

Outflows from Young Stellar Objects

MHD, Radiation & Chemistry

Bhargav Vaidya¹

¹ School of Physics and Astronomy, University of Leeds, Leeds.

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Collaborators:

Paola Caselli (Uni. Leeds), Thomas Douglas (Uni. Leeds), Oliver Porth (Uni. Leeds),
Christian Fendt (MPIA), Henrik Beuther (MPIA), Somayeah Sheiknezami (MPIA),
Ciraco Goddi (JIVE), Andrea Mignone (Uni. Torino).

Outline

- 1 Introduction
- 2 Motivation
- 3 Methods : Numerical Simulations
- 4 Outflow Dynamics : Launching
- 5 Outflow dynamics : Propagation
- 6 Summary

Star formation : What do we know

Star formation : Feedback

Talk about winds and outflows

Present challenges

Outflow Evolutionary Picture

Chemistry in outflows

Molecular bullets and EHV emission

Numerical code

Launching and Propagation

Radiation force

Chemistry and Cooling

MHD Acceleration

Collimation and Radiation

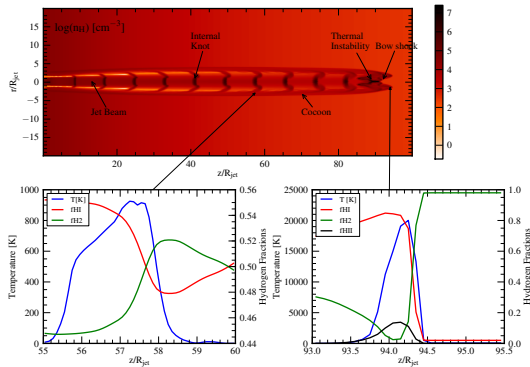
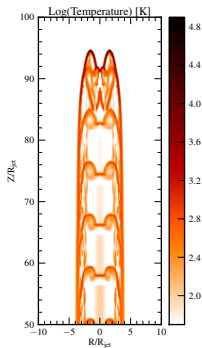
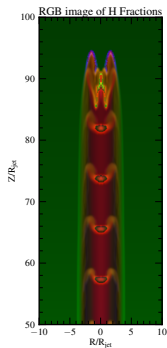
Force Parameters and its impact

Resistive effects

A case of Orion Source I

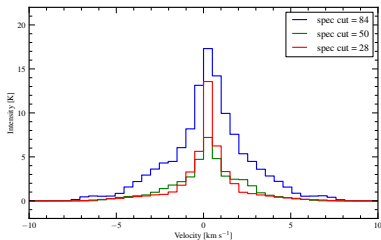
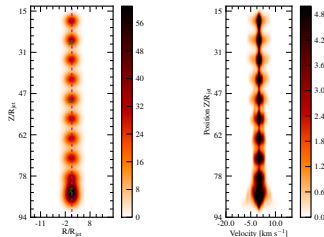
Cooling in Jets

Molecular Interplay

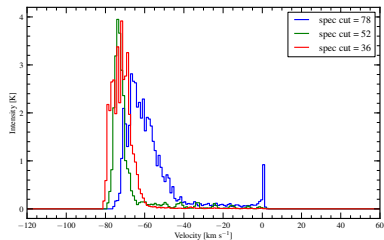
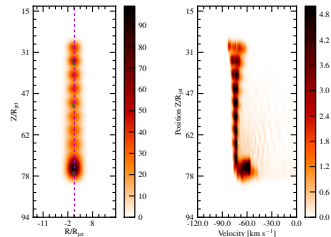


SiO Abundance and Jet Velocity

Spectra and PV diagrams : SiO (2-1)



$\phi = \pi/2$ (Plane of Sky)



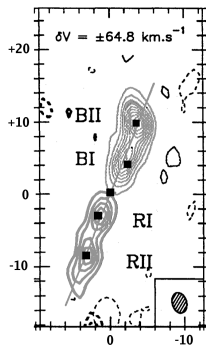
$\phi = \pi/4$

Multi-Line survey : Emission I

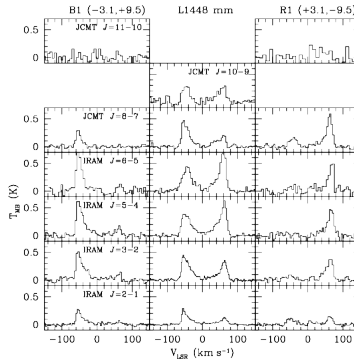
Movies Here.

Multi-Line survey : Emission II

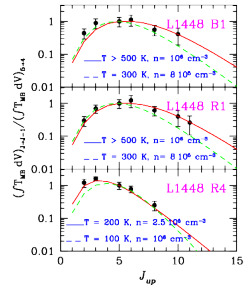
Case of L1448 (Nisini 2007)



Contours SiO 2-1



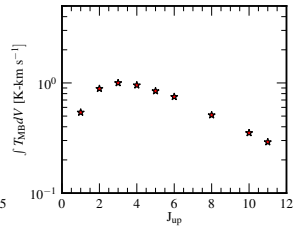
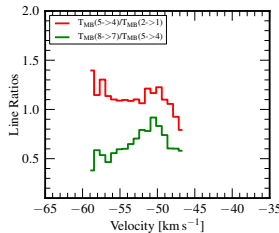
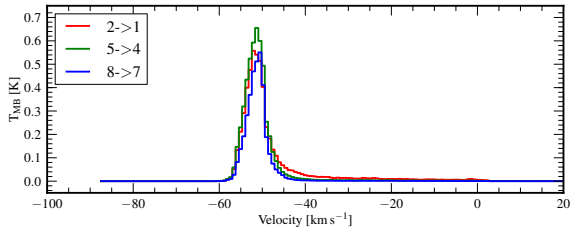
Spectral Features



Kinematic Study (LVG)

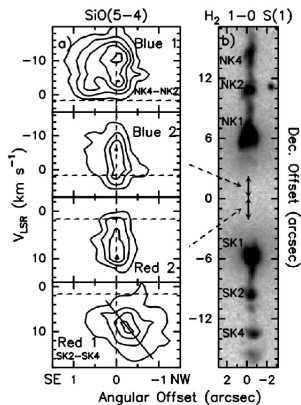
Multi-line survey : Line Ratios

- EHV emission of 0.5 K.
- Line ratios close to Unity.
- Multi-line emission show a distinct fall at high J_{up} .

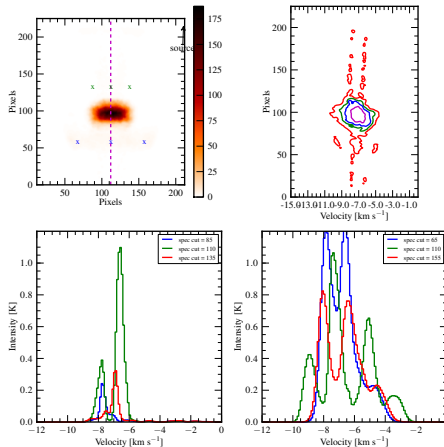


Predictions for ALMA

Focussing on a single knot



HH 212 (Codella 2007)



Rotation or Wiggles?

Conclusions