Motion of Rigid Bodies

Rigid bodies are system of particles having constant interparticle distances

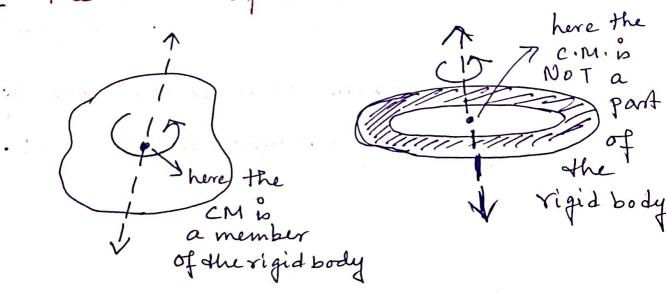
The degrees of freedom of a rigid body is = 6 as we discussed in the lecture

(all the particles move)

+ 3 rotational d.o.f.

(at least one point or a set of points passing through the rigid body do not move)

Mose that the fixed point may not be a real member of the system of particles but it represents somehow the 'society' of particles e.g. the centre of mass.

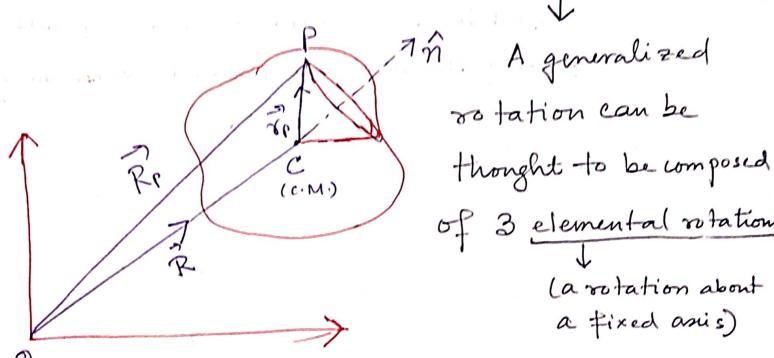


Charle's Theorem: Any general motion of a rigid body can be decomposed as the motion of a the following types!

(i) Translational motion of a representative point of the rigid body (e.g. Centre of man

(ii) Rotational motion w. r.t. an asis

passing through that fixed point. (can
with time)



Here, we assume one such axis (amentioned in (ii)) passing through the C.M. C.

Let us first take the situation where the C.M. is clamped (cannot translate) or we put an observer at C.M. with a coordinate system.

Ju both cases, the body will experience only rotational motion (or undergo)

I Let us do it analytically:

$$\overrightarrow{R_p} = \overrightarrow{r_p} + \overrightarrow{R} \Rightarrow \overrightarrow{R_p} = \overrightarrow{r_p} + \overrightarrow{R}$$

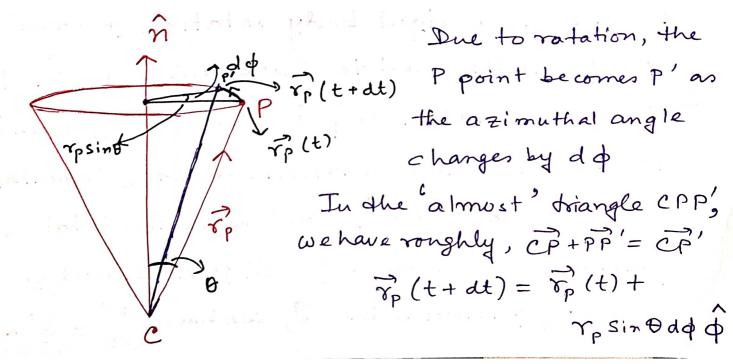
$$\overrightarrow{rotation}$$

$$\overrightarrow{rotation}$$

$$\overrightarrow{about}$$

$$\overrightarrow{c.M.}$$

First we take the example of an elemental rotation about \hat{n} by an angle ϕ



$$\Rightarrow \overrightarrow{r_p}(t+dt) - \overrightarrow{s_p}(t) = r_p \sin\theta d\phi \hat{\phi}$$

$$\Rightarrow \frac{d\hat{\gamma_p}}{dt} = \gamma_p \sin\theta \frac{d\phi}{dt} \hat{\phi} = (\frac{d\phi}{dt} \hat{\gamma}) \times \gamma_p \hat{\gamma_p}$$

$$\Rightarrow \frac{d \vec{r_p}}{dt} = \vec{\omega_p} \times \vec{r_p}$$
one can write

that

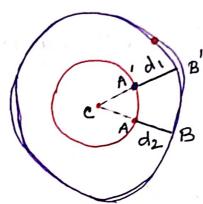
as $\angle \hat{r_p}, \hat{n} = \theta$

(to designate the angular velocity corresponding to the elemental rotation by the angle ϕ)

If the point undergoes three elemental rotations (sufficient to superesent a generalized rotation) by angles ϕ , θ and ψ , the total angular velocity of the combined so tation will be $\vec{\omega} = \vec{\omega}_{\phi} + \vec{\omega}_{\theta} + \vec{\omega}_{\phi}$

Since, in a rigid body rotation (elemental), see all particles need to have uniform $\vec{\omega}$, for Next page the first rotation $\vec{\omega}_{\phi}$ is same for all points. For the other two rotations about (elemental), similarly $\vec{\omega}_{\theta}$ and $\vec{\omega}_{\psi}$ and hence the total $\vec{\omega}$ will also be the same for all particles and for an arbitrary position vector $\vec{\tau}$, we have $\frac{d\vec{\tau}}{dt} = \vec{\omega} \times \vec{\tau}$

1 To understand why wo is uniform for all points, we need to take a top view of the plane of rotation:



Let us assume the and making circles

and body in rotating

and making circles

by in the plane of rotation

with C as the centre.

Rigid body points A and B, after rotation become A' and B'.

Rigidity constraint: AB = A'B'

=) This is only possible when both A and Bare rotated by the same angle do and hence have equal $\frac{d\phi}{dt} = \omega \phi$.

1 Coming back to our initial discussion; We have $\frac{d\vec{7}}{dt} = \vec{\omega} \times \vec{7}$ (with $\vec{\omega}$ being independent of $\vec{7}$)

and the total angular momentum wirt the

centre of mass
$$\vec{L} = \iint \vec{r} \times \frac{d\vec{r}}{dt} d^3\vec{r}$$

$$= \iint \vec{r} \times (\vec{\omega} \times \vec{r}) d^3\vec{r}$$

$$6) = \int \int \left[r^2 \vec{\omega} - (\vec{r} \cdot \vec{\omega}) \vec{r} \right] d^3 \vec{r}$$

Now $\gamma^2 \vec{\omega} - (\vec{\gamma}, \vec{\omega}) \vec{\gamma}$ is a vector. We want to express it Something times $\vec{\omega}$ (If we can do that then since, linear momentum = $\vec{p} = m\vec{v}$, we can say angular momentum = $\vec{l} = \vec{\omega}$ $\vec{\omega}$ measure of inertia for the rotational motion)

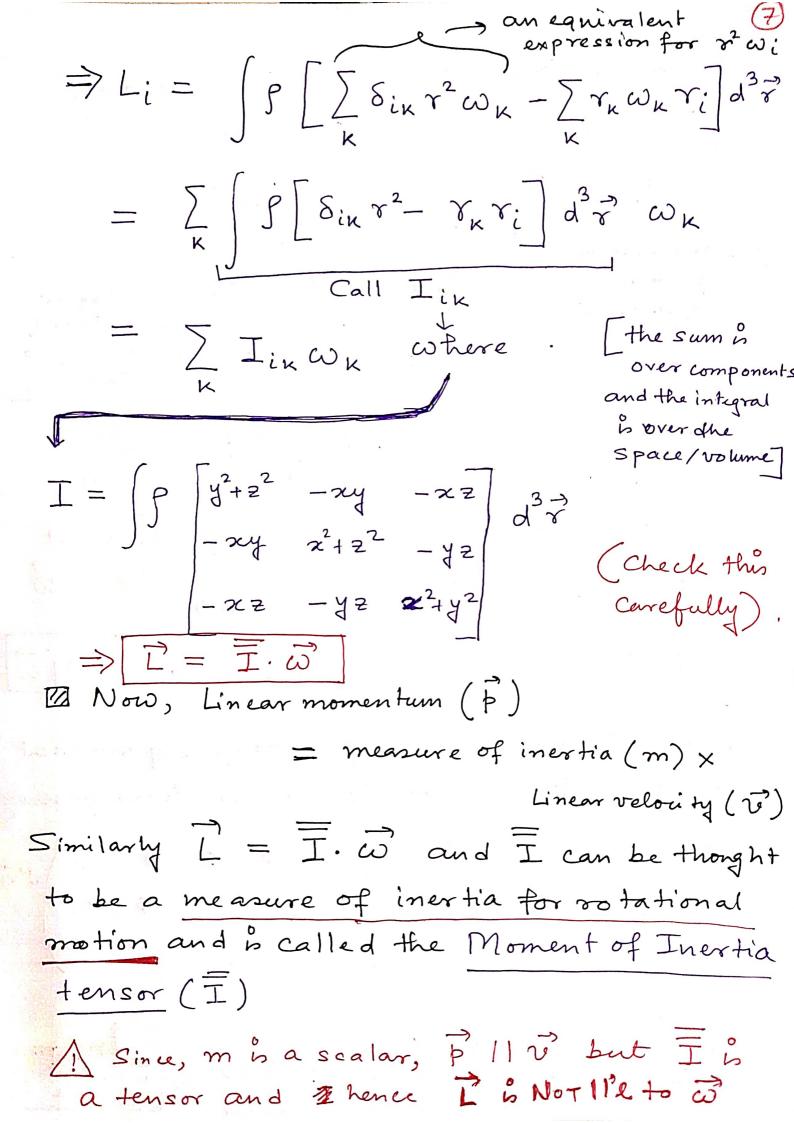
Clearly one can understand [?] cannot be a scalar. However, just like a vector can be obtained by multiplying a vector by a scalar, a vector can be obtained by contracting (3x1) a tensor (3x3) by a vector (3x1) to its right.

Maybe [?] is a tensor of rank 2.

(a planar matrix).

Itom to find that?

For that, we first write, the ith component of \vec{L} $L_i = \int g \left[\vec{r}^2 \omega_i - \sum \vec{r}_k \omega_k \vec{r}_i \right] d^3 \vec{r}$



De For an elemental rotation about z axis, evidenty $\vec{\omega} = \omega_2 \hat{z}$ and then (off diagonal terms) T_{xy} T_{yy} T_{yy} T_{yy} T_{yy} T_{yy} T_{yy} T_{yz} T_{yz} T_{zz} T_{zz} $\vec{L} = I_{xz} \omega_z \hat{z} + I_{yz} \omega_z \hat{y} + I_{zz} \omega_z^z$ For a point mass moving in x-y plane, we can effectively set S+111. Z=0 and hence Izz=0=Iyz [is not 11'2 to W but $I_{22} = \iint (x^2 + y^2) d^3 \vec{r} \neq 0$ and therefore $\vec{L} = \vec{J}_{zz} \, \omega_z \, \hat{z}$ and $\vec{L} \, l \, l \, \vec{\omega}$ In 3d space motion of a perfectly spherical body about 3 mutually I'r direction passing through its centre, $I = \int \int \int \Phi \circ \circ \circ \int_{3^{2}r} \Phi \circ \circ \circ \circ = \int_{3^{2}r} \Phi \circ \circ \circ = \int_{3^{2}r} \Phi \circ \circ \circ \circ = \int_{3^{2}r} \Phi \circ \circ \circ \circ = \int_{3^{2}r} \Phi \circ \circ = \int_{$ i.e. $\overline{I} = \begin{pmatrix} I & O & O \\ O & I & O \\ O & O & I \end{pmatrix}$ and then, $\vec{L} = \begin{pmatrix} \vec{I} & \vec{O} & \vec{O} \\ \vec{O} & \vec{I} & \vec{O} \end{pmatrix} \begin{pmatrix} \vec{\omega}_{\chi} \\ \vec{\omega}_{\gamma} \\ \vec{\omega}_{z} \end{pmatrix} = \vec{I} \vec{\omega} \Rightarrow \vec{L} |\vec{I}| \vec{\omega}$ Scalar

Let us find the moment of inertia tensor for a uniform sphere about an axis passing through its centre.

Of course, here only (any) so tational motions about an axis (changing with time) possing through the centre of the rigid body is concerned. Also note that the centre of the sphere is the C.M. of the sphere.

(How would you show that?)

Calculate the mass moments of all the particles of the sphere (rigid) w.r.t. the centre O and we obtain

 $\int \vec{y} \, d^3\vec{r} = \int (r\hat{r}) \, r^2 \sin\theta \, dr \, d\theta \, d\phi$ Uniform
Sphere $= 4\pi \, f \, r^3 \, dr \, \hat{r}$ Contains $\theta \, \theta \, \phi$ $= \int \vec{r} \, dr \, (\sin\theta \cos\phi \, \hat{x} + \sin\theta \sin\phi \, \hat{y} + \cos\theta \, \hat{z}) \, d\theta \, d\theta \, d\phi$

$$= S \left[r^{3} dr \sin^{2}\theta d\theta \cos\phi d\phi \right] \hat{x}$$

$$+ S \left[r^{3} dr \sin^{2}\theta d\theta . \sin\phi d\phi \right] \hat{y}$$

$$+ S \left[r^{3} dr \cos\theta \sin\theta d\theta . d\phi \right] \hat{z}$$

Taking limit $0 \le r \le R$, $0 \le \theta \le \Pi$ and $0 \le \phi \le 2\pi \rightarrow \omega e$ get $\int g \vec{r} d^3 \vec{r} = 0$.

Now, we have to calculate; the sphere is moment of inertia tensorw.r.t an man of it.

are axis passing through the centre (also the C.M).

From symmetry, Ixx = Iyy = Izz.

$$I_{zz} = \int g(x^{2}+y^{2}) r^{2} \sin \theta \, dr \, d\theta \, d\phi$$

$$= \int g(r^{2}-z^{2}) r^{2} \sin \theta \, dr \, d\theta \, d\phi$$

$$= g \int r^{4} \, dr \sin \theta \, d\phi \, d\theta + g \int r^{4} \, dr \cos^{2}\theta \, d(\omega s \theta) \, d\phi$$

$$= g \cdot 4\pi \frac{R^{5}}{5} + g \cdot 2\pi \frac{R^{5}}{5} \left[-\frac{2}{3} \right] = \frac{4\pi g R^{5}}{5} \left[1 - \frac{1}{3} \right]$$

$$(\omega th M = \frac{4}{3}\pi R^{3}g) = \frac{2}{5}MR^{2}$$

(11)

and Similarly we can write, $I_{2x} = J_{yy} = \frac{2}{5}MR^2$ and $I_{2y} = -\int g xy r^2 sin0 dr d0 d\phi$ $= -\int g r(\cos\phi sin\theta) \cdot r(\sin\phi sin\theta)$ $r^2 sin\theta dr d0 d\phi$ = 0 [check it assuming $r^2 sin\theta dr d\theta d\phi$]

and similarly for Iyz and Izz and hence the moment of inertia tensor is given by

$$\overline{T} = \begin{pmatrix} \frac{2}{5}MR^2 & 0 & 0 \\ 0 & \frac{2}{5}MR^2 & 0 \\ 0 & 0 & \frac{2}{5}MR^2 \end{pmatrix}.$$

De Please note that I is not in general a diagonal matrix.

Only when we consider the elemental rotations about particular axes, then I becomes diagonal and the three particular directions / axes are known as Principal axes or principal directions

(12)

Relation between the total time derivatives of a vector in the C.M. frame and the body frame which is rotating w.r.t C.M:

Ket us now assume two coordinate systems. For one, which we call an unprimed system, the observer is situated at C and if C is clamped, then that coordinate system is just equivalent to the lab frame. An arbitrary vector in this system can be written as:

(Cartesian)

 $\vec{A} = Ax\hat{i} + Ay\hat{j} + Az\hat{k}$ where the unit vectors are fixed in time.

For the second system, the observer is at P and Rence is rotating with wabout c. Ket us assume at t=0, both the coordinate Aystems Rave unit vectors i, i and i & & i', i' and i' being parallel to each other.

i, j' and k' being paralles to each other.

(Not needed

But with time, i', j', k' can be

changes w.r.t. unprimed system and

are fixed w.r.t othe primed system. We also

have, $\overrightarrow{A} = \overrightarrow{A}_{\times}\hat{i}' + \overrightarrow{A}_{Y}\hat{j}' + \overrightarrow{A}_{Z}\hat{k}'$

As it is showed earlier that any vector M' which rotates about C and is traving constant magnitude,

 $\frac{d\vec{M}}{dt} = \vec{\Omega} \times \vec{M}, \text{ we can show}$ $\frac{d\hat{i}' = \vec{\Omega} \times \hat{i}' \text{ and the same for } \hat{i}' \text{ and } \hat{k}'$

This result is essential in deriving the Euler's Egnations for rigid bodies.

[all the unit vectors in primed coordinates are nothing both Vectors with constant magnitude unity and rotating about an axis through C]

We there fore have, for a vector ?

$$\frac{d\vec{C}}{dt} = \left(\frac{d\vec{C}_{x}}{dt}\hat{i}' + \frac{d\vec{C}_{y}'}{dt}\hat{i}' + \frac{d\vec{C}_{z}'}{dt}\hat{k}'\right)$$

clamped C.M. + $\left(\frac{C_{z}'}{dt} + \frac{d\hat{i}'}{dt} + \frac{C_{z}'}{dt} + \frac{d\hat{k}'}{dt}\right)$

= $\frac{d\vec{C}}{dt}$ | + $\left(\frac{C_{z}'}{dt} + \frac{d\hat{k}'}{dt} + \frac{C_{z}'}{dt} + \frac{d\hat{k}'}{dt}\right)$

= $\frac{d\vec{C}}{dt}$ | rotating frame

(as in this frame

 \hat{i}', \hat{j}' and \hat{k}' are

tixed in time) + $\frac{d\vec{C}}{dt}$ | rotating