

# Physics-105-Lecture-Notes-02-05-2019

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### Abstract

A single PDF with all lectures in a single document can be downloaded at <https://www.dropbox.com/sh/8sqzvxghvbjifco/AAC9LoSRnsRQDp7pYedgWpQMa?dl=0>. The password is 'analytic.mech.dsp'. This file was automatically generated using a script, so there might be some errors. If there are, you can contact me at <mailto:ctdunc@berkeley.edu>.

## 0.1 Examples of Lagrangian Mechanics

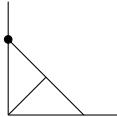
### 0.1.1 Cone?

Missed the first one, but we know that *angular momentum is conserved*. Basically, just a whole lot of algebran happening here, with something rotating in a conic shape, or on the surface of a cone (maybe like throwing a coin into one of those things at McDonalds). Stable solutions can be given by  $\ddot{r} = 0$ . I.e. the coin doesn't ever go into the money receptacle. Gives that  $\theta^2 \tan \alpha = \frac{g}{r_0} \Rightarrow \dot{\theta}^2 = \frac{\omega_0}{\tan \alpha}$ . Which implies that in order to have some stable orbit in a cone at a certain angle  $\alpha$ , you have explicit angular momentum dependence.



### 0.1.2 Mass/Spring on a T

Imagine some  $T$  on a tabletop, that looks a bit like this



Where the dot is connected to the  $T$  by a spring that's hooked up at the juncture. Let  $\omega t$  be the angle between the x-axis and the  $T$ . then we can write  $\vec{r} = (l \cos \omega t - \rho \sin \omega t)\hat{x} + (l \sin \omega t + \rho \cos \omega t)\hat{y}$

$$T = \frac{1}{2}m(\dot{r})^2 = \frac{1}{2}m(\omega^2(l^2 + \rho^2) + \dot{\rho}^2 + 2\omega l\dot{\rho})$$

put in to the euler lagrange equation

$$\begin{aligned}\frac{\partial L}{\partial \rho} &= m\omega^2\rho - k\rho \\ \frac{\partial L}{\partial \dot{\rho}} &= m\dot{\rho} + \omega l\end{aligned}$$

So, equating these two things, gives us that

$$\ddot{\rho} + \left(\frac{k}{m} - \omega^2\right)\rho = 0$$

which yields 3 sort of 'classes' of solutions. First is where  $\omega < \sqrt{\frac{k}{m}}$ , which yields a simple harmonic oscillator, very fun! We also could have  $\omega > \sqrt{\frac{k}{m}}$ , which gives us that  $\rho(t) = Be^{\alpha t} + Ce^{-\alpha t}$ . There's also the case of equality, which gives us resonant oscillation, or just a growth term  $\rho(t) \sim t$ .

### 0.1.3 Now with Gravity!

Take the previous problem, and just add gravity into the mix, since we all like to have fun. Now we have  $V = mgy$ , and we have  $y$  from the previous problem, so the lagrangian becomes some really long wild thing, that I cannot see (Bale didn't do the whole thing out, but the principle of the problem is similar to what we did above).

## 1 Symmetries (Formally)

Considering changes to  $L$  (the lagrangian) when we perturb one of the coordinates. Say it's  $q_i \rightarrow \tilde{q}_i = q_i + k_i$

### 1.1 Linear Momentum

consider

$$\begin{aligned}\tilde{x} &= x + \epsilon & \dot{\tilde{x}} &= \dot{x} \\ T &= \frac{1}{2}m\dot{x}^2 = \frac{1}{2}m\dot{\tilde{x}}^2 \\ L &= \frac{1}{2}m\dot{x}^2 - V(x) & \tilde{L} &= \frac{1}{2}m\dot{\tilde{x}}^2 - \tilde{V}(x)\end{aligned}$$

If we apply the constraint that  $L(x, \dot{x}) = \tilde{L}(\tilde{x}, \dot{\tilde{x}})$ , then  $V(x)$  has to be invariant to spatial perturbation, which implies that  $F_x = 0$ , since  $-F = \frac{\partial V}{\partial x}$ . This is just a meme'd way of writing conservation of linear momentum, since it boils down to

$$\frac{d}{dt}(m\dot{x}) = 0$$

### 1.2 Noethers Theorem (intro)

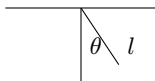
$$\begin{aligned}L(q, \dot{q}) &= L(q + \epsilon k, \dot{q} + \epsilon \dot{k}) \\ L(q, \dot{q}) &= L(q + \epsilon k, \dot{q} + \epsilon \dot{k}) = L(q, \dot{q}) + \epsilon \sum_i \dot{k}_i \frac{\partial L}{\partial \dot{q}_i} + \epsilon \sum_i k_i \frac{\partial L}{\partial q_i} \dots\end{aligned}$$

This just applies the constraint that the sum of the first n taylor expanded terms has to be zero, which is *Noethers Theorem*. Another, simpler way of writing this is that

$$\sum K \frac{\partial L}{\partial \dot{q}} = C$$

where  $C$  is constant.

#### Example



we just apply

$$L = \frac{1}{2}m\dot{l}^2 + \frac{1}{2}ml^2\dot{\theta}^2 - mgl \cos \theta$$

And derive properties from this, like conservation of angular momentum.

### 1.3 Probably the Hamiltonian

consider

$$\begin{aligned}H &= \dot{q} \frac{\partial L}{\partial \dot{q}} - L \\ \frac{dH}{dt} &= - \frac{dL}{dt}\end{aligned}$$

some long thing using the chain rule. has it simplify down to the above form, which implies that  $H$  is a conserved quantity with respect to time. Take

$$\begin{aligned}L &= \frac{1}{2}m\dot{x}^2 - V(x) \\ H &= \dot{x} \frac{dL}{d\dot{x}} - L = m\dot{x}^2 - (\frac{1}{2}m\dot{x}^2 + V(x)) = \text{total energy}\end{aligned}$$

which lets us say  $H$  is a total energy, the *hamiltonian*.