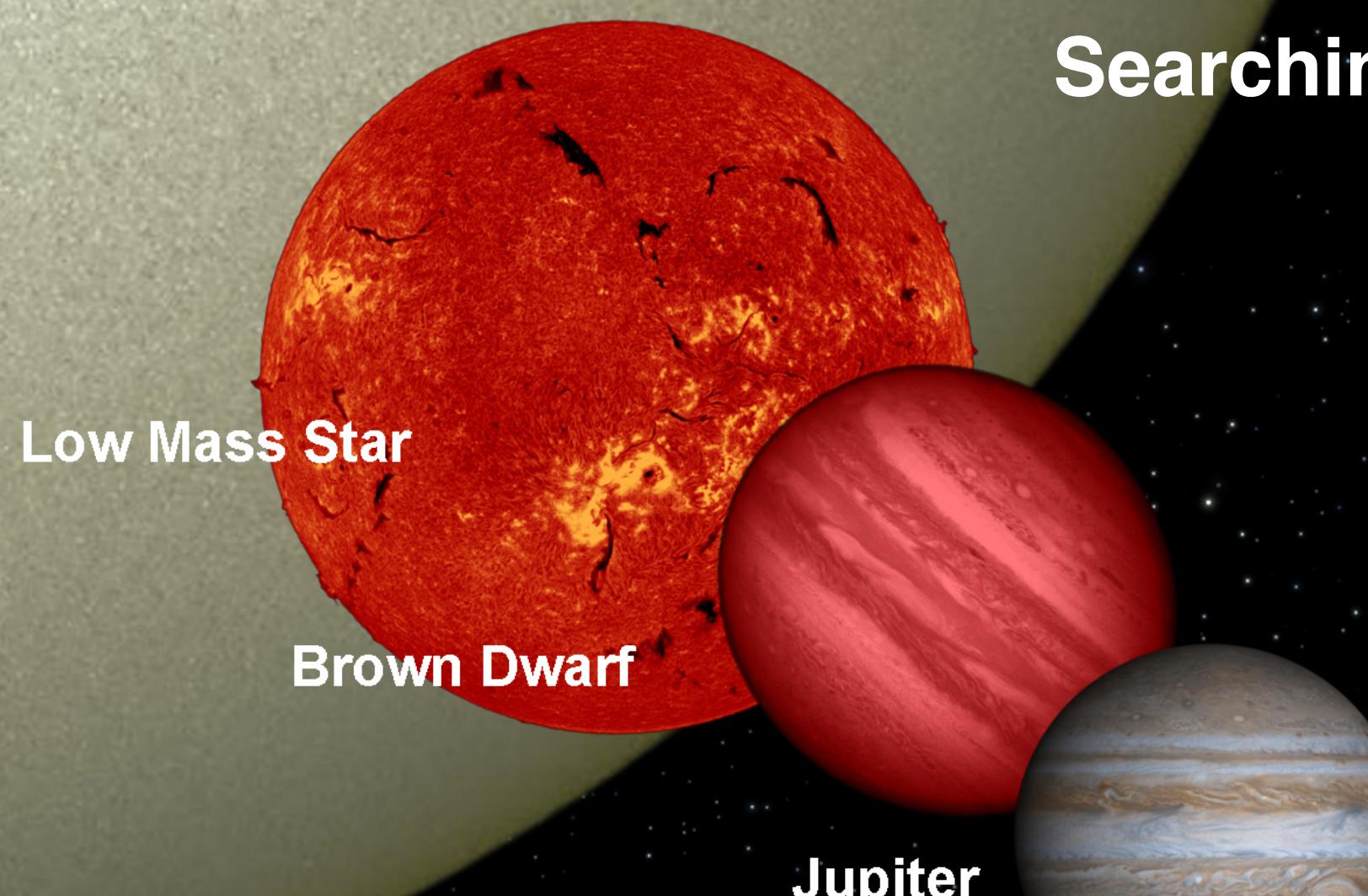


# Searching for Warm Dust around Cool Stars from the Sloan Digital Sky Survey

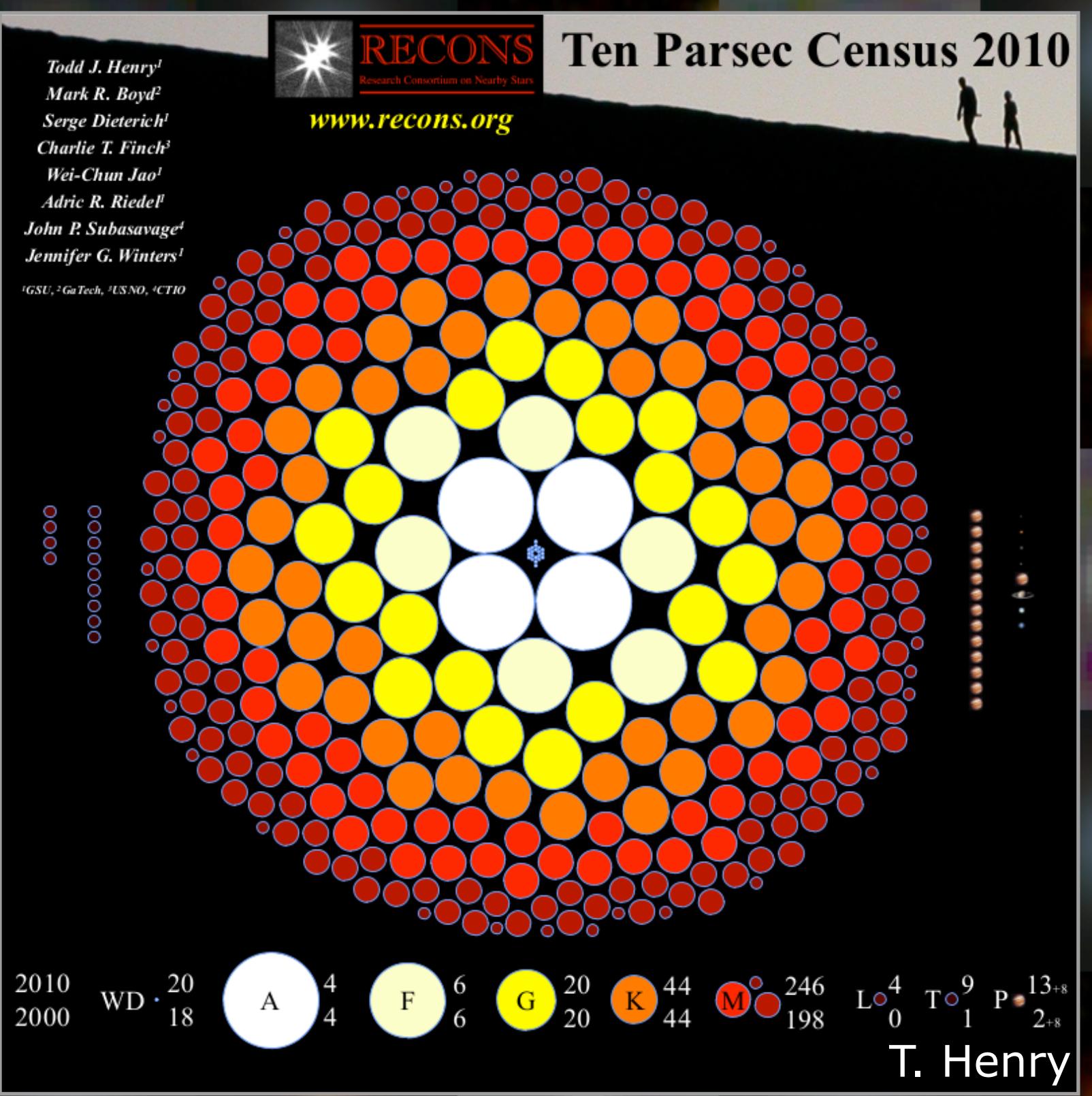
Sun



NASA

## Why low-mass stars?

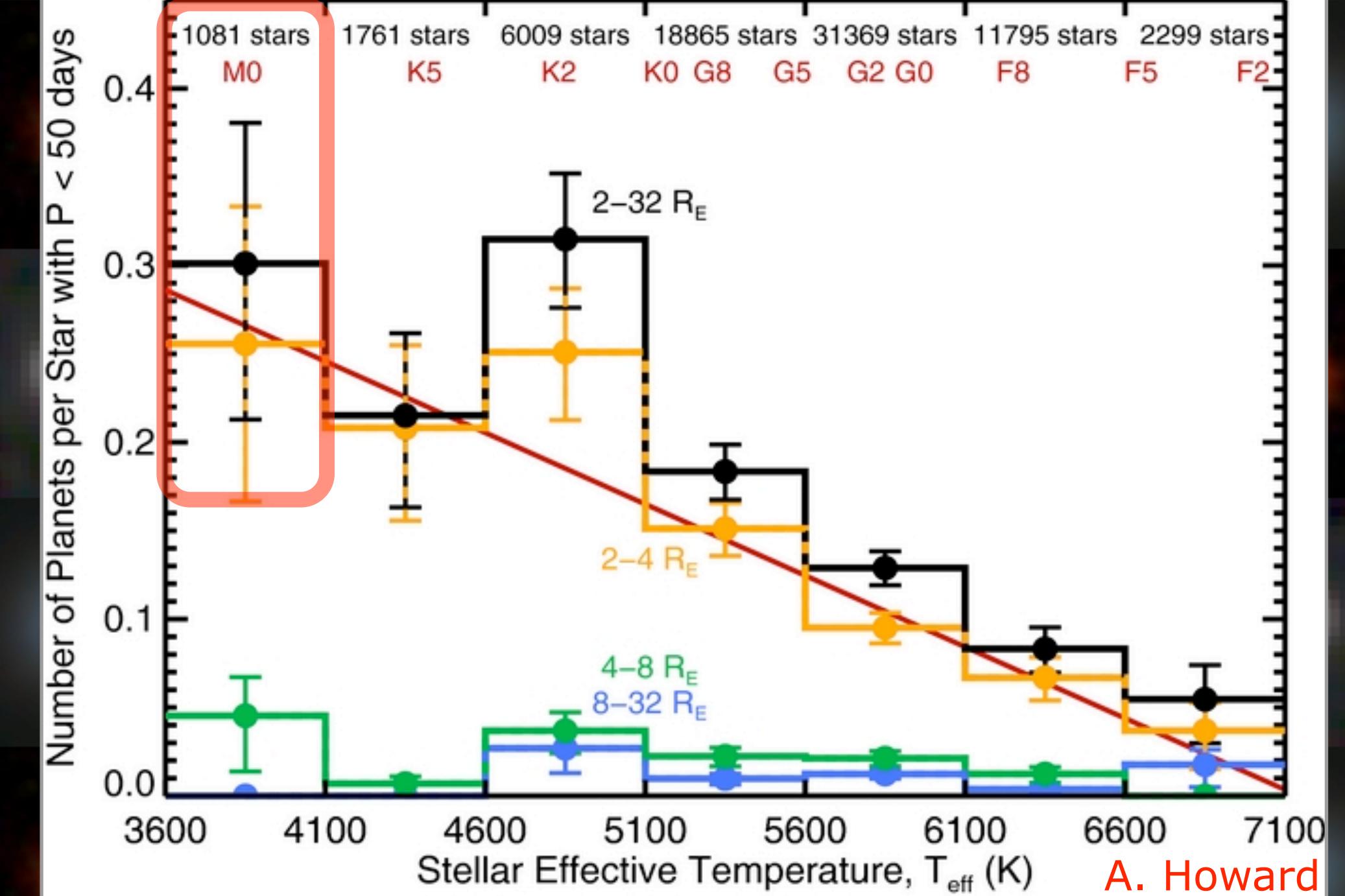
### What is in a star's size?



Large and small censuses of stars indicate that low mass stars are easier to form than high mass stars.

**Low-mass stars make up ~70% of the stellar population<sup>1</sup>.**

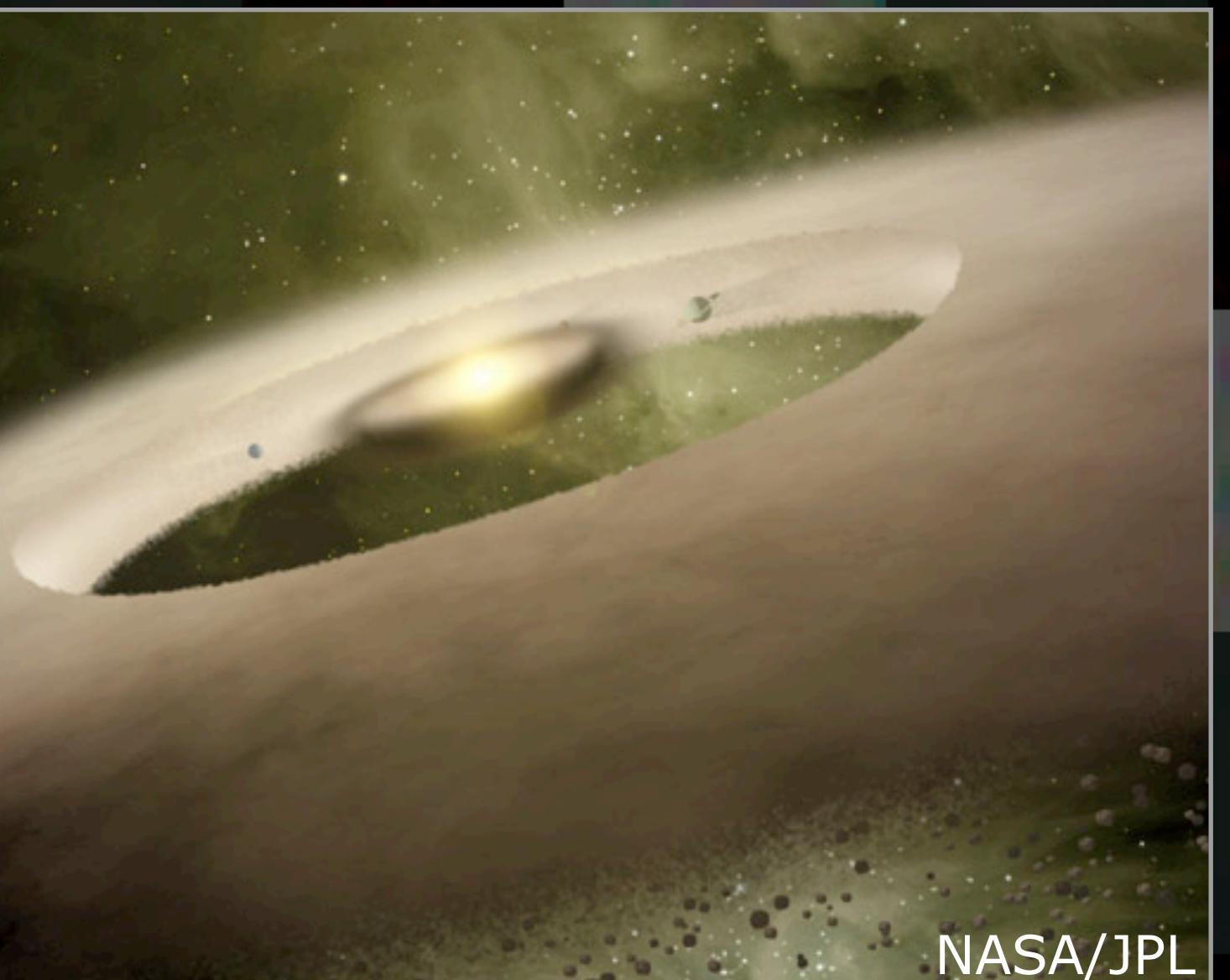
### Do we see planets around them?



**Low-mass stars have a high likelihood for hosting Earth sized planets<sup>3</sup>.**

### What are we looking for?

#### What can we find in the infrared?



Stars form from the collapse of a cloud of gas, which in turn creates a disk where planets form. Disks, and gaps within the disk, can be observed at infrared wavelengths.

**Planet formation signatures exist within the infrared wavelength regime.**

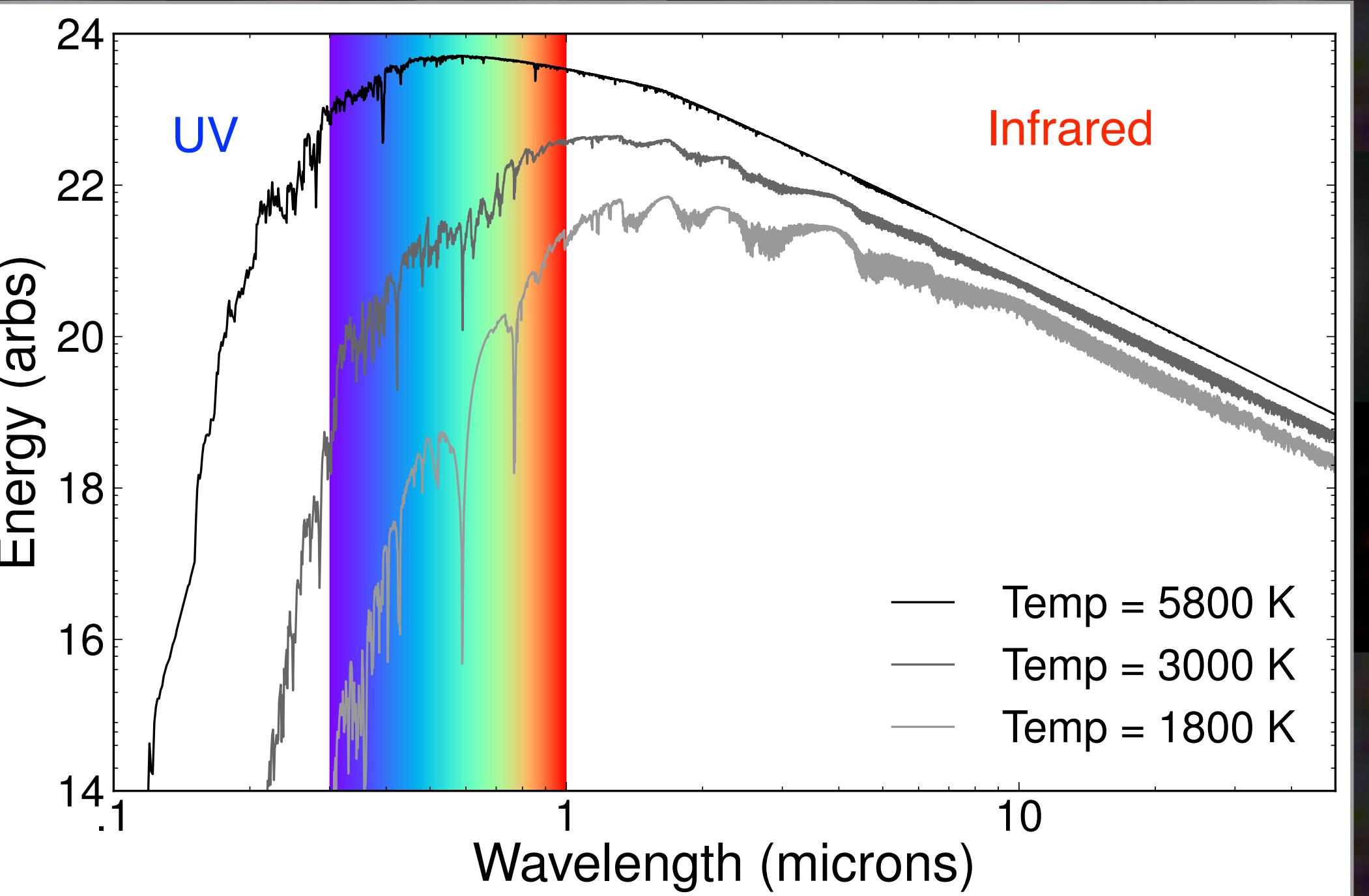
# Boston University

Christopher A. Theissen, Andrew A. West  
Department of Astronomy

**BOSTON**  
**UNIVERSITY**

## Where do we look?

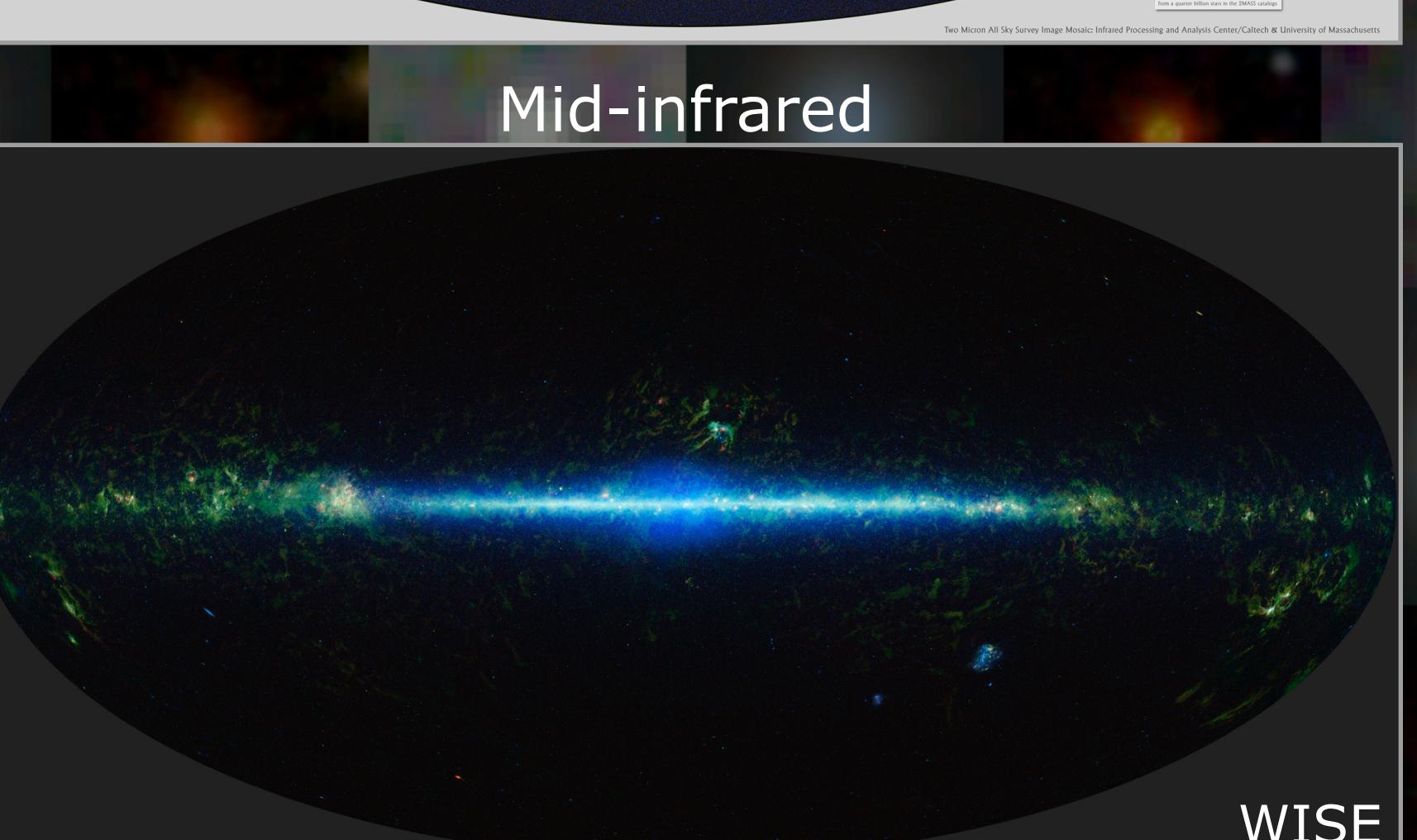
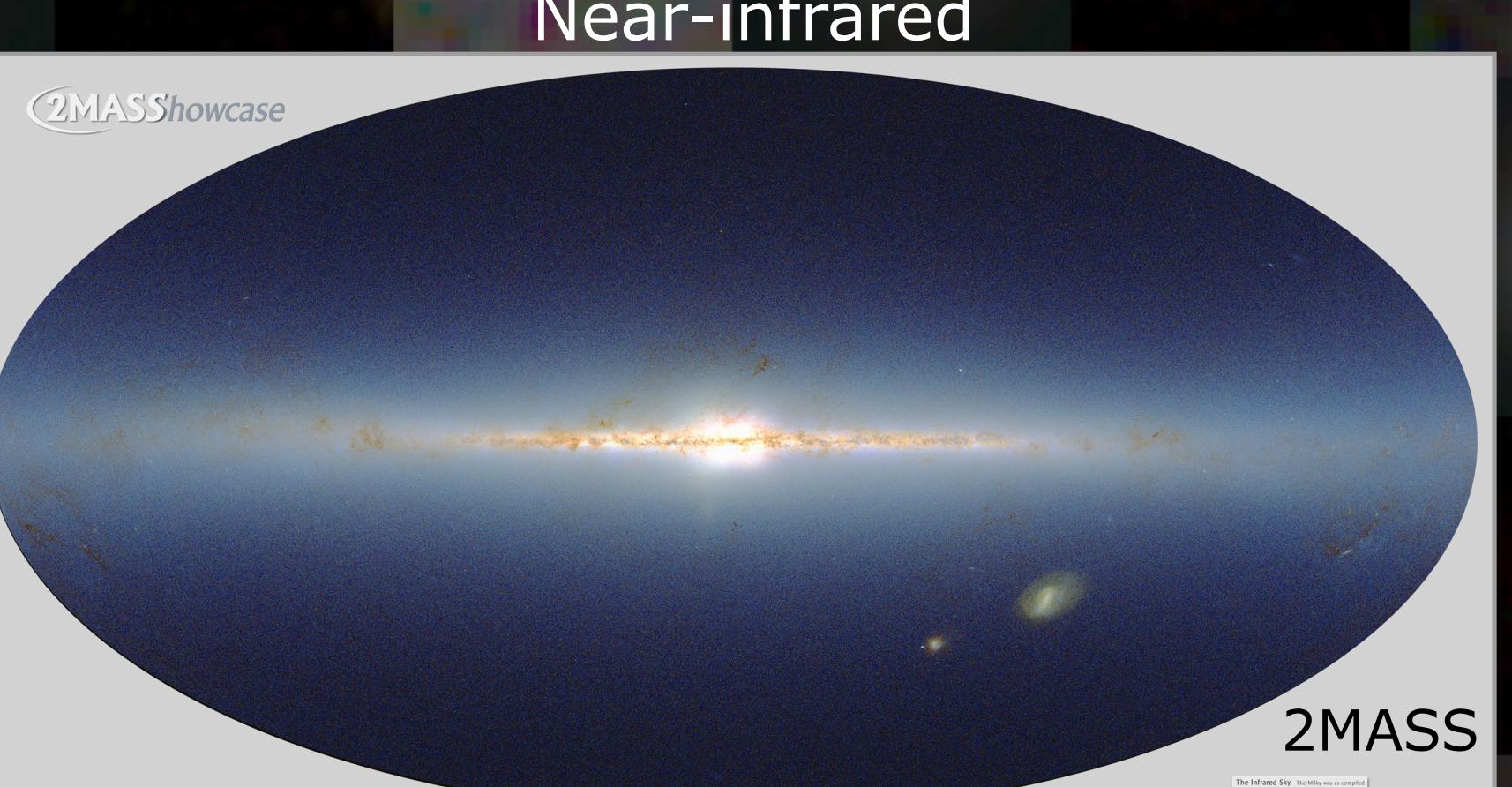
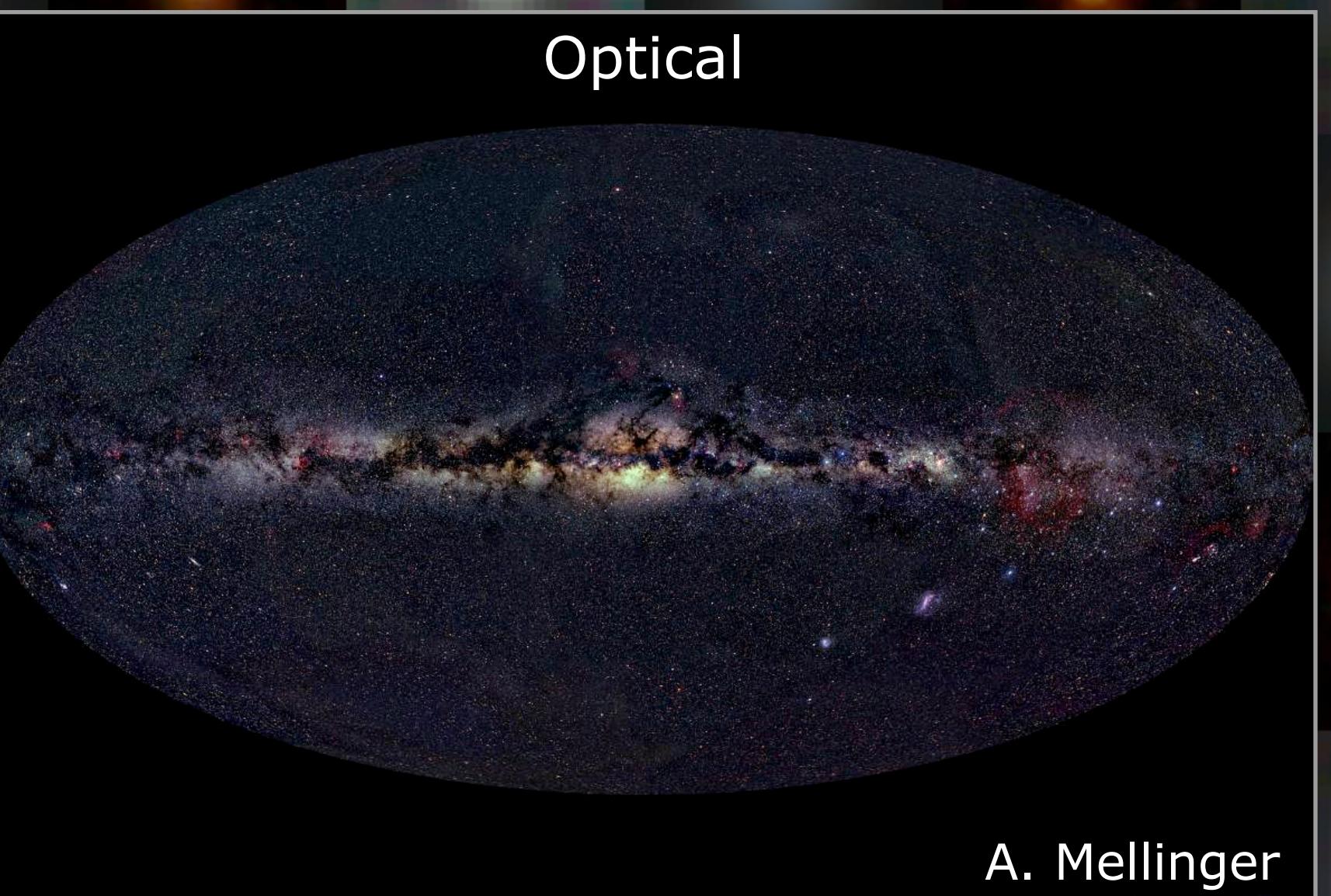
### Observational differences between stars?



Model stellar photospheres<sup>2</sup>. Photospheres are the part of the stellar atmosphere from which light we observe is produced. Temperature primarily defines the shape of the spectrum.

**M dwarfs are so cool that molecules begin forming in their atmospheres. This is extremely noticeable in the regime of visible light.**

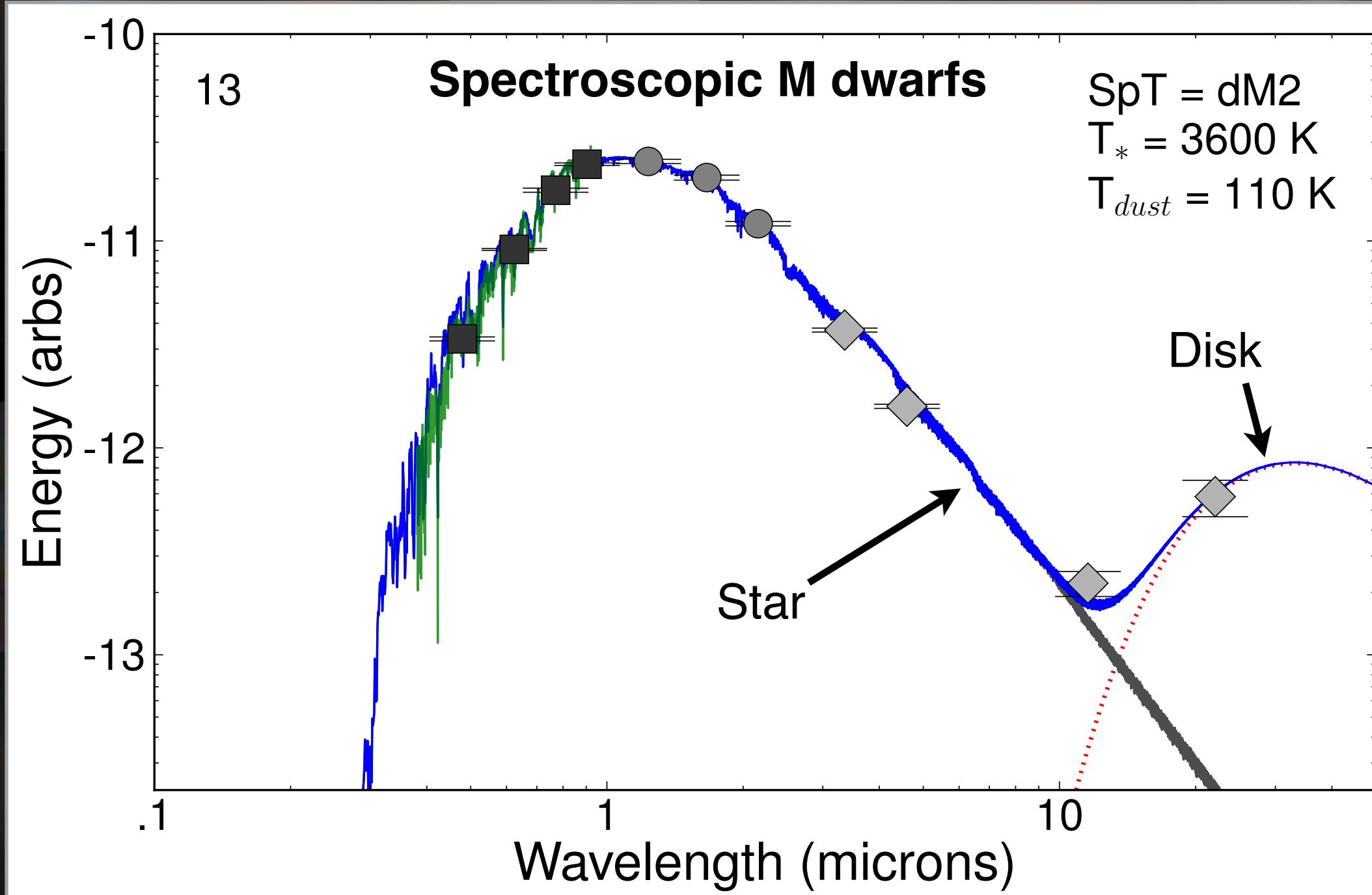
### What data is available to Astronomers?



**There is an abundance of data coming from all-sky surveys, both from space, and the northern and southern hemispheres of Earth.**

## What have we found?

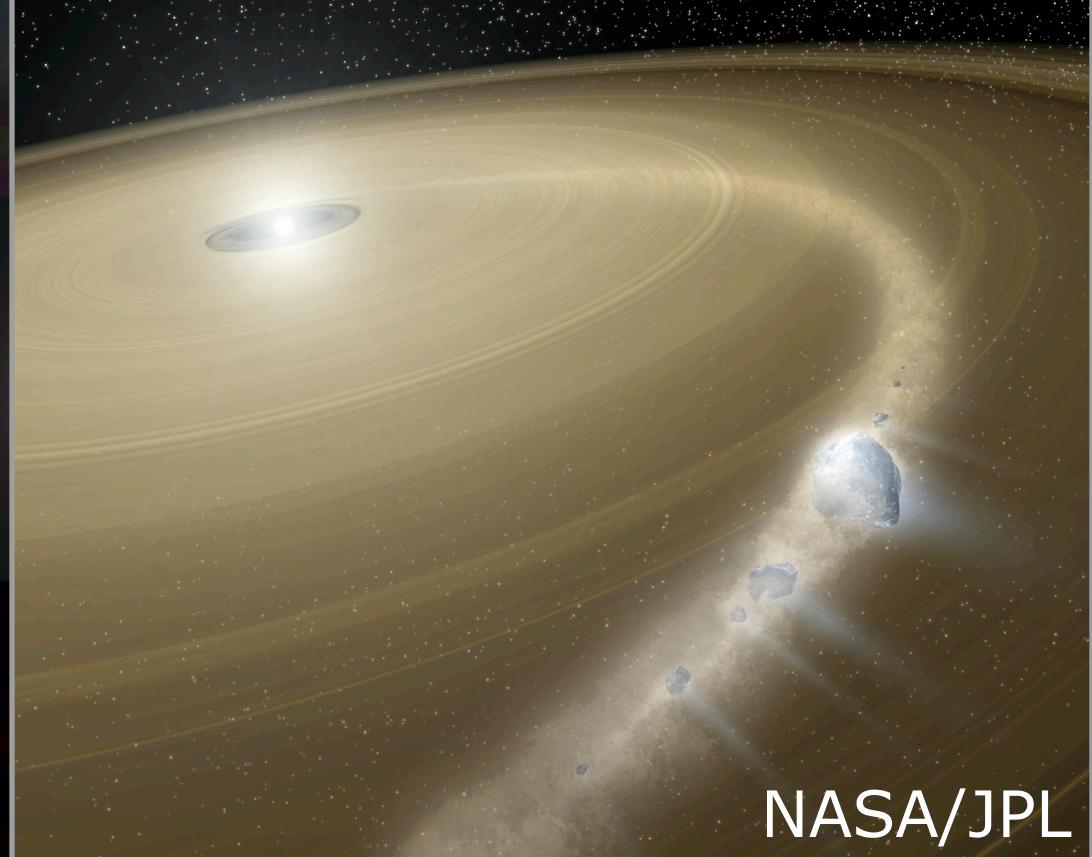
### 446 M dwarfs with signatures of dust



**These stars are mostly older stars (>1 billion years) that should not have disks.**

### What is the interpretation?

#### Possible scenarios include:

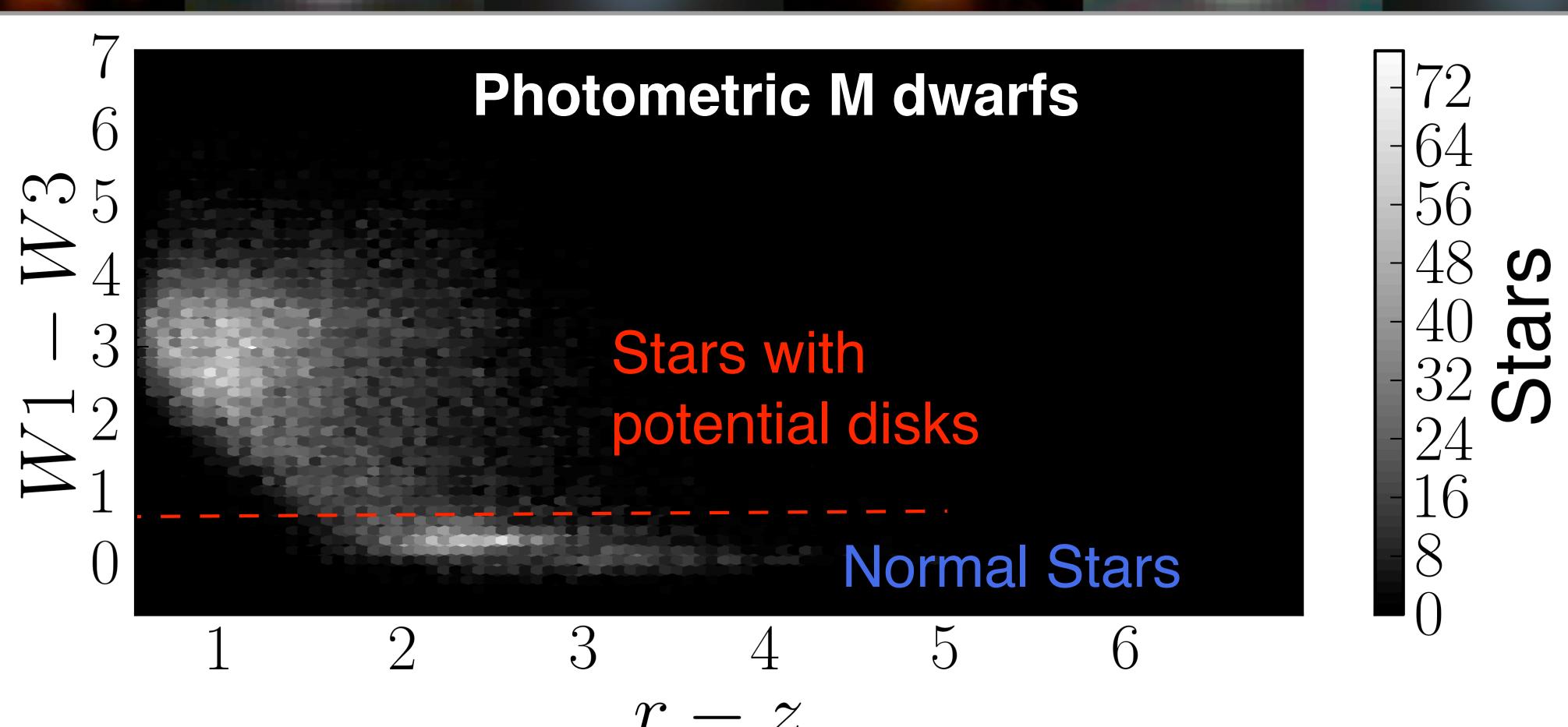


- 1) Planets or asteroids may stray too close to their host star and be ripped apart by the force of gravity, creating a disk of material similar to that found around young stars.
- 2) Failed planet formation, similar to our own Asteroid and Kuiper belt.
- 3) A second onset of planet formation.

**Some of these ideas are supported by the fact that low-mass stars have lifetimes longer than the current age of the Universe (~13.7 billion years).**

### What is the next step?

**There are still countless stars out there to study.**



### References

- [1] Bochanski, J. J., Hawley, S. L., Covey, K. R., et al. 2010, AJ, 139, 2679
- [2] Allard, F., Homeier, D., & Freytag, B. 2012, Royal Society of London Philosophical Transactions Series A, 370, 2765
- [3] Howard, A. W., Marcy, G. W., Bryson, S. T., et al. 2012, ApJS, 201, 15

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