

Tools for data analyses in Cosmology

- Aula 9 -

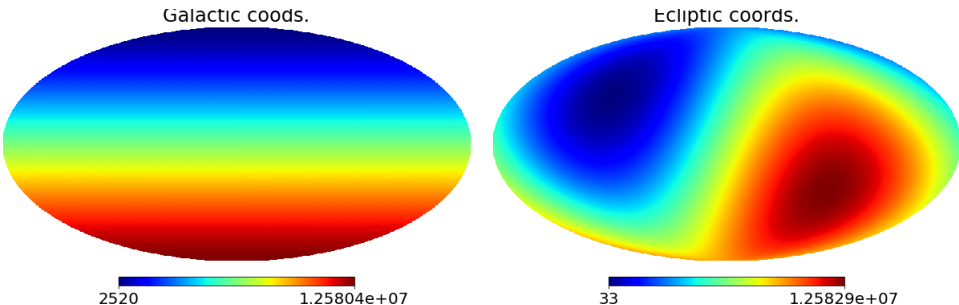
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Observatório Nacional

June 6, 2017

Healpy

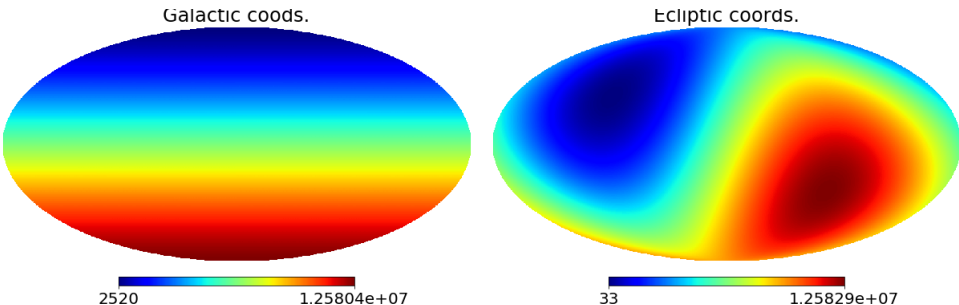
Previously on `healpy` class ... Rotation and geometry functions



How to apply this coordinate transform in a map?

How to perform a rotation instead of a coordinate transform?

Previously on `healpy` class ... Rotation and geometry functions

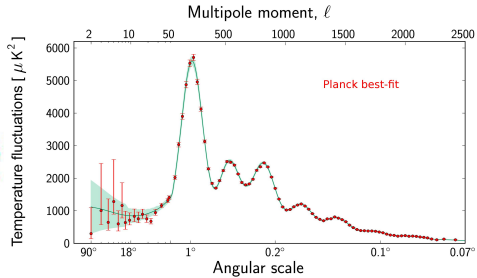
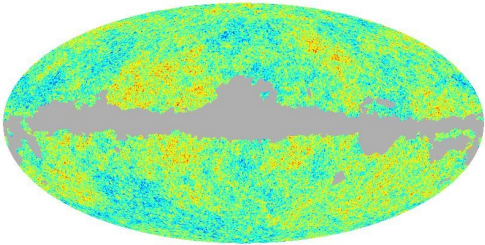


How to apply this coordinate transform in a map?

How to perform a rotation instead of a coordinate transform?

Lets see ...

- Spherical harmonic transforms -



C_ℓ , $a_{\ell m}$, smooth a map, smooth the $a_{\ell m}$, window function, ...

Spherical harmonic transforms

HEALPix conventions

A bandlimited function f on the sphere can be expanded in spherical harmonics, $Y_{\ell m}$, as

$$f(\gamma) = \sum_{\ell=0}^{\ell_{max}} \sum_m a_{\ell m} Y_{\ell m}(\gamma), \quad (4)$$

where γ denotes a unit vector pointing at polar angle $\theta \in [0, \pi]$ and azimuth $\phi \in [0, 2\pi)$. Here we have assumed that there is insignificant signal power in modes with $\ell > \ell_{max}$ and introduce the notation that all sums over m run from $-\ell_{max}$ to ℓ_{max} but all quantities with index ℓm vanish for $m > \ell$. Our conventions for $Y_{\ell m}$ are defined in subsection A.4 below.

The Spherical Harmonics are defined as

$$Y_{\ell m}(\theta, \phi) = \lambda_{\ell m}(\cos \theta) e^{im\phi}$$

where the

$$\begin{aligned} \lambda_{\ell m}(x) &= \sqrt{\frac{2\ell+1}{4\pi} \frac{(\ell-m)!}{(\ell+m)!}} P_{\ell m}(x), \quad \text{for } m \geq 0 \\ \lambda_{\ell m} &= (-1)^m \lambda_{\ell|m|}, \quad \text{for } m < 0, \\ \lambda_{\ell m} &= 0, \quad \text{for } |m| > \ell. \end{aligned}$$

Spherical harmonic transforms

HEALPix conventions

Pixelating $f(\gamma)$ corresponds to sampling it at N_{pix} locations γ_p , $p \in [0, N_{\text{pix}} - 1]$. The sample function values f_p can then be used to estimate $a_{\ell m}$. A straightforward estimator is

$$\hat{a}_{\ell m} = \frac{4\pi}{N_{\text{pix}}} \sum_{p=0}^{N_{\text{pix}}-1} Y_{\ell m}^*(\gamma_p) f(\gamma_p), \quad (5)$$

where the superscript star denotes complex conjugation, and an equal weight was assumed for each pixel. This zeroth order estimator, as well as higher order estimators, are imple-

Spherical harmonic transforms

$$\frac{T(\mathbf{r}_0, \hat{\mathbf{n}}) - T_0(\mathbf{r}_0)}{T_0(\mathbf{r}_0)} = \sum_{l>0} \sum_{m=-l}^{+l} a_{lm}(\mathbf{r}_0) Y_l^m(\hat{\mathbf{n}})$$

$$\langle |a_{lm}(\mathbf{r})|^2 \rangle = \langle a_{lm}(\mathbf{r}) a_{l'm'}^*(\mathbf{r}) \rangle \equiv \delta_{ll'} \delta_{mm'} C_l$$

$$\begin{aligned} C_l &= \langle |a_{lm}(\mathbf{r})|^2 \rangle = && \text{isotropy} \\ &= \left\langle \frac{\sum_m |a_{lm}(\mathbf{r})|^2}{2l+1} \right\rangle = && \text{ergodicity} \\ &= \left\langle \frac{\sum_m |a_{lm}(\mathbf{r})|^2}{2l+1} \right\rangle_{\text{space}} \simeq \\ &\simeq \frac{\sum_m |a_{lm}(\mathbf{r}_0)|^2}{2l+1}. \end{aligned}$$

Spherical harmonic transforms

Cosmic Variance

$$\delta C_\ell = \sqrt{\frac{2}{2\ell + 1}} C_\ell$$

Spherical harmonic transforms

Cosmic Variance

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What if we have a partial sky???

Spherical harmonic transforms

Cosmic Variance

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What if we have a partial sky???

$$\delta C_\ell = \sqrt{\frac{2}{(2\ell + 1)f_{sky}}} C_\ell$$

Spherical harmonic transforms

healpy.sphtfunc.anafast

```
healpy.sphtfunc.anafast(map1, map2=None, nspec=None, lmax=None, mmax=None, iter=3,  
alm=False, pol=True, use_weights=False, datapath=None)
```

Computes the power spectrum of an Healpix map, or the cross-spectrum between two maps if *map2* is given. No removal of monopole or dipole is performed.

$$\text{Map} \Rightarrow C_\ell$$

Spherical harmonic transforms

How to use:

```
In [32]: # Just for temperature:::  
...: mapa = hp.read_map('COM_CMB_IQU-smica_1024_R2.02_full.fits')  
...:  
...: cls_TT = hp.anafast(mapa)  
...:
```

Which is the ℓ_{max} ?

Spherical harmonic transforms

How to use:

```
In [32]: # Just for temperature:::  
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...:
```

Which is the ℓ_{max} ?

$$\ell_{max} = 3072$$

```
In [53]: res = hp.kside2resol(1024, arcmin = True)
```

```
In [54]: print('Resolution=', res, 'arcmin')  
Resolution= 3.43548641182 arcmin
```

```
In [55]: theta = (180./lmax)*60.
```

```
In [56]: print('Theta=', theta, 'arcmin')  
Theta= 3.515625 arcmin
```

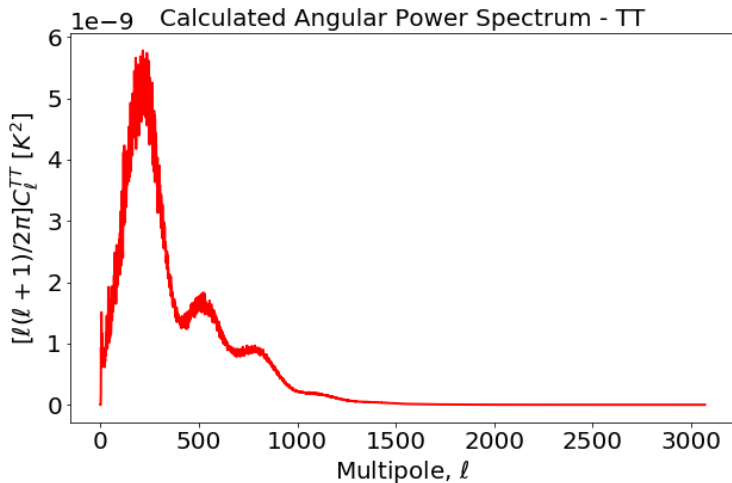
Spherical harmonic transforms

Script to plot the C_ℓ :

```
In [34]: #####
...: # Plot:::
...:
...: # TT:::
...: lmax = len(Cls_TT)
...: ell = np.arange(lmax)
...:
...: Dls_TT = ell*(ell+1)*Cls_TT/(2.*np.pi)
...:
...: pyplot.plot(ell, Dls_TT,linewidth=2.0, color="red") #label="CMB Cls"
...: #
...: #plt.ylim(-300., 1000.)
...: #pyplot.xscale('log')
...: pyplot.title('Calculated Angular Power Spectrum - TT',fontsize=20)
...: pyplot.xlabel('Multipole,  $\ell$ ',fontsize=20)
...: pyplot.ylabel('[$\ell(\ell + 1)/2\pi$] C_$\ell^{TT}$ $K^2$',fontsize=20)
...: #pyplot.legend(loc='best')
...: #
...: fig = pyplot.gcf()
...: fig.set_size_inches(10, 6)
...: pyplot.savefig("Cls.png")
...: pyplot.show()
...:
```

Spherical harmonic transforms

Result:



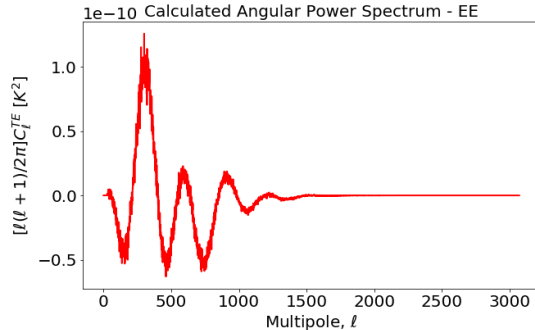
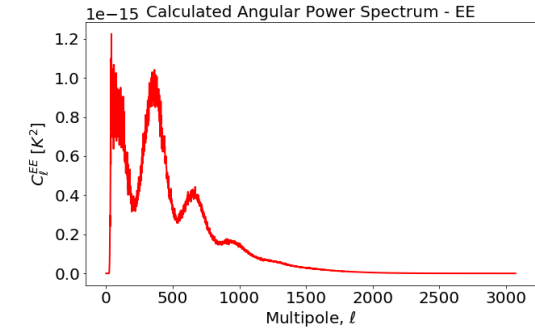
Spherical harmonic transforms

How to use:

```
In [45]: # For temperature and polarization::
...: mapa = hp.read_map('COM_CMB_IQU-smica_1024_R2.02_full.fits', field=[0,1,2])
...:
...: Cls = hp.anafast(mapa)
...:
...: Cls_TT = Cls[0]
...: Cls_EE = Cls[1]
...: Cls_BB = Cls[2]
...: Cls_TE = Cls[3]
...: Cls_EB = Cls[4]
...: Cls_TB = Cls[5]
...:
```

Spherical harmonic transforms

Result:



Spherical harmonic transforms

healpy.fitsfunc.write_cl

```
healpy.fitsfunc.write_cl(filename, cl, dtype=<type 'numpy.float64'>)
```

Writes Cl into an healpix file, as IDL cl2fits.

$C_\ell \Rightarrow$ fits file

Spherical harmonic transforms

How to use:

```
In [105]: # For one Cls array:
...: hp.write_cl('calc_Cls0.fits', Cls[0])

In [106]: # For all:
...: hp.write_cl('calc_Cls1.fits', Cls)
```

Spherical harmonic transforms

healpy.fitsfunc.read_cl

```
healpy.fitsfunc.read_cl(filename, dtype=<type 'numpy.float64'>, h=False)
```

Reads Cl from an healpix file, as IDL fits2cl.

fits file $\Rightarrow C_\ell$

Spherical harmonic transforms

How to use:

```
In [98]: # To open just the FIRST extension:  
...: Cls2 = hp.read_cl('COM_PowerSpect_CMB_R2.02.fits', h=True)
```

```
In [99]: len(Cls2)  
Out[99]: 4
```

```
In [100]: len(Cls2[0])  
Out[100]: 28
```

```
In [101]: # To access the header and chose the extension you want to open:  
...: from astropy.io import fits  
...: hdulist = fits.open('COM_PowerSpect_CMB_R2.02.fits')  
...: hdulist.info()  
...:
```

Filename: COM_PowerSpect_CMB_R2.02.fits

No.	Name	Type	Cards	Dimensions	Format
0	PRIMARY	PrimaryHDU	4	()	
1	TTLOLUNB	BinTableHDU	46	28R x 4C	[I, E, E, E]
2	TELOLUNB	BinTableHDU	46	28R x 4C	[I, E, E, E]
3	EELOLUNB	BinTableHDU	46	28R x 4C	[I, E, E, E]
4	TBLOLUNB	BinTableHDU	46	28R x 4C	[I, E, E, E]
5	EBLOLUNB	BinTableHDU	46	28R x 4C	[I, E, E, E]
6	BBLOLUNB	BinTableHDU	46	28R x 4C	[I, E, E, E]
7	TTHILBIN	BinTableHDU	46	83R x 5C	[E, I, I, E, E]
8	TTHILUNB	BinTableHDU	42	2479R x 3C	[I, E, E]
9	TEHILBIN	BinTableHDU	46	66R x 5C	[E, I, I, E, E]
10	TEHILUNB	BinTableHDU	42	1967R x 3C	[I, E, E]
11	EEHILBIN	BinTableHDU	46	66R x 5C	[E, I, I, E, E]
12	EEHILUNB	BinTableHDU	42	1967R x 3C	[I, E, E]

Spherical harmonic transforms

How to use:

```
In [48]: cls_TTLOLUNB = hp.read_cl(hdulist[1])  
...: cls_TTHILUNB = hp.read_cl(hdulist[8])  
...:
```

```
In [49]: len(cls_TTLOLUNB), len(cls_TTHILUNB)  
Out[49]: (4, 3)
```

```
In [50]: len(cls_TTLOLUNB[0]), len(cls_TTHILUNB[0])  
Out[50]: (28, 2479)
```

```
In [51]: cls_TTLOLUNB[0]  
Out[51]: array([ 2,  3,  4, ..., 27, 28, 29], dtype=int16)
```

```
In [52]: cls_TTHILUNB[0]  
Out[52]: array([ 30,  31,  32, ..., 2506, 2507, 2508], dtype=int16)
```

Spherical harmonic transforms

How to use:

```
In [84]: n1 = len(Cls_TTL0LUNB[0])  
In [85]: n2 = len(Cls_TTHILUNB[0])  
In [86]: n = n1 + n2  
In [87]: ell=np.zeros(n)  
In [88]: Dls=np.zeros(n)  
In [89]: ell[0:n1] = Cls_TTL0LUNB[0]  
In [90]: ell[n1:n] = Cls_TTHILUNB[0]  
In [91]: Dls[0:n1] = Cls_TTL0LUNB[1]  
In [92]: Dls[n1:n] = Cls_TTHILUNB[1]
```

Plot!

Spherical harmonic transforms

Result:

Angular Power Spectrum

