Fitting Baselines to Spectra

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1 Introduction

This document discusses baseline correction methods that can be used with hyperSpec. hyperSpec provides two fitting functions for polynomial baselines, spc.fit.poly and spc.fit.poly.below. Another possibility is spc.rubberband, a "rubberband" method that determines support points by finding the convex hull of each spectrum. The baselines are then piecewise linear or (smoothing) splines through the support points.

Please note that a specialized package for baseline fitting, baseline [1], exists that provides many more methods to fit baselines to spectroscopic data. Using baseline with hyperSpec objects is demonstrated in vignette ("hyperspec").

2 Polynomial Baselines

In contrast to many other programs that provide baseline correction methods, hyperSpec's polynomial baseline functions do least squares fits. However, the baselines can be forced through particular

points, if this behaviour is needed.

The main difference between the two functions is that spc.fit.poly returns a least squares fit through the complete spectrum that is given in *fit.to* whereas spc.fit.poly.below tries to find appropriate spectral regions to fit the baseline to.

2.1 Syntax & parameters

fit.to: hyperSpec object with the spectra whose baselines are to be fitted.

apply.to: hyperSpec object giving the spectral range, on which the baselines should be eval-

uated.

If apply is NULL, a hyperSpec object with the polynomial coefficients is returned

instead of evaluated baselines.

poly.order: polynomial order of the baselines

npts.min: minimal number of data points per spectrum to be used for the fit.

npts.min defaults to the larger of 3 times (poly.order + 1) or $\frac{1}{20th}$ of the number

of data points per spectrum.

If npts.min ≤ poly.order, a warning is issued and npts.min <- poly.order + 1 is

used.

noise: a vector giving the amount of noise, see below.

2.2 General Use

Both functions fit the polynomial to the spectral range given in *hyperSpec* object *fit.to*. If *apply.to* is not NULL, the polynomials are evaluated on the spectral range of *apply.to*. Otherwise, the polynomial coefficients are returned.

Subtracting the baseline is up to the user, it is easily done as *hyperSpec* provides the - (minus) operator.

2.3 Fitting polynomial baselines using least squares

Commonly, baselines are fit using (single) support points that are specified by the user. Also, usually n+1 support point is used for a polynomial of order n. This approach is appropriate for spectra with high signal to noise ratio.

Such a baseline can be obtained by restricting the spectra in *fit.to* to the respective points (see figure 1):

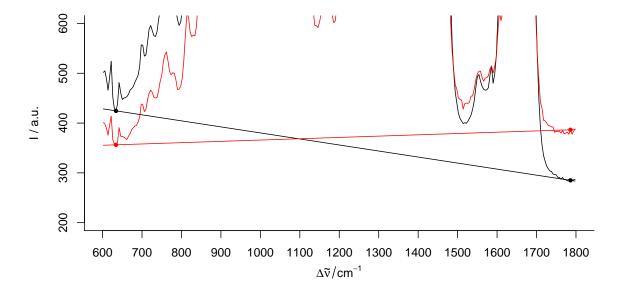


Figure 1 Fitting a linear baseline through two points. If the signal to noise ratio is not ideal, wavelengths that work fine for one spectrum (black) may not be appropriate for another (red).

However, if the signal to noise ratio is not ideal, a polynomial with n+1 supporting points (i.e. with zero degrees of freedom) is subject to a considerable amount of noise. If on the other hand, more data points consisting of baseline only are available, the uncertainty on the polynomial can be reduced by a least squares fit.

Both spc.fit.poly and spc.fit.poly.below therefore provide least squares fitting for the polynomial.

spc.fit.poly fits to the whole spectral region of *fit.to*. Thus, for baseline fitting the spectra need to be cut to appropriate wavelength ranges that do not contain any signal.

In order to speed up calculations, the least squares fit is done by using the Vandermonde matrix and solving the equation system by qr.solve.

This fit is not weighted. A spectral region with many data points therefore has greater influence on the resulting baseline than a region with just a few data points. It is up to the user to decide whether this should be corrected for by selecting appropriate numbers of data points (e.g. by using replicates of the shorter spectral region).

2.4 The mechanism of automatically fitting the baseline in spc.fit.poly.below

spc.fit.poly.below tries to automatically find appropriate spectral regions for baseline fitting. This is done by excluding spectral regions that contain signals from the baseline fitting. The idea is that all data points that lie above a fitted polynomial (initially through the whole spectrum, then through the remaining parts of the spectrum) will be treated as signal and thus be excluded from the baseline fitting.

The supporting points for the baseline polynomials are calculated iteratively:

1. A polynomial of the requested order is fit to the considered spectral range, initially to the whole spectrum given in fit.to

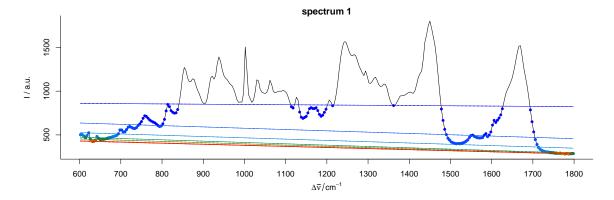


Figure 2 Iterative fitting of the baseline. The dots give the supporting points for the next iteration's baseline, color: iterations.

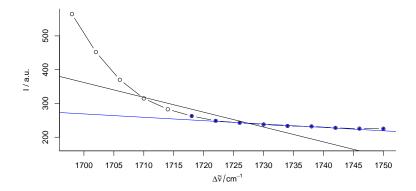


Figure 3 Influence of fit.to on the baseline polynomial. The black baseline is fit to the spectral range $1700 - 1750 \,\mathrm{cm}^{-1}$, the blue to $1720 - 1750 \,\mathrm{cm}^{-1}$ only.

2. Only the parts of the spectrum that lie below this polynomial plus the noise are retained as supporting points for the next iteration.

These two steps are repeated until either

- no further points are excluded, or
- the next polynomial would have less than npts.min supporting points.

The baselines and respective supporting points for each iteration of spc.fit.poly.below (chondro [1], poly.order = 1) are shown in figure 2.

2.5 Specifying the spectral range

It is possible to exclude spectral regions that do not contribute to the baseline from the fitting, while the baseline is used for the whole spectrum. This selection of appropriate spectral regions is essential for spc.fit.poly. But also spc.fit.poly.below can benefit from narrower spectral ranges: the fitting gains speed. The default value for *npts.min* depends on the number of data points per spectrum. Thus one should also consider requiring more support points than the default value suggests.

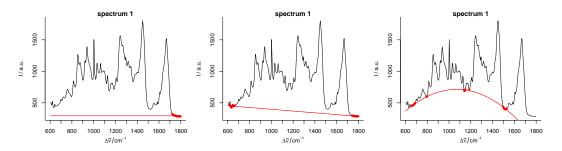


Figure 4 Baseline polynomial fit to the first spectrum of the chondro data set of order 0-2 (left to right). The dots indicate the points used for the fitting of the polynomial.

```
> system.time (spc.fit.poly.below (chondro, NULL, npts.min = 20))
   user   system elapsed
   0.426   0.012   0.437
> system.time (spc.fit.poly.below (chondro [,, c (min ~ 700, 1700 ~ max)], NULL, npts.min = 20))
   user   system elapsed
   0.235   0.000   0.234
```

The choice of the spectral range in fit.to influences the resulting baselines to a certain extent, as becomes clear from figure 3.

2.6 Fitting polynomials of different orders

Figure 4 shows the resulting baseline polynomial of spc.fit.poly.below (chondro [1], poly.order = order) with order = 0 to 3 for the first spectrum of the chondro data set.

2.7 The noise level

Besides defining a minimal number of supporting points, a "noise level" may be given. Consider a spectral range consisting only of noise. The upper part of figure 5 illustrates the problem. As the baseline fitting algorithm cannot distinguish between noise and real bands appearing above the fitted polynomial, the resulting baseline (black) is too low if the noise parameter is not given.

Setting the noise level to 4 (2 standard deviations), the fitting converges immediately with a much better result. The resulting baselines for spc.fit.poly.below (chondro [1], poly.order = 1, noise = 12) of the whole spectrum are shown in the middle and lower part of figure 5

noise may be a single value for all spectra, or a vector with the noise level for each of the spectra separately.

3 Rubberband Method

Particularly Raman spectra often show increasing background towards $\Delta \tilde{v} = 0$. In this case, polynomial baselines often either need high order or residual background is left in the spectra.

In that case, smoothing splines fitted through the supporting points are a good alternative.

For the paracetamol spectrum (fig. 6), a noise level of 300 counts and the equivalent of 20 degrees of freedom work well.

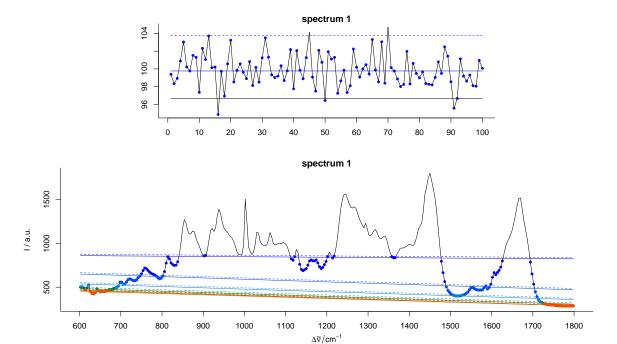


Figure 5 Iterative fitting of the baseline with noise level. Upper part: effects of the noise parameter on the baseline of a spectrum consisting only of noise and offset: without giving noise the resulting baseline (black) is clearly too low. A noise level of 4 results in the blue baseline. The lower part show the baseline fitting with noise level on the complete spectrum. Colour: iterations, dots: supporting points for the respectively next baseline. Dashed: baseline plus noise. All points above this line are excluded from the next iteration.

> bl <- spc.rubberband (paracetamol [,, 175 ~ 1800], noise = 300, df = 20)

However, there is possibly some background left between 1200 and 1750 cm⁻¹ where the original spectrum is slightly concave. This can be corrected by bending the spectrum before applying the rubberband correction (fig. 7):

> bend <- 5e4 * wl.eval (paracetamol [,, 175 ~ 1800], function (x) x^2, normalize.wl=normalize01) > bl <- spc.rubberband (paracetamol [,, 175 ~ 1800] + bend) - bend

References

[1] Kristian Hovde Liland, Trygve Almøy, and Bjørn-Helge Mevik. Optimal choice of baseline correction for multivariate calibration of spectra. *Applied Spectroscopy*, 64:1007–1016, 2010.

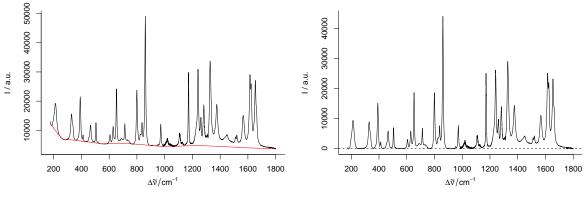
Session Info

R version 3.6.2 (2019-12-12)

Platform: x86_64-pc-linux-gnu (64-bit) Running under: Ubuntu 18.04.4 LTS

Matrix products: default

BLAS: /usr/lib/x86_64-linux-gnu/openblas/libblas.so.3 LAPACK: /usr/lib/x86_64-linux-gnu/libopenblasp-r0.2.20.so



(a) paracetamol with the rubberband fitted baseline.

(b) Corrected spectrum.

Figure 6 Rubberband baselines for the paracetamol spectrum.

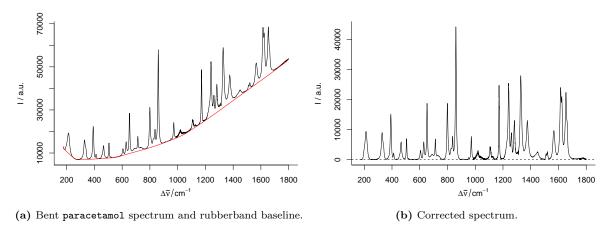


Figure 7 Rubberband baselines for the paracetamol spectrum after bending.

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[10] LC_TELEPHONE=C LC_MEASUREMENT=de_DE.UTF-8 LC_IDENTIFICATION=C attached base packages: [1] grid stats graphics grDevices utils datasets methods base							
other attached packages: [1] hyperSpec_0.99-20200213 xml2_1.2.2 ggplot2_3.2.1 lattice_0.20-38							
loaded via a namespace (and not attached):							
[1] Rcpp_1.0.3 ma	agrittr_1.5	tidyselect_0.	2.5 munsell_0.	5.0			
[5] colorspace_1.4-1 R.	.cache_0.14.0	R6_2.4.1	rlang_0.4.	2			
[9] dplyr_0.8.3 to	ools_3.6.2	gtable_0.3.0	R.oo_1.23.0	0			
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- -	ColorBrewer_1.1-2	•	R.utils_2.9				
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[29] scales_1.1.0 R.	.methodsS3_1.7.1	pkgconfig_2.0	.3				