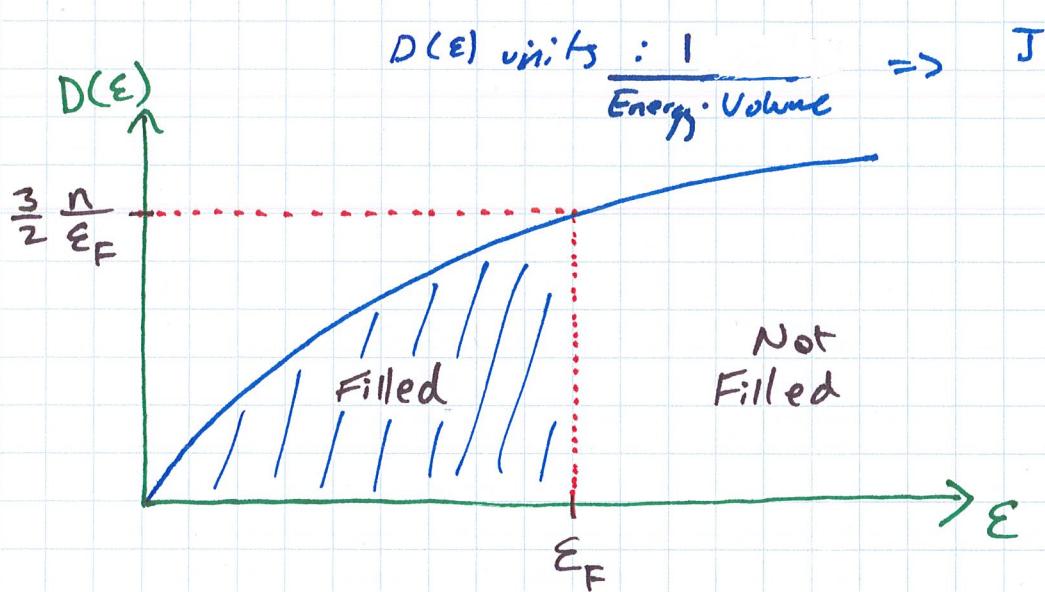


Note:

$$D(\epsilon_F) = \frac{3}{2} \frac{n}{\epsilon_F} = \frac{m k_F}{h^2 \pi^2}$$

$$D(\epsilon) = D(\epsilon_F) \sqrt{\frac{\epsilon}{\epsilon_F}}$$



$D(\epsilon)$ units: $\frac{1}{\text{Energy} \cdot \text{Volume}} \Rightarrow \text{J}^{-1} \text{m}^{-3}$

Quantum Free \bar{e} (Sommerfeld) @ finite temperatures

→ moving to $T > 0$ means \bar{e} will fill our single \bar{e} allowed states in a slightly different manner

Grand canonical ensemble:

→ expected single-state occupancy $\langle N_i \rangle \Rightarrow (f(\epsilon_k))$

$$\langle N_i \rangle = k_B T \frac{1}{Z} \left(\frac{\partial Z}{\partial \mu} \right)_{V,T}$$

$\mu \rightarrow$ chemical potential

$Z \rightarrow$ partition function

For any state (allowed \vec{k}) the partition function is given by a sum over all allowed occupancies $f(N_i)$: i.e. microstates.

$$Z = \sum_i e^{(N_i \mu - E_i)/k_B T}$$

For e^- 's (Fermions), there are only two microstates:

$$N_i = 0, \therefore E_i = 0$$

$$N_i = 1, \therefore E_i = \epsilon_i$$

$$\therefore Z = e^0 + e^{(\mu - \epsilon_i)/k_B T}$$

$$Z = 1 + \exp[(\mu - \epsilon_i)/k_B T]$$

\therefore For state i :

$$\langle N_i \rangle = k_B T \left(\frac{1}{1 + \exp[(\mu - \epsilon_i)/k_B T]} \right)$$

$$\times \frac{\partial}{\partial \mu} \left(1 + \exp[(\mu - \epsilon_i)/k_B T] \right)_{V,T}$$

$$\langle N_i \rangle = \frac{1}{\exp[(\epsilon_i - \mu)/k_B T] + 1}$$

We can drop the index & write as a continuous function.

$$\langle N \rangle = f(\epsilon) = \frac{1}{\exp[(\epsilon - \mu)/k_B T] + 1}$$

↓
Fermi-Dirac distribution



Previously :

$$@ T=0 \quad f(\epsilon) = \begin{cases} 1, \epsilon < \epsilon_F \\ 0, \epsilon > \epsilon_F \end{cases}$$

$$\therefore \lim_{T \rightarrow 0} \mu = \epsilon_F \quad \text{in general} \Rightarrow \mu = \mu(T)$$

$\epsilon_F = \mu$ is not strictly true @ $T > 0$, but it is a good approximation even up to room temperature

$$\text{Recall: chemical potential } \mu_i = \left. \frac{\partial U}{\partial N_i} \right|_{S, V, N_j \neq i}$$

is the energy cost to add/remove a particle to/from the system/ensemble.

What is the energy of this state @ finite T?

Same approach as $T=0$:

Same approach as $T=0$:

$$u = \frac{U}{V} = \frac{2}{V} \sum_k \epsilon_k f(\epsilon_k) \xrightarrow{\text{2 } \bar{e}'\text{'s per k}} 2 \bar{e}'\text{'s per k } (\uparrow \downarrow)$$

↓

energy density

↓

energy of state/k

↓

expected occupancy of state k (FD-dist.)

$$\text{using same trick: } \frac{1}{V} \sum_{\mathbf{k}} F(\mathbf{k}) = \int \frac{d^3 k}{8\pi^3} F(\mathbf{k})$$

$$u = \int \frac{d^3 k}{4\pi^3} \epsilon_k f(\epsilon_k) \quad \rightarrow \text{both } \epsilon_k \text{ & } f(\epsilon_k) \text{ are functions of } k.$$

Note from the dfn of DOS $\Rightarrow D(\epsilon)$, u can be expressed as:

$$u = \int D(\varepsilon) \varepsilon f(\varepsilon) d\varepsilon$$

\rightarrow drop k label
since $\int dE$ now.

$$\text{Likewise: } n = \frac{N}{V} = \frac{2}{V} \sum_k f(\varepsilon_k)$$

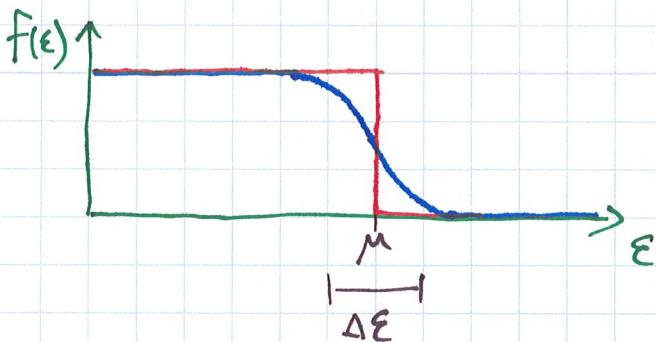
$$n = \int D(\varepsilon) f(\varepsilon) d\varepsilon$$

Note: $D(\varepsilon) \propto \sqrt{\varepsilon}$

These integrals have no closed form (try in your fav. solver).

and, in general, $D(\epsilon)$ may be much more complicated than the $\sqrt{\epsilon}$ form...

BUT



$f(T=0) \neq f(T>0)$ only differ by the small region around $\epsilon = \mu$.

∴ if $T \ll T_F$ ⇒ check this condition for yourself.

We can approximate the integral's by the first couple terms of an expansion about $\epsilon = \mu$:

Sommerfeld Expansion:

for a function of ϵ , $H(\epsilon)$:

$$\boxed{\int_0^{\infty} H(\epsilon) f(\epsilon) d\epsilon = \int_0^{\mu} H(\epsilon) d\epsilon + H'(\mu) \frac{\pi^2}{6} (k_B T)^2 + O(k_B T)^4 + \dots}$$

$$\text{Back to } u \text{ & } n: D(\epsilon) = D(\epsilon_F) \sqrt{\frac{\epsilon}{\epsilon_F}}$$

$$u = \frac{D(\epsilon_F)}{\sqrt{\epsilon_F}} \int \epsilon^{3/2} f(\epsilon) d\epsilon \Rightarrow H(\epsilon) = \epsilon^{3/2}$$

$$n = \frac{D(\epsilon_F)}{\sqrt{\epsilon_F}} \int \epsilon^{1/2} f(\epsilon) d\epsilon \Rightarrow I(\epsilon) = \epsilon^{1/2}$$

Apply S.E.:

$$u \approx \frac{D(\epsilon_F)}{\sqrt{\epsilon_F}} \left[\int_0^{\mu} \epsilon^{3/2} d\epsilon + \frac{3}{2} \mu^{1/2} \left| \frac{\pi^2}{6} (k_B T)^2 \right| \right]$$

$$n \approx \frac{D(\epsilon_F)}{\sqrt{\epsilon_F}} \int_0^{\mu} \epsilon^{1/2} d\epsilon + \frac{1}{2} \mu^{-1/2} \left| \frac{\pi^2}{6} (k_B T)^2 \right|$$

$$\therefore u \approx \frac{D(\epsilon_F)}{\sqrt{\epsilon_F}} \left[\frac{2}{5} \mu^{5/2} + \frac{3}{2} \mu^{1/2} \left| \frac{\pi^2}{6} (k_B T)^2 \right| \right]$$

$$n \approx \frac{D(\epsilon_F)}{\sqrt{\epsilon_F}} \left[\frac{2}{3} \mu^{3/2} + \frac{1}{2} \mu^{-1/2} \left| \frac{\pi^2}{6} (k_B T)^2 \right| \right]$$

Recall: $D(\epsilon_F) = \frac{3}{2} \frac{n}{\epsilon_F}$ and $n \neq n(T)$
(constant volume).

$$\therefore n = \frac{3}{2} \frac{n}{\epsilon_F^{3/2}} \left[\frac{2}{3} M^{3/2} + \frac{1}{2} \mu^{-1/2} \frac{\pi^2}{6} (k_B T)^2 \right]$$

$$\therefore \frac{2}{3} \epsilon_F^{3/2} = \frac{2}{3} \mu^{3/2} \left[1 + \frac{3}{4} \frac{\pi^2}{6} \left(\frac{k_B T}{\mu} \right)^2 \right]$$

or $\mu = \epsilon_F \left[\underbrace{\frac{1}{1 + \frac{3}{4} \frac{\pi^2}{6} \left(\frac{k_B T}{\mu} \right)^2}}_{\text{assume } \left(\frac{k_B T}{\mu} \right)^2 \ll 1} \right]^{2/3}$

and $\left(\frac{k_B T}{\mu} \right)^2 \approx \left(\frac{k_B T}{\epsilon_F} \right)^2$

$$\mu = \epsilon_F \left(1 - \frac{2}{3} \frac{3}{4} \frac{\pi^2}{6} \left(\frac{k_B T}{\epsilon_F} \right)^2 \right)$$

$$\therefore \mu = \epsilon_F + \Delta \mu$$

where $\frac{\Delta \mu}{\epsilon_F} \approx -\frac{\pi^2}{12} \left(\frac{k_B T}{\epsilon_F} \right)^2$

For $T > 0$ but $T \ll T_F$, μ differs by ϵ_F on the order of $\left(\frac{k_B T}{\epsilon_F} \right)^2$ per (fractional)

$$\rightarrow = \left(\frac{T}{T_F} \right)^2$$

Recall: $T_F \approx 10^5 \text{ K}$

room temp: $T \approx 3 \times 10^2 \text{ K}$

\Rightarrow large range of validity.

Likewise, we find:

$$U = U_0 + D(\epsilon_F) \frac{\pi^2}{6} (k_B T)^2 = U_0 - 3\Delta \mu n$$

where U_0 is the g.s. ($T=0$) energy: $U_0 = \frac{3}{5} \epsilon_F n$

Heat capacity

→ Now very simple to calculate:

$$C_V = \frac{1}{V} \left. \frac{\partial U}{\partial T} \right|_V = \left. \frac{\partial U}{\partial T} \right|_V$$

$$\therefore C_V = D(\epsilon_F) \frac{\pi^2}{3} k_B^2 T = \gamma T$$

γ - Sommerfeld constant

$$\boxed{\gamma = D(\epsilon_F) \frac{\pi^2}{3} k_B^2 = \frac{n\pi^2}{2\epsilon_F} k_B^2}$$

Application of v_0 & C_V to calculated parameters:

Mean free path:

$$\ell = v_0 \tau \approx v_F \tau$$

Thermal conductivity:

$$\frac{K}{\text{JT}} = \frac{1}{3} v_0^2 \tau C_V$$

$$\therefore \frac{K}{\text{JT}} = \frac{\pi^2}{3} \left(\frac{k_B}{e} \right)^2 \approx 2.4 \times 10^{-8} \text{ W} \cdot \Omega \cdot \text{K}^{-2}$$

good agreement w exp. (Table 1.6)

Thermopower:

$$\vartheta = -\frac{\pi^2}{6} \frac{k_B}{e} \left(\frac{k_B T}{\epsilon_F} \right) \approx -1.42 \left(\frac{k_B T}{\epsilon_F} \right) \times 10^{-4} \frac{\text{V}}{\text{K}}$$

Much better agreement than Drude

\approx smaller than Drude by $\partial \left(\frac{k_B T}{\epsilon_F} \right) \approx 0.01$

@ room temp.