

# Effective Area Studies with a Toy Generator

Cosmin Deaconu

UChicago/KICP

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# Context

- In the fall, I became worried that the  $A_{eff} = V_{eff}/L_{int}$  approximation may not be right for a non-infinite disc geometry.
  - ▶ In the non-infinite disc case, cannot factor out the projected length from the trajectory.
  - ▶ Also, two different generators for ANITA MC gave factor of two difference (much more complicated geometry there, though). One generator uses  $V_{eff}/L_{int}$ , the other computes  $A_{eff}$  directly.
- So I wrote a toy MC to test the two generators.

# Toy MC

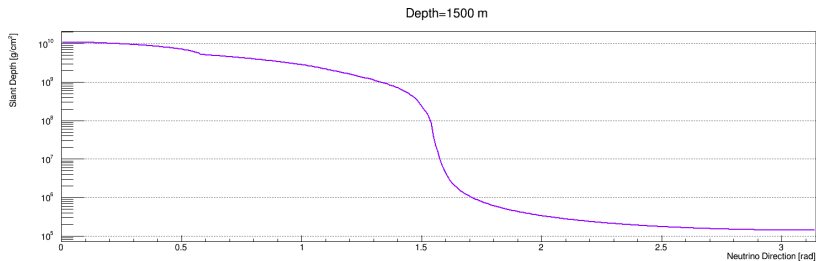
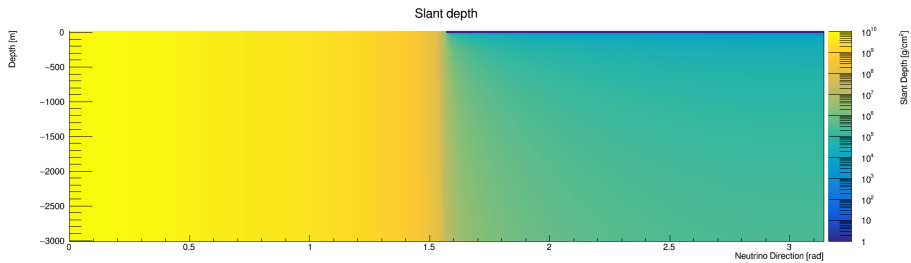
- <https://github.com/cozzyd/ToyDisc>
- From-scratch very basic Askaryan neutrino simulation in a cylinder of uniform ice, with two different generators and  $A_{eff}$  procedures..
- Implements simple interaction / inelasticity / Askaryan emission model and electric field threshold at detector for trigger.
  - ▶ Connolly et al inelasticity, interaction model.
  - ▶ Same Askaryan emission parameterization as used in `icemc` .
  - ▶ Since uniform ice, no raytracing. Use constant attenuation length.
- **Can't compare effective areas with a real simulation** (due to simple trigger and uniform ice model). The point is to compare the two generators in a simple setup that can run fast (100 million events in about 10 minutes on my workstation).

## Procedure A (the standard forced method)

- Generate random direction and random position within cylinder.
- Compute “absorption weight,”  $w_a$ : the chance of the neutrino not having interacted before entering the simulation volume.
  - ▶ Use PREM model with ice layer on top. Generate histogram of column density vs. angle and depth for interpolation (and subtract off column density from entrance point in volume to interaction).
- Also keep track of interaction length for each neutrino (since flavor, charge-dependent).  
Leads to:

$$\langle A_{eff} \rangle |_{\Omega} = \left\langle \frac{V_{eff}}{L_{int}} \right\rangle = \frac{V_{mc}}{N} \sum_i \frac{w_{a,i}}{L_{int,i}} pass_i \quad (1)$$

# Particle depth histogram

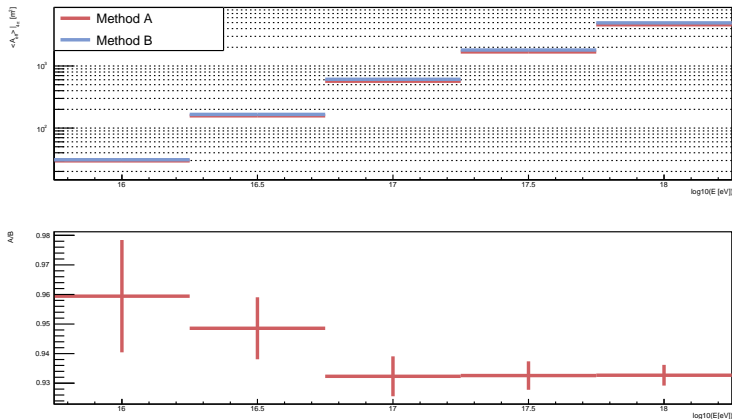


## Procedure B (area method)

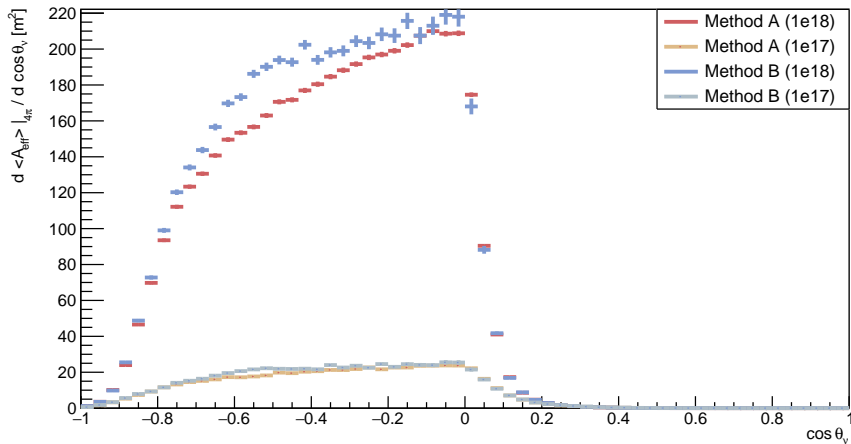
- Generate a random direction.
- Sampling area is a box passing in center of volume normal to neutrino direction. Box is big enough to cover largest projected area ( $L = \sqrt{4 * R^2 + h^2}$ ). Randomly pick offset within sampling box.
- Check for intersection of trajectory with the simulation volume. If no intersection, then continue.
- If trajectory intersects volume, pick a point for interaction (exponential weighting, but linear would work fine too with the interaction lengths considered here) and calculate two weights:
  - ▶  $w_a$ , identical to before.
  - ▶  $w_x$ , probability of interacting in volume ( $1 - \exp(-L_{proj}/L_{int})$ )
- In this case,

$$\langle A_{eff} \rangle |_{\Omega} = \frac{A_{box}}{N} \sum_i w_{a,i} w_{x,i} pass_i \quad (2)$$

# Effective area results ( $R = 15$ km, $h=3$ km)



# Elevation Dependence





# Discussion

- Small, but statistically significant difference between the two generators.
  - ▶ Size of effect on real simulation might be different when including raytracing, more sophisticated trigger.
  - ▶ Size of effect on ANITA, which samples a much narrower angular range with a more complex geometry, might be much larger.
- Note that since  $w_x \approx L_{proj}/L_{int}$ , the two expressions for  $\langle A_{eff} \rangle |_{\Omega}$  are not so different.
  - ▶ If instead of a box independent of the incoming angle, we used the projected area for each angle (so no trajectories miss), then you would get almost the same expression as A (since  $V_{mc} = \langle L_{proj} \rangle A_{proj}$ ), but I don't think you can factor out the integral over  $L_{proj}$  out of the other factors in the general case, possibly accounting for the difference.
  - ▶ Suggests possible correction factor of  $L_{proj}/(\langle L_{proj} \rangle (\theta))$  (or something like that) that might bring the two together (haven't tried this yet).