A practical guide to ATSP2K

P. Jönsson

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# Chapter 1

# The ATSP2K package

## 1.1 Application programs and tools

The new version of the ATSP2K package [1] consists of a number of application programs and tools. The application programs and tools, along with the underlying theory, are described in the original write-ups [2-13]. See also the book Computational Atomic Structure by Froese Fischer, Brage, and Jönsson [14].

Below is a partial list of programs in the package:

- 1. Routines that generate a configuration state list (CSL):
  - (a) gencl generate a configuration state list (CSL) from lists of reference configurations
  - (b) lsgen generate a CSL using rules
  - (c) lsreduce include only CSFs that have at least one non-zero matrix element with a CSF of a reference list
- 2. nonh compute angular coefficients
- 3. nonhz compute angular coefficients assuming a partition of the configuration state list into zero- and first order spaces (optional)
- 4. mchf determine orbitals and mixing coefficients
- 5. mchf\_C determine orbitals and mixing coefficients, assuming c.lst on disc (medium size calculations, to be used ONLY if ordinary mchf fails on current operating system)
- 6. mchf\_HC determine orbitals and mixing coefficients, assuming c.lst and hmx.lst (large size calculations, to be used ONLY if ordinary mchf fails on current operating system)
- 7. bpic perform Breit-Pauli CI calculation, where matrix elements computed on the fly
- 8. bp\_ang + bp\_mat + bp\_eiv perform Breit-Pauli CI calculation, by separating angular integration and saveon disc, the construct matrix and diagonalize (fine eigenvalues). Suitable in smaller cases for an iso-electronic sequence.
- 9. Routines for computing transition probabilities:
  - (a) trans perform transition calculations in a single orthonormal basis
  - (b) biotr perform biorthogonal transformation and subsequent transition calculation between states.
- 10. hfs compute hyperfine interactions

A number of generally short programs have been developed as tools to facilitate a computational procedure.

- 1. comp: Given the name of a set of files, and a tolerance, this program displays the composition of each state in the file by listing the CSFs and their expansion coefficients in order of decreasing magnitude provided the magnitude is greater than the tolerance. The name.c option usually may not apply (seen condens below). This composition information may be of importance in determining the multi-reference set or the extent of correlation. It is also essential for resolving naming conflicts.
- 2. condens: Given the name of a case, the source for the expansion coefficients (usually .l or .j) and whether results are to be sorted, the program produces a file cfg.out that includes only those CSFs for which  $\sqrt{\sum_i c_i^2} \geq T$ , where T is the cut-off tolerance. Since mchf no longer produces a cfg.out with the expansion coefficients included, the name.c option cannot usually be applied. The coefficients are included in the condensed output file, cfg.out. To restore this file to the present .c format (without expansion coefficients), move the file to clist.inp and run lsgen selecting the "r" mode (for restore). Note that condensing may destroy the "closed under de-excitation" requirement of biotr but when the tolerance is sufficiently small, no problems have been encountered. In a systematic method, where maximum principal quantum number is increased by unity from one iteration to the next, it is unlikely that a CSF with a high n will remain and the equivalent CSF with a lower n be deleted.
- 3. tables: This is a program written in C++ that takes a name.lsj file, usually a concatenated file of all the .lsj transition files for a given atom or ion, and finds the energy level structure of the levels and the multiplet transition arrays. The tables posted at the website http://atoms.vuse.vanderbilt.edu are examples of tables produced by the tables program. When an energy level is present in the .lsj file with two different energies, the higher level is considered to be unphysical. It and all data associated with this level are removed from the table. The user may also specify certain levels as "unphysical" in which case they will be removed. Finally, the program computes the lifetimes of the levels from the transition data provided in the file.
- 4. T\_dependence: Given a .c and a corresponding .j file, this program displays the term dependence of each state included in the file [?]. This gives an indication of LS term mixing in the wave function of a state. The Breit-Pauli programs all assign labels to states according to the largest expansion coefficient. When this process produces the same label for two different states a careful analysis is needed. The LS value should be that term with the largest composition and within that LS, the largest expansion coefficient identifies the label. For more on labeling, see [?].
- 5. relabel This utility reads radial functions in turn from wfn.inp. For each function, the user may enter either a blank, "d", or a new 3-character label for which the respective action is to write the radial function unchanged into wfn.out, skip (or delete) the radial function, write the radial function changing the displayed label to the entered label.
- 6. select This routine selects those CSFs from a designated list that contain orbitals specified by the user. This may be useful when expansions that have been condensed are extended with new CSFs to be included to first-order. The selected CSFs may be appended to the condensed list.
- 7. w\_format and w\_unformat: Binary file formats are not always compatible in going from one system environment to another. w\_format takes a wfn.inp file (in .w format) and produces a formatted wfn.fmt file. The program w\_unformat reads wfn.fmt and produces the binary file (in .w format) wfn.out. Some loss of accuracy can be expected in the process but the radial functions are tabulated to 11-digits of accuracy.

### References

- 1. ATSP2K: C. Froese Fisc her, G. Tachiev, G. Gaigalas, M.R. Godefroid Comput. Phys. Commun. 176 559 (2007)
- 2. MCHF LIBRARIES: C.Froese Fischer, Comput. Phys. Commun. 64 399 (1991)
- 3. MCHF GENCL: C. Froese Fischer, B. Liu, Comput. Phys. Commun. 64 406 (1991)
- 4. MCHF NONH: A. Hibbert, C. Froese Fischer, Comput. Phys. Commun. 64 417 (1991)
- 5. MCHF\_88: C. Froese Fischer, Comput. Phys. Commun. 64 431 (1991)
- 6. MCHF\_BREIT: A. Hibbert, R. Glass, C. Froese Fischer Comput. Phys. Commun. 64 455 (1991)
- 7. MCHF CI: C. Froese Fischer, Comput. Phys. Commun. 64 473 (1991)
- 8. MCHF\_MLTPOL: C. Froese Fischer, M.R. Godefroid, A. Hibbert, Comput. Phys. Commun. 64 486 (1991)
- 9. MCHF\_LSTR AND MCHF\_LSJTR: C.Froese Fischer, M.R. Godefroid, Comput. Phys. Commun. 64 501 (1991)
- 10. MCHF AUTO: C.Froese Fischer, T. Brage, Comput. Phys. Commun. 74 381 (1993)
- 11. MCHF\_HFS: P. Jönsson, C.-G. Wahlstrom, C.Froese Fischer, Comput. Phys. Commun. 74 399 (1993)
- 12. MCHF\_ISOTOPE: C. Froese Fischer, L. Smentek-Mielczarek, N. Vaeck, G. Miecznik, Comput. Phys. Commun. 74 415 (1993)
- 13. MCHF\_LSGEN: L. Sturesson and C. Froese Fischer, Comput. Phys. Commun. 74 432 (1993)
- 14. C. Froese Fischer, T. Brage, P. Jönsson, Computational Atomic Structure an MCHF approach, IoP, (1997)

## 1.2 File naming conventions, program and data flow

Passing of information between different programs is done through files. This process is greatly facilitated through file naming conventions. A name is associated with the results and an extension that defines the contents and format of the file. Thus the file name becomes name.extension. Common extensions are listed in Table 1.1.

To run ATSP2K a number of programs need to be run in a pre-determined sequence. Figure 1.1 displays a typical sequence of block version program calls to evaluate different expectation values. The resulting flow of files is displayed in Figure 1.2.

Table 1.1: Table of common extensions.

extension	data in the file
.с	configuration state function (CSF) expansion
.W	radial wave functions (numerical values in binary form)
.1	expansion coefficients from a non-relativistic $(LS)$ calculation
·j	expansion coefficients from a Breit-Pauli $(LSJ)$ calculation
.h	hyperfine data
.t	term dependence of a .j file
.ls	transition probability data a non-relativistic $(LS)$ calculation
.lsj	transition probability data a Breit-Pauli $(LSJ)$ calculation

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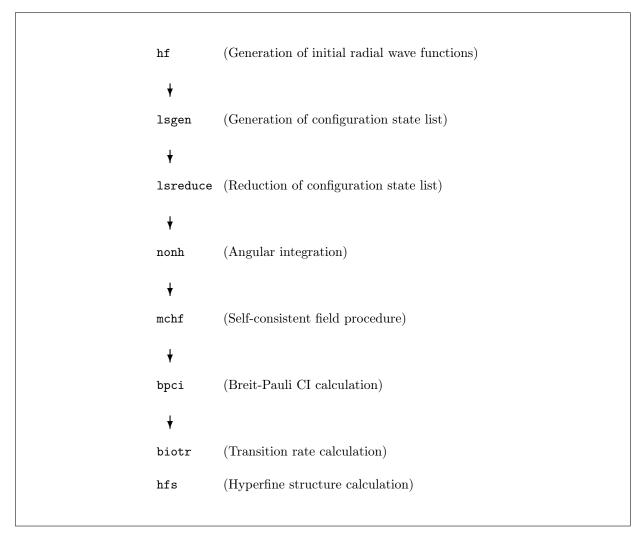


Figure 1.1: Typical sequence of block version program calls to evaluate different expectation values such as transition rates and hyperfine structure constants.

```
Output; wfn.out
hf
lsgen
            Output; clist.out
lsreduce
            Input; mrlist, cfg.inp
            Output; cfg.out
            Input; cfg.inp
nonh
            Output; number of different angular files
mchf
            Input; cfg.inp, wfn.inp
            Output; LS\pi.1, wfn.out, summry
            Input; name.c, name.w
bpci
            Output; name.j
biotr
            Input; name1.c, name1.w, name1.j, name2.c, name2.w, name2.j
            Output; name1.name2.lsj
hfs
            Input; name.c, name.w, name.1
            Output; name.h
```

Figure 1.2: Flow of files for a normal sequence of program runs.

### 1.3 Generating lists of configuration state functions

Exploring different correlation models and generating lists of configuration state functions (CSFs) is a major task of the computation. The ATSP2K provides several programs for performing this task. For generating small lists of CSFs it is often best to use the gencl program. To generate expansion based on the notion of excitations from subshells to an active set of orbitals it is often advantageous to use the lsgen program. Different restrictions can be put on the excitations and it is possible to describe core-valence correlation where at most one excitation is allowed from subshells of the core. To make sure that the generated CSFs interacts with the CSFs in multireference the program lsreduce should be used. Before continuing the reader is advised to study the write-up of the lsreduce program [L. Sturesson and C. Froese Fischer, Comput. Phys. Commun. 74, 432 (1993)].

## 1.4 Spectroscopic orbitals

The "spectroscopic orbitals" are those where node counting is required to ensure that the self-consistent field procedure converges to the desired solution [C. Froese Fischer, T. Brage, P. Jönsson, Computational Atomic Structure - an MCHF approach, IoP, 1997]. Spectroscopic orbitals build the reference CSFs and often have occupation numbers near unity or more. All other orbitals are "correlation orbitals". If the self-consistent field procedure fails for spectroscopic orbitals, i.e. the wrong number of nodes are obtained with a subsequent program halt.

# Chapter 2

# Running the application programs

In this chapter we demonstrate the use of the application programs of ATSP2K in three cases described below. The use of the tools of the ATSP2K package is described in the next chapter. The data written to the output files are explained and discussed in chapter 4.

## 2.1 First example: $1s^22s$ $^2S$ and $1s^22p$ $^2P$ in Li I

The first example is for  $1s^22s$   $^2S_{1/2}$  and  $1s^22p$   $^2P_{1/2,3/2}$  in Li.

#### Overview

- 1. Perform HF calculation for  $1s^22s$   $^2S$ .
- 2. Generate n = 3 CAS configuration list for  $1s^2 2s$   $^2 S$ .
- 3. Perform angular integration.
- 4. Perform self-consistent field calculation.
- 5. Save output to 2S\_CAS\_3.
- 6. Calculate hfs.
- 7. Generate n = 3 CAS Breit configuration list for  $1s^2 2s {}^2 S_{1/2}$ .
- 8. Perform CI Breit-Pauli calculation
- 9. Perform HF calculation for  $1s^22p$   $^2P$ .
- 10. Generate n = 3 CAS configuration list for  $1s^2 2p$   $^2P$ .
- 11. Perform angular integration.
- 12. Perform self-consistent field calculation.
- 13. Save output to 2P\_CAS\_3.
- 14. Calculate hfs.
- 15. Generate n=3 CAS Breit configuration list for  $1s^2 2p^2 P_{1/2,3/2}$ .
- 16. Perform CI Breit-Pauli calculation
- 17. Compute the transition rates from the CI wave functions.

### Program input

In the test-runs input is marked by >> and >>3, for example, indicate that the user should input 3 and then strike the return key. When >> is followed by blanks just strike the return key.

RUN HF FOR 1s(2)2s 2S OUTPUT FILE: wfn.out, hf.log >>HF \_\_\_\_\_\_ HARTREE-FOCK.96 \_\_\_\_\_ THE DIMENSIONS FOR THE CURRENT VERSION ARE: NWF= 20 NO=220 START OF CASE ========= Enter ATOM, TERM, Z Examples: 0,3P,8. or Oxygen,AV,8. >>Li,2S,3. List the CLOSED shells in the fields indicated (blank line if none) ... ... ... ... ... ... etc. >>1s Enter electrons outside CLOSED shells (blank line if none) Example: 2s(1)2p(3) >>2s(1)There are 2 orbitals as follows: 1s 2s Orbitals to be varied: ALL/NONE/=i (last i)/comma delimited list/H >>all Default electron parameters ? (Y/N/H) >>y Default values for remaining parameters? (Y/N/H)

. . . . . . . . . . . . . . . . . . .

Highest principal quantum number, n? (1..15)

>>3

```
______
        SCF CONVERGENCE CRITERIA (SCFTOL*SQRT(Z*NWF)) = 1.6D-06
        ED
                                        ΑZ
                                                     NORM
                                                              DPM
                  EL ED AZ NORT 5...

1s 4.9554830 9.2603703 1.0000000 1.43D-08

2s 0.3926457 1.4468006 0.9999999 1.28D-08
     < 1s| 2s >= 8.7D-09
    TOTAL ENERGY (a.u.)
    _____

        Non-Relativistic
        -7.43272693
        Kinetic
        7.43272693

        Relativistic Shift
        -0.00054376
        Potential
        -14.86545385

        Relativistic
        -7.43327069
        Ratio
        -2.000000000

Additional parameters ? (Y/N/H)
Do you wish to continue along the sequence ?
>>n
END OF CASE
========
************************************
        COPY FILES
       Keep a copy of the wfn.out in 2SeDF.w file for reference
       Keep a copy of hf.log in 2SDF.log
cp wfn.out wfn.inp
cp wfn.out 2SeDF.w
cp hf.log 2SeDF.log
RUN LSGEN TO GENERATE N = 3 CAS CONFIGURATION LIST FOR 2S
        OUTPUT FILES: clist.out, clist.log
New list, add to existing list, expand existing list, optimized sorting, restored order or quit? (*
Breit or MCHF? (B/*)
Default, symmetry or user specified ordering? (*/s/u)
```

```
Highest orbital angular momentum, 1? (s..d)
Are all these nl-subshells active? (n/*)
Limitations on population of n-subshells? (y/*)
Highest n-number in reference configuration? (1..3)
Number of electrons in 1s? (0..2)
Number of electrons in 2s? (0..2)
Number of electrons in 2p? (0..6)
Resulting term? (1S, 3P, etc.)
>>2S
Number of excitations = ? (0..3)
27 configuration states have been generated.
Generate a second list? (y/*)
27 configuration states in the final list.
The generated file is called clist.out.
FORTRAN STOP
************************************
      IT IS ADVISABLE TO SAVE THE LSGEN LOG-FILE TO HAVE A RECORD ON
       HOW THE CONFIGURATION LISTS GENERATION WAS DONE
************************************
>>cp clist.log 2SeCAS3.log
>>cp clist.out cfg.inp
RUN NONH TO GENERATE ENERGY EXPRESSION
       INPUT FILES: cfg.inp
      OUTPUT FILES: cfg.h, yint.lst, c.lst, ih.=n.lst
>>nonh
input file is cfg.inp ...
                 _____
                     N O N H
                               2000
                 ______
```

1s 2s 2p 3s 3p 3d

THERE ARE O CLOSED SUBSHELLS COMMON TO ALL CONFIGURATIONS AS FOLLOWS:

Allocating space for 214 integrals processing 2Se with 27 configurations

277 non-zero matrix elements

92 NF 76 NG 446 NR 109 NL

723 Total number of integrals

end-of-file clist!!!

\* INPUT FILES: wfn.inp (optional), cfg.inp, angular files

\* OUTPUT FILES: wfn.out, 2Se.1, summry

\* NOTE1: Mixing coefficients written to a file named by LS-symmetry

and parity, in this case 2Se.1

\* NOTE2: We force the program to iterate to higher precision than

\* default which is a good thing for small calculations

>>mchf

M C H F ... 2000

\_\_\_\_\_

THE DIMENSIONS FOR THE CURRENT VERSION ARE:

NWD= 60 NO=220 Lagrange Multipliers=800

START OF CASE

\_\_\_\_\_

ATOM, Z in FORMAT(A, F):

>>Li,3.

cfg.inp has configurations for

1 terms

Enter eigenvalues and weights: one line per term, eigenvalues with weights in parenthesis and separated by commas, default is  $1.0\,$ 

2Se

>>1

There are 6 orbitals as follows:

1s 2s 2p 3s 3p 3d

```
Enter orbitals to be varied: (ALL, NONE, SOME, NIT=, comma delimited list)
Enter those that are spectroscopic
>>1s,2s
Default electron parameters ? (Y/N)
Default values for other parameters? (Y/N)
>>n
Default values (NO, REL, STRONG) ? (Y/N)
>>y
Li
          3. 220 6 6 0 F
Default values for PRINT, CFGTOL, SCFTOL ? (Y/N)
Input free FORMAT(L, F, F)
>>.f.,1.e-50,1.e-50
Default values for NSCF, IC ? (Y/N)
Input free FORMAT(I, I)
>>300,0
Default values for ACFG, LD, TRACE? (Y/N)
>>y
 . . . . . . . . . . . . . . . . .
         ITERATION NUMBER 111
         _____
         CONVERGENCE CRITERIA: ENERGY (CFGTOL)
                                                      = 1.0D-50
                            :FUNCTION(SCFTOL*SQRT(Z*NWF))= 4.2D-50
         E(2s 1s) =
                       0.02212 E( 1s 2s) =
                                                 0.01109
         E(3s 1s) =
                       28.98616 E(1s 3s) =
                                                 0.03558
                                 E(2s 3s) =
         E(3s 2s) =
                       4.20140
                                                 0.01028
         E(3p 2p) =
                       51.03003 E( 2p 3p) =
                                                 5.25729
                   EL
                              ED
                                           ΑZ
                                                         NORM
                                                                  DPM
     < 1s| 2s >= 4.2D-11
     < 1s| 3s>=-4.5D-10
                           5.0342669
                                         9.2695740
                                                    1.0000000
                                                                 5.17D-10
     < 2s| 3s > = -7.3D - 09
     < 1s| 2s>= 1.0D-08
                          2s
     < 2p \mid 3p > = 6.0D-09
                                                    1.0000000** 9.70D-09
                          15.9592124
                                        4.7810460
     < 1s| 3s > = -4.4D - 07
     < 2s| 3s > = -5.6D - 08
```

3s

16.2077493

14.8282834

1.0000000\*\* 5.32D-07

```
< 2p| 3p > = -6.7D - 09
                       30.8279905 66.8115267 1.0000000** 2.45D-08
                  Зр
                       35.1829408 88.8117857 1.0000000** 2.94D-10
       ETOTAL= -7.47318427 Loops, DeltaE, Res.: 2 3.955D-16 7.343D-09
     1 \quad 0.9984151 \qquad \quad 2 \quad 0.0004030 \qquad \quad 3 \quad 0.0002912 \qquad \quad 4 \quad -0.0024959
     5 0.0016210
                  6 -0.0009009
                                  7 0.0029624
                                                 8 -0.0015623
       Sum of ETOTAL : -7.47318427
 DeltaE = 1.77635683940025046E-015 Sum_Energy = -7.4731842652568732
    ENERGY (a.u.)
         Total
                            -7.473184265
         Potential
                          -14.946368479
         Kinetic
                             7.473184213
         Ratio
                              2.00000007
SAVE OUTPUT FILES
>>cp wfn.out 2SeCAS3.w
>>cp cfg.inp 2SeCAS3.c
>>cp 2Se.l 2SeCAS3.l
>>cp summry 2SeCAS3.s
************************************
        RUN HFS TO COMPUTE HYPERFINE INTERACTION CONSTANTS
        INPUT FILES: 2SeCAS3.c, 2SeCAS3.w, 2SeCAS3.1
        OUTPUT FILES: 2SeCAS3.h
        NOTE1: THE OPTION MCHF IS OBSOLETE AND ADERES TO THE OLD OUTPUT
        FORMAT OF MCHF. NOW YOU NEED ALWAYS ASK FOR INPUT FROM A CI
        CALCULATION.
        NOTE2: THE HFS PROGRAM ONLY COMPUTES THE HYPERFINE INTERACTION
        FROM ONE STATE AND YOU HAVE TO SPECIFY THE POSITION OF THE LEADING *
        CSF OF THIS STATE. AT SOME POINT SOMEONE SHOULD RECODE THE MODULE
*************************************
Name of state ...
>>2SeCAS3
                  Hyperfine structure calculation
Electron density at the nucleus ? (Y/N)
>>y
Indicate the type of calculation
0 => diagonal A and B factors only;
1 => diagonal and off-diagonal A and B factors;
Input from an MCHF (M) or CI (C) calculation ?
```

>>

```
Is the CI calculation J dependant ? (Y/N)
Give the index of the dominant cfg in the CI
expansion for which the hfs is to be calculated ?
Give 2*I and nuclear dipole and quadrupole moments (in n.m. and barns)
>>2,1,1
                 The configuration set
               27 CONGIGURATIONS):
STATE (WITH
THERE ARE 6 ORBITALS AS FOLLOWS:
      1s 2s 2p 3s 3p 3d
THERE ARE O CLOSED SUBSHELLS COMMON TO ALL CONFIGURATIONS AS FOLLOWS:
 10
   ja =
  ja =
              20
PER TEST
END OF CASE
Total CPU time was 0.000 minutes
RUN LSGEN TO GENERATE N = 3 CAS BREIT CONFIGURATION LIST FOR 2S
        OUTPUT FILES: clist.out, clist.log
*************************************
New list, add to existing list, expand existing list, optimized sorting, restored order or quit? (*
Breit or MCHF? (B/*)
Default, symmetry or user specified ordering? (*/s/u)
Highest principal quantum number, n? (1..15)
Highest orbital angular momentum, 1? (s..d)
Are all these nl-subshells active? (n/*)
```

```
Limitations on population of n-subshells? (y/*)
>>
Highest n-number in reference configuration? (1..3)
Number of electrons in 1s? (0..2)
Number of electrons in 2s? (0..2)
Number of electrons in 2p? (0..6)
Maximum 2*J-value? (0..)
Minimum 2*J-value? (0..1)
Maximum (2*S+1)-value? (1..9)
Minimum (2*S+1)-value? (1..9)
Maximum resulting angular momentum? (S..N/N=*)
Minimum resulting angular momentum? (S..N/S=*)
Number of excitations = ?(0..3)
>>3
79 configuration states have been generated.
Generate a second list? (y/*)
79 configuration states in the final list.
The generated file is called clist.out.
FORTRAN STOP
*************************************
       COPY FILES
       IT IS ADVISABLE TO SAVE THE LSGEN LOG-FILE TO HAVE A RECORD ON
       HOW THE CONFIGURATION LISTS GENERATION WAS DONE
***********************************
>>cp clist.log 2SeCAS3BREIT.log
>>cp clist.out 2SeCAS3BREIT.c
*************************************
        COPY WAVE FUNCTION FILES TO PREPARE FOR THE BREIT-PAULI RUN
>>cp 2SeCAS3.w 2SeCAS3BREIT.w
***********************************
        RUN CI BREIT-PAULI
       INPUT FILES: 2SeCAS3BREIT.c, 2SeCAS3BREIT.w
       OUTPUT FILES: 2SeCASBREIT.j
********************************
>>bpci
Enter ATOM, relativistic (Y/N) with mass correction (Y/N)
```

```
>>2SeCAS3BREIT,y,n
Restarting (Y/y) ?
>>n
Use existing Matrix and \frac{1}{j} initial guess (\frac{y}{y})?
Enter Maximum and minimum values of 2*J
>>1,1
Enter eigenvalues: one line per term, eigenvalues separated by commas
2*J = 1
>>1
                   BREIT-PAULI
                  _____
Indicate the type of calculation
0 => non-relativistic Hamiltonian only;
1 => one or more relativistic operators only;
2 => non-relativistic operators and selected relativistic:
All relativistic operators ? (Y/N)
>>y
THE TYPE OF CALCULATION IS DEFINED BY THE FOLLOWING PARAMETERS -
     BREIT-PAULI OPERATORS IREL = 2
     PHASE CONVENTION PARAMETER
                                ICSTAS = 1
                          THE CONFIGURATION SET
STATE (WITH 79 CONGIGURATIONS):
 -----
THERE ARE 6 ORBITALS AS FOLLOWS:
      1s 2s 2p 3s 3p 3d
THERE ARE O CLOSED SUBSHELLS COMMON TO ALL CONFIGURATIONS AS FOLLOWS:
All Interactions? (Y/N):
Default Rydberg constant (y/n)
>>y
```

659

. . . . .

Allocating space for 783 integrals
Alcmat allocations for idisk= 0 nze =

J = 1 1

Entering LSJMAT with 2J = 1 NUME = 1

7.8543171982727995E-009

LSJMAT with idisk= 0 Nze = 2636

Summary of Davidson Performance

Number of Iterations: 10
Shifted Eigval's: -0.2830587230149465
Delta Lambda: 3.3306690738754696E-016

1 Eigenvalues found
Finished with Davidson
Leaving LSJMAT
onlydvd F
ILS = 0
Finished with the file

\begin{verbatim}

Residuals:

\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*

\* RUN HF FOR 1s(2)2p 2P

>>HF

H A R T R E E - F O C K . 96

THE DIMENSIONS FOR THE CURRENT VERSION ARE:

NWF= 20 NO=220

START OF CASE

Enter ATOM, TERM, Z

Examples: 0,3P,8. or Oxygen,AV,8.

>>Li,2P,3.

List the CLOSED shells in the fields indicated (blank line if none)

... ... ... ... ... ... etc.

>>1s

```
Enter electrons outside CLOSED shells (blank line if none)
Example: 2s(1)2p(3)
>>2p(1)
There are 2 orbitals as follows:
  1s 2p
Orbitals to be varied: ALL/NONE/=i (last i)/comma delimited list/H
>>all
Default electron parameters ? (Y/N/H)
>>y
Default values for remaining parameters? (Y/N/H)
 . . . . . . . . . . . . . . . . . . .
        ITERATION NUMBER 5
          -----
         SCF CONVERGENCE CRITERIA (SCFTOL*SQRT(Z*NWF)) = 3.9D-07
                                             AZ
                    EL
                              ED
                                                           NORM
                                                                     DPM
                    1s 5.0614517 9.2618326 0.9999999 2.29D-08
2p 0.2573450 0.4261374 1.0000000 9.86D-09
    TOTAL ENERGY (a.u.)
     -----

        Non-Relativistic
        -7.36506966
        Kinetic
        7.36506968

        Relativistic Shift
        -0.00053279
        Potential
        -14.73013934

          Relativistic -7.36560245 Ratio -1.999999997
Additional parameters ? (Y/N/H)
>>n
Do you wish to continue along the sequence ?
END OF CASE
*************************************
         COPY FILES
         Keep a copy of the wfn.out in 2PoDF.w file for reference
         Keep a copy of hf.log in 2PoDF.log
```

RUN NONH TO GENERATE ENERGY EXPRESSION

```
cp wfn.out wfn.inp
cp wfn.out 2PoDF.w
cp hf.log 2PoDF.log
************************************
        RUN LSGEN TO GENERATE N = 3 CAS CONFIGURATION LIST FOR 2P
        OUTPUT FILES: clist.out, clist.log
New list, add to existing list, expand existing list, optimized sorting, restored order or quit? (*
Breit or MCHF? (B/*)
Default, symmetry or user specified ordering? (*/s/u)
Highest principal quantum number, n? (1..15)
Highest orbital angular momentum, 1? (s..d)
Are all these nl-subshells active? (n/*)
Limitations on population of n-subshells? (y/*)
>>
Highest n-number in reference configuration? (1..3)
Number of electrons in 1s? (0..2)
Number of electrons in 2s? (0..2)
Number of electrons in 2p? (0..6)
Resulting term? (1S, 3P, etc.)
>>2P
Number of excitations = ? (0..3)
44 configuration states have been generated.
Generate a second list? (y/*)
44 configuration states in the final list.
The generated file is called clist.out.
FORTRAN STOP
************************************
       COPY FILES
        IT IS ADVISABLE TO SAVE THE LSGEN LOG-FILE TO HAVE A RECORD ON
        HOW THE CONFIGURATION LISTS GENERATION WAS DONE
>>cp clist.log 2PoCAS3.log
>>cp clist.out cfg.inp
***********************************
```

* INPUT FILES: cfg.inp * OUTPUT FILES: cfg.h, yint.lst, c.lst, ih.=n.lst ************************************
>>nonh
input file is cfg.inp
NONH 2000
THERE ARE 6 ORBITALS AS FOLLOWS:
1s 2s 2p 3s 3p 3d
THERE ARE O CLOSED SUBSHELLS COMMON TO ALL CONFIGURATIONS AS FOLLOWS:
Allocating space for 214 integrals processing 2Po with 44 configurations 649 non-zero matrix elements 152 NF 166 NG 1083 NR 191 NL 1592 Total number of integrals
end-of-file clist!!!
**************************************
*  NOTE2: We force the program to iterate to higher precision than  default which is a good thing for small calculations
**************************************
======================================

```
START OF CASE
=========
ATOM, Z in FORMAT(A, F):
>>Li,3.
cfg.inp has configurations for
                                1 terms
Enter eigenvalues and weights: one line per term, eigenvalues with weights
in parenthesis and separated by commas, default is 1.0
2Po
>>1
There are 6 orbitals as follows:
  1s 2s 2p 3s 3p 3d
Enter orbitals to be varied: (ALL,NONE,SOME,NIT=,comma delimited list)
Enter those that are spectroscopic
>>1s,2p
Default electron parameters ? (Y/N)
Default values for other parameters? (Y/N)
Default values (NO, REL, STRONG) ? (Y/N)
>>y
Li
          3. 220 6 6 0 F
Default values for PRINT, CFGTOL, SCFTOL ? (Y/N)
Input free FORMAT(L, F, F)
>>.f.,1.e-50,1.e-50
Default values for NSCF, IC ? (Y/N)
>>n
Input free FORMAT(I, I)
>>300,0
Default values for ACFG, LD, TRACE? (Y/N)
>>y
 . . . . . . . . . . . . . . . . .
         ITERATION NUMBER 62
          _____
         CONVERGENCE CRITERIA: ENERGY (CFGTOL)
                                                         = 1.0D-50
                             :FUNCTION(SCFTOL*SQRT(Z*NWF))= 4.2D-50
```

```
E( 2s 1s) = 25.62110 E( 1s 2s) = 0.03566
E( 3s 1s) = 125.27085 E( 1s 3s) = 0.00805
         E(3s 2s) = 72.76330 E(2s 3s) = 3.36051
         E(3p 2p) = 2.18254 E(2p 3p) = 0.00707
                        ED AZ NORM
                                                                 DPM
                   EL
     < 1s| 2s >= 9.9D-10
     < 1s| 3s>=-5.8D-11
                  1s 5.1380025 9.2694030 1.0000000 1.23D-09
     < 2s \mid 3s > = -2.8D - 07
     < 1s| 2s > = -3.6D - 07
                        15.1297623
                                       9.8051412 1.0000000** 8.16D-07
                   2s
     < 2p \mid 3p > = -7.7D - 11
                         2p
     < 1s| 3s>= 4.6D-08
     < 2s| 3s > = -3.2D - 07
                        32.0103498 25.1021580 1.0000000** 3.78D-07
                   3s
     < 2p| 3p>=-2.0D-10
                   3p 17.5853816 23.8010905 1.0000000** 2.18D-09
                   3d 32.5391413 81.3500754 1.0000000** 1.10D-09
        ETOTAL= -7.40458774 Loops, DeltaE, Res.: 2 3.608D-15 5.517D-08

      1
      0.9984119
      2
      0.0003758
      3
      -0.0006572
      4
      -0.0014719

      5
      -0.0042385
      6
      -0.0027196
      7
      -0.0003308
      8
      -0.0001325

        Sum of ETOTAL :
                           -7.40458774
 DeltaE = 6.21724893790087663E-015 Sum_Energy = -7.4045877355330409
    ENERGY (a.u.)
                            -7.404587736
-14.809175438
          Total
          Potential
          Kinetic
                               7.404587702
                                2.000000005
          Ratio
**********************************
         SAVE OUTPUT FILES
*************************************
>>cp wfn.out 2PoCAS3.w
>>cp cfg.inp 2PoCAS3.c
>>cp 2Po.1 2PoCAS3.1
>>cp summry 2PoCAS3.s
**********************************
         RUN HFS TO COMPUTE HYPERFINE INTERACTION CONSTANTS
        INPUT FILES: 2PoCAS3.c, 2PoCAS3.w, 2PoCAS3.1
        OUTPUT FILES: 2PoCAS3.h
         NOTE1: THE OPTION MCHF IS OBSOLETE AND ADERES TO THE OLD OUTPUT
         FORMAT OF MCHF. NOW YOU NEED ALWAYS ASK FOR INPUT FROM A CI
         CALCULATION.
```

```
NOTE2: THE HFS PROGRAM ONLY COMPUTES THE HYPERFINE INTERACTION
         FROM ONE STATE AND YOU HAVE TO SPECIFY THE POSITION OF THE LEADING *
         CSF OF THIS STATE. AT SOME POINT SOMEONE SHOULD RECODE THE MODULE
*************************************
Name of state ...
>>2PoCAS3
                   Hyperfine structure calculation
Electron density at the nucleus ? (Y/N)
>>y
Indicate the type of calculation
0 => diagonal A and B factors only;
1 => diagonal and off-diagonal A and B factors;
>>0
Input from an MCHF (M) or CI (C) calculation ?
Is the CI calculation J dependant ? (Y/N)
Give the index of the dominant cfg in the CI
expansion for which the hfs is to be calculated ?
>>1
Give 2*I and nuclear dipole and quadrupole moments (in n.m. and barns)
>>2,1,1
                   The configuration set
STATE (WITH 44 CONGIGURATIONS):
THERE ARE 6 ORBITALS AS FOLLOWS:
      1s 2s 2p 3s 3p 3d
THERE ARE O CLOSED SUBSHELLS COMMON TO ALL CONFIGURATIONS AS FOLLOWS:
 ja =
                20
   ja =
   ja =
                30
                40
   ja =
PER TEST
```

Total CPU time was 0.000 minutes

END OF CASE

COPY FILES

```
**********************************
        RUN LSGEN TO GENERATE N = 3 CAS BREIT CONFIGURATION LIST FOR 2P
        OUTPUT FILES: clist.out, clist.log
New list, add to existing list, expand existing list, optimized sorting, restored order or quit? (*
>>
Breit or MCHF? (B/*)
Default, symmetry or user specified ordering? (*/s/u)
Highest principal quantum number, n? (1..15)
Highest orbital angular momentum, 1? (s..d)
Are all these nl-subshells active? (n/*)
Limitations on population of n-subshells? (y/*)
Highest n-number in reference configuration? (1..3)
Number of electrons in 1s? (0..2)
Number of electrons in 2s? (0..2)
Number of electrons in 2p? (0..6)
>>1
Maximum 2*J-value? (0..)
Minimum 2*J-value? (0..1)
Maximum (2*S+1)-value? (1..9)
Minimum (2*S+1)-value? (1..9)
Maximum resulting angular momentum? (S..N/N=*)
Minimum resulting angular momentum? (S..N/S=*)
Number of excitations = ? (0..3)
114 configuration states have been generated.
Generate a second list? (y/*)
>>
114 configuration states in the final list.
The generated file is called clist.out.
FORTRAN STOP
```

IT IS ADVISABLE TO SAVE THE LSGEN LOG-FILE TO HAVE A RECORD ON

```
HOW THE CONFIGURATION LISTS GENERATION WAS DONE
>>cp clist.log 2PoCAS3BREIT.log
>>cp clist.out 2PoCAS3BREIT.c
COPY WAVE FUNCTION FILES TO PREPARE FOR THE BREIT-PAULI RUN
************************************
>>cp 2PoCAS3.w 2PoCAS3BREIT.w
************************************
       RUN CI BREIT-PAULI
       INPUT FILES: 2PoCAS3BREIT.c, 2PoCAS3BREIT.w
       OUTPUT FILES: 2PoCASBREIT.j
*******************************
>>bpci
Enter ATOM, relativistic (Y/N) with mass correction (Y/N)
>>2PoCAS3BREIT, y, n
Restarting (Y/y) ?
Use existing Matrix and <atom>.1/j initial guess (Y/y)?
Enter Maximum and minimum values of 2*J
>>3,1
Enter eigenvalues: one line per term, eigenvalues separated by commas
2*J = 3
Enter eigenvalues: one line per term, eigenvalues separated by commas
2*J = 1
>>1
                BREIT-PAULI
               _____
Indicate the type of calculation
0 => non-relativistic Hamiltonian only;
1 => one or more relativistic operators only;
2 => non-relativistic operators and selected relativistic:
All relativistic operators ? (Y/N)
>>y
THE TYPE OF CALCULATION IS DEFINED BY THE FOLLOWING PARAMETERS -
    BREIT-PAULI OPERATORS
                              IREL = 2
                             ICSTAS = 1
    PHASE CONVENTION PARAMETER
```

-----

### THE CONFIGURATION SET

-----

STATE (WITH 114 CONGIGURATIONS): THERE ARE 6 ORBITALS AS FOLLOWS: 1s 2s 2p 3s 3p 3d THERE ARE O CLOSED SUBSHELLS COMMON TO ALL CONFIGURATIONS AS FOLLOWS: All Interactions? (Y/N): >>y Default Rydberg constant (y/n) . . . . . . Allocating space for 783 integrals jb = 100Alcmat allocations for idisk= 0 nze = 1124 J = 3 1 Entering LSJMAT with 2J = 3 NUME = LSJMAT with idisk= 0 Nze = 8992 Summary of Davidson Performance Number of Iterations: 9 Shifted Eigval's: -3.95974915438215003E-002 Delta Lambda: 6.10622663543836097E-016 Residuals: 1.09932870175402541E-008 1 Eigenvalues found Finished with Davidson Leaving LSJMAT J = 1 1 Entering LSJMAT with 2J = 1 NUME = LSJMAT with idisk= 0 Nze = 8992 Summary of Davidson Performance \_\_\_\_\_ Number of Iterations: 9 Shifted Eigval's: -3.95976088106196750E-002

Delta Lambda: 3.33066907387546962E-016 Residuals: 1.10031822657079236E-008

1 Eigenvalues found Finished with Davidson Leaving LSJMAT onlydvd F ILS = 0 Finished with the file

\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*

RUN BIOTR TO COMPUTE TRANSITION RATES

INPUT FILES: 2SoCAS3BREIT.c, 2SoCAS3BREIT.w, 2SoCAS3BREIT.j

2PoCAS3BREIT.c, 2PoCAS3BREIT.w, 2PoCAS3BREIT.j OUTPUT FILES: 2SoCAS3BREIT.2PoCASBREIT.1sj

\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*

>>biotr

TRANSBIO 99

\_\_\_\_\_

Name of Initial State

>>2SeCAS3BREIT

Name of Final State

>>2PoCAS3BREIT

intermediate printing (y or n) ?

Relativistic calculation ? (y/n)

Type of transition ? (E1, E2, M1, M2, .. or \*)

>>E1

Pair number 1

J = 0.5Initial CSF : 1s(2).2s\_2S Final  $CSF : 1s(2).2p_2P$ J = 1.5

-7.4738052 -7 405 2\*j = 1 lbl = 1 total energy =2\*j = 3 lbl = 1 total energy =

LENGTH FORMALISM:

-----

FINAL OSCILLATOR STRENGTH (GF) = 1.0308323D+00 TRANSITION PROBABILITY IN EMISSION (Aki) = 3.8973684D+07

### VELOCITY FORMALISM:

-----

SV = 2.1851565D+01FINAL OSCILLATOR STRENGTH (GF) = 9.9944136D-01TRANSITION PROBABILITY IN EMISSION (Aki) = 3.7786856D+07

npair = 2

-----

Pair number 2

2\*j = 1 lbl = 1 total energy = -7.47380522\*j = 1 lbl = 1 total energy = -7.4051945

LENGTH FORMALISM:

-----

SL = 1.1268927D+01FINAL OSCILLATOR STRENGTH (GF) = 5.1540601D-01TRANSITION PROBABILITY IN EMISSION (Aki) = 3.8971507D+07

### VELOCITY FORMALISM:

-----

SV = 1.0925995D+01 FINAL OSCILLATOR STRENGTH (GF) = 4.9972136D-01 TRANSITION PROBABILITY IN EMISSION (Aki) = 3.7785540D+07

Type of transition ? (E1, E2, M1, M2, .. or \*)

STOP END OF CASE