

The Locus Algorithm

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ABSTRACT

Aims. This Paper describes the design, implementation and operation of a new algorithm, The Locus Algorithm; which enables optimised differential photometry.

Methods. For a given target, The Locus Algorithm identifies the pointing for which the resultant FoV includes the target and the maximum number of similar reference stars available, thus enabling optimised differential photometry of the target.

Results. The application of The Locus Algorithm to a target from the Sloan Digital Sky Survey to provide optimum differential photometry for that target is also described. The algorithm was also used to generate catalogues of pointings to optimise Quasars variability studies and to generate catalogues of optimised pointings in the search for Exoplanets via the transit method.

Key words. Algorithm

1. Introduction

Photometric variability studies involve identifying variations in brightness of a celestial point source over time. Such studies are hampered by the Earth's atmosphere, which causes first order and second order extinction (Milone & Pel 2011; Young et al. 1991). Differential Photometry mitigates the effect of the Earth's atmosphere by comparing the brightness of a target to reference stars in the same Field of View (FoV). Differential photometry can be optimised for the target by choosing a pointing whose Field of View (FoV) includes the target and the maximum number of reference stars of similar magnitude and colour. (Milone & Pel 2011; Young et al. 1991; Howell 2006; Honeycutt 1992)).

The Locus Algorithm enables optimised differential photometry by identifying the pointing for which the resultant FoV includes the target and the best set of similar reference stars available.

2. Conceptual basis to The Locus Algorithm

A locus can be defined around any star such that a FoV centred on any point on the locus will include the star at the edge of the FoV. For fields containing stars close to one another, if one locus intersects with another, they produce Points of Intersection (PoI) as shown in Figure 1.

A FoV centred on any such PoI will include both stars associated with creating it. At Points of Intersection the set of stars that can be included in a FoV changes.

The Locus Algorithm considers candidate reference stars in what is termed a Candidate Zone (CZ) - the zone of sky centred on the target within which a FoV can be selected which includes both the reference star and the target. For Candidate Reference Stars within the CZ, loci are determined, and all relevant PoI are identified. Each PoI is

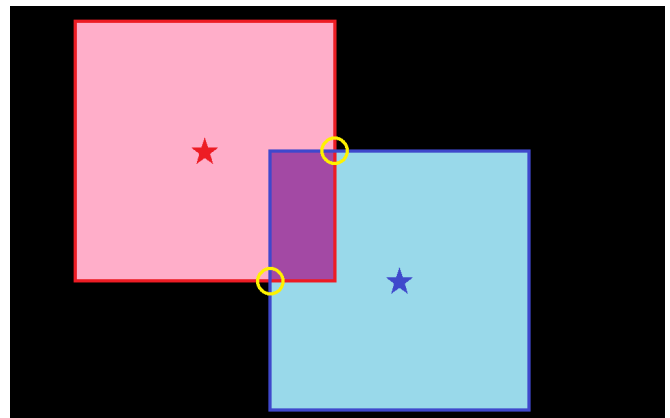


Fig. 1: Diagrammatic representation of two stars with loci (red and blue perimeter lines), which intersect and produce two Points of Intersection (PoI) circled in yellow. Modified from Creaner (2016)

assigned a score derived from the number and similarity of reference stars included in a FoV centred on that PoI. The PoI with the highest score becomes the pointing for the target.

3. Locus Algorithm Design

Based on the conceptual outline above, this section provides a mathematical definition of the Locus Algorithm and an explanation of the terms used in it. Section 4 below describes a worked example of this algorithm applied to a sample star, *SDSS1237680117417115655*.

3.1. Definition of Coordinate System and Locus

For computational efficiency, The Locus Algorithm considers a Field of View to be a rectangular area on the sky orientated such that the edges are aligned with the primary x and y axes of the Cartesian coordinate system. Movement of the field is restricted to x or y translations.

However, the Celestial coordinate system is defined by the Equatorial coordinate system, with coordinates specified by Right Ascension (RA) and Declination (Dec). Because this is a spherical coordinate system, unit angle in RA is foreshortened, with the degree of foreshortening defined in Expression 1

$$\text{angle in RA} = \frac{\text{True Angle}}{\cos(\text{Dec})} \quad (1)$$

Expression 1: Right Ascension foreshortening with Declination

By using this conversion, it is possible to approximate to a high degree of accuracy a Cartesian coordinate system using RA and Dec; with a small FoV of East-West size R and North-South size S about a target located at point RA_t and Dec_t . Expression 2 defines a corrected angular size in RA direction (R')

$$R' = \frac{R}{\cos(Dec_t)} \quad (2)$$

Expression 2: Definition of a corrected angular size along the RA direction (R')

Given these terms, Expression 3 defines the FoV.

$$\begin{aligned} RA_t - \frac{R'}{2} &\leq RA \leq RA_t + \frac{R'}{2} \\ Dec_t - \frac{S}{2} &\leq Dec \leq Dec_t + \frac{S}{2} \end{aligned} \quad (3)$$

Expression 3: Definition of a FoV of size $R \times S$ centred on a target at (RA_t, Dec_t)

This definition is accurate to approximately 1% for a FoV of area $15'$ square outside celestial polar regions as shown in Expression 4.

$$\begin{aligned} \text{Given : } R, S &= 15', |Dec| \leq 66.5^\circ \\ \left| \frac{\frac{R}{\cos(Dec-S)} - \frac{R}{\cos(Dec+S)}}{\frac{R}{\cos(Dec)}} \right| &\leq 0.01 \end{aligned} \quad (4)$$

Expression 4: Evaluation of the accuracy of the R' for areas away from the celestial pole.

As expressed here, the formula does not consider RA “loop around” from 359.99° to 0.00° ; resulting, for example,

in the exclusion of 0.23% of the SDSS catalogue. Planned enhancements to The Locus Algorithm will resolve these shortcomings.

We can therefore define the locus about any star on the sky located at RA_t and Dec_t as the values of Right Ascension and Declination as defined in Equation 2.

3.2. Candidate Zone

A Candidate Zone is defined as a region centred on the target, equal to four times the area of the Field of View, within which any reference star can be included in a Field of View with the target and can therefore be considered as a candidate reference star in identifying the optimum pointing. Conversely, stars outside the candidate zone cannot be included in a Field of View with the target and cannot therefore be considered as candidates reference stars. Hence the Candidate Zone is the maximum region of sky centred on the target from which to choose candidate reference stars when identifying an optimum pointing for a given target. For a target positioned at coordinates RA_t and Dec_t the resulting Candidate Zone is defined by Expression 5.

$$\begin{aligned} RA_t - R' &\leq RA_r \leq RA_t + R' \\ Dec_t - S &\leq Dec_r \leq Dec_t + S \end{aligned} \quad (5)$$

Expression 5: Definition of a Candidate Zone of size $2R \times 2S$ centred on a target with coordinates (RA_t, Dec_t) , in which zone reference stars with coordinates (RA_r, Dec_r) can be found.

3.3. Identification and Filtering of Reference Stars

For each target, a list of candidate reference stars in its Candidate Zone is produced based on the following criteria:

Position: the reference star must be in the Candidate Zone as defined in Expression 5.

Magnitude: the magnitude of the reference star (mag_r) must be within a user-defined limit (Δmag) of the target's magnitude (mag_t) as shown in Expression 6.

Colour: the colour index (e.g. $g - r$) of the reference star (col_r) must match the colour of the target (col_t) to within a user-specified limit (Δcol) as shown in Expression 7.

Resolvability: the reference star must be resolvable, i.e. no other star that would impact a brightness measurements within a user-specified resolution limit.

All stars in the Candidate Zone which pass these initial filters become the list of candidate reference stars for which loci will be identified.

$$mag_t - \Delta mag \leq mag_r \leq mag_t + \Delta mag \quad (6)$$

Expression 6: Definition of the limits of mag difference between the target and references.

$$col_t - \Delta col \leq col_r \leq col_t + \Delta col \quad (7)$$

Expression 7: Definition of the limits of colour difference between the target and references.

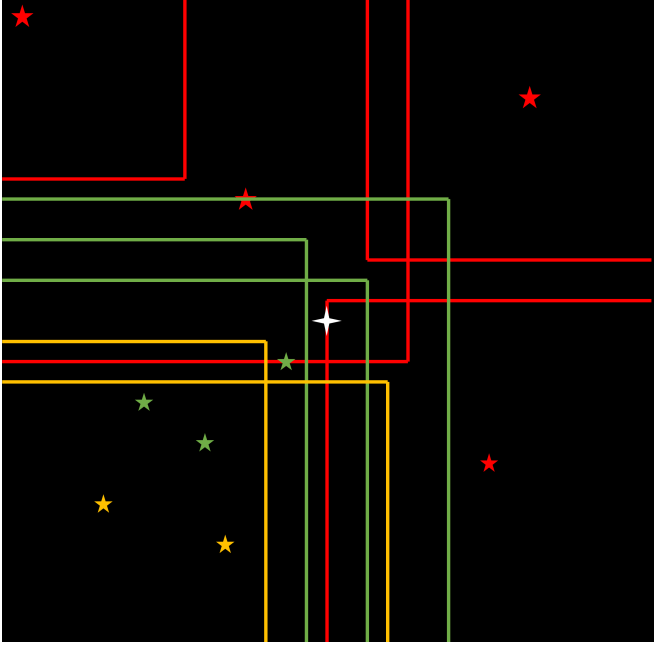


Fig. 2: Each effective locus is defined by assigning a pair of RA and Dec coordinates for a corner point and a pair of lines North or South and East or West from the corner point. In this diagram, each candidate reference star is assigned a colour, and the effective locus that corresponds to it is drawn in the same colour. Modified from Creaner (2016)

3.4. Identifying the Effective Locus for each Candidate Reference Star

The locus associated with each candidate reference star must be identified based on Equation 3. For the purposes of identifying Points of Intersection, only the side surrounding a given candidate reference star closest to the target need be considered. Hence, we can define the effective locus for such a candidate reference star as a single line of constant RA and a single line of constant Dec nearest the target star as shown in Figure 2.

Specifically, the effective locus can be defined as a corner point of the locus and two lines: one of constant RA and the other of constant Dec emanating from the corner point.

Using the Equatorial Coordinate System discussed in Section 3.1, with coordinates of the target specified by (RA_t, Dec_t) and coordinates of the candidate reference star defined by (RA_r, Dec_r) and a size of FoV of horizontal length R and vertical length S, the coordinates of the corner point (RA_c, Dec_c) are defined as shown in Expression 8. The directions $DirRA$ (the direction of the line of constant RA) and $DirDec$ (the direction of the line of constant Dec) of the lines is determined by the RA and Dec of the candidate reference star relative to that of the target are given in Expression 9 and as described below.

$$\begin{aligned} RA_t \leq RA_r &\Rightarrow RA_c = RA_r - \frac{R'}{2} \\ RA_t > RA_r &\Rightarrow RA_c = RA_r + \frac{R'}{2} \\ Dec_t \leq Dec_r &\Rightarrow Dec_c = Dec_r - \frac{S}{2} \\ Dec_t > Dec_r &\Rightarrow Dec_c = Dec_r + \frac{S}{2} \end{aligned} \quad (8)$$

Expression 8: Definition of the corner point (RA_c, Dec_c) of the effective locus for a FoV of size R x S for a candidate reference star at (RA_r, Dec_r) and a target at (RA_t, Dec_t)

- If the RA of the candidate is greater than the target, the line of constant Dec is drawn in the direction of increasing RA
- If the RA of the candidate is less than the target, the line of constant Dec is drawn in the direction of decreasing RA
- If the Dec of the candidate is greater than the target, the line of constant RA is drawn in the direction of increasing Dec
- If the Dec of the candidate is less than the target, the line of constant RA is drawn in the direction of decreasing Dec.

$$\begin{aligned} RA_t \leq RA_r &\Rightarrow DirDec = +ive \\ RA_t > RA_r &\Rightarrow DirDec = -ive \\ Dec_t \leq Dec_r &\Rightarrow DirRA = +ive \\ Dec_t > Dec_r &\Rightarrow DirRA = -ive \end{aligned} \quad (9)$$

Expression 9: Definition the directions $(DirRA, DirDec)$ of the lines from the corner point of that define the effective locus for a FoV of size R x S for a candidate reference star at (RA_c, Dec_c) and given a target at (RA_t, Dec_t) . In current implementations, these values are encoded as a binary switch, with 1 representing increasing (+ive) direction and 0 representing decreasing (-ive) direction.

3.5. Identifying and Scoring Points of Intersection and identifying the pointing.

The points where lines from any two loci are identified. This involves comparing the corner point RA and Dec and direction of lines for one locus with the corner point RA and Dec and direction of lines for a second locus. In total eight variable associated with each two loci are checked:

- For Locus 1: $RA_{c1}, Dec_{c1}, DirRA_1, DirDec_1$
- For Locus 2: $RA_{c2}, Dec_{c2}, DirRA_2, DirDec_2$

Using these parameters, a check as to whether an intersection between the two loci occurs is achieved as follows:

- A line of constant Dec in the positive RA direction from the corner point of locus 1 will intersect with a line of constant RA in the positive Dec direction from the corner point of locus 2 if locus 1 has a lower RA than locus 2 and locus 1 has a higher Dec than locus 2.

- A line of constant RA in the positive Dec direction from the corner point of locus 1 will intersect with a line of constant Dec in the positive RA direction from the corner point of locus 2 if locus 1 has a lower Dec than locus 2 and locus 1 has a higher RA than locus 2.

... and so on. By checking all such possible combinations, all pairs of loci in the field which result in a Point of Intersection are identified and their RA and Dec noted. The above cases are mathematically expressed in Expression 10.

if : $DirDec_1 = +ive, DirRA_2 = +ive$
and : $RA_{r1} < RA_{r2}$
and : $Dec_{r1} > Dec_{r2}$
PoI exists at $RA_p = RA_{c2}, Dec_p = Dec_{c1}$

if : $DirRA_1 = +ive, DirDec_2 = +ive$
and : $RA_{r1} > RA_{r2}$
and : $Dec_{r1} < Dec_{r2}$
PoI exists at $RA_p = RA_{c1}, Dec_p = Dec_{c2}$

...

(10)

Expression 10: Definition of a PoI (RA_p, Dec_p) given several sample cases.

Subsequent to identification, each Point of Intersection is then scored. This is achieved as follows:

- The number of reference stars in the Field of View centred on the Point of Intersection is counted.
- Each reference star is assigned a *rating* value between 0 and 1 based on its similarity in colour to the target.
- The ratings from all counted reference stars in the Field of View are combined into one overall *score* for the field (Figure 3).
- The Point of Intersection with the highest score becomes the pointing for the target (Figure 4).

Scenarios can arise which result in an inability to identify an optimum pointing for a given target for example if there are no, or a maximum of one reference stars in the candidate zone; and if no points of intersection arise – a scenario which can arise if two (or more) reference fall in one quadrant of the candidate zone resulting in concentric loci, or where reference stars are too far apart in different quadrants of the candidate zone in order for their loci to intersect. All four of these scenarios are considered in practical implementations of the locus algorithm aimed at identifying the optimum pointings for a set of targets in a catalogue or list of targets.

In summary, the Locus Algorithm successfully identifies the RA and Dec coordinates of the optimum pointing for a given target, where optimum means a field of view with the maximum number of reference stars which are similar in magnitude and colour to the target.

4. Example Implementation of the Locus Algorithm

To illustrate the operations of the Locus Algorithm, a worked example is given here. The process described here in producing an optimal pointing for a given star follows the same

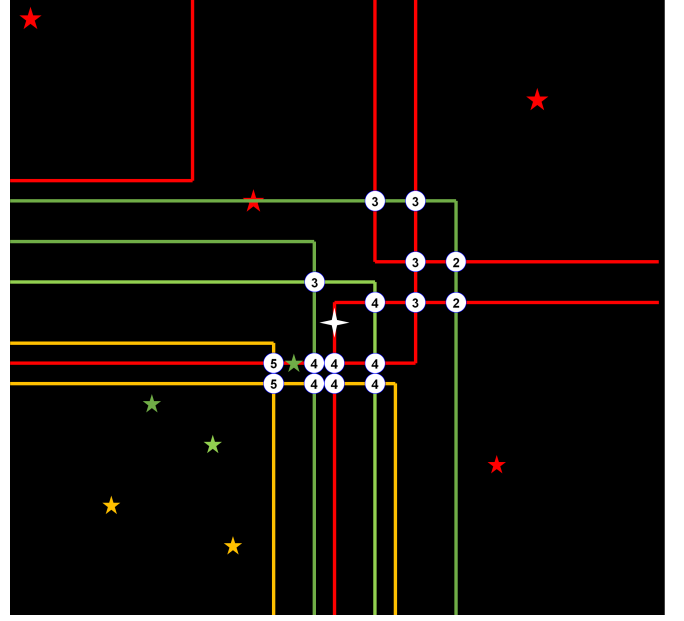


Fig. 3: Points of Intersection (PoI), and their associated score. In this diagram each star has a rating of 1, hence the score associated with each PoI is equal to the number of reference stars within a FoV centred at that PoI. Modified from Creaner (2016)

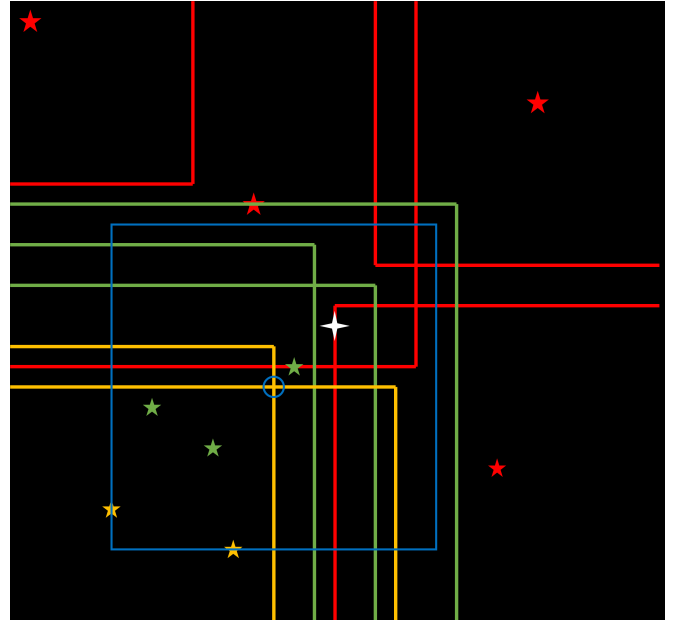


Fig. 4: Locus Algorithm. Target: white star. Pointing & FoV: blue. Reference stars and their loci: Fully in the FoV: greens. On the edge of the FoV: yellows. Outside FoV: reds. Modified from Creaner (2016)

sequence of steps described in the first part of this paper. The process is implemented in the R programming language and is geared for reproducible research. The code is available on Creaner et al. (2007) It can be trivially adapted for different target stars and telescope parameters.

4.1. Target

The star *SDSS1237680117417115655*, henceforth called the target, ($RA = 346.6500^\circ$, $DEC = -5.0393^\circ$) is used as the example. This star, in the constellation Aquarius, has SDSS magnitudes as given on Table 1

Band	SDSS Magnitude
u	17.20
g	15.38
r	14.65
i	14.40
z	14.28

Table 1: SDSS *ugriz* Magnitudes for *SDSS1237680117417115655*, the target used in the sample implementation of the Locus Algorithm

The observational parameters are taken from the telescope at Blackrock Castle Observatory. This telescope has parameters given on Table 2

Parameters	Values
Field of View in degrees	0.1667
Resolution Limit in degrees	0.0030
Dynamic Range in magnitudes	2.0000
Colour Match Limit	0.1000

Table 2: Observational parameters for the example use of the Locus Algorithm

4.2. Candidate Zone

The size of the FoV when corrected for shortening by declination is given in expression 11, by substituting into Expression 2 above.

$$\begin{aligned}
 R' &= \frac{R}{\cos(Dec_c)} \\
 R' &= \frac{0.1667^\circ}{\cos(-5.0393^\circ)} \\
 R' &= 0.16731^\circ
 \end{aligned}
 \tag{11}$$

Expression 11: Definition of R' for the target.

The locus of the target is given in Expression 12 by substituting into Expression 3.

$$\begin{aligned}
 346.5664^\circ &\leq RA \leq 346.7337^\circ \\
 -5.1226^\circ &\leq Dec \leq -4.9560^\circ
 \end{aligned}
 \tag{12}$$

Expression 12: Definition of the locus centred on the target.

The candidate zone as defined above is the area of sky within which reference stars can possibly be included in the same field of view as the target. This is four times the size of the FoV and is given in Expression 13 by substituting into Expression 5 above,

$$\begin{aligned}
 346.4827^\circ &\leq RA \leq 346.8173^\circ \\
 -5.2060^\circ &\leq Dec \leq -4.8726^\circ
 \end{aligned}
 \tag{13}$$

Expression 13: Definition of the Candidate Zone (CZ) centred on the target.

4.3. Identification and Filtering of Reference Stars

The potential reference stars are selected as follows:

- Position: Within the Candidate Zone defined in Expression 13, SDSS records 1345 separate objects with clean photometry (Aguado et al. 2019). These are downloaded by an SQL query run on the CAS database, release DR15
- Magnitude: as shown in Expression 6 the reference star must be within the dynamic range, 2, of the target’s magnitude of 14.648, i.e. $12.648 \leq r \leq 16.648$. This leaves 41 potential references.
- Colour: as shown in Expression 7 the reference star must match the colour of the target to within a user-specified limit of 0.1 magnitudes. In this case this means $0.634 \leq (g - r) \leq 0.834$ and $0.149 \leq (r - i) \leq 0.349$. This leaves 15 potential references.
- Resolvability: the reference star must be resolvable, i.e. no other star that would impact a brightness measurements within a user-specified resolution limit, in this case $11''$ (0.003°). Any object this close to a potential reference star and with an *r*-band magnitude which is 5 magnitudes greater than the potential reference or brighter will pollute the light from the potential reference star. This leaves 14 potential references.

These numbers are presented in Table 3, accessed using Lang & the CRAN team (2019).

filters	numbers
Position, in Field of View	1345
Correct Magnitude	41
Correct Colour	15
Resolvable	14
In Final Field of View	7

Table 3: A summary of the number of candidate reference stars remaining at each stage.

Table 4 gives the 15 stars in the candidate zone. 14 of these are candidate reference stars. 7 of those 14 are used in the calculation of score for the pointing that is ultimately selected. Also shown in this table is the target itself, highlighted in green.

Table 4 also includes ratings for each star. This gives a measure of how close, spectrally, each reference star is to the target. This rating hasn’t been optimised and is the subject of further research by our group, but for the purpose of this work was calculated as shown in Expression 14:

Where Δr stand for the difference between the *r*-band magnitude for the target and the reference star and Δcol_{max} is the colour match limit. Similarly for Δg and Δi .

For example, the star *SDSS1237680117417050120* (Star #8 on Table 4) has $g_r = 16.498$, $r_r = 15.771$, and

ref.	objID	ra	dec	u	g	r	i	z	ratings
1	1237680117417115692	346.7237	-5.0473	18.36	16.58	15.84	15.57	15.46	0.7345
2	1237680065885241391	346.4833	-4.9462	18.97	17.14	16.42	16.16	15.99	0.8080
3	1237680117417115683	346.7128	-5.0499	17.59	15.78	15.11	14.87	14.80	0.3610
4	1237680117417050117	346.5378	-5.0185	18.40	16.52	15.77	15.53	15.39	0.7006
5	1237680117417181202	346.8104	-4.9749	17.46	15.62	14.85	14.58	14.44	0.5353
6	1237680065348435996	346.7072	-5.1987	16.70	14.70	13.90	13.68	13.52	0.3220
7	1237680065348501526	346.8037	-5.2019	15.88	14.15	13.39	13.15	12.99	0.6440
8	1237680117417050120	346.5626	-5.1530	18.46	16.50	15.77	15.53	15.40	0.8298
10	1237680117417115655	346.6500	-5.0393	17.20	15.38	14.65	14.40	14.28	1.0000
11	1237680117417115762	346.6755	-5.1195	18.92	17.02	16.28	15.97	15.85	0.3804
12	1237680117417050133	346.5944	-5.1611	16.70	14.83	14.07	13.89	13.65	0.2410
13	1237680065885241371	346.5598	-4.9312	17.68	15.93	15.23	15.01	14.90	0.4846
14	1237680065885306903	346.6695	-4.9149	17.72	15.65	14.90	14.69	14.45	0.4292
15	1237680117417115701	346.7433	-5.0054	18.46	16.41	15.58	15.32	15.16	0.0683

Table 4: Details of the candidate reference stars from SDSS. The seven stars highlighted in red are found in the final field of view. The target star itself is shown highlighted in green. Then numbers given in the first column “ref” are used to identify the candidate reference stars in the worked examples shown in Subsection 4.4. Data accessed using RCurl by Lang & the CRAN team (2019), table generated with knitr by Xie (2019, 2015, 2014) and kableExtra by Zhu (2019)

$$\begin{aligned}
 \text{rating} &= \left(1 - \left| \frac{\Delta g - \Delta r}{\Delta \text{col}_{\max}} \right| \right) \times \left(1 - \left| \frac{\Delta r - \Delta i}{\Delta \text{col}_{\max}} \right| \right) \quad (14) \quad \text{if :} \quad RA_r < RA_t \\
 &\implies RA_c = RA_r + \frac{R'}{2} \\
 \text{Expression 14: Definition of the rating of a reference star} \quad &\text{because :} \quad 346.5626^\circ < 346.6500^\circ - +0.0837^\circ \\
 &\implies RA_{c,8} = 346.5626^\circ \\
 &\quad RA_{c,8} = 346.6463^\circ \\
 i_r = 15.533. \text{ This compares to the target magnitudes of} \quad &\text{and} \\
 g_t = 15.382, r_t = 14.648, \text{ and } i_t = 14.399. \quad &\text{if :} \quad Dec_r < Dec_t \\
 \text{Substituting these into Expression 14 gives the calculation demonstrated in Expression 15.} \quad &\implies Dec_c = Dec_r + \frac{S}{2} \\
 &\text{because :} \quad -5.1530^\circ < -5.0393^\circ \\
 &\implies Dec_{c,8} = -5.1530^\circ + 0.0833^\circ \\
 &\quad Dec_{c,8} = -5.0697^\circ \quad (16) \\
 \Delta g = 16.498 - 15.382 = 1.116 \\
 \Delta r = 15.771 - 14.648 = 1.122 \\
 \Delta i = 15.533 - 14.399 = 1.134 \\
 \therefore \\
 \text{rating} &= \left(1 - \left| \frac{1.116 - 1.122}{0.1} \right| \right) \times \left(1 - \left| \frac{1.122 - 1.134}{0.1} \right| \right) \\
 \text{rating} &= 0.93 \times 0.89 \\
 \text{rating} &= 0.83 \quad (15) \\
 \text{Expression 15: Definition of the rating of Star \#8} \quad &\text{if : } RA_r < RA_t \implies DirDec = -ive \\
 &\text{because : } 346.5626^\circ < 346.6500^\circ \implies DirDec_8 = -ive \\
 &\text{and} \\
 &\text{if : } Dec_r < Dec_t \implies DirRA = -ive \\
 &\text{because : } -5.1530^\circ < -5.0393^\circ \implies DirRA_8 = -ive \quad (17)
 \end{aligned}$$

4.4. Identifying the Effective Locus for each Candidate Reference Star

Expression 8 defines how the corner points for a given reference can be calculated. Substituting in for Star #8, the resulting corner point is calculated as shown in Expression 16

From Expression 9, the direction the locus must be drawn from the corner point to generate the locus can be determined. Applying this to Star #8 gives the values shown in Expression 17

Expression 17: Definition of the directions of the lines drawn from the corner-point of the effective locus for Star #8

Repeating this calculation for *SDSS1237680065348435996* (Star #6 from Table 4) gives the values shown on Table 5. These two reference stars are used to calculate Points of Intersection (PoI) in Subsection 4.5

quantity	Ref #6	Ref #8
RA_c	346.6235°	346.6463°
Dec_c	-5.1153°	-5.0697°
$DirRA$	<i>-ive</i>	<i>-ive</i>
$DirDec$	<i>+ive</i>	<i>-ive</i>

Table 5: A summary of the defining parameters of the effective loci for example Stars #6 and #8

4.5. Identifying and Scoring Points of Intersection and identifying the pointing.

We can compare these two effective loci to see if they intersect to form a valid PoI. Expression 10 shows some examples of how a valid PoI can be determined. Applying this process to Stars #6 and #8 gives Expression 18:

$$DirRA_8 = +ive, DirDec_6 = -ive$$

$$RA_{r,8} < RA_{r,6} \text{ and } Dec_{r,8} > Dec_{r,6}$$

\Rightarrow

PoI exist at :

$$(RA, Dec) = (RA_{c,6}, Dec_{c,8})$$

$$(RA, Dec) = (346.6463^\circ, -5.1153^\circ) \quad (18)$$

Expression 18: Definition of the PoI between the effective Loci of Stars #6 and #8

From Expression 3, the FoV centred on the PoI ($346.6463^\circ, -5.1153^\circ$) can be calculated to be as shown in Expression 19

$$RA_c - \frac{R'}{2} \leq RA \leq RA_c + \frac{R'}{2}$$

$$\Rightarrow 346.6463^\circ - 0.0837^\circ \leq RA \leq 346.6463^\circ + 0.0837^\circ$$

$$346.5626^\circ \leq RA \leq 346.7299^\circ$$

and :

$$Dec_c - \frac{S}{2} \leq Dec \leq Dec_c + \frac{S}{2}$$

$$-5.1153^\circ - 0.0833^\circ \leq Dec \leq -5.1153^\circ + 0.0833^\circ$$

$$-5.1987^\circ \leq Dec \leq -5.0320^\circ \quad (19)$$

Expression 19: Definition of the FoV centred on the PoI between the effective Loci of Stars #6 and #8

The locus associated with each of these 14 potential references is identified and the Points of Intersection associated with these loci are calculated. This leads to the situation shown in Figure 5.

All these 38 points of intersection are checked in turn as potential pointings. For each one, a field of view is constructed and all the potential reference stars within each one are identified. This is used to calculate a score for each point of intersection. For the star, *SDSS1237680117417115655*, an optimised pointing was thus discovered with $RA = 346.6463^\circ, Dec = -5.1153^\circ$. The field

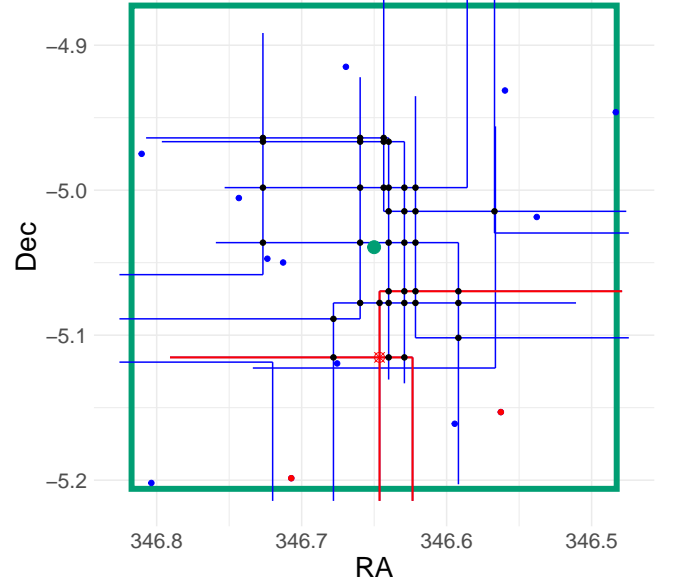


Fig. 5: A plot of the Candidate Zone (green box), the target (green dot), the candidate reference stars (blue dots), the effective loci for each candidate reference star (blue lines) the Points of Intersection between those loci (black dots). Highlighted in red are the candidate reference stars, loci and PoI used to identify the final pointing. Plot generated using tidyverse by Wickham (2017)

of view centred on this pointing included both the target and 7 reference stars. This pointing is calculated to have a score of 3.87 and is illustrated in Figure 6.

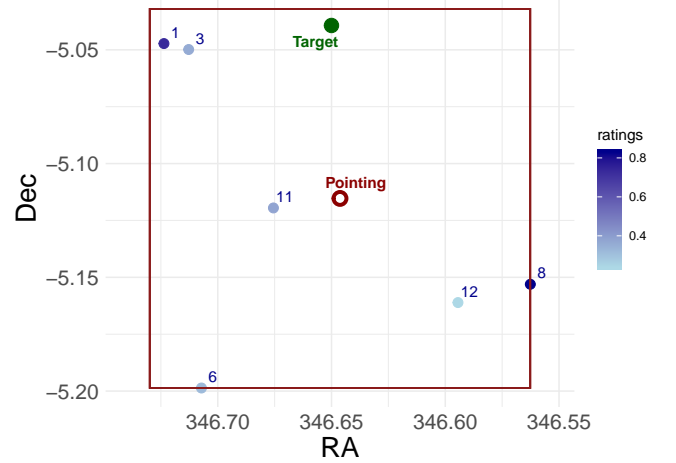


Fig. 6: A plot illustrating the final pointing. On this plot, the target is shown in green, the pointing and the resulting final Field of View are shown in red, and the final reference stars are shown in shades of blue that vary from light to dark based on the rating of the reference star. Plot generated using tidyverse by Wickham (2017)

5. Applications of the Algorithm

The pointings generated in this algorithm are the optimum pointing for each target given the observational parameters and scoring system used. This is of use to any observer

aiming to carry out differential photometry observations as it automates the selection of pointing and identification of reference stars. Two main use-cases are envisaged for this system: targeted use (where an observer wishes to identify the optimum pointing for a pre-determined target or set of targets) and catalogue generation (where many targets are submitted to the system and a set of scores and pointings are generated for each).

5.1. Targeted Use

This scenario considers an observer who wishes to perform differential photometry observations of a pre-determined target(s). The observer may use the algorithm to identify the optimum pointing for their target(s). As illustrated for the target above, this pointing may be offset from the target but will always include the target and the reference stars with the maximum combined rating. Software to identify this optimum pointing and select the reference stars centred on that pointing is available online at Creaner et al. (2007). A web interface to this software is planned.

5.2. Catalogue Generation

By submitting many targets at once, the optimum pointing for each can be determined and scores calculated for each. These are output together and can be collated into a catalogue as demonstrated in Creaner et al. (2019c). The catalogues can then be used by an observer to select targets suitable for their needs by filtering the catalogue. For example, by using the scores for each target, targets with a higher score (and thus a better set of reference stars) can be selected for observation over targets with worse scores. The catalogue production software generates lists of targets, their pointings and the scores associated with those pointings. Users who select targets from the catalogues can then follow up their choice by identifying the set of reference stars with the SQL queries found at Creaner et al. (2007).

6. Conclusions

This paper presents the locus algorithm, a novel system for the identification of optimal pointings for differential photometry given a set of parameters of the planned observation provided by the user. The algorithm is presented in two stages. In the first stage, the concept of the algorithm is laid out and the steps of the algorithm are defined. In the second stage, a fully-worked example is shown applying the algorithm to the star *SDSS1237680117417115655* for a 10' FoV. The premise of the algorithm is that a locus can be defined about any target or reference star, upon which a Field of View (FoV) can be centred and include that object at the edge of the FoV. At the Points of Intersection (PoI) between these loci, the set of targets which can be included in a FoV changes if the FoV moves in any direction. Therefore, these PoI are the essential points to compare when determining the maximum number and quality of reference stars which can be included in a FoV of a given size.

6.1. Summary of Algorithm

The algorithmic description can be summarised in the following nine steps:

Identify the target: The target is identified by its coordinates in the RA/Dec coordinate system.

Provide observational parameters: The algorithm requires a FoV size, magnitude and colour difference limits and a resolution parameter to be provided.

Define a Candidate Zone: identify all stars which could be included in a FoV with the target by translating the position of the FoV in accordance with Expression (ref to CZ), which defines the Candidate Zone (CZ).

Filter Candidate References: for each star in the CZ, apply the filtering criteria magnitude, colour and resolution (Expression (ref to filters) to identify the candidate reference stars

Calculate Loci: The effective locus around each candidate reference star (i.e. the path upon which the FoV may be centred and include the candidate and the target) is calculated as per Expressions (ref to expressions).

Identify Points of Intersection: The points where the effective loci for two candidate reference stars intersect with one another are identified by combining their coordinates as shown in Expression (ref to PoI expression).

Calculate Rating: for each candidate reference star, a rating is calculated to indicate how closely its colour matches that of the target.

Calculate Score: For each PoI, a score is calculated by combining the ratings for each candidate reference star which can be included in a FoV centred on that target.

Output Optimum Pointing: The PoI with the best score is then selected as the optimised pointing for that target

6.2. Summary of Demonstration

For demonstration purposes, this algorithm has been applied to *SDSS1237680117417115655* (referred to as the target). The same nine steps are applied as outlined below

Identify the target: The target is found at ($RA = 346.6500^\circ$, $Dec = -5.0393^\circ$)

Provide observational parameters: The FoV for the demonstration is 10' (0.1667°). The magnitude difference limit is ± 2.0 mag. The colour difference limit is ± 0.1 mag. The resolution selected is 11" (0.003°)

Define a Candidate Zone: The CZ around the target is defined by ($346.4827^\circ \leq RA \leq 346.8173^\circ$, $-5.2060^\circ \leq Dec \leq -4.8726^\circ$) and contains 1345 objects

Filter Candidate References: A candidate reference must meet the following filtering criteria ($12.648 \leq r \leq 16.648$), ($0.634 \leq g - r \leq 0.834$) and ($0.149 \leq r - i \leq 0.349$) and have no star within 0.003° of it. Applying the filtering criteria leaves 14 Candidate References

Calculate Loci: For each of the 14 candidate reference stars, a locus is calculated, for example for Star #8, the locus is defined by $RA_c = 346.6463^\circ$ $Dec_c = 5.0697^\circ$, $DirRA = -ive$ and $DirDec = -ive$.

Identify Points of Intersection: Given 14 candidate reference stars and their corresponding Loci, there are 364 possible combinations that could be a PoI, which are generated by combining the coordinates of the cornerpoints of the Loci in pairs and checking the directions of the lines from each to determine whether a PoI actually exists. PoI actually exist in 38 cases. For example, a PoI exists between the loci for Stars #6 and #8 at ($RA = 346.6463^\circ$, $Dec = -5.1153^\circ$).

Calculate Rating: for each of the candidate reference stars a rating is calculated based on the difference in colour between the target and the reference star, for example, the rating of Star #8 is 0.83.

Calculate Score: Score is calculated by identifying the candidate reference stars which can be included in an FoV centred on each PoI and summing their ratings. For example, for the PoI between stars #6 and #8 above, seven candidate reference stars can be included in the FoV and their ratings sum to 3.87.

Output Optimum Pointing: By comparing the scores for each PoI, the one with the highest score can be determined to be the pointing used in this example at ($RA = 346.6463^\circ$, $Dec = -5.1153^\circ$) and $Score = 3.87$.

7. Further developments

A software system has been developed to implement this algorithm and is available from Creaner et al. (2007). A paper describing the software has been submitted at Creaner et al. (2019a). Scaling that software to allow for large-scale use required the use of a Grid Computing solution, described at Creaner et al. (2019b). By supplying a set of quasars from the 4th Quasar Catalogue of the Sloan Digital Sky Survey (SDSS) as targets and the remainder of SDSS as potential reference stars, it was possible to generate a catalogue of optimised pointings for 26779 quasars as per Creaner et al. (2019c,f). Using all of the stars in SDSS as targets allowed for the optimum pointing to be determined for all such stars. These pointings are of use, for example, in the search for extrasolar planets by the transit method, where high-precision differential photometry is required (Creaner et al. 2019d,e).

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