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Anonymous CVPR submission

Paper ID ****

Abstract

The ABSTRACT is to be in fully-justified italicized text, at the top of the left-hand column, below the author and affiliation information. Use the word "Abstract" as the title, in 12-point Times, boldface type, centered relative to the column, initially capitalized. The abstract is to be in 10-point, single-spaced type. Leave two blank lines after the Abstract, then begin the main text. Look at previous CVPR abstracts to get a feel for style and length.

1. Introduction

contributions

- We introduce the idea of augmenting an RGB camera with a single-pixel SPAD to address scale ambiguity error in monocular depth estimators.
- We analyze our approach on indoor scenes using the NYU Depth v2 dataset. We demonstrate that our approach is able to resolve scale ambiguity while being fast and easy to implement.
- We build a hardware prototype and evaluate the efficacy of our approach on real-world data.

2. Related Work

Depth Imaging

- stereo and multiview
- structured illumination and random patterns (kinect, etc.), active stereo
- time of flight (continuous wave and pulsed)
- what we do: like pulsed but much simpler setup; no scanning, no spad array, ...

Monocular Depth Estimation

- summary of architectures and cost functions: u-net type architecture with reverse huber loss
- what we do: same thing, but augment with global hints (inspired by these approaches, we do ...)

Deep Sensor Fusion global hints for super-resolution, colorization, depth estimation

- colorization
- david's 2018 paper for depth estimation and denoising (see david's 2019 sig paper for related work)
- what we do: slightly different application

Histogram Matching Histogram matching as an image processing technique

- Exact histogram matching paper used in this work
- Wasserstein-based optimization techniques for histogram-based regularization

3. Method

In this section, we describe the measurement model for a single-pixel time-of-flight lidar sensor under diffuse, pulsed laser illumination.

3.1. Measurement Model

Consider a laser which emits a pulse at time $t = 0$ with time-varying intensity $g(t)$ uniformly illuminating some 3D scene. We parameterize the geometry of the scene as a height map $z(x, y)$. Neglecting albedo and falloff effects, an ideal detector counting photon events from a location (x, y) in the time interval $(n\Delta t, (n + 1)\Delta t)$ would record

$$\lambda_{x,y}[n] = \int_{n\Delta t}^{(n+1)\Delta t} (f * g)(t - 2z(x, y)/c)dt \quad (1)$$

where c is the speed of light, and f is a function that models the temporal uncertainty in the detector. Single-photon avalanche diodes (SPADs) are highly sensitive photodetectors which are able to record single photon events with high temporal precision [**?**]. Since the detection of each photon can be described with a Bernoulli random variable, the total number of accumulated photons in this time interval follows a Poisson distribution according to

$$h[n] \sim \mathcal{P}\left(\sum_{x,y} \alpha_{x,y}\eta\lambda_{x,y}[n] + b\right) \tag{2}$$

where $\alpha_{x,y} = r_{x,y}/z(x,y)^2$ captures the attenuation of the photon counts due to the reflectance $r(x,y)$ of the scene and due to the inverse square falloff $1/z(x,y)^2$. In addition, η is the detection probability of a photon triggering a SPAD event, and $b = \eta a + d$ is the average number of background detections resulting from ambient photons a and erroneous "dark count" events d resulting from noise within the SPAD.

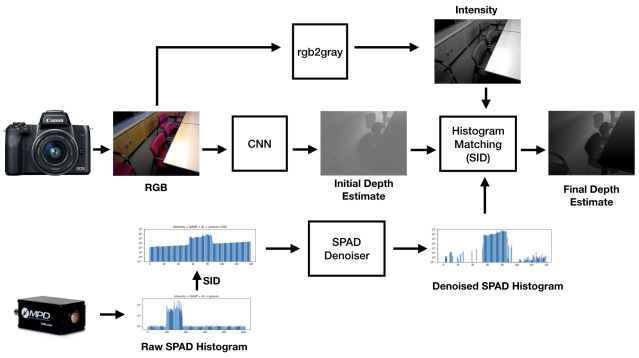

Figure 1: **Overview of the full pipeline** We use a CNN to get an initial per-pixel depth estimate. We then perform gradient descent to optimize that estimate using the SPAD forward model and the dual-Sinkhorn distance .

3.2. Monocular depth estimation with global depth hints

Given a single RGB image $I(x,y)$ and a vector of photon arrivals $h[n]$ described by equation 2, we seek to reconstruct the ground truth depth map $z(x,y)$. Our method has two parts. First, we initialize our estimate of the depth map from the single RGB image via a monocular depth estimator described below. Second, we refine this depth map using the captured measurements $h[n]$ via a process we call Differentiable Histogram Matching (DHM). Differentiable histogram matching is a tool for post-processing the image to match the depth map to the statistics we capture from the SPAD.

Initialization via CNN Convolutional Neural Networks have become increasingly capable of leveraging monocular depth cues to produce accurate estimates of depth from only a single image. We therefore choose to initialize our depth map estimate $\hat{z}^{(0)}(x,y)$ using a CNN. However, any depth estimator reliant on only a single view will be unable to resolve the inherent scale ambiguity in the scene resulting from the tradeoff between size of and distance to an object. The next step, differentiable histogram matching, will resolve this ambiguity using the depth information present in the SPAD histogram.

SPAD Denoising

- Discuss MLE for SPAD denoising
- Write optimization problem for SPAD denoising
- Show performance on a few examples

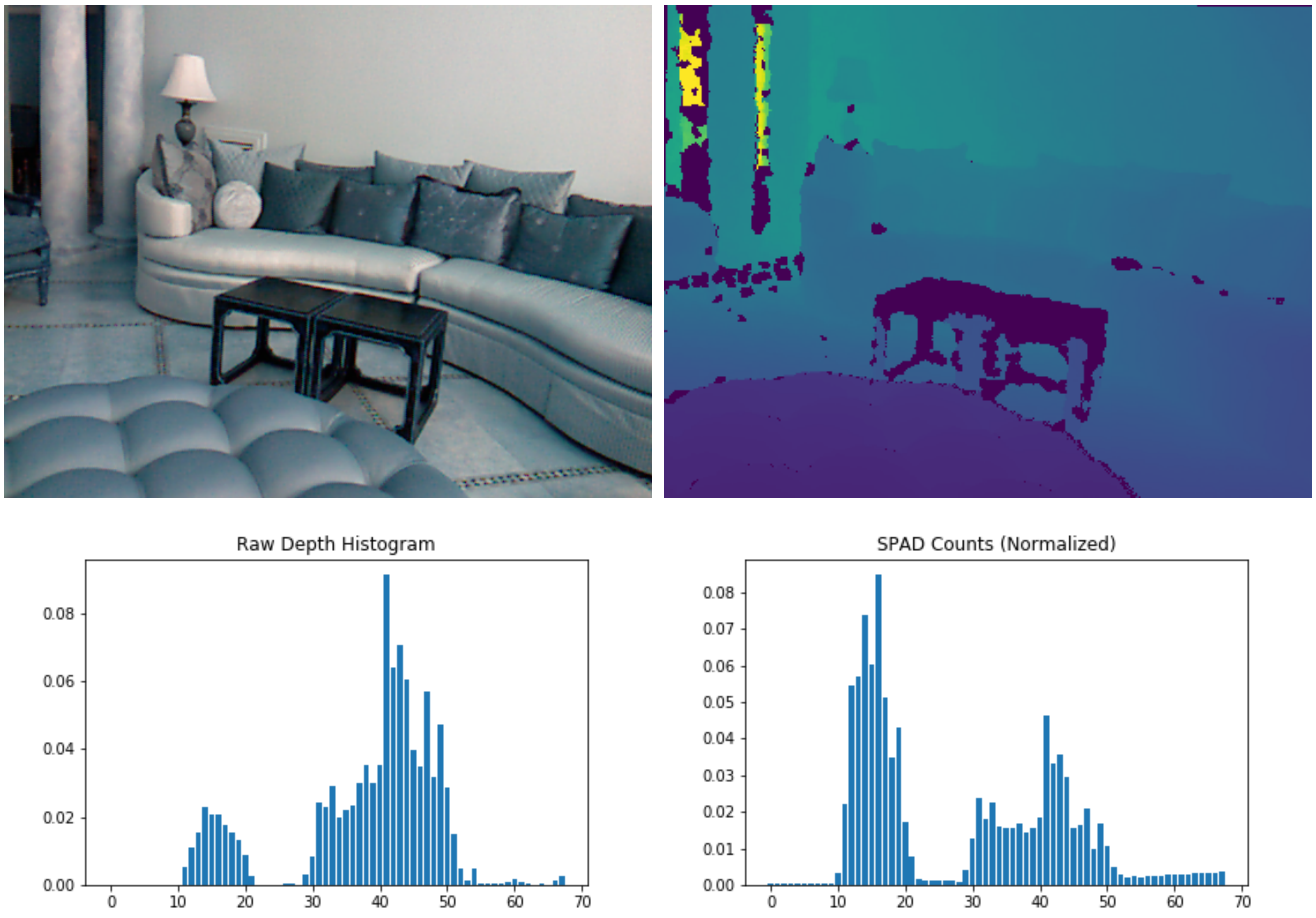

Table 1: Sample Image. Top Left is the RGB image. Top Right is ground truth depth. Bottom Left is Raw ground truth depth histogram. Bottom Right is simulated SPAD measurements. Notice how closer depths are magnified and far depths are attenuated.

Exact Histogram Matching An image's *histogram* is a pair of vectors (h, b) where h_i is the number of pixels of the image whose value lies in the range $[b_i, b_i + 1)$. Then, given a source image S with histogram (h_s, b) and a target histogram (h_t, b), histogram matching generates a new image M such that $h_m \approx h_t$ and the pixel values in M are in the same relative order as in S.

3.3. Implementation Details

For the Monocular Depth Estimator, we use pretrained versions of the the Deep Ordinal Regression Network (DORN) [] and the DenseDepth Network. The exact histogram matching method is as described in [].

4. Simulation

4.1. Implementation Details

- Number of bins used, depth range, laser parameters, use of intensity image.
- Using

NYU Depth v2 The NYU Depth v2 Dataset consists of 249 training and 215 testing scenes of RGB-D data captured using a Microsoft Kinect. We used a version of DORN pretrained according to [**?**] as our CNN.

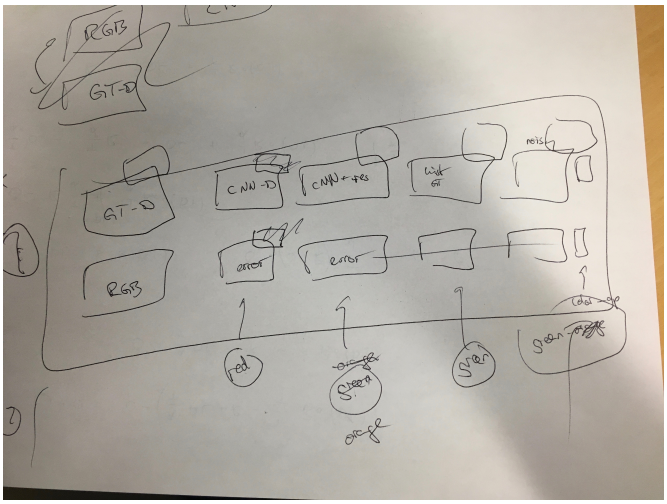

Figure 2: **Comparing our results with other methods**

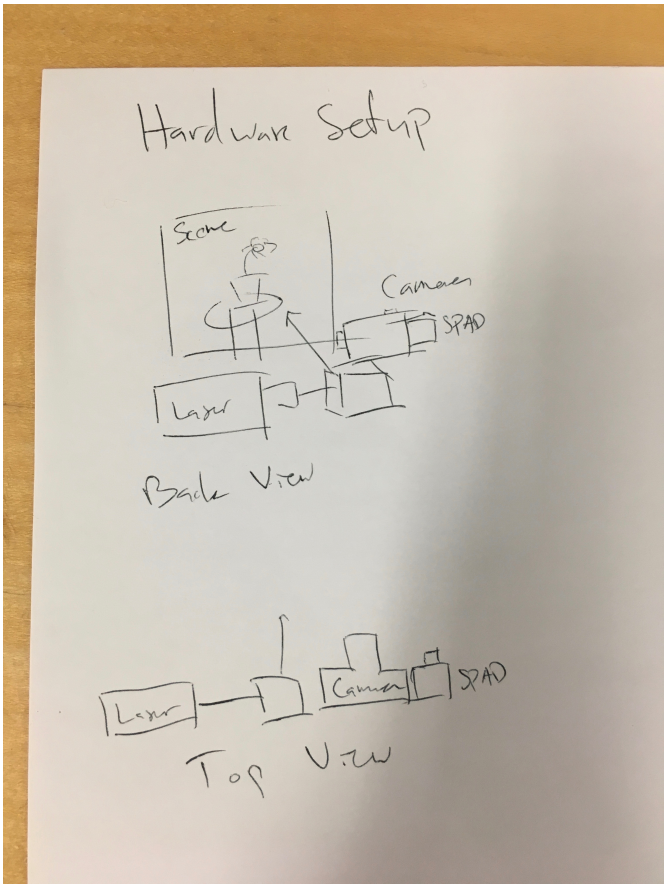

Figure 3: **Hardware setup**

5. Hardware Prototype

5.1. Setup

- Description of hardware used

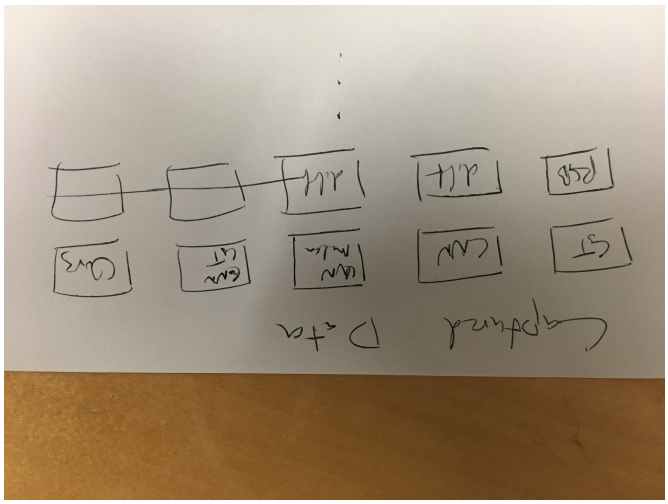

Figure 4: **Hardware results**

- Images of scenes used

6. Discussion

	$\delta^1 \uparrow$	$\delta^2 \uparrow$	$\delta^3 \uparrow$	RMSE $\downarrow$	rel $\downarrow$	$log_{10} \downarrow$
Eigen et. al.	0.769	0.950	0.988	0.641	0.158	-
Laina et. al.	0.811	0.953	0.988	0.573	0.127	0.055
DORN	0.818	0.950	0.982	0.620	0.137	0.063
DORN (rescaled)	0.872	0.967	0.989	0.548	0.111	0.048
Alhashim, Wonka (2019)	0.847	0.973	0.994	0.548(0.461)	0.123	0.053
Alhashim, Wonka (2019) rescaled using GT depth"	0.888	**0.978**	**0.995**	**0.499**(0.409)	**0.106**	**0.045**
Ours (raw depth counts)	**0.899**	0.970	0.990	0.529	0.199	0.055
Ours (DORN) (intensity/falloff)	0.835	0.953	0.984	0.521	0.129	0.060
Ours (DenseDepth) (intensity/falloff)	0.867	0.974	0.994	0.445	0.114	0.050

Table 2: Results on the NYU Depth v2 test set [**?**].

model	hyperparams	delta1	delta2	delta3	rel_abs_diff	rmse	log10
dorn	CNN	0.846	0.954	0.983	0.120	0.501	0.053
	CNN + median rescaling	0.871	0.964	0.988	0.111	0.473	0.048
	CNN + GT histogram matching	0.906	0.972	0.990	0.095	0.419	0.040
	CNN + WHM intensity only (SBR=infinite)	0.904	0.970	0.989	0.091	0.414	0.040
	CNN + WHM (SBR=10)	0.903	0.970	0.989	0.091	0.422	0.040
	CNN + WHM (SBR=50)	0.906	0.971	0.990	0.089	0.410	0.039
	CNN + WHM (SBR=100)	0.906	0.971	0.990	0.090	0.408	0.039
densedepth	CNN	0.847	0.973	0.994	0.123	0.461	0.053
	CNN + median rescaling	0.888	0.978	0.995	0.106	0.409	0.045
	CNN + GT histogram matching	0.930	0.984	0.995	0.079	0.338	0.034
	CNN + WHM intensity only (SBR=infinite)	0.926	0.983	0.995	0.081	0.346	0.035
	CNN + WHM (SBR=10)	0.922	0.982	0.994	0.082	0.361	0.036
	CNN + WHM (SBR=50)	0.925	0.983	0.995	0.081	0.348	0.035
	CNN + WHM (SBR=100)	0.926	0.983	0.995	0.081	0.346	0.035