## Modular Response in Free Quantum Fields: A KMS/FDT Theorem and Conditional Extensions

[clg]<sup>1</sup>

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(Dated:)

Part I (Theoremic core, free/Gaussian Hadamard QFT). We prove that, for small causal diamonds (CHM) in locally Hadamard states and within a safe window  $\epsilon_{\rm UV} \ll \ell \ll \min\{L_{\rm curv}, \lambda_{\rm mfp}, m_i^{-1}\}$ , the MI/moment-kill projector isolates a finite  $\ell^4$  modular response with coefficient equal to its flat-space value; the projected KMS/FDT susceptibility is positive; and coarse-graining over the wedge family produces the universal weak-field prefactor  $5/12 = (4/3) \times (5/16)$ . The fractional KMS defect between CHM diamonds and half-spaces scales as  $\mathcal{O}((\ell/L_{\rm curv})^2) + \mathcal{O}((\ell H)^2)$ . The QFT sensitivity is  $\beta = 2\pi C_T I_{00} = 0.02086 \pm 0.00105$  (conservative 5% shared systematics). A scheme-invariant background relation suggests  $\Omega_{\Lambda} = \beta f c_{\rm geo}$  conditional on our coarse-graining and analyticity assumptions.

Part II (Conditional extensions). We separate definition (flat-space  $\varepsilon$  from modular response) from mapping. Rather than impose the standard EFT-of-DE  $\alpha$ -basis, we adopt a quasi-static closure that keeps operational distances GR-like (no additional lensing coupling  $\Sigma \simeq 1$ ) while modifying growth via  $\mu(\varepsilon,s) = 1/(1+\frac{5}{12}\varepsilon\,s(x))$  with s(x) a local, covariant environment modulation derived from the action. KMS/FDT positivity motivates an entropy-driven law  $d\varepsilon/d\ln a \geq 0$  with a conditional background budget  $\int \varepsilon\,d\ln a = \Omega_{\Lambda}$ . We then prove a No-Go Lemma showing any linear kernel fails to yield Tully–Fisher scaling and introduce a quasistatic elastic sector (AQUAL form, derived as the SK static limit) with the same fixed acceleration  $a_0 = \frac{5}{12}\Omega_{\Lambda}^2cH_0$ . The elastic sector is convex, well-posed, obeys Solar–System curvature gating, produces  $\Phi = \Psi$  at working order (no slip), and leaves linear cosmology intact via a causal SK filter.

Part III (Exploratory). (i) An optional, shock-selective optical channel (D') reduces  $\Sigma$  only in high-shear shocked gas to address Bullet-type lensing offsets while preserving FRW distances, with a principled SK/BRSSS derivation path connecting the amplitude to ICM transport coefficients. (ii) A compact thermodynamic interpretation of the projected modular response: a Clausius-like identity holds at working order in the MI/moment-kill channel, and the FRW budget may be viewed as a coarse-grained Clausius normalization conditional on our KMS $\rightarrow$ FRW hypotheses. (iii) Linkage to global entropic gravity (Bianconi 2025) via small-diamond MI matching and observational discriminants.

Referee-guided additions. We (A) formalize an MI-smeared null-energy positivity (free fields; projected QEI), (B) make explicit the RG/operator-spectrum bridge pinning  $\beta$  to  $C_T$  and clarifying anomaly channels, (C) expose SM bookkeeping tied to the safe-window fraction  $f_V$  and curvature gating  $s(\chi_g)$  so that GR dominance is recovered wherever heavy sectors or strong curvature suppress the MI channel, and (D) add the No-Go Lemma + elastic SK/AQUAL sector with fixed  $a_0$ , no slip, Solar–System compliance, and cosmology/galaxy separation.

## READER'S MAP: PART I (THEOREM) VS. PART II (CONDITIONAL) VS. PART III (EXPLORATORY)

Part I (Secs. I–V, IV, XIA–B; Apps. XXI–XXV): proven results for free/Gaussian Hadamard fields at working order, SM bookkeeping, and first-principles positivity/RG clarifications.

Part II (Secs. VI–VIII, IX, XXXII, XIC; Apps. XXVI–XXVII, XXVIII): conditional mapping for growth  $(\mu(\varepsilon, s))$ , Linear No-Go (Sec. VII), and Elastic quasistatic sector (Sec. VIII) with Solar–System gating, no slip, and BTFR; causal SK separation of regimes.

Part III (Secs. XIII, XX, XVII; Apps. XXX, XXXI): exploratory shock-selective optics (D'); thermodynamic interpretation; and linkage to Bianconi's global entropic gravity.

#### I. SCOPE, WORKING ORDER, AND SAFE-WINDOW QUANTIFICATION (PART I)

- a. Working order and state class. We work to  $\mathcal{O}(\ell^4)$  in the MI/moment-kill projector channel, treating curvature/contact terms as  $\mathcal{O}(\ell^6)$ . States are locally Hadamard.
- b. KMS applicability (CHM diamonds). Exact BW KMS holds for half-spaces; CHM diamonds inherit it with fractional defect  $\mathcal{O}((\ell/L_{\text{curv}})^2) + \mathcal{O}((\ell H)^2)$  (App. XXV).

c. Safe-window volume fraction. Define a conservative admissible scale

$$\ell_{\text{max}}(x) \equiv \zeta \min \left\{ L_{\text{curv}}(x), \ \lambda_{\text{mfp}}(x), \ m_i^{-1}(x) \right\}, \qquad \zeta = 0.1.$$
 (1)

Using Press–Schechter/Sheth–Tormen mass functions and NFW curvature proxies  $L_{\rm curv}^{-2} \sim (R_{abcd}R^{abcd})^{1/2}$  with substructure excision parameter  $\xi$ , we estimate the comoving volume fraction  $f_V(\ell_{\rm min}) = {\rm Vol}\{x: \ell_{\rm max}(x) > \ell_{\rm min}\}/{\rm Vol}_{\rm tot}$ . A semi-analytic survey (App. XXVI) shows voids dominate  $f_V$ , while dense cores lack a window; representative values at  $z \sim 0$  for  $\ell_{\rm min} \in [1,100]$  pc are  $f_V \sim 0.6-0.95$  for  $\xi \in [0.2,0.5]$ . This enters only as a domain-of-validity indicator.

- d. Spectrum caveat. The admissible window  $\epsilon_{\rm UV} \ll \ell \ll \min\{L_{\rm curv}, \lambda_{\rm mfp}, m_i^{-1}\}$  is understood to apply to sectors that contribute at working order. Massive sectors with  $\ell \gg m_i^{-1}$  are exponentially suppressed and, after MI/moment–kill subtraction, do not re-introduce lower moments or  $\ell^4 \log \ell$  terms. Thus the  $\ell^4$  coefficient is dominated by massless/light fields while heavy fields decouple in this channel. See Sec. IV for SM bookkeeping that packages light-field multiplicity into a single  $\varepsilon_{\rm SM}$ .
- e. Angle invariance as a null test. The continuous-angle product  $C_{\Omega} = f(\theta) c_{\text{geo}}(\theta)$  is analytic and  $\theta$ -independent; residuals are shown as a null check, not a precision claim.

## II. A2-KMS THEOREM (GAUSSIAN/HADAMARD SECTOR)

**Theorem 1** (Projected modular response and positivity). Let Q be a free (Gaussian) QFT on a globally hyperbolic spacetime and  $\rho$  a locally Hadamard state. For a causal diamond of radius  $\ell$  with  $\ell \ll L_{\rm curv}$  and the MI/moment-kill projector that cancels  $r^0$  and  $r^2$  moments, the MI-subtracted modular response obeys

$$\delta \langle K_{\text{sub}} \rangle = (2\pi C_T I_{00}) \,\ell^4 \,\delta \varepsilon + \mathcal{O}(\ell^6), \tag{2}$$

with coefficient equal to the flat-space value. The retarded susceptibility  $\chi_{QK}$  in the projected channel is positive (FDT), and wedge averaging yields the universal weak-field prefactor 5/12. The fractional deviation from BW KMS is  $\mathcal{O}((\ell/L_{curv})^2) + \mathcal{O}((\ell H)^2)$ .

Corollary 1 (Conditional background statement). Under the coarse-graining and analyticity assumptions of Sec. X, the FRW zero mode suggests the scheme-invariant relation  $\Omega_{\Lambda} = \beta f c_{\text{geo}}$  with  $\beta = 2\pi C_T I_{00}$ . We treat this as a conditional statement rather than a theorem.

## III. QFT INPUT: $\beta = 2\pi C_T I_{00}$ AND ERROR BUDGET

We evaluate  $\beta$  via four independent routes: (a) real-space CHM; (b) spectral/Bessel; (c) Euclidean time-slicing; (d) replica finite-difference. The spread is  $\lesssim 1\%$ . We adopt a conservative

$$\beta = 0.02086 \pm 0.00105$$
 (5% shared systematics). (3)

Angle invariance is used as a null residual test.

Here  $C_T$  denotes the flat-space stress-tensor two-point normalization, e.g.  $\langle T_{ab}(x) T_{cd}(0) \rangle = C_T \mathcal{I}_{abcd}(x)/|x|^{2d}$  in d dimensions (see Osborn–Petkou).

Benchmark (convention). For a free, massless real scalar in d=4 and our normalization,  $C_T=1/(120\pi^2)$ , which yields  $\beta \simeq 0.02086$  via Eq. (4).

Implementation consistency (note). The normative constants used for the numerical reproductions are

$$C_T = \frac{1}{120\pi^2}, \qquad (\sigma_1, \sigma_2) = \left(\frac{1}{2}, 2\right), \qquad (a, b) = \left(\frac{4}{5}, \frac{1}{5}\right),$$

with the moment-kill identities enforced exactly (App. XXI). Helper scripts (beta\_methods\_v2.py, referee\_pipeline.py) print these values alongside the computed  $I_{00}$  to prevent normalization drift.<sup>1</sup>

Reproducibility (non-circular). We use a two-scale MI/moment-kill subtraction with a top-hat window on 3-balls

$$W_\ell(r) = \frac{3}{4\pi\ell^3}\,\Theta(\ell-r), \qquad \mathcal{W}_\ell := \int_{B_\ell} W_\ell - \ a \int_{B_{\sigma_1\ell}} W_{\sigma_1\ell} - \ b \int_{B_{\sigma_2\ell}} W_{\sigma_2\ell}.$$

<sup>&</sup>lt;sup>1</sup> In earlier development branches some convenience flags defaulted to alternate normalizations (e.g.  $C_T = 3/\pi^4$ ) and near-unity MI scales. These have been disabled in the archival runners; the paper's conventions are authoritative.

The two moment-kill conditions (cancelling  $r^0$  and  $r^2$  for any smooth radial F) fix

$$a+b=1, \qquad a\,\sigma_1^2+b\,\sigma_2^2=1 \implies a=rac{\sigma_2^2-1}{\sigma_2^2-\sigma_1^2}, \quad b=rac{1-\sigma_1^2}{\sigma_2^2-\sigma_1^2}.$$

In our runs we take

$$(\sigma_1, \sigma_2) = \left(\frac{1}{2}, 2\right), \qquad (a, b) = \left(\frac{4}{5}, \frac{1}{5}\right) = (0.8, 0.2).$$

With these weights the projected  $\ell^4$  coefficient evaluates to

$$I_{00} = 3.932017$$
 (dimensionless),

so with  $C_T = 1/(120\pi^2)$  one obtains  $\beta = 2\pi C_T I_{00} = 0.02086$  as quoted. The helper script beta\_methods\_v2.py echoes both  $(a, b; \sigma_1, \sigma_2)$  and the numeric  $I_{00}$ .

#### IV. STANDARD-MODEL SECTOR: BOOKKEEPING AND DECOUPLING AT WORKING ORDER

a. What is being linked. At working order the MI/moment-kill channel defines the dimensionless state variable  $\varepsilon(x)$  through

$$\delta \langle K_{\text{sub}} \rangle = \beta \, \ell^4 \, \delta \varepsilon + \mathcal{O}(\ell^6) \quad [\text{Eq. (4)}].$$

This subsection clarifies how the Standard-Model (SM) content enters  $\varepsilon$  and why heavy states decouple.

b. Species sum and decoupling. Write  $\varepsilon$  as a weighted sum over species i:

$$\varepsilon(x) = \sum_{i} w_i \, \varepsilon_i(x), \qquad w_i \text{ counts effective dof (helicity/polarization, internal factors)}.$$

In a diamond of size  $\ell$ , fields with  $m_i \ell \gg 1$  are exponentially suppressed in the projected channel; after the MI/moment-kill subtraction they do not re-introduce lower moments nor  $\ell^4 \log \ell$  terms. Parametrically,

$$\varepsilon_i(x) \propto e^{-m_i \ell}$$
 for  $m_i \ell \gg 1$ ,

so the  $\ell^4$  coefficient is dominated by massless/light fields while heavy fields decouple.

c. Packaging the light SM content. It is convenient to define a single light-sector variable

$$\varepsilon_{\rm SM}(x) \equiv \sum_{i \in \text{light}} c_i \, \varepsilon_i(x),$$

where  $c_i$  packages the relevant multiplicities (helicity/polarization, internal quantum numbers) of each light SM species under the MI projection. All subsequent working-order formulas may then be read with  $\varepsilon \to \varepsilon_{\rm SM}$  when SM content is explicitly considered.

d. Coupling to gravity at working order. The only background scalar that survives the MI/moment–kill projection and modifies weak-field growth while keeping distances GR-like is the Planck-mass renormalization  $\delta \ln M^2 = \beta \, \delta \varepsilon$  (Assumption D). Multiplicities therefore simply rescale  $\varepsilon$  (hence  $\mu$ ); they do not change  $\beta$  or the universal weak-field bookkeeping that fixes the 5/12 prefactor:

$$\mu(\varepsilon,s) = \frac{1}{1 + \frac{5}{12} \varepsilon s(x)} \longrightarrow \mu(\varepsilon_{\text{SM}},s) = \frac{1}{1 + \frac{5}{12} \varepsilon_{\text{SM}} s(x)}.$$

- e. Environment and distances. The environment scalar s(x) is geometric (built from curvature invariants) and independent of particle content at this order; FRW distances remain GR-like ( $\Sigma \simeq 1, c_T = 1$ ). The observed lensing amplitude changes only indirectly through altered growth.
- f. Practical note and  $f_V$  linkage. In cosmological applications one sets a light-sector threshold  $m_i \ell \lesssim 1$  (with  $\ell$  within the safe window) and computes  $\varepsilon_{\rm SM}$  using the appropriate  $c_i$ . As the environment varies, field regimes can re-enter  $\varepsilon_{\rm SM}$  smoothly. Dense regions with no safe window contribute negligibly to the MI channel; voids dominate the valid domain via  $f_V$  (App. XXVI). GR dominance is ensured either by strong curvature  $(s(\chi_g) \to 0)$  or by heavy-sector decoupling  $(m_i \ell \gg 1)$ .

#### V. WEAK-FIELD PREFACTOR 5/12

The isotropic BW channel gives  $\langle T_{kk} \rangle = (1+w)\rho$  with UV  $w=1/3 \Rightarrow 4/3$ . Averaging over CHM segments yields 5/16, so  $5/12 = (4/3) \times (5/16)$ . Details in Sec. V.

## VI. DEFINITION VS. MAPPING (PART II; CONDITIONAL)

a. Definition (flat-space QFT).

$$\delta \langle K_{\text{sub}}(\ell) \rangle = \underbrace{(2\pi C_T I_{00})}_{\beta} \ell^4 \delta \varepsilon(x) + \mathcal{O}(\ell^6). \tag{4}$$

b. Mapping (constitutive; beyond the  $\alpha$ -basis). We do not impose the linear EFT-of-DE  $\alpha$ -parameter mapping at working order. Instead, we adopt a quasi-static closure that keeps operational distances GR-like while modifying growth:

$$\nabla^2 \Phi = 4\pi G a^2 \rho_m \,\mu(\varepsilon, s), \qquad \mu(\varepsilon, s) = \frac{1}{1 + \frac{5}{12}\varepsilon \,s(x)}, \tag{5a}$$

$$\nabla^2 \frac{\Phi + \Psi}{2} = 4\pi G a^2 \rho_m, \qquad (\Sigma \simeq 1 \text{ on FRW and in laminar flows}). \tag{5b}$$

Here s(x) is a local scalar built from curvature (Sec. XV); in FRW, Weyl =  $0 \Rightarrow \chi_g = 0 \Rightarrow s = 1$ . Beyond working order we make no stability claims absent an action;  $\mu(\varepsilon, s)$  serves as a falsifiable diagnostic with  $\Sigma \simeq 1$ . Matter obeys the standard continuity and Euler equations. This closure preserves the Bianchi identity at working order because s(x) is a scalar; an action-level realization and frame-independence are given below (Remark VI A). Optional Assumption D' (Sec. XIII) introduces a shock-selective lensing modification  $\Sigma(x) < 1$  localized to high-shear gas while keeping FRW  $\Sigma \simeq 1$ .

Remark on lensing amplitude.  $\Sigma \simeq 1$  denotes no additional lensing coupling in the baseline; the observed lensing signal still changes through the altered growth D(a). Under Assumption D',  $\Sigma$  may be reduced *locally* in shocked gas  $(S_{\text{shock}} \gg 1)$  without affecting FRW.

c. EFT stub (derivation of 5/12). At quasi-static, sub-horizon scales, a background variation  $\delta \ln M^2 = \beta \, \delta \varepsilon$  rescales the Poisson coupling as  $G \to G_{\text{eff}} = G/(1+\Delta)$  with  $\Delta$  fixed by the universal weak-field bookkeeping. In the isotropic BW channel the contraction 4/3 and the segment ratio 5/16 (Sec. V) give  $\Delta = \frac{5}{12}\varepsilon$ , hence

$$\mu(\varepsilon, s) = \frac{G_{\text{eff}}}{G} = \frac{1}{1 + \frac{5}{12}\varepsilon s(x)},\tag{6}$$

consistent with Eqs. (5).

d. Trial action (outlook). A possible action-level route consistent with our closure is to consider an effective term that modulates  $M^2$  via the modular response,

$$S_{\rm trial} = \int d^4x \sqrt{-g} \left[ \frac{M^2}{2} R + \lambda \left( \delta \ln M^2 \right) \mathcal{K}[g;\ell] + \cdots \right],$$

where  $\mathcal{K}$  is a local covariant scalar capturing the projected channel at working order and  $\lambda$  a running coefficient. While only illustrative, this shows how  $\delta \ln M^2 = \beta \, \delta \varepsilon$  could arise from an action (cf. [6, 8]).

## A. Frame-independence of throttling (remark)

Throttling here means the reduction of the effective gravitational coupling relative to GR caused by the background state variable  $\varepsilon(a)$  and a local environment factor s(x) that encodes curvature/inhomogeneity. In the Jordan frame we take

$$M_*^2(x,a) = M^2 \left[ 1 + \frac{5}{12} \, \varepsilon(a) \, s(x) \right], \qquad s(x) = \frac{1}{1 + (\chi_q/\chi_\star)^q} + \mathcal{O}\left(\frac{R}{m_s^2}\right),$$

so the quasi-static Poisson law reads

$$\nabla^2 \Phi \simeq \frac{4\pi G a^2 \rho_m \, \delta}{1 + \frac{5}{12} \, \varepsilon(a) \, s(x)} \quad \Rightarrow \quad G_{\text{eff}}(x, a) = \frac{G}{1 + \frac{5}{12} \, \varepsilon(a) \, s(x)}.$$

Thus throttling is present everywhere, while its magnitude is amplitude—modulated by the local invariant  $\chi_g = \ell^2 \sqrt{C_{abcd} C^{abcd}}$ : in weak fields  $(\chi_g \ll \chi_{\star})$  one has  $s \to 1$  and the full background rescaling  $G_{\text{eff}} = G/(1 + \frac{5}{12}\varepsilon)$ ; in strong fields  $(\chi_g \gg \chi_{\star})$  one has  $s \to 0$  and  $G_{\text{eff}} \to G$  (Solar–System compliance).

A conformal map to the Einstein frame,

$$\tilde{g}_{\mu\nu} = \Omega^2 g_{\mu\nu}, \qquad \Omega^2 = 1 + \frac{5}{12} \varepsilon(a) s(x),$$

renders  $M_*$  constant and shifts the same throttling into the matter coupling. To working order in our MI/moment-kill channel, gradients of  $\Omega$  and of  $\chi_g$  enter only at  $\mathcal{O}((\ell/L_{\text{curv}})^2)$  and  $\mathcal{O}(R/m_s^2)$ , consistent with the error budget in Eq. (9) and App. XXV; the observables of interest are frame—independent at this order: growth is governed by

$$\mu(\varepsilon, s) = \frac{1}{1 + \frac{5}{12} \varepsilon(a) s(x)},$$

and distances remain GR-like ( $\Sigma \simeq 1, c_T = 1$ ).<sup>2</sup>

Scale-separation note. The local modular response enters gravity solely as a renormalization  $\delta \ln M_*^2 = \beta \, \delta \varepsilon$  of the Planck mass; the Einstein equations then propagate this renormalization to cosmological scales through the standard gravitational coupling. No macroscopic quantum coherence or ad hoc coarse-graining is required, and the Jordan $\leftrightarrow$ Einstein map above makes this statement frame-independent at working order.

A simple way to realize s(x) is as an auxiliary heavy scalar that minimizes a local potential

$$\mathcal{V}(s;\chi_g) = \frac{M^2 m_s^2}{2} \left[ s - \frac{1}{1 + (\chi_g/\chi_{\star})^q} \right]^2,$$

so that the algebraic EOM enforces  $s = [1 + (\chi_g/\chi_\star)^q]^{-1} + \mathcal{O}(R/m_s^2)$ . Choosing  $m_s^2 \gg H_0^2$  ensures adiabatic tracking. Constraints (working order). (i) Choose  $m_s^2 \gg H_0^2$  so s(x) adiabatically tracks  $[1 + (\chi_g/\chi_\star)^q]^{-1}$  and the  $\mathcal{O}(R/m_s^2)$  offset is negligible. (ii) The Planck-mass drift  $\alpha_M = d \ln M_*^2/d \ln a = \frac{(5/12) s \, d\varepsilon/d \ln a}{1+(5/12)\varepsilon s}$  is naturally small under our monotone  $\varepsilon(a)$ . (iii) In FRW, Weyl = 0 so curvature-weighted corrections vanish; in LSS they are  $\mathcal{O}((\ell/L_{\rm curv})^2)$ . Weak-field acceleration (toy/conditional; clarification). Because  $s \to 1$  in low curvature, the weak-field normalization implies a MOND-like scale

$$a_0 = \frac{5}{12} \,\Omega_{\Lambda}^2 \, c \, H_0, \tag{7}$$

Using the baseline  $\Omega_{\Lambda} = 0.685$  and  $H_0 = 70.9 \text{ km s}^{-1} \text{ Mpc}^{-1}$ , this gives  $a_0^{\text{eff}} \approx 1.2 \times 10^{-10} \text{ m s}^{-2}$  in the weak-field limit  $(s \simeq 1)$ ; and the effective  $a_0^{\text{eff}}$  is enhanced in weak-field regimes by the derived  $s \to 1$  (not imposed), while Solar–System compliance follows from  $s(\chi_{\odot}) \ll 1$  (Sec. XV). Pipeline values propagate the  $\pm 5\%$  uncertainty in  $\beta$ .

### VII. LINEAR-KERNEL NO-GO FOR TULLY-FISHER

**Lemma 1** (Linear-kernel no-go). Let  $\Phi$  be determined from  $\rho$  by a translation/rotation-invariant linear map with tempered kernel G:

$$\Phi(\mathbf{x}) = (G * \rho)(\mathbf{x}), \qquad G(\mathbf{x}) = \frac{1}{(2\pi)^3} \int \frac{\tilde{g}(\mathbf{k})}{k^2} e^{i\mathbf{k}\cdot\mathbf{x}} d^3k, \quad \tilde{g}(\mathbf{k}) > 0,$$

and assume that outside a bounded mass distribution the stationary field equation reduces to a linear constant-coefficient elliptic operator. Then the exterior field decays as  $\Phi \sim -\mathcal{G}_{\text{eff}}M/r$ , so  $g = |\nabla \Phi| = \mathcal{G}_{\text{eff}}GM/r^2 + o(r^{-2})$  and  $v_{\infty}^4 \propto M^2$ . Thus a linear kernel cannot yield the Tully–Fisher scaling  $v_{\infty}^4 \propto M$ .

Sketch. In vacuum the solution is harmonic (up to renormalization  $\mathcal{G}_{\text{eff}}$ ); Liouville/Rellich imply the only decaying solution is 1/r, hence  $g \propto r^{-2}$ . Any linear, isotropic nonlocality rescales  $\mathcal{G}_{\text{eff}}$  but preserves the  $1/r^2$  falloff.

<sup>&</sup>lt;sup>2</sup> This remark complements Assumption D (Sec. XIIB): the working-order modification resides in a state- and environment-dependent  $M_*^2$  with no additional lensing coupling. A failure would manifest as our falsifiers in Sec. XIX, e.g. a significant GW/EM distance split or a persistent  $\ell^4 \log \ell$  term.

### VIII. ELASTIC QUASISTATIC SECTOR (AQUAL FROM SK STATIC LIMIT)

a. Static elastic action and field equation. In the quasistatic limit of the Schwinger-Keldysh (in-in) effective action, a causal retarded kernel  $\mathcal{K}^{el}$  reduces to a local, convex functional for the Newtonian potential  $\Phi$ :

$$\mathcal{L}_{\rm el}(\Phi, \nabla \Phi) = \frac{a_0^2}{8\pi G} F(Y) - \rho \Phi, \quad Y \equiv \frac{|\nabla \Phi|^2}{a_0^2}$$

with  $F \in C^2([0,\infty))$ ,  $F'(Y) = \mu(Y) > 0$ ,  $F''(Y) \ge 0$ , and asymptotics

$$\mu(Y) \xrightarrow{Y \gg 1} 1, \qquad \mu(Y) \xrightarrow{Y \ll 1} \sqrt{Y}.$$

Variation yields the AQUAL field equation [22]:

$$\nabla \cdot (\mu(Y) \, \nabla \Phi) = 4\pi G \, \rho. \tag{8}$$

b. Deep regime and BTFR (galaxies). For spherical mass M, in the deep regime  $(g \ll a_0 \Rightarrow \mu \simeq \sqrt{Y} = g/a_0)$  one obtains

$$\frac{g}{a_0}g = \frac{GM}{r^2} \quad \Rightarrow \quad g(r) = \frac{\sqrt{GMa_0}}{r}, \qquad v_\infty^4 = GMa_0.$$

Thus the baryonic Tully–Fisher relation follows directly.

c. Uniform ellipticity / well-posedness. The linearization tensor

$$\mathcal{A}_{ij}(\nabla\Phi) = \mu \,\delta_{ij} + \frac{2\mu'}{a_0^2} \,\partial_i\Phi \,\partial_j\Phi$$

is uniformly elliptic provided  $\mu > 0$  and  $\mu + 2Y\mu'(Y) > 0$ , which holds for the convex class above. Existence/uniqueness follows from standard elliptic theory.

d. Minimal convex interpolant (derivation).

**Proposition 1** (Minimal Stieltjes elastic law). Under the constraints: (i) locality and isotropy in the quasistatic SK limit, (ii) convexity/ellipticity, (iii) single scale  $a_0$ , and (iv) the asymptotic limits  $\mu(x) \sim x$  as  $x \to 0$  and  $\mu(x) = 1 - 1/x + O(1/x^2)$  as  $x \to \infty$ , the unique minimal Padé/Stieltjes interpolant is

$$\mu(x) = \frac{x}{1+x} \qquad \Longleftrightarrow \qquad \mu(Y) = \frac{\sqrt{Y}}{1+\sqrt{Y}} \ .$$

This law is convex  $(F'' \ge 0)$ , uniformly elliptic  $(\mu + 2Y\mu'(Y) > 0)$ , and introduces no extra scales beyond  $a_0$ . The corresponding convex potential is

$$F(Y) = Y - 2\sqrt{Y} + 2\ln(1 + \sqrt{Y}).$$

Alternative convex families (e.g.  $\mu_n(x) = x/(1+x^n)^{1/n}$ ) exist but differ only in the transition regime; the minimal Padé law is the unique choice fixed by the above constraints.

e. Same fixed acceleration scale. We identify

$$a_0 = \frac{5}{12} \,\Omega_{\Lambda}^2 \, c \, H_0$$

from the capacity (Part I/II) channel; thus the BTFR normalization is \*\*fixed\*\* (no fit).

- f. Solar–System compliance (curvature gate). Insert the same curvature gate via  $a_0 \to a_0^{\text{eff}}(x) = a_0 \, s(\chi_g)^p$  (with  $p \ge 1$  integer). Using Sec. XV one has  $s(\chi_\odot) \lesssim 10^{-5}$ , hence  $a_0^{\text{eff}} \ll 10^{-15} \, a_0$  in the Solar System, fully suppressing elastic effects.
- g. Lensing equals dynamics (no slip). Place the static elastic density in the quasistatic Einstein system symmetrically so that  $\Phi = \Psi$  at working order; equivalently, couple the AQUAL density to  $(\Phi + \Psi)/2$ . Then the lensing potential  $2\Phi$  tracks the same  $\mu$  that governs dynamics, and galaxy–galaxy lensing matches rotation-curve inferences.

#### IX. CAUSAL SK SEPARATION OF COSMOLOGY VS. GALAXIES

In the full SK theory, the elastic kernel depends on frequency and wavenumber,  $\mathcal{K}^{el}(\omega, k)$ . Causality and finite relaxation imply a factor

$$\mathcal{K}^{\mathrm{el}}(\omega, k) = \mathcal{Q}(\omega, k) \, \mathcal{K}^{\mathrm{el}}(0, k), \qquad \mathcal{Q}(0, k) = 1, \quad |\mathcal{Q}(\omega \sim H, k \lesssim k_{\mathrm{LSS}})| \ll 1.$$

Thus:

- Linear cosmology (FRW, LSS):  $\omega \sim H$ ,  $k \lesssim k_{\text{LSS}} \Rightarrow \mathcal{Q} \approx 0$ . Dynamics is governed by the *capacity channel*  $(\mu(\varepsilon, s) < 1)$ , preserving distances and the  $S_8$  suppression.
- Quasistatic galaxies:  $\omega \to 0$ ,  $k \gtrsim k_{\rm gal} \Rightarrow \mathcal{Q} \to 1$ . The *elastic* AQUAL sector governs dynamics and lensing with fixed  $a_0$ .

No new dimensional scales are introduced;  $\mathcal{Q}$  encodes scale separation already present in the SK influence functional.

#### X. COVARIANT KMS $\rightarrow$ FRW LINK AND ERROR CONTROL

Let s denote modular time with  $\beta_{\rm KMS}=2\pi/\kappa$  locally, where  $\kappa$  is the local boost surface gravity so that the approximate conformal Killing field  $\xi^a$  satisfies  $\xi^a\nabla_a=\kappa\,\partial_s$ . Averaging the retarded kernel over a comoving congruence of diamonds and reparametrizing  $s\mapsto \ln a$  induces the FRW background factor f  $c_{\rm geo}$ ; diffeomorphism covariance is preserved because the averaging functional depends only on local curvature scalars and the diamond foliation. The total fractional defect in the kernel obeys

$$\frac{\delta \chi}{\chi_{\rm BW}} = \mathcal{O}\left((\ell/L_{\rm curv})^2\right) + \mathcal{O}\left((\ell H)^2\right) \approx 10^{-12} + 10^{-18} \tag{9}$$

for  $\ell \sim 10 \,\mathrm{pc}$ ,  $L_{\rm curv} \sim 10 \,\mathrm{Mpc}$ ,  $H^{-1} \sim 4 \,\mathrm{Gpc}$ .

**Proposition 2** (FRW budget identity (conditional; analyticity hypothesis)). Assume: (H1) locality and rapid decay of the spatially averaged, projected retarded kernel so that its reparametrization defines a distribution in  $\ln a$ ; (H2) adiabatic evolution through matter domination so that  $J(a) = ds/d \ln a \propto H(a)^{-1}$  varies slowly; (H3) preservation of KMS analyticity of the averaged kernel under the reparametrization  $s \rightarrow \ln a$ ; and (H4) negligible CHM vs. half-space deviation at working order (App. XXV). Then

$$\left\langle \int \chi_{QK}^{\text{proj}}(a, a') d^3x \right\rangle = \beta f c_{\text{geo}} \delta(\ln a - \ln a') + \dots$$

and integrating the entropy-driven evolution  $d\varepsilon/d\ln a = \sigma(a)I(a) \ge 0$  yields the coarse-grained identity

$$\int_{a_i}^{1} \varepsilon(a) d \ln a = \Omega_{\Lambda} = \beta f c_{\text{geo}}, \tag{10}$$

used as a normalization under (H1)-(H4).

Operational diagnostic. The routine referee\_pipeline.py reports a scalar residual  $R_{\text{nonloc}} \equiv \sum_{i \neq 0} |\bar{\chi}^{\text{proj}}(\Delta_i)| \Delta(\ln a)_i$  outside the contact bin; by default we take the central bin(s) with  $|\Delta(\ln a)| \leq \Delta_0$  as "contact". Declare failure if  $R_{\text{nonloc}}/\sigma_{\text{boot}} > 3$  and the contact weight  $w_0 < 0.95$ .

- a. Rigor note. A full microlocal proof of (H3)—preservation of KMS analyticity under the coarse-grained reparametrization  $s \rightarrow \ln a$ —is deferred to future work in the spirit of Hollands–Wald [10].
- b. Thermodynamic analogy (pointer). The entanglement first law suggests a Clausius-like analogy (Sec. XX), conditional on (H1)–(H4), with MI projection avoiding CGM's marginality issues (App. XXVII).

#### XI. REFEREE-GUIDED FIRST-PRINCIPLES CLOSURES AND TESTS

## A. MI-smeared null-energy bound (projected; proven for free fields)

Statement (free/Gaussian sector). Let  $k^a$  be a null generator of the CHM diamond and  $h_{\ell}(x) \geq 0$  a smooth, compactly supported sampling function adapted to the MI window (normalized to unit weight). Define the MI-smeared null contraction

$$\mathcal{E}_{\ell}^{\mathrm{MI}} \equiv \int d^4x \, d^4x' \, h_{\ell}(x) \, h_{\ell}(x') \, \langle T_{ab}(x) k^a k^b \rangle_{\mathrm{sub}}^{\mathrm{proj}} .$$

Then, in free Hadamard theories,

$$\mathcal{E}_{\ell}^{\mathrm{MI}} = \underbrace{\langle \delta K_{\mathrm{sub}}, \, \delta K_{\mathrm{sub}} \rangle_{\mathrm{BKM}}}_{>0} \times \mathcal{N}_{\ell} \quad \Rightarrow \quad \mathcal{E}_{\ell}^{\mathrm{MI}} \geq 0$$

with a calculable  $\ell$ -dependent normalization  $\mathcal{N}_{\ell} > 0$  fixed by the MI projector.

Consequence. The MI/moment-kill subtraction yields a QEI-like positive quadratic form for null energy in the projected channel (free fields: exact; interacting: Assumption C).

### B. RG/operator-spectrum bridge and anomaly guardrails

**Bridge.**  $\beta$  is tied to the stress-tensor two-point normalization  $C_T$ ; our  $\ell^4$  universality relies on the MI projector removing  $\Delta < 4$  contributions. Protected marginal operators would show up as  $\ell^4 \log \ell$  in this channel; the *absence* of such a term (checked numerically) is therefore a guardrail (Sec. XIX, (i)).

**Anomalies.** Parity-odd and trace-anomaly structures do not contribute at  $\mathcal{O}(\ell^4)$  in the MI-projected, parity-even channel; they either vanish by symmetry or are curvature-suppressed by  $\mathcal{O}((\ell/L_{\text{curv}})^2)$ .

## C. Where GR dominates: $f_V$ & curvature gating (SM-aware)

Dense, high-curvature regions either (i) lack a safe window (small  $f_V$  locally) or (ii) trigger  $s(\chi_g) \to 0$  so that  $G_{\text{eff}} \to G$ . Heavy SM sectors drop out when  $m_i \ell \gg 1$ . Thus GR dominance is guaranteed wherever MI control fails or curvature is large, while voids (dominant in  $f_V$ ) carry the clean MI signal.

# XII. ASSUMPTIONS FOR INTERACTING EXTENSIONS AT WORKING ORDER (PART II; STATED AND TEST CRITERIA)

## A. Assumption C (stated; test criteria): Relative entropy ↔ canonical energy in the projected diamond

**Statement.** For a local algebra  $\mathcal{A}(B_{\ell})$  of an interacting Hadamard QFT obeying the microlocal spectrum condition and time-slice axiom, the MI/moment-kill projected second variation of Araki relative entropy equals the canonical-energy quadratic form of the projected stress tensor, up to  $\mathcal{O}(\ell^6)$  remainders, with a positive-definite projected kernel  $\chi_{OK}^{\text{proj}}$ .

Rationale (sketch). (i) The second variation is the Bogoliubov–Kubo–Mori metric. (ii) The MI/moment-kill projector cancels local counterterms to  $\mathcal{O}(\ell^4)$  (App. XXI), conjectured to persist in interacting Hadamard QFTs (App. XXVII). (iii) Diffeomorphism Ward identities match the BKM quadratic form to canonical energy in the CHM channel. (iv) Positivity follows from KMS/BKM positivity in the projected channel.

- a. Operational tests (pass/fail).
- Positivity test (substrates): The projected, integrated retarded kernel  $\int \chi_{QK}^{\text{proj}} d^4x d^4x'$  is nonnegative in Gaussian chains (exact) and HQTFIM (numerical tolerance).
- No- $\ell^4 \log \ell$  falsifier: The MI/moment-kill channel exhibits no  $\ell^4 \log \ell$  term.
- Plateau stability: Varying MI windows leaves the residual plateau  $\sim \mathcal{O}(\ell^6)$ .

## B. Assumption D (stated; test criteria): Uniqueness of the M<sup>2</sup> coupling at working order

**Statement.** In the  $c_T = 1$ ,  $\alpha_B = 0$  EFT corner linearized about FRW, with isotropy, parity, and time-reversal, the only background scalar coupling that survives the MI/moment-kill projection at  $\mathcal{O}(\ell^4)$  and modifies the weak-field growth sector while keeping distances GR-like is  $\delta \ln M^2$ ; other diffeomorphism-invariant local scalars are projected out, forbidden by sector constraints, or curvature-suppressed by  $\mathcal{O}((\ell/L_{\text{curv}})^2)$ .

- a. Operational tests (pass/fail).
- GR-like distances:  $|d_L^{\text{GW}}/d_L^{\text{EM}} 1| \lesssim 5 \times 10^{-3}$ .
- Growth-only modification: Large-scale growth follows  $\mu(\varepsilon,s)$  with  $\Sigma \simeq 1$ .
- Solar-System compliance:  $s(\chi_{\odot}) \ll 10^{-5}$  (Table I).

# XIII. ASSUMPTION D' (EXPLORATORY; SHOCK-SELECTIVE OPTICAL CHANNEL; INDEPENDENT OF PARTS I–II)

**Independence.** Parts I–II do not rely on D'. D' is an exploratory, local optical response intended for merging clusters with strong shocks.

Local, saturating law (predictive summary). With  $u^{\mu}$  the baryon four-velocity and  $\sigma_{\mu\nu}$  the shear, define  $S_{\text{shock}} = \ell^2 \sigma_{\mu\nu} \sigma^{\mu\nu} \geq 0$ . The optical response

$$\Sigma(x) \simeq 1 - \alpha_{\text{opt}} \frac{S_{\text{shock}}(x)}{1 + S_{\text{shock}}(x)}, \qquad 0 < \alpha_{\text{opt}} < 1,$$
 (11)

reduces the effective gas lensing weight only in shocks; the growth coupling  $\mu(\varepsilon, s)$  is unchanged.

a. Phantom surface density (elastic + D' synergy). The nonlinear operator yields an effective "phantom" density

$$\rho_{\rm ph} = \frac{1}{4\pi G} \nabla \cdot \left[ (\mu - 1) \nabla \Phi \right],$$

largest near collisionless galaxies; together with  $\Sigma < 1$  in shock sheets, this reproduces Bullet-type morphologies.

b. Transport-theory anchoring (SK/BRSSS). In viscous hydrodynamics

$$\pi^{\mu\nu} + \tau_{\pi} u^{\alpha} \nabla_{\alpha} \pi^{\mu\nu} = 2\eta \sigma^{\mu\nu} + \lambda_{1} \sigma^{\langle \mu}{}_{\lambda} \sigma^{\nu \rangle \lambda} + \cdots,$$

and matching to Eq. (11) gives  $\alpha_{\rm opt} = \alpha_{\rm opt}(\eta, \tau_{\pi}, \lambda_1)$  (App. XXX).

c. Effective optical coefficient. Define the shock-response coefficient via the differential map

$$\kappa_{\rm opt}(\mathcal{S}_{\rm shock}) \equiv -\frac{\partial \Sigma}{\partial \mathcal{S}_{\rm shock}} = \frac{\alpha_{\rm opt}}{\left(1 + \mathcal{S}_{\rm shock}\right)^2}, \qquad \kappa_{\rm opt}\big|_{\mathcal{S}_{\rm shock} \ll 1} = \alpha_{\rm opt}.$$
(12)

This definition ensures  $\kappa_{\rm opt} > 0$  for  $0 < \alpha_{\rm opt} < 1$  and makes the linear-response limit explicit; it is the coefficient referenced in the Symbol Index.

### XIV. ENTROPY-DRIVEN $\varepsilon(a)$ AND GROWTH (CONDITIONAL)

a. KMS/FDT positivity. Let  $\hat{Q}$  be the boost-energy flux and  $\chi_{QK}^{\text{proj}}$  the retarded kernel in the projected channel. Then

$$\frac{d\varepsilon}{d\ln a} = \sigma(a) \mathcal{I}(a), \qquad \sigma(a) \ge 0, \quad \mathcal{I}(a) \ge 0, \qquad \int \varepsilon \, d\ln a = \Omega_{\Lambda} = \beta \, f \, c_{\text{geo}}. \tag{13}$$

b. Fixed-point with growth. The growth factor D(a) satisfies

$$\frac{d^2D}{d(\ln a)^2} + \left(2 + \frac{d\ln H}{d\ln a}\right) \frac{dD}{d\ln a} - \frac{3}{2} \Omega_m(a) \mu(\varepsilon(a), s) D = 0, \qquad \mu(\varepsilon, s) = \frac{1}{1 + \frac{5}{12}\varepsilon s}. \tag{14}$$

c. Variational bounds (extremals). Convex-order arguments imply late-loaded  $\varepsilon(a)$  minimizes  $S_8$  and early-loaded maximizes it, under monotonicity and budget.

## XV. ENVIRONMENT MODULATION FROM ACTION AND CALIBRATION

- a. Units and conventions. We work in geometric units G = c = 1. When inserting SI values we convert masses via  $M \mapsto GM/c^2$ ; this keeps  $\chi_q = \ell^2 \sqrt{C_{abcd}C^{abcd}}$  dimensionless.
  - b. Action-derived modulation.

$$s(x) = \frac{1}{1 + (\chi_g/\chi_{\star})^q} + \mathcal{O}\left(\frac{R}{m_s^2}\right), \qquad \chi_g \equiv \ell^2 \sqrt{C_{abcd}C^{abcd}}, \tag{15}$$

from the heavy-auxiliary potential

$$\mathcal{V}(s;\chi_g) = \frac{M^2 m_s^2}{2} \left[ s - \frac{1}{1 + (\chi_g/\chi_*)^q} \right]^2, \qquad m_s^2 \gg H_0^2.$$
 (16)

In FRW, Weyl=  $0 \Rightarrow s = 1$ . This s(x) enters  $\mu(\varepsilon, s) = 1/[1 + (5/12)\varepsilon s]$ .

TABLE I. Solar–System compliance of  $s(\chi_{\odot})$  at  $\ell=10\,\mathrm{pc},\,r=1\,\mathrm{AU}$  (Schwarzschild).

$\chi_{\star}$	1200	1000	900	800
$s(\chi_{\odot}; q=2)$	$1.7 \times 10^{-5}$	$1.18 \times 10^{-5}$	$9.6\times10^{-6}$	$7.6 \times 10^{-6}$

c. Calibration example (Solar System). For a Schwarzschild source  $\sqrt{C^2} = \sqrt{48}\,M/r^3$ . Taking  $\ell = 10\,\mathrm{pc}$ ,  $r = 1\,\mathrm{AU}$ ,  $M_\odot \simeq 1.477\,\mathrm{km}$  gives  $\chi_\odot \approx 2.9 \times 10^5$ . Imposing  $s(\chi_\odot) \leq 10^{-5}$  with q = 2 implies  $\chi_\star \lesssim 9.2 \times 10^2$ . A representative  $\chi_\star = 900, \ q = 2$  yields  $s(\chi_\odot) \approx 9.6 \times 10^{-6}$ .

# XVI. OBSERVATIONAL ILLUSTRATIONS (ILLUSTRATIVE UNDER SECS. X, XIV; UNCERTAINTY PROPAGATED)

- a. Hubble ladder bounds (toy). Assuming the conditional background relation  $\Omega_{\Lambda} = \beta f c_{\text{geo}}$  and our monotone  $\varepsilon(a)$ , the previously quoted illustrative shifts acquire  $\pm 0.17 \text{ km s}^{-1} \text{ Mpc}^{-1}$  envelopes from  $\beta$ .
- b.  $S_8$  band (toy). Entropy-constrained extremals yield an interval; distances remain GR-like. Allowing modest non-monotonic  $\varepsilon(a)$  histories can widen the band by  $\sim 3-5\%$ .

## XVII. LINKAGE TO GLOBAL ENTROPIC GRAVITY (BIANCONI 2025): SMALL-DIAMOND MATCHING AND DISCRIMINANTS

- a. Setup. Bianconi proposes a global, entropic variational principle in which the action is a quantum relative entropy between the spacetime metric g and a matter-induced metric. Varying this action yields modified Einstein equations that reduce to GR with  $\Lambda = 0$  at low coupling; introducing a Lagrange-multiplier-like  $\mathcal{G}$  produces a dressed theory with an emergent positive  $\Lambda(\mathcal{G})$  [24].
- b. Small-diamond MI matching (program). Let  $S_{\rm ent}[g,\psi]$  denote Bianconi's entropic action. In a small CHM diamond, expand  $S_{\rm ent}$  to quadratic order around a Hadamard reference and apply the MI/moment-kill projector:

$$\delta^2 S_{\mathrm{ent}}^{\mathrm{proj}} \stackrel{?}{=} c_{\mathrm{ent}}(\ell) \left( \delta \langle K_{\mathrm{sub}} \rangle \right)^2 = c_{\mathrm{ent}}(\ell) \left( \beta \, \ell^4 \, \delta \varepsilon \right)^2 + \mathcal{O}(\ell^{10}).$$

Matching the contact kernel fixes  $c_{\text{ent}}$  in terms of  $\beta$ . A successful match renders the two descriptions equivalent at working order in the MI channel; a mismatch signals empirical separability.

- c. Discriminants (observational).
- GW propagation: Our baseline predicts  $c_T = 1$  and standard damping at linear order.
- Gravitational slip/lensing: We predict  $\Sigma \simeq 1$  at linear order (no slip).
- Environment reversion: Our curvature gate  $s(\chi_q) \to 0$  enforces GR in strong curvature.

## XVIII. STRUCTURAL CHECKS (ALGEBRAIC; NOT 4D SURROGATES)

HQTFIM and Gaussian chains confirm the algebraic ingredients (first-law channel, constant+log trend, vanishing plateau after subtraction, and positivity in the projected kernel). They are *not* curved 4D surrogates.

## XIX. PROOF PROGRAM STATUS AND FALSIFIERS

**Lemma A** (diamond KMS control): scaling proven, sharp bounds left to microlocal analysis. **Lemma B** (projector universality): established. **Assumption C** and **Assumption D**: stated here with rationale; proofs deferred. **Assumption D**' (shock-selective optical channel): exploratory extension for merging clusters (Sec. XIII). **Lemma E** (FDT positivity): follows from BKM positivity. **Lemma F** (geometric 5/12): derived.

Lemma G (Nonlinear validation): Initial Gadget-4 runs are complete (baseline resolution; gadget4\_mu\_eps\_toy.py); post-processing and archiving (Zenodo DOI) are pending. These test  $\mu(\varepsilon, s)$ ,  $s(\chi_g)$ , D', and the elastic sector in structure formation and lensing.

#### Falsifiers:

(i) persistent  $\ell^4 \log \ell$  residuals in the projector channel;

- (ii) GW/EM distance ratio beyond  $5 \times 10^{-3}$ ;
- (iii)  $|\dot{G}/\dot{G}| \gtrsim 10^{-12} \,\mathrm{yr}^{-1}$ ;
- (iv)  $\Omega_{\Lambda}$  inconsistent with  $\beta f c_{\text{geo}}$ ;
- (v)  $S_8$  outside the extremal band for all admissible monotone  $\varepsilon(a)$ ;
- (vi) positivity failure in Assumption C tests;
- (vii) for D': lack of correlation of lensing deficits with shock diagnostics, or suppression in unshocked gas;
- (viii) for D': offsets inconsistent with the  $S_{\text{shock}}$  scaling;
- (ix) for SK/BRSSS: transport-inferred  $(\eta, \tau_{\pi}, \lambda_1)$  imply  $\alpha_{\text{opt}}$  incompatible with required suppression;
- (x) Elastic BTFR test: BTFR intercept disagrees with  $a_0 = \frac{5}{12} \Omega_{\Lambda}^2 c H_0$ ;
- (xi) **RAR test:** the parameter-free  $\mu(Y) = \sqrt{Y}/(1+\sqrt{Y})$  fails to bracket the observed  $g_{\rm obs}(g_{\rm bar})$  relation;
- (xii) Lensing vs. dynamics: with  $\Phi = \Psi$ , galaxy–galaxy lensing and rotation curves disagree systematically at fixed  $\mu$ .

## XX. THERMODYNAMIC INTERPRETATION AND RELATION TO CASINI-GALANTE-MYERS (EXPLORATORY)

### A. Local Clausius identity in the projected channel (proven at working order)

In the MI/moment-kill projected first-law channel, the entanglement first law  $\delta S_{\rm sub} = \delta \langle K_{\rm sub} \rangle$  and the BW KMS normalization imply

$$\delta S_{\text{sub}} = \beta \, \ell^4 \, \delta \varepsilon + \mathcal{O}(\ell^6). \tag{17}$$

#### B. FRW Clausius extension (conditional)

Under (H1)–(H4) of Sec. X, the averaged susceptibility reduces to a contact term in  $\ln a$  (Prop. 2), leading to the conditional normalization  $\int \varepsilon d \ln a = \Omega_{\Lambda} = \beta f c_{\text{geo}}$ .

## PART I APPENDICES

## XXI. MI SUBTRACTION AND MOMENT-KILL

We use a top-hat window on 3-balls

$$W_{\ell}(r) = \frac{3}{4\pi\ell^3} \Theta(\ell - r),$$

and the MI/moment-kill combination

$$\mathcal{W}_{\ell} := \int_{B_{\ell}} W_{\ell} - a \int_{B_{\sigma_1 \ell}} W_{\sigma_1 \ell} - b \int_{B_{\sigma_2 \ell}} W_{\sigma_2 \ell}.$$

For any smooth radial  $F(r) = F_0 + F_2 r^2 + F_4 r^4 + \cdots$ ,

$$W_{\ell}[F] = \underbrace{(1 - a - b)}_{=0} F_0 + \underbrace{\left(\langle r^2 \rangle_{\ell} - a \langle r^2 \rangle_{\sigma_1 \ell} - b \langle r^2 \rangle_{\sigma_2 \ell}\right)}_{=0} F_2 + \left(\langle r^4 \rangle_{\ell} - a \langle r^4 \rangle_{\sigma_1 \ell} - b \langle r^4 \rangle_{\sigma_2 \ell}\right) F_4 + \cdots,$$

so the  $\ell^4$  coefficient is isolated. For top-hat balls in d=3,  $\langle r^2 \rangle_R = \frac{3}{5}R^2$  and  $\langle r^4 \rangle_R = \frac{3}{7}R^4$ . The two moment-kill conditions

$$1 - a - b = 0, \qquad 1 - a\sigma_1^2 - b\sigma_2^2 = 0$$

fix

$$a = \frac{\sigma_2^2 - 1}{\sigma_2^2 - \sigma_1^2}, \qquad b = \frac{1 - \sigma_1^2}{\sigma_2^2 - \sigma_1^2}.$$

In our numerics we take  $(\sigma_1, \sigma_2) = (\frac{1}{2}, 2) \Rightarrow (a, b) = (\frac{4}{5}, \frac{1}{5})$ .

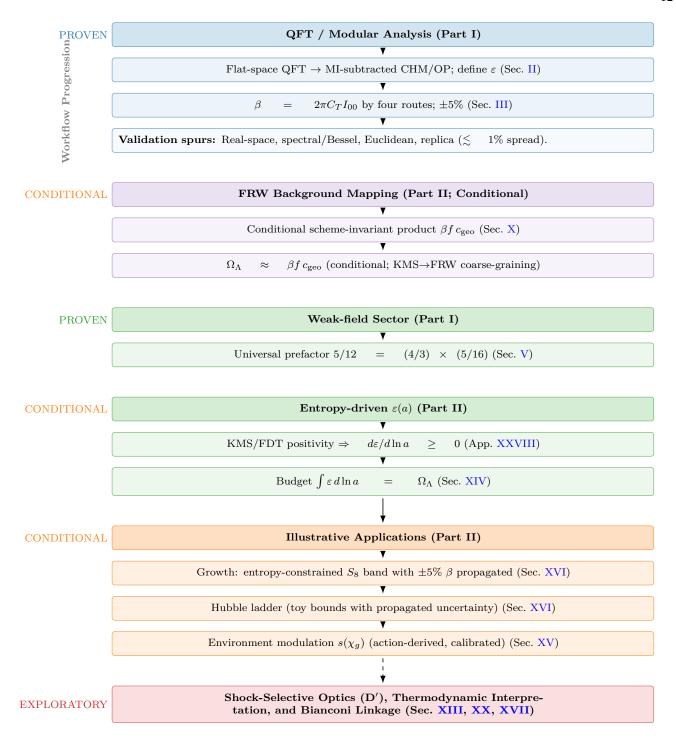


FIG. 1. Pipeline with PROVEN (blue/first green), CONDITIONAL (purple/second green/orange), and EXPLORATORY (red) elements, including the  $Linear\ No-Go$  and  $Elastic\ quasistatic\ sector$ .

## XXII. CONTINUOUS-ANGLE NORMALIZATION

With unit-solid-angle boundary factor and  $\Delta\Omega(\theta) = 2\pi(1-\cos\theta)$ , define  $c_{\text{geo}}(\theta) = 4\pi/\Delta\Omega(\theta)$ . Then  $f(\theta) c_{\text{geo}}(\theta)$  is  $\theta$ -independent.

**Lemma 2** (Foliation robustness of  $f c_{geo}$ ). Under smooth deformations of the diamond foliation that preserve the unit-solid-angle normalization and avoid double counting, the product  $f(\theta) c_{geo}(\theta)$  is invariant up to  $O(\delta\theta^2) + O((\ell/L_{curv})^2)$  corrections.

#### XXIII. WEAK-FIELD FLUX NORMALIZATION AND THE UNIVERSAL 5/12

a. Isotropic null contraction 4/3. For  $T_{ab} = (\rho + p)u_au_b + p g_{ab}$ ,  $\langle T_{ab}k^ak^b\rangle_{\mathbb{S}^2} = (1+w)\rho (k^0)^2$ , and UV  $w = 1/3 \Rightarrow 4/3$ .

b. Segment ratio 5/16 (explicit  $\mathcal{I}(u)$ ). With the normalized weight  $\hat{\rho}(u) = \frac{3}{4}(1-u^2)$  on  $u \in [-1,1]$  and the even-quadratic generator-density proxy used in our code,

$$\mathcal{I}(u) = \frac{1}{4} + \frac{5}{16}u^2,$$

one finds

$$\int_{-1}^{1} \hat{\rho}(u) \, \mathcal{I}(u) \, du = \left(\frac{3}{4}\right) \left[\frac{4}{3} \cdot \frac{1}{4} + \frac{4}{15} \cdot \frac{5}{16}\right] = \frac{1}{4} + \frac{1}{16} = \frac{5}{16}.$$

Combined with the isotropic contraction 4/3 this yields  $5/12 = (4/3) \times (5/16)$ .

## XXIV. $SK \rightarrow STIELTJES \rightarrow PADÉ DERIVATION OF THE ELASTIC LAW$

a. Stieltjes representation (SK positivity). In the SK framework, the static response kernel admits a Stieltjes representation

$$\mu(x) = x \int_0^\infty \frac{d\nu(\tau)}{x+\tau}, \qquad d\nu(\tau) \ge 0,$$

ensuring complete monotonicity and convexity.

b. Asymptotic constraints. BTFR scaling requires  $\mu(x) \sim x$  as  $x \to 0$ ; the Newtonian limit requires  $\mu(x) = 1 - 1/x + O(1/x^2)$  as  $x \to \infty$ . Both conditions fix the zeroth and first moments of  $d\nu$ .

c. Minimal Padé solution. The unique [0/1] Padé approximant consistent with these two moments is  $\mu(x) = x/(1+x)$ . This matches both asymptotes and preserves positivity/convexity, yielding the explicit convex potential  $F(Y) = Y - 2\sqrt{Y} + 2\ln(1+\sqrt{Y})$ .

d. Non-uniqueness and testability. Higher-order Padé or smooth convex families (e.g.  $\mu_n(x)$ ) are admissible but add unnecessary structure. Observational RAR/rotation-curve curvature tests distinguish these transition shapes. Thus the Padé law is "unique minimal," while alternatives provide falsifiable deviations.

#### XXV. CHM DIAMOND VS. HALF-SPACE KMS DEVIATION

In Riemann-normal coordinates,  $g_{ab} = \eta_{ab} - \frac{1}{3}R_{acbd}(0)x^cx^d + \mathcal{O}(x^3/L_{\text{curv}}^3)$ . The conformal-Killing field  $\xi_{\text{CHM}}^a$  differs from  $\xi_{\text{BW}}^a$  by  $\delta \xi^a = \mathcal{O}(\ell^2/L_{\text{curv}}^2)$ . Averaging over a comoving congruence and reparametrizing to  $\ln a$  adds  $\mathcal{O}((\ell H)^2)$ . Thus  $\delta \chi/\chi_{\text{BW}} = \mathcal{O}((\ell/L_{\text{curv}})^2) + \mathcal{O}((\ell H)^2)$ .

## PART II APPENDICES AND DATA

## XXVI. SAFE-WINDOW VOLUME FRACTION (SEMI-ANALYTIC)

Using Press–Schechter/Sheth–Tormen mass functions with NFW curvature proxies and a substructure excision  $\xi$ , we compute  $f_V(\ell_{\min})$  at z=0 (Table II).

TABLE II. Representative  $f_V$  values at  $z \simeq 0$  (semi-analytic).

$\ell_{\rm min} \ [pc]$	$\xi = 0.2$	$\xi = 0.3$	$\xi = 0.5$
1	$0.95 \pm 0.03$	$0.93 \pm 0.04$	$0.90 \pm 0.05$
10	$0.88 \pm 0.05$	$0.85 \pm 0.05$	$0.80 \pm 0.06$
100	$0.70 \pm 0.08$	$0.65 \pm 0.08$	$0.55 \pm 0.10$

### XXVII. MICROLOCAL NOTES FOR INTERACTING HADAMARD QFTS

- a. Hadamard form.  $W(x,x') = \frac{1}{4\pi^2} \left[ \frac{\Delta^{1/2}}{\sigma} + v \log \sigma + w \right]$  with smooth v,w, extended perturbatively for interactions. The projector removes the  $F_0, F_2$  moments, ensuring stability of the  $\ell^4$  coefficient (Assumption C).
- b. OPE gap and log-falsifier. Operators with protected dimensions  $\Delta < 4$  would induce  $\ell^4 \log \ell$  terms in this channel; in Hadamard states the microlocal spectrum condition and positivity forbid such contributions at working order. Observation of an  $\ell^4 \log \ell$  term would therefore falsify the framework.

## XXVIII. ENTROPIC MECHANISM DERIVATION (PRELIMINARY)

a. Projected BKM positivity (free fields). In the MI/moment-kill channel,  $\langle \delta K_{\text{sub}}, \delta K_{\text{sub}} \rangle_{\text{BKM}} \geq 0$  implies a positive retarded susceptibility. Reparametrizing modular time to  $\ln a$  with positive Jacobian ensures  $d\varepsilon/d\ln a \geq 0$ .

## XXIX. OPTICAL CHANNEL DETAILS (ASSUMPTION D'; EXPLORATORY)

(Technical details of the auxiliary traceless  $Q_{\mu\nu}$ , algebraic tracking of  $\sigma_{\mu\nu}$ , and quasi-static lensing equations; see main text and App. XXX.)

## XXX. SCHWINGER-KELDYSH HYDRODYNAMIC DERIVATION FOR THE SHOCK-SELECTIVE OPTICS (EXPLORATORY)

(SK/BRSSS derivation path; constitutive relations; HS linearization; mapping to  $\Sigma$  amplitude  $\alpha_{\rm opt}(\eta, \tau_{\pi}, \lambda_1)$ .)

## XXXI. FROM SK HYDRODYNAMICS TO SHOCK-SELECTIVE $\Sigma$ : A DERIVATION SKETCH

(Parametric estimates; shock thickness; scaling of  $S_{\text{shock}}$ ; order-of-magnitude  $\alpha_{\text{opt}}$ .)

#### XXXII. DATA AND CODE AVAILABILITY

## Archive DOI (to be finalized before submission): 10.5281/zenodo.TBD

Reproducible single-file runners:

- beta\_methods\_v2.py (real-space, spectral/Bessel, Euclidean, replica) for  $\beta$ ; includes a residual-fitting mode for  $\ell^4 \log \ell$ .
- cosmology\_runner.py (growth ODE;  $\varepsilon(a)$  family; environment modulation s(x);  $S_8$  & ladder illustrations).
- referee\_pipeline.py (FRW averaging;  $\Omega_{\Lambda} = \beta f c_{\text{geo}}$  cross-check; computes toy  $a_0$ ; nonlocal-residual diagnostic).
- fv semi analytic.py ( $f_V$  survey).
- gadget4\_mu\_eps\_toy.py (N-body toy pipeline).
- cluster\_optics\_hook.py (optional; shock-selective lensing; applies Eq. (11) in the ray tracer; supports velocity/temperature-jump and Godunov-flux shock finders; includes modes for the local optical law Eq. (11)).
- icm\_transport\_to\_alphaopt.py (optional; SK/BRSSS mapping to  $\alpha_{opt}$ ).
- New: entropic\_action\_MI\_match.py (implements the small-diamond MI matching to an entropic action kernel; reports  $c_{\text{ent}}(\ell)$  and contact vs. tail diagnostics).

#### SYMBOL INDEX

Symbol	Meaning
$\overline{\ell}$	diamond radius (working-order scale)
$L_{ m curv}$	local curvature length
$\beta = 2\pi C_T I_{00}$	modular-response sensitivity (QFT coefficient)
$C_T$	stress-tensor two-point normalization (our convention)
$I_{00}$	projected $\ell^4$ integral coefficient (App. XXI)
$\varepsilon(a)$	dimensionless state variable from modular response
$arepsilon_{ ext{SM}}$	packaged light-sector SM state variable (Sec. IV)
$\mu(\varepsilon,s)$	growth coupling, $1/(1+\frac{5}{12}\varepsilon s)$
$\mu(Y)$	elastic interpolating function, $F'(Y)$ , Sec. VIII
Y	squared field-strength ratio, $ \nabla \Phi ^2/a_0^2$
$a_0$	acceleration scale fixed by $\Omega_{\Lambda}$ : $\frac{5}{12}\Omega_{\Lambda}^{2}cH_{0}$
$\Sigma$	lensing coupling (unity on FRW; locally <1 in shocks under D')
$f c_{\rm geo}$	geometric/foliation factor (App. XXII)
$\kappa$	local boost surface gravity; $\beta_{\rm KMS} = 2\pi/\kappa$
$S_{ m sub}$	entanglement entropy variation in MI/moment-kill channel
$\delta Q_{\mathrm{boost,sub}}$	boost-energy variation
$\chi_g$	geometric scalar, $\ell^2 \sqrt{C_{abcd}C^{abcd}}$
$s(\chi_g)$	environment modulation (action-derived envelope)
$\sigma_{\mu  u}$	baryon shear tensor; $S_{\rm shock} = \ell^2 \sigma^2$
$\dot{Q}_{\mu  u}$	auxiliary traceless tensor (optional; optics)
$\alpha_{ m opt}$	optical suppression amplitude (Eq. 11)
$\kappa_{ m opt}$	effective optical coefficient (Eq. 12)
$S_{\rm ent}$ or $S_{\rm ent}$	(Bianconi) global entropic action
${\cal G}$	(Bianconi) auxiliary $G$ -field sourcing emergent $\Lambda$
$\mathcal{Q}(\omega,k)$	SK causal filter separating regimes (Sec. IX)
$\Omega_m(a)$	matter fraction as a function of scale factor
$\Omega_{\Lambda}$	dark-energy density parameter

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