

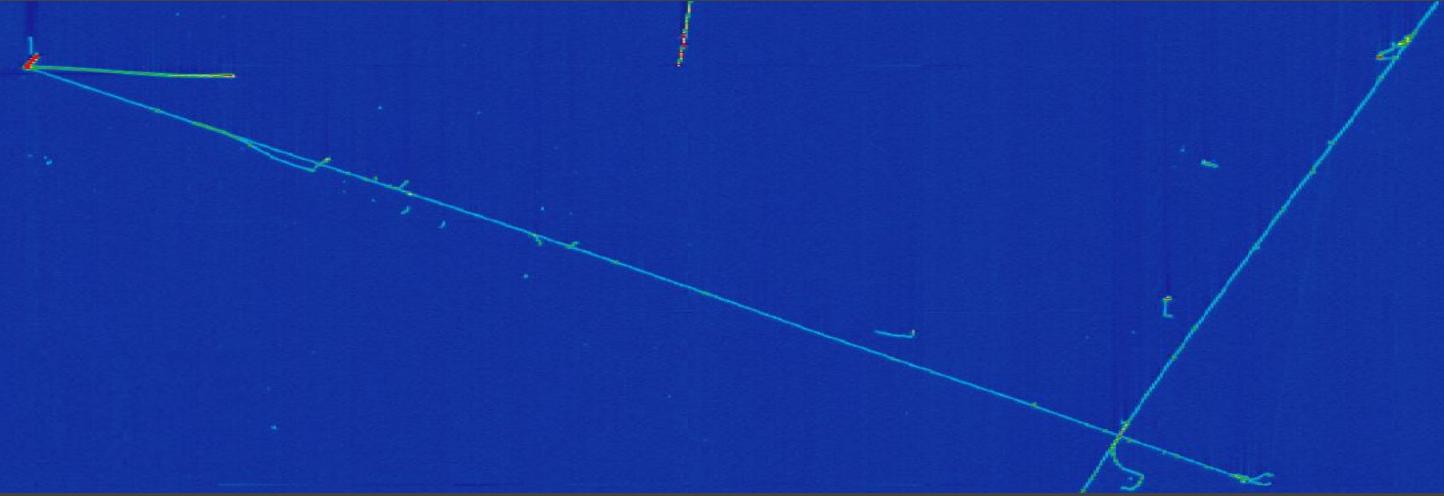
# Machine Learning for Particle Imaging Detectors in Experimental Neutrino Physics

Kazuhiro Terao

SLAC National Accelerator Lab.

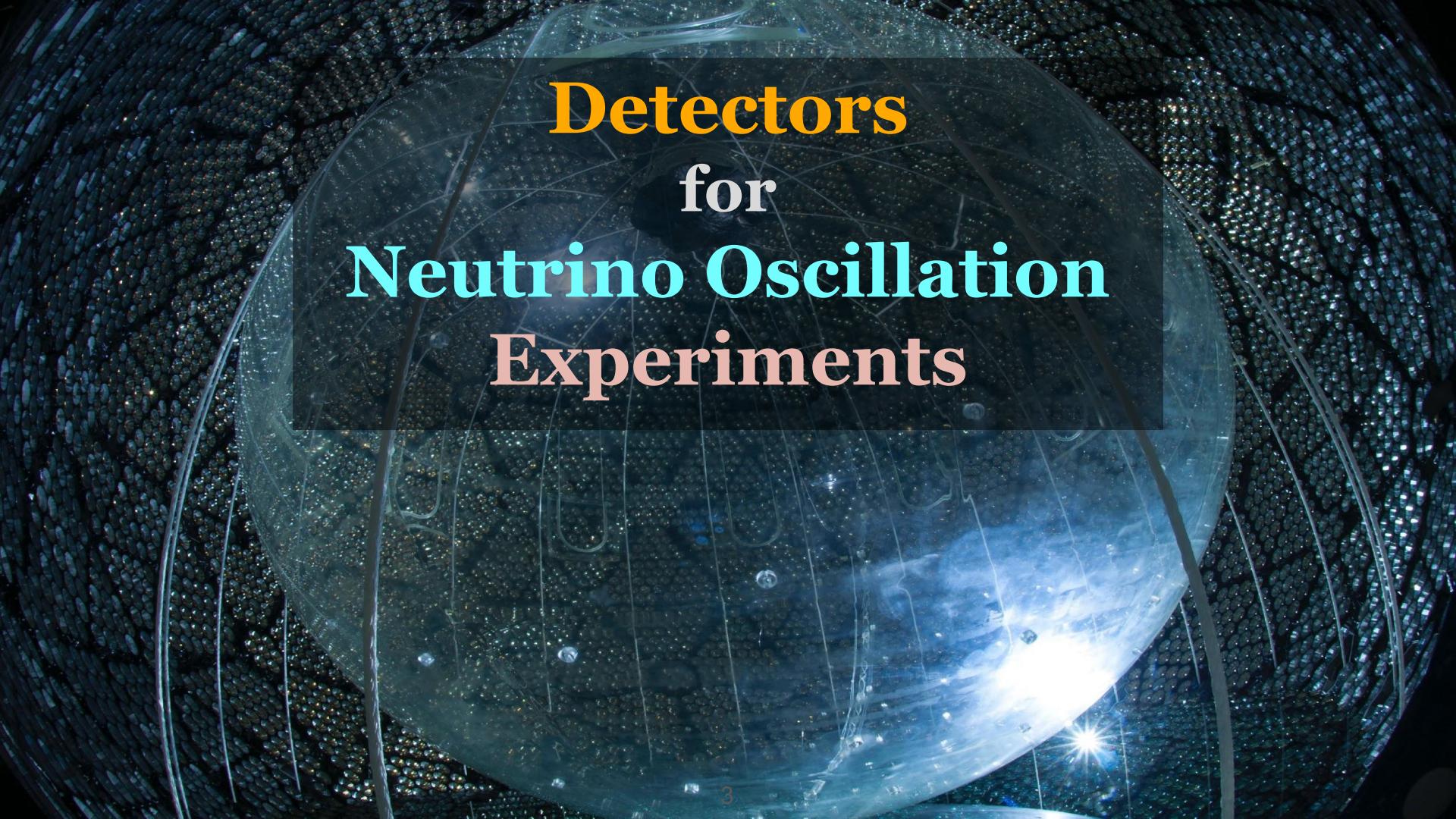
Deep learning and Physics 2020 Seminar Series





# Outline

1. Neutrino detectors
2. Machine Learning & Computer Vision Applications
3. ML-based Neutrino Data Reconstruction Chain
4. Summary

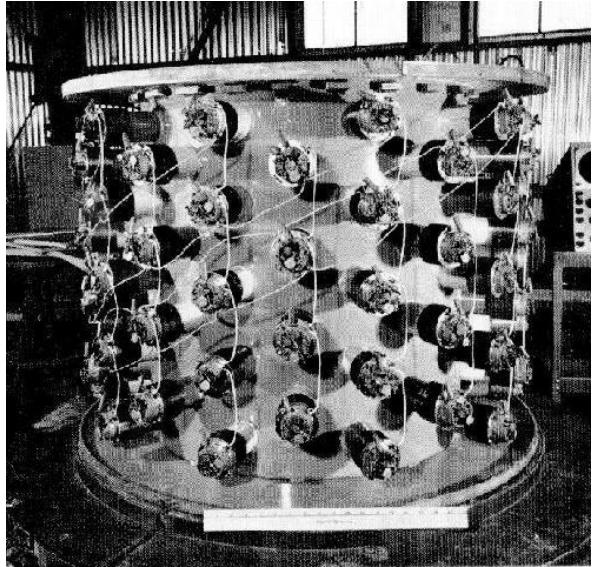


# Detectors for Neutrino Oscillation Experiments

# Machine Learning & Computer Vision in Neutrino Physics

## Neutrino Detectors: Early Days

SLAC



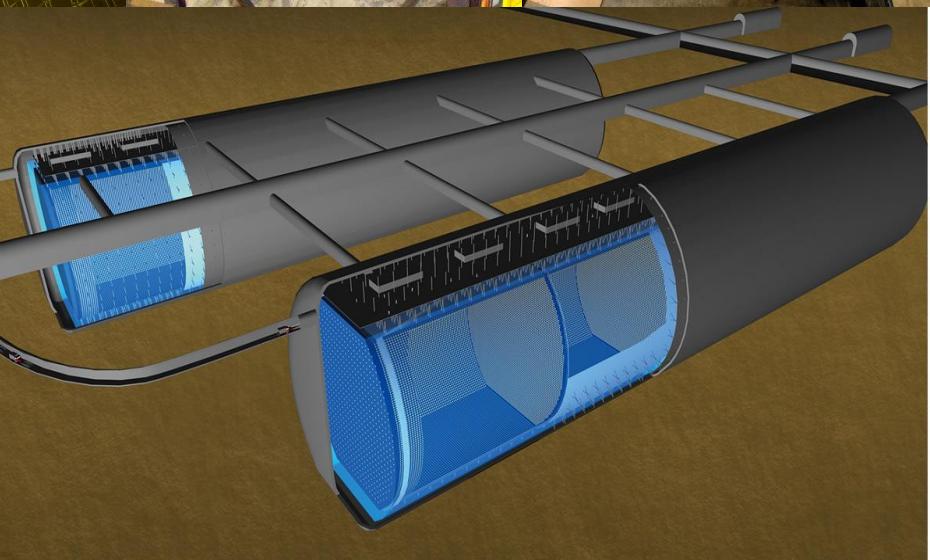
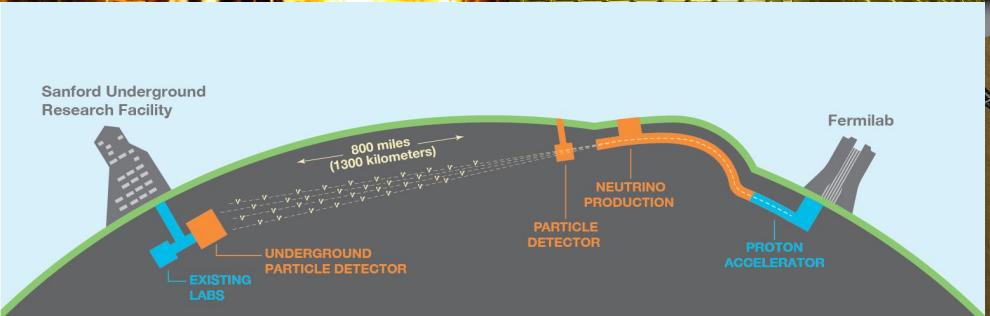
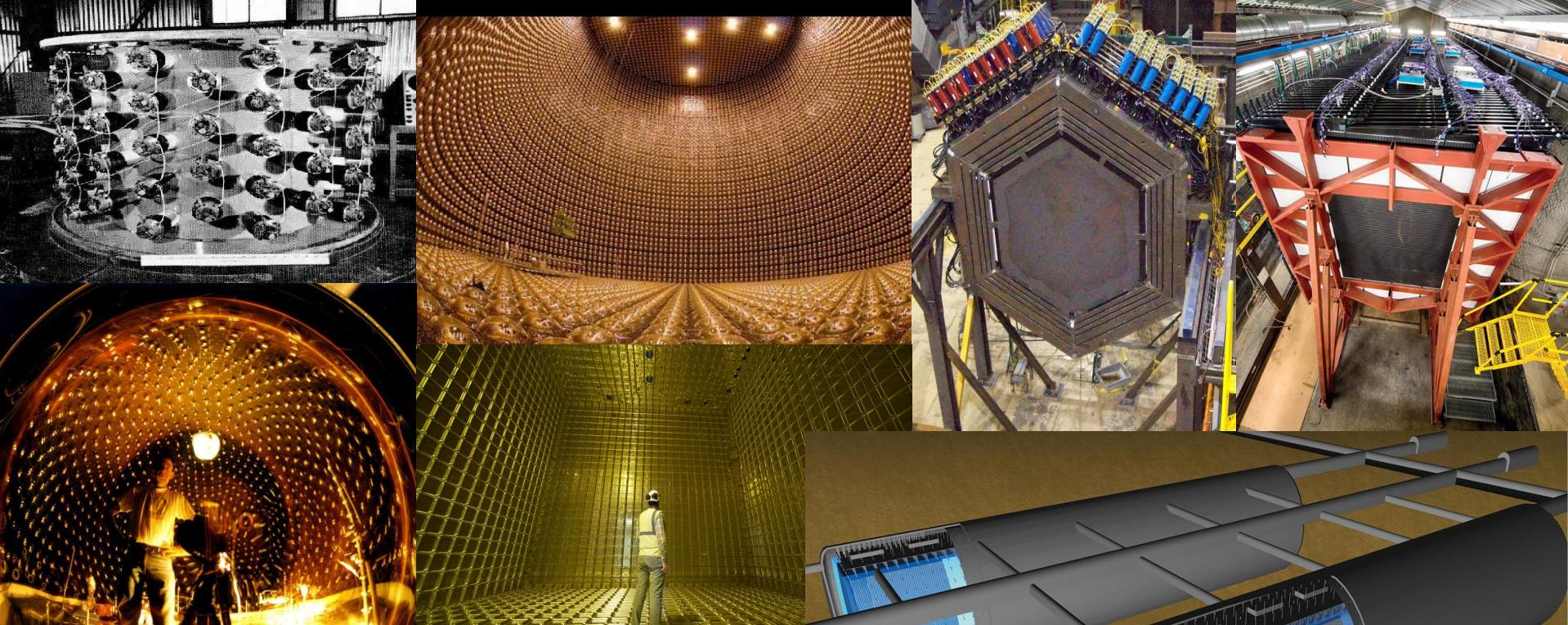
Cd-doped water  
**0.4 ton**, 100 PMTs  
(1956)



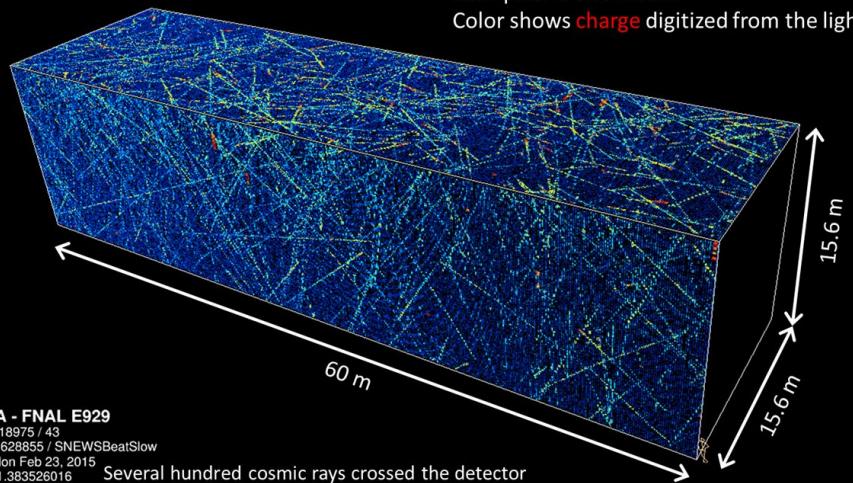
Inverse Beta Decay (IBD)

$\bar{\nu}_e + p \rightarrow e^+ + n$   
by Reines & Cowan (Nobel Prize 1995)

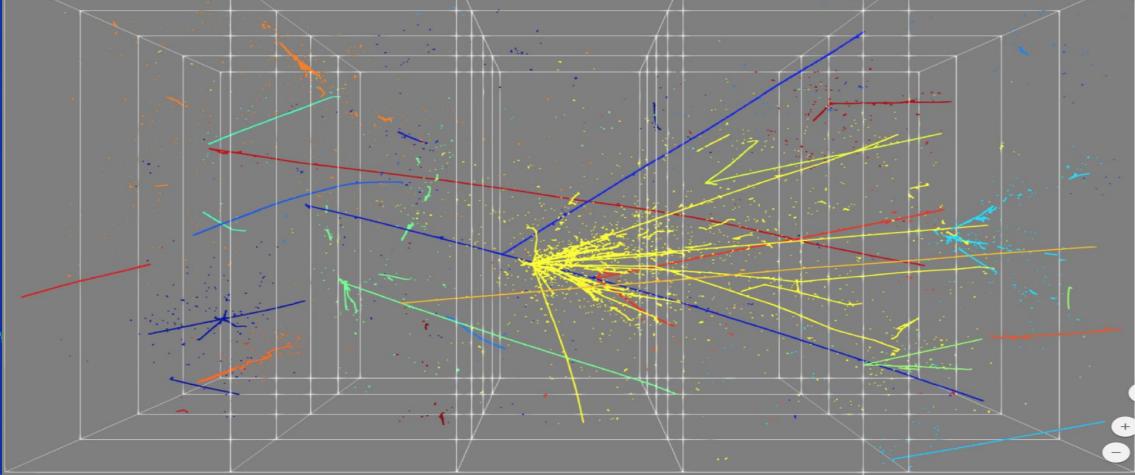
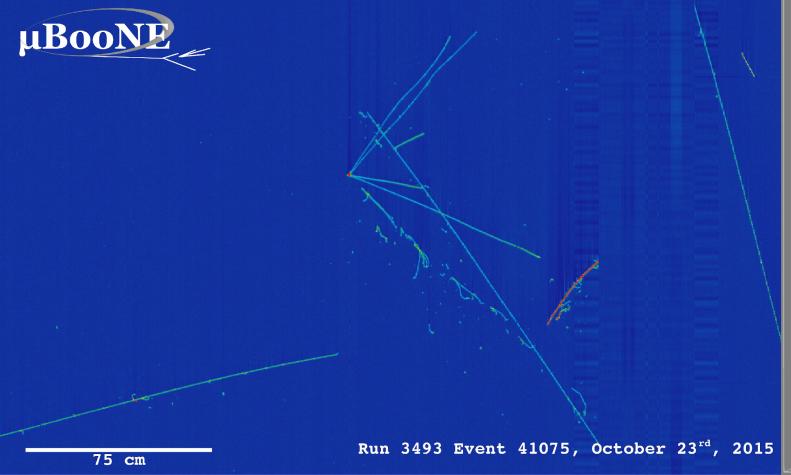
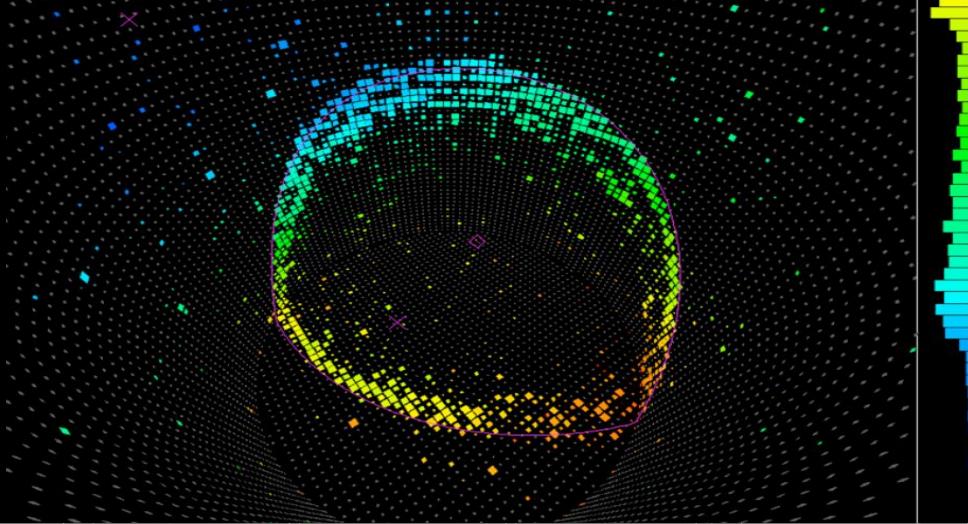
**First neutrino detection**



5ms of data at the NOvA Far Detector  
Each pixel is one hit cell  
Color shows charge digitized from the light

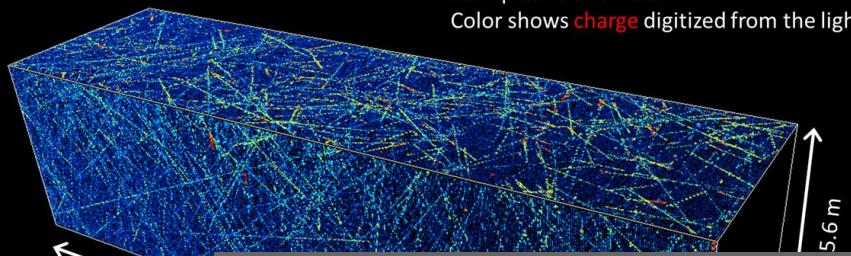


## A 603MeV muon in Super-K.

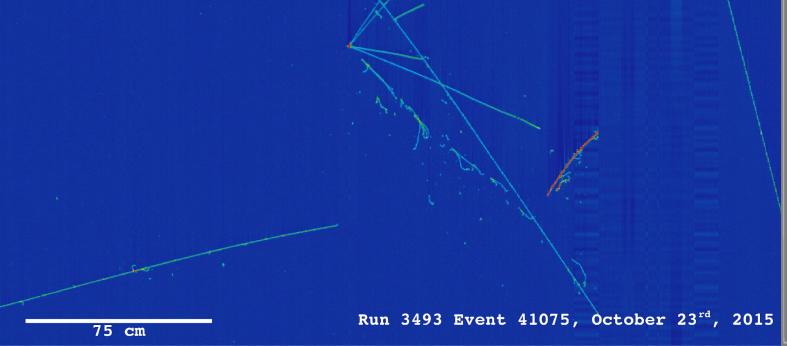


5ms of data at the NO<sub>ν</sub>A Far Detector  
Each pixel is one hit cell  
Color shows charge digitized from the light

## A 603MeV muon in Super-K.



**Need for advanced algorithms**  
for analyzing **high resolution data** with  
**complex topologies**.  
**(goal:** maximize physics output)



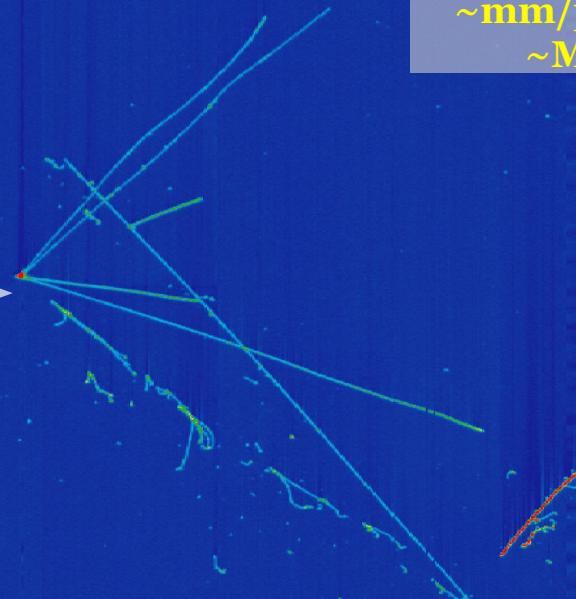
# Machine Learning & Computer Vision in Neutrino Physics

## Time Projection Chambers

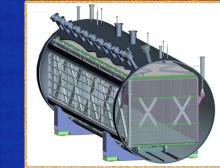
SLAC



$\nu_\mu$  - - - - - →



~mm/pixel spatial resolution  
~MeV level sensitivity



MicroBooNE  
~87 ton (school bus size)

## Liquid Argon Time Projection Chamber

- High resolution photograph of charged particle trajectories
- Calorimetric measurement + scalability to a large mass

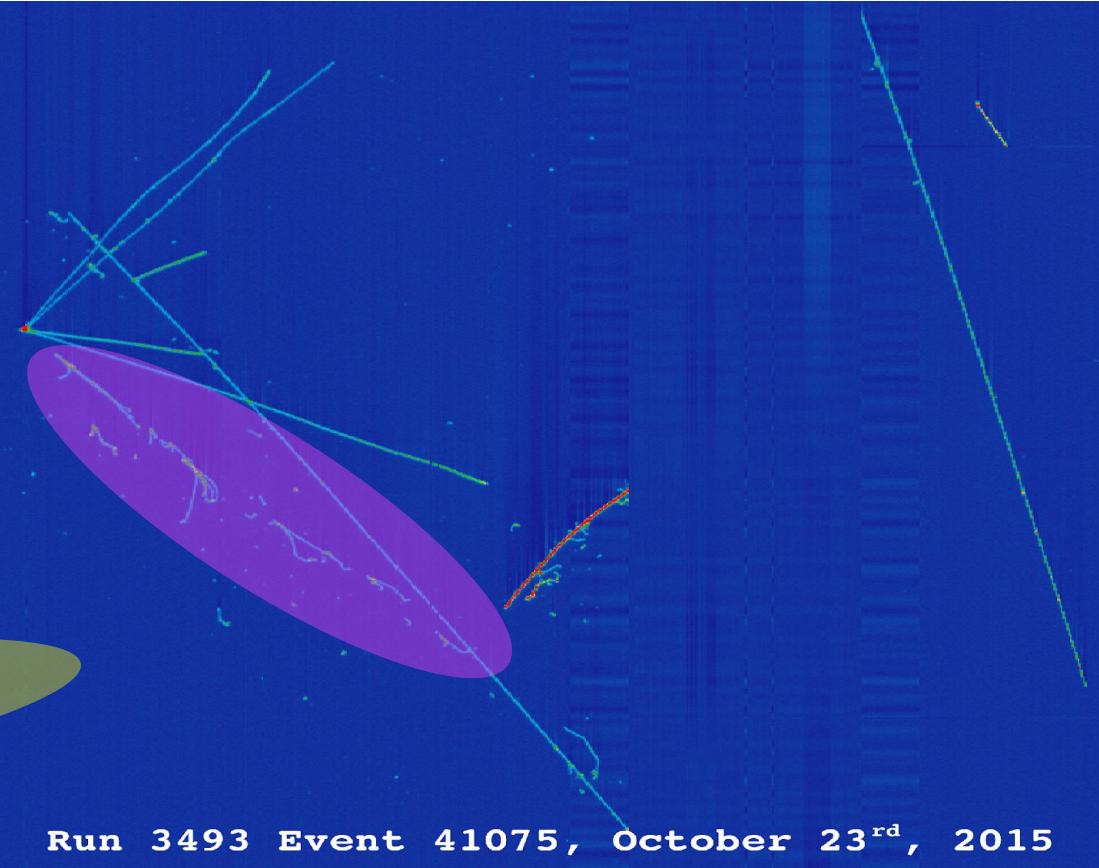
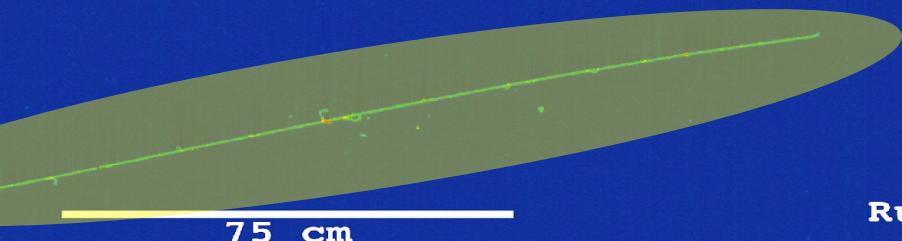
# Machine Learning & Computer Vision in Neutrino Physics

## Time Projection Chambers

SLAC



**Topological shape**  
difference is a major  
distinction for “shower”  
particles



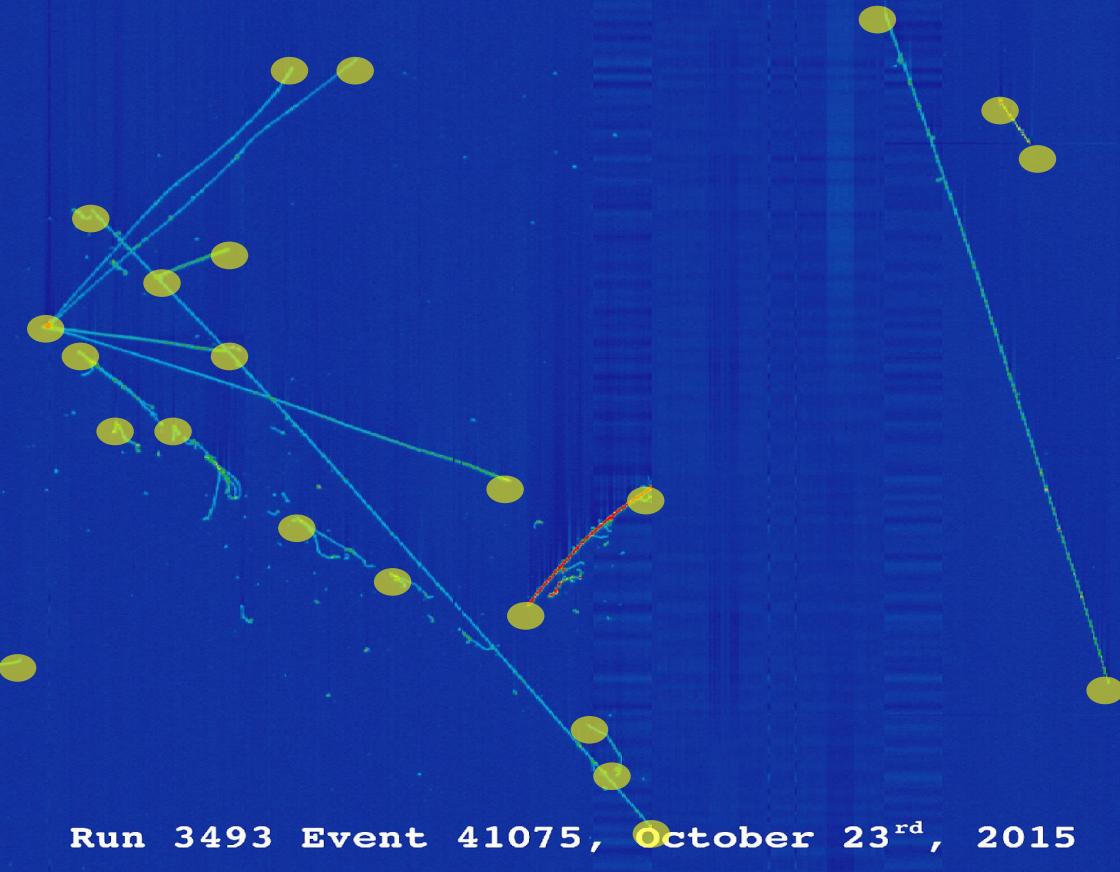
# Machine Learning & Computer Vision in Neutrino Physics

## Time Projection Chambers

SLAC



**Trajectory ends** are distinct, and useful for seeding particle clustering and trajectory fitting



# Machine Learning & Computer Vision in Neutrino Physics

## Time Projection Chambers

SLAC



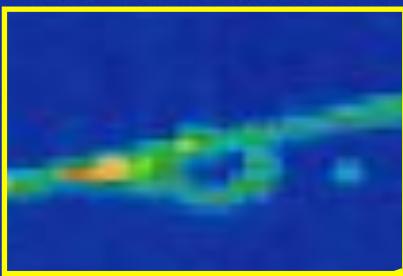
75 cm

# Machine Learning & Computer Vision in Neutrino Physics

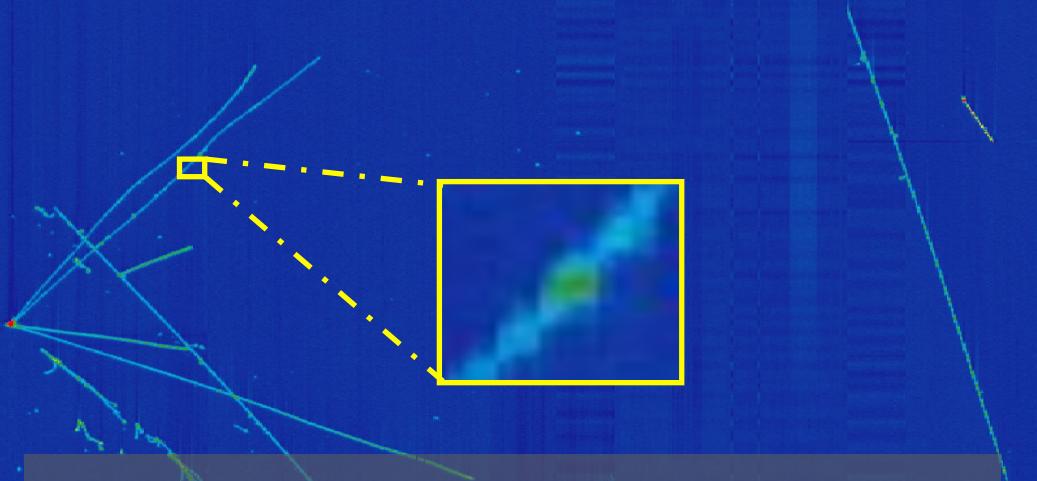
## Time Projection Chambers

SLAC

**μBooNE**



75 cm



**Small branches** on muon-like trajectories are knocked-off electrons, useful key for the direction

Run 3493 Event 41075, October 23<sup>rd</sup>, 2015

# Machine Learning & Computer Vision in Neutrino Physics

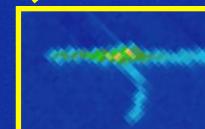
## Time Projection Chambers

SLAC

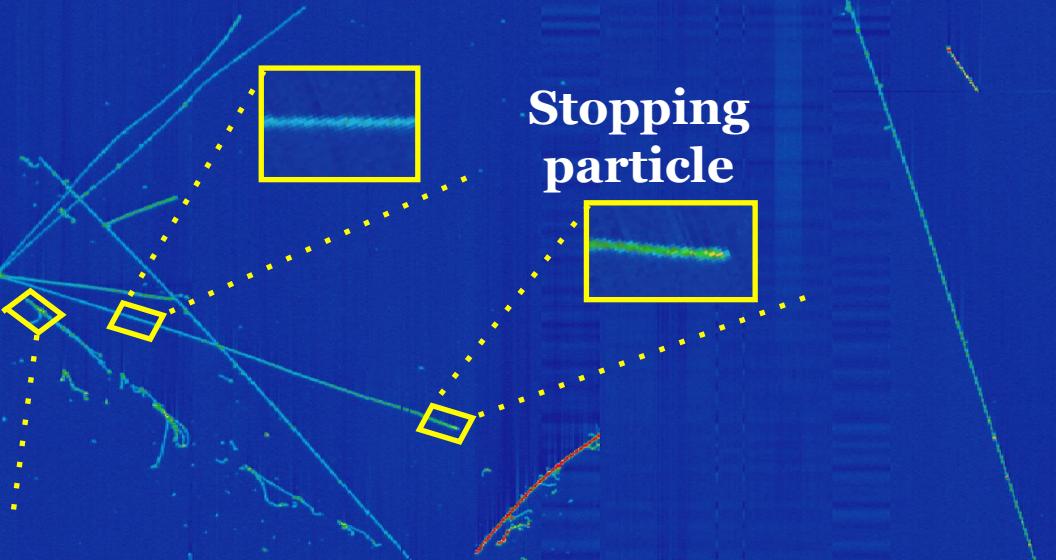
$\mu$ BooNE

**Energy deposition  
patterns ( $dE/dX$ )**

vary with particle mass  
& momentum, useful  
for analysis

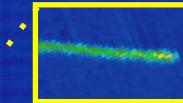


75 cm



e- vs.  $\gamma$   
using  $dE/dX$

Stopping  
particle



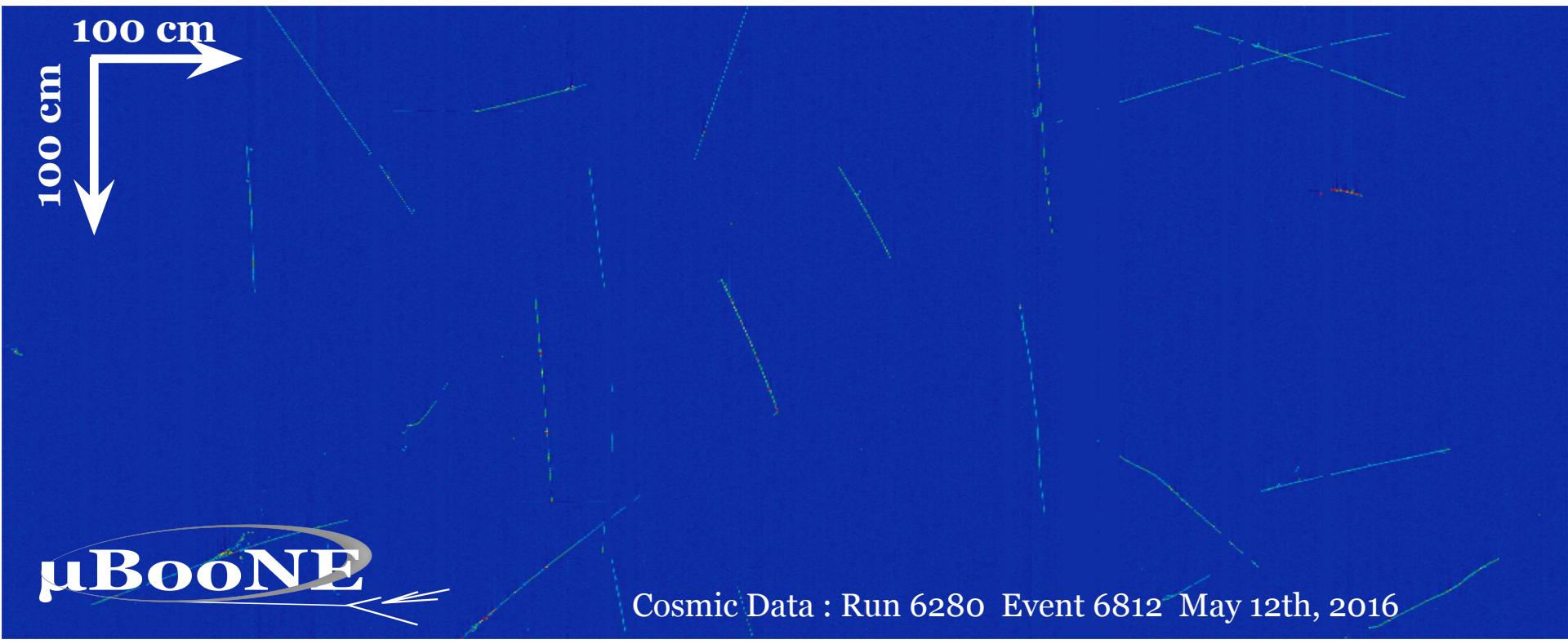
Run 3493 Event 41075, October 23<sup>rd</sup>, 2015

# Machine Learning & Computer Vision in Neutrino Physics

## Time Projection Chambers

SLAC

Do you see neutrino interaction here?

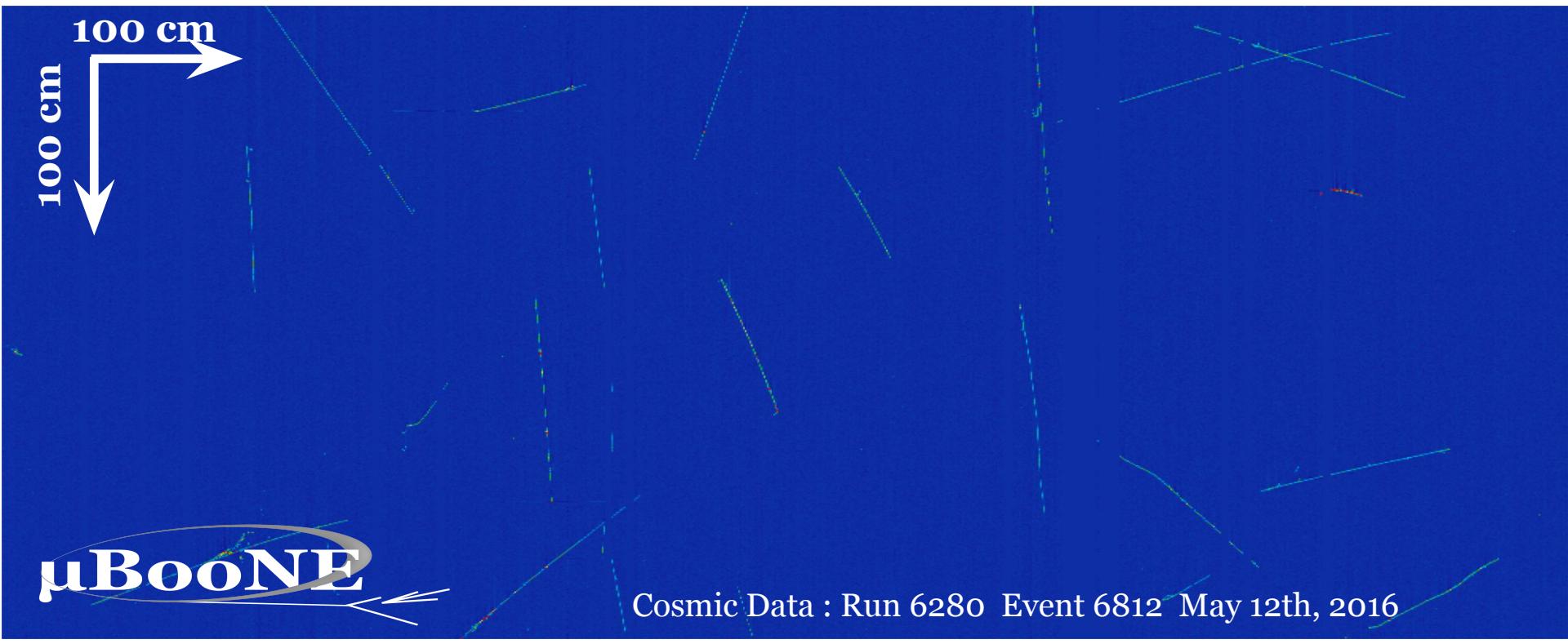


# Machine Learning & Computer Vision in Neutrino Physics

## Time Projection Chambers

SLAC

Nope :) In this detector, only  $\sim 1/700$  beam neutrino interacts

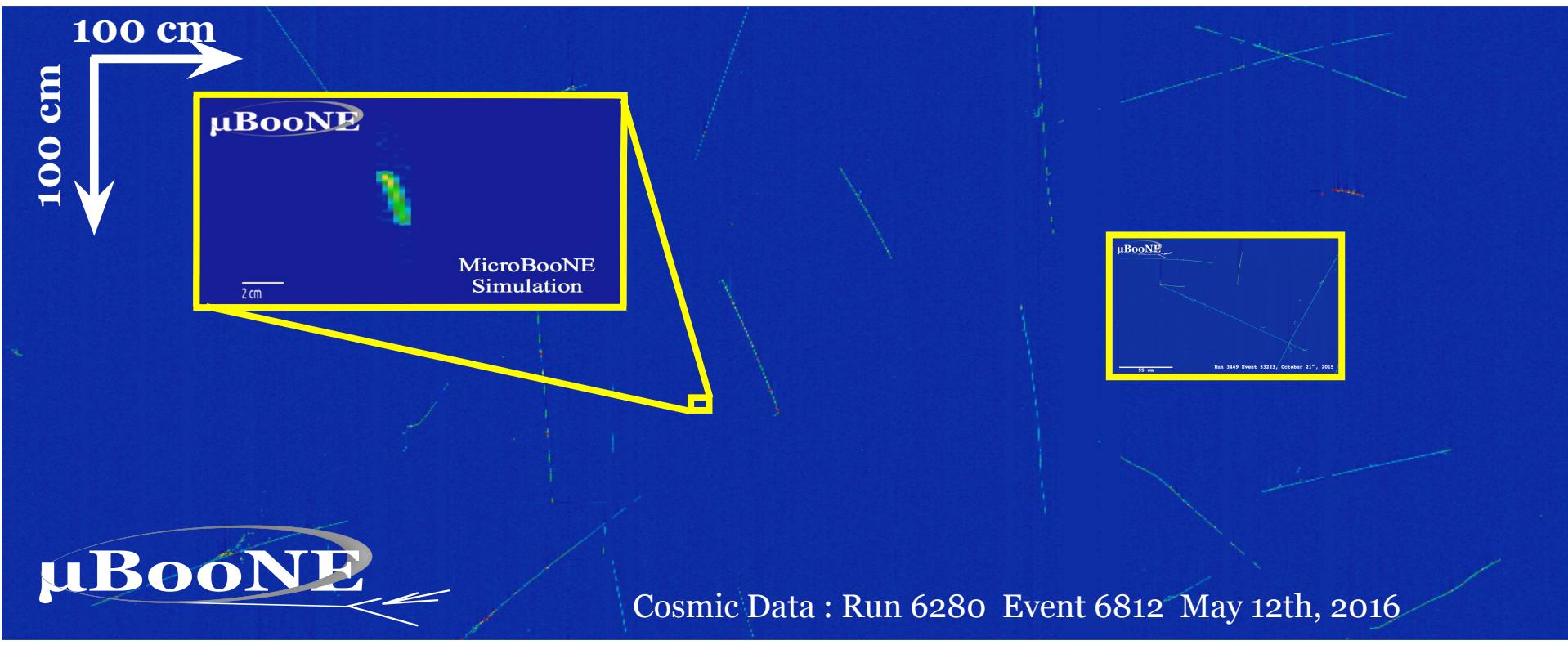


# Machine Learning & Computer Vision in Neutrino Physics

## Time Projection Chambers

SLAC

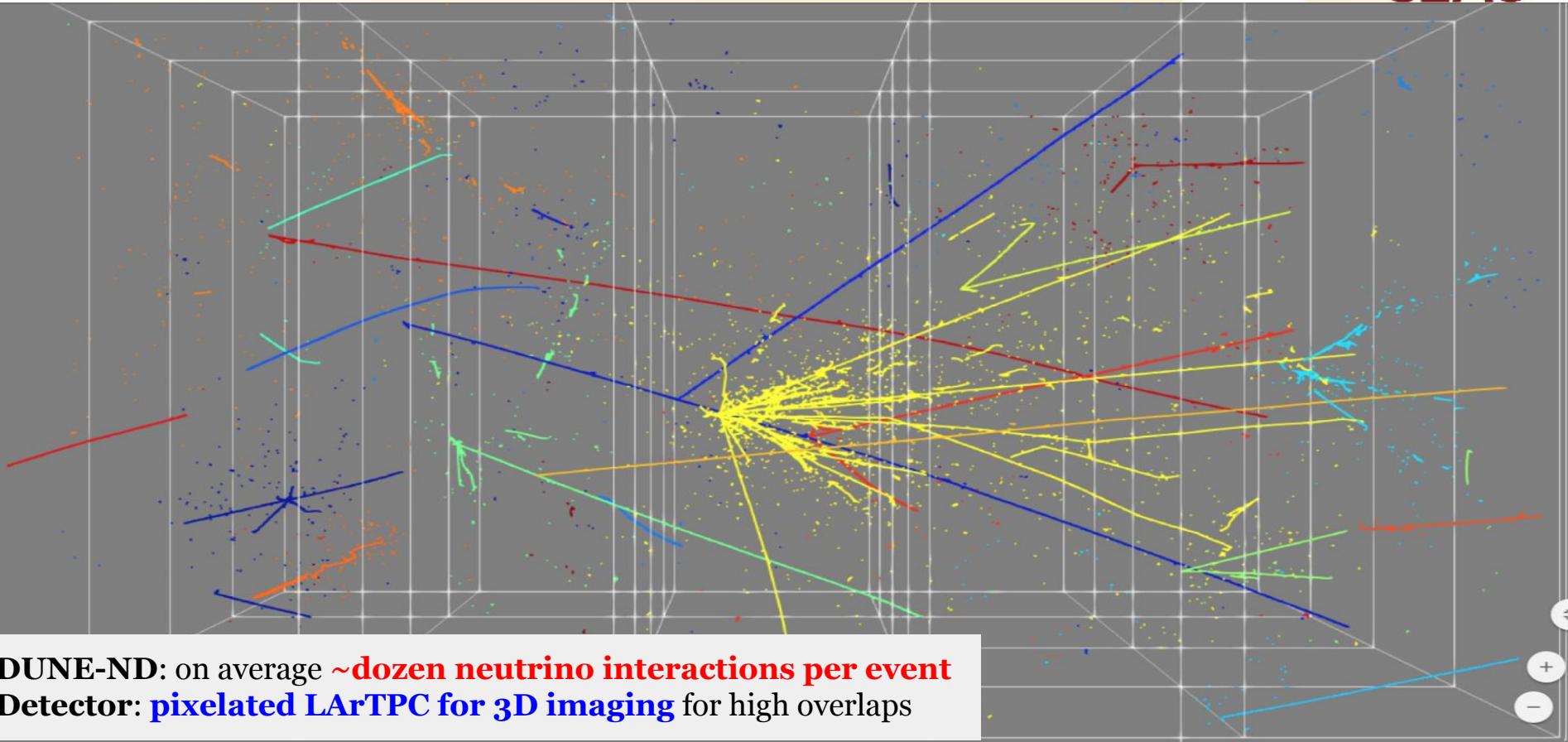
... and 1/700 have many variations in hi-resolution imaging...



# Machine Learning & Computer Vision in Neutrino Physics

## Time Projection Chambers (3D ones)

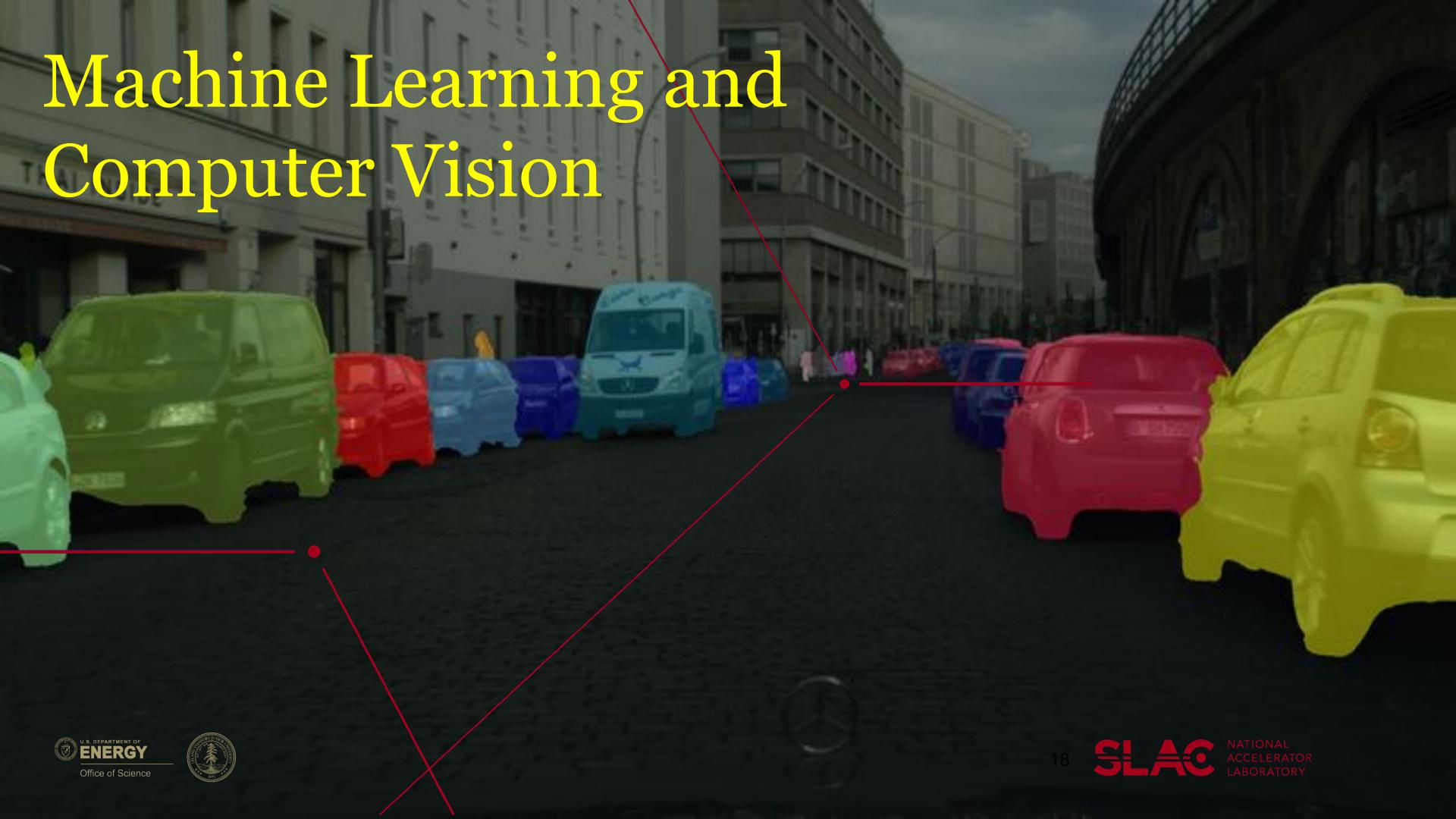
SLAC



DUNE-ND: on average ~dozen neutrino interactions per event

Detector: pixelated LArTPC for 3D imaging for high overlaps

# Machine Learning and Computer Vision





How to write an algorithm to  
identify a cat?

... very hard task ...

16	08	67	15	83	09
37	52	77	23	22	74
35	42	48	72	85	27
68	36	43	54	21	33
79	60	10	25	54	71
18	55	38	73	50	47

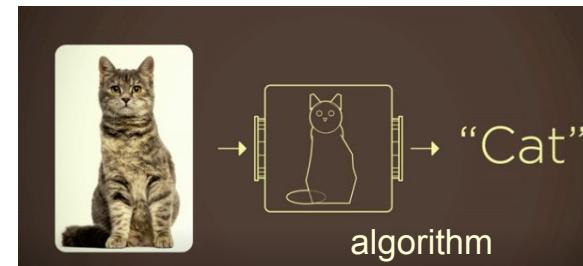
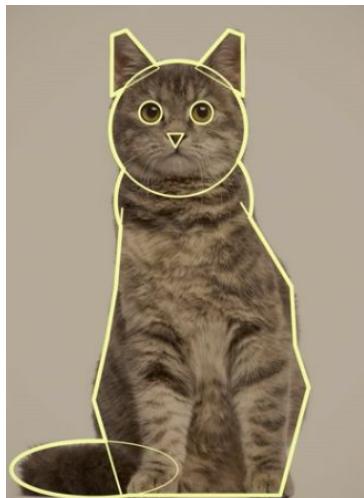
# Machine Learning & Computer Vision in Neutrino Physics

You can find a cat? You can find a neutrino!

SLAC

## Development Workflow for non-ML reconstruction

1. Write an algorithm based on physics principles



A cat = collection of  
(or, a neutrino) certain shapes

### Development Workflow for non-ML reconstruction

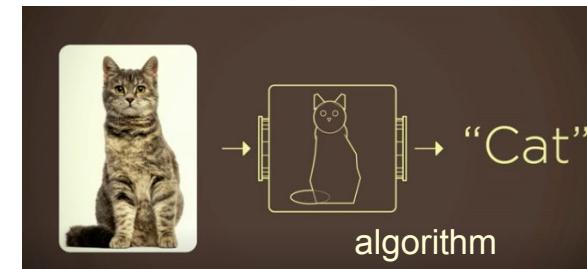
1. Write an algorithm based on physics principles
2. Run on simulation and data samples
3. Observe failure cases, implement fixes/heuristics
4. Iterate over 2 & 3 till a satisfactory level is achieved
5. Chain multiple algorithms as one algorithm, repeat 2, 3, and 4.



Partial cat  
(escaping the detector)



Stretching cat (Nuclear FSI)



A cat = collection of  
(or, a neutrino) certain shapes

### Development Workflow for non-ML reconstruction

1. Write an algorithm based on physics principles
2. Run on simulation and data samples
3. Observe failure cases, implement fixes/heuristics
4. Iterate over 2 & 3 till a satisfactory level is achieved
5. Chain multiple algorithms as one algorithm, repeat 2, 3, and 4.

#### “Machine learning”

- Automatization of step 2, 3, and 4.
- Well-defined error propagation (step 5).
- Can optimize the whole chain for physics.

**Next:** what kind of ML algorithms?

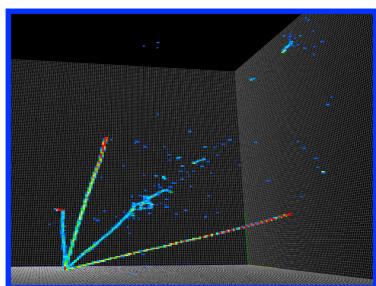
# Machine Learning & Computer Vision in Neutrino Physics

## My Research

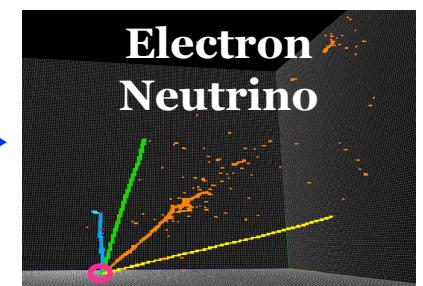
SLAC

### Machine Learning for Data Reconstruction

- **Goal:** high level abstract information (like image classification)



Input Data



High-level  
Output

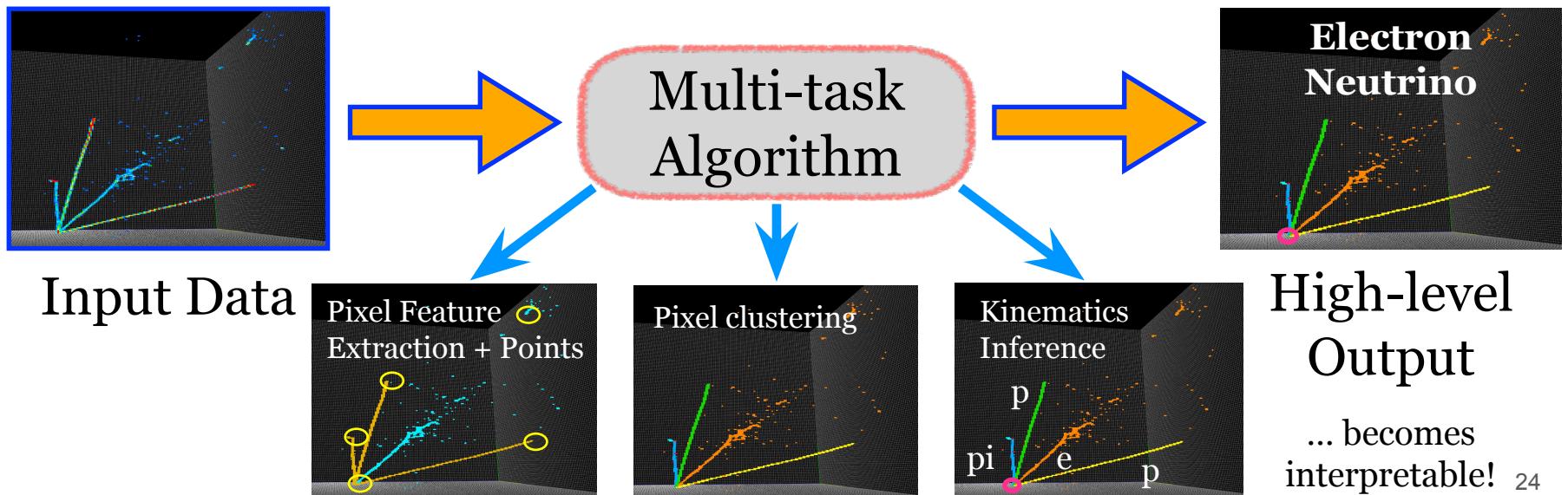
# Machine Learning & Computer Vision in Neutrino Physics

## My Research

SLAC

### Machine Learning for Data Reconstruction

- **Goal:** high level abstract information (like image classification)
- **How:** design the algorithm = data transformation architecture that extracts a hierarchy of physically meaningful features (evidences)



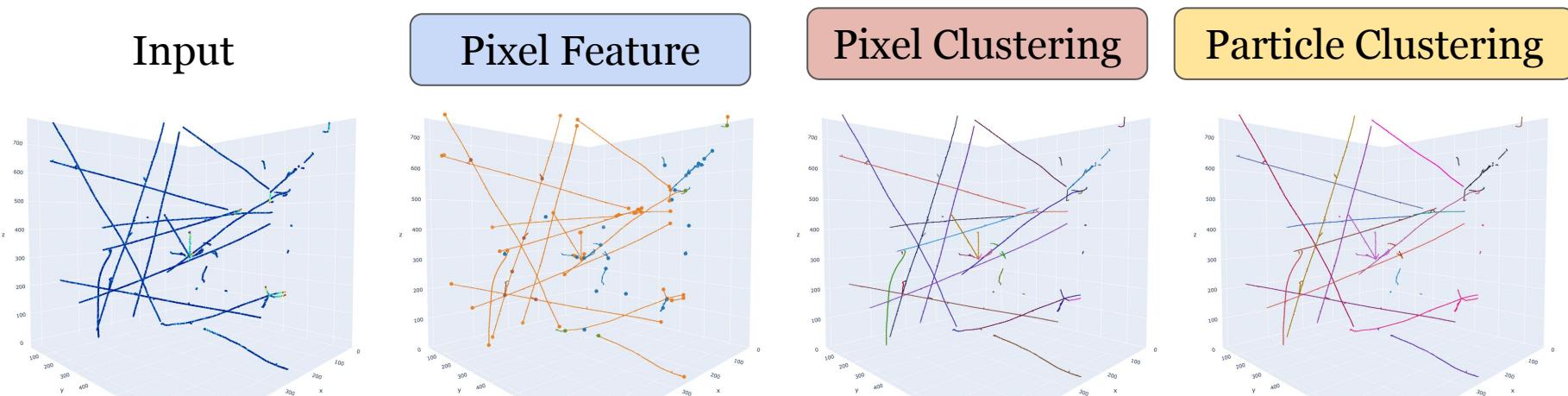
# Machine Learning & Computer Vision in Neutrino Physics

## My Research

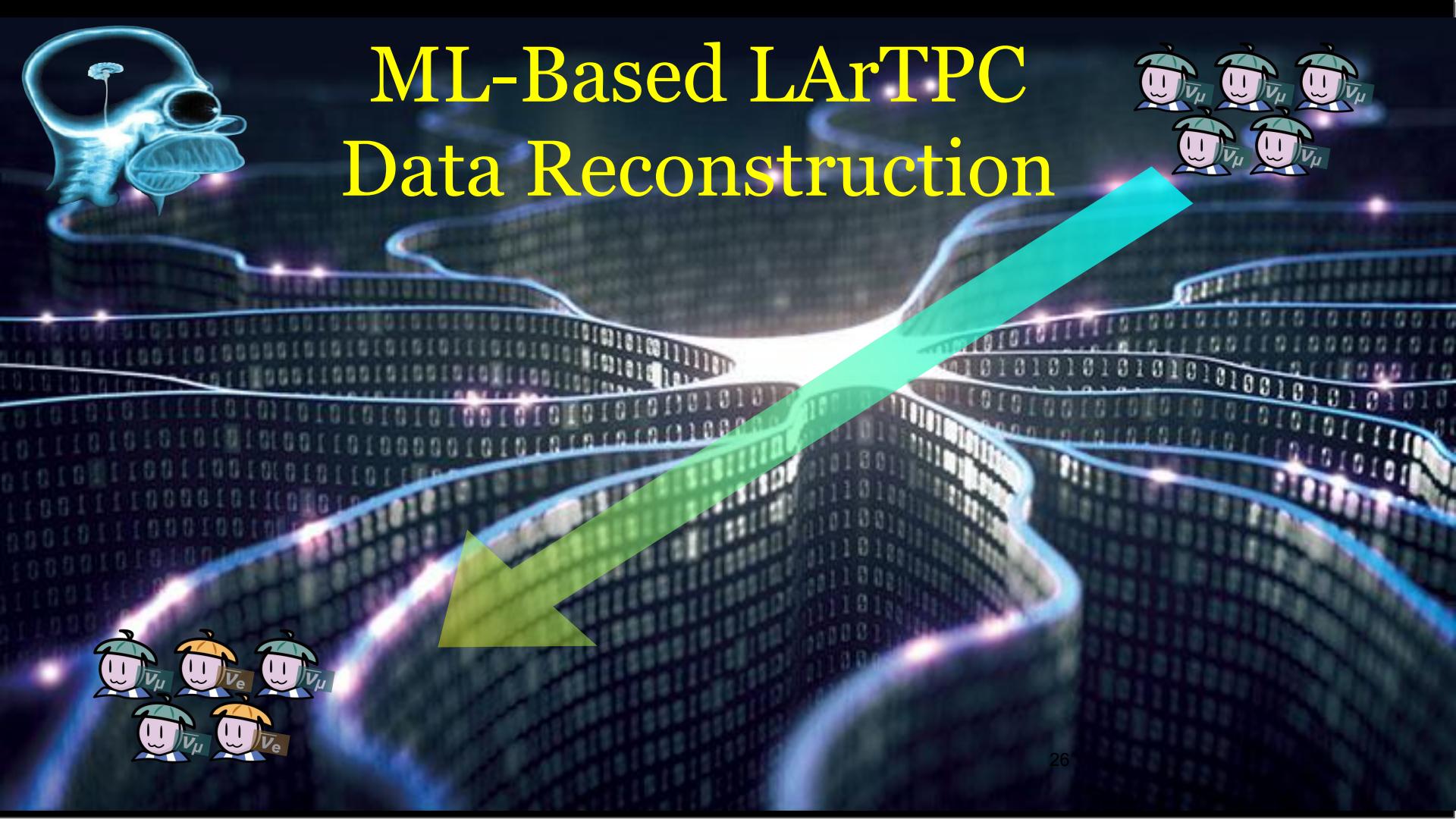
SLAC

### Machine Learning for Data Reconstruction

- **Goal:** high level abstract information (like image classification)
- **How:** design the algorithm = data transformation architecture that extracts a hierarchy of physically meaningful features



The Rest: describe the chain



# ML-Based LArTPC Data Reconstruction



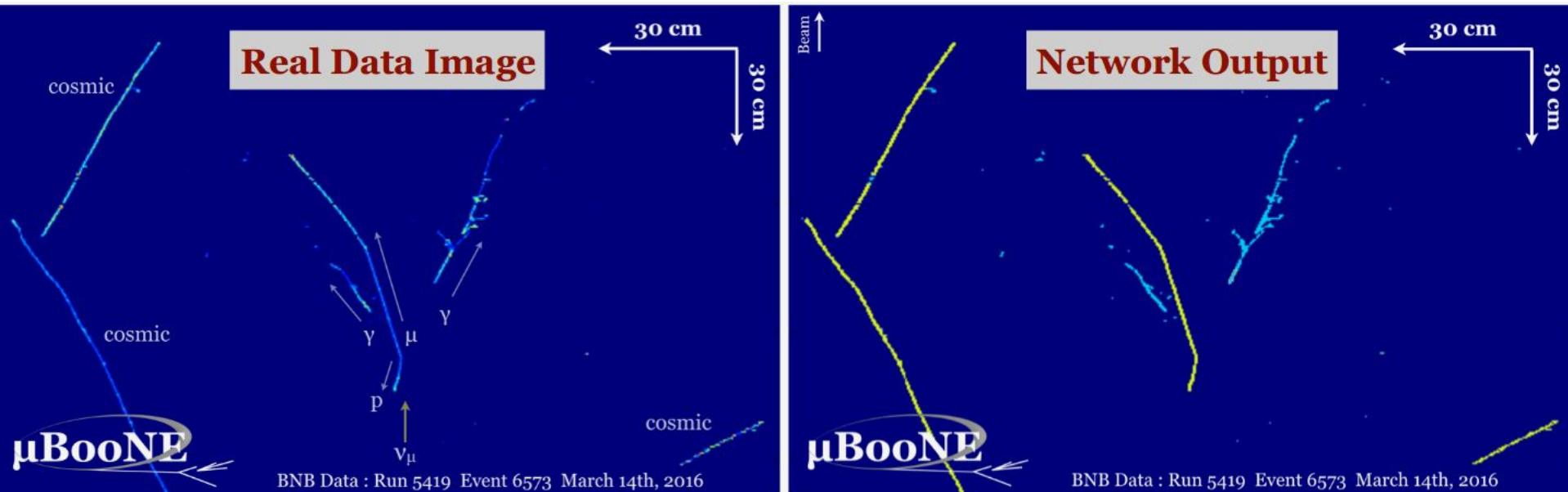
# ML-based Neutrino Data Reconstruction Chain

## Stage 1-a: Pixel Feature Extraction + Scalability

SLAC

Distinguish 2 distinct particle topologies: **showers v.s. tracks**

Critical to deploy different algorithms for clustering pixels in the next stage.



**Network Input**

[PRD 99 092001](#)  
[arXiv:1808.07269](#)

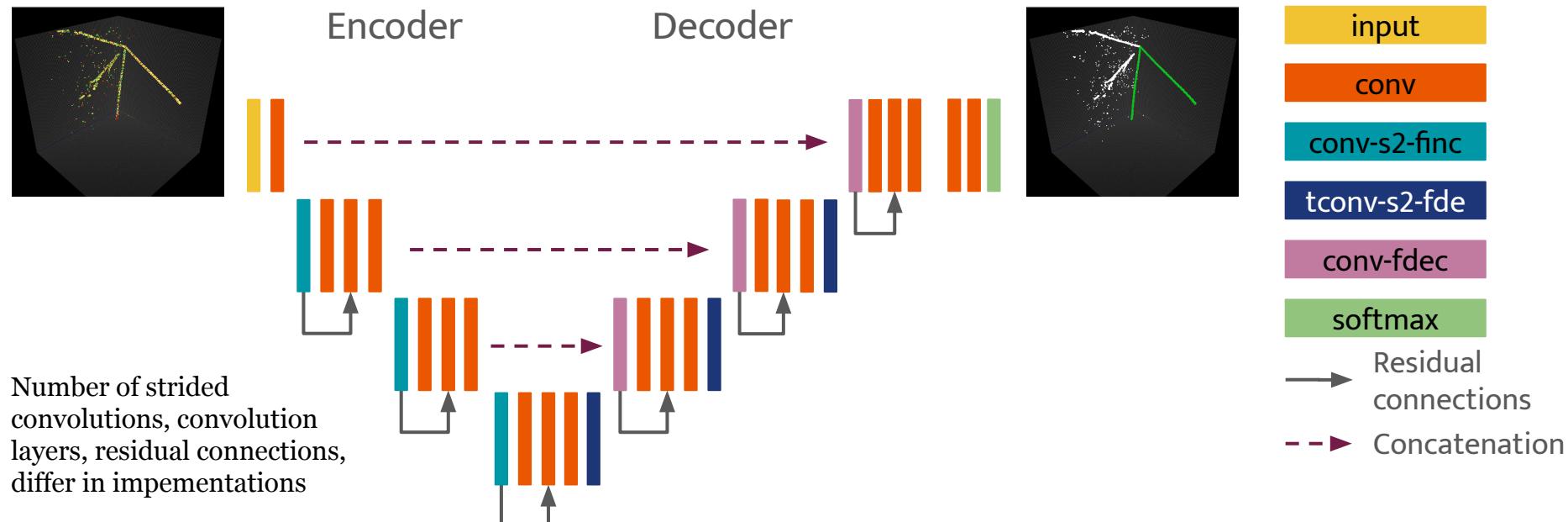
**Network Output**

# ML-based Neutrino Data Reconstruction Chain

## Stage 1-a: Pixel Feature Extraction + Scalability

SLAC

### Architecture: U-Net + Residual Connections



## Stage 1-a: Pixel Feature Extraction + Scalability

“Applying CNN” is simple, but **is it scalable for us?**

CNN applies  
**dense matrix**  
**operations**

In photographs,  
**all pixels are**  
**meaningful**



grey pixels = dolphins,  
blue pixels = water, etc...

# ML-based Neutrino Data Reconstruction Chain

## Stage 1-a: Pixel Feature Extraction + Scalability

SLAC

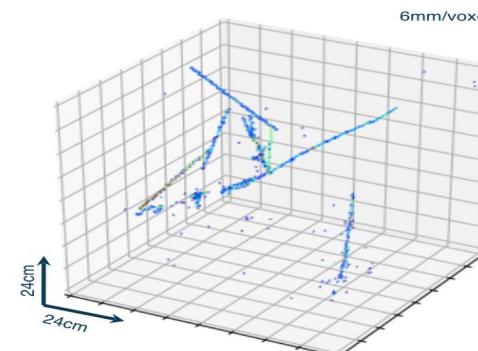
**“Applying CNN” is simple, but **is it scalable for us?**  
LArTPC data is generally sparse, but locally dense**

CNN applies  
**dense matrix**  
**operations**

In photographs,  
**all pixels are**  
**meaningful**



grey pixels = dolphins,  
blue pixels = water, etc...



Empty pixels = no energy

<1% of pixels  
are non-zero in  
LArTPC data

**Zero pixels are  
meaningless!**

Figures/Texts: courtesy of  
Laura Domine @ Stanford

# ML-based Neutrino Data Reconstruction Chain

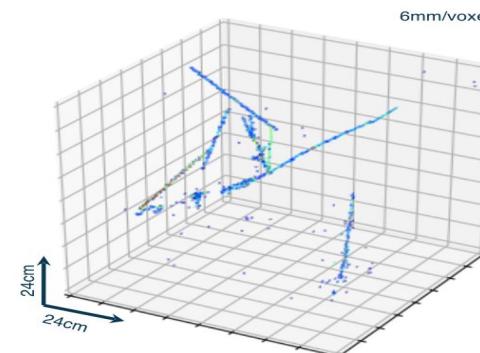
## Stage 1-a: Pixel Feature Extraction + Scalability

SLAC

**“Applying CNN” is simple, but **is it scalable for us?**  
LArTPC data is generally sparse, but locally dense**

CNN applies  
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In photographs,  
**all pixels are**  
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<1% of pixels  
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LArTPC data

**Zero pixels are  
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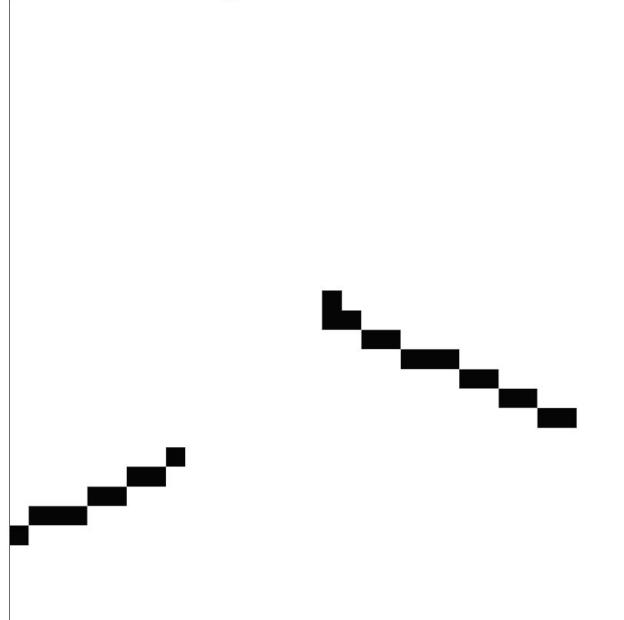
Figures/Texts: courtesy of  
Laura Domine @ Stanford

- **Scalability for larger detectors**

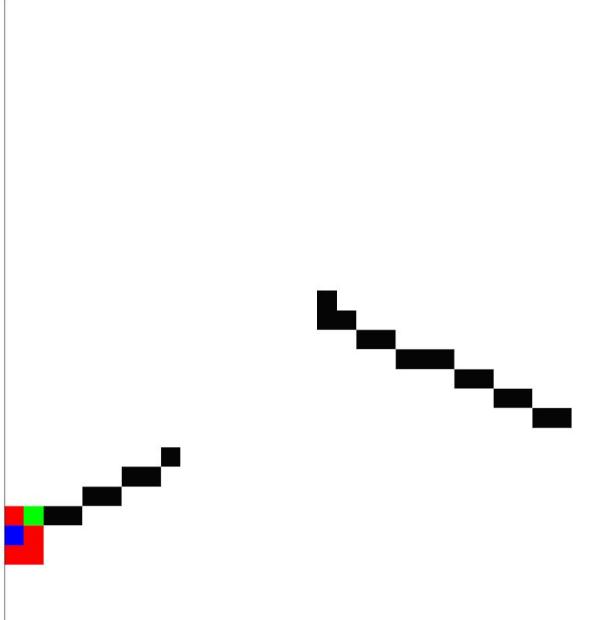
- Computation cost increases linearly with the volume
- But the number of non-zero pixels does not

### Sparse Submanifold Convolutions

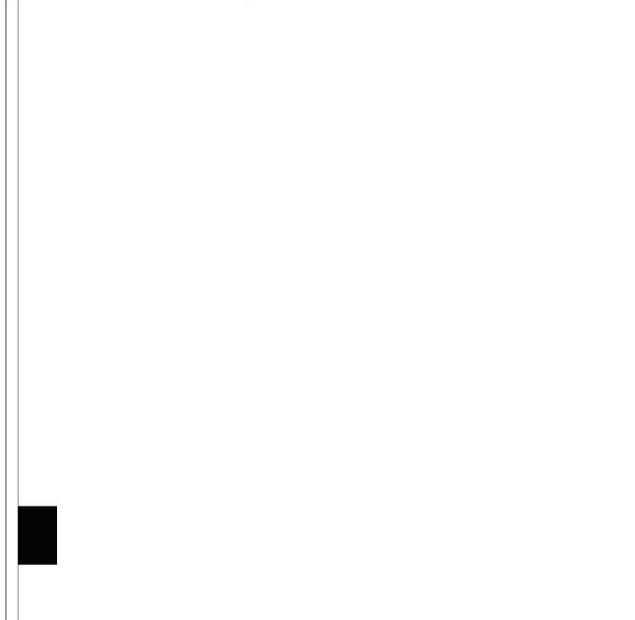
Original Activations



Filter Reach

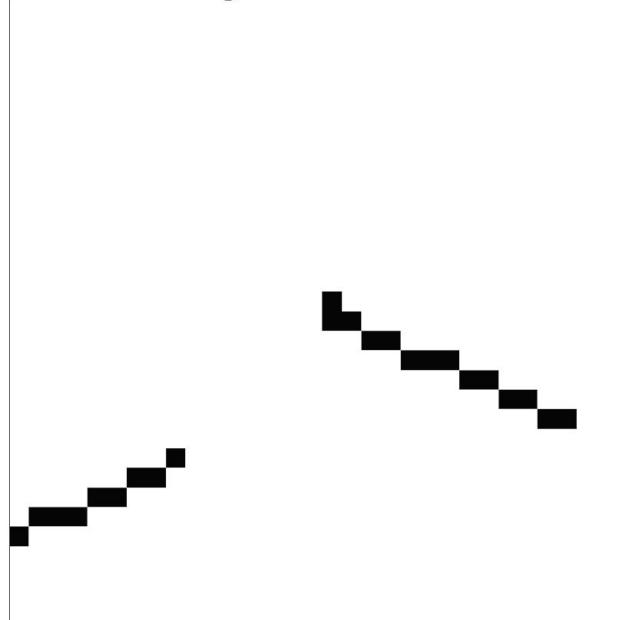


Output Activations

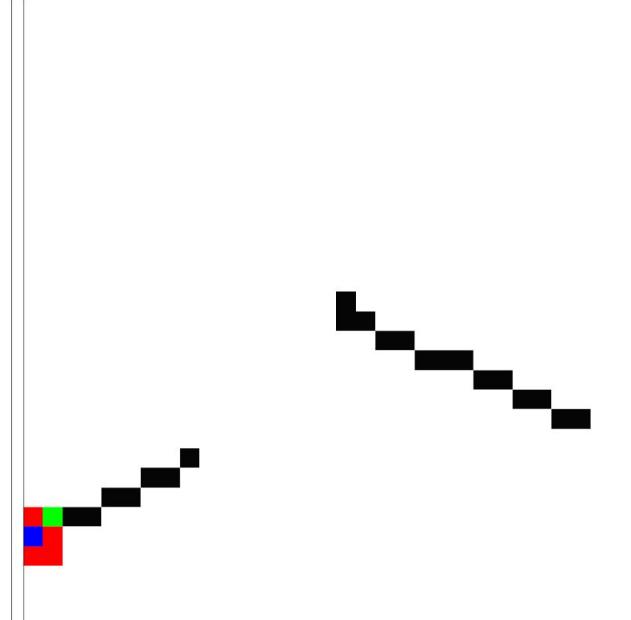


### Sparse Submanifold Convolutions

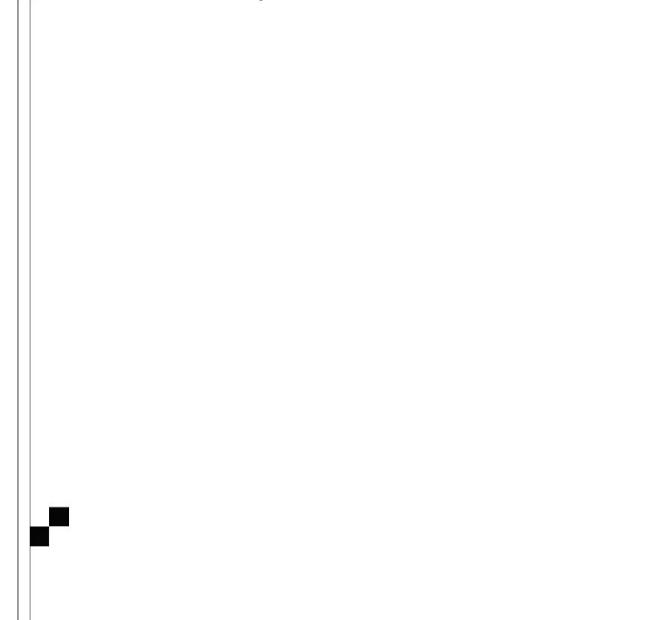
Original Activations



Filter Reach



Output Activations



# ML-based Neutrino Data Reconstruction Chain

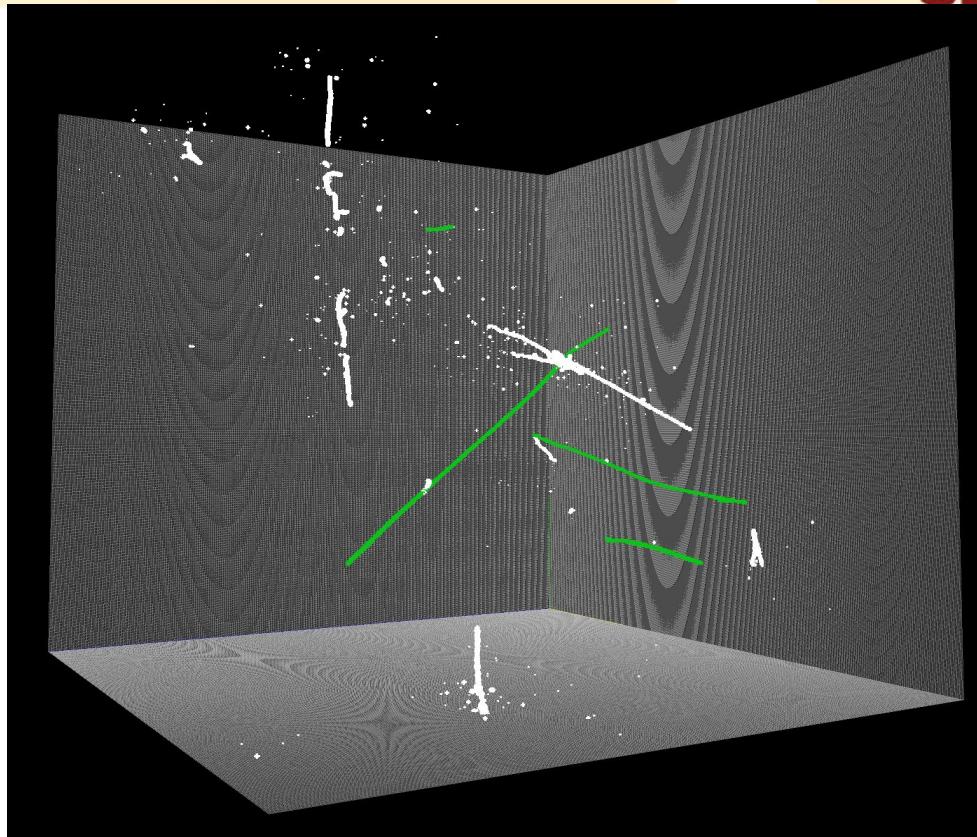
## Stage 1-a: Pixel Feature Extraction + Scalability

SLAC

Our data is locally much more dense  
than ShapeNet 3D dataset



... which makes convolution filter  
more effective on our data as long as  
the sparsity issue is handled

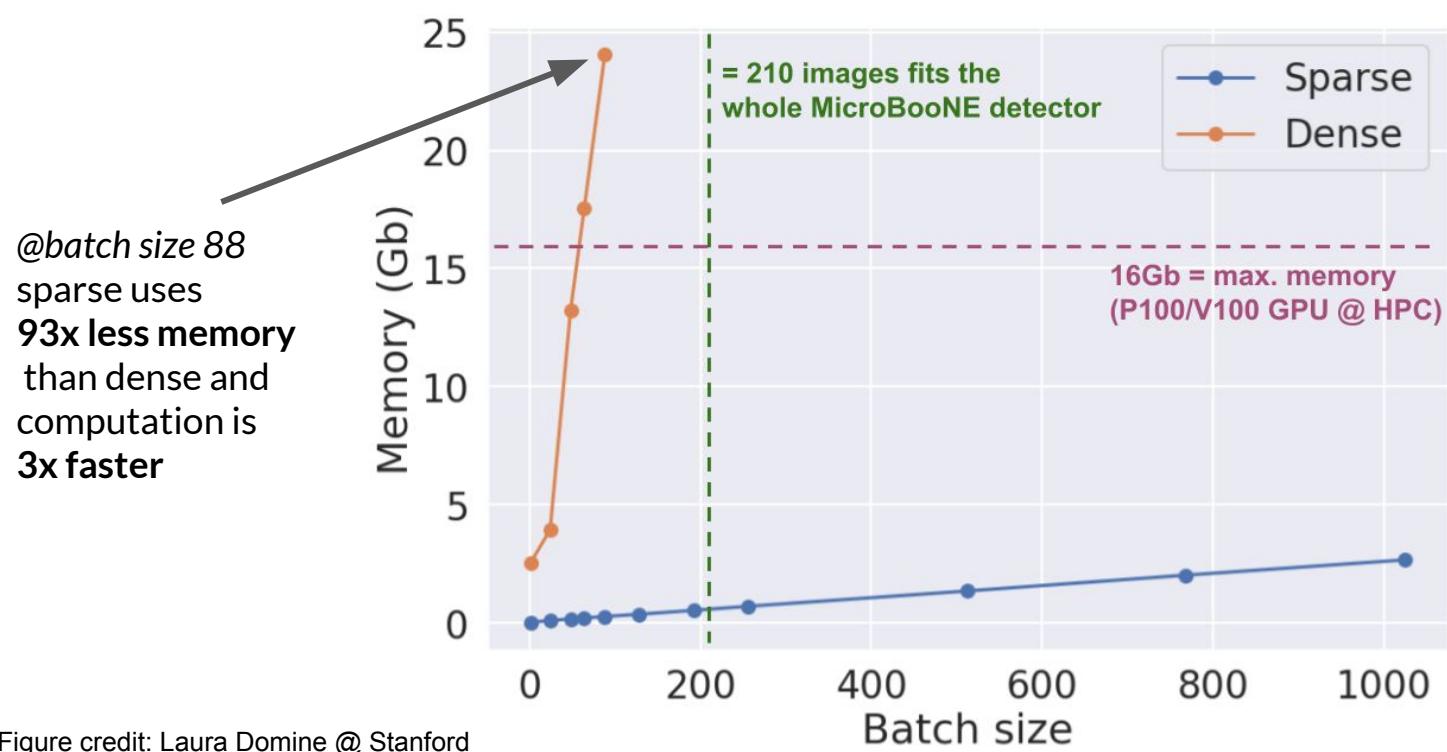


# ML-based Neutrino Data Reconstruction Chain

## Stage 1-a: Pixel Feature Extraction + Scalability

SLAC

Sparse U-ResNet fits more data in GPU + good scalability



Can handle easily the whole ICARUS detector which is x6 larger than MicroBooNE.

DUNE-FD is piece of cake (larger volume but less non-zero pixels)

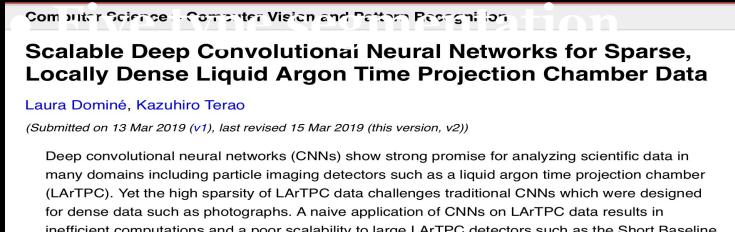
# ML-based Neutrino Data Reconstruction Chain

## Stage 1-a: Pixel Feature Extraction + Scalability

SLAC

### Sparse Sub-manifold Convolutional NN

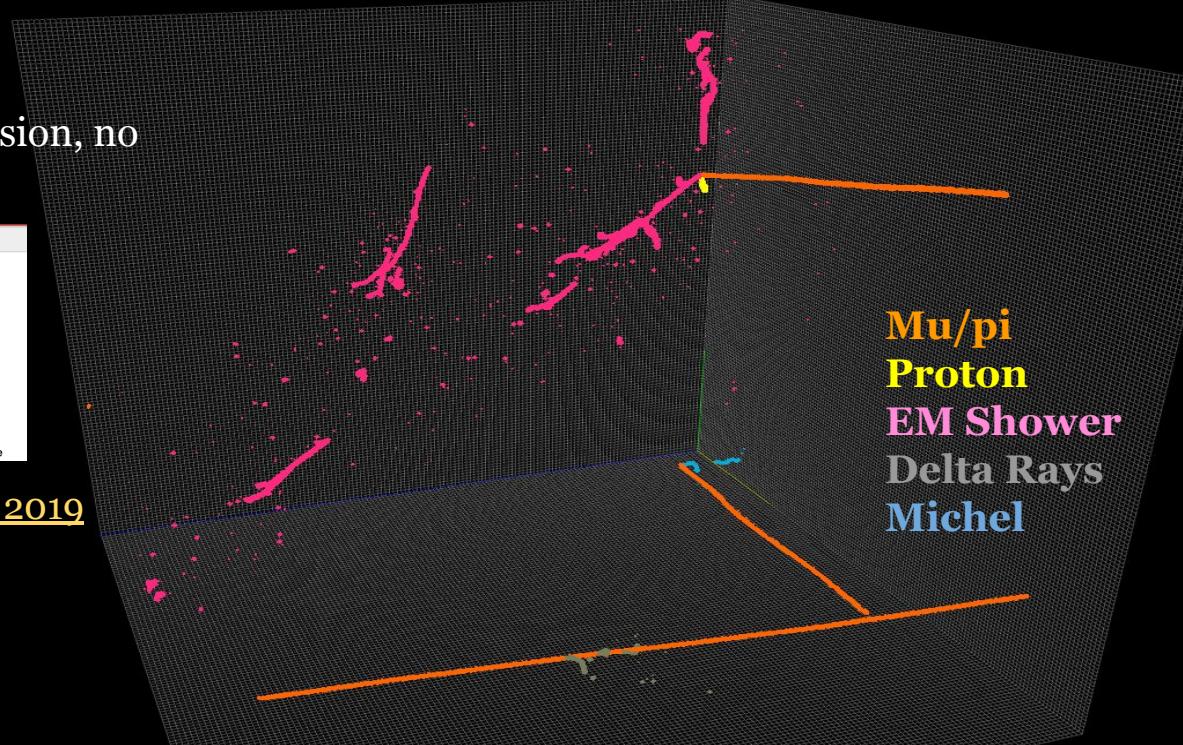
- Public LArTPC simulation
  - Particle tracking (Geant4) + diffusion, no noise, true energy



[PhysRevD.102.012005](#) presented @ [ACAT 2019](#)

- Memory reduction  $\sim 1/360$
- Compute time  $\sim 1/30$
- Handles large future detectors

Type	Proton	Mu/Pi	Shower	Delta	Michel
Acc.	0.99	0.98	0.99	0.97	0.96

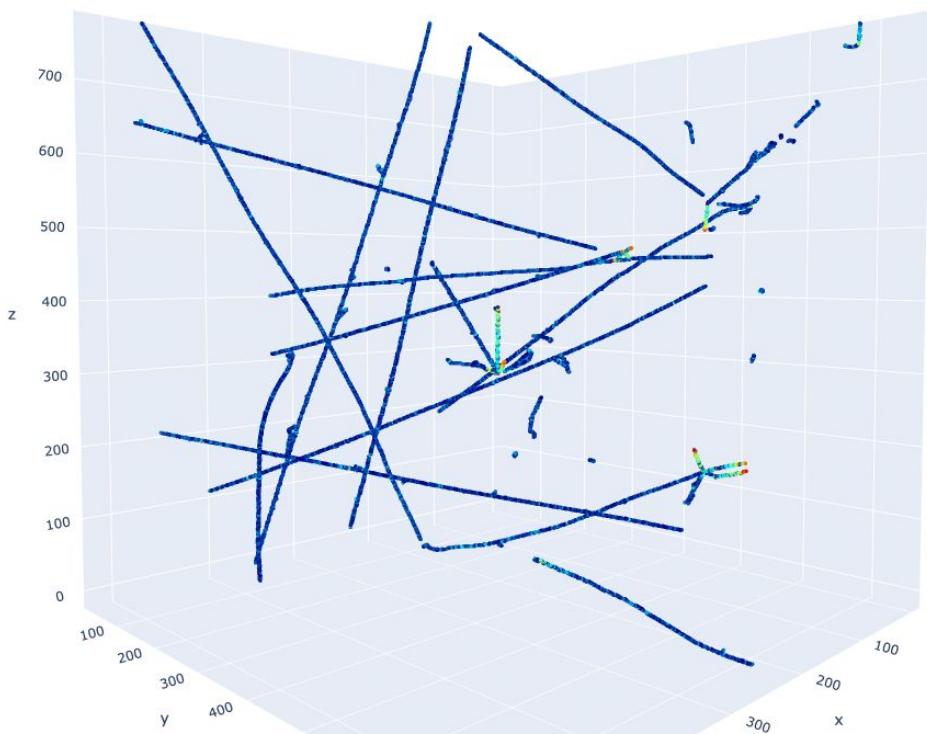


# ML-based Neutrino Data Reconstruction Chain

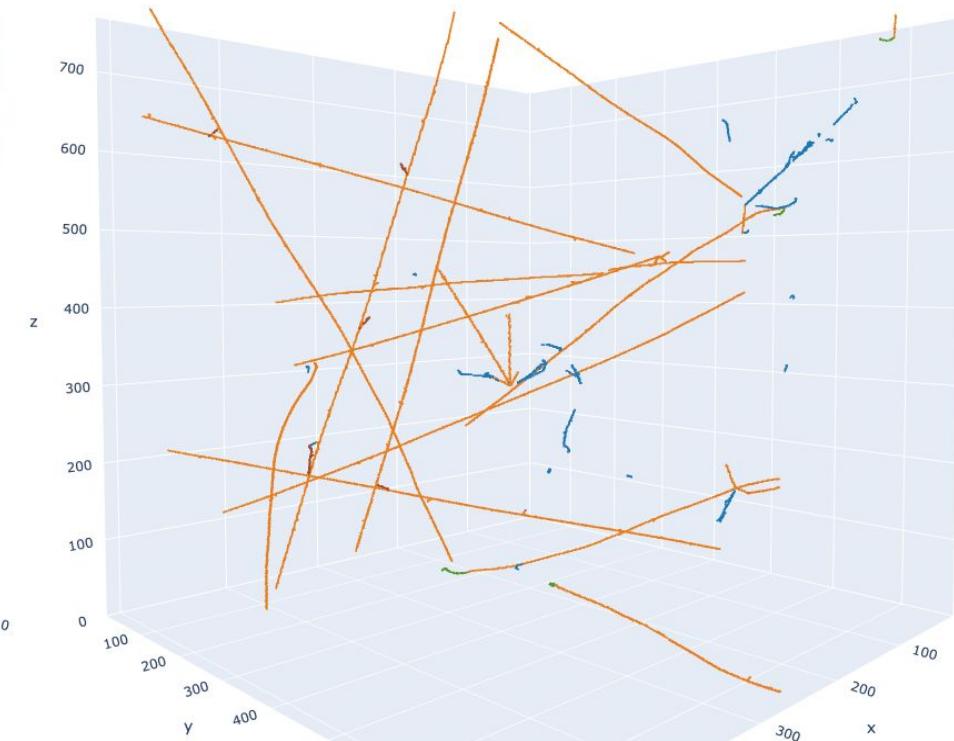
## Stage 1-a: input & output

SLAC

Stage 1-a Input



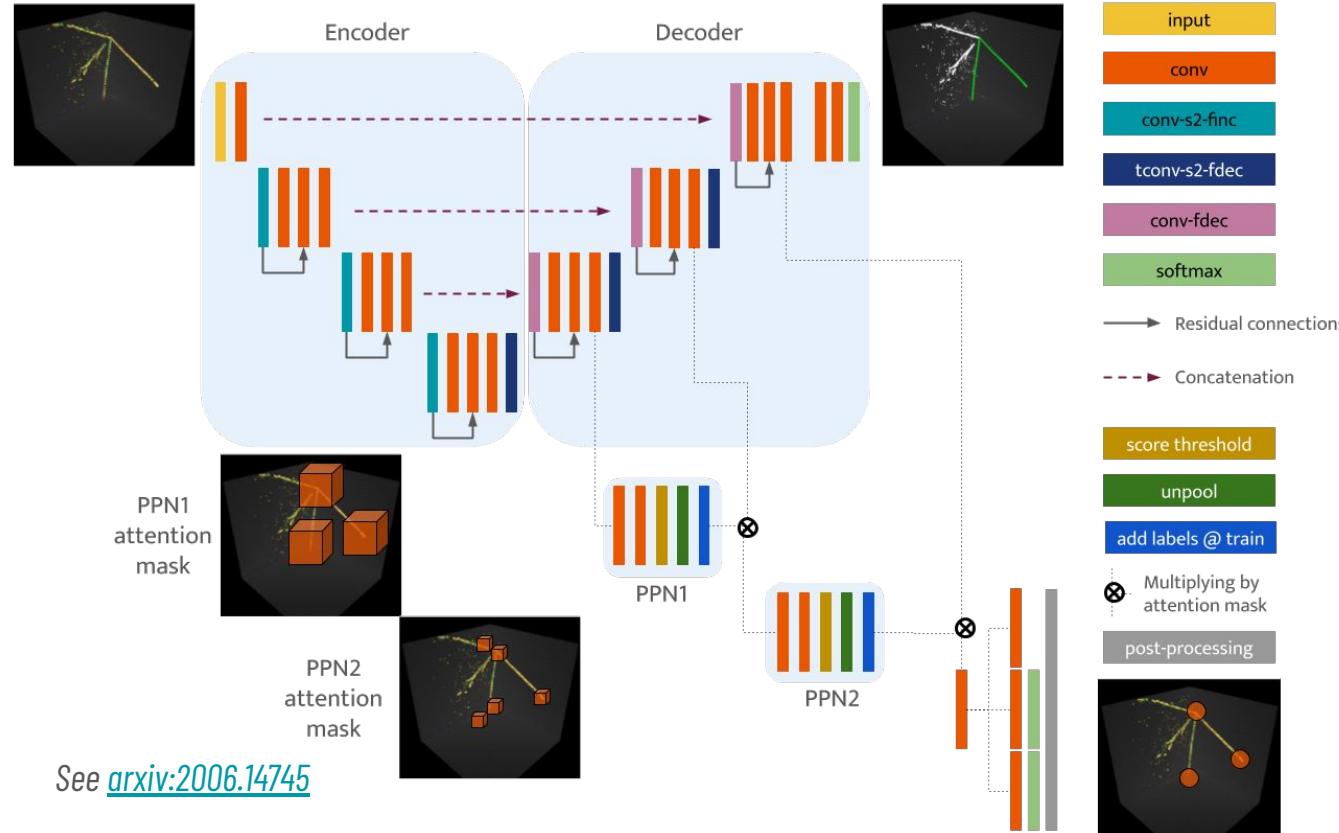
Stage 1-a Output



# ML-based Neutrino Data Reconstruction Chain

## Stage 1-b: Particle Endpoint Prediction

SLAC



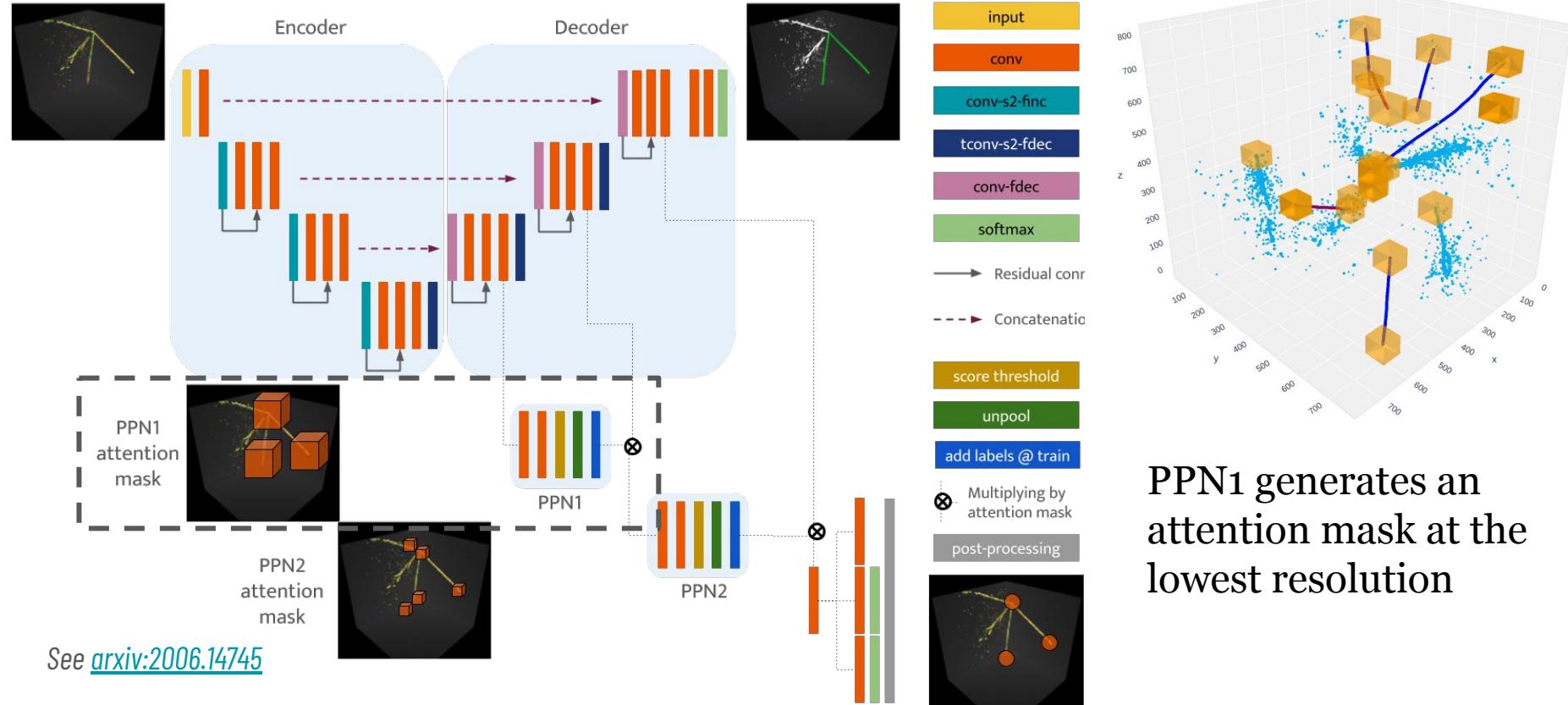
Point Proposal  
Network (PPN)

... extension of U-ResNet  
with 3 CNN blocks

# ML-based Neutrino Data Reconstruction Chain

## Stage 1-b: Particle Endpoint Prediction

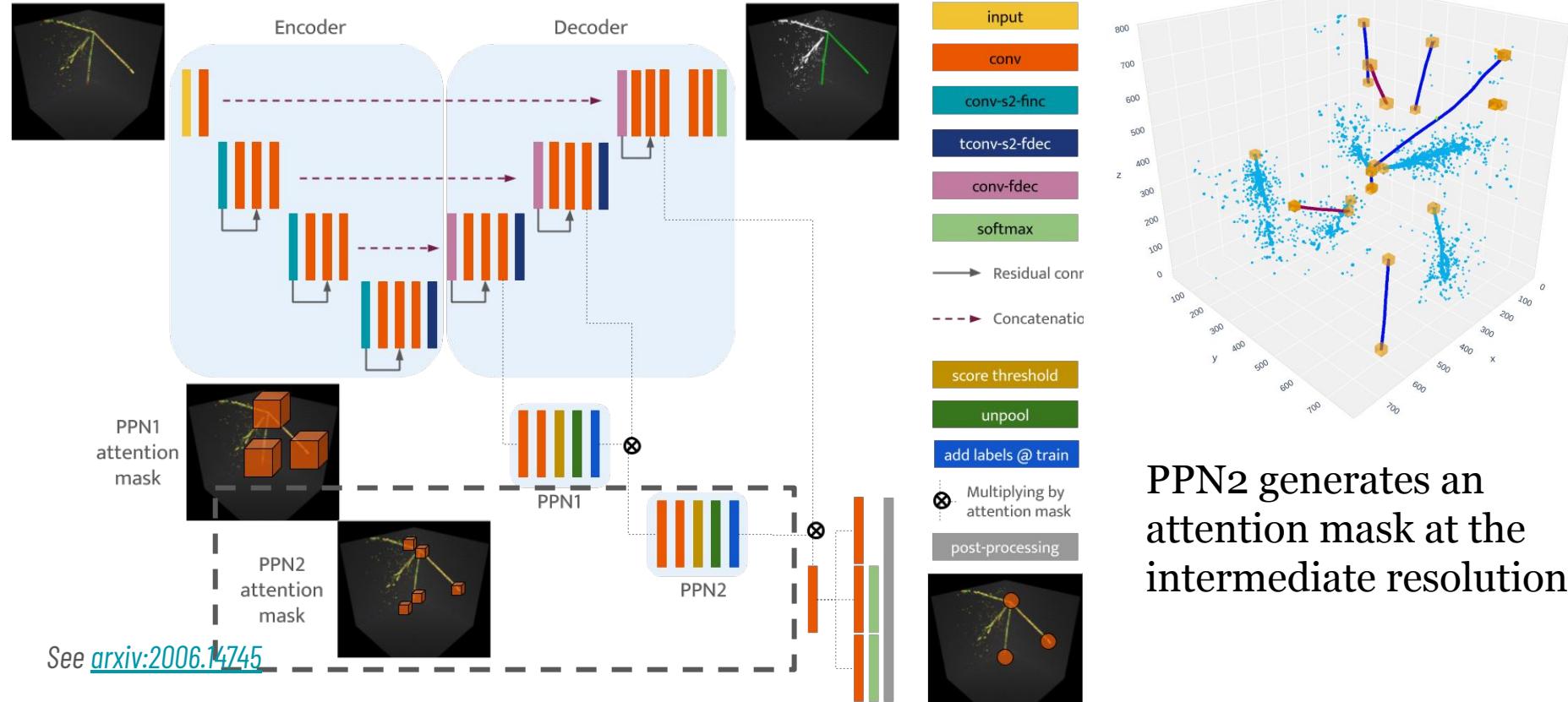
SLAC



# ML-based Neutrino Data Reconstruction Chain

## Stage 1-b: Particle Endpoint Prediction

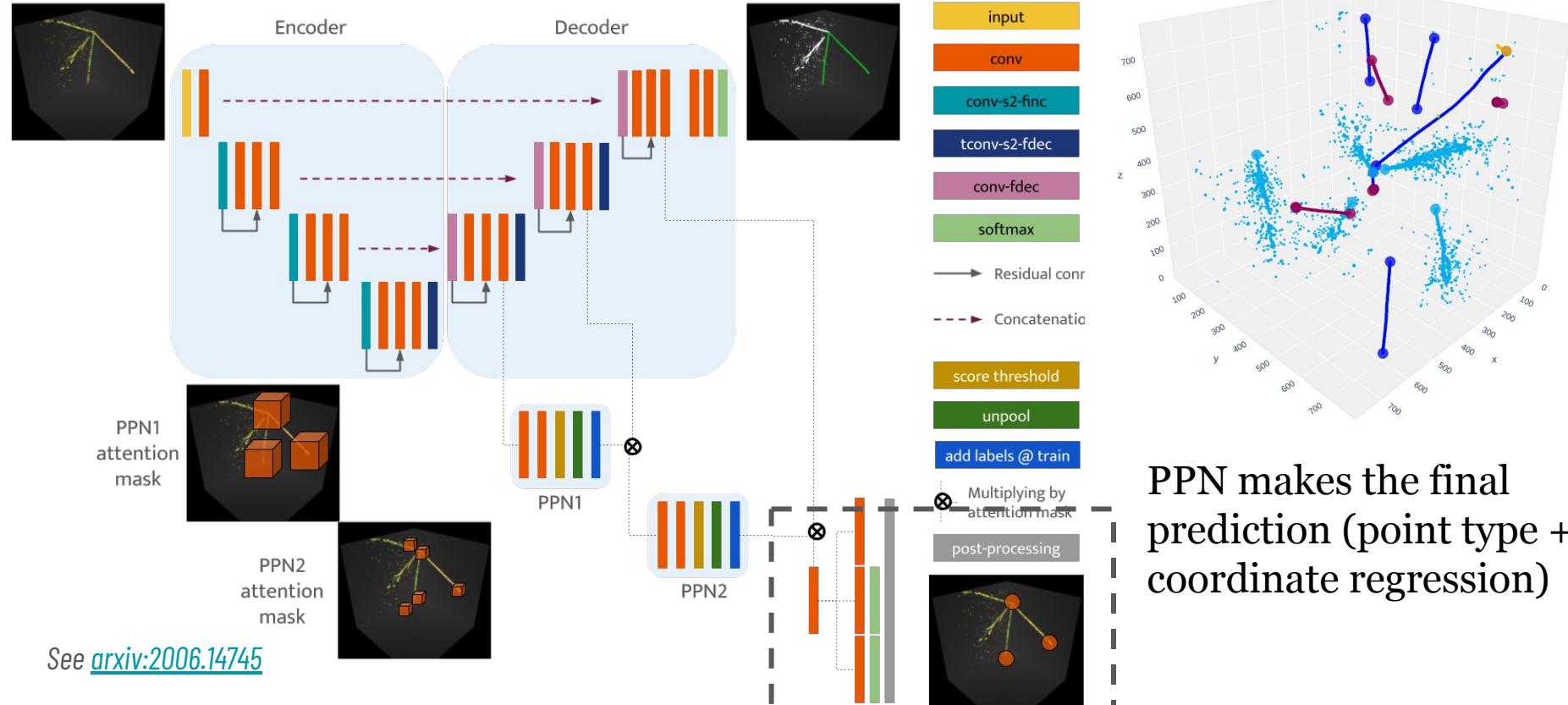
SLAC



# ML-based Neutrino Data Reconstruction Chain

## Stage 1-b: Particle Endpoint Prediction

SLAC



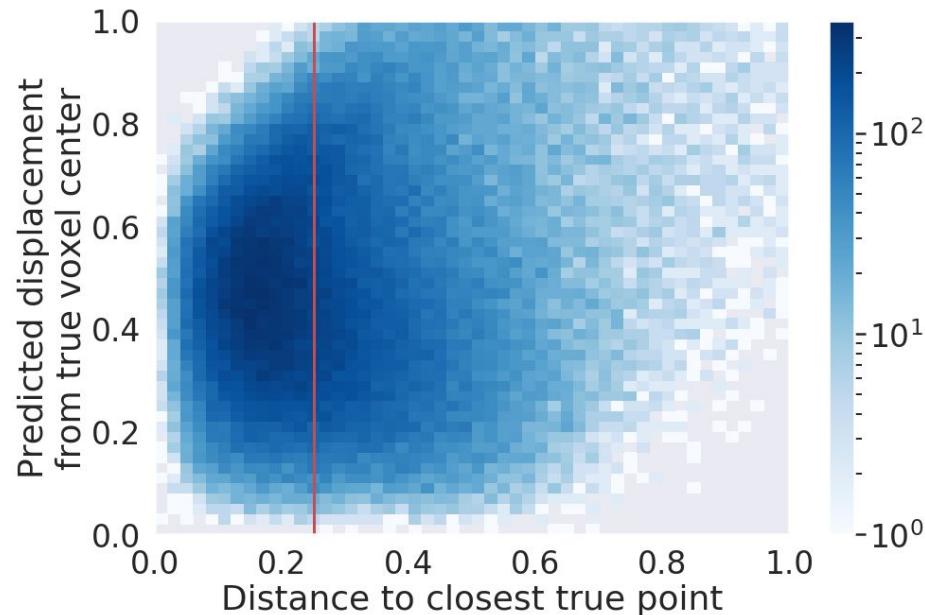
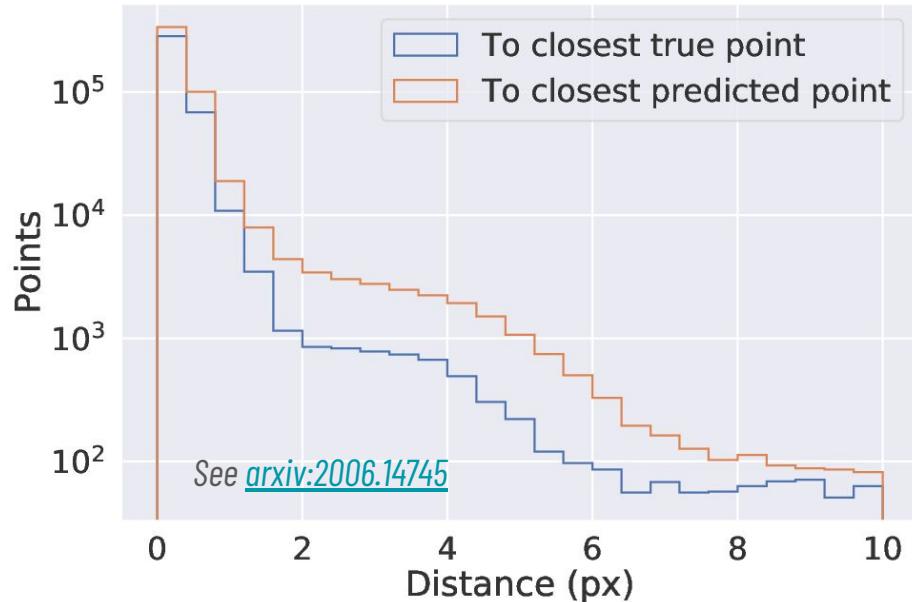
# ML-based Neutrino Data Reconstruction Chain

## Stage 1-b: Particle Endpoint Prediction

SLAC

96.8% of predicted points within 3 voxels of a true point

- 68% of true points found within the radius of 0.12 cm
- Traditional (nominal) reconstruction method finds 90% of predicted points within 17 voxels, and 68% of true points found within the radius of 0.74cm

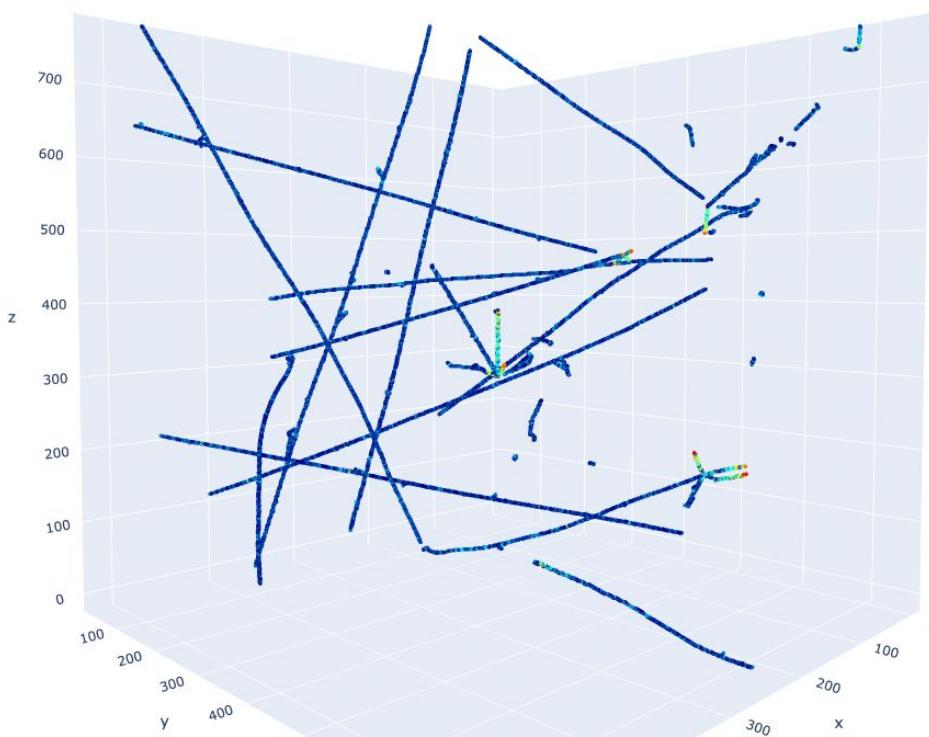


# ML-based Neutrino Data Reconstruction Chain

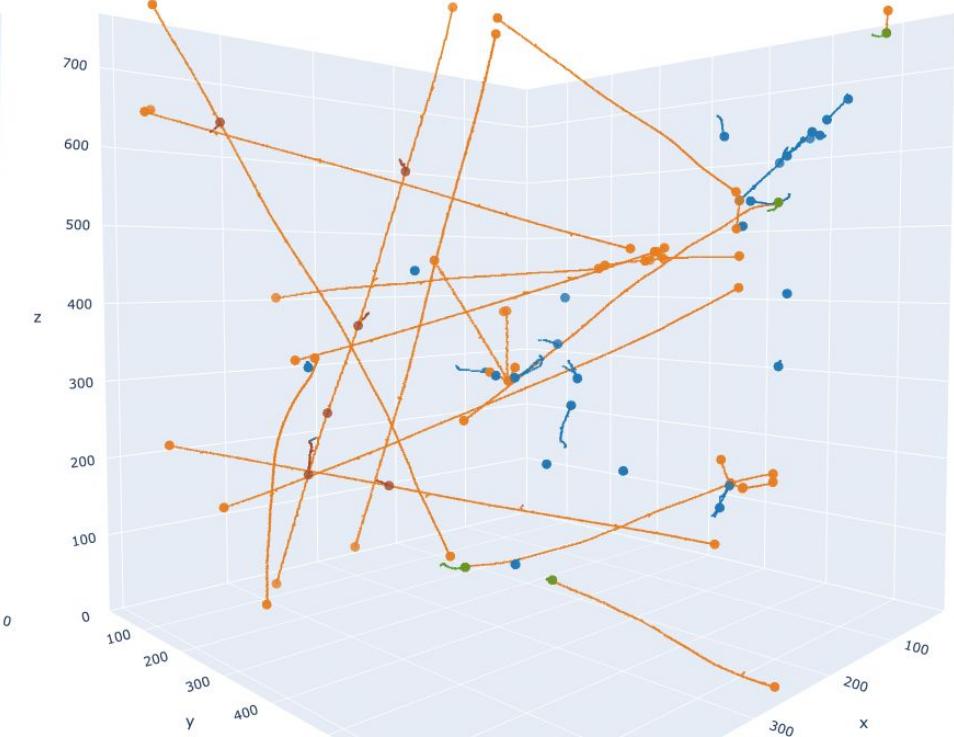
## Stage 1: input & output

SLAC

Stage 1 Input



Stage 1 Output



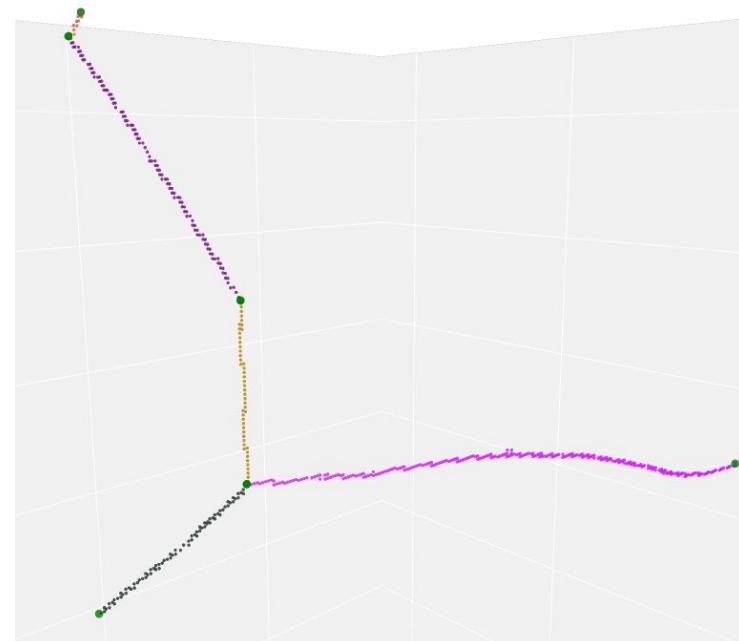
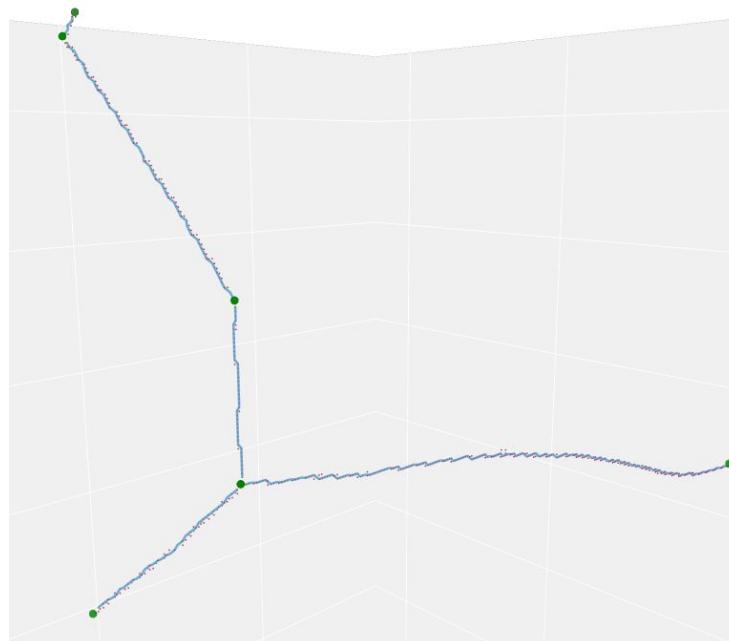
# ML-based Neutrino Data Reconstruction Chain

## Stage 2-a: Dense Pixel Clustering

SLAC

### Simple approach: path-finding between PPN points

- MST to find the “shortest” path between PPN points to cluster pixels
- **Works well! BUT** it depends on PPN performance directly + not learnable



# ML-based Neutrino Data Reconstruction Chain

## Stage 2: Particle & Interaction Clustering

SLAC

**Learnable approach:** clustering in the embedding space

- Use CNN to learn a transformation function from the 3D voxels to the embedding space where clustering can be performed in a simple manner

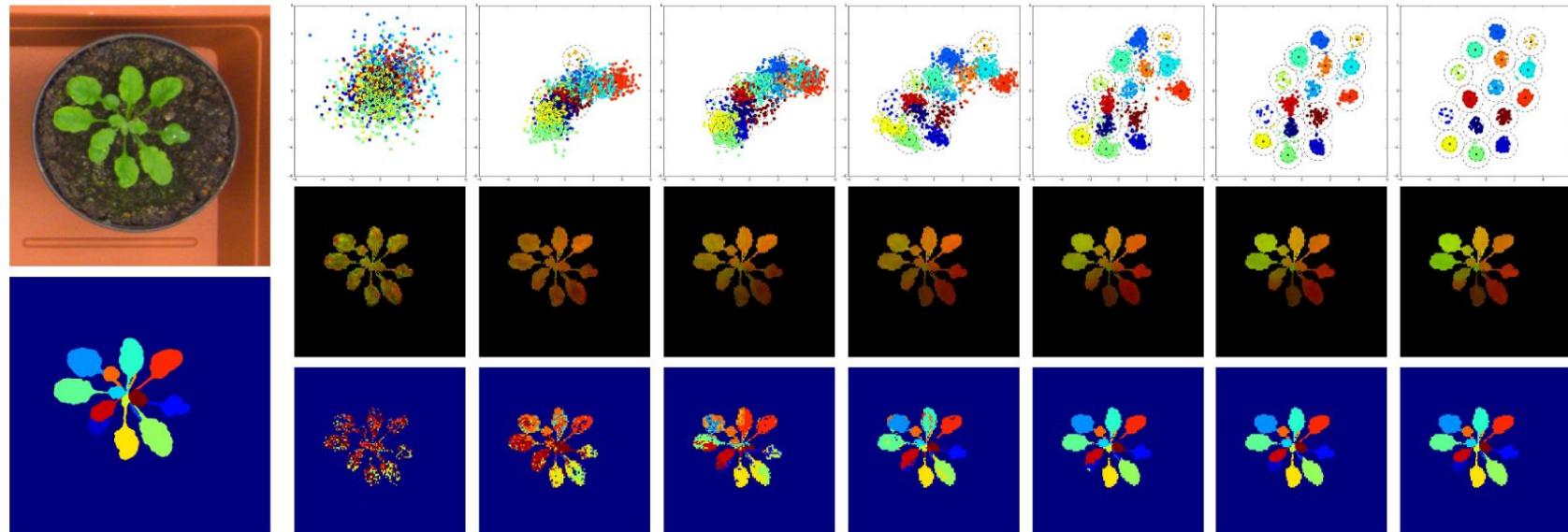
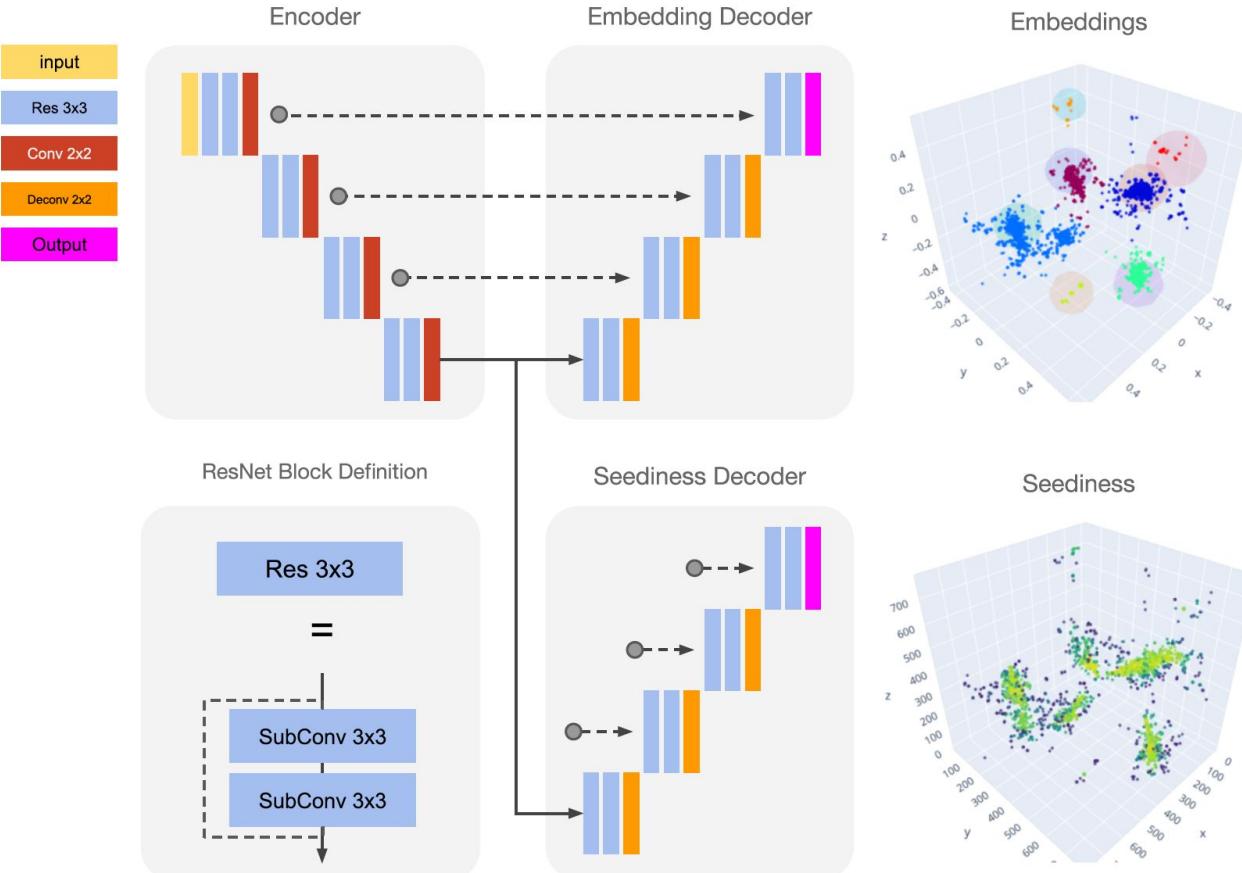


Image credit: [arXiv 1708.02551](https://arxiv.org/abs/1708.02551)

# ML-based Neutrino Data Reconstruction Chain

## Stage 2-a: Dense Pixel Clustering

SLAC



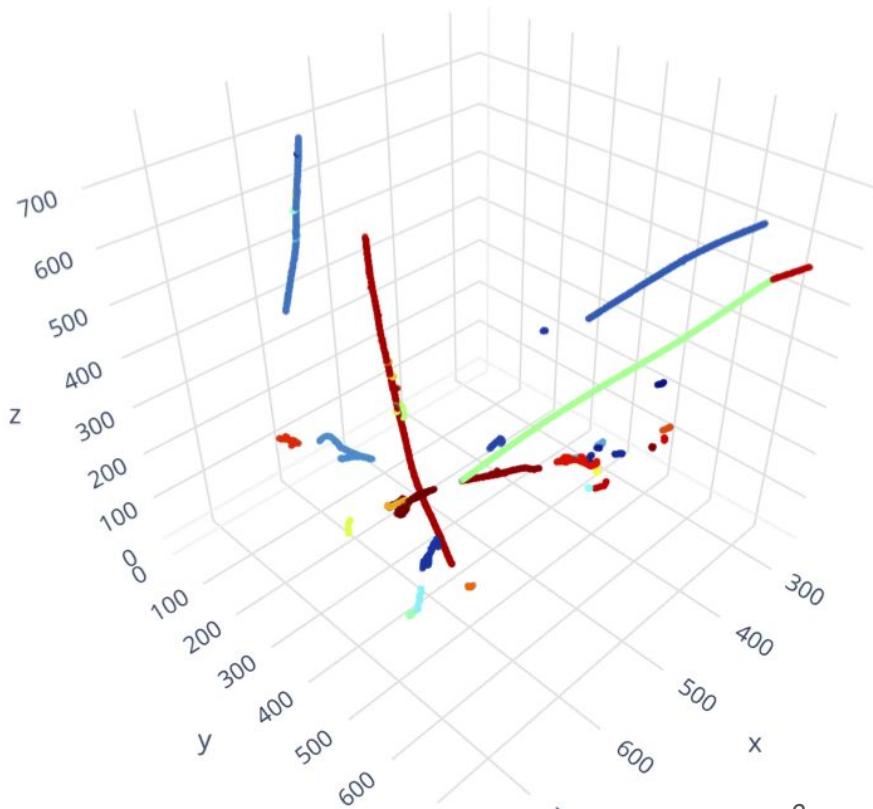
## Scalable Particle Instance Clustering using Embedding (SPICE)

- Embedding decoder learns transformation
- Seediness decoder identifies the centroids
- During training, loss is conditioned so that the points that belong to the same cluster follow a normal distribution

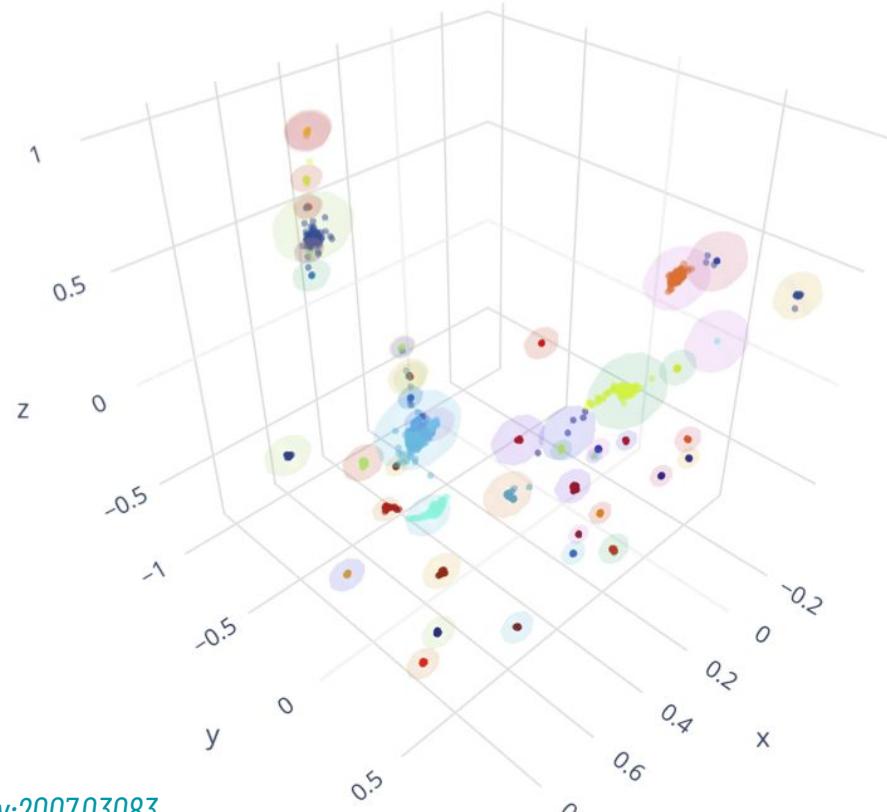
# ML-based Neutrino Data Reconstruction Chain

## Stage 2-a: Dense Pixel Clustering

SLAC



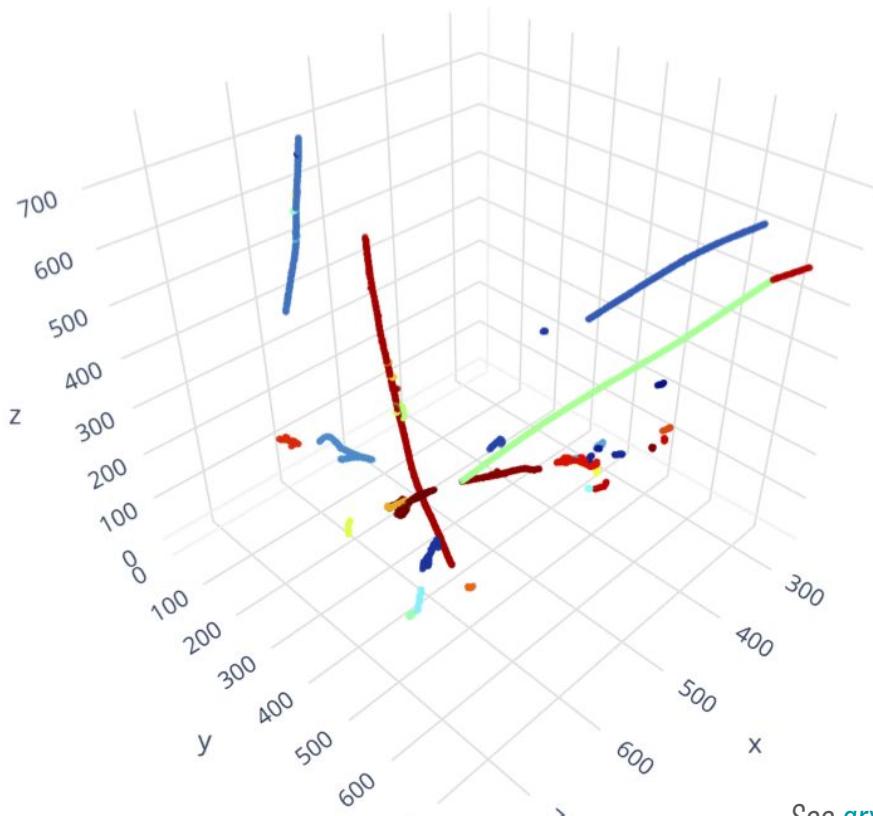
See [arxiv:2007.03083](https://arxiv.org/abs/2007.03083)



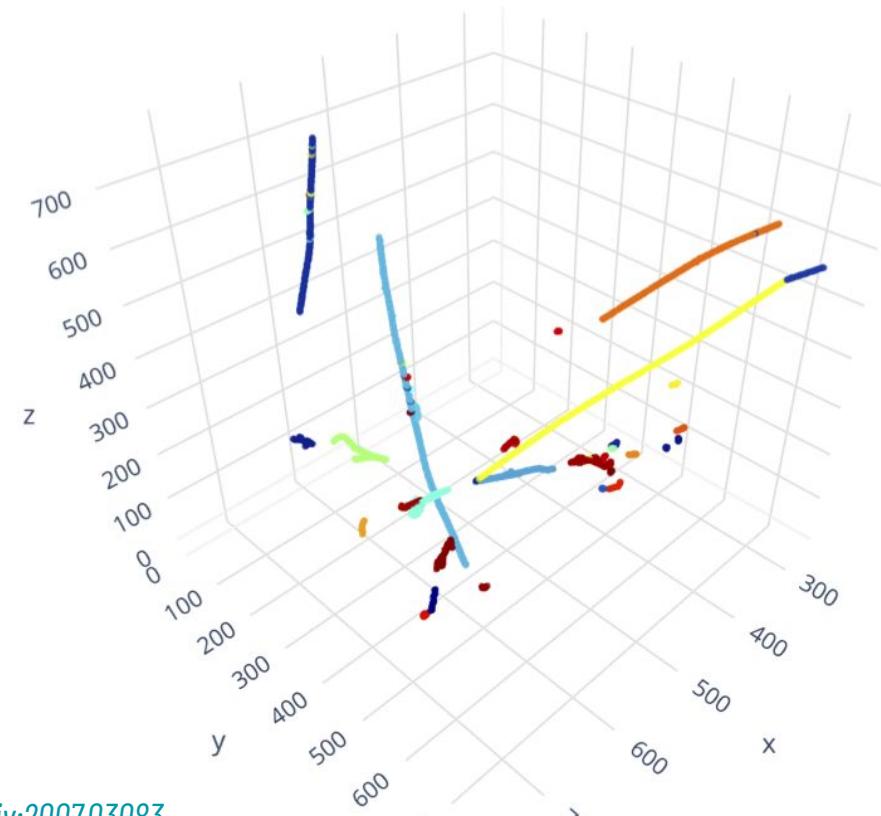
# ML-based Neutrino Data Reconstruction Chain

## Stage 2-a: Dense Pixel Clustering

SLAC



See [arxiv:2007.03083](https://arxiv.org/abs/2007.03083)

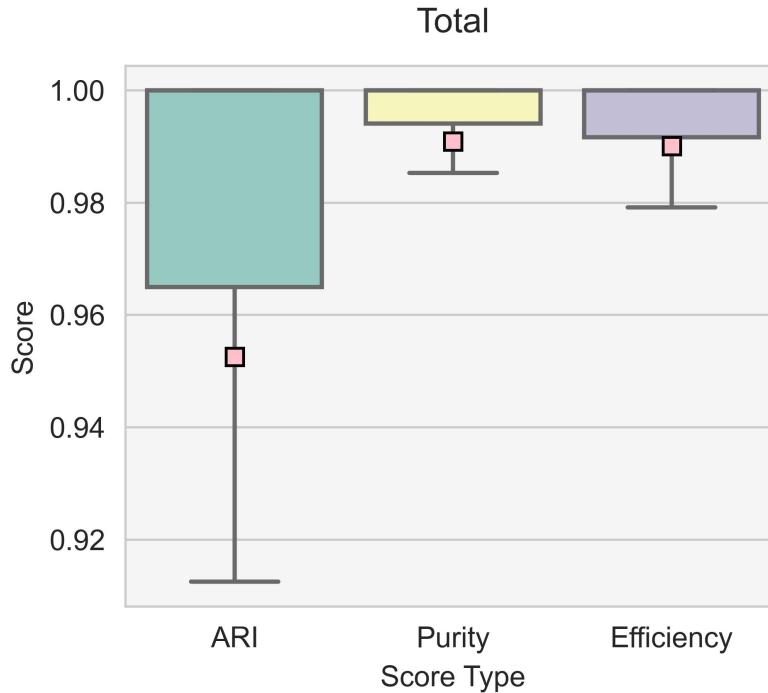


# ML-based Neutrino Data Reconstruction Chain

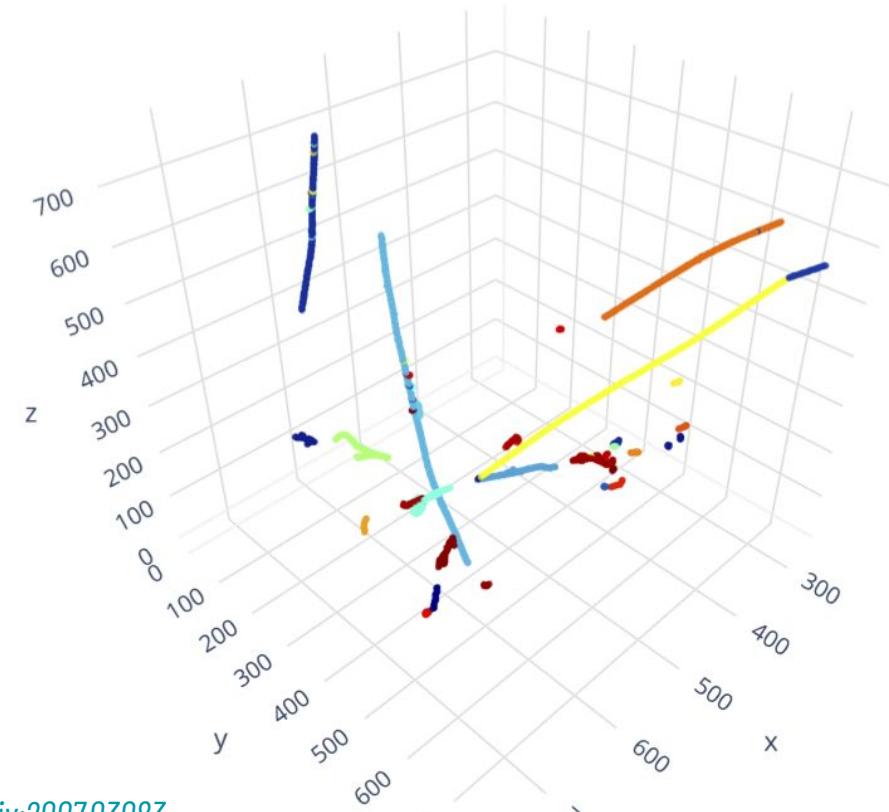
## Stage 2-a: Dense Pixel Clustering

SLAC

Pixels clustered into trajectory  
fragments using SPICE



See [arxiv:2007.03083](https://arxiv.org/abs/2007.03083)

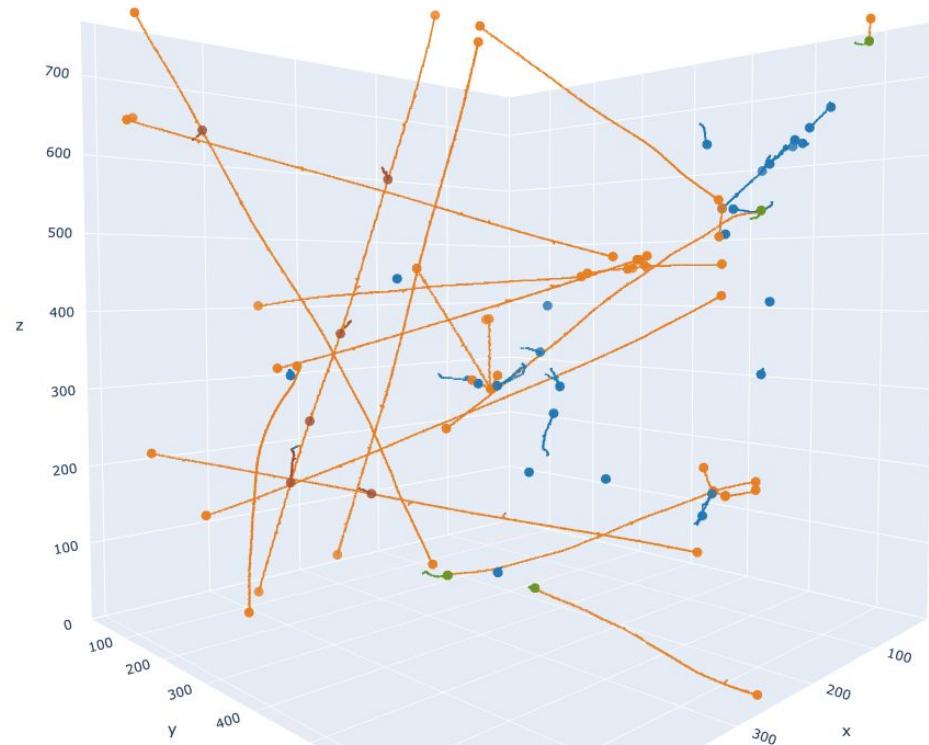


# ML-based Neutrino Data Reconstruction Chain

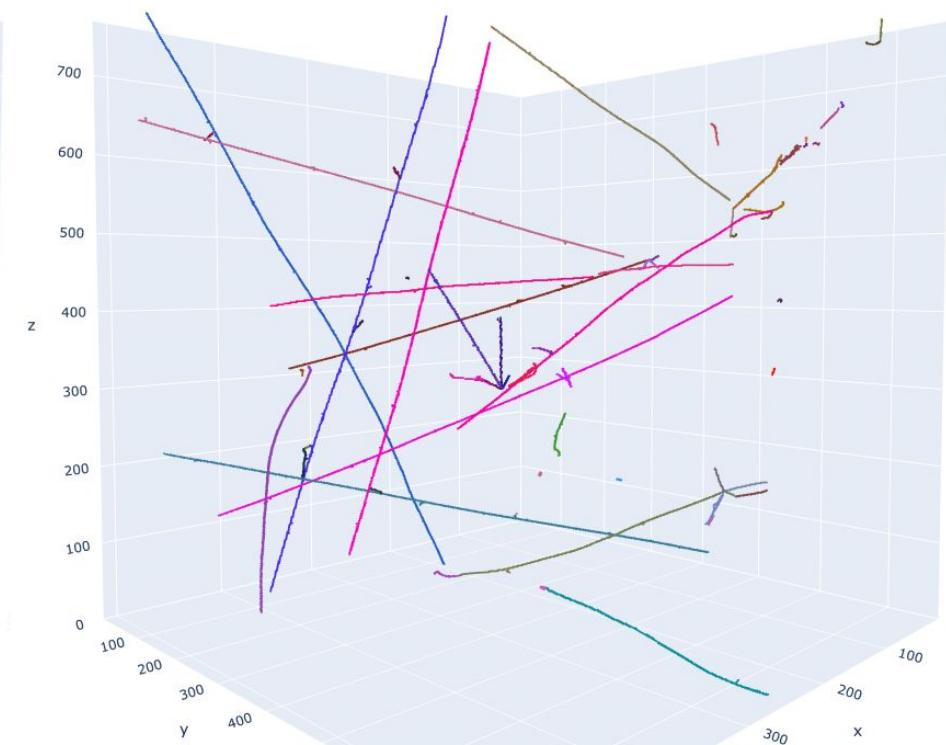
## Stage 2-a: input & output

SLAC

Stage 2-a Input



Stage 2-a Output



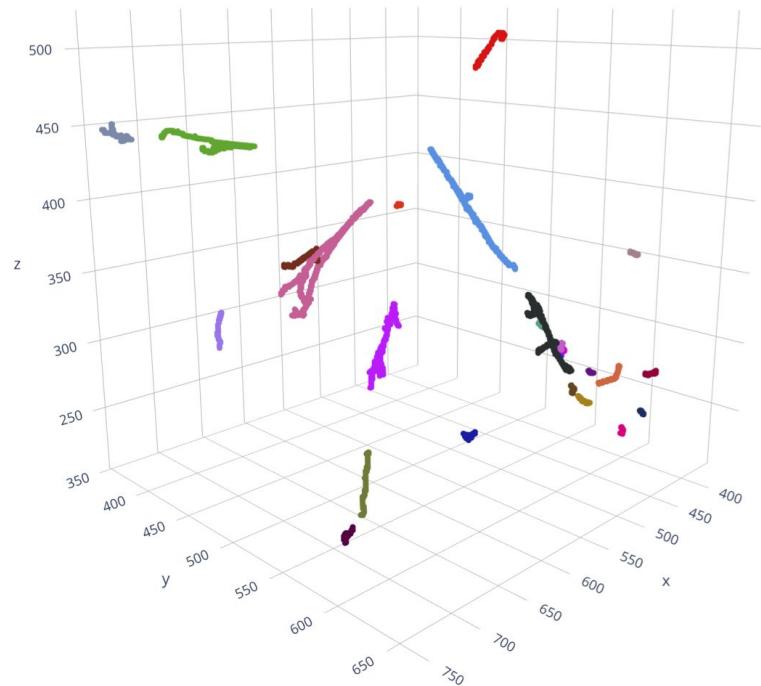
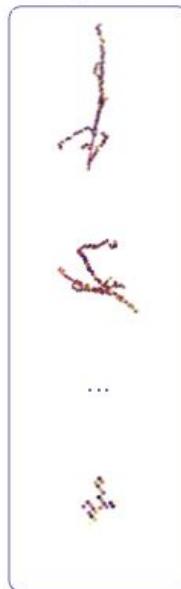
# ML-based Neutrino Data Reconstruction Chain

## Stage 2-b: Sparse Fragment Clustering

SLAC

Identifying **1 shower** ... which consists of **many fragments**

Fragments



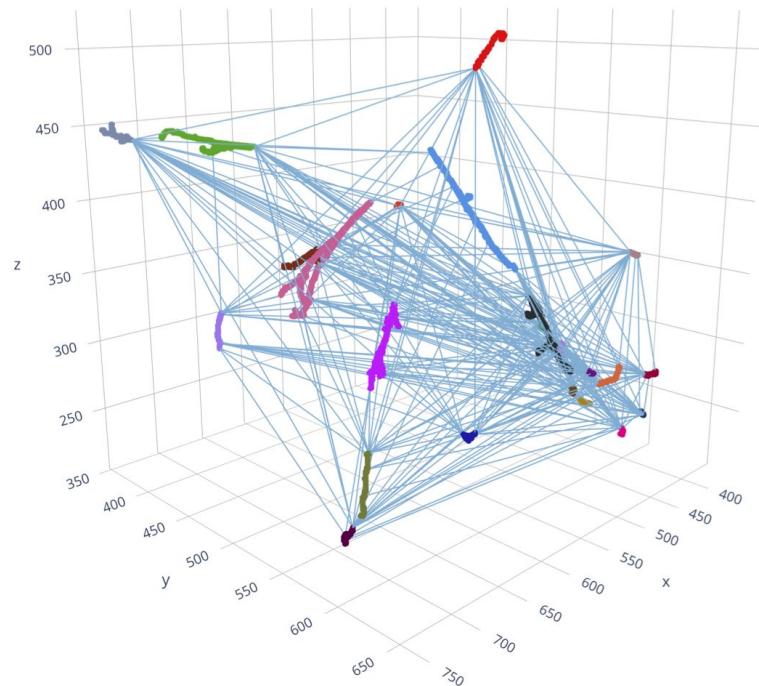
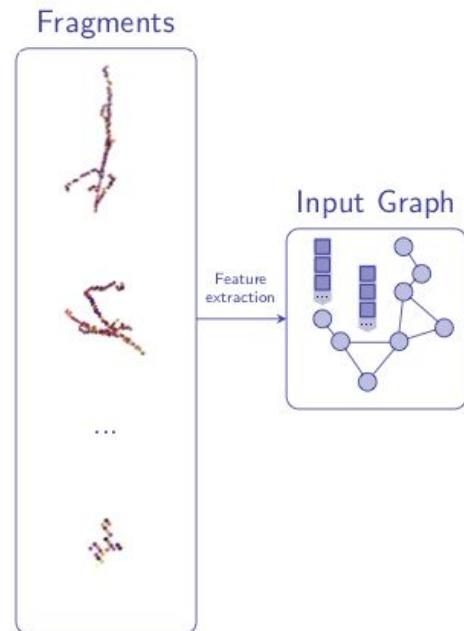
# ML-based Neutrino Data Reconstruction Chain

## Stage 2-b: Sparse Fragment Clustering

SLAC

**Identifying 1 shower** ... which consists of **many fragments**

- Interpret each fragment as a graph node + edges connect nodes in the same cluster



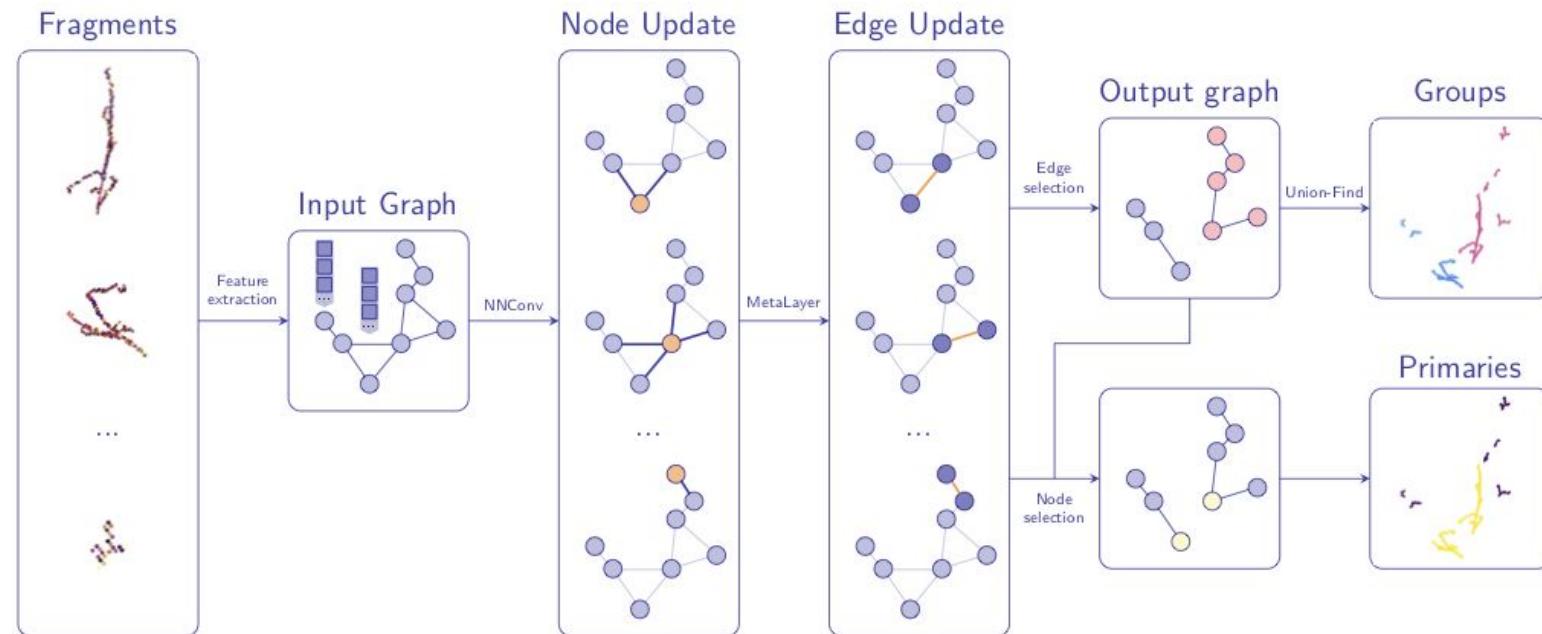
# ML-based Neutrino Data Reconstruction Chain

## Stage 2-b: Sparse Fragment Clustering

SLAC

**Identifying 1 shower** ... which consists of **many fragments**

- Interpret each fragment as a graph node + edges connect nodes in the same cluster
- Cast the problem to a classification of node (e.g. particle type) and edge (clustering)



# ML-based Neutrino Data Reconstruction Chain

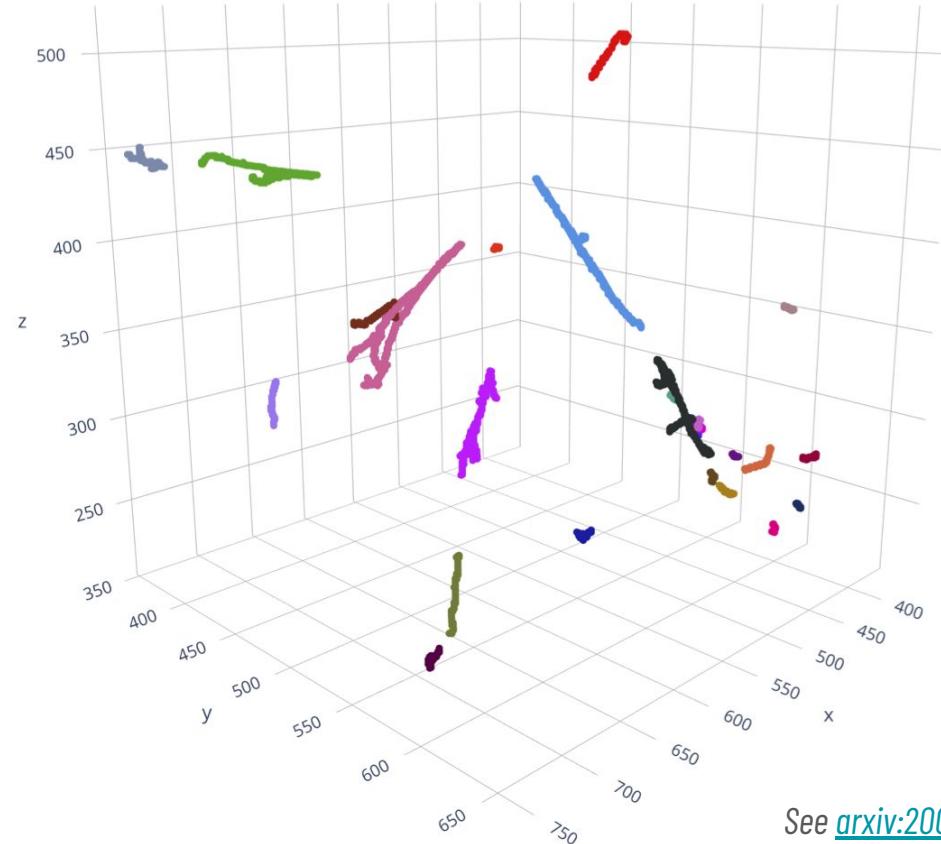
## Stage 2-b: Sparse Fragment Clustering

SLAC

### Graph-NN for Particle Aggregation (GrapPA)

Input:

- Fragmented EM showers



See [arxiv:2007.01335](https://arxiv.org/abs/2007.01335)

# ML-based Neutrino Data Reconstruction Chain

## Stage 2-b: Sparse Fragment Clustering

SLAC

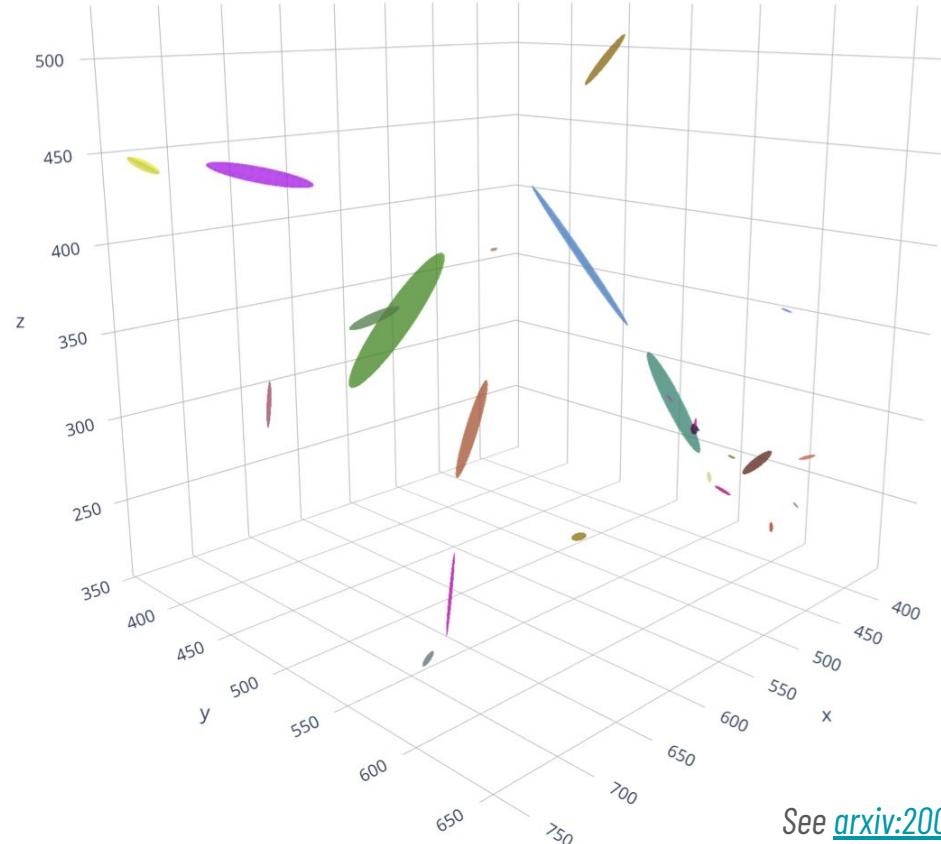
### Graph-NN for Particle Aggregation (GrapPA)

Input:

- Fragmented EM showers

Node features:

- Centroid, Covariance matrix, PCA
- Start point, direction (PPN)



See [arxiv:2007.01335](https://arxiv.org/abs/2007.01335)

# ML-based Neutrino Data Reconstruction Chain

## Stage 2-b: Sparse Fragment Clustering

SLAC

### Graph-NN for Particle Aggregation (GrapPA)

Input:

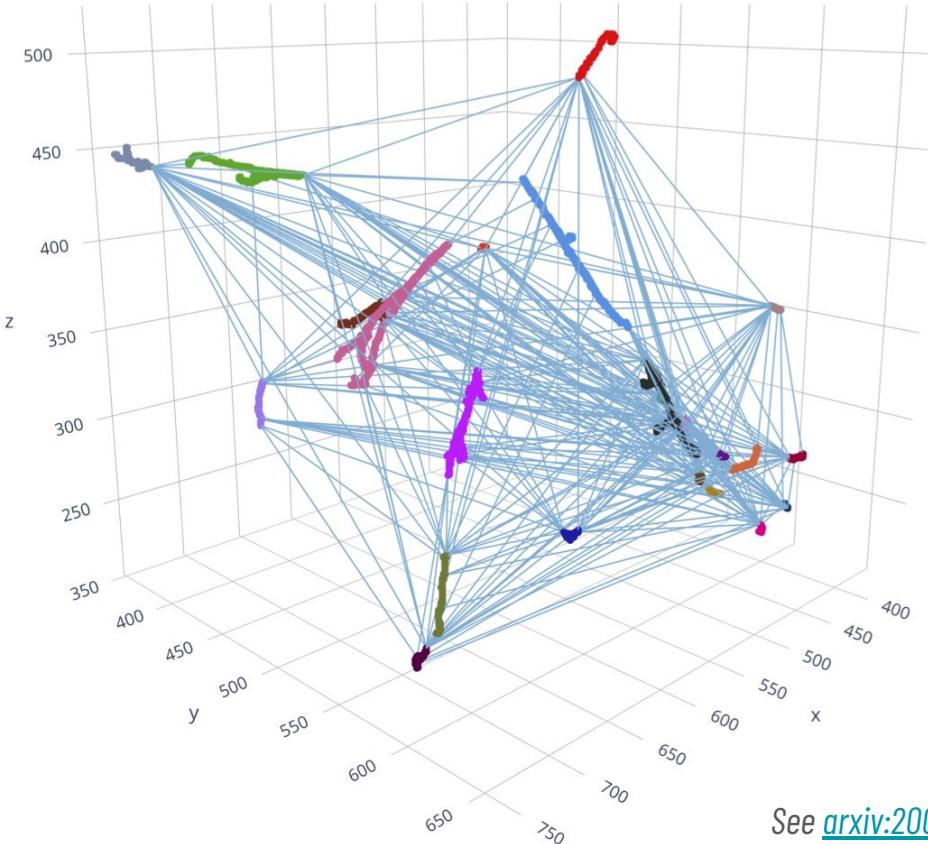
- Fragmented EM showers

Node features:

- Centroid, Covariance matrix, PCA
- Start point, direction (PPN)

Input graph:

- Connect every node with every other node (complete graph)



See [arxiv:2007.01335](https://arxiv.org/abs/2007.01335)

# ML-based Neutrino Data Reconstruction Chain

## Stage 2-b: Sparse Fragment Clustering

SLAC

### Graph-NN for Particle Aggregation (GrapPA)

Input:

- Fragmented EM showers

Node features:

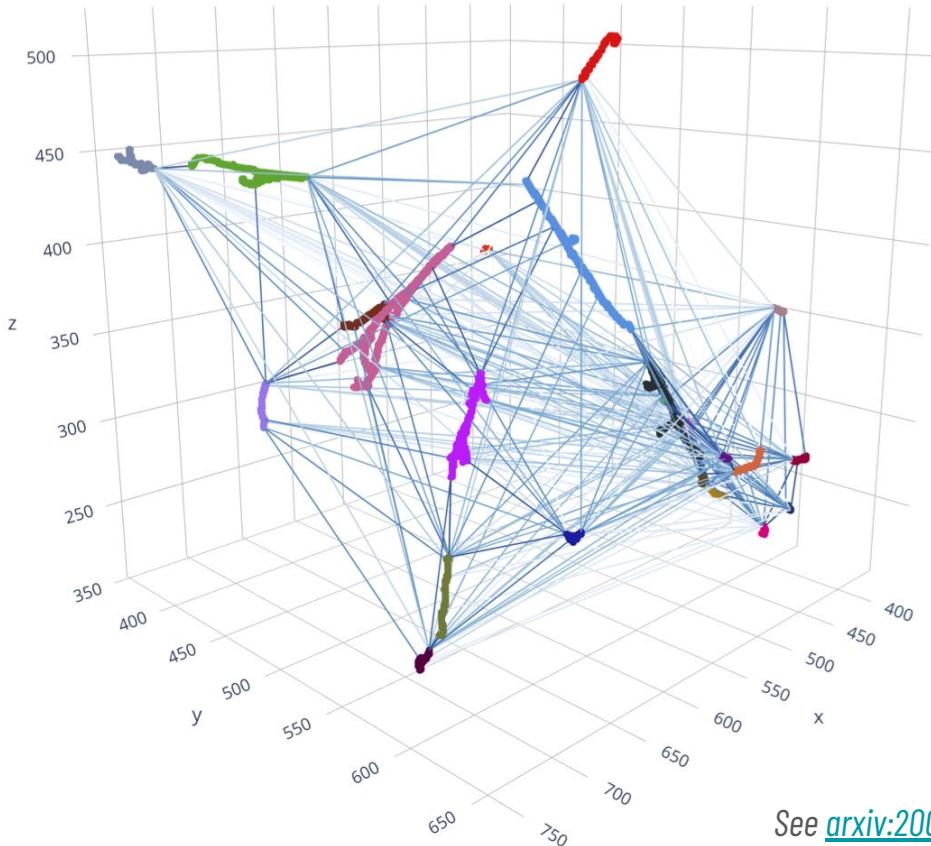
- Centroid, Covariance matrix, PCA
- Start point, direction (PPN)

Input graph:

- Connect every node with every other node (complete graph)

Edge features:

- Displacement vector
- Closest points of approach



See [arxiv:2007.01335](https://arxiv.org/abs/2007.01335)

# ML-based Neutrino Data Reconstruction Chain

## Stage 2-b: Sparse Fragment Clustering

SLAC

### Graph-NN for Particle Aggregation (GrapPA)

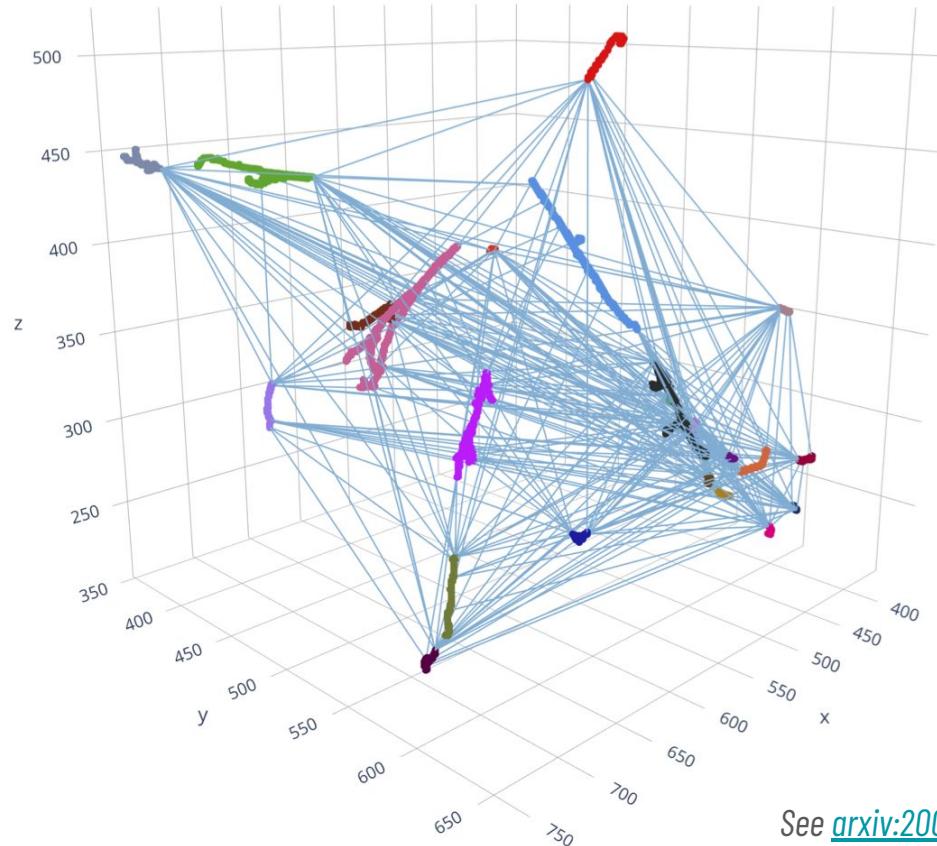
#### Message passing (MP):

- Meta layer ([arxiv:1806.01261](https://arxiv.org/abs/1806.01261))
- Essentially two 3-layer MLPs (BatchNorm + LeakyReLU) for edge feature update and node feature update
- 3 times MP (=Edge+Node feature update)

#### Target:

- Prediction of adjacency matrix representing valid edges (=true partition)
- Apply cross-entropy loss

For more studies, see [our paper](https://arxiv.org/abs/2007.01335)



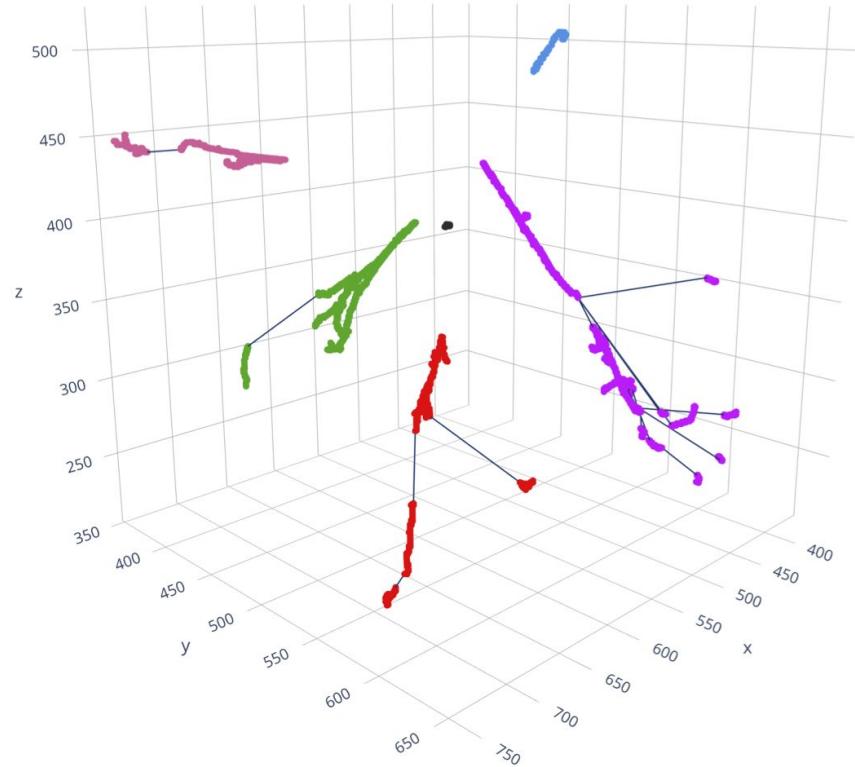
See [arxiv:2007.01335](https://arxiv.org/abs/2007.01335)

# ML-based Neutrino Data Reconstruction Chain

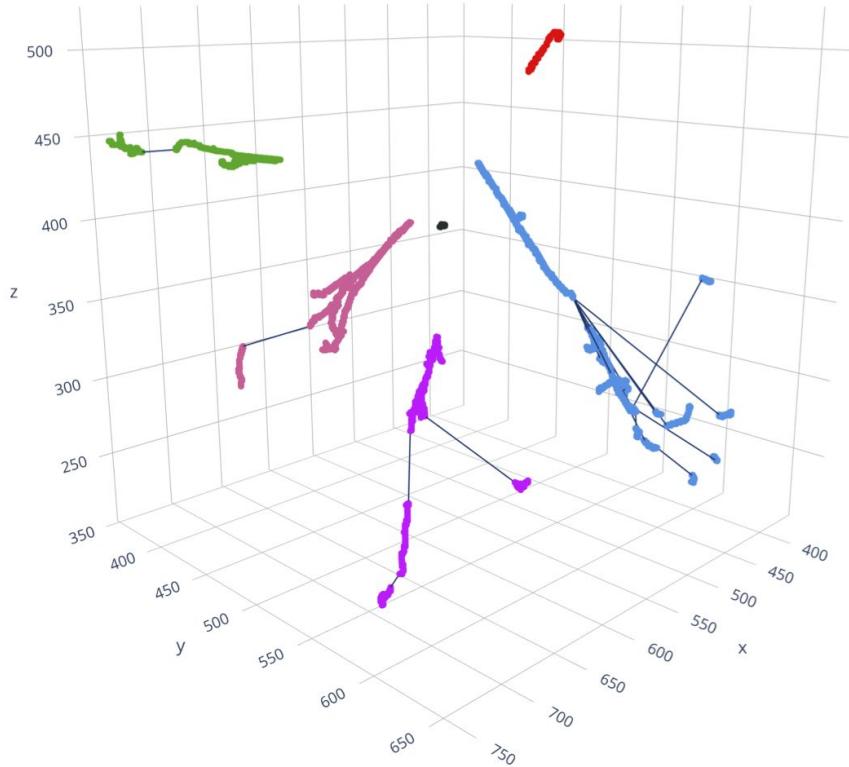
## Stage 2-b: Sparse Fragment Clustering

SLAC

Target Label



Edge Prediction



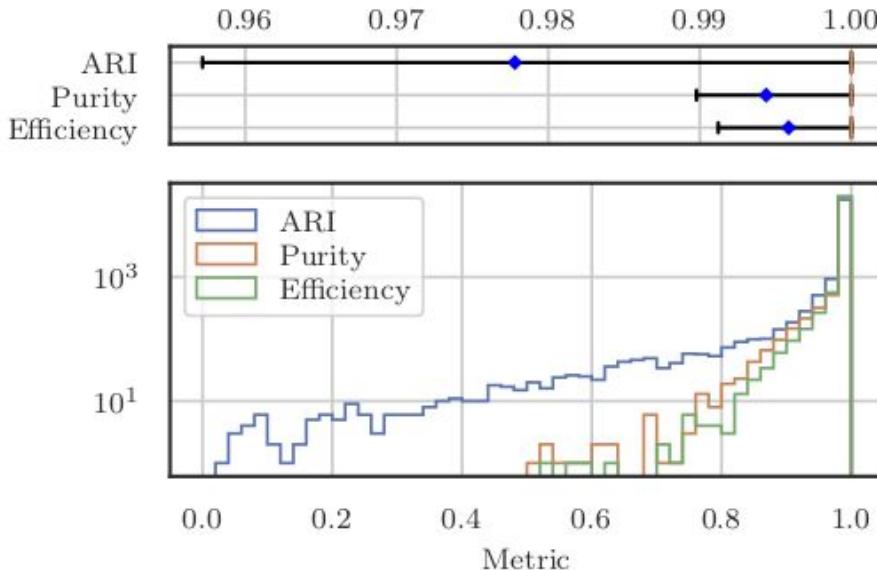
# ML-based Neutrino Data Reconstruction Chain

## Stage 2-b: Sparse Fragment Clustering

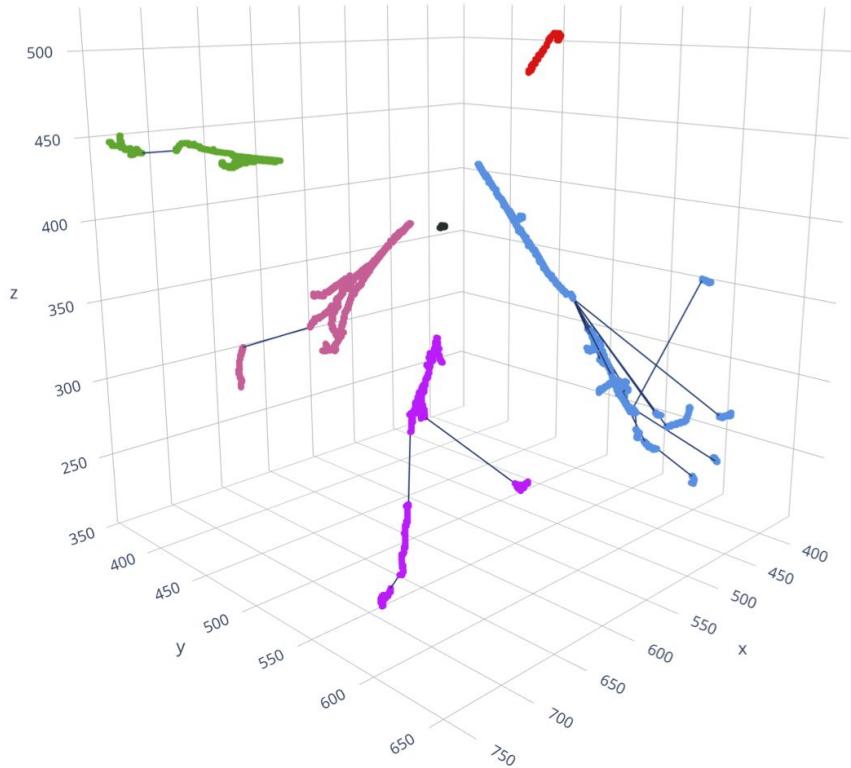
SLAC

### Clustering using GrapPA

- Mean purity and efficiency > 99%
- Sufficient for moving to the next stage (particle analysis)



### Edge Prediction



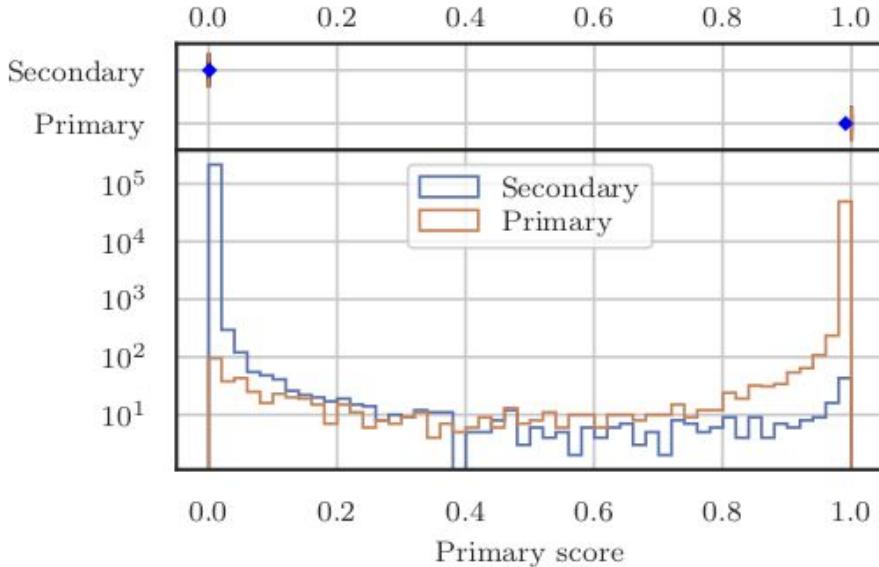
# ML-based Neutrino Data Reconstruction Chain

## Stage 2-b: Sparse Fragment Clustering

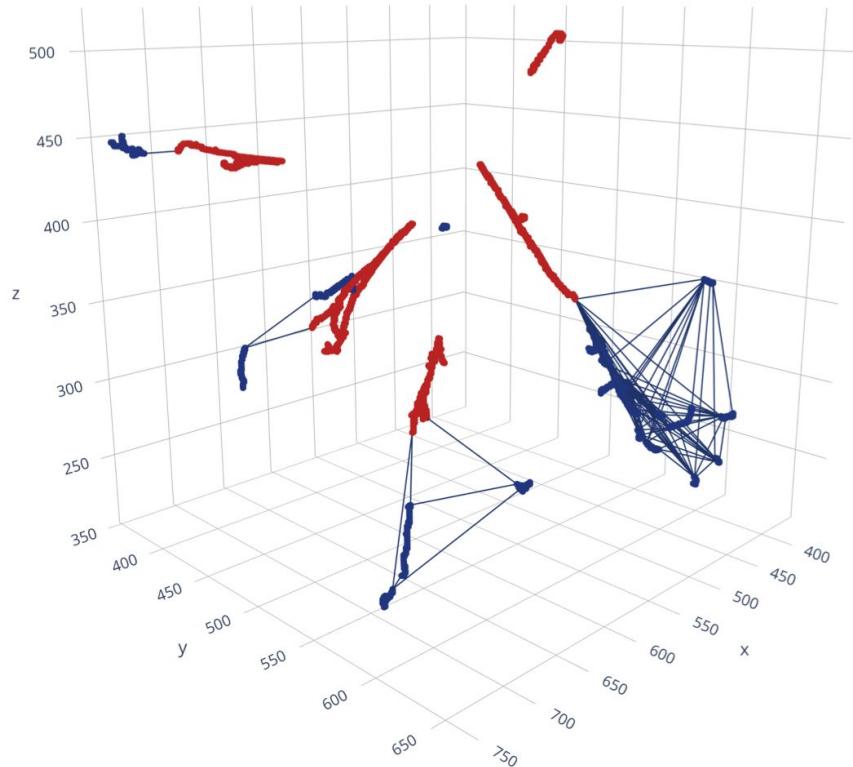
SLAC

### Start ID using GrapPA

- Important to identify the “primary fragment” (=shower start)
- >99% classification accuracy



### Node prediction

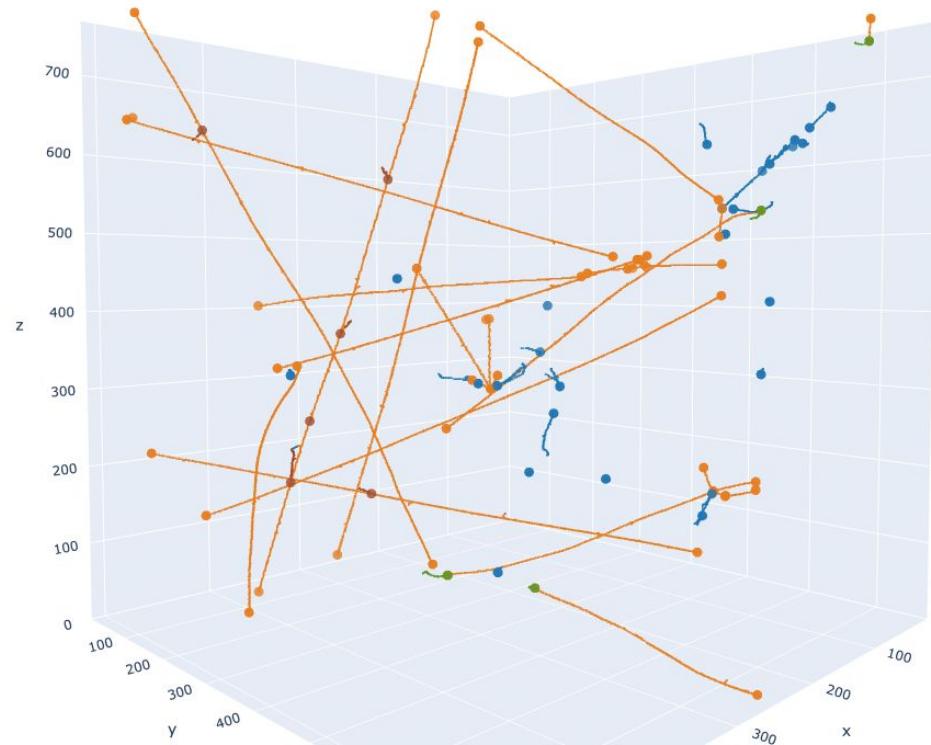


# ML-based Neutrino Data Reconstruction Chain

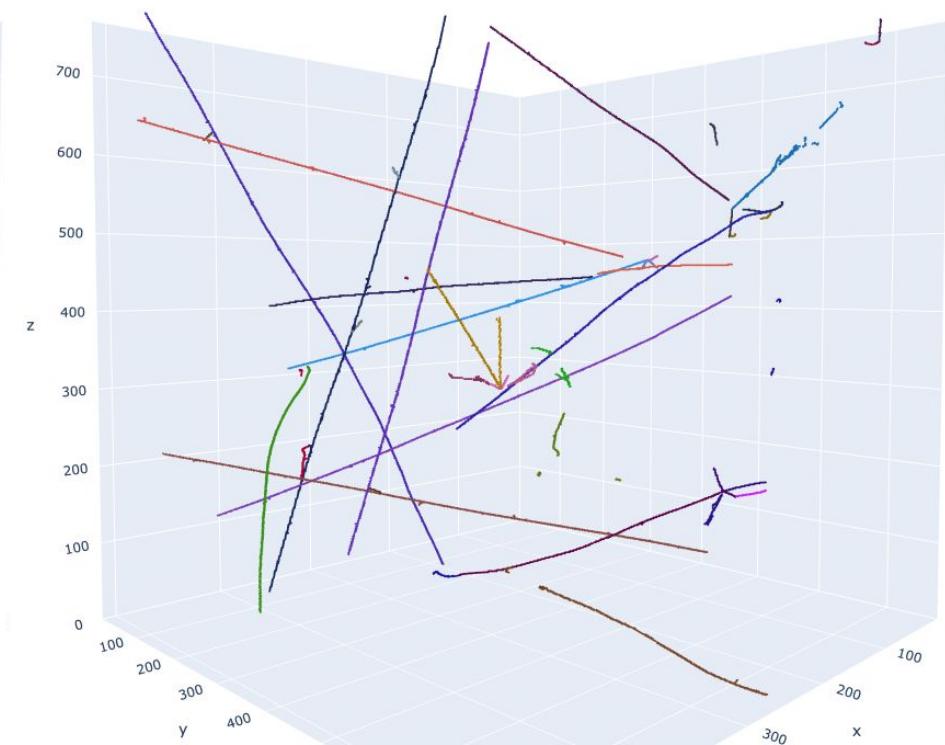
## Stage 2: input & output

SLAC

Stage 2 Input



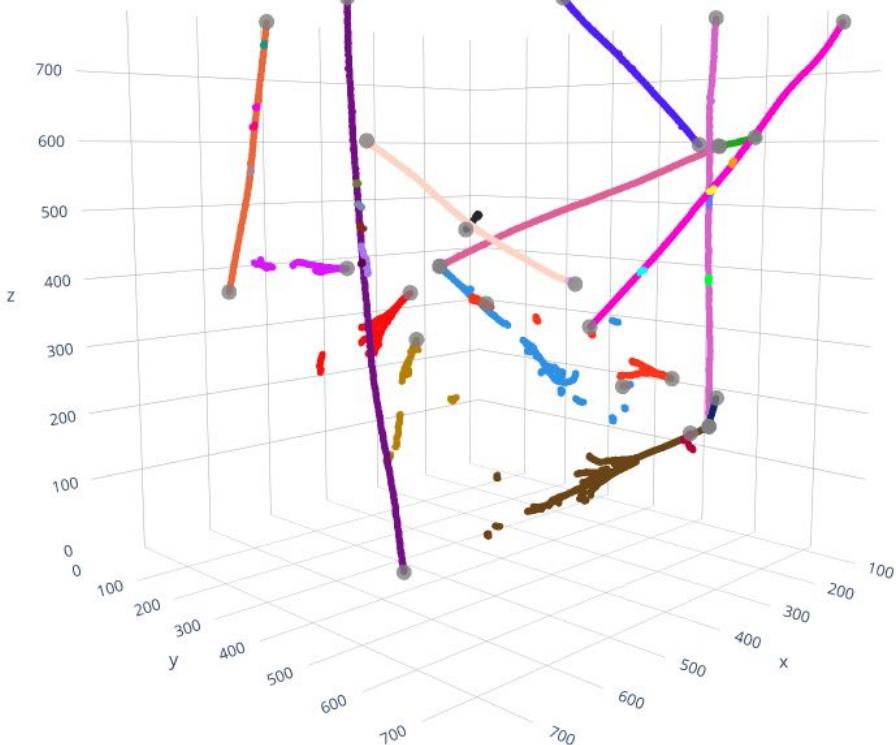
Stage 2 Output



# ML-based Neutrino Data Reconstruction Chain

## Stage 3: Interaction Clustering

SLAC



### Identifying Each Interaction?

This task can be casted to the same task already solved using GrapPA!

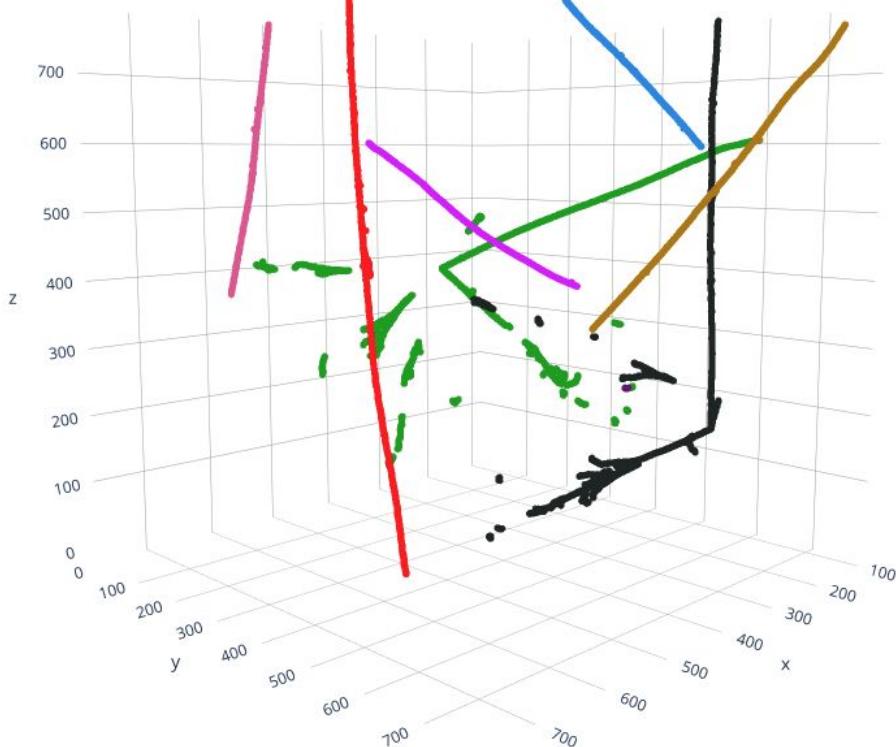
- Interaction = a group of particles that shared the same origin (i.e. neutrino interaction)
- Edge classification to identify an interaction
- Node classification for particle type ID

# ML-based Neutrino Data Reconstruction Chain

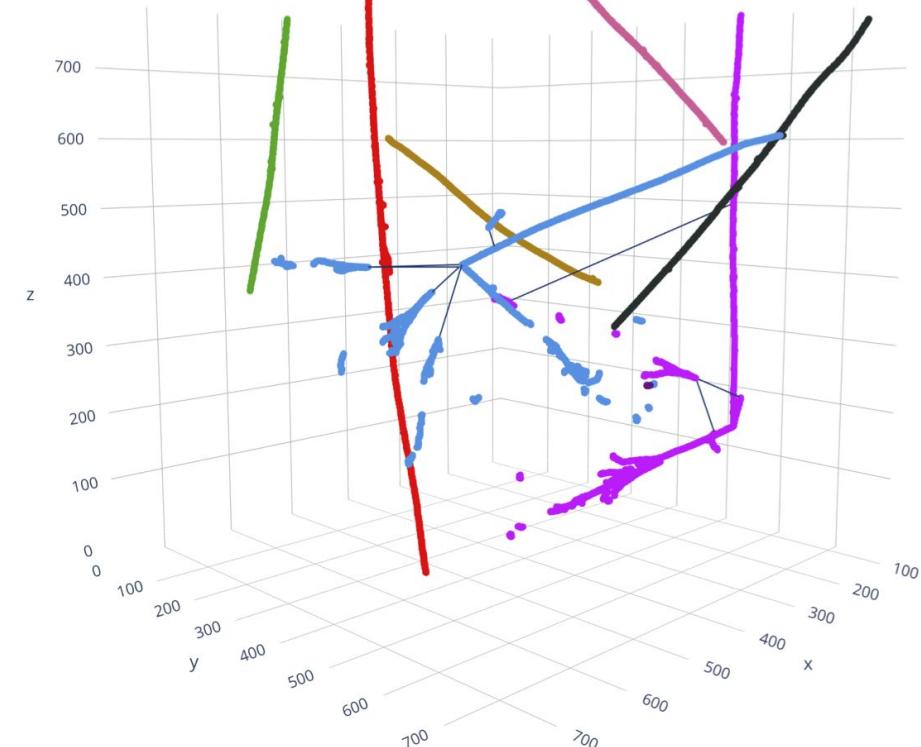
## Stage 3: Interaction Clustering

SLAC

Target Group



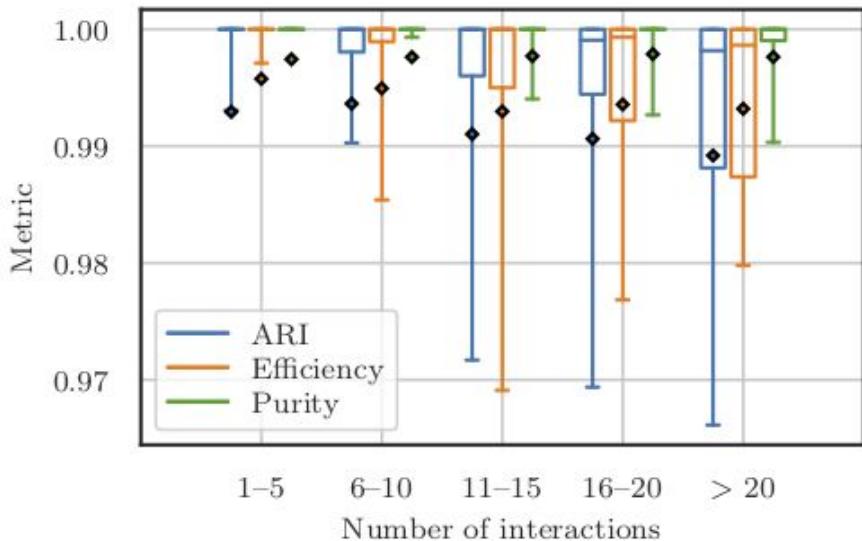
Predicted Interaction



# ML-based Neutrino Data Reconstruction Chain

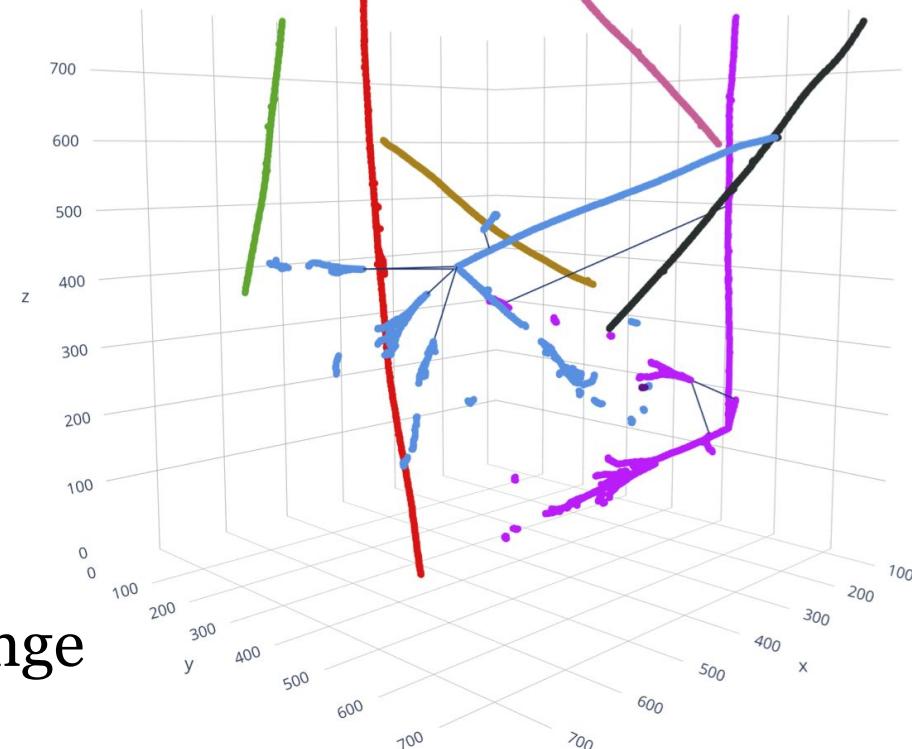
## Stage 3: Interaction Clustering

SLAC



Promising result to address  
DUNE-ND reconstruction challenge  
(~20 neutrino pile-up)

## Predicted Interaction

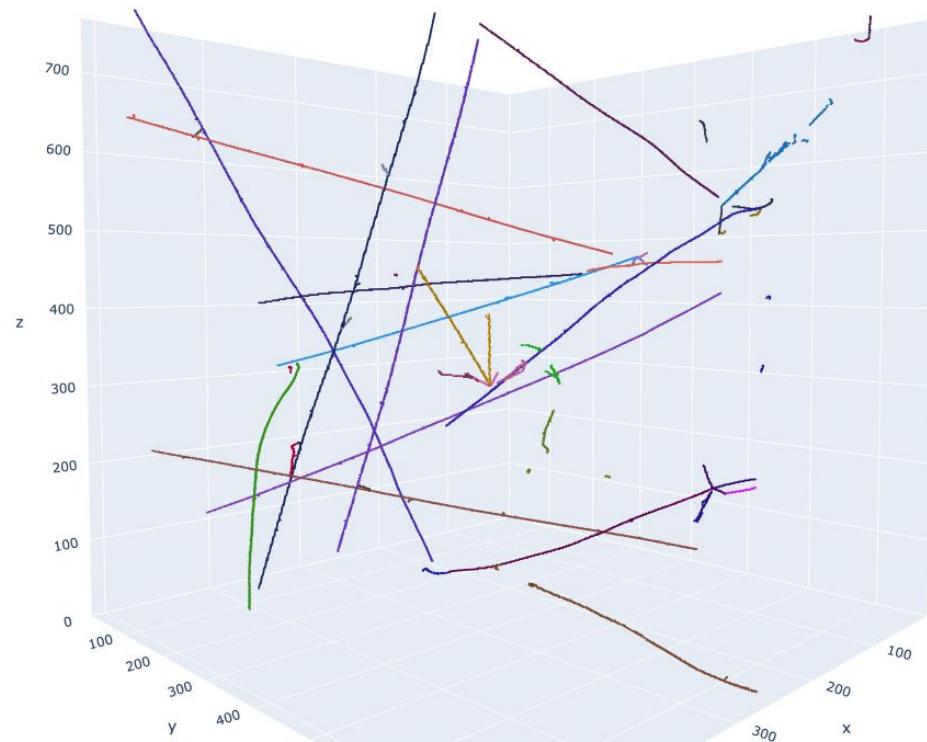


# ML-based Neutrino Data Reconstruction Chain

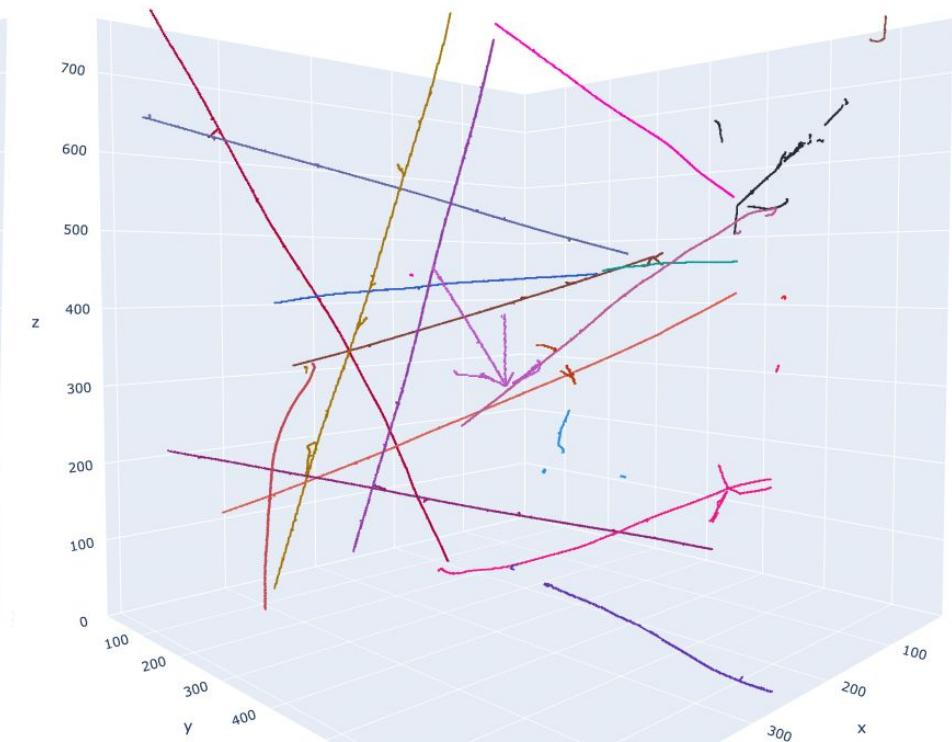
## Stage 3: input & output

SLAC

Stage 3 Input



Stage 3 Output



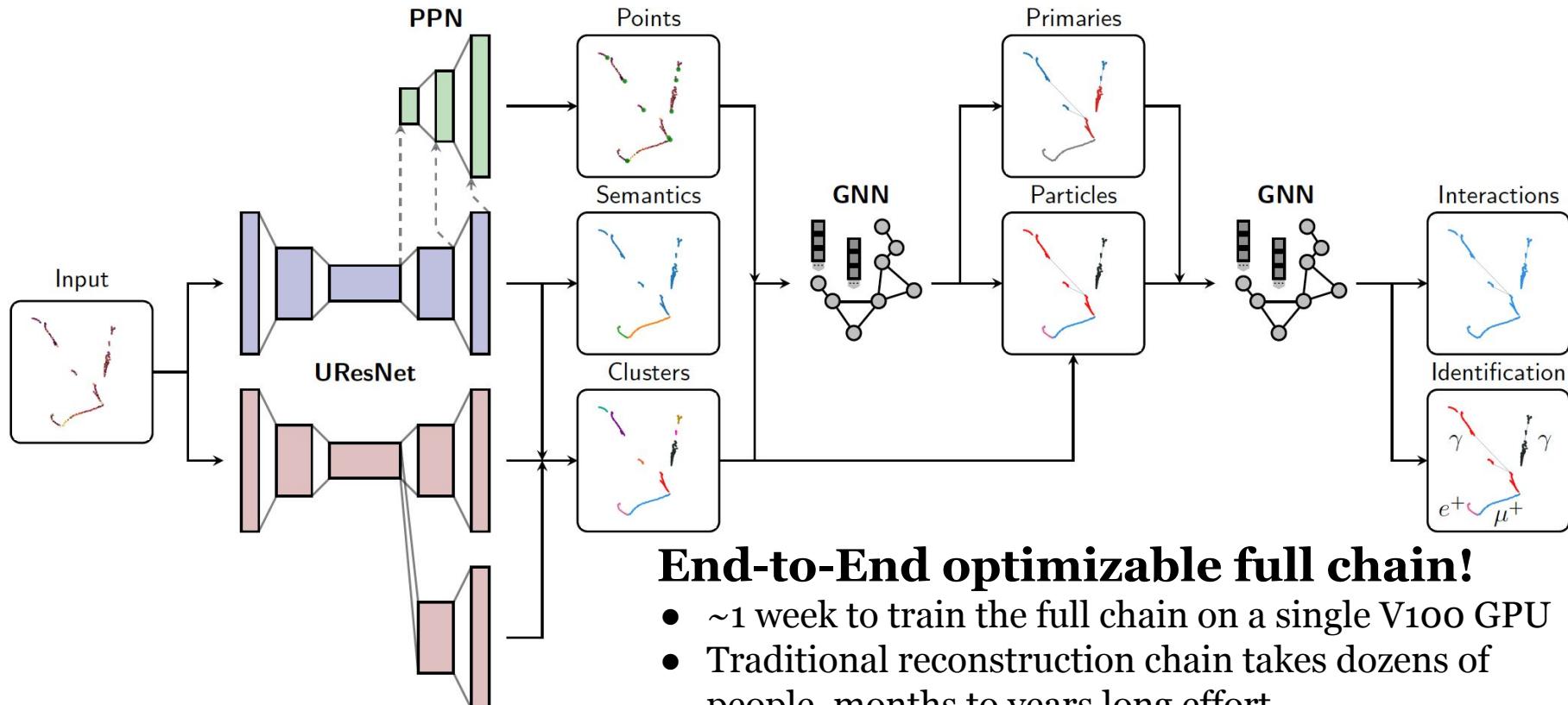


... wrapping up ...

# ML-based Neutrino Data Reconstruction Chain

## Wrapping up...

SLAC



### Summary

- **Neutrino detector trend: hi-res. particle imaging**
- **Analysis trend: computer vision algorithms**
  - Benefit the hi-resolution image = lots of heuristics (in non-ML)
  - ML-based approach has shown strong promise
- **ML-based data reconstruction approach**
  - especially for “busy” detectors ... my research :)
  - Working on implementing inductive-bias/causality (“physics”)
- **Other active areas:** data/sim domain discrepancy adaptation
  - minimize the discrepancy, identify the source, quantify uncertainty

**FIN**

# Machine Learning for Particle Image Analysis

**SLAC**

# Questions?

# SPICE

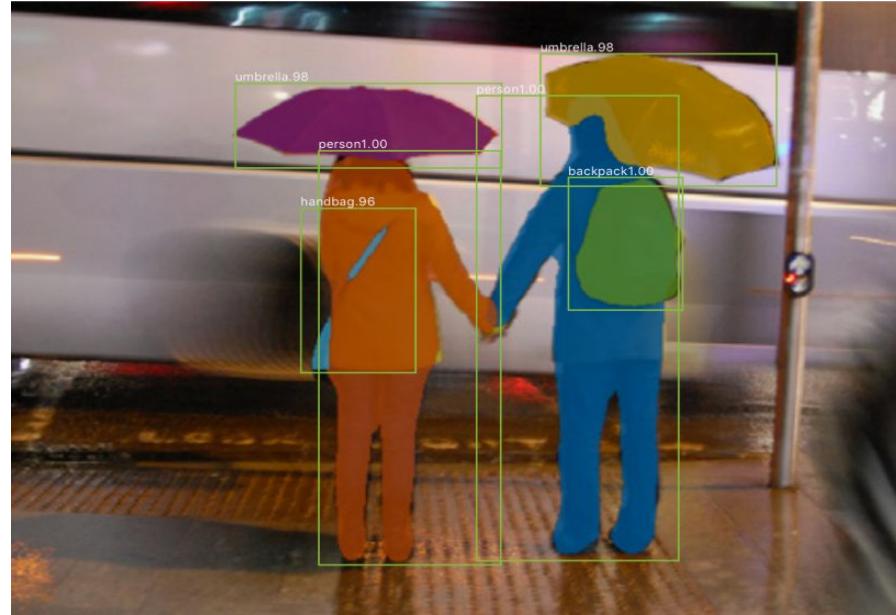
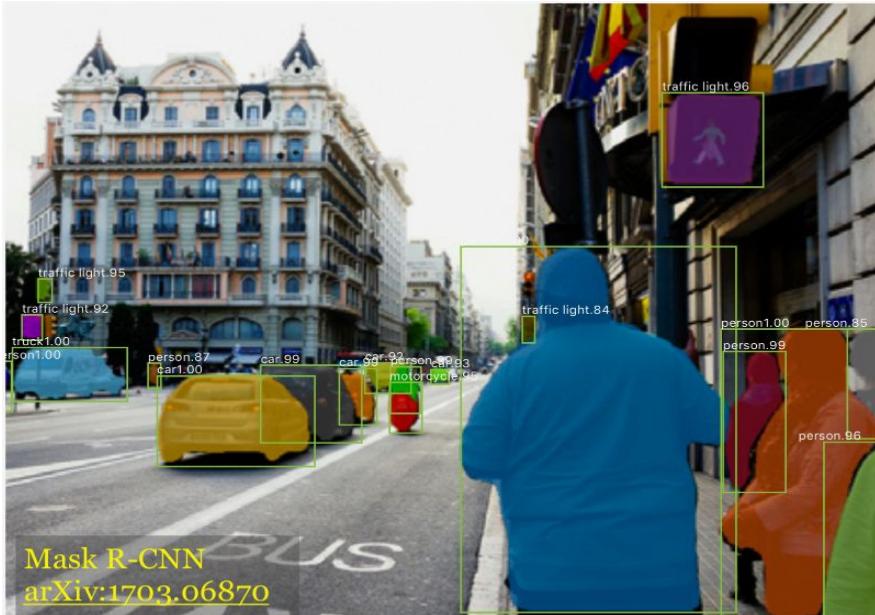
# ML-based Neutrino Data Reconstruction Chain

## Stage 2: Particle & Interaction Clustering

SLAC

### Instance+Semantic Segmentation

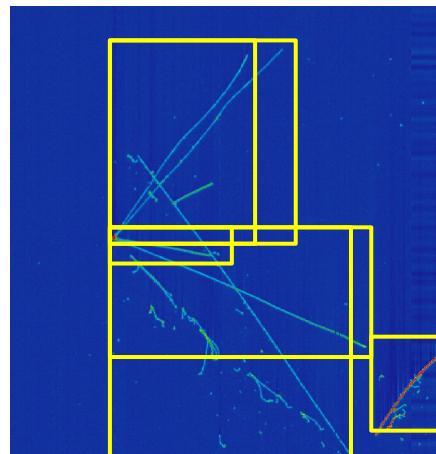
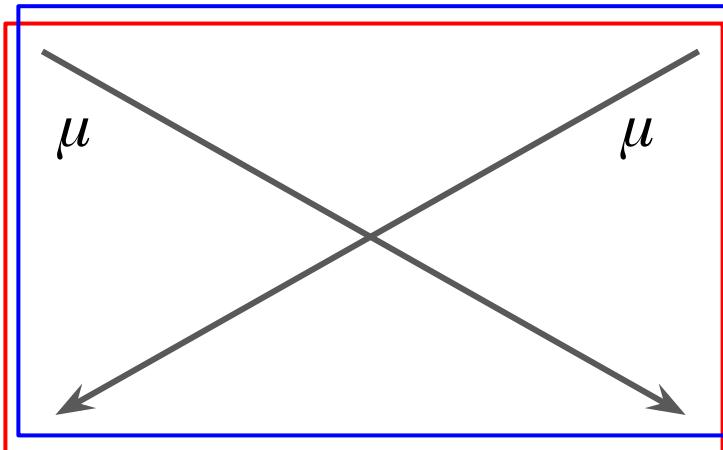
- **Mask R-CNN** ... a popular solution, many applications in science/industries
  - Object (=instance) detection + instance pixel masking in a bounding box



## Stage 2: Particle & Interaction Clustering

### Instance+Semantic Segmentation

- **Mask R-CNN** ... a popular solution, many applications in science/industries
  - Object (=instance) detection + instance pixel masking in a bounding box
  - **Issue:** instance distinction is affected by BB position/size
  - Another family: Single-Shot-Detection (SSD) based (not covered here)



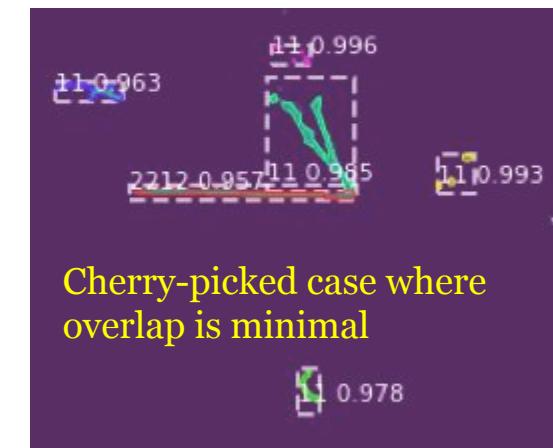
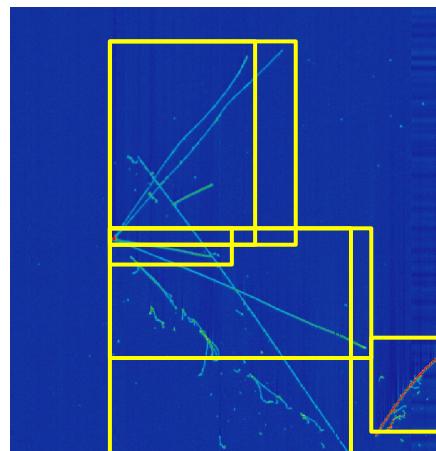
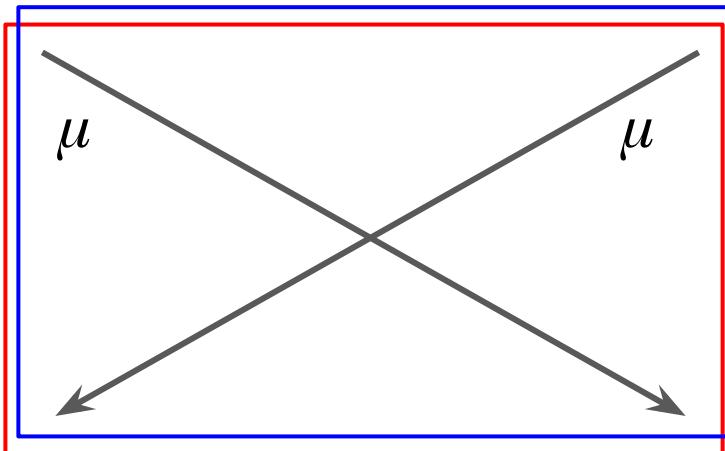
### Occlusion issue

The overlap rate of particles is very high especially for our signal (neutrinos) with an event vertex.

## Stage 2: Particle & Interaction Clustering

### Instance+Semantic Segmentation

- **Mask R-CNN** ... a popular solution, many applications in science/industries
  - Object (=instance) detection + instance pixel masking in a bounding box
  - **Issue:** instance distinction is affected by BB position/size
  - Another family: Single-Shot-Detection (SSD) based (not covered here)



# ML-based Neutrino Data Reconstruction Chain

## Stage 2: Particle & Interaction Clustering

SLAC

### Instance+Semantic Segmentation

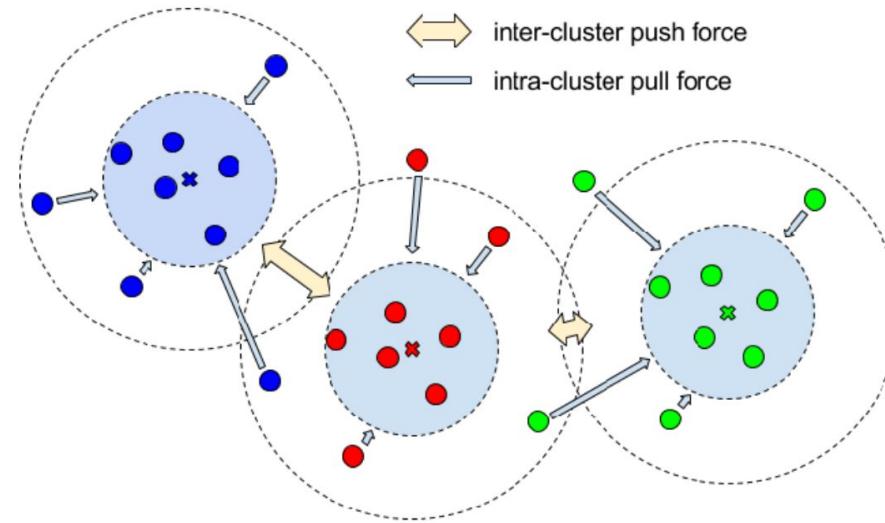
- Three component loss: pull together points that belong to the same cluster, keep distance between clusters, and regularization

$$L = \alpha L_{var} + \beta L_{dist} + \gamma L_{reg},$$

$$L_{var} = \frac{1}{C} \sum_{c=1}^C \frac{1}{N_c} \sum_{i=1}^{N_c} [\max(0, \|\mu_c - x_i\| - \delta_v)]^2$$

$$L_{dist} = \frac{1}{C(C-1)} \sum_{\substack{c_A, c_B=1 \\ c_A \neq c_B}} [\max(0, 2\delta_d - \|\mu_{c_A} - \mu_{c_B}\|)]^2$$

$$L_{reg} = \frac{1}{C} \sum_{c=1}^C \|\mu_c\|$$



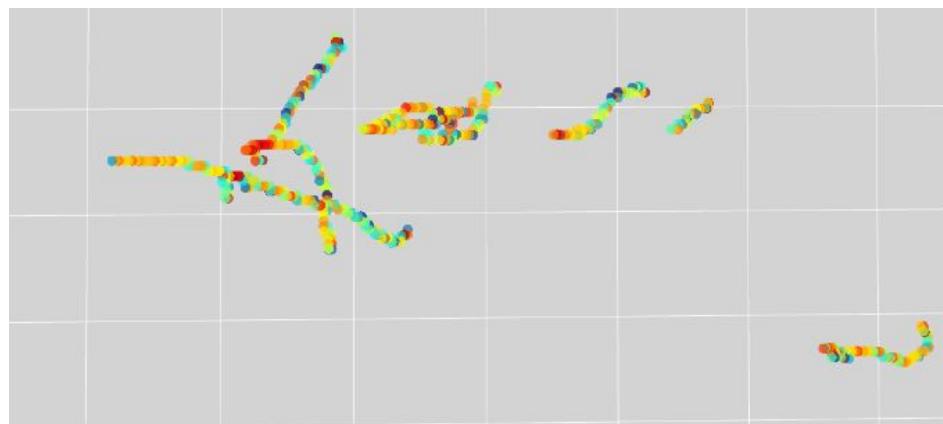
Equation credit: Dae Hyun K. @ Stanford

Image credit: arXiv 1708.02551

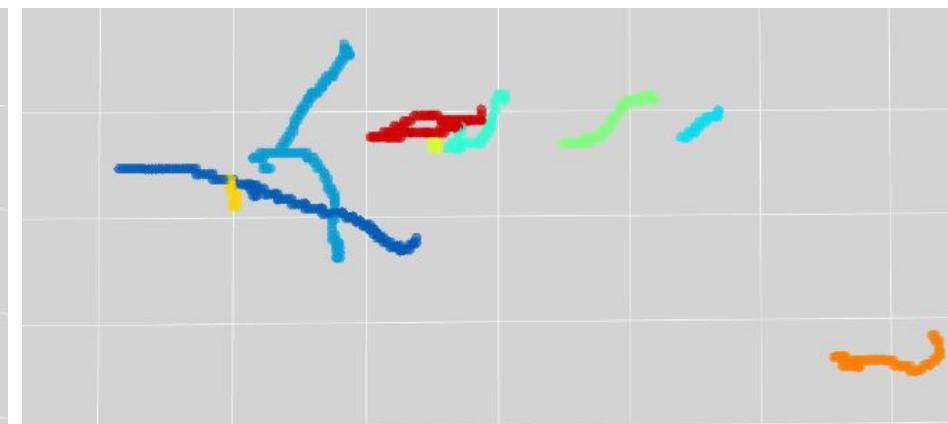
## Stage 2: Particle & Interaction Clustering

### Instance+Semantic Segmentation

- Three component loss: pull together points that belong to the same cluster, keep distance between clusters, and regularization



**Input:** 3D pixel energy depositions



**Output:** 3D pixel clusters  
(DBScan in hyperspace)

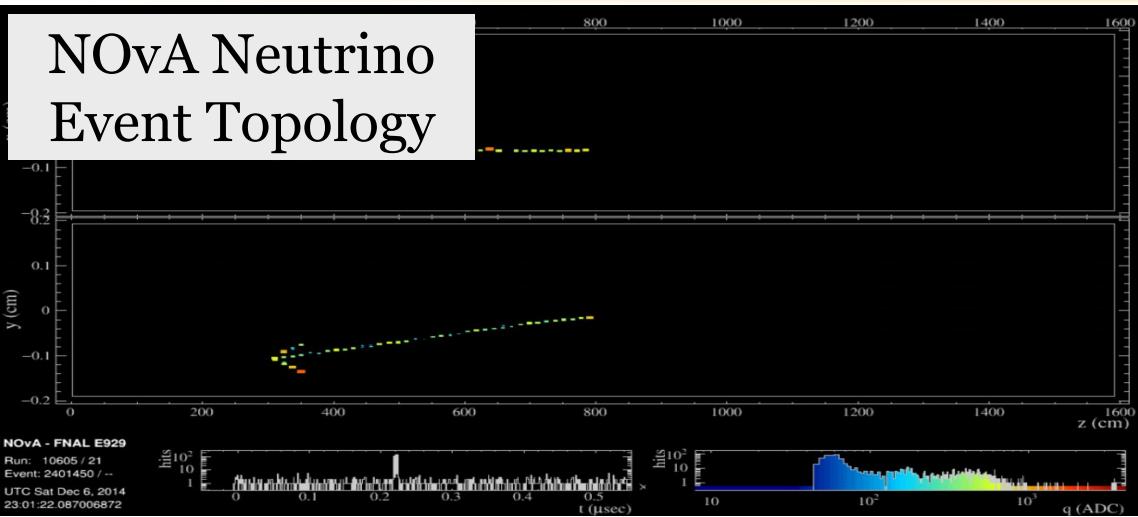
## Image Classification?

# Machine Learning & Computer Vision in Neutrino Physics

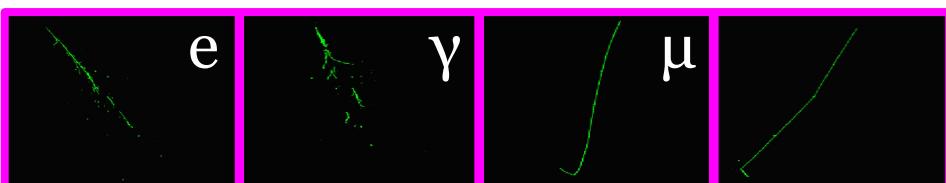
## Image Classifications: a lot of applications

SI AC

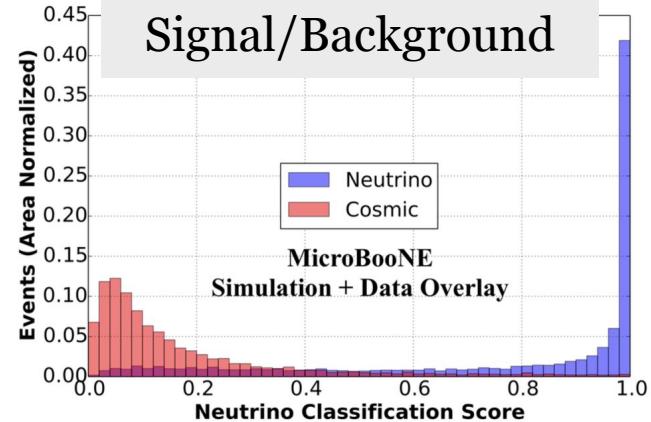
### NOvA Neutrino Event Topology



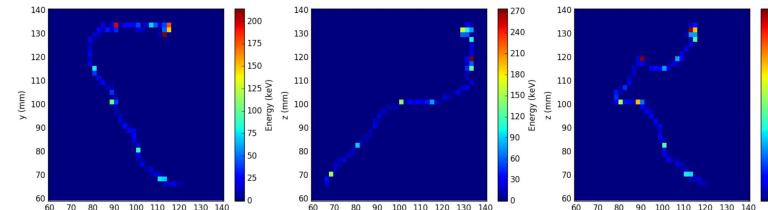
### LArLIAT Particle Type Identification



### MicroBooNE Signal/Background



### NEXT Signal vs. Background



## Image Classifications: a lot of applications

Especially great for: “**a rare event in a quiet detector**”

- **Quiet** = can assume “almost always neutrino”
  - e.g.) no cosmic-ray background
- **Rare** = “only 1 neutrino”

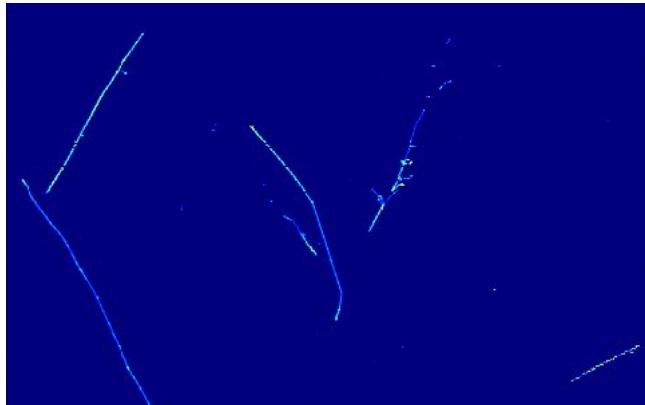
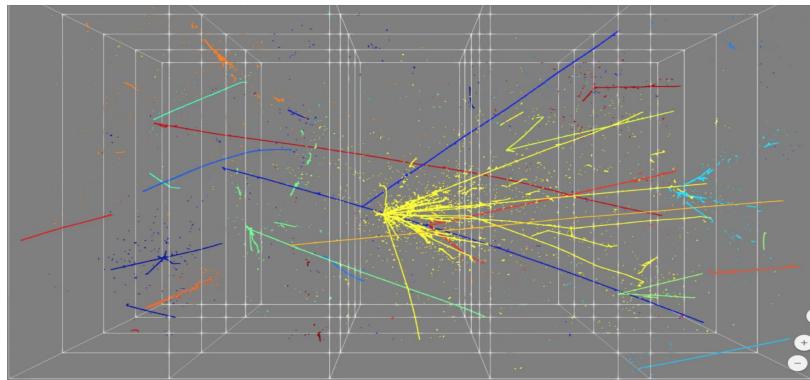
## Image Classifications: a lot of applications

Especially great for: “**a rare event in a quiet detector**”

- **Quiet** = can assume “almost always neutrino”
  - e.g.) no cosmic-ray background
- **Rare** = “only 1 neutrino”
  - the same “image classification architecture” can be applied for...
    - neutrino flavor (topology) classification
    - energy regression (image to one FP32 value)
    - vertex regression (image to three FP32 value)
    - etc. ...

## Image Classifications: a lot of applications

Especially great for: “**a rare event in a quiet detector**”

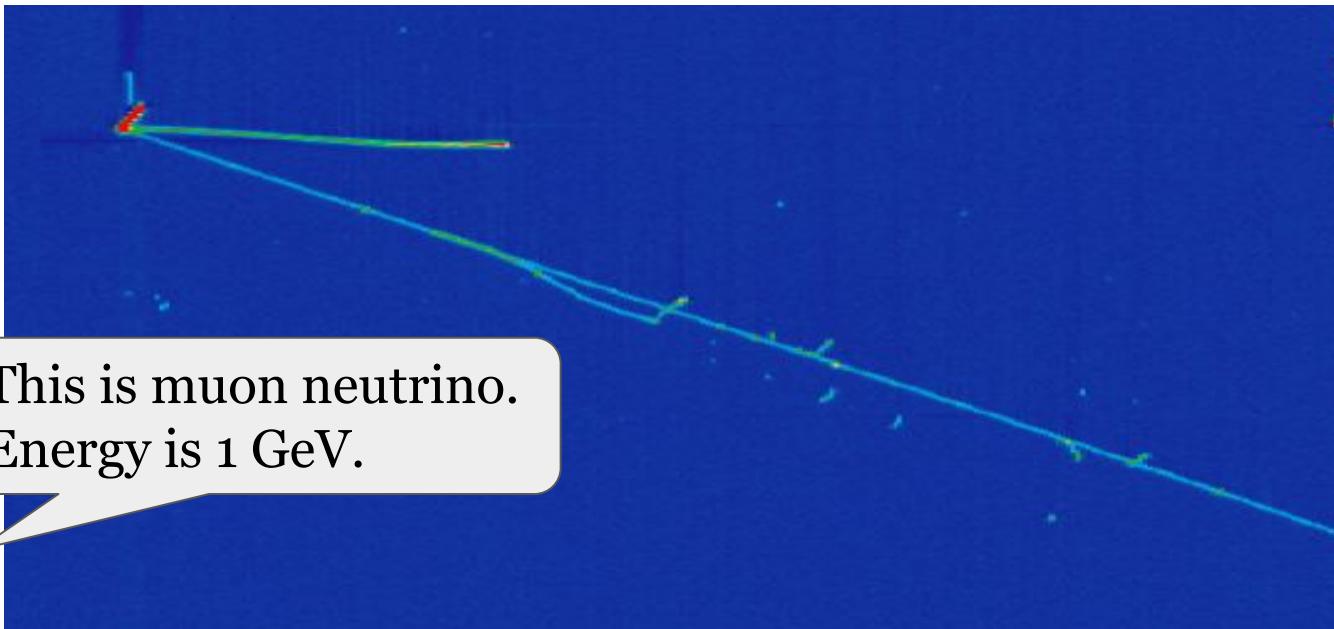


... **but most of LArTPC detectors are not ...**

- MicroBooNE, ICARUS, SBND, ProtoDUNE ... physics in next 5 years
  - Busy: typically dozens of cosmic rays in each event
- DUNE-ND
  - Not rare (busy): a dozen of neutrino interaction pile-up in each event

## Why Data Reconstruction

Image classification/regression: straight to “flavour & energy”



This is muon neutrino.  
Energy is 1 GeV.

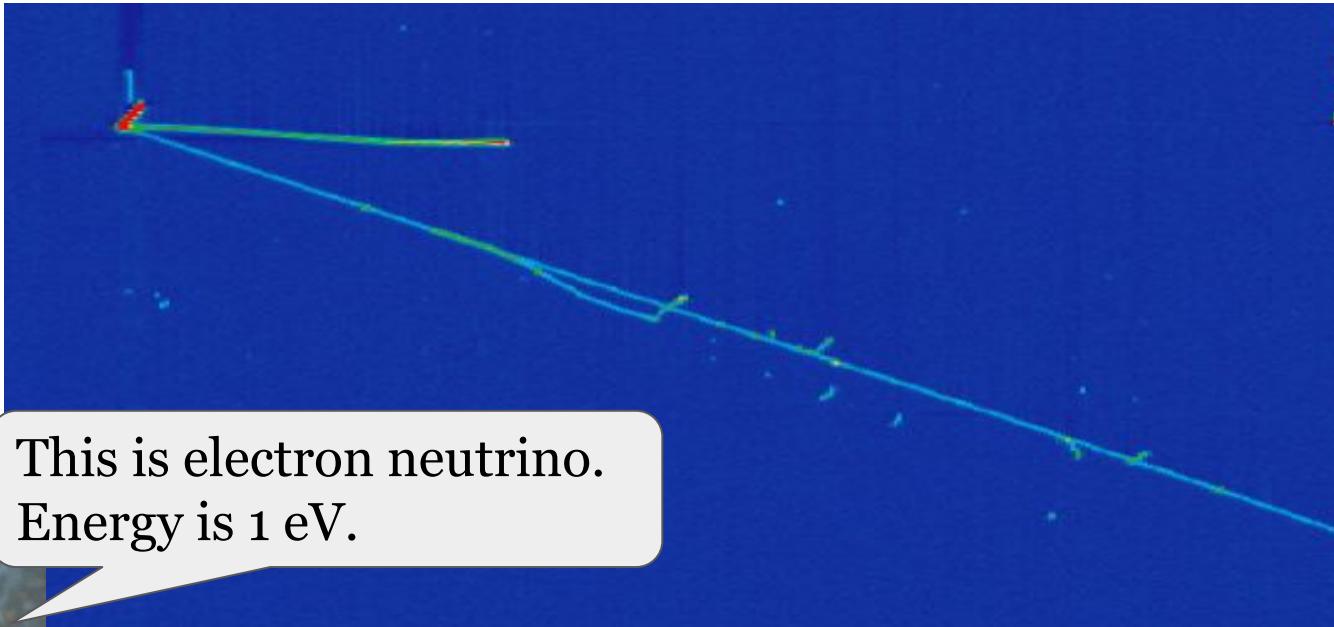


# Machine Learning & Computer Vision in Neutrino Physics

## Why Data Reconstruction

SLAC

... but also challenging: a huge single-step of information reduction



... would be nice to know why you thought so ...

# Machine Learning & Computer Vision in Neutrino Physics

## Why Data Reconstruction

SLAC

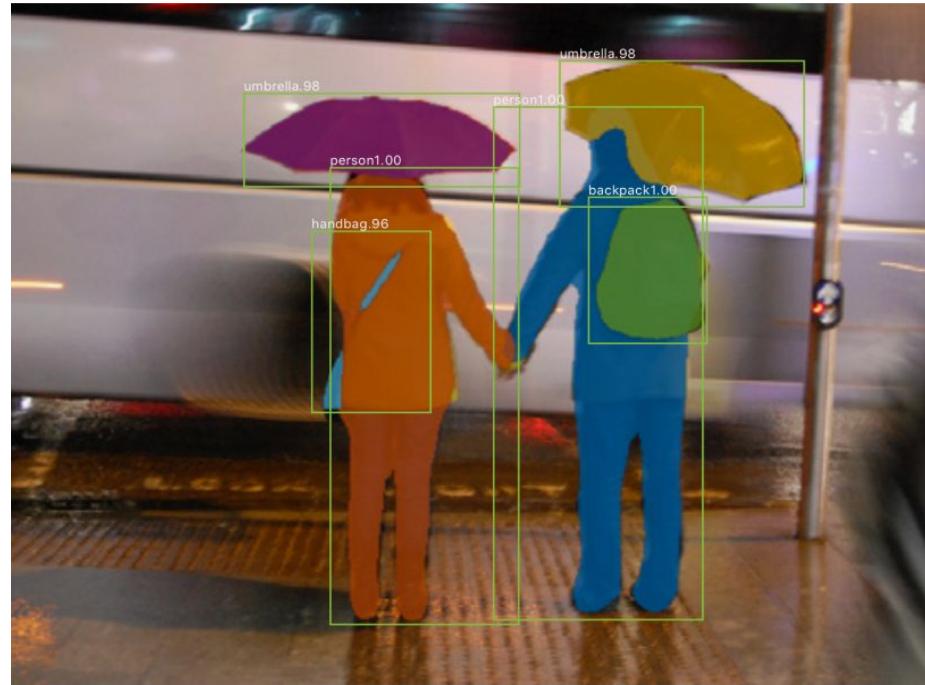
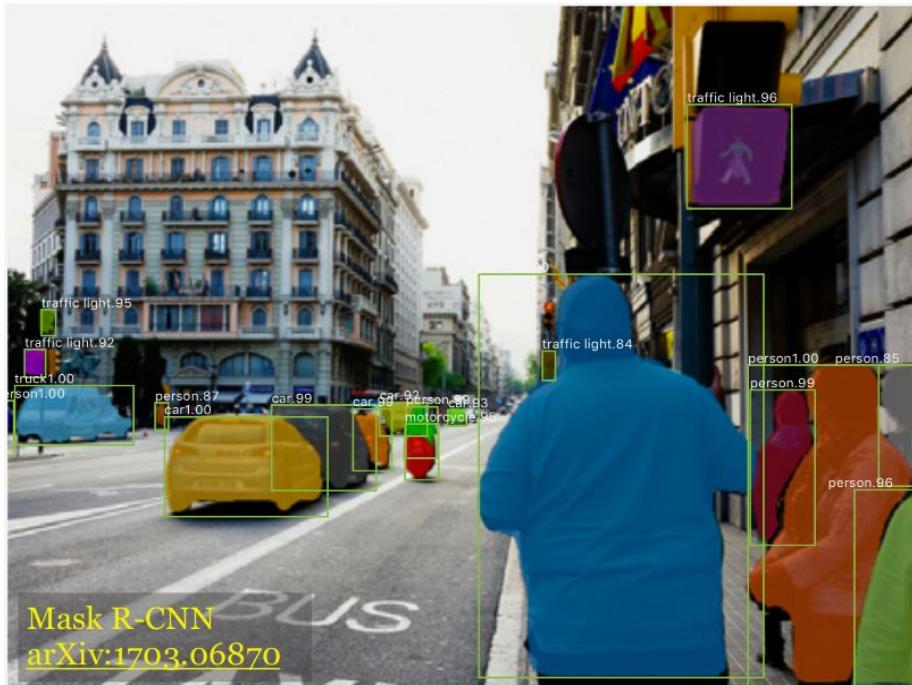


Image Context Identification

# Machine Learning & Computer Vision in Neutrino Physics

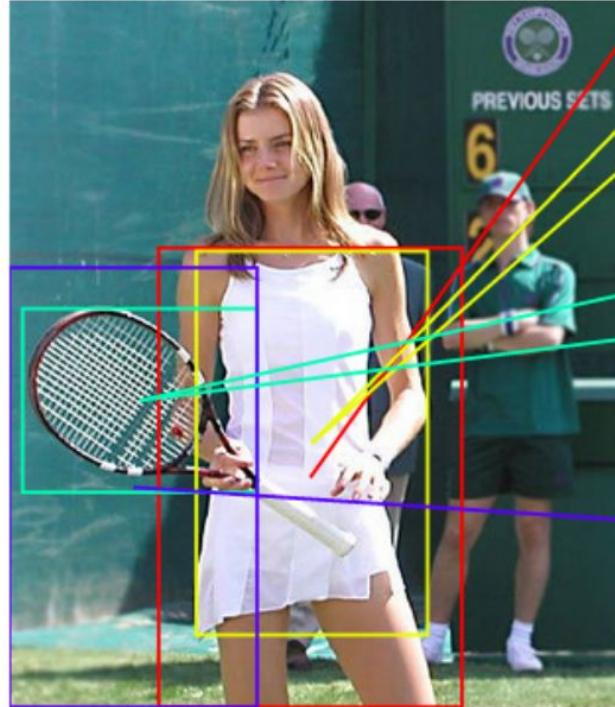
## Why Data Reconstruction

SLAC



NeuralTalk  
[github:karpathy/neuraltalk2](https://github.com/karpathy/neuraltalk2)

"girl in pink dress is jumping in air."



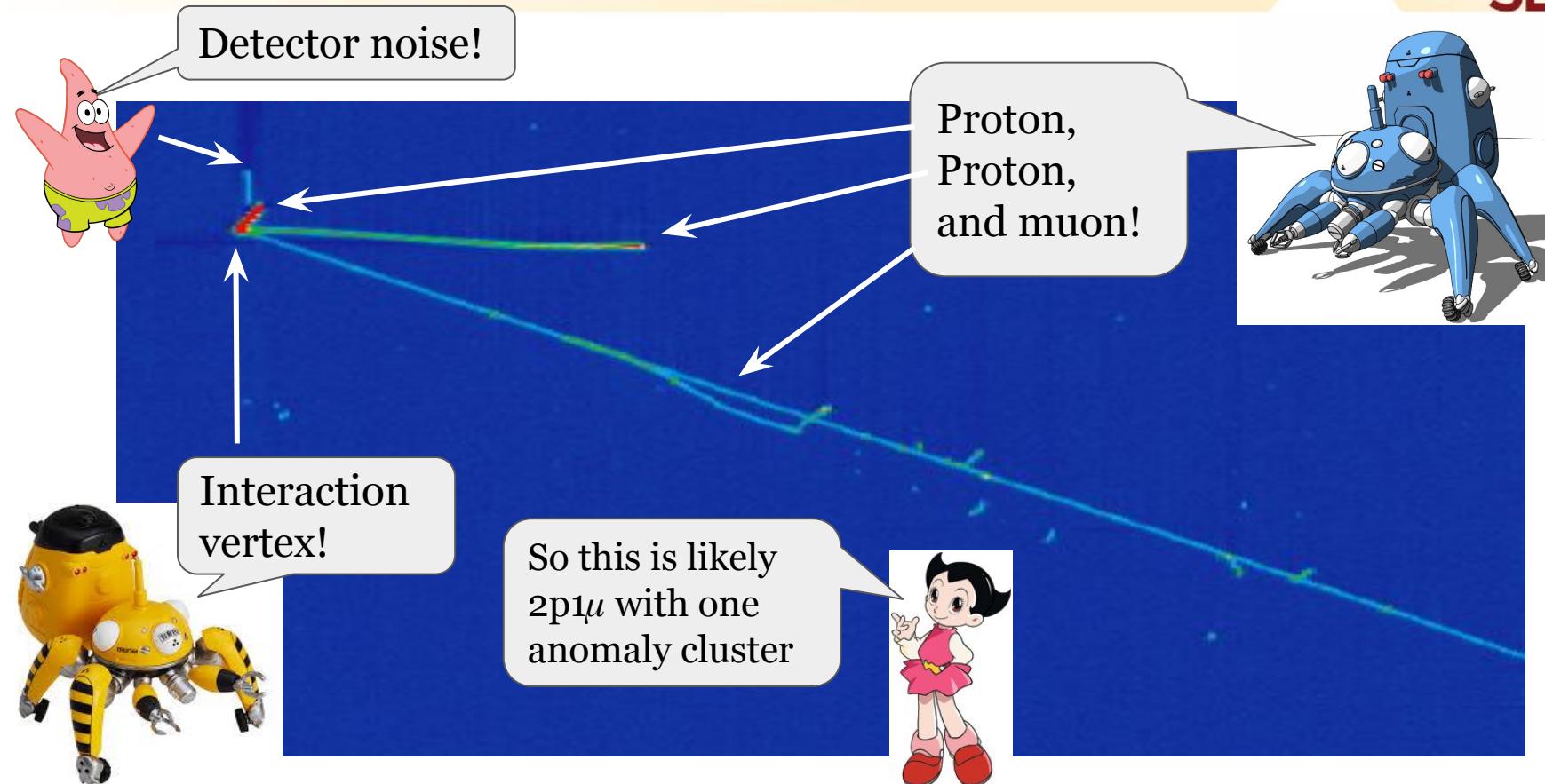
1.12 woman
-0.28 in
1.23 white
1.45 dress
0.06 standing
-0.13 with
3.58 tennis
1.81 racket
0.06 two
0.05 people
-0.14 in
0.30 green
-0.09 behind
-0.14 her

Image Context Correlation/Hierarchy Analysis

# Machine Learning & Computer Vision in Neutrino Physics

## Why Data Reconstruction

SLAC



## Segmentation Data

# Machine Learning & Computer Vision in Neutrino Physics

## Object Detection & Semantic Segmentation

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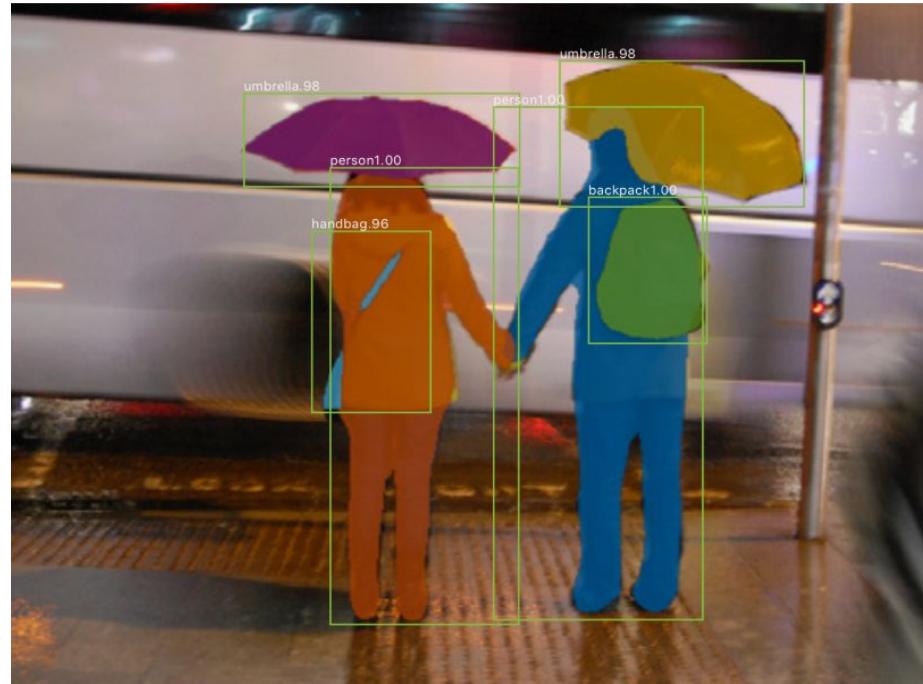
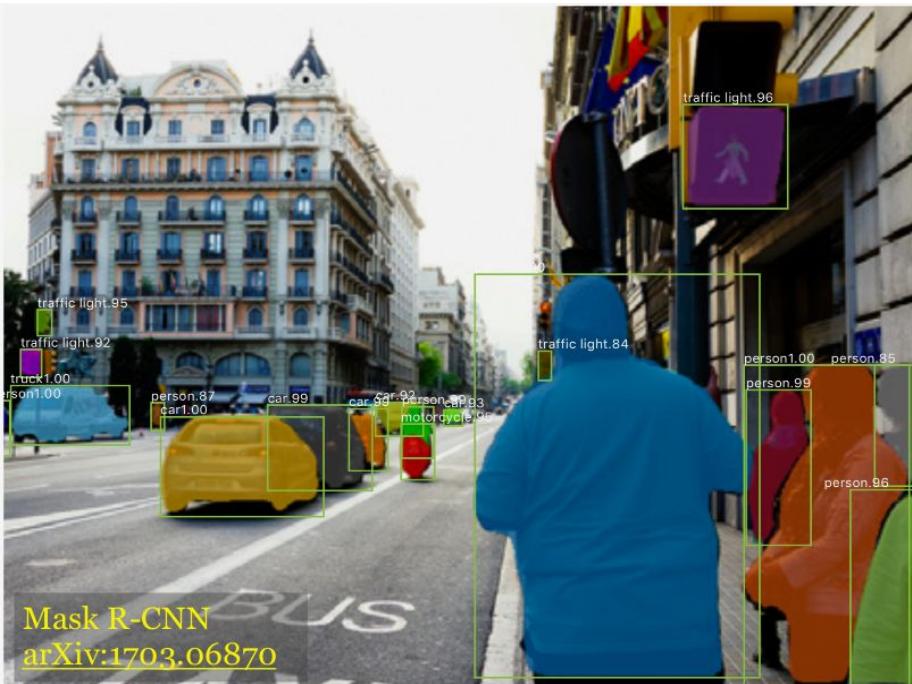


Image Context Identification

# Machine Learning & Computer Vision in Neutrino Physics

## Hierarchy and Correlation of Context

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NeuralTalk  
[github:karpathy/neuraltalk2](https://github.com/karpathy/neuraltalk2)

"girl in pink dress is jumping in air."

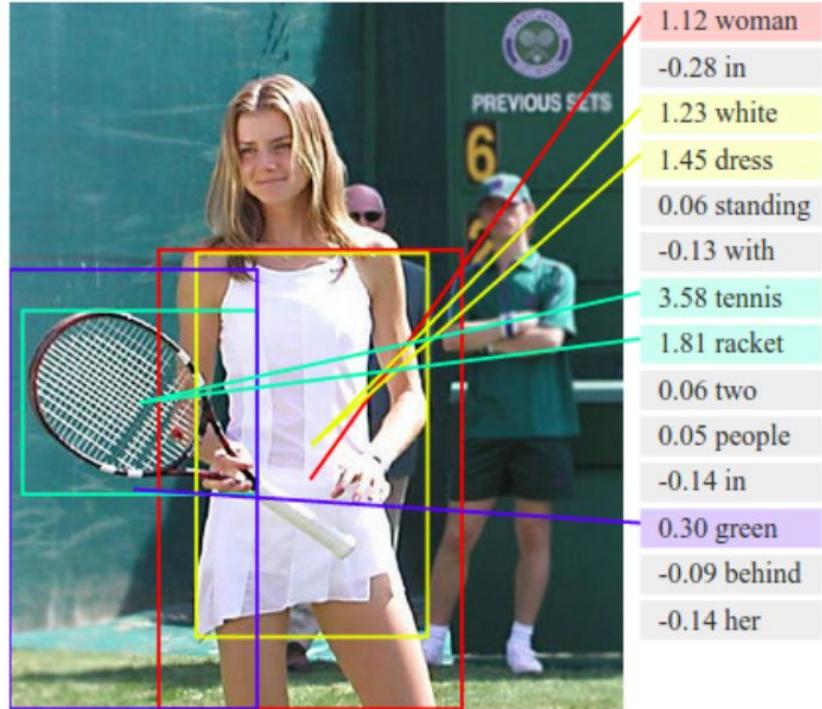
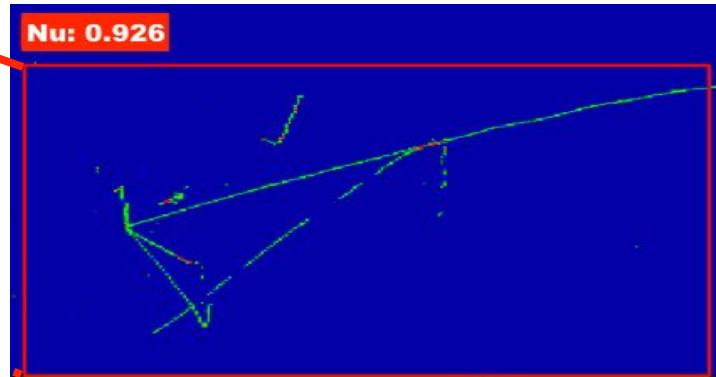
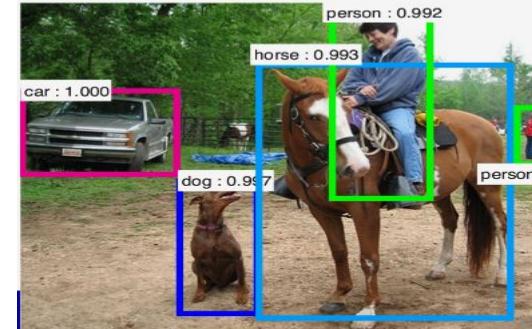
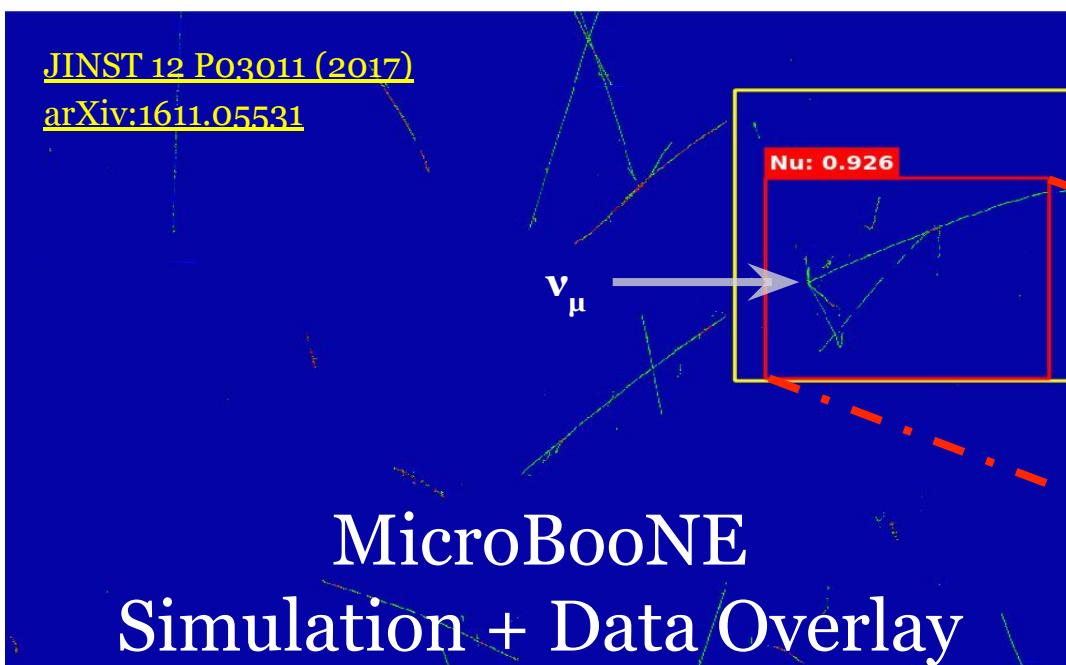


Image Context Correlation/Hierarchy Analysis

## Object Detection for Neutrino ID

### Neutrino Detection w/ R-CNN (MicroBooNE LArTPC)



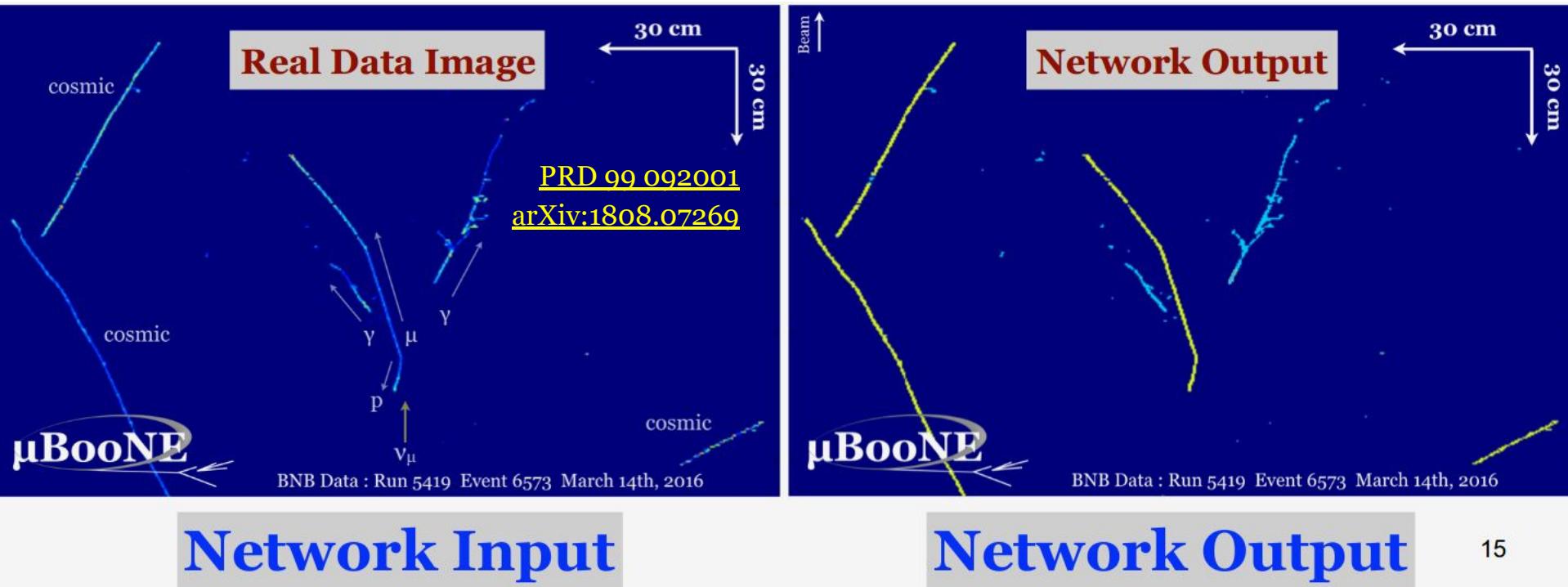
**Task:** propose a rectangular box that contains neutrino interaction (location & size)

# Machine Learning & Computer Vision in Neutrino Physics

## Semantic Segmentation for Pixel-level Particle ID

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Separate electron/positron energy depositions from other types at raw waveform level.  
Helps the downstream clustering algorithms (**data/sim comp.** @ [arxiv:1808.07269](https://arxiv.org/abs/1808.07269))

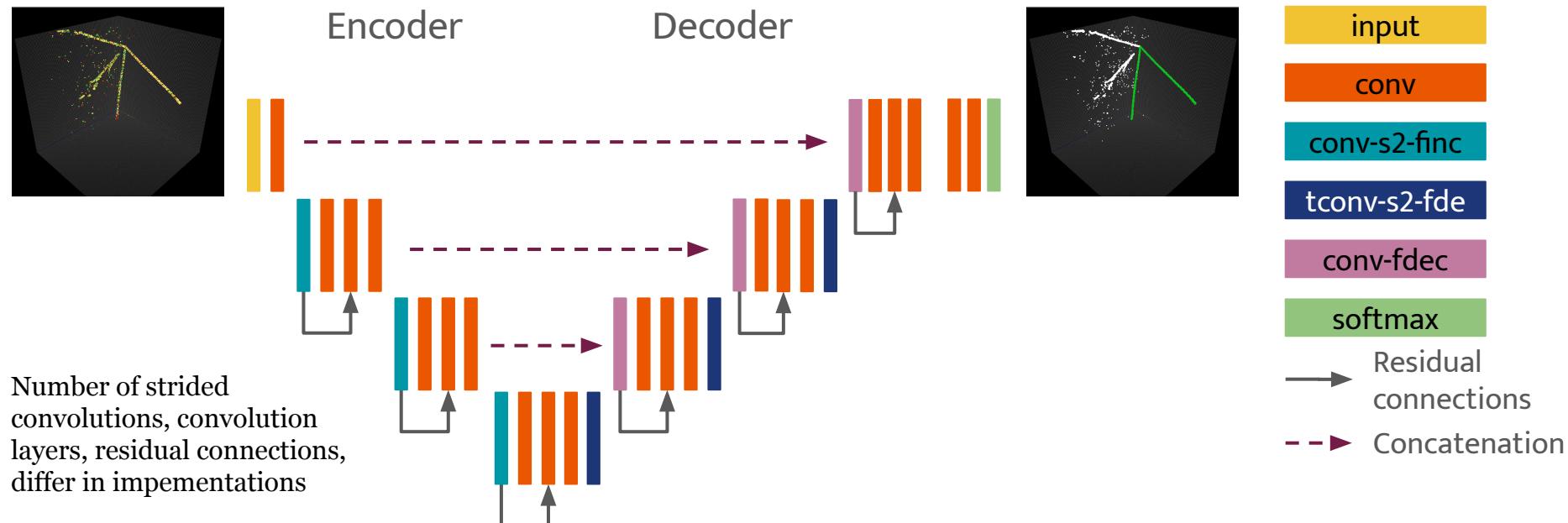


# Machine Learning & Computer Vision in Neutrino Physics

## Semantic Segmentation for Pixel-level Particle ID

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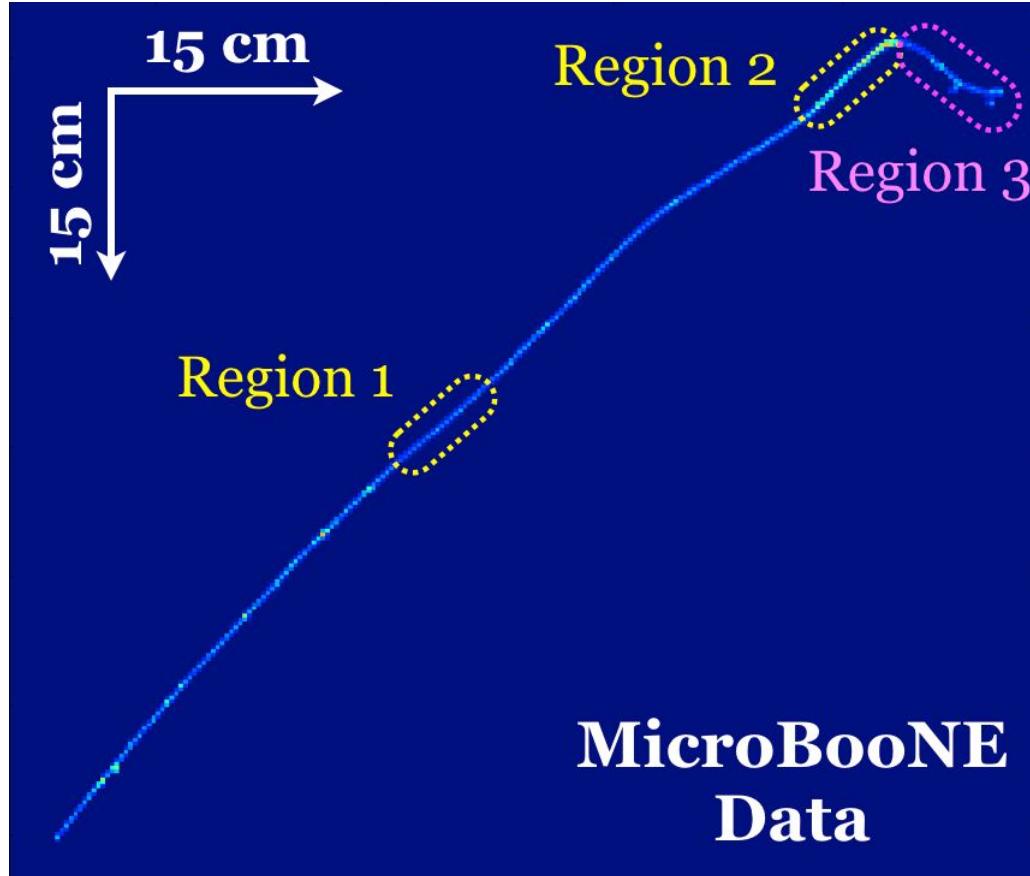
### Architecture: U-Net + Residual Connections



# Machine Learning & Computer Vision in Neutrino Physics

## Fun Playing with Semantic Segmentation

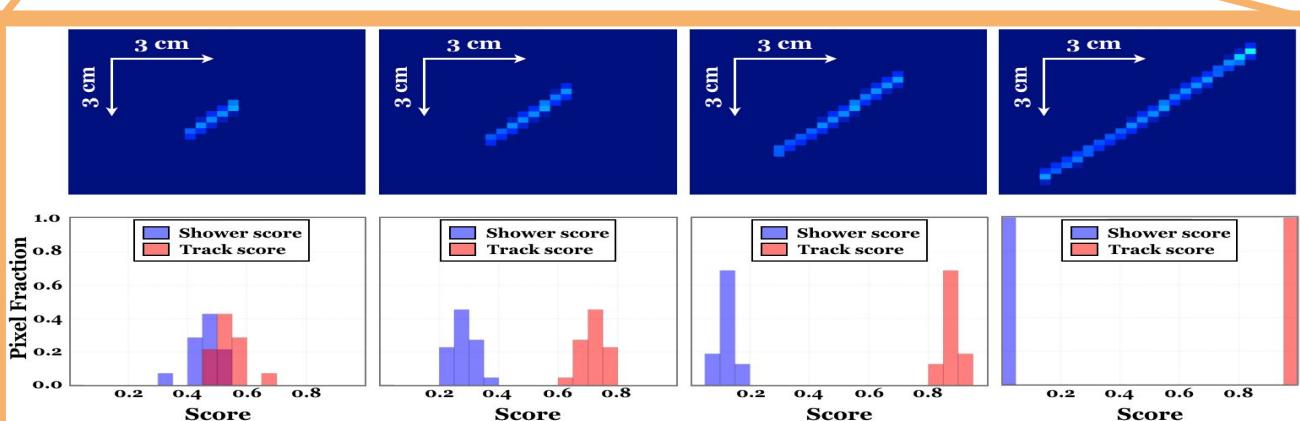
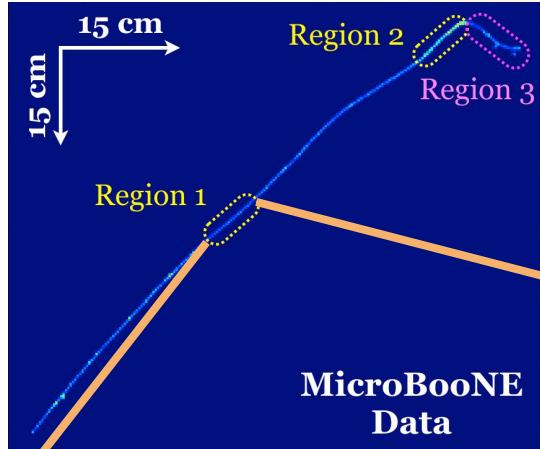
SLAC



# Machine Learning & Computer Vision in Neutrino Physics

## Fun Playing with Semantic Segmentation

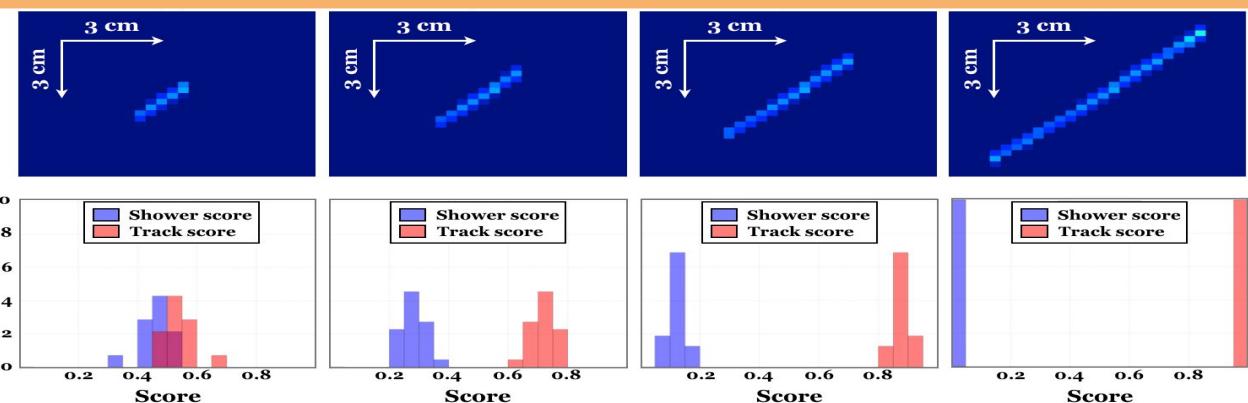
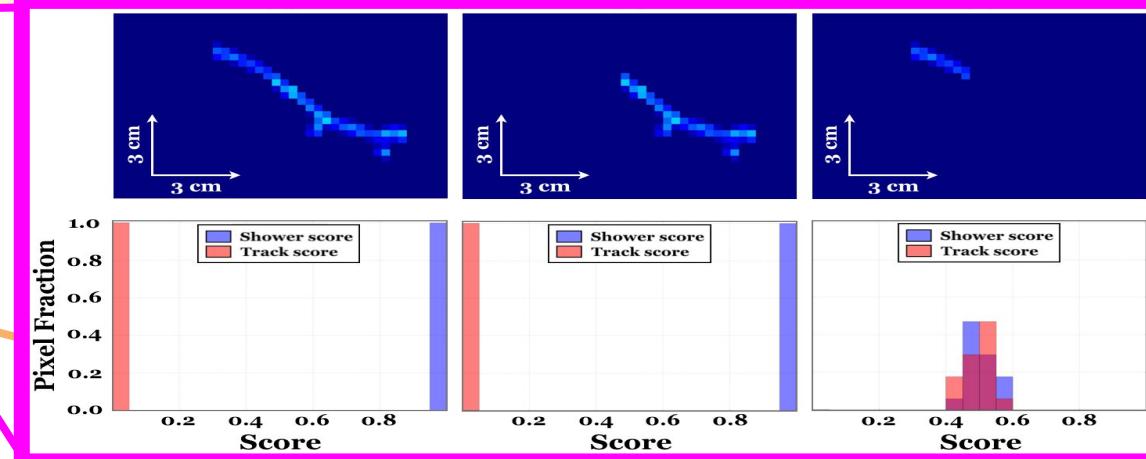
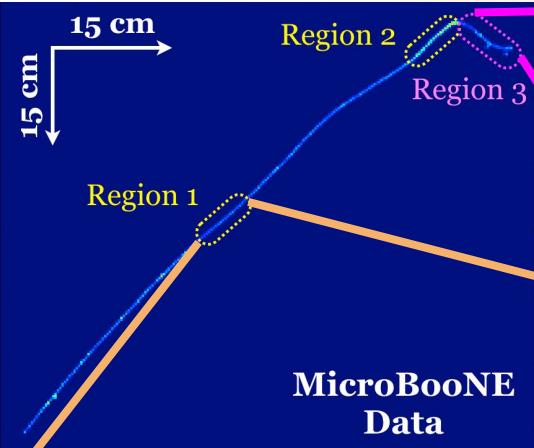
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**Localized features** at the pixel-level are useful to inspect **correlation of data features & algorithm responses**

# Machine Learning & Computer Vision in Neutrino Physics

## Fun Playing with Semantic Segmentation



**Localized features** at the pixel-level are useful to inspect **correlation of data features & algorithm responses**

## Why Neutrino

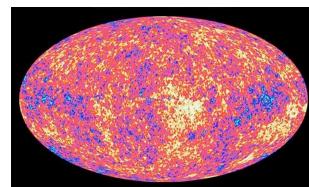
# Machine Learning & Computer Vision in Neutrino Physics

## Why neutrinos?

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### Neutrinos are everywhere!

... which makes them the **natural probe to the universe and its history**



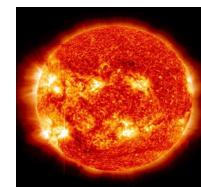
Early Universe



Supernova



AGN



Stars



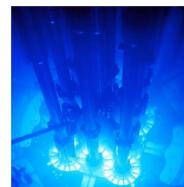
Atmosphere



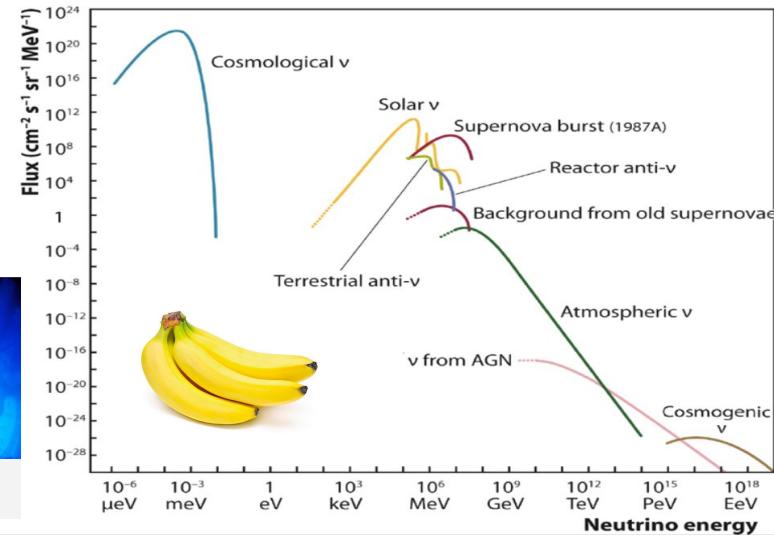
Earth (planets)



Accelerators



Reactors



**Want to detect & understand more of them**

**First**, understand how neutrinos travel over spacetime (neutrino oscillations)

## Neutrino Detectors: What's Important

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### Neutrino Oscillation Measurement

Use a neutrino source (flavour X), measure flavour Y at the detector

### What's important?

Three important detector features for oscillation measurement

$$P(\nu_\mu \rightarrow \nu_e) = \sin^2 2\theta \sin^2 \left( \frac{1.27 \Delta m^2 L}{E_\nu} \right)$$

#### Good Energy Resolution

Precise  $E_\nu$  reduce oscillation uncertainty

#### Large Mass (scalable)

“More” statistics to measure rare physics process

#### Particle ID Capability

Better ν identification  
background rejection