

Searching of habitable alternative planets suitable for Space Colonization



Ooops, It's **not Earth**. It's *Gliese 581g* or you may call it as **Zarmina**

```
In [1]: import pandas as pd
import numpy as np

import matplotlib.pyplot as plt
import matplotlib.colors as clr
import matplotlib.cm as cm
from mpl_toolkits.mplot3d import Axes3D
import seaborn as sns

from scipy.stats.stats import pearsonr
```

Requirements Specification and analysis

The dataset(PHL's Exoplanet Catalog of the Planetary Habitability Laboratory) I've used is created and maintained by

The Planetary Habitability Laboratory, University of Puerto Rico at Arecibo

It can be found at the following url: <http://phl.upr.edu/projects/habitable-exoplanets-catalog/data/database> (<http://phl.upr.edu/projects/habitable-exoplanets-catalog/data/database>)

The PHL's Exoplanets Catalog (PHL-EC) contains observed and modeled parameters for all currently confirmed exoplanets from the Extrasolar Planets Encyclopedia and NASA Kepler candidates from the NASA Exoplanet Archive, including those potentially habitable. It also contains a few still unconfirmed exoplanets of interest. The main difference between PHL-EC and other exoplanets databases is that it contains more estimated stellar and planetary parameters, habitability assessments with various habitability metrics, planetary classifications, and many corrections. Some interesting inclusions are the identification of those stars in the Catalog of Nearby Habitable Systems (HabCat aka HabStar Catalog), the apparent size and brightness of stars and planets as seen from a vantage point (i.e. moon-Earth distance), and the location constellation of each planet.

Reasons behind choosing this dataset over any other is because of its expanded target list combining measures and modeled parameters from various sources. Hence, it provides a good metric for visualization and statistical analysis.

PHL-EC consists of 68 features and 3875 confirmed exoplanets.

```
In [2]: allExoplanets = pd.read_csv('confirmed_exoplanets.csv',low_memory=False)

print('Features, Data Points = '+str(allExoplanets.shape))

Features, Data Points = (3875, 68)
```

```
In [3]: print('All Features of PHL-EC:\n\n')
        for i in allExoplanets:
            print("{feature}".format(feature=i),sep='\t')
```

All Features of PHL-EC:

- P. Name
- P. Name Kepler
- P. Name KOI
- P. Zone Class
- P. Mass Class
- P. Composition Class
- P. Atmosphere Class
- P. Habitable Class
- P. Min Mass (EU)
- P. Mass (EU)
- P. Max Mass (EU)
- P. Radius (EU)
- P. Density (EU)
- P. Gravity (EU)
- P. Esc Vel (EU)
- P. SFlux Min (EU)
- P. SFlux Mean (EU)
- P. SFlux Max (EU)
- P. Teq Min (K)
- P. Teq Mean (K)
- P. Teq Max (K)
- P. Ts Min (K)
- P. Ts Mean (K)
- P. Ts Max (K)
- P. Surf Press (EU)
- P. Mag
- P. Appar Size (deg)
- P. Period (days)
- P. Sem Major Axis (AU)
- P. Eccentricity
- P. Mean Distance (AU)
- P. Inclination (deg)
- P. Omega (deg)
- S. Name
- S. Name HD
- S. Name HIP
- S. Constellation
- S. Type
- S. Mass (SU)
- S. Radius (SU)
- S. Teff (K)
- S. Luminosity (SU)
- S. [Fe/H]
- S. Age (Gyrs)
- S. Appar Mag
- S. Distance (pc)
- S. RA (hrs)
- S. DEC (deg)
- S. Mag from Planet
- S. Size from Planet (deg)
- S. No. Planets
- S. No. Planets HZ
- S. Hab Zone Min (AU)
- S. Hab Zone Max (AU)
- P. HZD
- P. HZC
- P. HZA
- P. HZI
- P. SPH
- P. Int ESI
- P. Surf ESI
- P. ESI
- S. HabCat
- P. Habitable
- P. Hab Moon
- P. Confirmed
- P. Disc. Method
- P. Disc. Year

```
In [4]: ## object type scientific notation ke float korar portion ## Last checkpoint

allExoplanets['P. SFlux Max (EU)'] = pd.to_numeric(allExoplanets['P. SFlux Max (EU)'],errors='coerce')

allExoplanets['P. SFlux Mean (EU)'] = pd.to_numeric(allExoplanets['P. SFlux Mean (EU)'],errors='coerce')

allExoplanets['P. SFlux Min (EU)'] = pd.to_numeric(allExoplanets['P. SFlux Min (EU)'],errors='coerce')
```

```
In [5]: cat = len(allExoplanets.select_dtypes(include=['object']).columns)
num = len(allExoplanets.select_dtypes(include=['int64','float64']).columns)
print('Features of allExoplanets consists of ', cat, 'categorical', ' and ',
      num, 'numerical features')
print('\n\nCategorical Features:\n ')
for i in allExoplanets.select_dtypes(include=['object']).columns:
    print("{feature}".format(feature=i),end='\t')
print('\n\nNumerical Features:\n ')
for i in allExoplanets.select_dtypes(include=['int64','float64']).columns:
    print("{feature}".format(feature=i),end='\t')
```

Features of allExoplanets consists of 14 categorical and 54 numerical features

Categorical Features:

P. Name P. Name Kepler P. Zone Class P. Mass Class P. Composition Class P. Atmosphere Class P. Habitable Class S. Name S. Name HD S. Name HIP S. Constellation S. Type P. Disc. Method P. Discovery Year

Numerical Features:

P. Name KOI P. Min Mass (EU) P. Mass (EU) P. Max Mass (EU) P. Radius (EU) P. Density (EU) P. Gravity (EU) P. Esc Vel (EU) P. SFlux Min (EU) P. SFlux Mean (EU) P. SFlux Max (EU) P. Teq Min (K) P. Teq Mean (K) P. Teq Max (K) P. Ts Min (K) P. Ts Mean (K) P. Ts Max (K) P. Surface Press (EU) P. Mag P. Appar Size (deg) P. Period (days) P. Sem Major Axis (AU) P. Eccentricity P. Mean Distance (AU) P. Inclination (deg) P. Omega (deg) S. Mass (SU) S. Radius (SU) S. Temperature (K) S. Luminosity (SU) S. [Fe/H] S. Age (Gyrs) S. Appar Mag S. Distance (pc) S. RA (hrs) S. DEC (deg) S. Mag from Planet S. Size from Planet (deg) S. No. Planets S. No. Planet s HZ S. Hab Zone Min (AU) S. Hab Zone Max (AU) P. HZD P. HZC P. HZA P. HZI P. SPH P. Int ESI P. Surf ESI P. ESI S. HabCat P. Habitable P. Hab Moon P. Confirmed

correction:

P. Habitable is also a Categorical Variable, since it has only two unique values 0,1 indicates Yes and No respectively

```
In [6]: desc = pd.DataFrame()
for c in allExoplanets:
    desc[c]=(allExoplanets[c].describe())
desc.head()
```

Out[6]:

	P. Name	P. Name Kepler	P. Name KOI	P. Zone Class	P. Mass Class	P. Composition Class	P. Atmosphere Class	P. Habitable Class	P. Min Mass (EU)	P. Mass (EU)	...	P. SPH	P. Int ESI	P. Surf ESI
count	3875	2328	933.0	3829	3869	3834	3790	3875	1148.0	3842.0	...	1801.0	3875.0	3875.0
unique	3875	2328	NaN	3	6	5	3	5	NaN	NaN	...	NaN	NaN	NaN
top	HD 30177 c	Kepler-1410 b	NaN	Hot	Jovian	gas	metals-rich	non-habitable	NaN	NaN	...	NaN	NaN	NaN
freq	1	1	NaN	3251	1273	2074	2372	3820	NaN	NaN	...	NaN	NaN	NaN

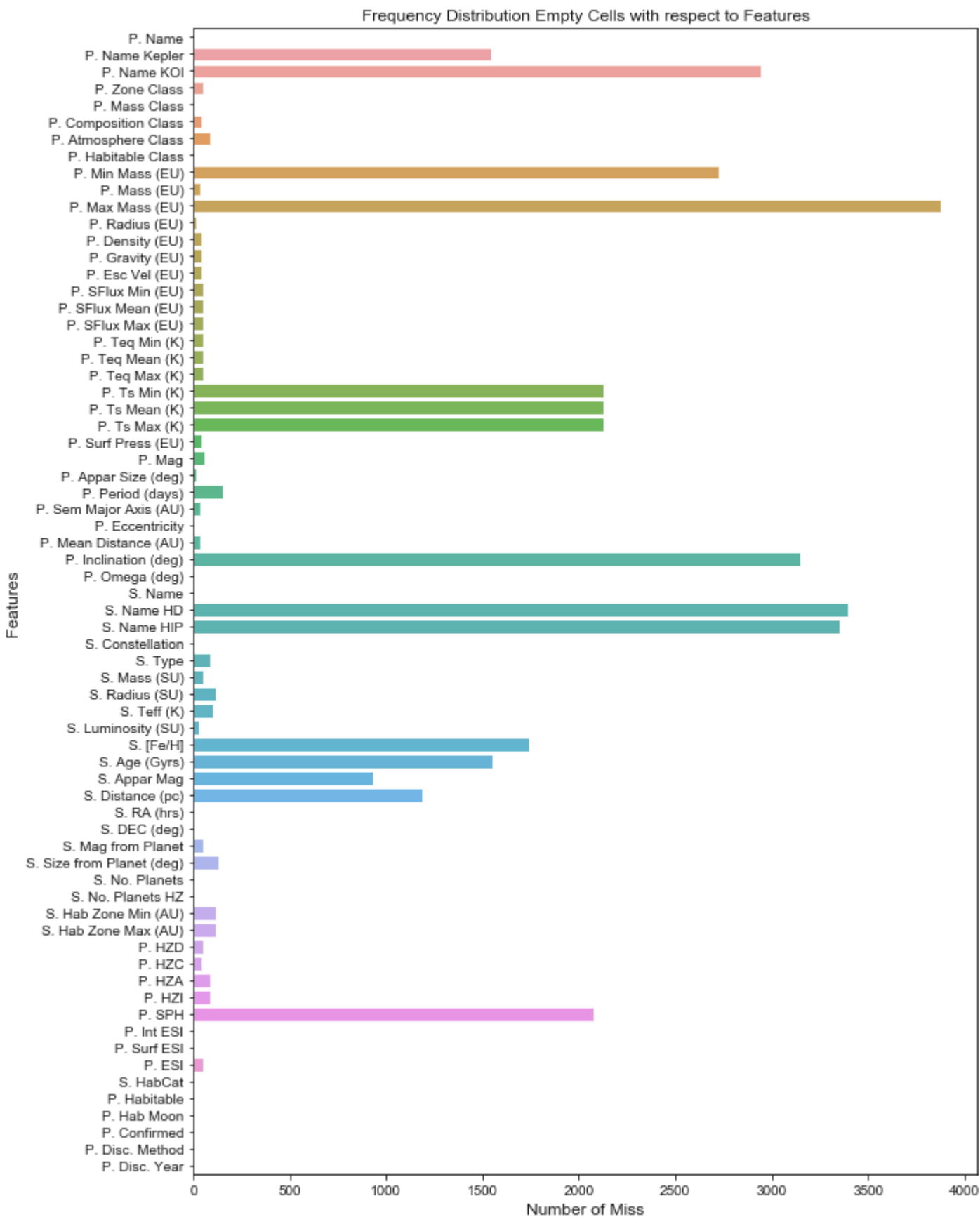
4 rows × 68 columns



Percentage of Null Value in each feature of the given data set

```
In [7]: #allExoplanets.isnull()
####print(allExoplanets.isnull().sum() * 100 / len(allExoplanets),end='\t\t\t\t')
f, ax = plt.subplots(figsize=(10, 15))

stat_count = allExoplanets.isnull().sum()
sns.set(style="darkgrid")
sns.barplot(stat_count.values,stat_count.index, alpha=0.9)
plt.title('Frequency Distribution Empty Cells with respect to Features')
plt.xlabel('Number of Miss', fontsize=12)
plt.ylabel('Features', fontsize=12)
plt.show()
```

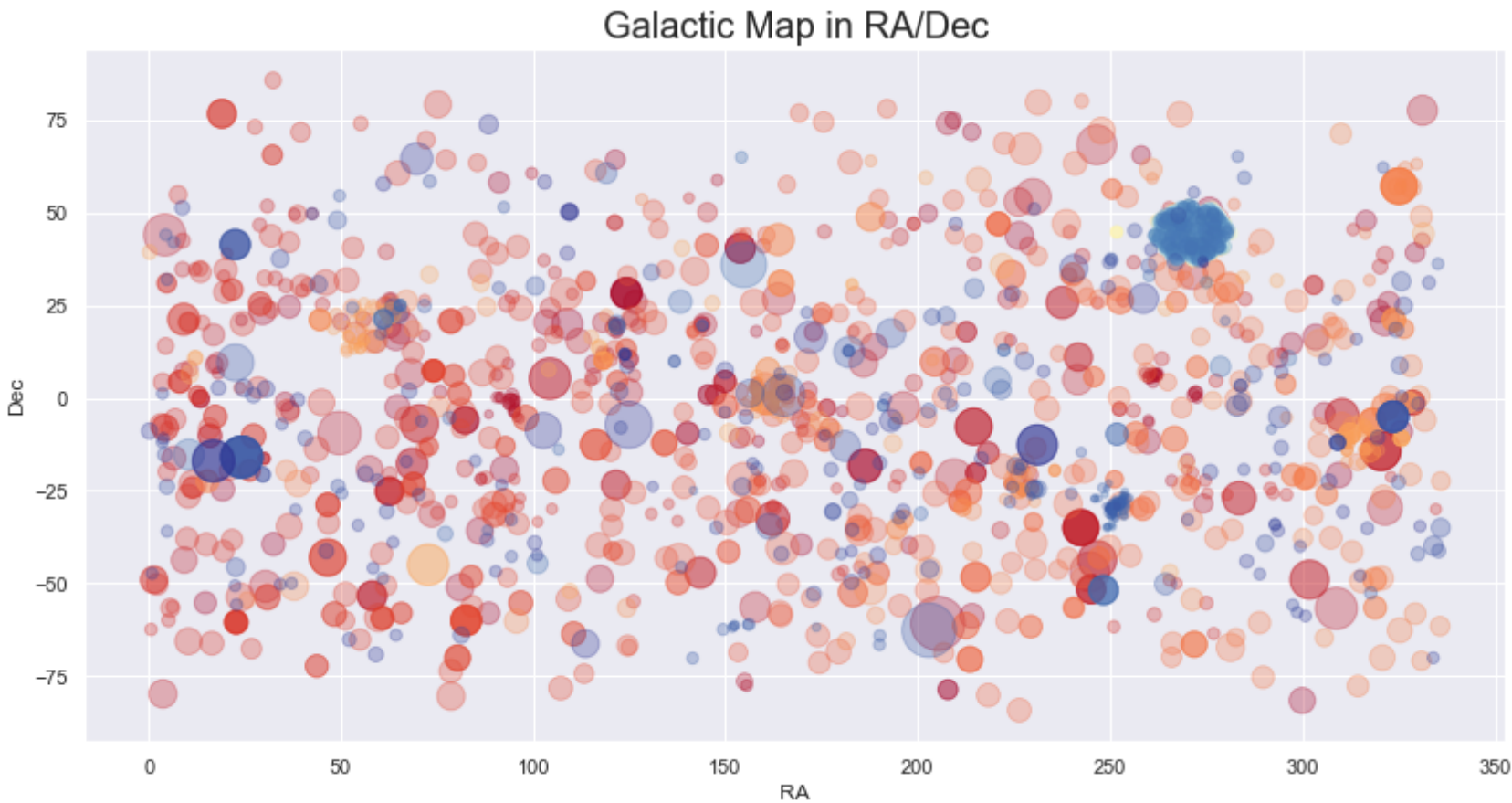


Some visualisations regarding the database

```
In [8]: x = allExoplanets['S. RA (hrs)']*14
y = allExoplanets['S. DEC (deg)']
#area = np.pi * allExoplanets['S. Radius (SU)']**2
csp=plt.cm.RdYlBu(np.linspace(0,1,len(allExoplanets)))

fig,ax = plt.subplots(figsize=(14,7))
dists = allExoplanets['S. Distance (pc)']
dists.fillna(value=np.mean(dists),inplace=True)
ax=plt.scatter(x,y,s=1000/dists**.5,alpha=0.3,c=csp,cmap=cm.coolwarm)
#ax=plt.scatter(x, y, s=area, c = colors, cmap = colormap, alpha=0.3)
plt.xlabel('RA')
plt.ylabel('Dec')
plt.title('Galactic Map in RA/Dec',size=20)
plt.figure(1,figsize=(16,12))

plt.show()
```



```
In [9]: ## x = allExoplanets['S. RA (hrs)']*14
y = allExoplanets['S. DEC (deg)']
#area = np.pi * allExoplanets['S. Radius (SU)']**2
csp=plt.cm.RdBu(np.linspace(0,1,len(allExoplanets)))

fig,ax = plt.subplots()
dists = allExoplanets['S. Distance (pc)']
dists.fillna(value=np.mean(dists),inplace=True)
ax=plt.scatter(x,y,s=1000/dists,alpha=0.3,c=csp,cmap=cm.coolwarm)
#ax=plt.scatter(x, y, s=area, c = colors, cmap = colormap, alpha=0.3)
plt.xlabel('RA')
plt.ylabel('Dec')
plt.title('Galactic Map in RA/Dec',size=20)
plt.figure(1,figsize=(16,12))

plt.annotate('Earth', xy = (0,0),
            xytext = (560, 0),
            textcoords = 'offset points', ha = 'right', va = 'bottom',
            bbox = dict(boxstyle = 'round,pad=0.5', fc = 'blue', alpha = 0.05),
            arrowprops = dict(arrowstyle = '->',connectionstyle = 'arc3,rad=0')
            )

plt.annotate('Mars', xy = (319.3208,18.6386),
            xytext = (350, 0),
            textcoords = 'offset points', ha = 'right', va = 'bottom',
            bbox = dict(boxstyle = 'round,pad=0.5', fc = 'blue', alpha = 0.05),
            arrowprops = dict(arrowstyle = '->',connectionstyle = 'arc3,rad=0')
            )

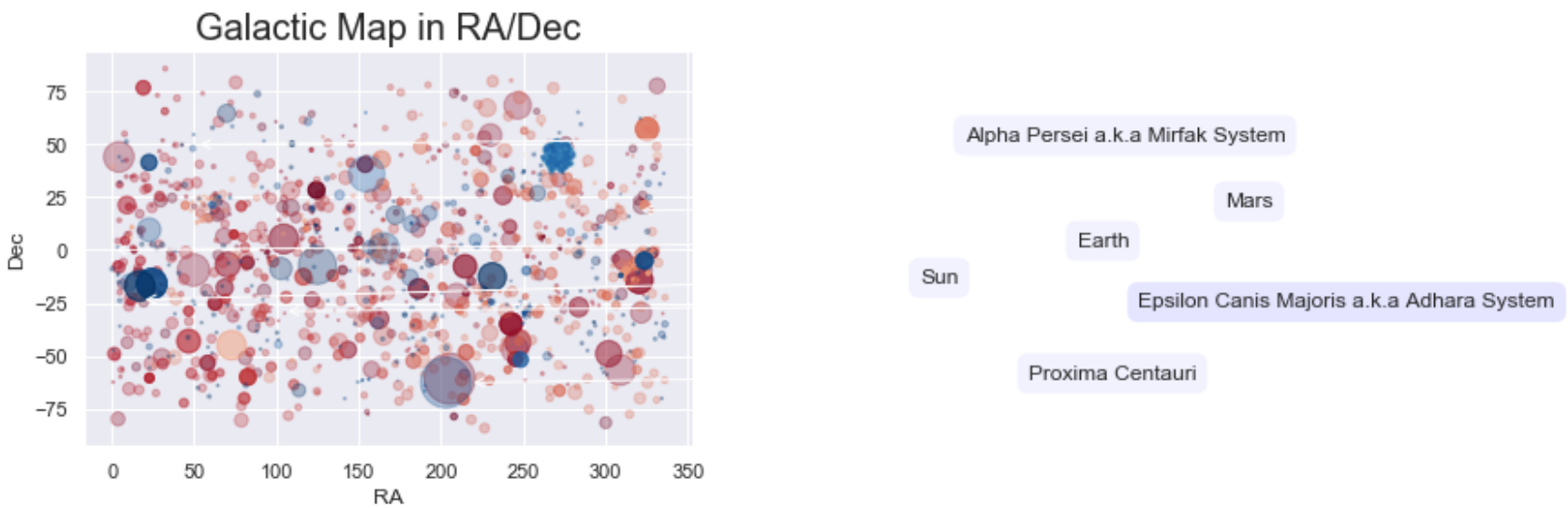
plt.annotate('Sun', xy = (18,-23.5),
            xytext = (450, 7),
            textcoords = 'offset points', ha = 'right', va = 'bottom',
            bbox = dict(boxstyle = 'round,pad=0.5', fc = 'blue', alpha = 0.05),
            arrowprops = dict(arrowstyle = '->',connectionstyle = 'arc3,rad=0')
            )

plt.annotate('Proxima Centauri', xy = (217.4292,-62.6794),
            xytext = (400, 0),
            textcoords = 'offset points', ha = 'right', va = 'bottom',
            bbox = dict(boxstyle = 'round,pad=0.5', fc = 'blue', alpha = 0.05),
            arrowprops = dict(arrowstyle = '->',connectionstyle = 'arc3,rad=0')
            )

plt.annotate('Alpha Persei a.k.a Mirfak System', xy = (51.0792,49.8611),
            xytext = (600, 0),
            textcoords = 'offset points', ha = 'right', va = 'bottom',
            bbox = dict(boxstyle = 'round,pad=0.5', fc = 'blue', alpha = 0.05),
            arrowprops = dict(arrowstyle = '->',connectionstyle = 'arc3,rad=0')
            )

plt.annotate('Epsilon Canis Majoris a.k.a Adhara System', xy = (104.6583,-28.9719),
            xytext = (700, 0),
            textcoords = 'offset points', ha = 'right', va = 'bottom',
            bbox = dict(boxstyle = 'round,pad=0.5', fc = 'blue', alpha = 0.1),
            arrowprops = dict(arrowstyle = '->',connectionstyle = 'arc3,rad=0')
            )

plt.show()
```



Visualization

- Size

Size is denoting the relative distance from earth. actual size is normalized for better visualisation

- Color

color incrementation is used for better visualisation and it increments over all datapoints.

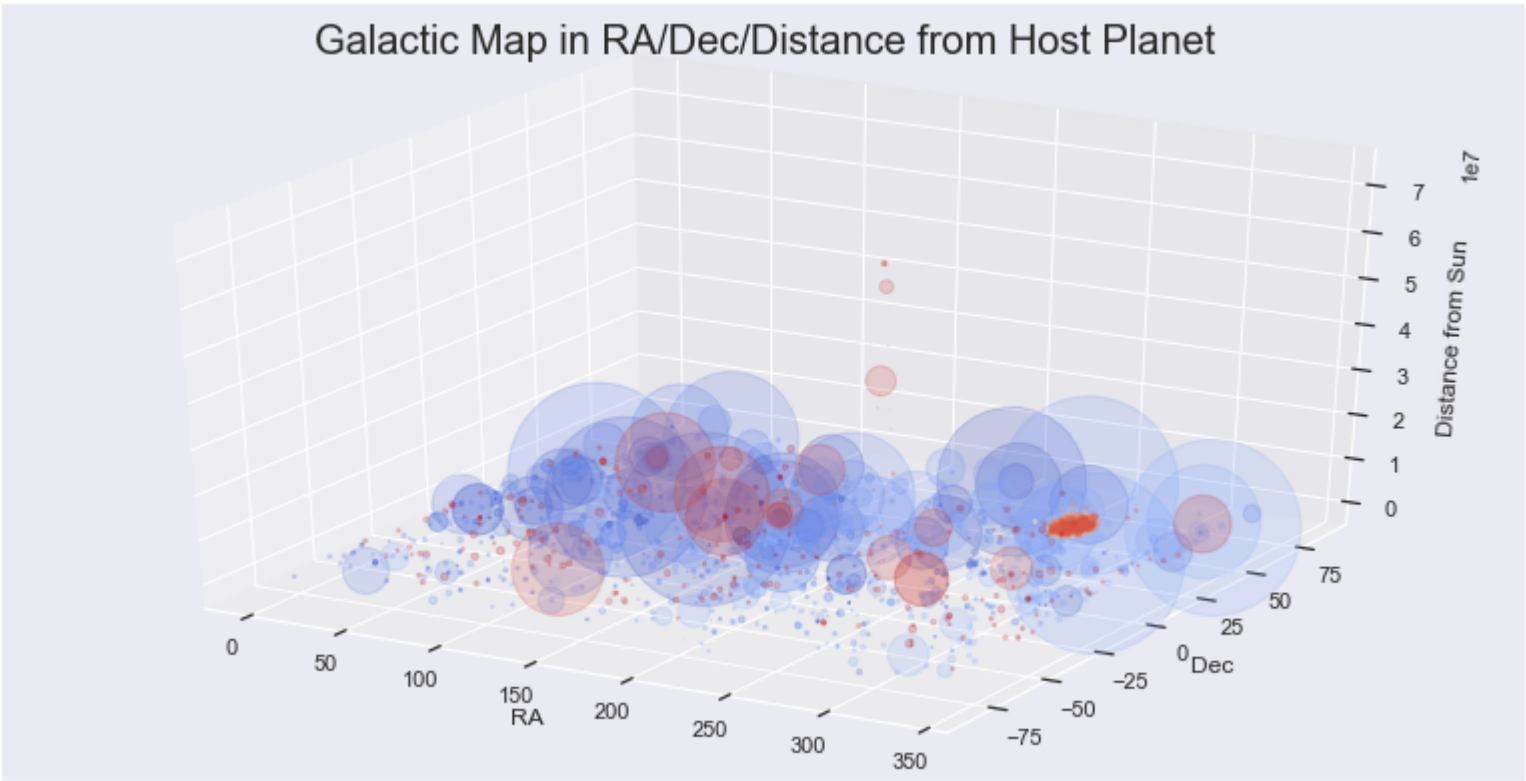
- Filling Up procedure

missing distance values are filled with mean value of distance

```
In [10]: x = allExoplanets['S. RA (hrs)']*14
y = allExoplanets['S. DEC (deg)']
z = allExoplanets['S. Distance (pc)']**2
area = np.pi * allExoplanets['S. Radius (SU)']**2
sl = plt.cm.coolwarm(np.linspace(0,1,len(allExoplanets)))
#csp2=plt.cm.(oec['HostStarTempK'])

fig = plt.figure(figsize=(30,50))
ax = fig.add_subplot(621, projection='3d')

ax.axis('on')
ax.scatter(x,y,z,cmap=cm.coolwarm,alpha=0.2,s=area,c=sl)
ax.set_xlabel('RA ')
ax.set_ylabel('Dec ')
ax.set_zlabel('Distance from Sun')
plt.title('Galactic Map in RA/Dec/Distance from Host Planet',size=20)
plt.show()
```



```
In [11]: x2 = allExoplanets['S. Mass (SU)']
y2 = allExoplanets['S. Age (Gyrs)']
z2 = allExoplanets['S. Teff (K)']

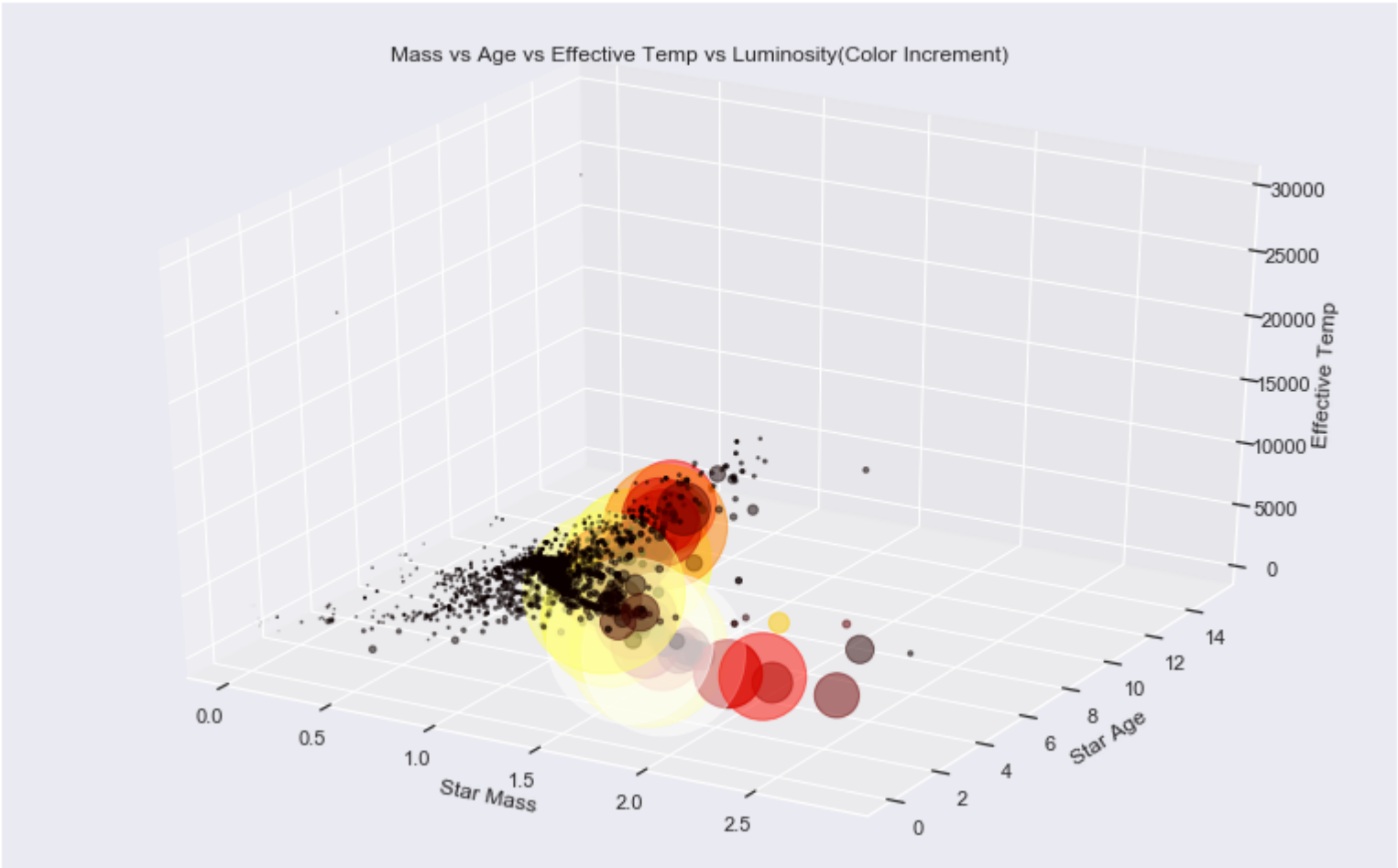
area2 = np.pi * allExoplanets['S. Radius (SU)']**2

#x1 = oec['RA_dd']
#y1 = oec['Dec_dd']
#z1 = dists/1000
#area = np.pi * oec['HostStarRadiusSlrRad']**2

#csp2=plt.cm.(oec['HostStarTempK'])

fig = plt.figure(figsize=(30,40))
ax = fig.add_subplot(421, projection='3d')

ax.axis('on')
ax.scatter(x2,y2,z2,cmap=cm.hot,alpha=0.5,s=area2,c=allExoplanets['S. Luminosity (SU)'])
ax.set_xlabel('Star Mass')
ax.set_ylabel('Star Age')
ax.set_zlabel('Effective Temp')
plt.title('Mass vs Age vs Effective Temp vs Luminosity(Color Increment)')
plt.show()
```



Data Preprocessing

PHL-EC is a very complex and sensitive dataset. Every observation is recorded with high accuracy, That's why we need to perform sensitivity analysis before cleaning, imputing, scaling any part of it.

Detection of Habitable Exoplanets

habstar or habitability, is currently defined as an area, such as a planet or a moon, where liquid water can exist for at least a short duration of time

- A "habitable" planet should:
 - Orbit a star that remains stable in output for billions of years
 - Be at a distance from the star that results in its achieving a suitable temperature so its surface water is liquid, not frozen
 - Have a circular orbit, so constant conditions prevail for its entire "year"
 - Not orbit a star that is too close to a cosmic explosion like a supernova
 - Be far enough from massive planets that they do not continually divert asteroids to hit it or perturb its orbit strongly
 - Probably not be so massive that it retains hydrogen and becomes a "gas giant"
 - Perhaps it is also essential to have a massive planet well outside its orbit, like Jupiter, to divert potential devastating asteroids away, or to make them destroy themselves (as in the asteroid belt).

Habitability

habstar or habitability, is currently defined as an area, such as a planet or a moon, where liquid water can exist for at least a short duration of time

Solar Twin

A true solar twins as noted by the Lowell Observatory should have a temperature within ~10 K of the Sun. Space Telescope Science Institute, Lowell Observatory, noted in 1996 that temperature precision of ~10 K can be measured. A temperature of ~10 K reduces the solar twin list to near zero, so ±50 K is used for the chart

Types of Planets

In PHL-EC, Planets are classified into five categories. This classification has been done on the basis of their thermal properties.

Mesoplanets [Asimov - 1989]:

The planetary bodies whose sizes lie between Mercury and Ceres falls under this category (smaller than Mercury and larger than Ceres). These are also referred to as M-planets [Méndez - 2011]. These planets have mean global surface temperature between 0°C to 50°C, a necessary condition for complex terrestrial life. These are generally referred as Earth-like planets.

Psychroplanets [Méndez - 2011]:

These planets have mean global surface temperature between -50°C to 0°C. Hence, the temperature is colder than optimal for sustenance of terrestrial life

Non-Habitable:

Planets other than mesoplanets and psychroplanets do not have thermal properties required to sustain life.

Description Of Targeted Region

```
In [12]: print(allExoplanets['P. Habitable Class'].describe())

count          3875
unique           5
top      non-habitable
freq           3820
Name: P. Habitable Class, dtype: object

In [13]: print('Our Dataset, or Target region if I be more precise, has been classified into 5 categories, which are:')
for i in allExoplanets['P. Habitable Class'].unique():
    print(i)

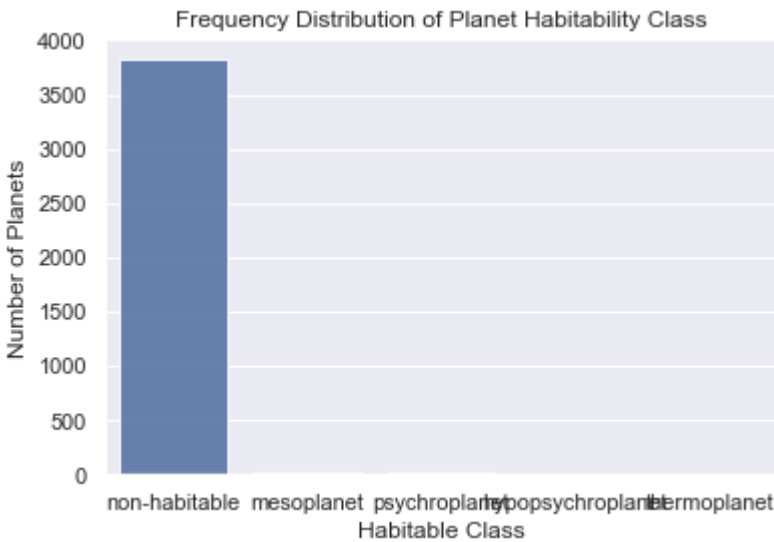
Our Dataset, or Target region if I be more precise, has been classified into 5 categories, which are:
non-habitable
psychroplanet
mesoplanet
thermoplanet
hypopsychroplanet
```

```
In [14]: print(allExoplanets['P. Habitable Class'].value_counts())
```

```
non-habitable      3820
mesoplanet         31
psychroplanet      18
hypopsychroplanet   3
thermoplanet        3
Name: P. Habitable Class, dtype: int64
```

Last three output clearly states that 'non - habitable' class is dominating over other classes. We should visualise the frequency to understand the issue more deeply.

```
In [15]: stat_count = allExoplanets['P. Habitable Class'].value_counts()
sns.set(style="darkgrid")
sns.barplot(stat_count.index, stat_count.values, alpha=0.9)
plt.title('Frequency Distribution of Planet Habitability Class')
plt.ylabel('Number of Planets', fontsize=12)
plt.xlabel('Habitable Class', fontsize=12)
plt.show()
```



Thermoplanet:

A class of planets, which has a temperature in the range of 50°C-100°C. This is warmer than the temperature range suited for most terrestrial life [Méndez2011].

Hypopsychroplanets:

A class of planets whose temperature is below −50°C. Planets belonging to this category are too cold for the survival of most terrestrial life [Méndez2011].

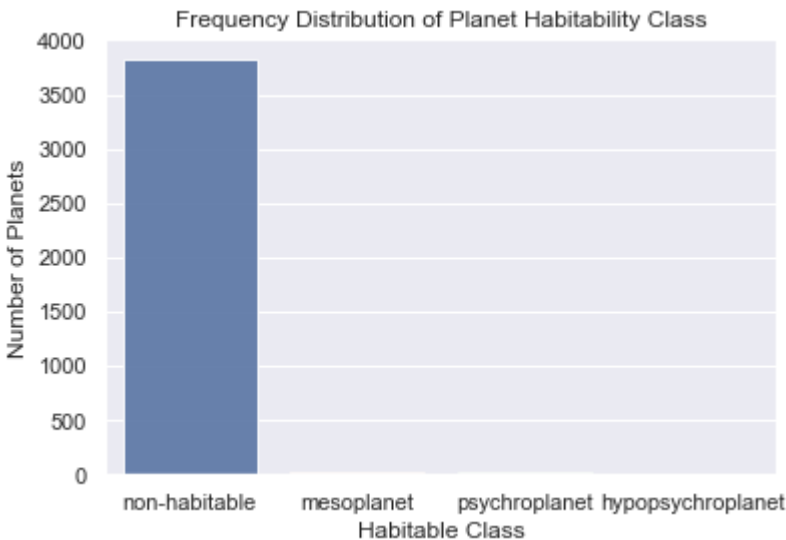
NOTE

The above two classes have three data entities each in the augmented data set used. This number is inadequate for the task of classification, and hence the total of six entities were excluded from the experiment.

```
In [16]: allExoplanets=allExoplanets[allExoplanets.iloc[:,7]!= 'thermoplanet']
allExoplanets['P. Habitable Class'].value_counts()
```

```
Out[16]: non-habitable      3820
mesoplanet         31
psychroplanet      18
hypopsychroplanet   3
Name: P. Habitable Class, dtype: int64
```

```
In [17]: stat_count = allExoplanets['P. Habitable Class'].value_counts()
sns.set(style="darkgrid")
sns.barplot(stat_count.index, stat_count.values, alpha=0.9)
plt.title('Frequency Distribution of Planet Habitability Class')
plt.ylabel('Number of Planets', fontsize=12)
plt.xlabel('Habitable Class', fontsize=12)
plt.show()
```



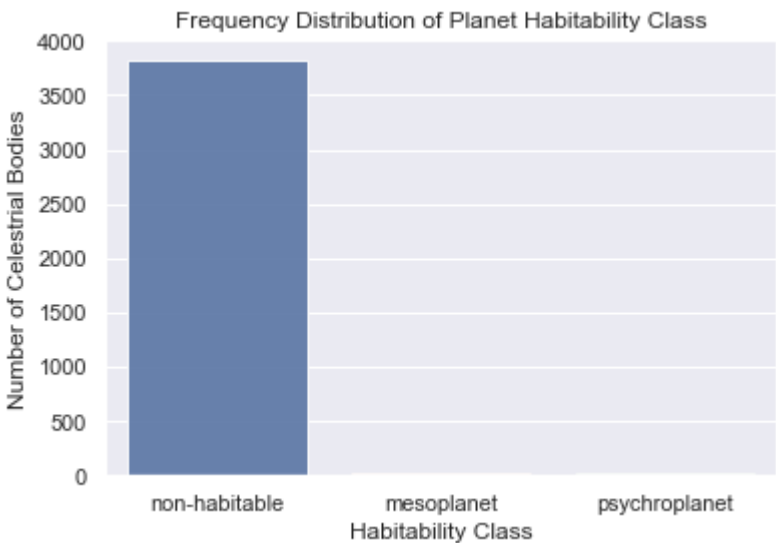
```
In [18]: allExoplanets = allExoplanets[allExoplanets.iloc[:,7]!= 'hypopsychroplanet']
allExoplanets['P. Habitable Class'].value_counts()
```

```
Out[18]: non-habitable    3820
mesoplanet         31
psychroplanet      18
Name: P. Habitable Class, dtype: int64
```

```
In [19]: allExoplanets['P. Habitable Class'].value_counts()
```

```
Out[19]: non-habitable    3820
mesoplanet         31
psychroplanet      18
Name: P. Habitable Class, dtype: int64
```

```
In [20]: stat_count = allExoplanets['P. Habitable Class'].value_counts()
sns.set(style="darkgrid")
sns.barplot(stat_count.index, stat_count.values, alpha=0.9)
plt.title('Frequency Distribution of Planet Habitability Class')
plt.ylabel('Number of Celestial Bodies', fontsize=12)
plt.xlabel('Habitability Class', fontsize=12)
plt.show()
```



thermoplanet and hypopsychroplanet instances are successfully removed

Note

Although thermoplanet and hypopsychroplanet class have been removed from the set for being the most submissive ones. Still, the ratio hasn't improved very much. This may result in High Accuracy and False Positives.

Feature Selection

P.NameKepler (planet’s name), Sname HD and Sname Hid (name of parent star), S.constellation (name of constellation), Stype (type of parent star), P.SPH (planet standard primary habitability), P.interior ESI (interior earth similarity index), P.surface ESI (surface earth similarity index), P.disc method (method of discovery of planet), P.disc year (year of discovery of planet), P. Max Mass, P. Min Mass, P.inclination and P.Hab Moon (flag indicating planet’s potential as a habitable exomoons) were removed as these attributes do not contribute to the nature of classification of habitability of a planet. Interior ESI and surface ESI, however, together contribute to habitability, but since the data set directly provides P.ESI, these two features were neglected. Following this, classification algorithms were applied on the processed data set. In all, 50 features are used.

```
In [21]: pred = ['P. Zone Class','P. Mass Class','P. Composition Class','P. Atmosphere Class',
               'P. Habitable Class','P. Habitable',

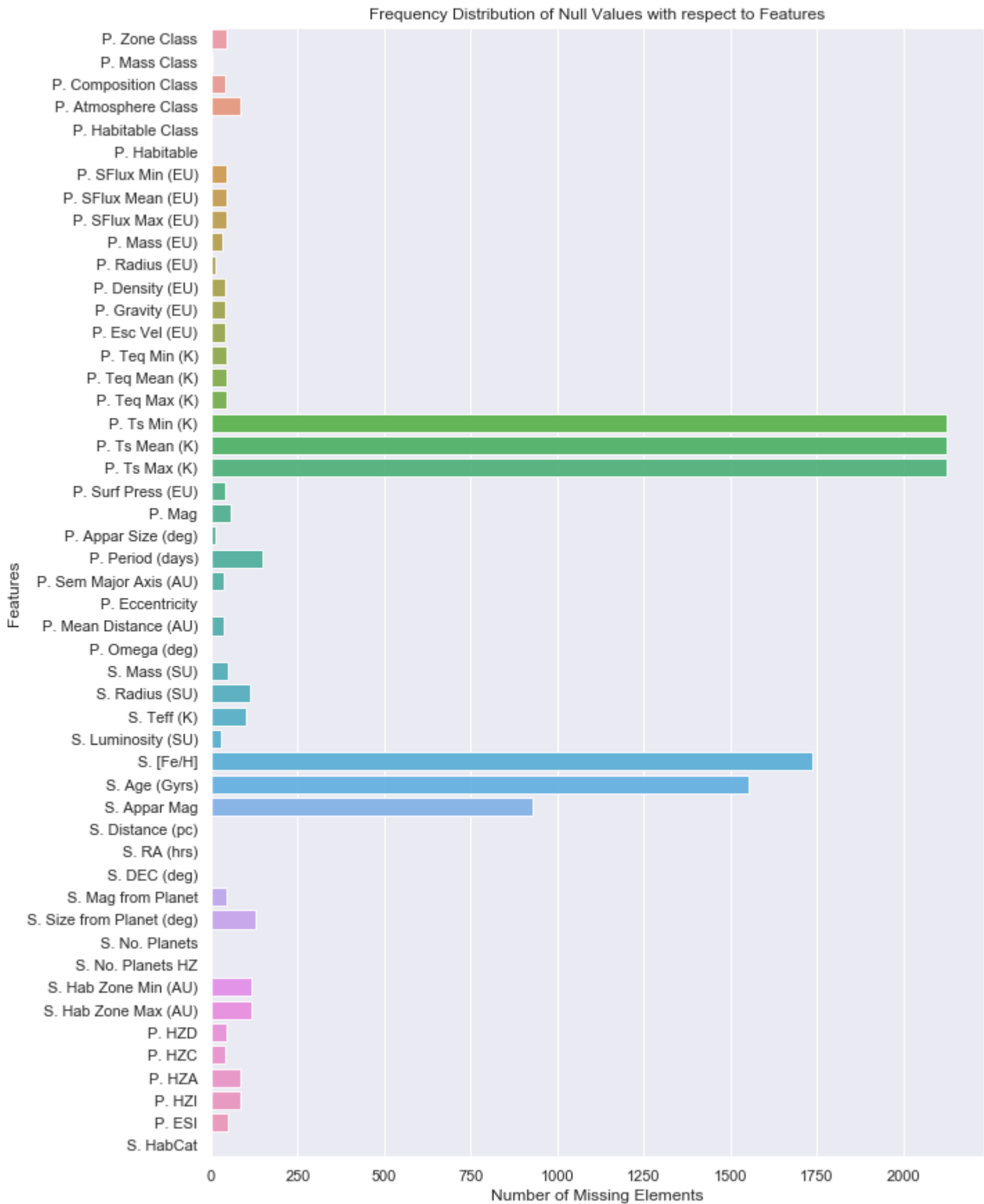
               'P. SFlux Min (EU)','P. SFlux Mean (EU)','P. SFlux Max (EU)',
               'P. Mass (EU)','P. Radius (EU)','P. Density (EU)','P. Gravity (EU)',
               'P. Esc Vel (EU)','P. Teq Min (K)','P. Teq Mean (K)','P. Teq Max (K)',
               'P. Ts Min (K)','P. Ts Mean (K)','P. Ts Max (K)','P. Surf Press (EU)',
               'P. Mag','P. Appar Size (deg)','P. Period (days)','P. Sem Major Axis (AU)',
               'P. Eccentricity','P. Mean Distance (AU)','P. Omega (deg)','S. Mass (SU)',
               'S. Radius (SU)','S. Teff (K)','S. Luminosity (SU)','S. [Fe/H]','S. Age (Gyrs)',
               'S. Appar Mag','S. Distance (pc)','S. RA (hrs)','S. DEC (deg)',
               'S. Mag from Planet','S. Size from Planet (deg)','S. No. Planets',
               'S. No. Planets HZ','S. Hab Zone Min (AU)','S. Hab Zone Max (AU)',
               'P. HZD','P. HZC','P. HZA','P. HZI','P. ESI','S. HabCat']

pred_cat = ['P. Zone Class','P. Mass Class','P. Composition Class','P. Atmosphere Class',
            'P. Habitable Class','P. Habitable']
pred_num = ['P. SFlux Min (EU)','P. SFlux Mean (EU)','P. SFlux Max (EU)','P. Mass (EU)',
            'P. Radius (EU)','P. Density (EU)','P. Gravity (EU)',
            'P. Esc Vel (EU)','P. Teq Min (K)','P. Teq Mean (K)','P. Teq Max (K)',
            'P. Ts Min (K)','P. Ts Mean (K)','P. Ts Max (K)','P. Surf Press (EU)',
            'P. Mag','P. Appar Size (deg)','P. Period (days)','P. Sem Major Axis (AU)',
            'P. Eccentricity','P. Mean Distance (AU)','P. Omega (deg)','S. Mass (SU)',
            'S. Radius (SU)','S. Teff (K)','S. Luminosity (SU)','S. [Fe/H]','S. Age (Gyrs)',
            'S. Appar Mag','S. Distance (pc)','S. RA (hrs)','S. DEC (deg)',
            'S. Mag from Planet','S. Size from Planet (deg)','S. No. Planets',
            'S. No. Planets HZ','S. Hab Zone Min (AU)','S. Hab Zone Max (AU)',
            'P. HZD','P. HZC','P. HZA','P. HZI','P. ESI','S. HabCat']

examine = pd.DataFrame()
examine = allExoplanets[pred]

f, ax = plt.subplots(figsize=(10, 15))

stat_count = examine.isnull().sum()
sns.set(style="darkgrid")
sns.barplot(stat_count.values,stat_count.index, alpha=0.9)
plt.title('Frequency Distribution of Null Values with respect to Features')
plt.xlabel('Number of Missing Elements', fontsize=12)
plt.ylabel('Features', fontsize=12)
plt.show()
```



```
In [22]: cat1 = len(examine.select_dtypes(include=['object']).columns)
num1 = len(examine.select_dtypes(include=['int64','float64']).columns)
print('Features of examine consists of ', cat, 'categorical', ' and ',
      num, 'numerical features')
print('\n\nCategorical Features:\n ')
for i in pred_cat:
    print("{feature}".format(feature=i),end='\t')
print('\n\nNumerical Features:\n ')
for i in pred_num:
    print("{feature}".format(feature=i))
```

Features of examine consists of 14 categorical and 54 numerical features

Categorical Features:

P. Zone Class P. Mass Class P. Composition Class P. Atmosphere Class P. Habitable Class P. Habitable

Numerical Features:

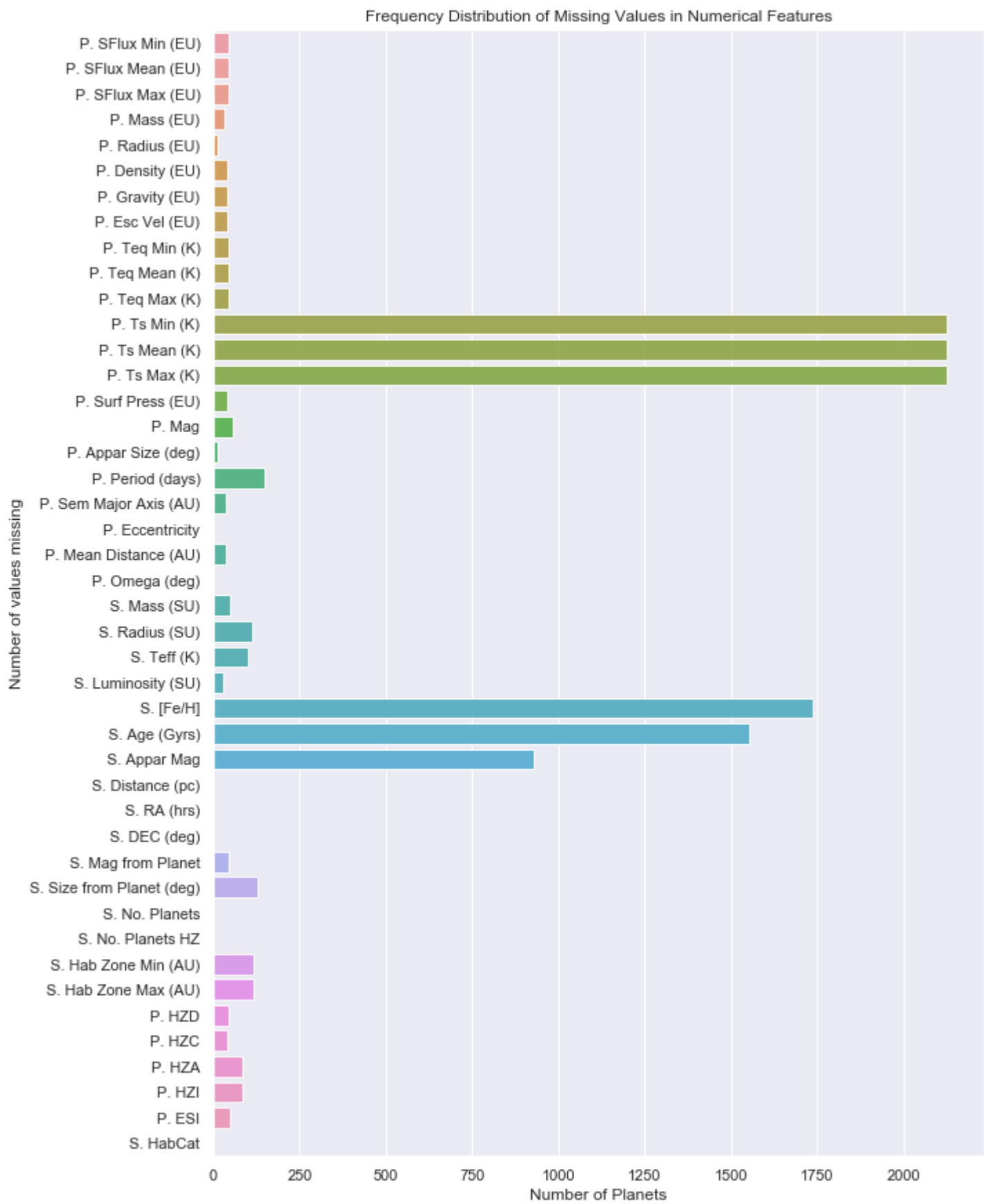
- P. SFlux Min (EU)
- P. SFlux Mean (EU)
- P. SFlux Max (EU)
- P. Mass (EU)
- P. Radius (EU)
- P. Density (EU)
- P. Gravity (EU)
- P. Esc Vel (EU)
- P. Teq Min (K)
- P. Teq Mean (K)
- P. Teq Max (K)
- P. Ts Min (K)
- P. Ts Mean (K)
- P. Ts Max (K)
- P. Surf Press (EU)
- P. Mag
- P. Appar Size (deg)
- P. Period (days)
- P. Sem Major Axis (AU)
- P. Eccentricity
- P. Mean Distance (AU)
- P. Omega (deg)
- S. Mass (SU)
- S. Radius (SU)
- S. Teff (K)
- S. Luminosity (SU)
- S. [Fe/H]
- S. Age (Gyrs)
- S. Appar Mag
- S. Distance (pc)
- S. RA (hrs)
- S. DEC (deg)
- S. Mag from Planet
- S. Size from Planet (deg)
- S. No. Planets
- S. No. Planets HZ
- S. Hab Zone Min (AU)
- S. Hab Zone Max (AU)
- P. HZD
- P. HZC
- P. HZA
- P. HZI
- P. ESI
- S. HabCat

```
cat1 = len(examine.select_dtypes(include=['object']).columns) num1 = len(examine.select_dtypes(include=['int64','float64']).columns) print('Features of
examine consists of ', cat, 'categorical', ' and ', num, 'numerical features') print('\n\nCategorical Features:\n ') for i in pred_cat: print("{
feature}".format(feature=i),end='\t') print('\n\nNumerical Features:\n ') for i in pred_num: print("{feature}".format(feature=i),end='\t')
```



```
In [23]: print("Total of null Values in each Feature of Numerical Features region: \n")
#imputed_exnum=examine.iloc[:,8:]
#print(examine.iloc[:,8:].isnull().sum())
fig , ax = plt.subplots(figsize=(10,15))
stat_count = examine.iloc[:,6:].isnull().sum()
sns.set(style="darkgrid")
sns.barplot(stat_count.values, stat_count.index, alpha=0.9)
plt.title('Frequency Distribution of Missing Values in Numerical Features')
plt.xlabel('Number of Planets', fontsize=12)
plt.ylabel('Number of values missing', fontsize=12)
plt.show()
```

Total of null Values in each Feature of Numerical Features region:



```
In [24]: from sklearn.preprocessing import Imputer
mean_imputer = Imputer(missing_values='NaN', strategy='mean', axis=0)

# Training imputer on the numerical region of 'Examine'
mean_imputer = mean_imputer.fit(examine.iloc[:,6:])
examine.iloc[:,6:] = mean_imputer.transform(examine.iloc[:,6:])

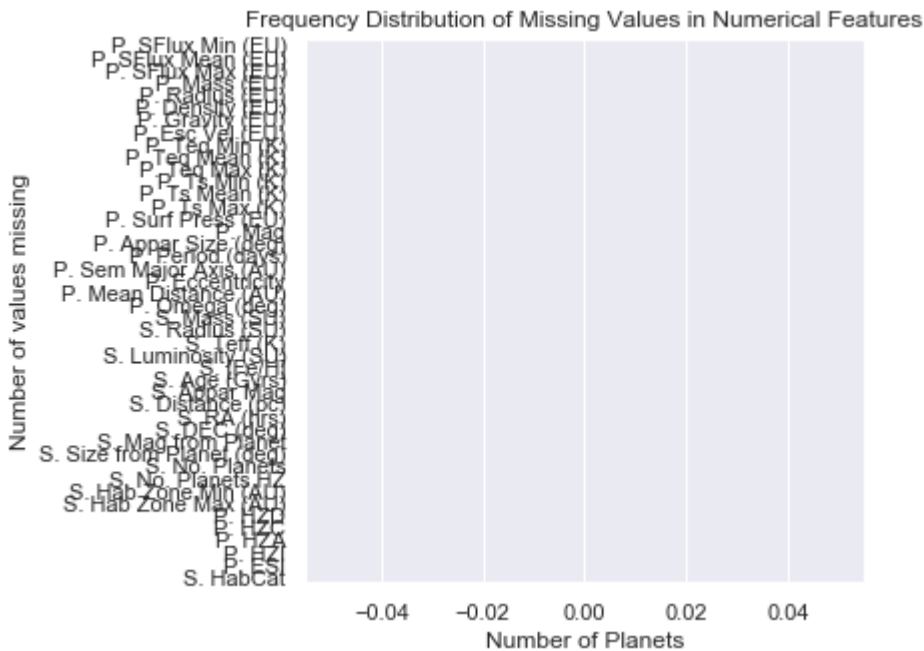
numpart = examine.iloc[:,6:]
```

C:\Users\User\Anaconda3\lib\site-packages\pandas\core\indexing.py:543: SettingWithCopyWarning: A value is trying to be set on a copy of a slice from a DataFrame. Try using .loc[row_indexer,col_indexer] = value instead

See the caveats in the documentation: <http://pandas.pydata.org/pandas-docs/stable/indexing.html#indexing-view-versus-copy>
self.obj[item] = s

```
In [25]: #print(examine.iloc[:,8:].isnull().sum())

fig , ax = plt.subplots(figsize=(5,5))
stat_count = examine.iloc[:,6:].isnull().sum()
sns.set(style="darkgrid")
sns.barplot(stat_count.values, stat_count.index, alpha=0.9)
plt.title('Frequency Distribution of Missing Values in Numerical Features')
plt.xlabel('Number of Planets', fontsize=12)
plt.ylabel('Number of values missing', fontsize=12)
plt.show()
```



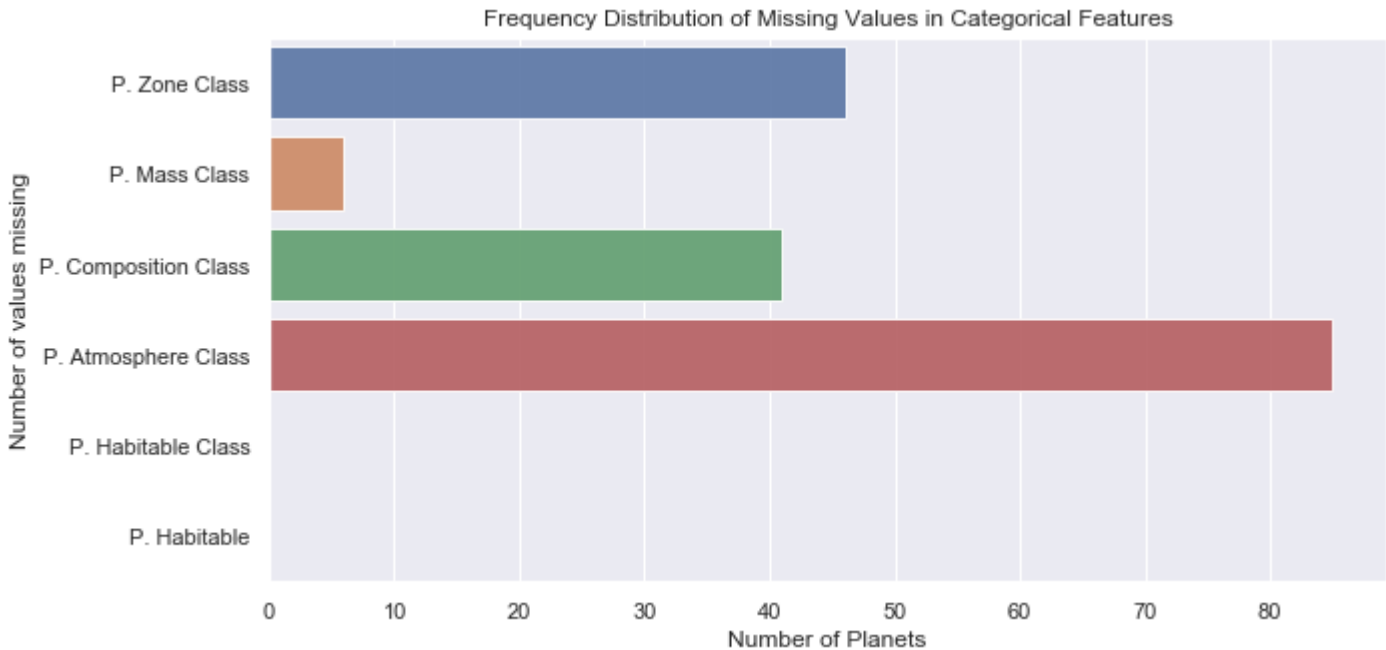
Imputing of Categorical Features

Categorical Features are imputed by most frequent strategy

```
In [26]: #imputed_exnum=examine.iloc[:,8:]
#print(examine.iloc[:,8:].isnull().sum())
#f, ax = plt.subplots(figsize=(5, 5))

###for i in examine[pred_cat]:
###    stat_count = examine[i].value_counts()
###    sns.set(style="darkgrid")
###    sns.barplot(stat_count.index, stat_count.values, alpha=0.9)
###    plt.title('Frequency Distribution of {col}'.format(col=i))
###    plt.ylabel('Number of Planets', fontsize=12)
###    plt.xlabel('{s}'.format(s=i), fontsize=12)
###plt.show()
print("Total of null Values in each Feature of Categorical Features region: \n")
#imputed_exnum=examine.iloc[:,8:]
#print(examine.iloc[:,8:].isnull().sum())
fig , ax = plt.subplots(figsize=(10,5))
stat_count = examine.iloc[:,6:].isnull().sum()
sns.set(style="darkgrid")
sns.barplot(stat_count.values, stat_count.index, alpha=0.9)
plt.title('Frequency Distribution of Missing Values in Categorical Features')
plt.xlabel('Number of Planets', fontsize=12)
plt.ylabel('Number of values missing', fontsize=12)
plt.show()
```

Total of null Values in each Feature of Categorical Features region:



```
In [27]: #dataframe imputer
from sklearn.base import TransformerMixin

class DataFrameImputer(TransformerMixin):
    def __init__(self):
        #cat -> most frq , num -> mean
        return None
    def fit(self, X, y=None):

        self.fill = pd.Series([X[c].value_counts().index[0]
                               if X[c].dtype == np.dtype('O')
                               else X[c].mean() for c in X],
                               index = X.columns)

        return self

    def transform(self,X,y=None):
        return X.fillna(self.fill)

examine.iloc[:,4] = DataFrameImputer().fit_transform(examine.iloc[:,4])
```

C:\Users\User\Anaconda3\lib\site-packages\pandas\core\indexing.py:543: SettingWithCopyWarning: A value is trying to be set on a copy of a slice from a DataFrame. Try using .loc[row_indexer,col_indexer] = value instead

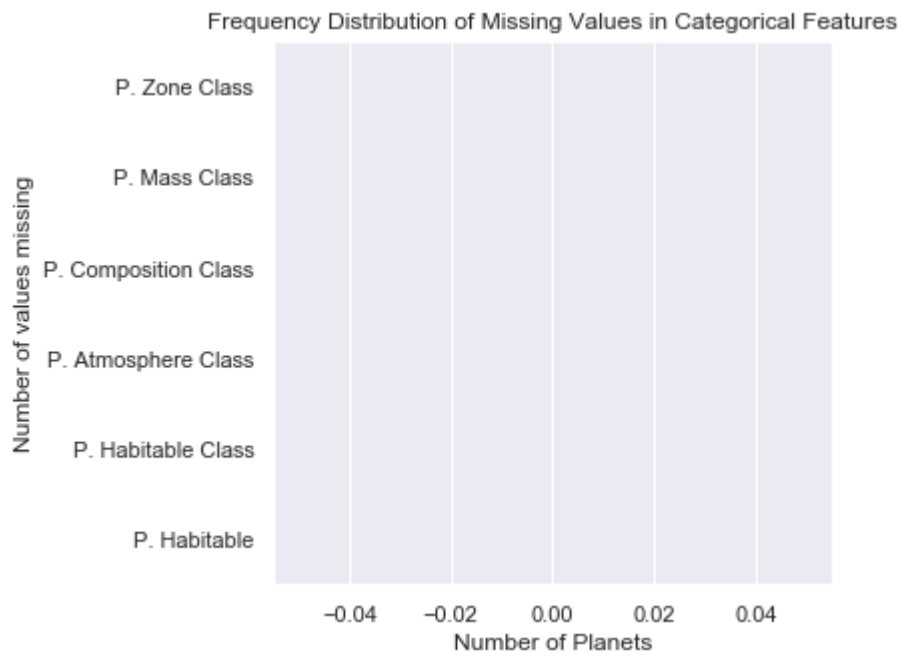
See the caveats in the documentation: <http://pandas.pydata.org/pandas-docs/stable/indexing.html#indexing-view-versus-copy>

```
self.obj[item] = s
```

```
In [28]: catpart = examine.iloc[:,6]
```

```
In [29]: #imputed_exnum=examine.iloc[:,8:]
#print(examine.iloc[:,8:].isnull().sum())

fig , ax = plt.subplots(figsize=(5,5))
stat_count = examine.iloc[:,6].isnull().sum()
sns.set(style="darkgrid")
sns.barplot(stat_count.values, stat_count.index, alpha=0.9)
plt.title('Frequency Distribution of Missing Values in Categorical Features')
plt.xlabel('Number of Planets', fontsize=12)
plt.ylabel('Number of values missing', fontsize=12)
plt.show()
```



Label Frequency in Categorical Features

```
In [30]: stat_count = allExoplanets['P. Zone Class'].value_counts()
sns.set(style="darkgrid")
sns.barplot(stat_count.index, stat_count.values, alpha=0.9)
plt.title('Frequency Distribution of P. Zone Class')
plt.ylabel('Number of Planets', fontsize=12)
plt.xlabel('P. Zone Class', fontsize=12)
plt.show()

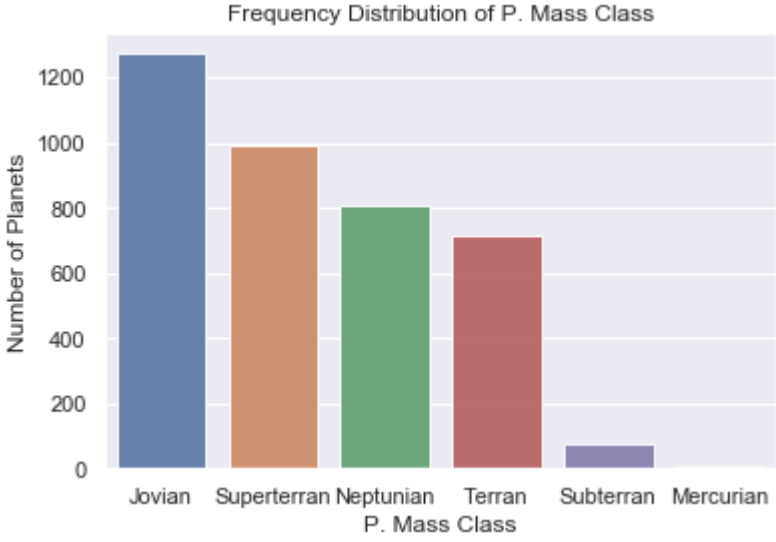
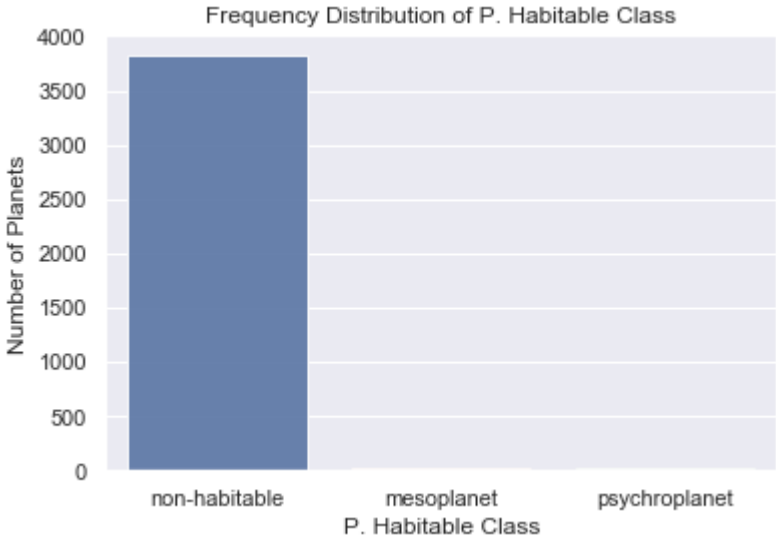
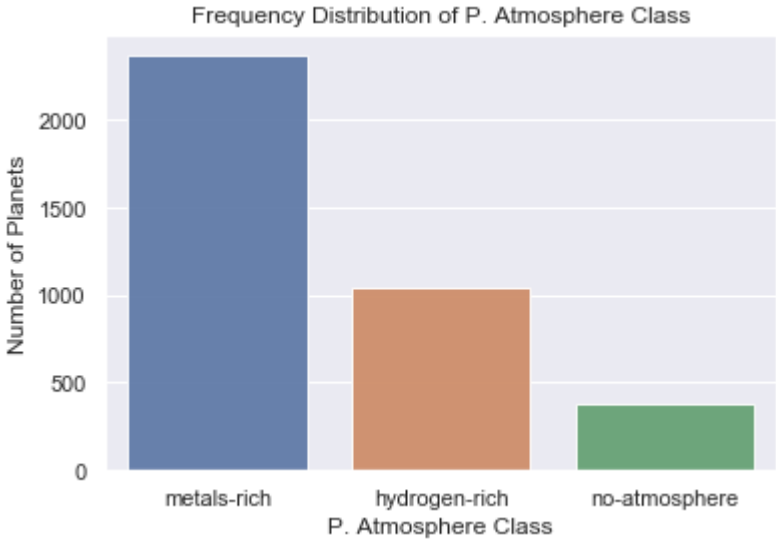
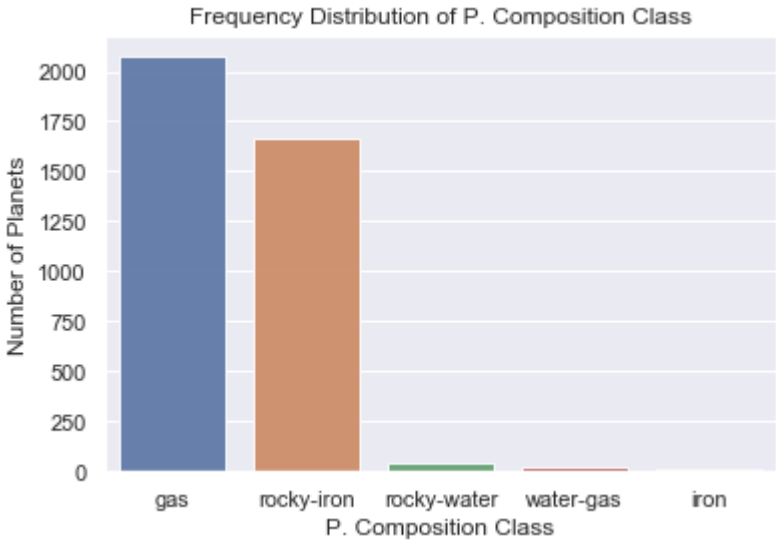
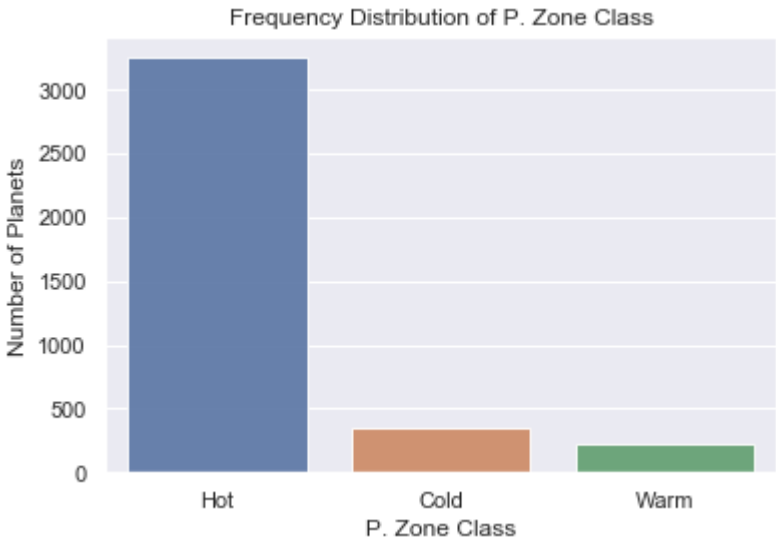
stat_count = allExoplanets['P. Composition Class'].value_counts()
sns.set(style="darkgrid")
sns.barplot(stat_count.index, stat_count.values, alpha=0.9)
plt.title('Frequency Distribution of P. Composition Class')
plt.ylabel('Number of Planets', fontsize=12)
plt.xlabel('P. Composition Class', fontsize=12)
plt.show()

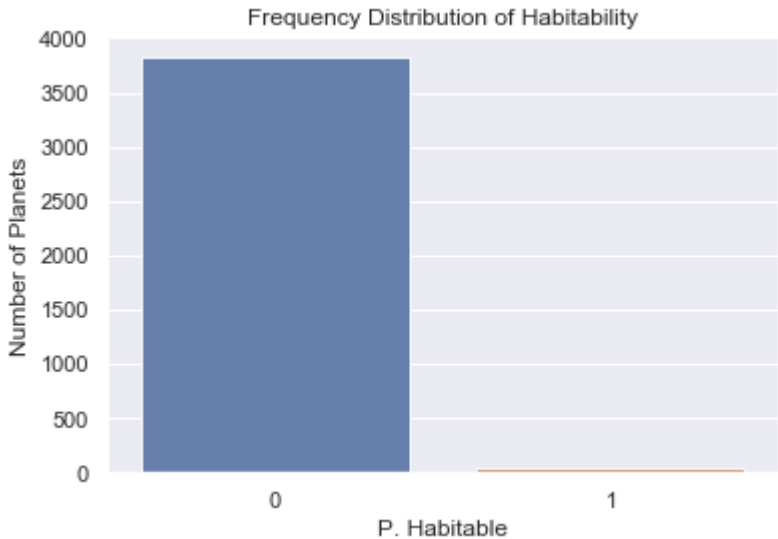
stat_count = allExoplanets['P. Atmosphere Class'].value_counts()
sns.set(style="darkgrid")
sns.barplot(stat_count.index, stat_count.values, alpha=0.9)
plt.title('Frequency Distribution of P. Atmosphere Class')
plt.ylabel('Number of Planets', fontsize=12)
plt.xlabel('P. Atmosphere Class', fontsize=12)
plt.show()

stat_count = allExoplanets['P. Habitable Class'].value_counts()
sns.set(style="darkgrid")
sns.barplot(stat_count.index, stat_count.values, alpha=0.9)
plt.title('Frequency Distribution of P. Habitable Class')
plt.ylabel('Number of Planets', fontsize=12)
plt.xlabel('P. Habitable Class', fontsize=12)
plt.show()

stat_count = allExoplanets['P. Mass Class'].value_counts()
sns.set(style="darkgrid")
sns.barplot(stat_count.index, stat_count.values, alpha=0.9)
plt.title('Frequency Distribution of P. Mass Class')
plt.ylabel('Number of Planets', fontsize=12)
plt.xlabel('P. Mass Class', fontsize=12)
plt.show()

stat_count = allExoplanets['P. Habitable'].value_counts()
sns.set(style="darkgrid")
sns.barplot(stat_count.index, stat_count.values, alpha=0.9)
plt.title('Frequency Distribution of Habitability')
plt.ylabel('Number of Planets', fontsize=12)
plt.xlabel('P. Habitable', fontsize=12)
plt.show()
```





Present Status of my database is: `#examine2 = DataFrameImputer().fit_transform(examine)` ### taking a backup
`#examine2.to_csv('exoplanets_imputed_notencoded.csv')` `#examine2` `#examine.isna().sum()`

All the Features have been imputed. Now, What I'm gonna do is to encode all the variables. Before that, separating the Independent variable, i.e, separating X and y.

Selection of X and y

```
In [31]: y=examine['P. Habitable Class']
dependent = ['P. Zone Class','P. Mass Class','P. Composition Class','P. Atmosphere Class',
            'P. Habitable',

            'P. SFlux Min (EU)','P. SFlux Mean (EU)','P. SFlux Max (EU)',
            'P. Mass (EU)','P. Radius (EU)','P. Density (EU)','P. Gravity (EU)',
            'P. Esc Vel (EU)','P. Teq Min (K)','P. Teq Mean (K)','P. Teq Max (K)',
            'P. Ts Min (K)','P. Ts Mean (K)','P. Ts Max (K)','P. Surf Press (EU)',
            'P. Mag','P. Appar Size (deg)','P. Period (days)','P. Sem Major Axis (AU)',
            'P. Eccentricity','P. Mean Distance (AU)','P. Omega (deg)','S. Mass (SU)',
            'S. Radius (SU)','S. Teff (K)','S. Luminosity (SU)','S. [Fe/H]','S. Age (Gyrs)',
            'S. Appar Mag','S. Distance (pc)','S. RA (hrs)','S. DEC (deg)',
            'S. Mag from Planet','S. Size from Planet (deg)','S. No. Planets',
            'S. No. Planets HZ','S. Hab Zone Min (AU)','S. Hab Zone Max (AU)',
            'P. HZD','P. HZC','P. HZA','P. HZI','P. ESI','S. HabCat']

X = examine[dependent]
```

```
In [32]: print('X and y are of shape '+str(X.shape)+' and '+str(y.shape)+' respectively')

X and y are of shape (3869, 49) and (3869,) respectively
```

Current Status of our dataset

```
In [33]: examine.head()
```

Out[33]:

	P. Zone Class	P. Mass Class	P. Composition Class	P. Atmosphere Class	P. Habitable Class	P. Habitable	P. SFlux Min (EU)	P. SFlux Mean (EU)	P. SFlux Max (EU)	P. Mass (EU)	...	S. No. Planets	S. HabCat
0	Cold	Jovian	gas	hydrogen-rich	non-habitable	0	4.080000e-06	4.080000e-06	4.080000e-06	4451.16	...	1.0	0.0
1	Cold	Jovian	gas	hydrogen-rich	non-habitable	0	2.166914e-02	2.166914e-02	2.166914e-02	6358.80	...	1.0	0.0
2	Cold	Jovian	gas	hydrogen-rich	non-habitable	0	3.960000e-06	3.960000e-06	3.960000e-06	4133.22	...	1.0	0.0
3	Cold	Jovian	gas	hydrogen-rich	non-habitable	0	1.030000e-05	1.030000e-05	1.030000e-05	6358.80	...	1.0	0.0
4	Cold	Jovian	gas	hydrogen-rich	non-habitable	0	2.370000e-07	2.370000e-07	2.370000e-07	4419.37	...	1.0	0.0

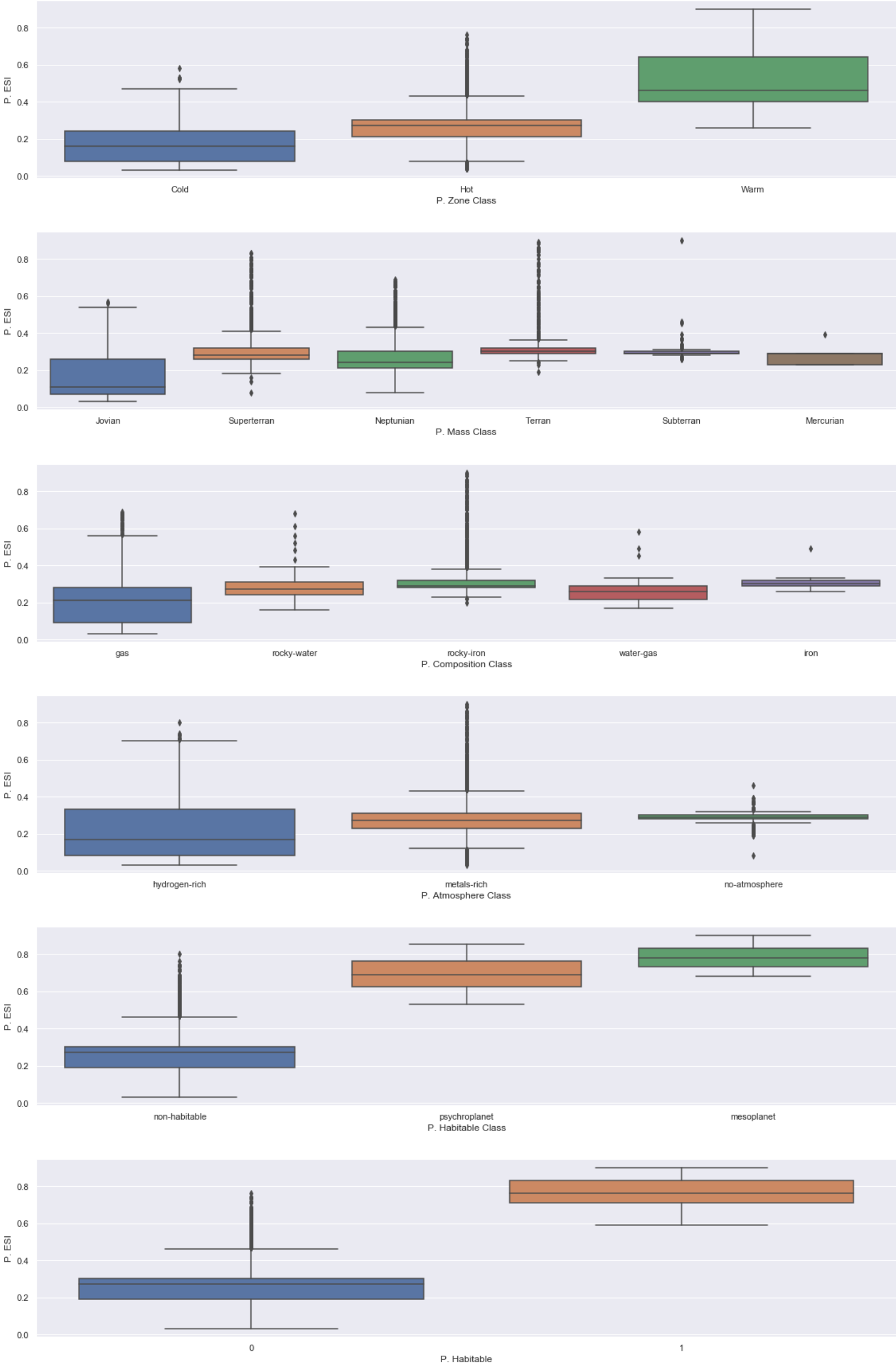
5 rows × 50 columns

How ESI is correlated with different features


```
In [34]: plt.figure(figsize=(20,5))

for i in pred_cat:
    sns.catplot(x=i, y="P. ESI", data=examine,height=4, aspect=4,kind="box")
```

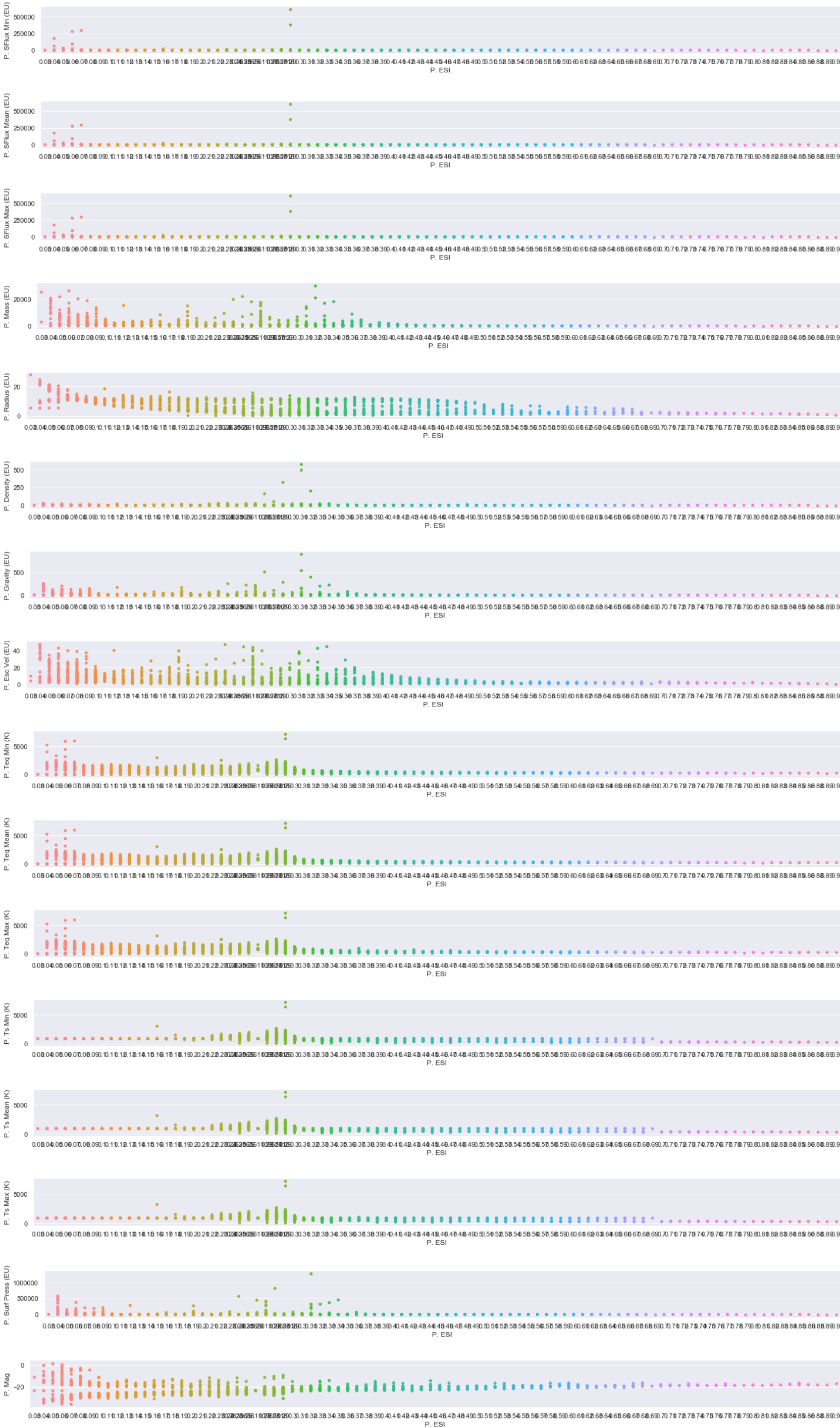
<Figure size 1440x360 with 0 Axes>

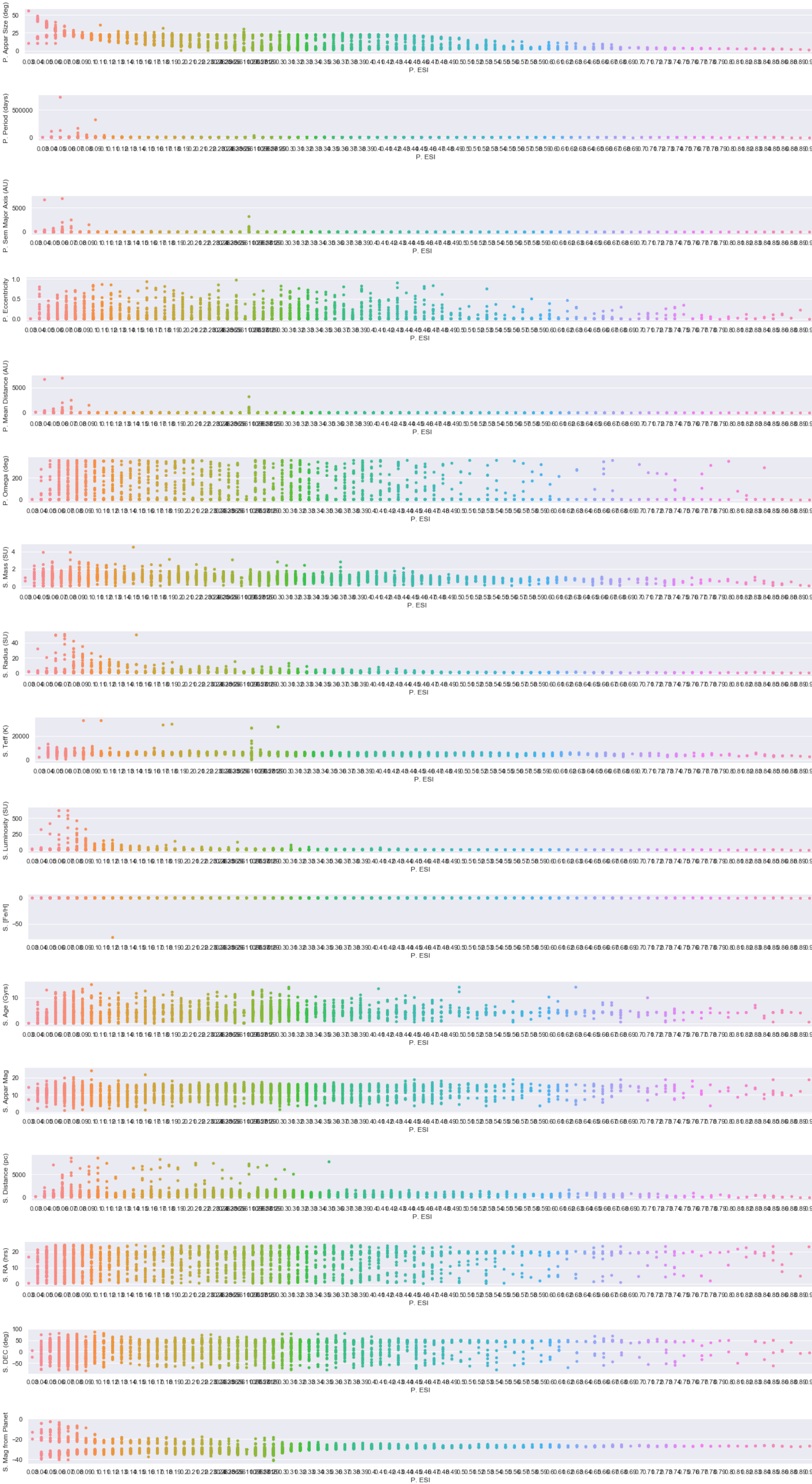


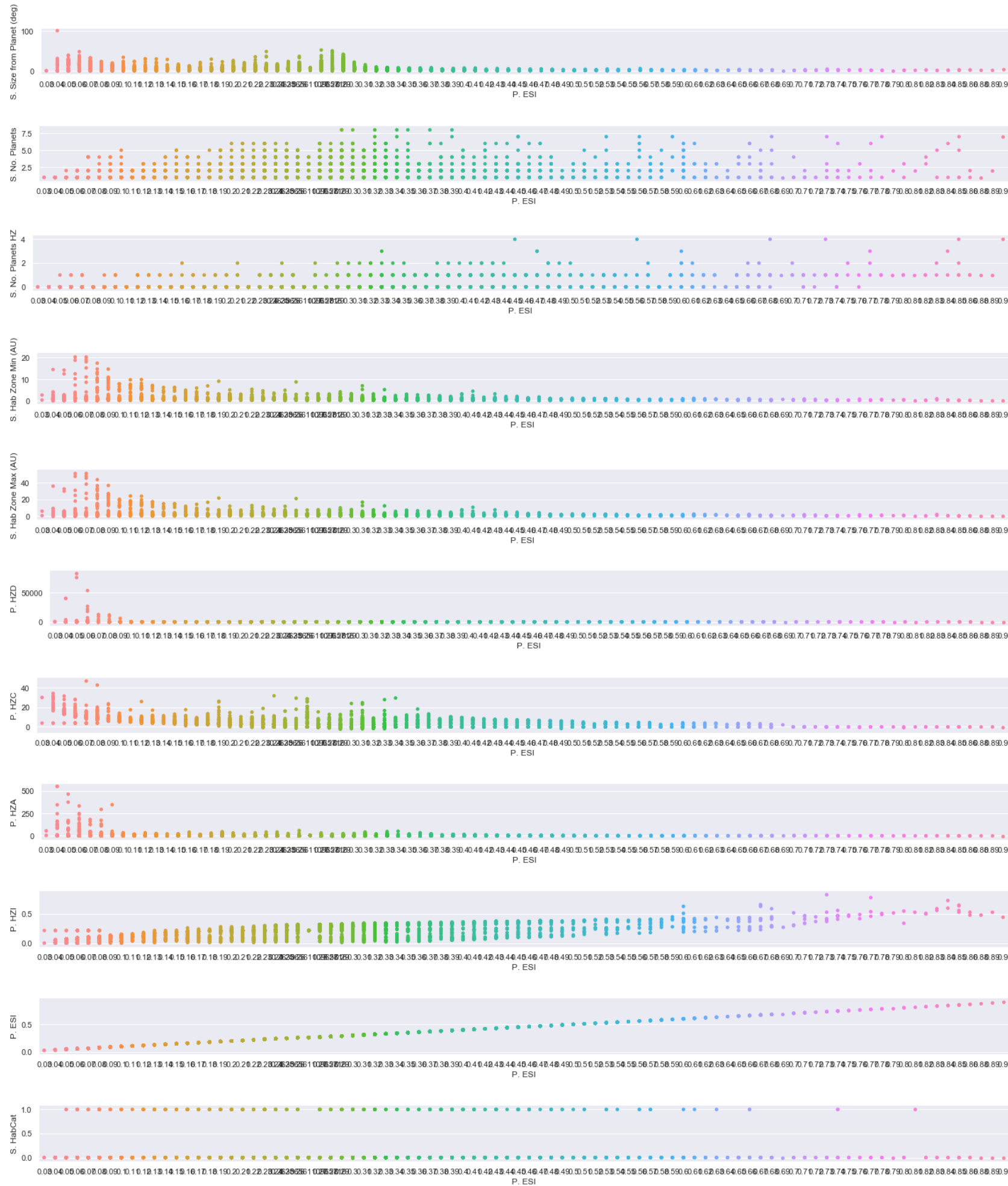
```
In [35]: for i in pred_num:
         sns.catplot(x='P. ESI', y=i, jitter=False, data=examine,height=2,aspect=10,sharex=True);
```

C:\Users\User\Anaconda3\lib\site-packages\matplotlib\pyplot.py:537: RuntimeWarning: More than 20 figures have been opened. Figures created through the pyplot interface (`matplotlib.pyplot.figure`) are retained until explicitly closed and may consume too much memory. (To control this warning, see the rcParam `figure.max_open_warning`).

max_open_warning, RuntimeWarning)







Encoding the data

```
In [36]: X.shape
Out[36]: (3869, 49)

In [37]: # Encoding the Independent Variable
from sklearn.preprocessing import LabelEncoder, OneHotEncoder

# Encoding the Dependent Variable
labelencoder_y = LabelEncoder()
y = labelencoder_y.fit_transform(y)

# Encoding the Independent Variable
X = pd.concat([X,pd.get_dummies(X['P. Zone Class'], prefix='P. Zone Class',drop_first=True)],axis=1)
X.drop(['P. Zone Class'],axis=1, inplace=True)

X = pd.concat([X,pd.get_dummies(X['P. Mass Class'], prefix='P. Mass Class',drop_first=True)],axis=1)
X.drop(['P. Mass Class'],axis=1, inplace=True)

X = pd.concat([X,pd.get_dummies(X['P. Composition Class'], prefix='P. Composition Class',drop_first=True)],axis=1)
X.drop(['P. Composition Class'],axis=1, inplace=True)

X = pd.concat([X,pd.get_dummies(X['P. Atmosphere Class'], prefix='P. Atmosphere Class',drop_first=True)],axis=1)
X.drop(['P. Atmosphere Class'],axis=1, inplace=True)
```

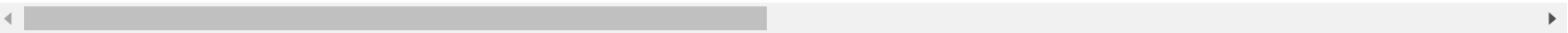
```
In [38]: print(X.shape)
print(X.columns)
X.head()
```

```
(3869, 58)
Index(['P. Habitable', 'P. SFlux Min (EU)', 'P. SFlux Mean (EU)',
      'P. SFlux Max (EU)', 'P. Mass (EU)', 'P. Radius (EU)',
      'P. Density (EU)', 'P. Gravity (EU)', 'P. Esc Vel (EU)',
      'P. Teq Min (K)', 'P. Teq Mean (K)', 'P. Teq Max (K)', 'P. Ts Min (K)',
      'P. Ts Mean (K)', 'P. Ts Max (K)', 'P. Surf Press (EU)', 'P. Mag',
      'P. Appar Size (deg)', 'P. Period (days)', 'P. Sem Major Axis (AU)',
      'P. Eccentricity', 'P. Mean Distance (AU)', 'P. Omega (deg)',
      'S. Mass (SU)', 'S. Radius (SU)', 'S. Teff (K)', 'S. Luminosity (SU)',
      'S. [Fe/H]', 'S. Age (Gyrs)', 'S. Appar Mag', 'S. Distance (pc)',
      'S. RA (hrs)', 'S. DEC (deg)', 'S. Mag from Planet',
      'S. Size from Planet (deg)', 'S. No. Planets', 'S. No. Planets HZ',
      'S. Hab Zone Min (AU)', 'S. Hab Zone Max (AU)', 'P. HZD', 'P. HZC',
      'P. HZA', 'P. HZI', 'P. ESI', 'S. HabCat', 'P. Zone Class_Hot',
      'P. Zone Class_Warm', 'P. Mass Class_Mercurian',
      'P. Mass Class_Neptunian', 'P. Mass Class_Subterran',
      'P. Mass Class_Superterran', 'P. Mass Class_Terran',
      'P. Composition Class_iron', 'P. Composition Class_rocky-iron',
      'P. Composition Class_rocky-water', 'P. Composition Class_water-gas',
      'P. Atmosphere Class_metals-rich', 'P. Atmosphere Class_no-atmosphere'],
      dtype='object')
```

Out[38]:

	P. Habitable	P. SFlux Min (EU)	P. SFlux Mean (EU)	P. SFlux Max (EU)	P. Mass (EU)	P. Radius (EU)	P. Density (EU)	P. Gravity (EU)	P. Esc Vel (EU)	P. Teq Min (K)	...	P. Mass Class_Neptunian	Class_...
0	0	4.080000e-06	4.080000e-06	4.080000e-06	4451.16	19.04	0.64	12.28	15.29	11.4	...	0	0
1	0	2.166914e-02	2.166914e-02	2.166914e-02	6358.80	10.94	4.86	53.12	24.11	97.7	...	0	0
2	0	3.960000e-06	3.960000e-06	3.960000e-06	4133.22	11.40	2.79	31.79	19.04	11.4	...	0	0
3	0	1.030000e-05	1.030000e-05	1.030000e-05	6358.80	11.20	4.53	50.69	23.83	14.4	...	0	0
4	0	2.370000e-07	2.370000e-07	2.370000e-07	4419.37	16.13	1.05	16.99	16.55	5.6	...	0	0

5 rows × 58 columns



DF to ndArr conversion of X

(I find it easier to perform the conversion after encoding the categorical variables)

```
In [39]: X = X.values
```

```
In [40]: X.shape
```

Out[40]: (3869, 58)

```
In [41]: print("X is: \n\n" + repr(X))
print("y is: \n\n" + repr(y))
```

```
X is:

array([[0.000000e+00, 4.080000e-06, 4.080000e-06, ..., 0.000000e+00,
        0.000000e+00, 0.000000e+00],
       [0.000000e+00, 2.166914e-02, 2.166914e-02, ..., 0.000000e+00,
        0.000000e+00, 0.000000e+00],
       [0.000000e+00, 3.960000e-06, 3.960000e-06, ..., 0.000000e+00,
        0.000000e+00, 0.000000e+00],
       ...,
       [0.000000e+00, 9.110547e+00, 9.110548e+00, ..., 0.000000e+00,
        1.000000e+00, 0.000000e+00],
       [0.000000e+00, 4.674975e+00, 5.060307e+00, ..., 0.000000e+00,
        1.000000e+00, 0.000000e+00],
       [0.000000e+00, 2.268338e+00, 2.915344e+00, ..., 0.000000e+00,
        1.000000e+00, 0.000000e+00]])

y is:

array([1, 1, 1, ..., 1, 1, 1], dtype=int64)
```

Feature Scaling


```
In [42]: from sklearn.preprocessing import StandardScaler
sc = StandardScaler()
X = sc.fit_transform(X)
```

```
In [43]: print("X is: \n\n" + repr(X))
print("y is: \n\n" + repr(y))
```

X is:

```
array([[ -0.11325734, -0.05969398, -0.0606261 , ..., -0.07900552,
        -1.31691899, -0.32857427],
       [ -0.11325734, -0.05969239, -0.06062451, ..., -0.07900552,
        -1.31691899, -0.32857427],
       [ -0.11325734, -0.05969398, -0.0606261 , ..., -0.07900552,
        -1.31691899, -0.32857427],
       ...,
       [ -0.11325734, -0.05902641, -0.05995858, ..., -0.07900552,
         0.75934815, -0.32857427],
       [ -0.11325734, -0.05935142, -0.06025533, ..., -0.07900552,
         0.75934815, -0.32857427],
       [ -0.11325734, -0.05952777, -0.06041249, ..., -0.07900552,
         0.75934815, -0.32857427]])
```

y is:

```
array([1, 1, 1, ..., 1, 1, 1], dtype=int64)
```

Splitting the dataset into the Training set and Test set

```
In [44]: from sklearn.cross_validation import train_test_split
#X = X[:,[10,11,24]]
X_train, X_test, y_train, y_test = train_test_split(X, y, test_size = 0.30, random_state = 120)
```

C:\Users\User\Anaconda3\lib\site-packages\sklearn\cross_validation.py:41: DeprecationWarning: This module was deprecated in version 0.18 in favor of the model_selection module into which all the refactored classes and functions are moved. Also note that the interface of the new CV iterators are different from that of this module. This module will be removed in 0.20.
"This module will be removed in 0.20.", DeprecationWarning)

Feature Scaling

```
from sklearn.preprocessing import StandardScaler
sc = StandardScaler()
X_train = sc.fit_transform(X_train)
X_test = sc.transform(X_test)
```

Fitting Naive Bayes to the Training set

```
In [45]: from sklearn.naive_bayes import GaussianNB
classifier = GaussianNB()
classifier.fit(X_train, y_train)
```

Out[45]: GaussianNB(priors=None)

Predicting the Test set results

```
In [46]: y_pred = classifier.predict(X_test)
```

Making the Confusion Matrix

```
In [47]: import itertools

from sklearn import svm
from sklearn.metrics import confusion_matrix

class_names = ['meso', 'non-hab', 'psychro']

def plot_confusion_matrix(cm, classes,
                          normalize=False,
                          title='Confusion matrix',
                          cmap=plt.cm.Blues):
    """
    This function prints and plots the confusion matrix.
    Normalization can be applied by setting `normalize=True`.
    """
    if normalize:
        cm = cm.astype('float') / cm.sum(axis=1)[:, np.newaxis]
        print("Normalized confusion matrix")
    else:
        print('Confusion matrix, without normalization')

    print(cm)

    plt.imshow(cm, interpolation='nearest', cmap=cmap)
    plt.title(title)
    plt.colorbar()
    tick_marks = np.arange(len(classes))
    plt.xticks(tick_marks, classes, rotation=45)
    plt.yticks(tick_marks, classes)

    fmt = '.2f' if normalize else 'd'
    thresh = cm.max() / 2.
    for i, j in itertools.product(range(cm.shape[0]), range(cm.shape[1])):
        plt.text(j, i, format(cm[i, j], fmt),
                 horizontalalignment="center",
                 color="white" if cm[i, j] > thresh else "black")

    plt.ylabel('True label')
    plt.xlabel('Predicted label')
    plt.tight_layout()

# Compute confusion matrix
cnf_matrix_SVM = confusion_matrix(y_test, y_pred)
cnf_matrix = cnf_matrix_SVM

np.set_printoptions(precision=2)

# Plot non-normalized confusion matrix
plt.figure(figsize=(8,5))
plot_confusion_matrix(cnf_matrix, classes=class_names,
                      title='Confusion matrix, without normalization')

# Plot normalized confusion matrix
plt.figure(figsize=(8,5))
plot_confusion_matrix(cnf_matrix, classes=class_names, normalize=True,
                      title='Normalized confusion matrix')

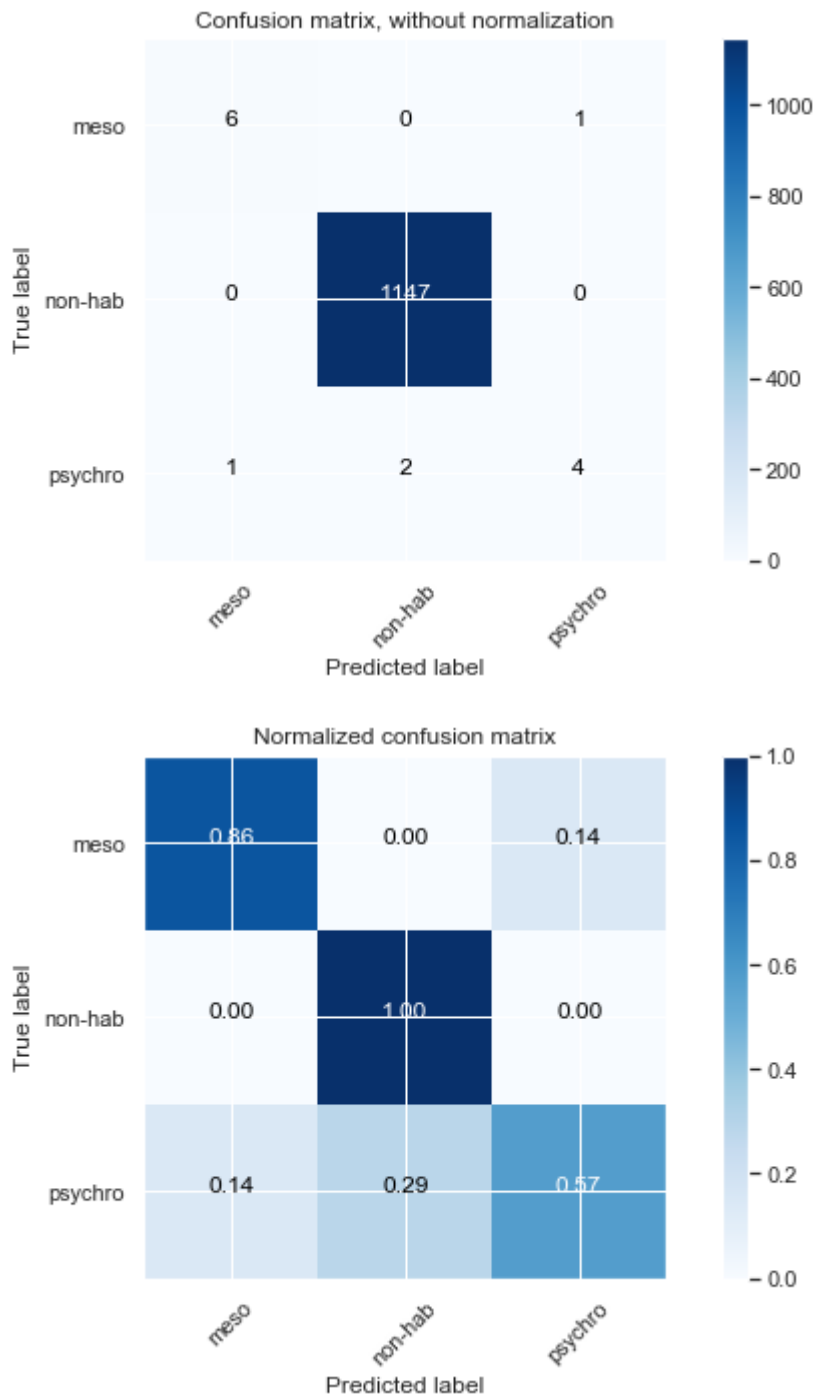
plt.show()
```

Confusion matrix, without normalization

```
[[ 6   0   1]
 [ 0 1147   0]
 [ 1   2   4]]
```

Normalized confusion matrix

```
[[0.86 0.   0.14]
 [0.   1.   0.   ]
 [0.14 0.29 0.57]]
```



Fitting Decision Tree to the training set

```
In [48]: # Fitting Decision Tree Classification to the Training set
from sklearn.tree import DecisionTreeClassifier
classifier = DecisionTreeClassifier(criterion = 'entropy', random_state = 0)
classifier.fit(X_train, y_train)

# Predicting the Test set results
y_pred = classifier.predict(X_test)
```

Confusion Matrix

```
In [49]: # Compute confusion matrix
cnf_matrix_DT = confusion_matrix(y_test, y_pred)
cnf_matrix = cnf_matrix_DT
np.set_printoptions(precision=2)

# Plot non-normalized confusion matrix
plt.figure(figsize=(8,5))
plot_confusion_matrix(cnf_matrix, classes=class_names,
                      title='Confusion matrix, without normalization')

# Plot normalized confusion matrix
plt.figure(figsize=(8,5))
plot_confusion_matrix(cnf_matrix, classes=class_names, normalize=True,
                      title='Normalized confusion matrix')

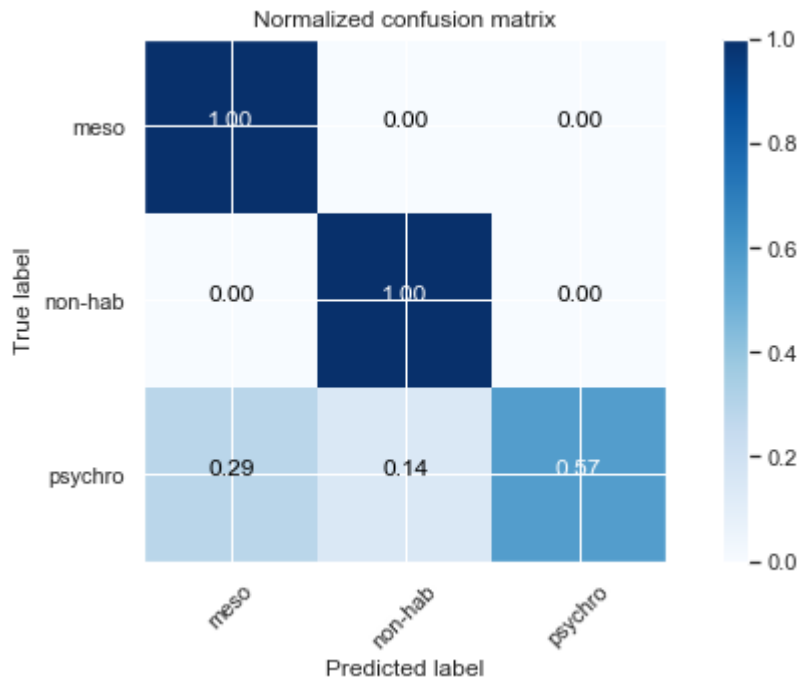
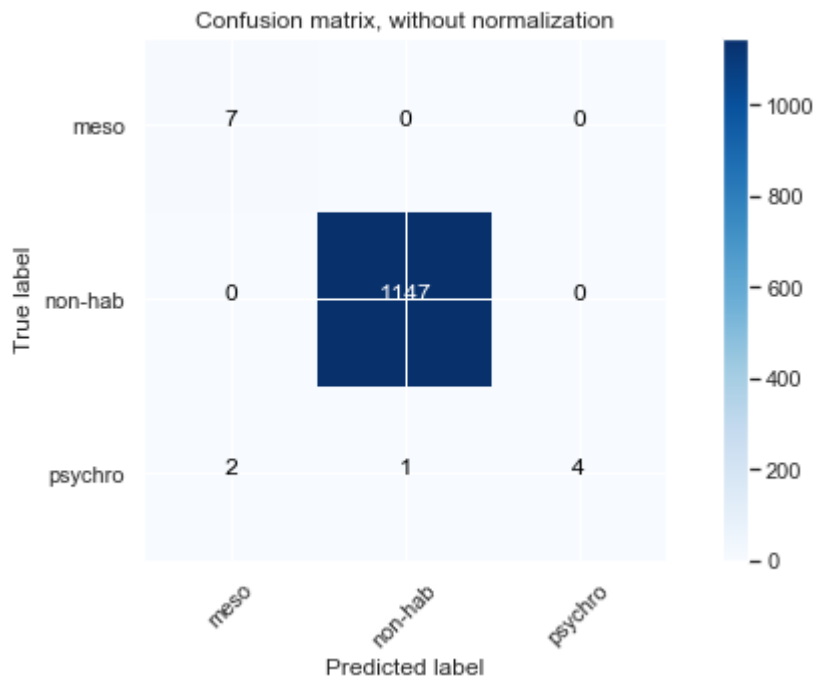
plt.show()
```

Confusion matrix, without normalization

```
[[ 7  0  0]
 [ 0 1147  0]
 [ 2  1  4]]
```

Normalized confusion matrix

```
[[1.  0.  0. ]
 [0.  1.  0. ]
 [0.29 0.14 0.57]]
```



Fitting Random Forest to the training set

```
In [50]: # Fitting Random Forest Classification to the Training set
from sklearn.ensemble import RandomForestClassifier
classifier = RandomForestClassifier(n_estimators = 10, criterion = 'entropy', random_state = 0)
classifier.fit(X_train, y_train)

# Predicting the Test set results
y_pred = classifier.predict(X_test)
```

```
In [51]: # Compute confusion matrix
cnf_matrix_RF = confusion_matrix(y_test, y_pred)
cnf_matrix = cnf_matrix_RF
np.set_printoptions(precision=2)

# Plot non-normalized confusion matrix
plt.figure(figsize=(8,5))
plot_confusion_matrix(cnf_matrix, classes=class_names,
                      title='Confusion matrix, without normalization')

# Plot normalized confusion matrix
plt.figure(figsize=(8,5))
plot_confusion_matrix(cnf_matrix, classes=class_names, normalize=True,
                      title='Normalized confusion matrix')

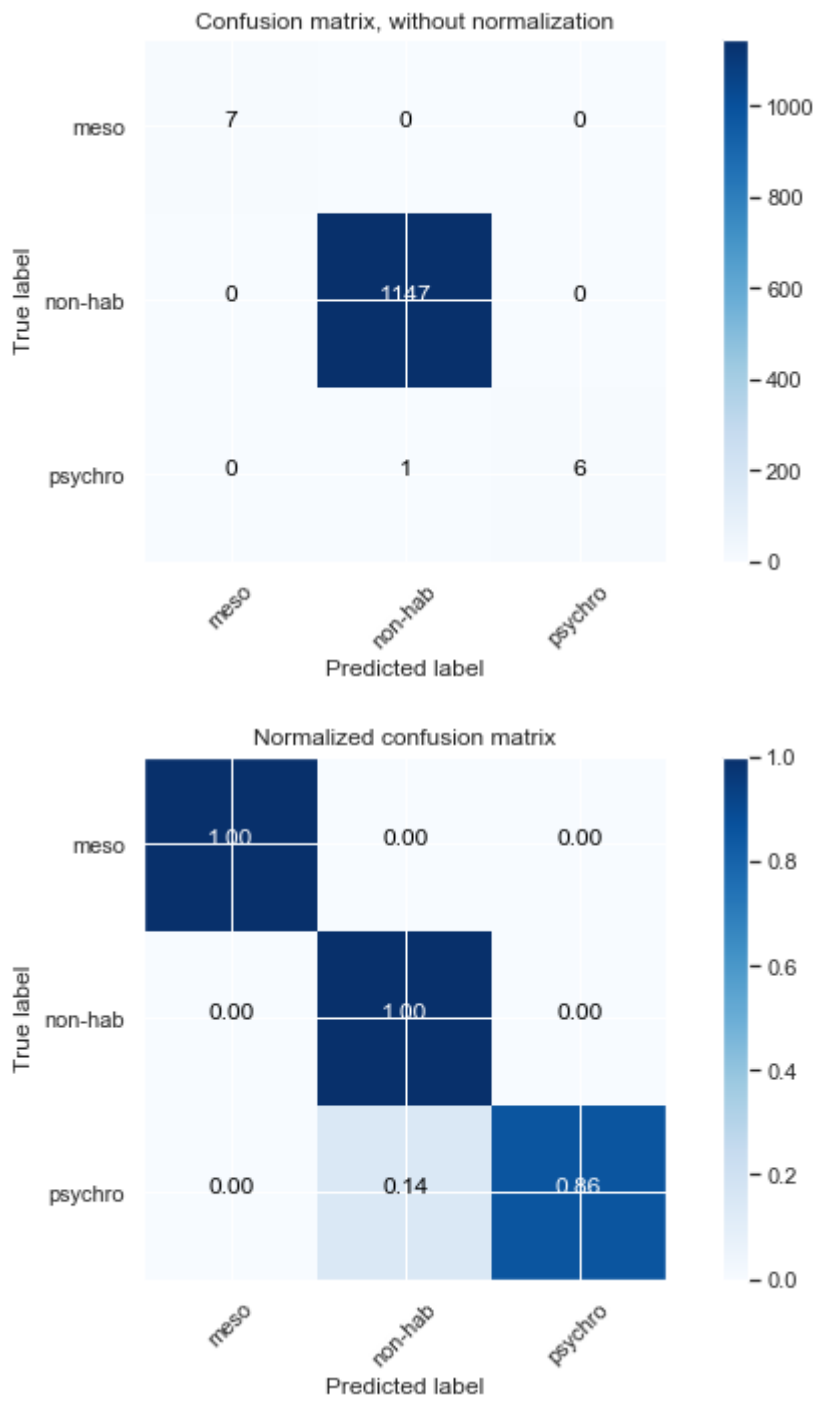
plt.show()
```

Confusion matrix, without normalization

```
[[ 7  0  0]
 [ 0 1147  0]
 [ 0  1  6]]
```

Normalized confusion matrix

```
[[1.  0.  0. ]
 [0.  1.  0. ]
 [0.  0.14 0.86]]
```



Fitting K-nearest Neighbours to the training set

```
In [52]: # Fitting K-NN to the Training set
from sklearn.neighbors import KNeighborsClassifier
classifier = KNeighborsClassifier(n_neighbors = 5, metric = 'minkowski', p = 2)
classifier.fit(X_train, y_train)

# Predicting the Test set results
y_pred = classifier.predict(X_test)
```

```
In [53]: # Compute confusion matrix
cnf_matrix_KNN = confusion_matrix(y_test, y_pred)
cnf_matrix = cnf_matrix_KNN
np.set_printoptions(precision=2)

# Plot non-normalized confusion matrix
plt.figure(figsize=(8,5))
plot_confusion_matrix(cnf_matrix, classes=class_names,
                      title='Confusion matrix, without normalization')

# Plot normalized confusion matrix
plt.figure(figsize=(8,5))
plot_confusion_matrix(cnf_matrix, classes=class_names, normalize=True,
                      title='Normalized confusion matrix')

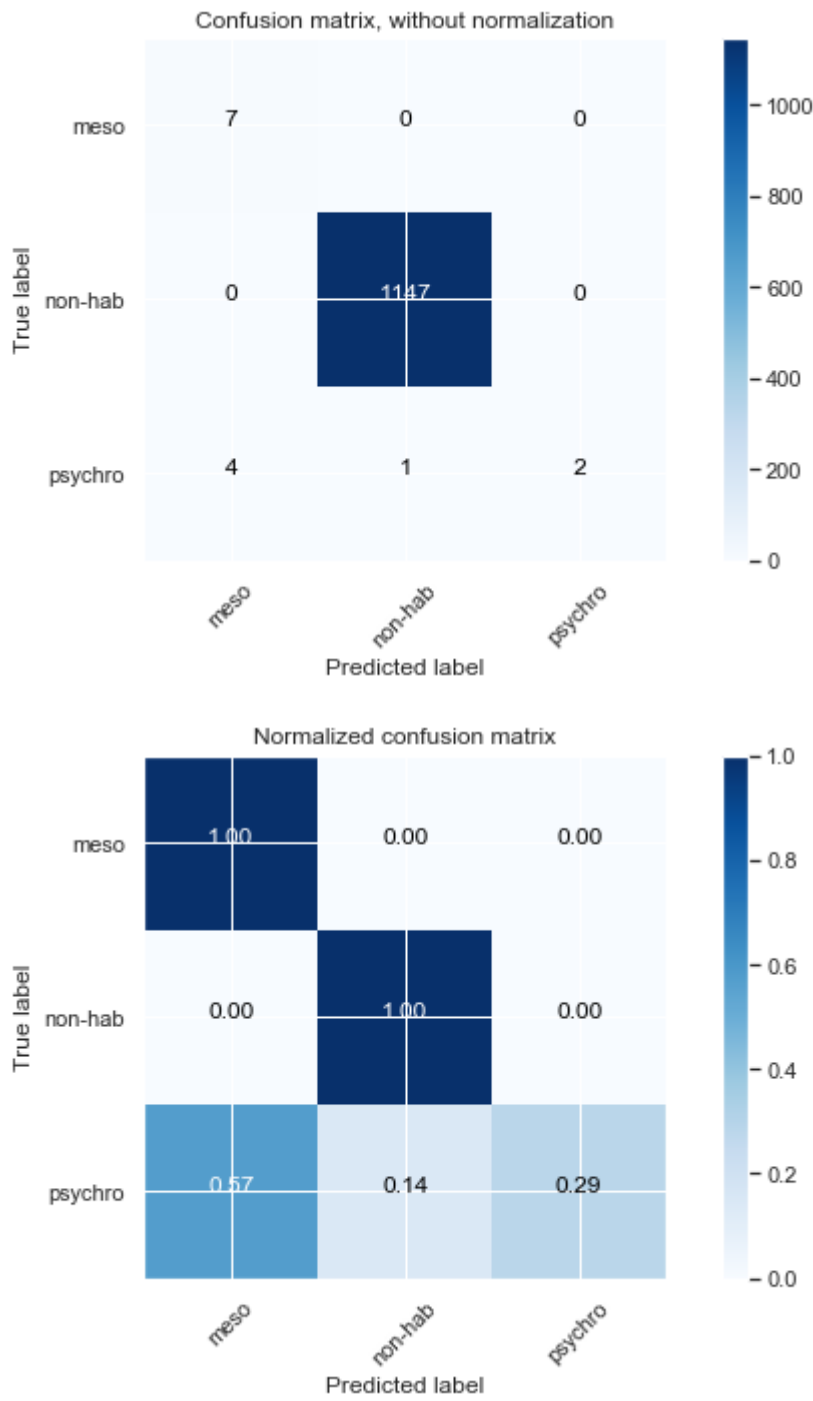
plt.show()
```

Confusion matrix, without normalization

```
[[ 7  0  0]
 [ 0 1147  0]
 [ 4  1  2]]
```

Normalized confusion matrix

```
[[1.  0.  0. ]
 [0.  1.  0. ]
 [0.57 0.14 0.29]]
```



NOTE

- 1. Since we have find only a few habitable planets, our dataset is heavily dominated by the Non-Habitable class.
- 2. High accuracy is mostly because of Non-Habitable class dominance which is influencing the rate of false positive predictions.
- 3. Combining other detection methods like Image Classification, Time Series Data Classification,Luminosity Calculation can make this project better by providing us better accuracy by reducing the number of false positives in our prediction.
- 4. Using different algorithms can also reduce the percentage of errors in our project.

Appendix:

[1] RAW Data Source: <http://phl.upr.edu/projects/habitable-exoplanets-catalog/data/database> (<http://phl.upr.edu/projects/habitable-exoplanets-catalog/data/database>) PHL's Exoplanets Catalog Last Update: July 2, 2018 Introduction The PHL's Exoplanets Catalog (PHL-EC) contains observed and modeled parameters for all currently confirmed exoplanets from the Extrasolar Planets Encyclopedia and NASA Kepler candidates from the NASA Exoplanet Archive, including those potentially habitable. It also contains a few still unconfirmed exoplanets of interest. The main difference between PHL-EC and other exoplanets databases is that it contains more estimated stellar and planetary parameters, habitability assessments with various habitability metrics, planetary classifications, and many corrections. Some interesting inclusions are the identification of those stars in the Catalog of Nearby Habitable Systems (HabCat), the apparent size and brightness of stars and planets as seen from a vantage point (i.e. moon-Earth distance), and the location constellation of each planet.

[2] A Comparative Study in Classification Methods of Exoplanets: Machine Learning Exploration via Mining and Automatic Labeling of the Habitability Catalog Surbhi Agrawal¹ ?, Suryoday Basak¹ , Snehanshu Saha^{1†}, Marialis Rosario-Franco² , Swati Routh³ , Kakoli Bora⁴ , Abhijit Jeremiel Theophilus¹ ¹Department of Computer Science and Engineering, PESIT Bangalore South Campus, Karnataka, India 560100 ²Physics Department, University of Texas at Arlington ³Physics Department, CPGS, Jain University ⁴Department of Information Science and Engineering, PESIT Bangalore South Campus, Karnataka, India 560100

[3] Emergence of a Habitable Planet - <https://link.springer.com/article/10.1007/s11214-007-9225-z> (<https://link.springer.com/article/10.1007/s11214-007-9225-z>)

[4] Honorary Mention: <http://curious.astro.cornell.edu/> (<http://curious.astro.cornell.edu/>)

This project can be found at my portfolio website www.kaustavbasu.me/minor-project-fifth-sem

END