

# Errata: “Spin Parity of Spiral Galaxies I – Corroborative Evidence for Trailing Spirals”(2019, ApJ, 886, 133)

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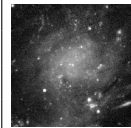
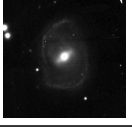
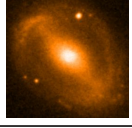





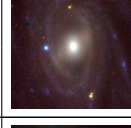




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In Figure 8, we correct our judgment of the near side of NGC772 to “?” due to possible intervening features.



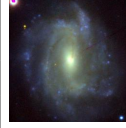

(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)	(9)	(10)
ID	S/Z	Dark Side	Appr. Side	T/L	Reference	Image 1	Image 2	Image 3	Image 4
NGC772	S	?	NW	?	Rhee (1996)				
NGC801	S	SW	SE	T	Rubin (1980)				
NGC871	Z	W	N	T	Lee-Waddell (2014)				
NGC877	Z	SW	NW	T	Lee-Waddell (2014)				

In Figure 9, the approaching side of NGC1 is NW. NGC1614 and NGC2207 show intervening features and we correct our identification of their near sides to “?”. The approaching side of NGC4536 is SE (Kuno 2007).

(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)	(9)
ID	S/Z	Dark Side	Appr. Side	Reference	Image 1	Image 2	Image 3	Image 4
NGC1	Z	sw	NW	Falcon-Barroso (2017)				
NGC23	S	se	NE	Falcon-Barroso (2017)				
NGC36	S	ne	NW	Falcon-Barroso (2017)				
NGC45	Z	sw	NW	Chemini. (2006)				
NGC1358	S	ne	NW	Schnorr-Muller (2017)				
NGC1512	Z	nw	NE	Buta (1988)				
NGC1614	S	?	?	Garcia-Burillo (2018)				
NGC1645	S	s	E	Garcia-Lorenzo (2015)				
NGC2139	S	ne	NW	Andersen (2008)				
NGC2207	Z	?	SE	Rubin (1983)				
NGC2276	S	s	E	Gruendl (1993)				
NGC2336	S	w	S	Wilke (1999)				
NGC4536	Z	ne	SE	Kuno (2007)				
NGC4540	S	w	S	Neumayer (2011)				

NGC4548	Z	sw	NW	Vollmer (1999)				
NGC4618	Z	sw	NW	Bush (2004)				

12 In Figure 10, The dark side of NGC3887 is W instead of E.

(1)	(2)	(3)	(4)	(5)	(6)	(7)	(8)
ID	S/Z	Dark Side	Appr. Side	Image 1	Image 2	Image 3	Image 4
NGC3813	S	N	w				
NGC3877	Z	SE	sw				
NGC3887	Z	<b>W</b>	<b>n</b>				
NGC3917	S	SE	ne				

13 We retract the near-side identification of three possible non-isolated spirals and correct data of three other spirals.  
 14 These corrections change the numbers 146 and 252 in the 1st and 5th entries of column (4) of Table 1 to 145 and 253,  
 15 respectively. Finally, the number 146 in the abstract should also be 145.

16 The authors thank Cora Schallock and Wayne B. Hayes for cross-checking our catalog and reporting some discrepant  
 17 results. The current errata reflects our reexamination of all the samples.

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