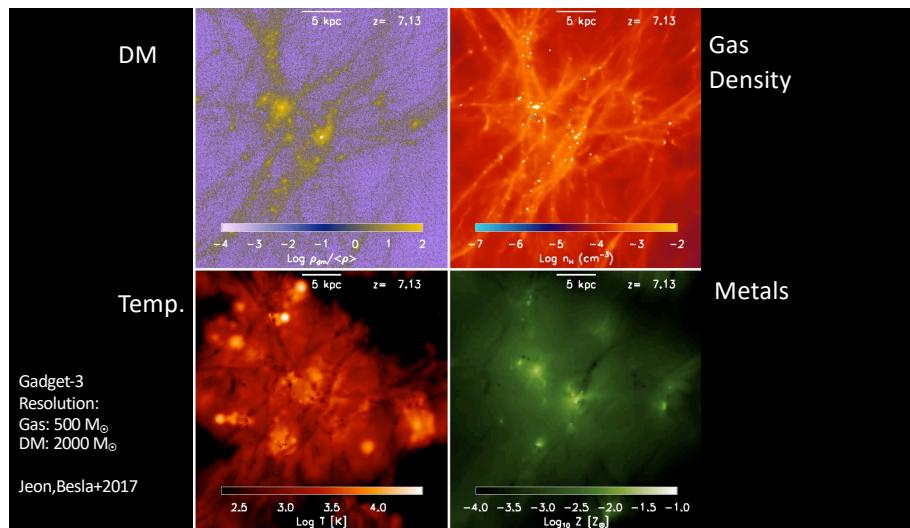


Reionization  
z ~ 10-6 ?

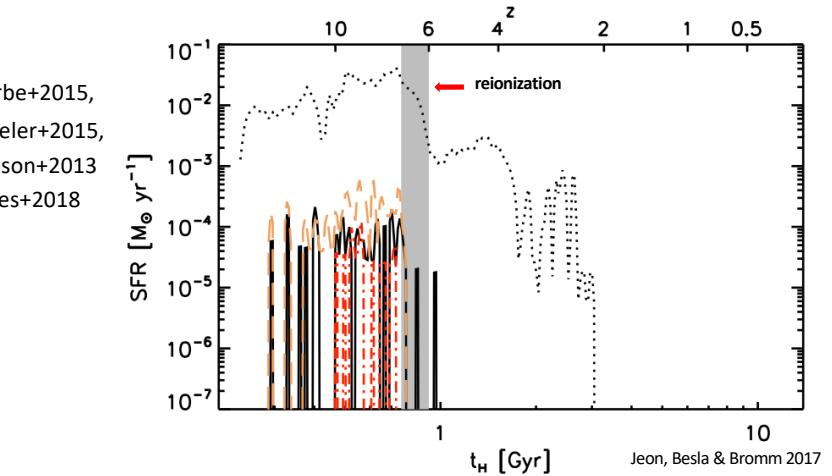


Reionization is Patchy: <https://www.youtube.com/watch?v=kifF3RYcfn0>

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SFHs: Reionization + Feedback

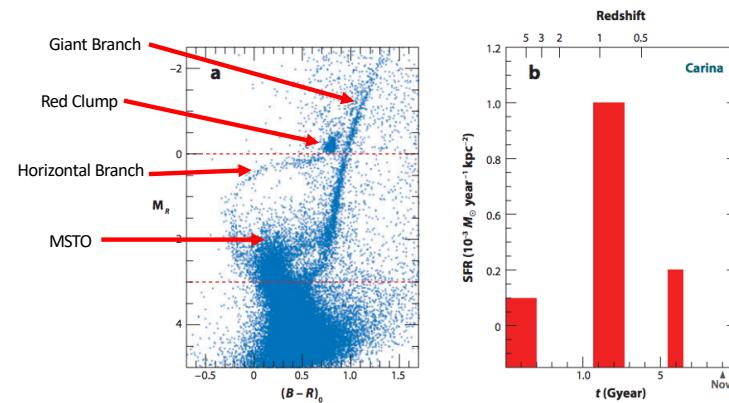


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## Reasons for why Satellite Galaxies are quenched

- Tidal Sharing
- Ram Pressure Sharing
- Preservation (cut off of gas supply)

But for the smallest galaxies,  
 Feedback from stars  
 &  
 Reionization  
 Also matter.



**Figure 4**

(a) A color-magnitude diagram of the Carina dwarf spheroidal (obtained by M. Mateo with the CTIO 4-m and MOSAIC camera, private communication) in the central 30' of the galaxy. This clearly shows the presence of at least three distinct MSTOs. (b) The star-formation history of the central region of Carina determined by Hurley-Keller, Mateo & Nemec (1998), showing the relative strength of the different bursts. The ages are also shown in terms of redshift.

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