

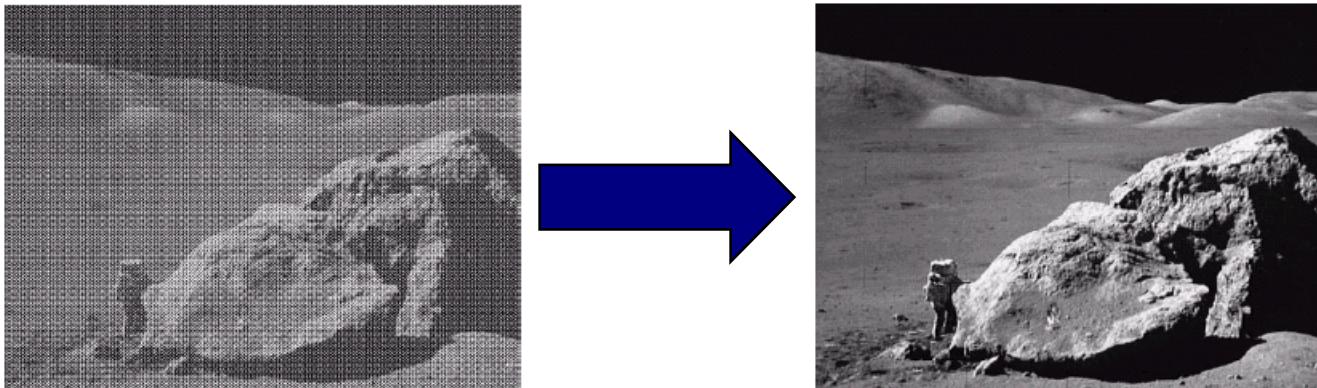
# Image & Video Processing

Image Restoration:  
Noise Removal

# What is Image Restoration?

Image restoration attempts to restore images that have been degraded

- Identify the degradation process and attempt to remove it in order to go back to the “original”
- Similar to image enhancement, but more objective

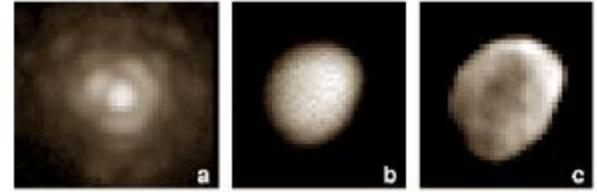


# Applications

Started from the 1950s

- Scientific explorations
- Legal investigations
- Film making and archival
- Image and video (de-)coding
- ...
- Consumer photography

Example of image restoration  
Asteroid Vesta



# Degradation causes

## Degradation examples:



- original



- optical blur



- motion blur



- spatial quantization (discrete pixels)

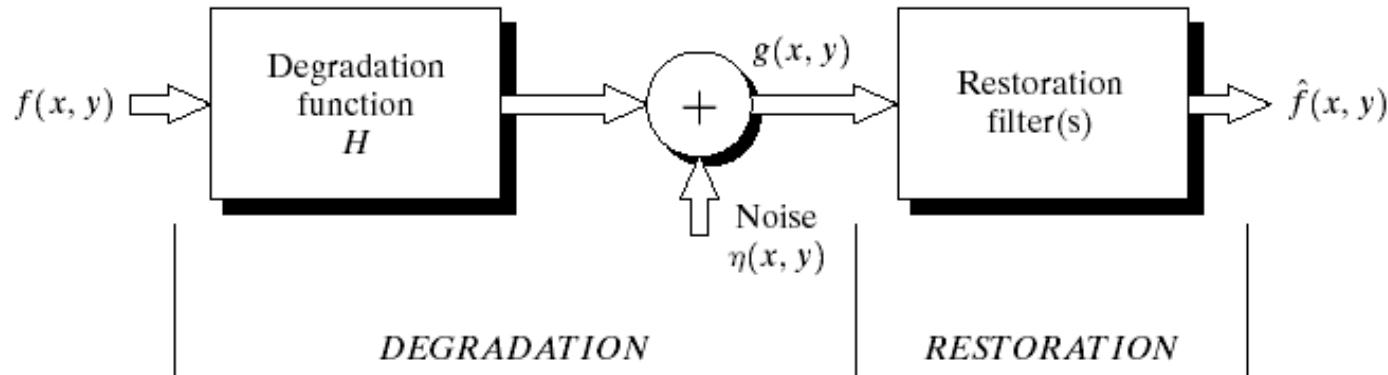


- additive intensity noise

### Causes:

- Camera: translation, shake, out-of-focus ...
- Environment: scattered and reflected light
- Device noise: CCD/CMOS sensor and circuitry
- Quantization noise
- Transmission error

# A Model for Image Distortion/Restoration



**FIGURE 5.1** A model of the image degradation/ restoration process.

$$g(x, y) = H[f(x, y)] + \eta(x, y)$$

where  $f(x, y)$  is the original image pixel,  $H$  is the degradation function,  $\eta(x, y)$  is the noise term and  $g(x, y)$  is the resulting noisy pixel

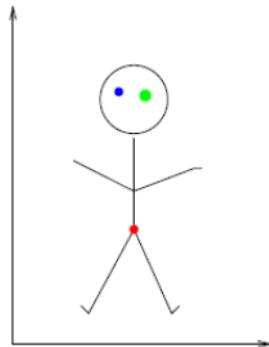
# Assumptions for the Distortion Model

- Noise
  - Independent of spatial location
    - Exception: periodic noise ..
  - Uncorrelated with image
- Degradation function  $H$ 
  - Linear
  - Position-invariant

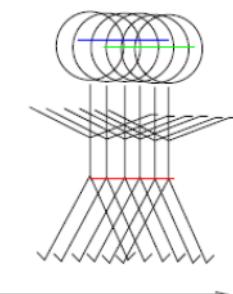
$$g(x, y) = h(x, y) * F(x, y) + \eta(x, y)$$

**Step #1: image degraded only by noise.**

Input Image

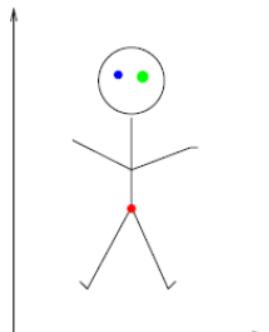


Blurred by Camera linear motion

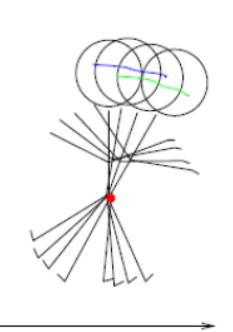


SPACE-INVARIANT RESPONSE - each point on image gives same response just shifted in position.

Input Image



Blurred by Camera Rotation

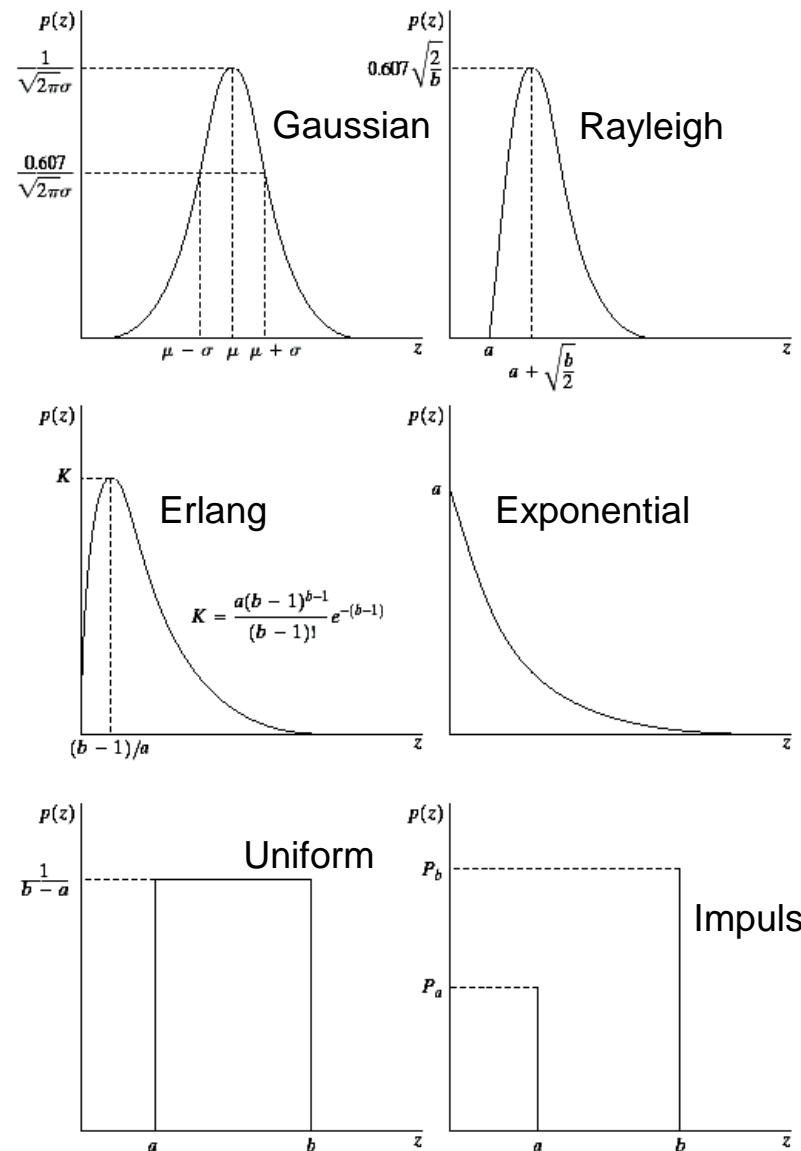


SPACE-VARIANT RESPONSE - each point on image gives a different response

# Noise Models

There are many different models for the image noise term  $\eta(x, y)$ :

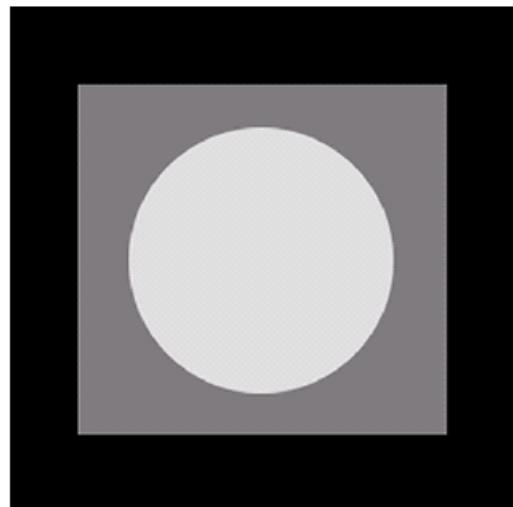
- Gaussian
  - Most common model
- Rayleigh
- Erlang
- Exponential
- Uniform
- Impulse
  - *Salt and pepper* noise



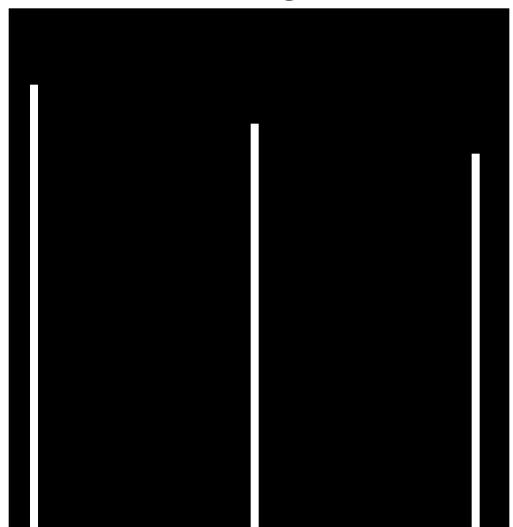
# Noise Example

The test pattern to the right is ideal for demonstrating the addition of noise

The following slides will show the result of adding noise based on various models to this image



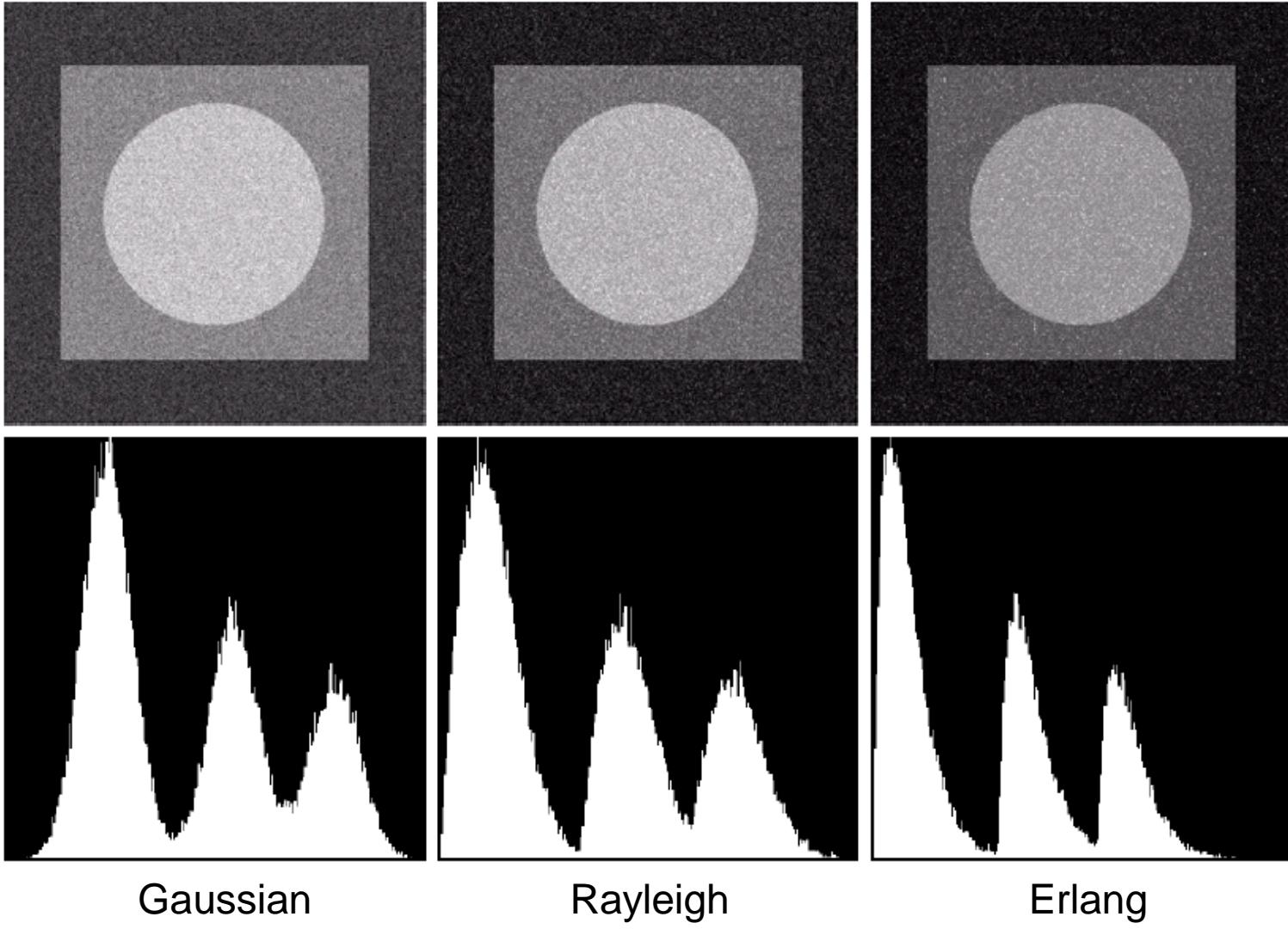
Image



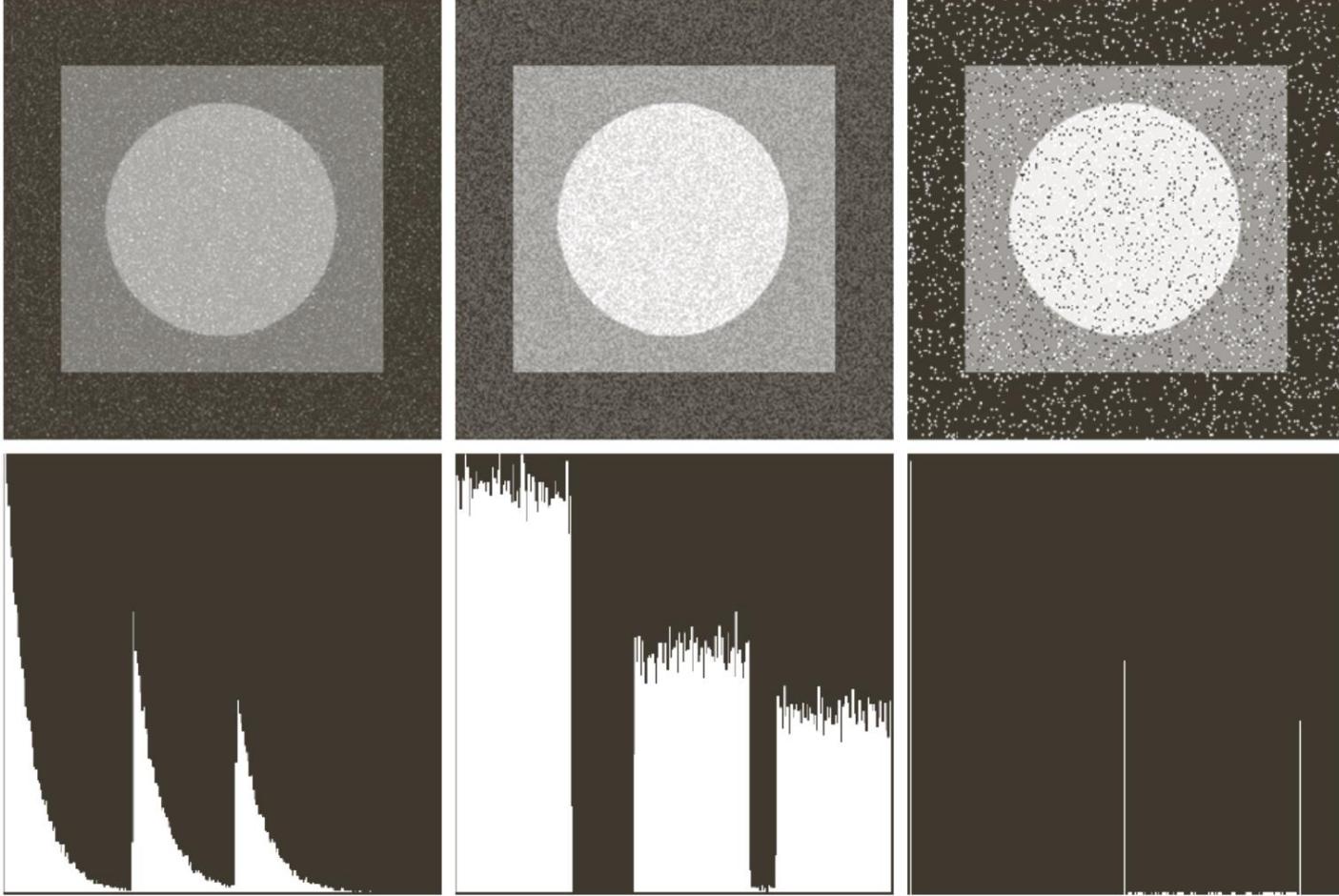
Histogram



# Noise Example (cont...)



# Noise Example (cont...)

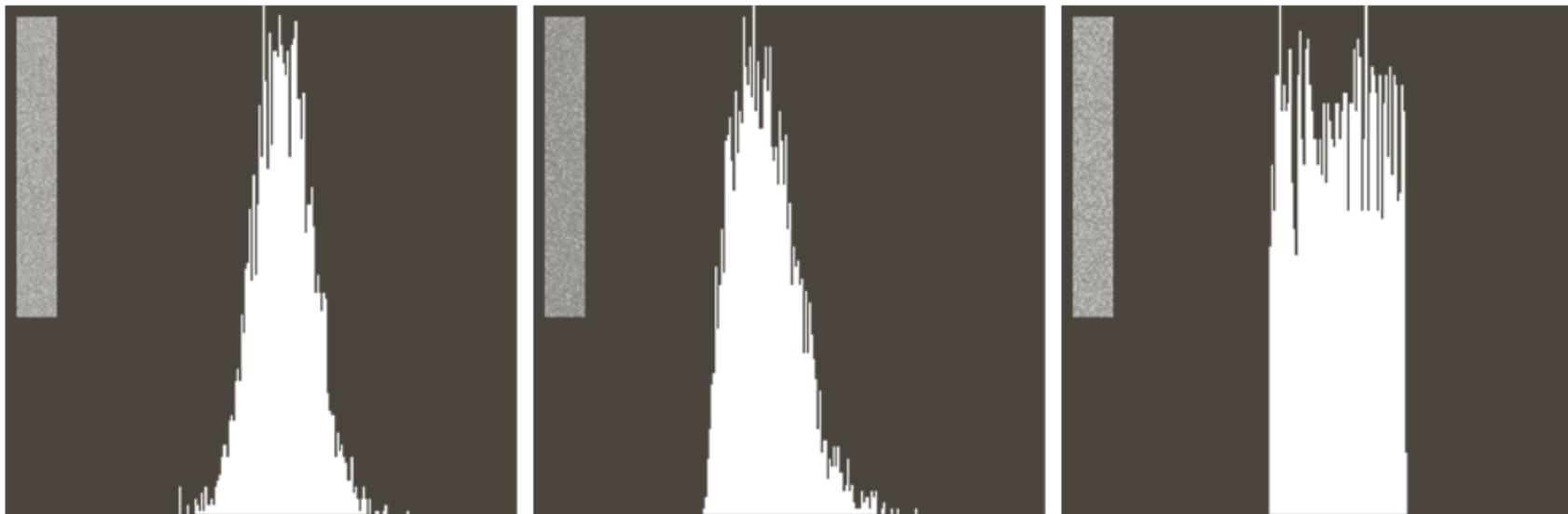


Exponential

Uniform

Impulse

# Estimation of noise parameters



a | b | c

**FIGURE 5.6** Histograms computed using small strips (shown as inserts) from (a) the Gaussian, (b) the Rayleigh, and (c) the uniform noisy images in Fig. 5.4.

# Filtering to Remove Noise

We can use spatial filters of different kinds to remove different kinds of noise

The *arithmetic mean* filter is a very simple one and is calculated as follows:

$$\hat{f}(x, y) = \frac{1}{mn} \sum_{(s,t) \in S_{xy}} g(s, t)$$

|     |     |     |
|-----|-----|-----|
| 1/9 | 1/9 | 1/9 |
| 1/9 | 1/9 | 1/9 |
| 1/9 | 1/9 | 1/9 |

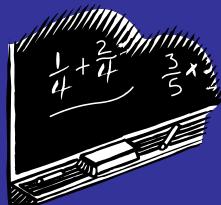
This is implemented as the simple smoothing filter

Blurs the image to remove noise

# Other Means

There are different kinds of mean filters all of which exhibit slightly different behaviour:

- Geometric Mean
- Harmonic Mean
- Contraharmonic Mean



# Other Means (cont...)

There are other variants on the mean which can give different performance

**Geometric Mean:**

$$\hat{f}(x, y) = \left[ \prod_{(s,t) \in S_{xy}} g(s, t) \right]^{\frac{1}{mn}}$$

Achieves similar smoothing to the arithmetic mean, but tends to lose less image detail

# Noise Removal Examples

Original  
Image

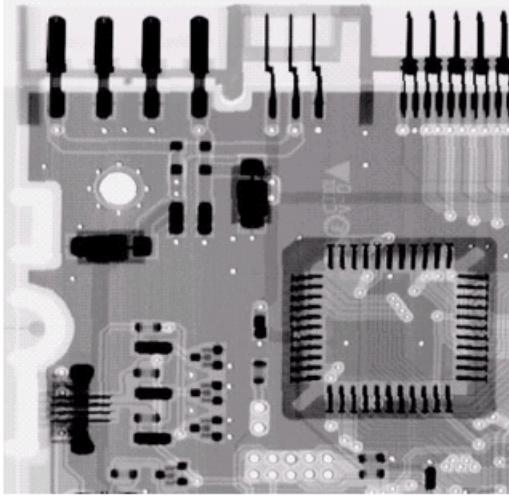
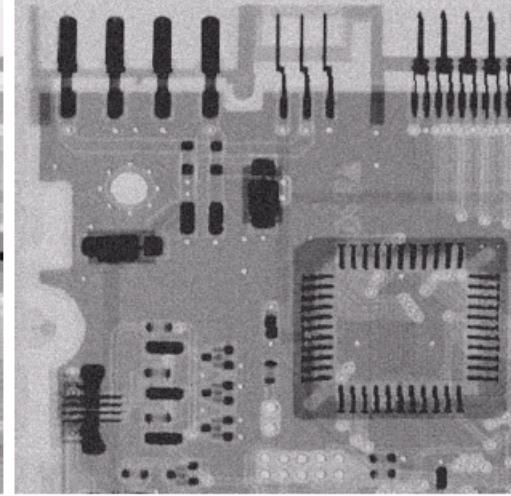
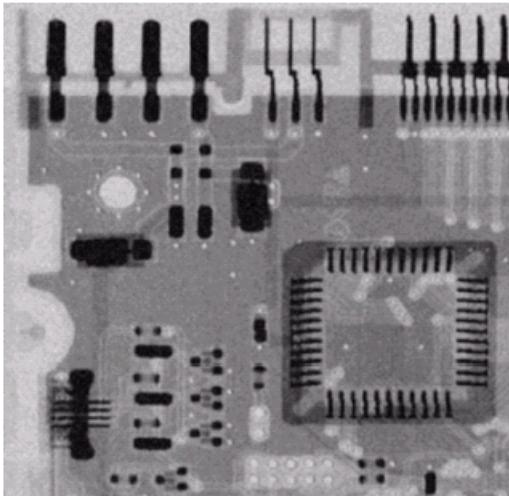


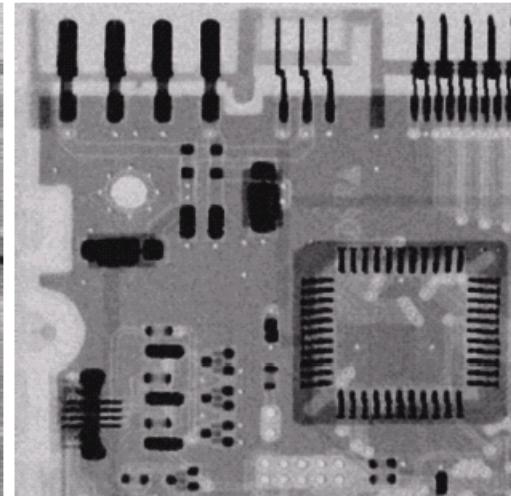
Image  
Corrupted  
By Gaussian  
Noise

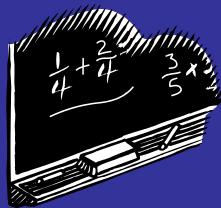


After A 3\*3  
Arithmetic  
Mean Filter



After A 3\*3  
Geometric  
Mean Filter





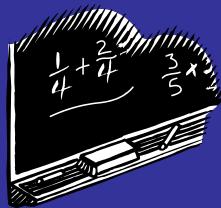
# Other Means (cont...)

## Harmonic Mean:

$$\hat{f}(x, y) = \frac{mn}{\sum_{(s,t) \in S_{xy}} \frac{1}{g(s,t)}}$$

Works well for salt noise, but fails for pepper noise

Also does well for other kinds of noise such as Gaussian noise



# Other Means (cont...)

## Contraharmonic Mean:

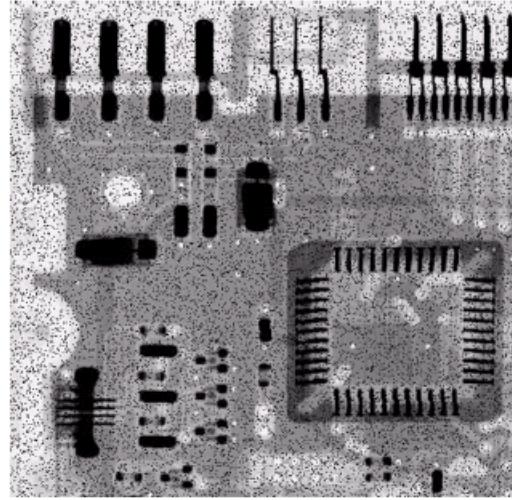
$$\hat{f}(x, y) = \frac{\sum_{(s,t) \in S_{xy}} g(s, t)^{Q+1}}{\sum_{(s,t) \in S_{xy}} g(s, t)^Q}$$

$Q$  is the *order* of the filter and adjusting its value changes the filter's behaviour

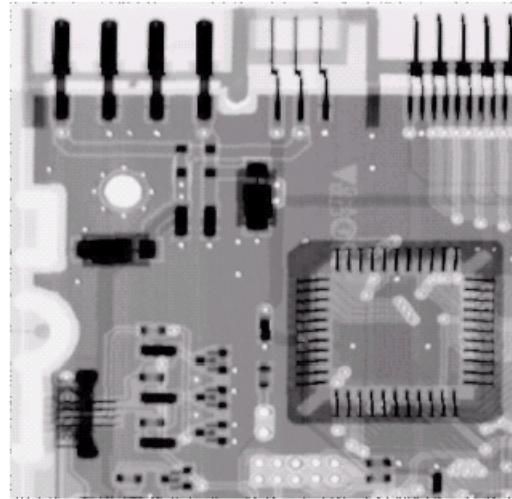
Positive values of  $Q$  eliminate pepper noise  
Negative values of  $Q$  eliminate salt noise

# Noise Removal Examples (cont...)

Image  
Corrupted  
By Pepper  
Noise



Result of  
Filtering Above  
With 3\*3  
Contraharmonic  
 $Q=1.5$



# Noise Removal Examples (cont...)

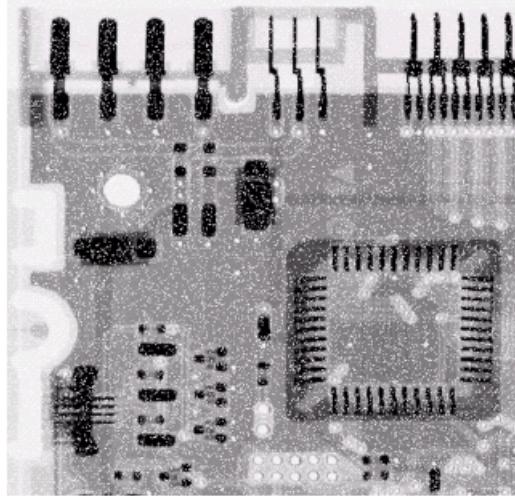
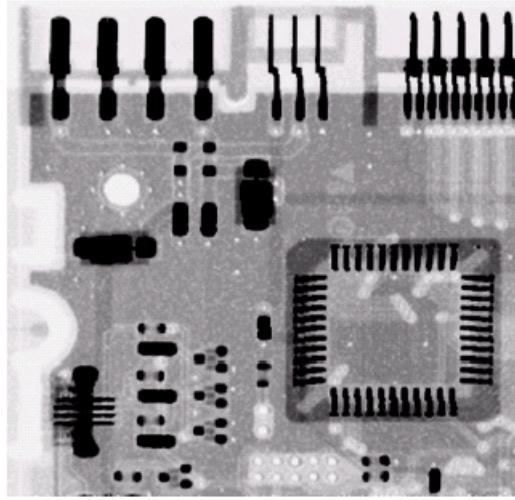


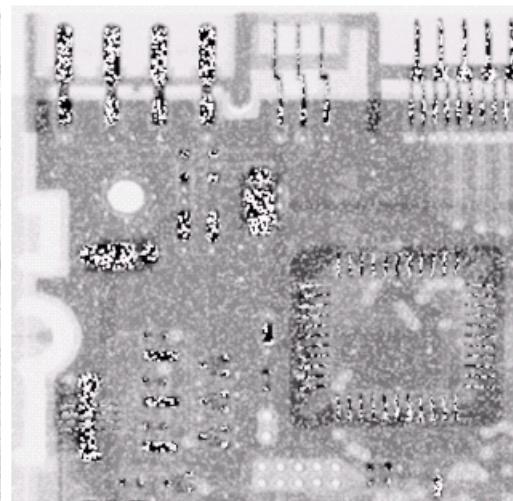
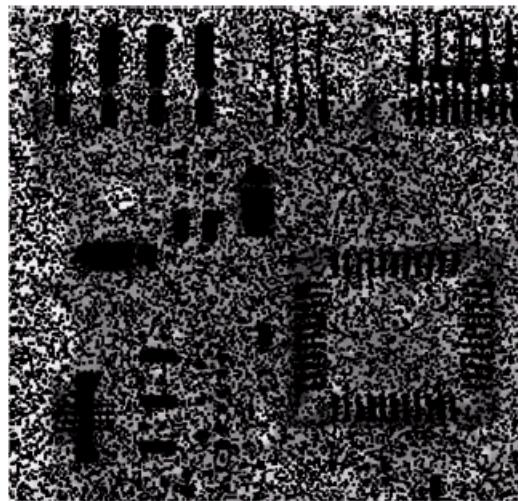
Image  
Corrupted  
By Salt  
Noise



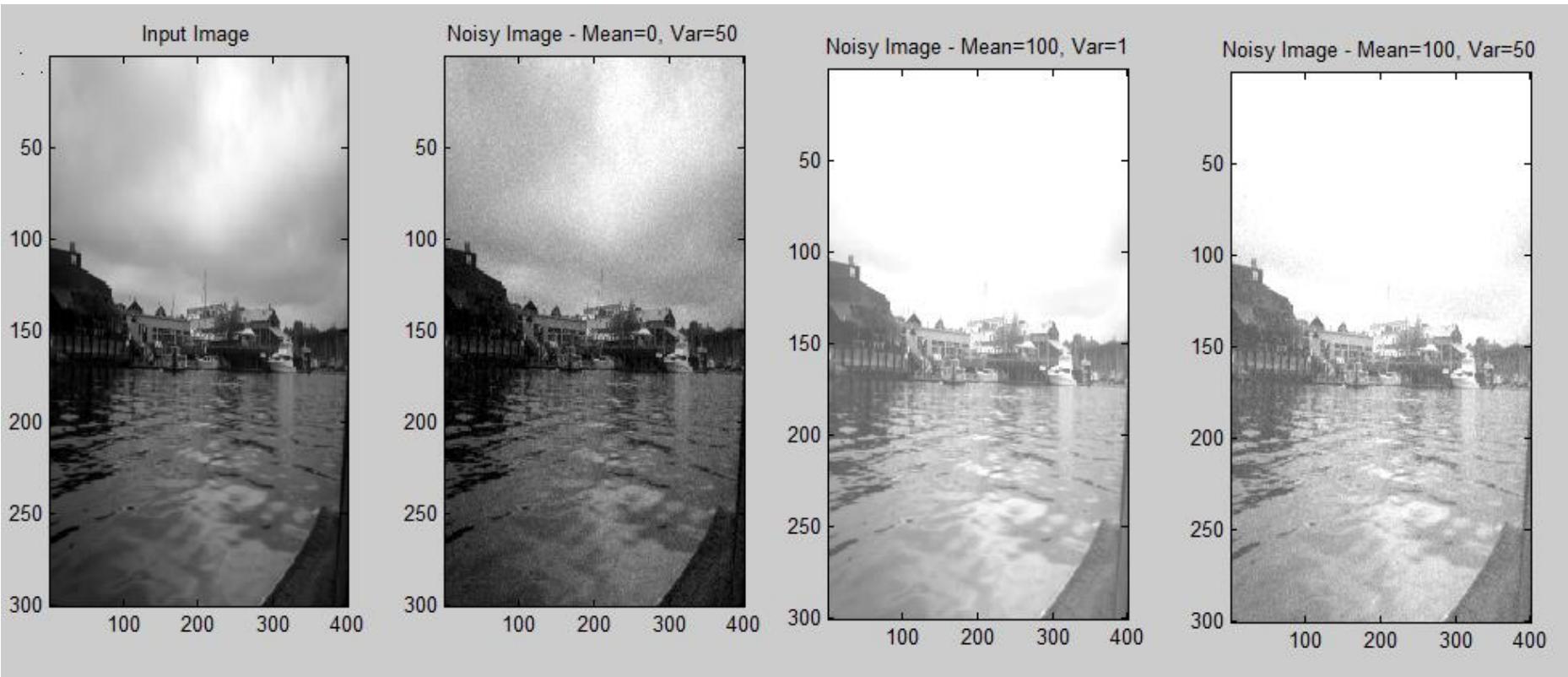
Result of  
Filtering Above  
With  $3 \times 3$   
Contraharmonic  
 $Q=-1.5$

# Contraharmonic Filter: Caution!

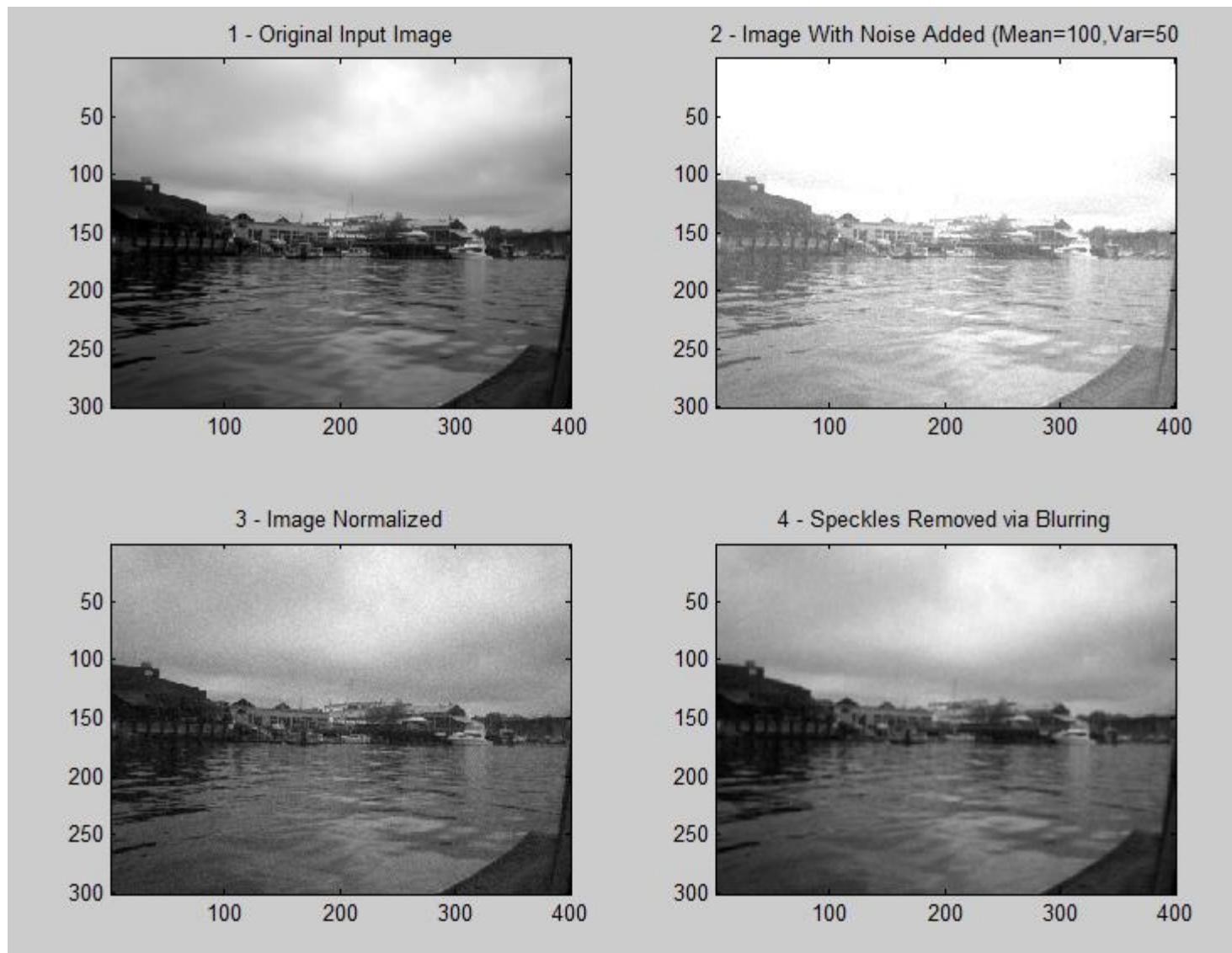
Choosing the wrong value for Q when using the contraharmonic filter can have drastic results



# Filtering of Gaussian noise



# Filtering of Gaussian noise (contd.)



# Rank Filters (non linear filtering)

Spatial filters that are based on ordering the pixel values that make up the neighbourhood operated on by the filter

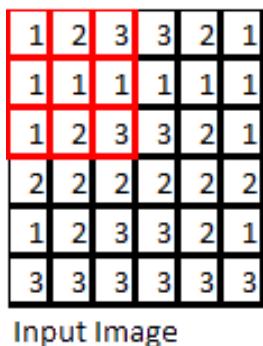
Useful spatial filters include

- Median filter
- Max and min filter
- Midpoint filter
- Alpha trimmed mean filter

## Median Filter:

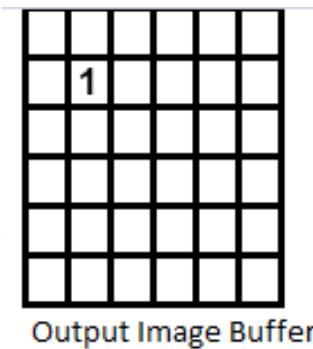
$$\hat{f}(x, y) = \underset{(s,t) \in S_{xy}}{\operatorname{median}}\{g(s, t)\}$$

- Excellent at noise removal, without the smoothing effects that can occur with other smoothing filters
  - Particularly good when salt and pepper noise is present



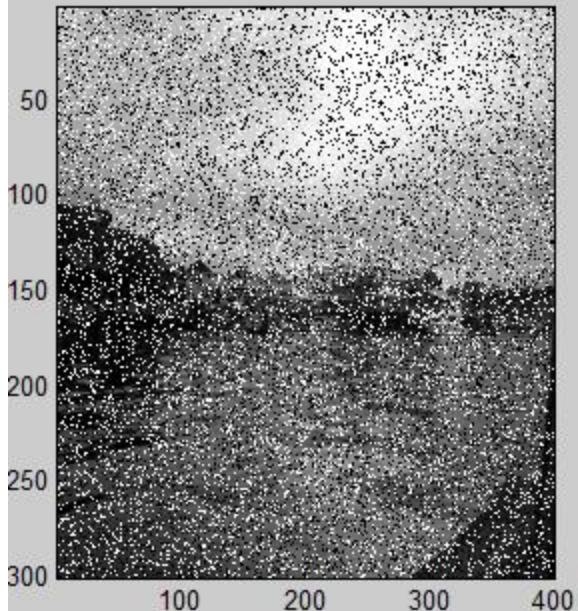
|   |   |   |
|---|---|---|
| 1 | 2 | 3 |
| 1 | 1 | 1 |
| 1 | 2 | 3 |

|   |                                                       |
|---|-------------------------------------------------------|
| 1 |                                                       |
| 1 |                                                       |
| 1 |                                                       |
| 1 |                                                       |
| 1 |                                                       |
| 1 | ← Median value<br>(Sort the pixels by<br>value FIRST) |
| 2 |                                                       |
| 2 |                                                       |
| 3 |                                                       |
| 3 |                                                       |

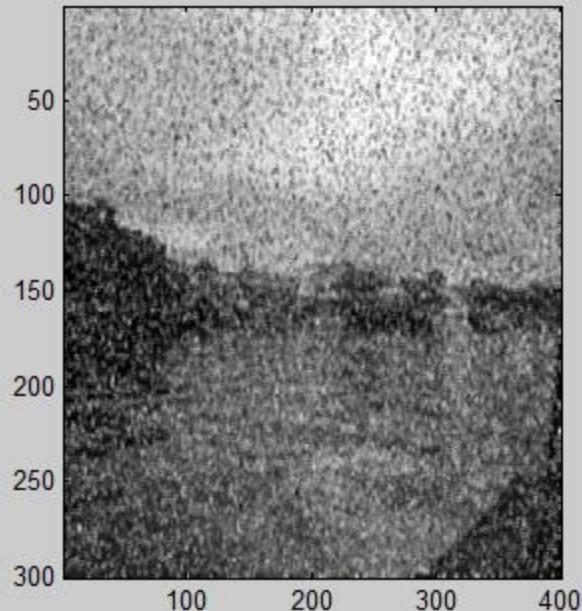


# Mean Vs. Median filter

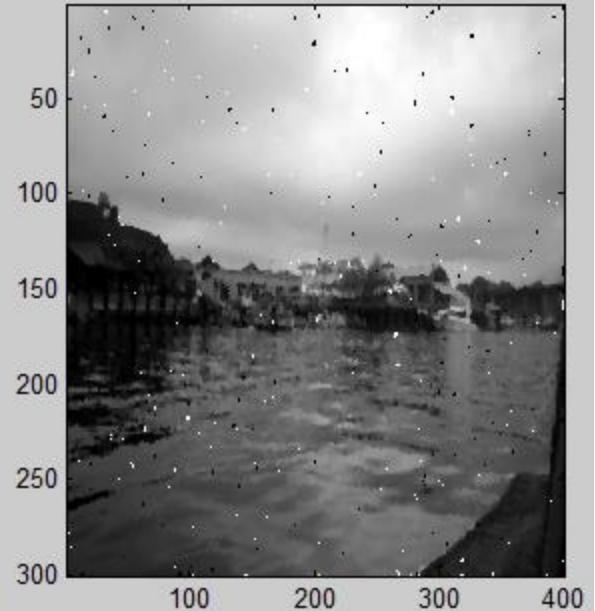
Input Image with 25% Salt & Pepper Noise



Noise Removed via 3x3 Smoothing Filter

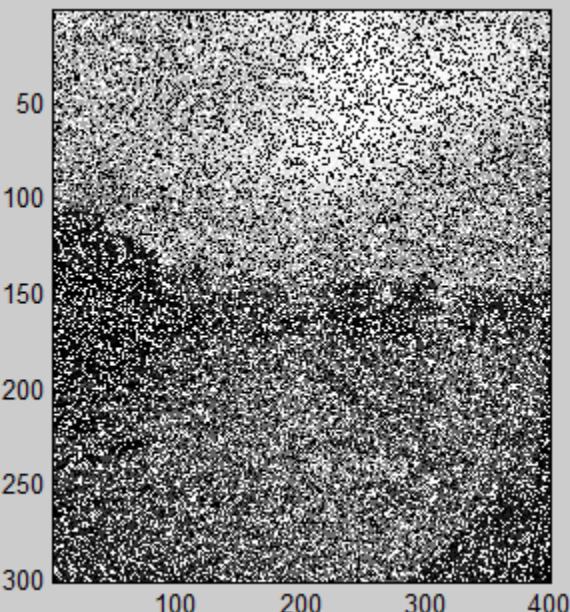


Noise Removed via 3x3 Median Filter

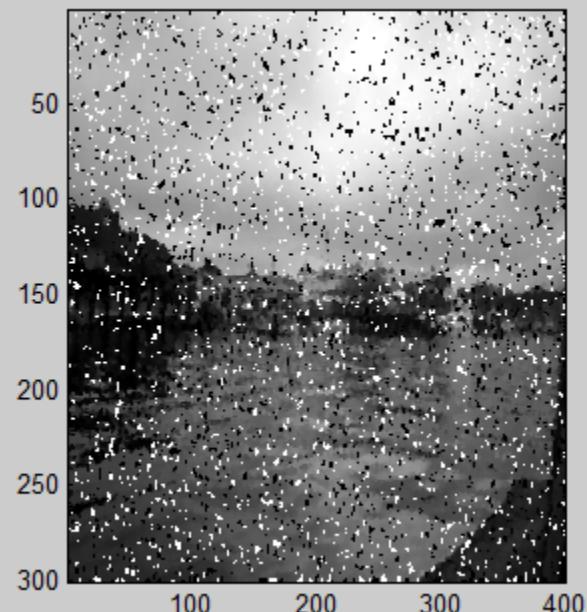


# Noise Removal Examples

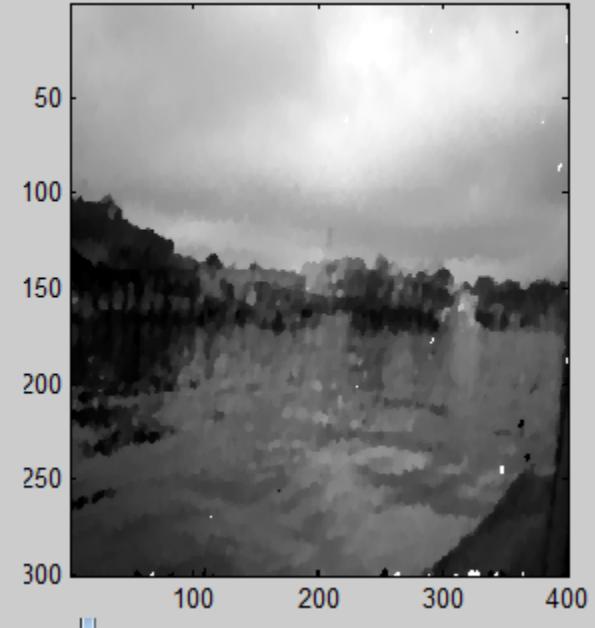
Input Image with 50% Salt & Pepper Noise



Noise Removed via 3x3 Median Filter

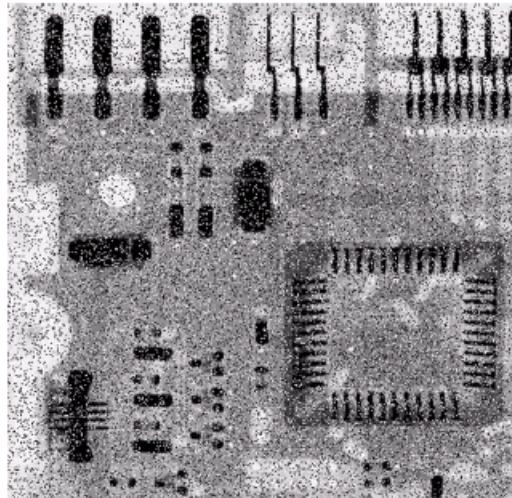


Noise Removed via 5x5 Median Filter

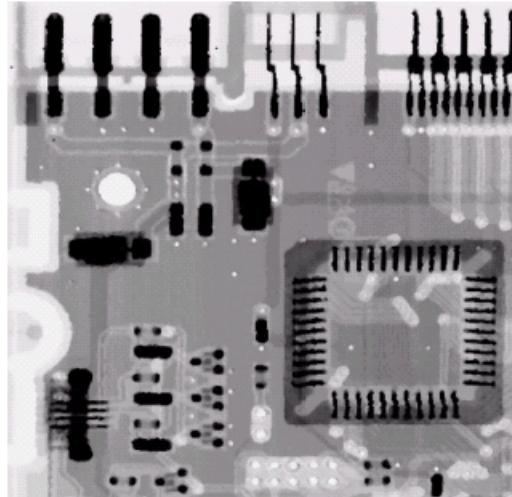


# Noise Removal Examples (contd)

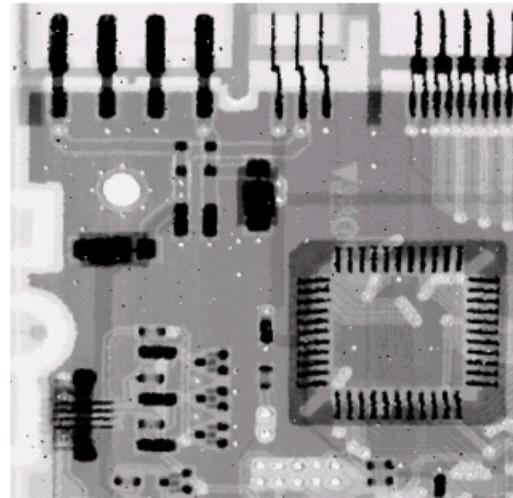
Image  
Corrupted  
By Salt And  
Pepper Noise



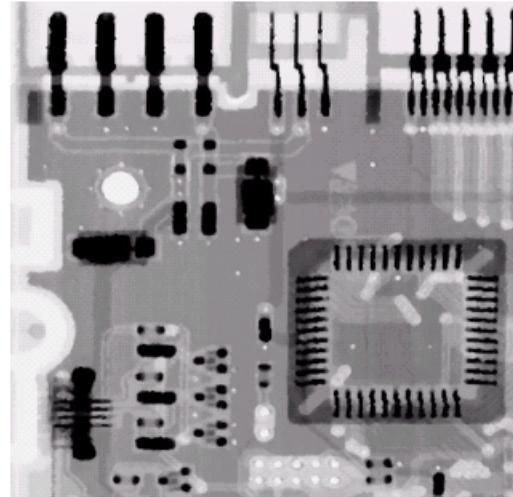
Result of 2  
Passes With  
A  $3 \times 3$  Median  
Filter



Result of 1  
Pass With A  
 $3 \times 3$  Median  
Filter



Result of 3  
Passes With  
A  $3 \times 3$  Median  
Filter



# Max and Min Filter

## Max Filter:

$$\hat{f}(x, y) = \max_{(s,t) \in S_{xy}} \{g(s, t)\}$$

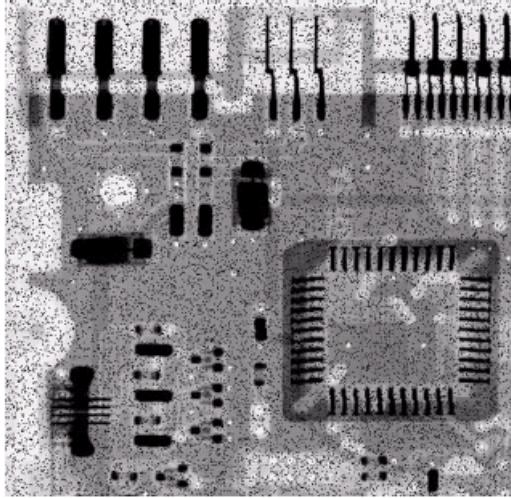
## Min Filter:

$$\hat{f}(x, y) = \min_{(s,t) \in S_{xy}} \{g(s, t)\}$$

Max filter is good for pepper noise and min is good for salt noise

# Noise Removal Examples (cont...)

Image  
Corrupted  
By Pepper  
Noise



Result Of  
Filtering  
Above  
With A  $3 \times 3$   
Max Filter

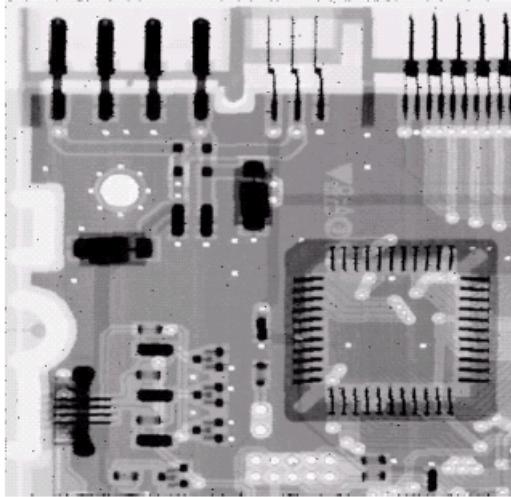
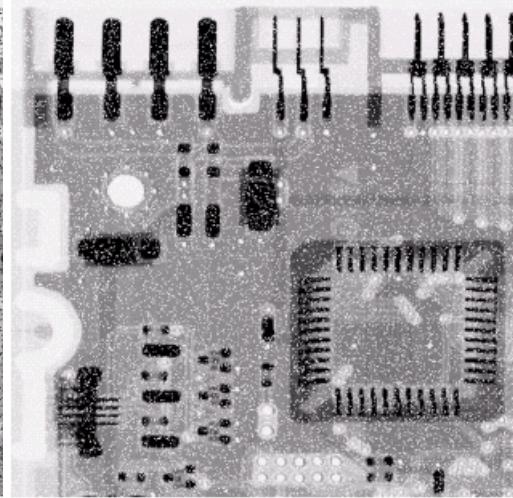
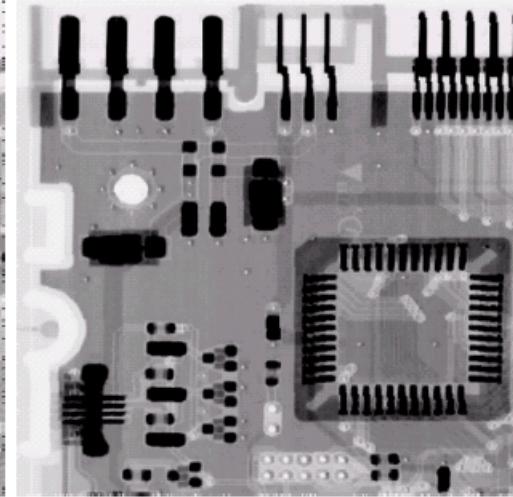


Image  
Corrupted  
By Salt  
Noise



Result Of  
Filtering  
Above  
With A  $3 \times 3$   
Min Filter



# Alpha-Trimmed Mean Filter

## Alpha-Trimmed Mean Filter:

$$\hat{f}(x, y) = \frac{1}{mn - d} \sum_{(s,t) \in S_{xy}} g_r(s, t)$$

We can delete the  $d/2$  lowest and  $d/2$  highest grey levels

So  $g_r(s, t)$  represents the remaining  $mn - d$  pixels

# Noise Removal Examples (cont...)

Image  
Corrupted  
By Uniform  
Noise

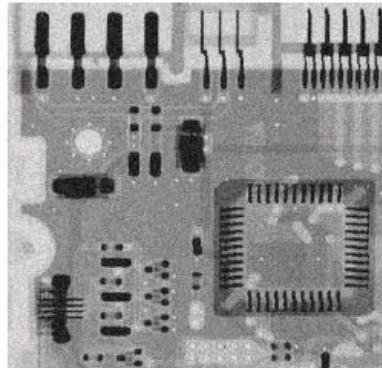
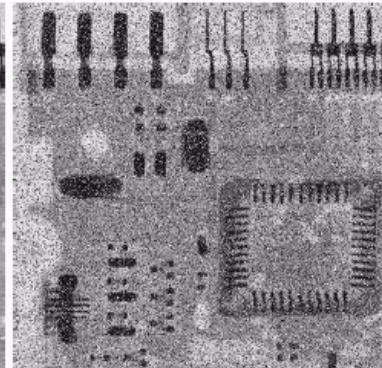
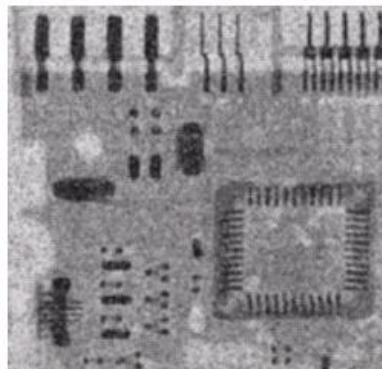


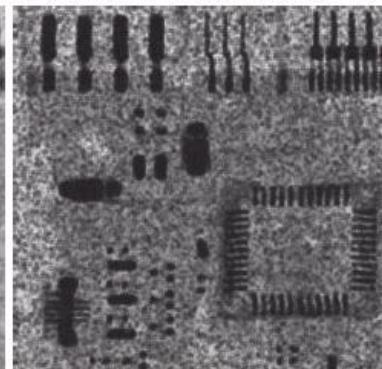
Image Further  
Corrupted  
By Salt and  
Pepper Noise



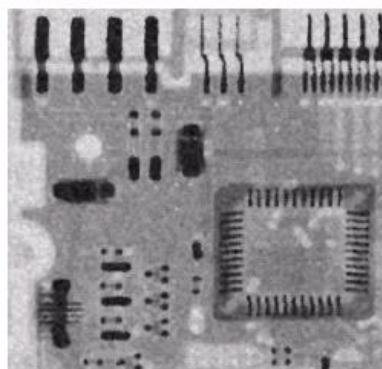
Filtered By  
5\*5 Arithmetic  
Mean Filter



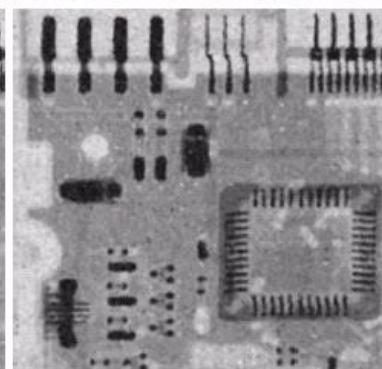
Filtered By  
5\*5 Geometric  
Mean Filter



Filtered By  
5\*5 Median  
Filter



Filtered By  
5\*5 Alpha-Trimmed  
Mean Filter



# Adaptive Filters

The filters discussed so far are applied to an entire image without any regard for how image characteristics vary from one point to another

The behaviour of **adaptive filters** changes depending on the characteristics of the image inside the filter region

# Adaptive filter: Neighborhood-based

## Adaptive local noise reduction filter

- Response based on 4 quantities
  - Local variance, **variance of noise**,  $g(x,y)$ , and local mean
- Behavior of filter
  - If variance of noise is zero, return  $g(x,y)$
  - If the local variance is high compared to the variance of noise, return a value close to  $g(x,y)$
  - If the two variances are equal, return the mean value

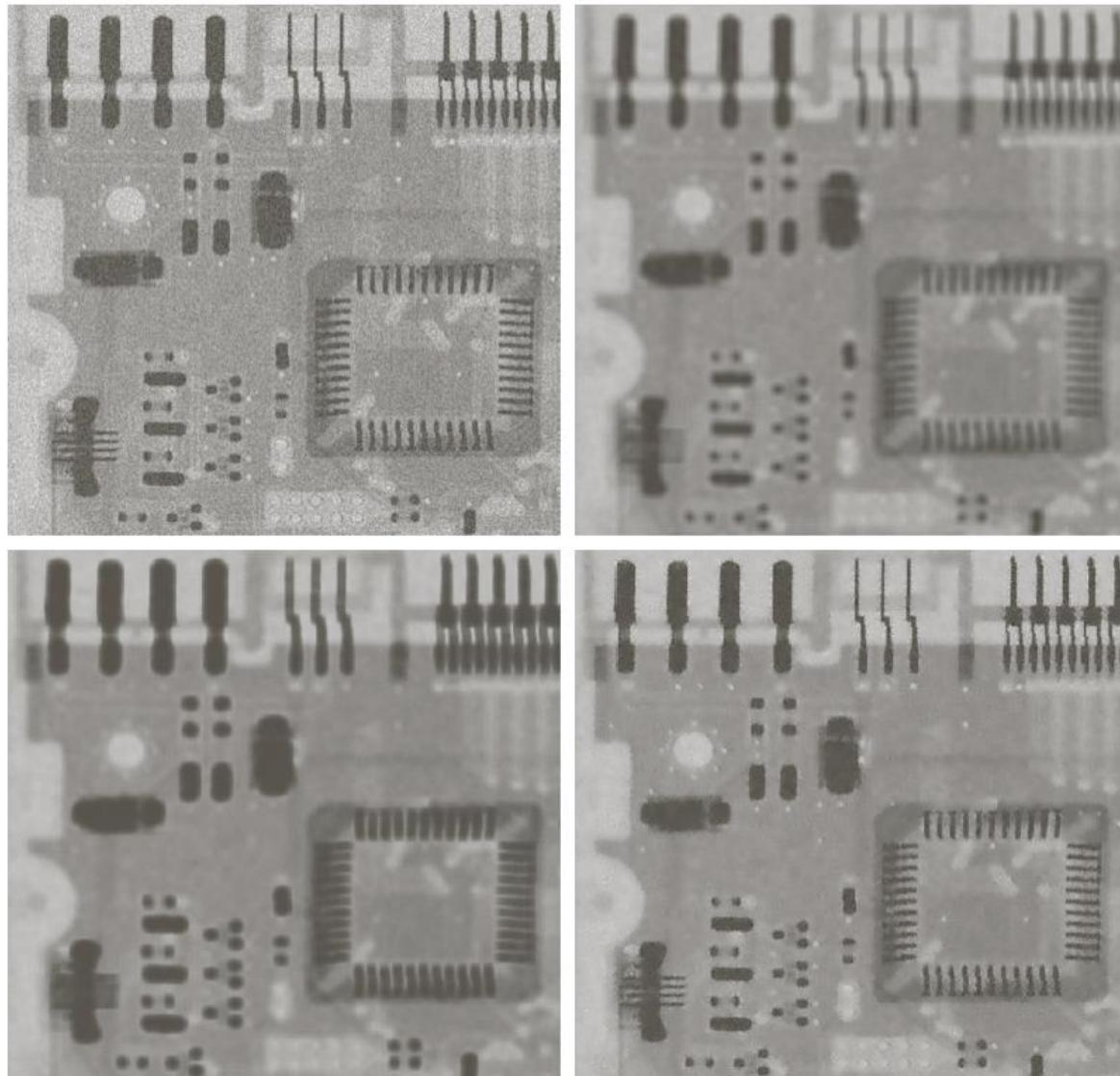
$$\hat{f}(x,y) = g(x,y) - \frac{S_h^2}{S_L^2} [g(x,y) - m_L]$$

# Adaptive filter: Neighborhood-based

a  
b  
c  
d

**FIGURE 5.13**

- (a) Image corrupted by additive Gaussian noise of zero mean and variance 1000.  
(b) Result of arithmetic mean filtering.  
(c) Result of geometric mean filtering.  
(d) Result of adaptive noise reduction filtering. All filters were of size  $7 \times 7$ .



# Adaptive Median Filtering

The median filter performs relatively well on impulse noise as long as the spatial density of the impulse noise is not large

The adaptive median filter can handle much more spatially dense impulse noise, and does less distortion

The key insight in the adaptive median filter is that the filter size changes depending on the characteristics of the image

# Adaptive Median Filtering (cont...)

The adaptive median filter has following purposes:

- Remove spatially dense impulse noise
- Reduce distortion

First examine the following notation:

- $z_{min}$  = minimum grey level in  $S_{xy}$
- $z_{max}$  = maximum grey level in  $S_{xy}$
- $z_{med}$  = median of grey levels in  $S_{xy}$
- $z_{xy}$  = grey level at coordinates  $(x, y)$
- $S_{max}$  = maximum allowed size of  $S_{xy}$

# Adaptive Median Filtering (cont...)

Level A:  $A1 = z_{med} - z_{min}$

$A2 = z_{med} - z_{max}$

If  $A1 > 0$  and  $A2 < 0$ , Go to level B

Else increase the window size

If window size  $\leq S_{max}$  repeat level A

Else output  $z_{med}$

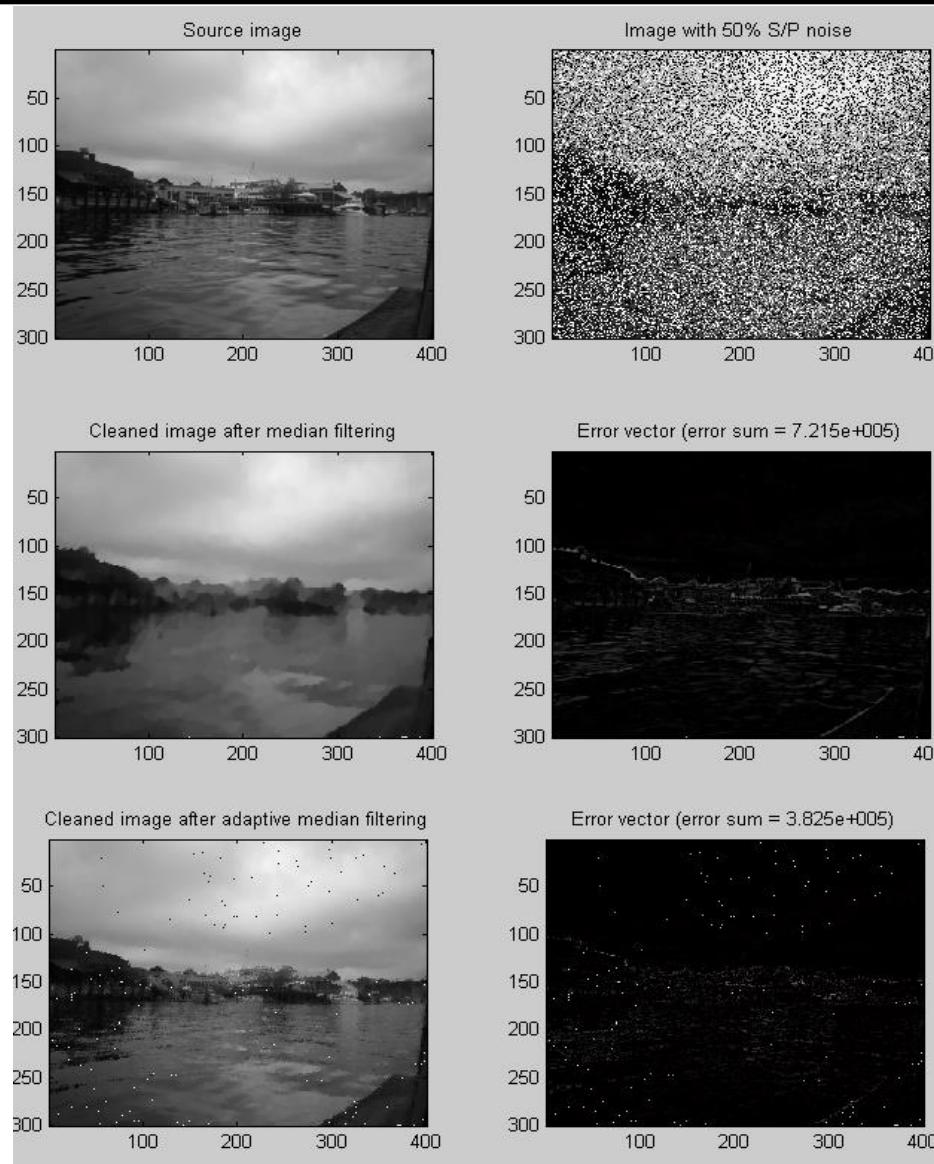
Level B:  $B1 = z_{xy} - z_{min}$

$B2 = z_{xy} - z_{max}$

If  $B1 > 0$  and  $B2 < 0$ , output  $z_{xy}$

Else output  $z_{med}$

# Simple Adaptive Median Filtering Example



Using only  
level B and  
filter 7 x 7

# Adaptive Median Filtering Examples

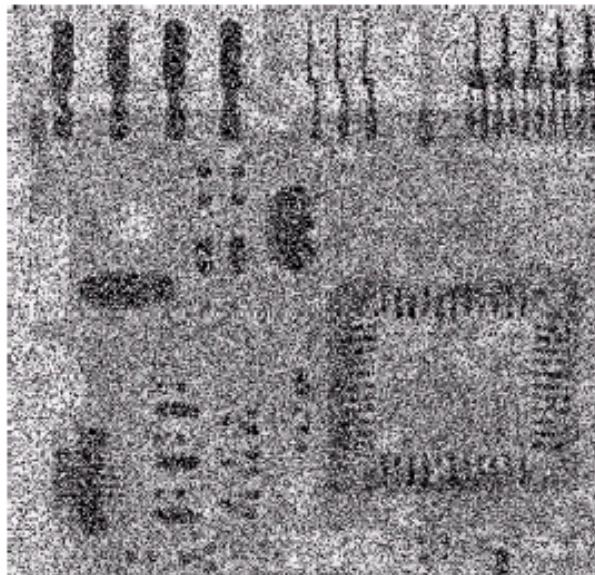
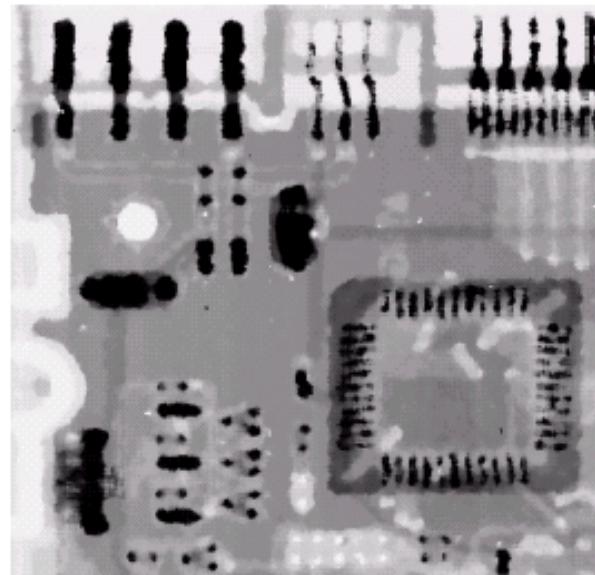
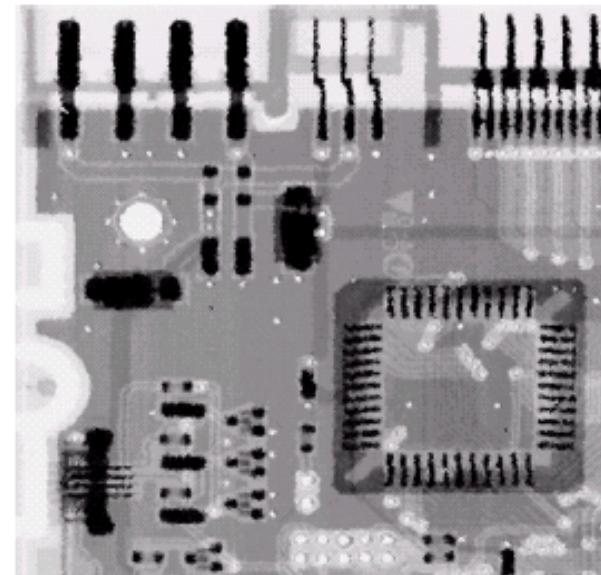


Image corrupted by salt  
and pepper noise with  
probabilities  $P_a = P_b = 0.25$



Result of filtering with a 7  
\* 7 median filter



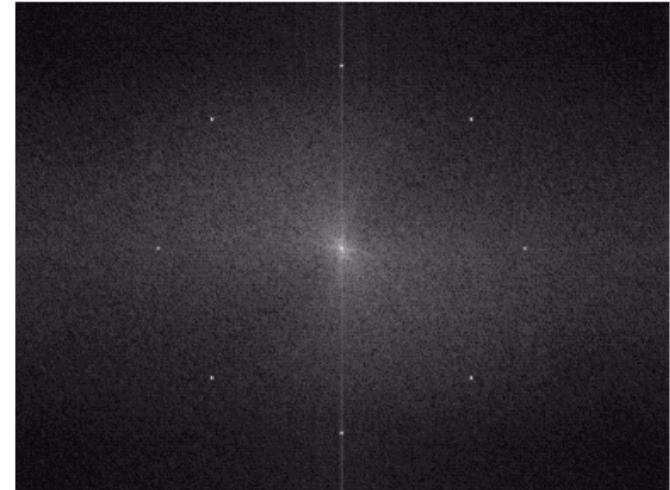
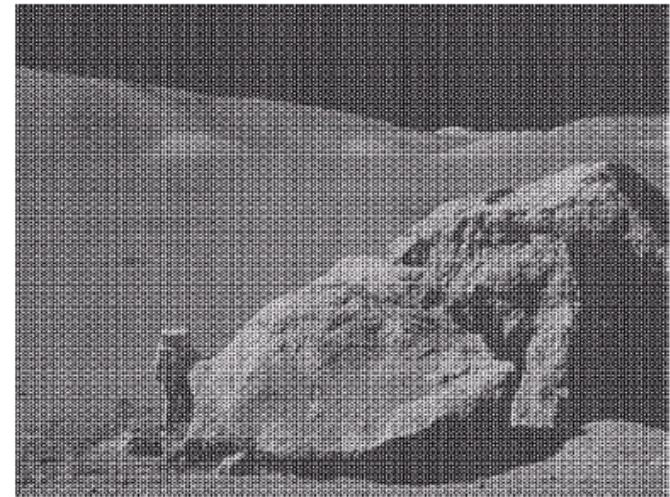
Result of adaptive median  
filtering with  $S_{max} = 7$

# Periodic Noise removal

Typically arises due to electrical or electromagnetic interference

Gives rise to regular noise patterns in an image

Frequency domain techniques in the Fourier domain are most effective at removing periodic noise



# Band Reject Filters

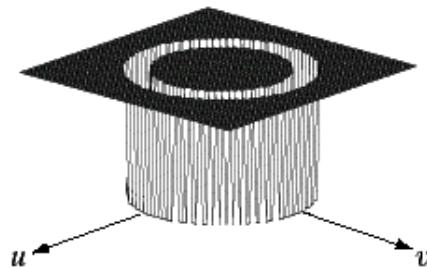
Removing periodic noise from an image involves removing a particular range of frequencies from that image

*Band reject* filters can be used for this purpose  
An ideal band reject filter is given as follows:

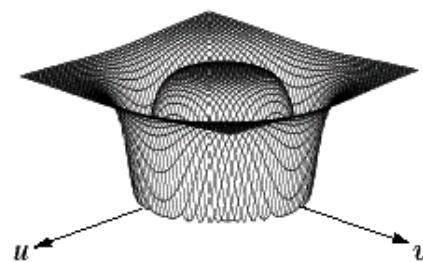
$$H(u, v) = \begin{cases} 1 & \text{if } D(u, v) < D_0 - \frac{W}{2} \\ 0 & \text{if } D_0 - \frac{W}{2} \leq D(u, v) \leq D_0 + \frac{W}{2} \\ 1 & \text{if } D(u, v) > D_0 + \frac{W}{2} \end{cases}$$

# Band Reject Filters (cont...)

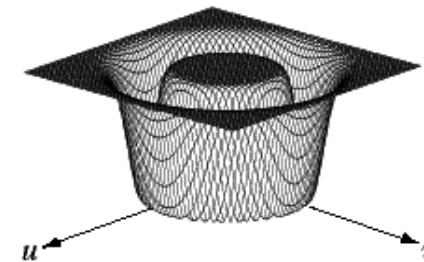
The ideal band reject filter is shown below, along with Butterworth and Gaussian versions



Ideal Band  
Reject Filter



Butterworth  
Band Reject  
Filter (of order 1)

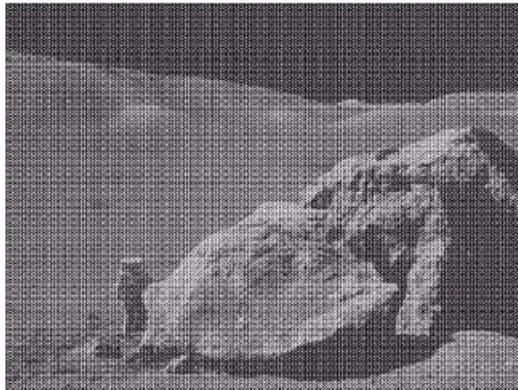


Gaussian  
Band Reject  
Filter

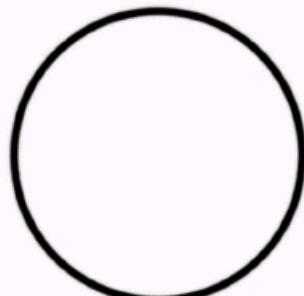
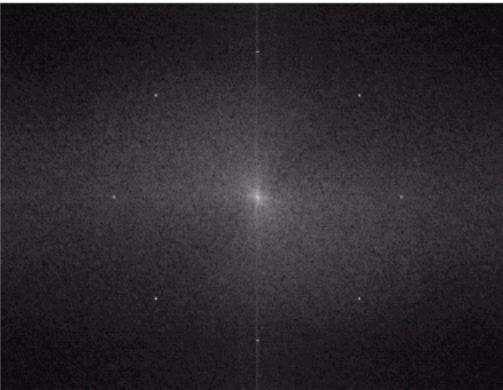
| Ideal                                                                                                                        | Butterworth                                                          | Gaussian                                                     |
|------------------------------------------------------------------------------------------------------------------------------|----------------------------------------------------------------------|--------------------------------------------------------------|
| $H(u, v) = \begin{cases} 0 & \text{if } D_0 - \frac{W}{2} \leq D \leq D_0 + \frac{W}{2} \\ 1 & \text{otherwise} \end{cases}$ | $H(u, v) = \frac{1}{1 + \left[ \frac{DW}{D^2 - D_0^2} \right]^{2n}}$ | $H(u, v) = 1 - e^{-\left[ \frac{D^2 - D_0^2}{DW} \right]^2}$ |

# Band Reject Filter Example

Image corrupted by sinusoidal noise



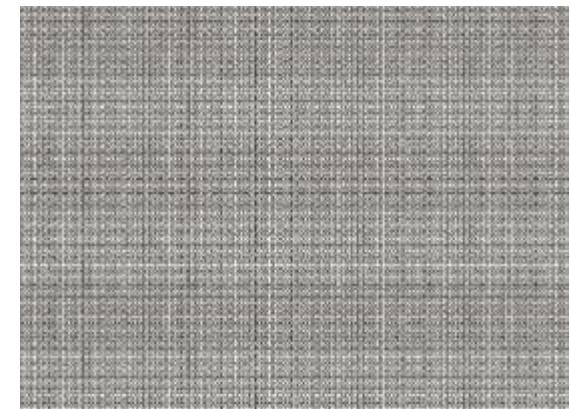
Fourier spectrum of corrupted image



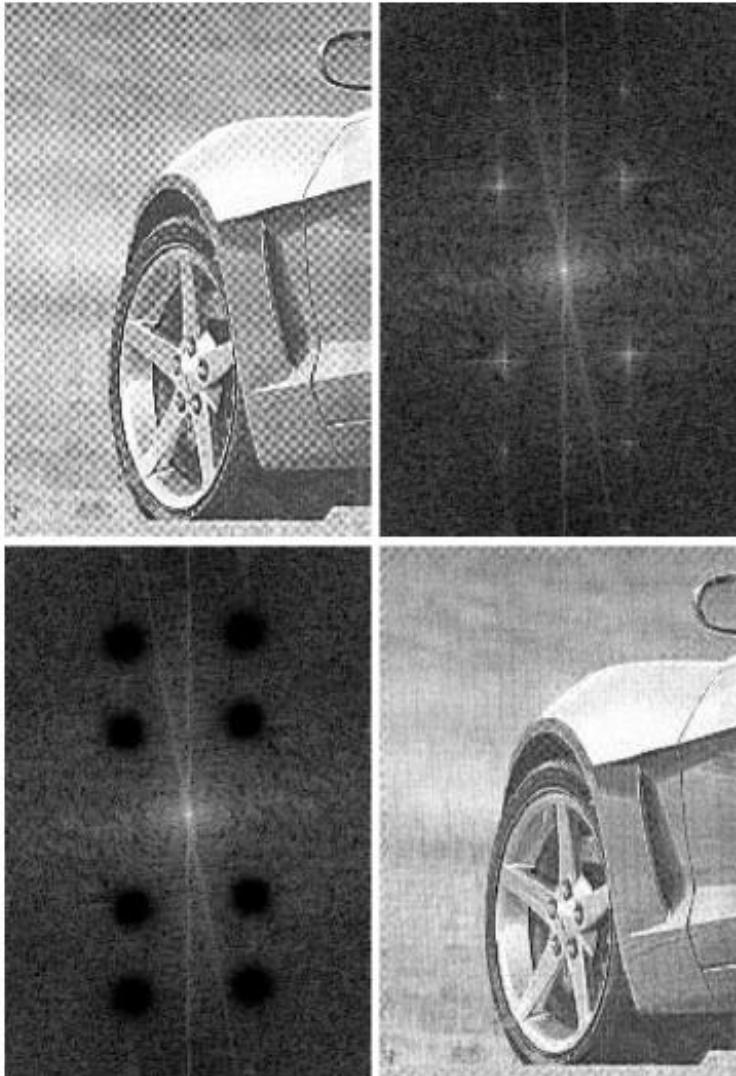
Butterworth band reject filter



Filtered image



# Notch Filter



a b  
c d

**FIGURE 4.64**

- (a) Sampled newspaper image showing a moiré pattern.
- (b) Spectrum.
- (c) Butterworth notch reject filter multiplied by the Fourier transform.
- (d) Filtered image.

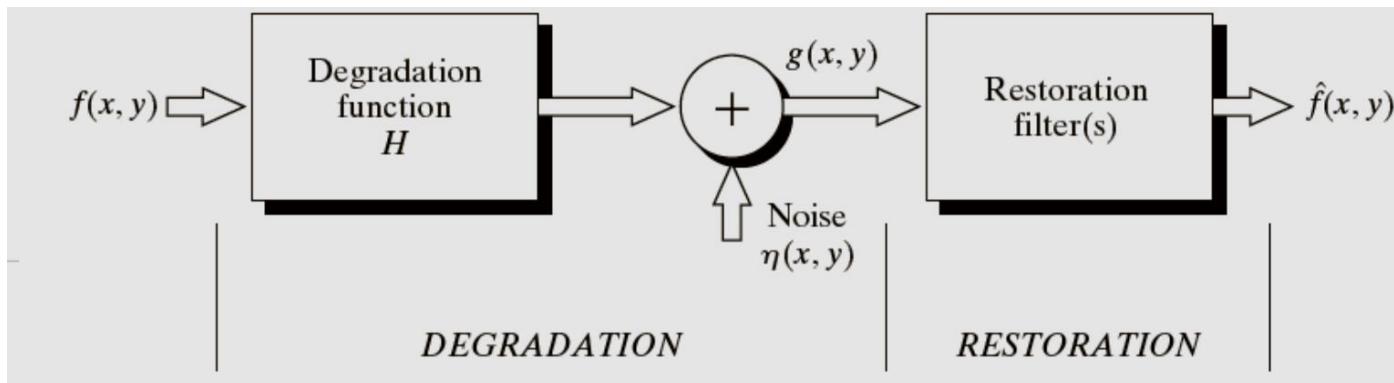
# Restoration From Degradation

# Linear, position-invariant degradation model

- Many types of degradation can be *approximated* by linear, position-invariant processes.

$$g(x,y) = f(x,y) * h(x,y) + \eta(x,y)$$

$$G(u,v) = F(u,v)H(u,v) + N(u,v)$$



- Image deconvolution: find  $H(u,v)$  and apply inverse process

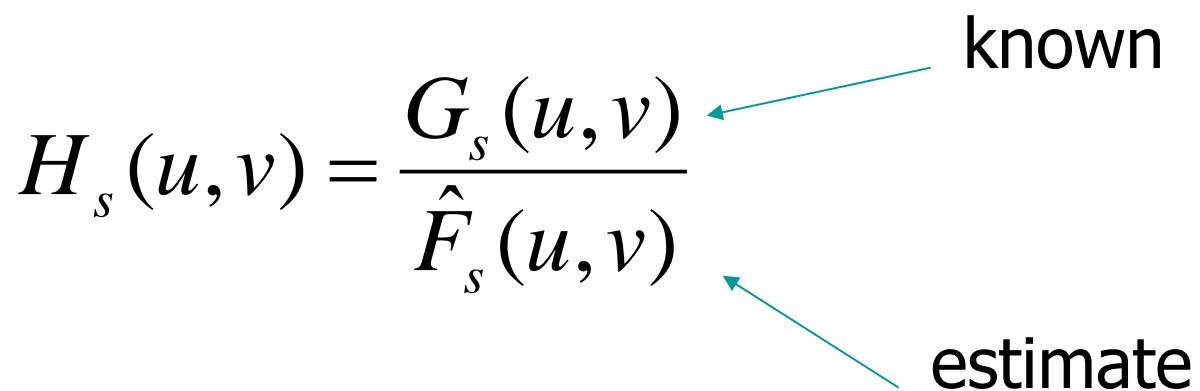
# Estimating the Degradation Function

- By image observation: Take a region in the image with
  - Simple structure
  - Strong signal content (negligible noise)
- Estimate the original image in the window

$$H_s(u, v) = \frac{G_s(u, v)}{\hat{F}_s(u, v)}$$

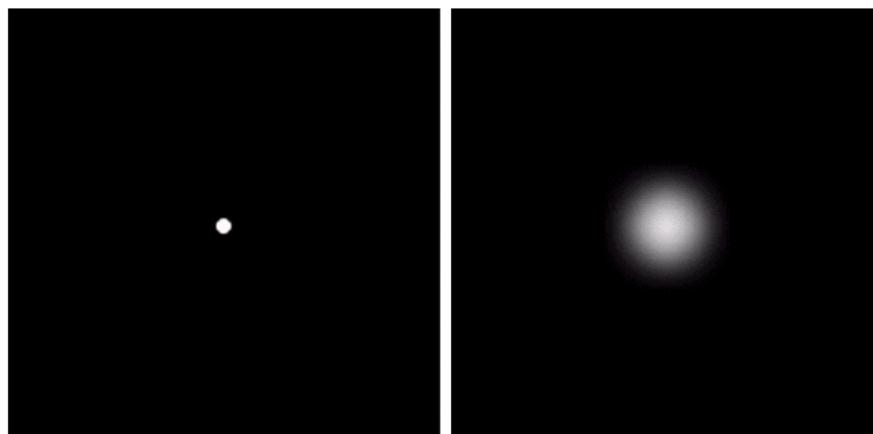
known

estimate



# Estimating the Degradation Function

- By experimentation: If the image acquisition system is available, then obtain **impulse response**
  - $G(u,v) = H(u,v)F(u,v) + N(u,v)$
  - $H(u,v) = G(u,v) / A$



a b

**FIGURE 5.24**  
Degradation estimation by impulse characterization.  
(a) An impulse of light (shown magnified).  
(b) Imaged (degraded) impulse.

# Estimating the Degradation Function

- By modeling: Ex. Atmospheric turbulence

$$H(u, v) = e^{-k(u^2 + v^2)^{5/6}}$$

original



$k=0.001$



$k=0.0025$



$k=0.00025$

# Estimating the Degradation Function

- Derive a **mathematical model**: ex. Motion blur

$$g(x, y) = \int_0^T f(x - x_0(t), y - y_0(t)) dt$$

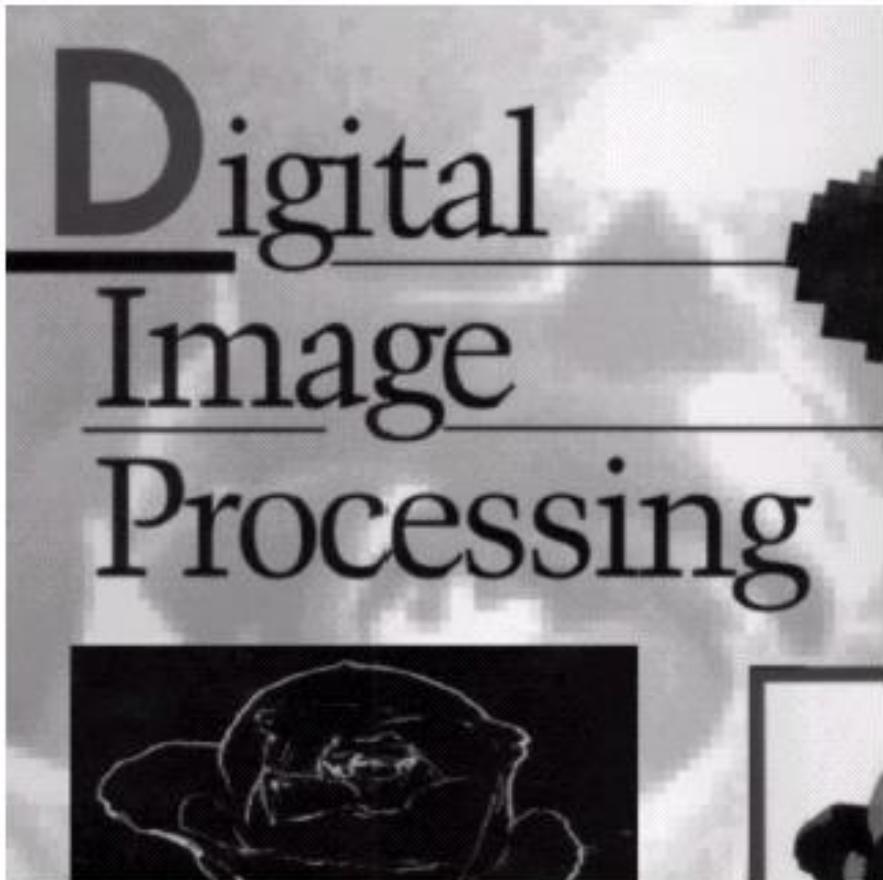
Fourier  
transform

Planar motion

$$G(u, v) = F(u, v) \int_0^T e^{-j2\pi[ux_0(t)+vy_0(t)]} dt$$

# Estimating the Degradation Function

original



Apply motion model



# Inverse Filtering

- Compute an estimate by simply dividing the transform of degraded image by the degradation function

$$\hat{F}(u, v) = \frac{G(u, v)}{H(u, v)} = F(u, v) + \frac{N(u, v)}{H(u, v)}$$

Estimate of  
original image

Unknown  
noise

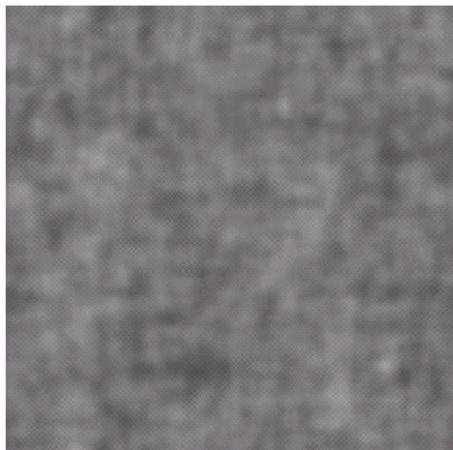
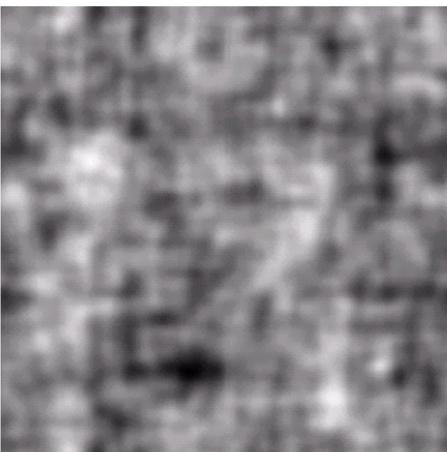
Problem: 0 or small values

- Suffers from noise amplification

# Psuedo inverse filtering

- Limiting the analysis to frequencies near origin

Full  
inverse  
filter  
for  
 $k=0.0025$



$$\hat{F}(u, v) = G(u, v)\hat{H}(u, v)$$

$$\hat{H}(u, v) = \begin{cases} 1/H(u, v), & |u^2 + v^2| \leq \eta \\ 0, & |u^2 + v^2| > \eta \end{cases}$$

$$H(u, v) = e^{-k(u^2+v^2)}$$

Cut  
Outside  
40%

Degraded  
Image



Cut  
Outside  
85%

# Wiener Filtering

- Incorporates both the degradation and statistical characteristics of noise
- Objective function: find an estimate  $\hat{f}$  of  $f$  such that the mean square error between them is minimized

$$e^2 = E\{(f - \hat{f})^2\}$$

$$\hat{F}(u, v) = \frac{H^*(u, v) S_f(u, v)}{S_f(u, v) |H(u, v)|^2 + S_n(u, v)} G(u, v)$$

$S_f(u, v) = |F(u, v)|^2 = \text{power spectrum of the undegraded image}$

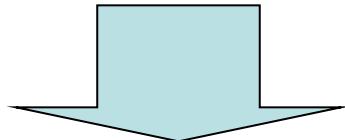
$S_n(u, v) = |N(u, v)|^2 = \text{power spectrum of the noise}$

# Wiener Filtering

$$\hat{F}(u, v) = \left[ \frac{1}{H(u, v)} \frac{|H(u, v)|^2}{|H(u, v)|^2 + S_\eta(u, v) / S_f(u, v)} \right] G(u, v)$$

Constant

Unknown



$$\hat{F}(u, v) = \left[ \frac{1}{H(u, v)} \frac{|H(u, v)|^2}{|H(u, v)|^2 + K} \right] G(u, v)$$

# Inverse vs. Wiener Filtering



a b c

**FIGURE 5.28** Comparison of inverse- and Wiener filtering. (a) Result of full inverse filtering of Fig. 5.25(b). (b) Radially limited inverse filter result. (c) Wiener filter result.

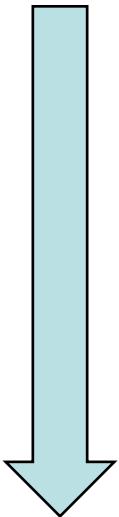
Full inverse  
filtering

Radially limited  
inverse filtering

Wiener filtering  
(K was chosen interactively)

# Further comparison

Reduced Variance



a b c  
d e f  
g h i



**FIGURE 5.29** (a) 8-bit image corrupted by motion blur and additive noise. (b) Result of inverse filtering. (c) Result of Wiener filtering. (d)–(f) Same sequence, but with noise variance one order of magnitude less. (g)–(i) Same sequence, but noise variance reduced by five orders of magnitude from (a). Note in (h) how the deblurred image is quite visible through a “curtain” of noise.

# Quantitative evaluation of image restoration

- Some Performance Measures for evaluating Image restoration method where the original image and reconstructed image from its degraded version is available are as follows:
  - Signal-to-Noise Ratio (SNR)
  - Mean Square Error (MSE)
  - Root Mean Square Error (RMSE)
  - Peak Signal-to-Noise Ratio (PSNR)

- Signal to noise ratio

$$SNR = \frac{\sum_{x=1}^m \sum_{y=1}^n |F(x, y)|^2}{\sum_{x=1}^m \sum_{y=1}^n |N(x, y)|^2}$$

- ***Mean square error:***

$$MSE = \frac{1}{m \times n} \sum_{i=1}^m \sum_{j=1}^n [I'(i, j) - I(i, j)]^2$$

Where  $I$  is the original image and  $I'$  is the reconstructed image.

- ***Root mean square error:***

$$RMSE = \sqrt{MSE}$$

- ***Peak signal-to-noise ratio:***

$$PSNR = 20 \log_{10} \left[ \frac{255}{RMSE} \right]$$