Homework 12

due May 10th

1. For what constant κ do the Euler-Lagrange equations of the following action:

$$S[g_{\mu\nu}] = \int d^4x \sqrt{-g} \left(\kappa R + \mathcal{L}_M\right)$$

using the relation

$$-\frac{2}{\sqrt{-g}}\frac{\delta\left(\sqrt{-g}\mathcal{L}_{M}\right)}{\delta g^{\mu\nu}}=T_{\mu\nu}$$

yield the canonical Einstein Field Equations.

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = \frac{8\pi G}{c^4}T_{\mu\nu}$$

2. Show via the Nöther procedure that spacetime translation invariance implies the conservation of a rank two tensor and express this tensor in terms of the Lagrangian for a scalar field:

$$S[\phi] = \int d^4x \mathcal{L}(\phi(x), \partial \phi(x)) \tag{1}$$

Is the resulting tensor necessarily symmetric? Any guesses as to why?¹

- 3. Complete the following two calculations that were skipped in class:
 - (a) Calculate G_{tt} for the static weakfield metric

$$ds^2 = -(1+2\phi)dt^2 + (1-2\phi)(dx^2 + dy^2 + dz^2)$$

(b) Calculate the $G_{\hat{t}\hat{t}}$, $G_{\hat{r}\hat{r}}$, $G_{\hat{\phi}\hat{\phi}}$, and $G_{\hat{\theta}\hat{\theta}}$ for the FRW metric, with spatial curvature k

$$ds^{2} = -dt^{2} + a(t)^{2} \left(\frac{dr^{2}}{1 - kr^{2}} + r^{2} (d\theta^{2} + \sin^{2} d\phi^{2}) \right)$$

(c) Calculate $G_{\hat{t}\hat{t}}$ and $G_{\hat{r}\hat{r}}$ for the ansatz we made in calculating the Schwarzchild metric

$$ds^{2} = -e^{\nu(r)}dt^{2} + e^{\lambda(r)}dr^{2} + r^{2}d\Omega$$

¹This problem is slightly outside the scope of the course, just do your best. I fear that the lack of dynamics requirement at NYU is handicapping students interested in graduate school. For a complete solution to this problem see chapter 2 of Peskin and Schröder's book on quantum field theory.