

# **RHODIUM: A post-processor for the **BIGSTICK** configuration-interaction calculations**

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October 23, 2025

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## **Abstract**

This is the manual for **RHODIUM**, a postprocessing code to be used with **BIGSTICK**. It can be used to compute density matrices, spectroscopic amplitudes, and other information, using wave function and basis files created by **BIGSTICK**.

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# Chapter 1

## Introduction

BIGSTICK Johnson et al. [2018, 2013] is an efficient and powerful configuration-interaction code for tackling the many-fermion problem, primarily modeling atomic nuclei. In addition to finding the eigenpairs of the many-nucleon system, BIGSTICK can carry out a variety of additional tasks, such as computing expectation values, transition density matrices, and so on.

BIGSTICK has one key restriction, which both enables its efficiency but also limits the tasks it can carry out: it uses a basis with certain fixed quantum numbers, such as total  $M/J_z$  ( $z$ -component of angular momentum) and  $N$  and  $Z$  (number of valence neutrons and protons, respectively). Thus, BIGSTICK cannot compute spectroscopic amplitudes, which are matrix elements of adding or removing one nucleon, nor of charge-changing transitions such as  $\beta$ -decay (or only through some assumptions such as isospin rotation). Note that one has the option to fix parity or to allow for both parities simultaneously.

RHODIUM addresses this limitation. RHODIUM reads in files generated by BIGSTICK, but unlike BIGSTICK, it can carry out operations between wave functions computed in two different bases with different quantum numbers. In particular, RHODIUM can:

- Compute the one-body density between wave functions (the name RHODIUM refers both to the symbol  $\rho$  for density matrices, as well as being a *dense* metal), even if the bases have different  $M$  and/or different parity  $\pi$ ;
- Apply a general one-body operator to a wave function (e.g., E1, E2, or  $\hat{J}_+$ ) or set of wave functions. The final state(s) can be in a basis with different quantum numbers, e.g.,  $M$  or  $\pi$ , from the initial state(s). They can also have different  $T_z$ .
- Compute the spectroscopic amplitudes between two wave functions with different particle number;
- Add or remove a particle or superposition of particles to a wave function;

- Compute the overlap between two wave functions with different many-body truncations;
- Project a wave function existing in one basis into another basis with a different truncation;

and so on.

The resulting wave functions are written back as files in a format usable by BIGSTICK. One can summarize the difference between BIGSTICK and RHODIUM as BIGSTICK is much more efficient, but limited to a fixed basis, while RHODIUM is more flexible, able to connect two different bases, but less efficient. RHODIUM does not diagonalize a Hamiltonian and currently does not deal with two-body operators.

## 1.1 Current version

The current version of the code this manual has been updated for is 1.0.0.

**NOTE:** This manual is still in progress. Some options may not be fully addressed.

## 1.2 License

This code is distributed under the MIT Open Source License.

The source code and sample inputs are found at [github.com/cwjsdsu/BigstickPublick](https://github.com/cwjsdsu/BigstickPublick).

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## Chapter 2

# Compiling the code

The code can be compiled with either the intel `ifort` or the GNU `gfortran` compiler. It does not require external libraries or packages; it carries with it all the libraries it needs.

RHODIUM has good OpenMP parallelism. We have partial MPI parallelism, but in its current version it is not yet fully implemented.

You can find the compile options by

```
make help
```

The current options are:

Here are some compile options:

Note default compiler is intel ifort compiler

```
-----  
default compiler is intel ifort
```

```
make serial -- default serial, rhodium.x
```

```
make openmp -- OpenMP parallelism, rhodium-openmp.x
```

```
make mpi -- MPI parallelism, rhodium-mpi.x
```

```
make openmp-mpi -- hybrid OpenMP+MPI, rhodium-mpi-omp.x  
-----
```

```
make gfortran -- serial with GNU compiler gfortran, rhodium.x
```

```
make gfortran-openmp -- OpenMP with GNU compiler gfortran, rhodium-openmp.x
```

Currently, the recommended compile option is the last one: `make gfortran-openmp`.

In the near future we plan to fully implement MPI distributed memory parallelism.

## Chapter 3

# Required input

Here I discuss the required input.

To prepare for a calculation, **RHODIUM** needs files with basis information for both the initial and final state(s). These are files with extension **.bas** and are created by **BIGSTICK** using the initial menu option '**(b)**'. If the initial and final spaces are the same, you may reuse the same **.bas** file.

In all applications you need an initial wave function file with extension **.wfn**; this can be generated by **BIGSTICK**. In some applications, **RHODIUM** will create a new final wave function file, which can then either be read back into **BIGSTICK** or perhaps reused by **RHODIUM**. In many applications, but not all, you will also need a final wave function file. **The initial and final wave functions files must be different files**; if the information is the same, you must make a copy with a different name. (This is because both the initial and the final wave function files are kept open simultaneously.)

### 3.1 Creating a basis file with BIGSTICK

Here is how a **BIGSTICK** session creating a basis looks like:

```
Enter choice
b
Create basis (.bas) file for later postprocessing

Enter file with s.p. orbit information (.sps or .sp)
(Enter "auto" to autofill s.p. orbit info for no-core shell model )
(Enter "?" for more information )
sd
Enter # of valence protons (max 12 ), neutrons (max 12)
4 4
Enter 2 x Jz of system
0
```



```

.... Building basis ...

Information about basis:
          495  SDs for species          1
          495  SDs for species          2
Total basis =                28503

.... Basis built ...
This will write information on the basis useful for postprocessing
(Written as a binary file, much like a .wfn file)
Enter name of file (without extension )
mg24
  assuming only a single fragment
  finished writing proton SDs
  finished writing neutron SDs
  Finished writing to file

```

This creates a binary file `mg24.bas`. (If you want to have an explicit, human-readable file that includes the basis, you can use the option ‘(t)’ in BIGSTICK to create a `.trwfn` file. That file is not usable in RHODIUM; you can, however, use it to write your own postprocessing code.

## 3.2 Creating a wave function file

There are multiple ways to get to a wavefunction file with extension `.wfn`. The first and simplest way is to carry out a default or normal run in BIGSTICK with initial menu option ‘(n)’. Any file with extension `.wfn`, whether created by BIGSTICK or RHODIUM, should be usable by RHODIUM (or can be fed back into BIGSTICK). The main difference is that RHODIUM requires `.bas` basis files, while BIGSTICK ‘knows’ how to construct the basis.

## 3.3 Other inputs

Other inputs are created by the user as needed, described in subsequent chapters. These are files that include matrix elements of operators applied to a wave function, such as a `.opme` file for one-body operators, or a `.spa` file for single-particle addition/removal.

## Chapter 4

# Density matrices

Much of what RHODIUM does is computing ‘density matrices,’ which I now discuss in some depth. This chapter only discusses the formal definition(s) of density matrices. Later chapters will guide you through computing them using RHODIUM.

There are many objects called ‘density matrices’ in quantum systems. Here we deal with one-body density matrices, which generically have the form

$$\rho_{ab} \sim \langle \Psi_f | \hat{c}_a^\dagger \hat{c}_b | \Psi_i \rangle, \quad (4.1)$$

where  $\hat{c}_a^\dagger, \hat{c}_b$  are nucleon creation and annihilation operators, respectively, creating and destroying nucleons in orbitals  $a$  and  $b$ . If  $|\Psi_f\rangle \neq |\Psi_i\rangle$ , this is a transition density matrix. Transition density matrix elements can be used, for example, for computing  $\beta$ -decay and  $\gamma$ -decay probabilities. On the other hand, if  $|\Psi_f\rangle = |\Psi_i\rangle$ , then one can extract either static properties, such as magnetic dipole (M1) or electric quadrupole (E2) moments, or elastic scattering, say of electrons, neutrinos, or dark matter.

If the initial and final numbers of protons and neutrons are the same, then one can compute (charge-conserving) density matrices.

For one-body operators, one generally defines the density matrix as

$$\rho_K^{f,i}(a,b) = \frac{\langle \Psi_f : J_f || [\hat{c}_a^\dagger \otimes \tilde{c}_b]_K || \Psi_i : J_i \rangle}{\sqrt{2K+1}}. \quad (4.2)$$

This means the initial state  $\Psi_i$  has angular momentum  $J_i$ , the final state  $\Psi_f$  has angular momentum  $J_f$ , and the nucleon creation/annihilation operators are coupled up (through a Clebsch-Gordan coefficient) to total angular momentum  $K$ . The indices  $a, b$  label single-particle orbitals, such as  $0d_{5/2}$ . The notation  $\tilde{c}$  denotes a “time-reversed” operator:

$$\tilde{c}_{j,m} = (-1)^{j+m} \hat{c}_{j,-m} \quad (4.3)$$

This is necessary to make the operators transform as “spherical tensors” (see Edmonds [1996] for details) as well as to correctly preserve selection rules. Fi-

nally the reduced matrix element (see Edmonds) is defined by

$$\langle J_f || \hat{O}_K || J_i \rangle = (-1)^{J_f - M_f} \frac{\langle J_f M_f | \hat{O}_{K, M_K} | J_i M_i \rangle}{\begin{pmatrix} J_f & K & J_i \\ -M_f & M_K & M_i \end{pmatrix}} \quad (4.4)$$

$$= (-1)^{2K} \sqrt{2J_f + 1} \frac{\langle J_f M_f | \hat{O}_{K, M_K} | J_i M_i \rangle}{(J_i M_i, K M_K | J_f M_f)}. \quad (4.5)$$

Note that the phase  $(-1)^{2K}$  in Eq. (4.5) only has an effect if  $K$  is half-integer, that is, the operator  $\hat{O}$  has an odd number of fermion operators, for example to create or destroy a particle. The reduced matrix element arises from the Wigner-Eckart theorem, which states that any dependence upon orientation, that is, initial  $M_i$  and final  $M_f$ , must be proportional to a Clebsch-Gordan coefficient. With the above definition, a useful symmetry relation is

$$\rho_K^{if}(b, a) = (-1)^{j_a - j_b + J_i - J_f} \rho_K^{fi}(a, b). \quad (4.6)$$

The reason for the definition Eq. (4.2) is to make easy computing the reduced matrix element of an arbitrary one-body operator:

$$\langle \Psi_f || \hat{O}_K || \Psi_i \rangle = \sum_{ab} \rho_K^{(fi)}(a, b) \langle a || \hat{O} || b \rangle. \quad (4.7)$$

A number of post-processing codes use one-body densities, for example to compute electromagnetic transitions.

Though BIGSTICK computes density matrices very efficiently, it has limitations. For example, BIGSTICK works in the  $M$ -scheme, that is, a fixed  $J_z$  or  $M$  basis. Therefore, to compute density matrices, BIGSTICK actually computes  $\langle J_f M_f | \hat{O}_{K, M_K} | J_i M_i \rangle$  and then divides by the appropriate Clebsch-Gordan coefficient to obtain the reduced density matrix element. Some Clebsch-Gordan coefficients vanish, however, usually due to some symmetry. In that case, the associated reduced density matrix element cannot be computed. This mostly happens for  $M = 0$ . Specifically,

$$(J_i M_i, K M | J_f M_f) = (-1)^{J_i + K_i - J_f} (J_i - M_i, K - M | J_f - M_f) \quad (4.8)$$

When  $M_i = M_f = M = 0$ , then implies that the Clebsch-Gordan coefficient must vanish when  $J_i + K - J_f$  is odd.

One inelegant solution is to simply re-run BIGSTICK with  $M \neq 0$ . For large cases this can be time-consuming. Furthermore, there can be arbitrary signs introduced. RHODIUM fixes this problem by applying a raising operator  $\hat{J}_+$  to change an  $M = 0$  state to a  $M = 1$  state. This breaks the symmetry and allows one to extract the matrix elements missing due to vanishing Clebsch-Gordan coefficients.

**It's important to remember that we must have  $K \geq |M| = |M_i - M_f|$ .** RHODIUM will automatically skip such density matrix elements.

Another limitation of **BIGSICK**, and addressed in **RHODIUM**, is the computation of density matrix elements between levels of opposite parity. **BIGSTICK** allows one to choose positive (+) or negative (−) parity, or both, (‘0’). Computing both parities simultaneously allows one to extract cross-parity matrix elements, but at the price of a larger (roughly  $2\times$ ) basis dimension. **RHODIUM** allows one to compute density matrix elements directly between states of different parities.

## Chapter 5

# Running RHODIUM

No matter what action you are taking, the first step is to read in the basis information files, with extension `.bas`, created by `BIGSTICK` with initial menu option `'(b)'`. Basis files are need for both the initial and final wave functions, but these can be the same. `RHODIUM` provides some confirming information, such as the basis dimension, valence number, and so on. (Note: as of version 0.9.1, `RHODIUM` can now utilize the hole formalism available in `BIGSTICK`, where entering, e.g., `'-4'` means four holes in a space. One must, however, use it consistently, that is, if one space uses the hole formalism, the other space must as well.)

```
Enter name of INITIAL .bas file
mg24
Basis file successfully opened
about to read in basis
dimbasischeck=          28503
Valence Z, N =          4          4
Single particle space :
  N    L  2xJ
  0    2    3
  0    2    5
  1    0    1
....
DONE READING IN
Enter name of FINAL .bas file
mg24
Basis file successfully opened
dimbasischeck=          28503
Valence Z, N =          4          4
Single particle space :
...
```

Depending upon the particle numbers, `RHODIUM` will automatically provide ap-

propriate options.

For example, if both initial and final spaces have the same number of protons and neutrons, here are the options provided:

Choose one of the following menu options

```
(dotwfn) Dot product between wave functions in different bases (different truncations)
(dnliso) 1-body density matrices, non-charge changing, good isospin
(dnlxpn) 1-body density matrices, non-charge changing, explicit proton-neutron
(entlbd) 1-body entropy
(applbd) Apply non-charge-changing 1-body operator to wave functions and write out
(projct) Project a state into a basis with different truncation
(lincom) construct linear combinations of wave functions
```

Enter the six-character code for your choice::

If the total number of nucleons is conserved, but  $N$  and  $Z$  both change by 1, RHODIUM provides charge-changing options:

```
(dn1bcc) Charge-changing 1-body densities (pn)
(dl1scc) Charge-changing 1-body densities (iso)
(applcc) Apply charge-changing 1-body operator to wave functions and write out
```

Finally, if there is a difference by one in  $Z$  or  $N$ , but not both, the options are

```
(spamp1) 1-particle spectroscopic amplitude (change in number of particles)
(appsp1) Add/remove a single particle (linear combination)
```

In the future additional options will be added, such as two-nucleon spectroscopic factors, etc.

## 5.1 Dot products of wave functions

Option ‘(dotwfn)’ allows one to take the inner product (or dot product or overlap) between two wave functions,  $\langle\psi_f|\psi_i\rangle$ . This can also be done using BIGSTICK with initial menu option ‘(v)’, but that requires both wave functions be in the same basis. In RHODIUM, one can compute the dot product as long as both wave functions have the same  $J_z$  (or  $M$ ) and parity; they can be embedded in different Hilbert spaces, such as different truncations. One could, for example, find the overlap of a no-core shell-model  $N_{\max} = 6$  wave function with one with  $N_{\max} = 10$ .

## 5.2 Projection of a state into a basis with a different truncation

Related to dot products, menu option ‘(projct)’ allows one to project wave functions into a different Hilbert space, as long as the initial and final bases

have the same  $J_z$  (or  $M$ ) and parity. This is useful, for example, to project states from a smaller space into a much larger space. This can be useful for bootstrapping a block Lanczos calculation in **BIGSTICK**, by using a block of wave functions computed in a small space as a pivot block (initial vectors) for a calculation in a larger space.

The output of this option is a **BIGSTICK**-compatible file with extension `.wfn`.

### 5.3 Charge-conserving one-body density matrices

If valence  $N$  and  $Z$  are the same, you get these options always (even if  $M$  and parity  $\pi$  are not the same):

```
(dnliso) 1-body density matrices, non-charge changing, good isospin
(dn1xpn) 1-body density matrices, non-charge changing, explicit proton-neutron
```

These options assume the sets of initial and final states are different. Indeed, you **cannot** use the same `.wfn` files when choosing these options; if you want the same data, you must make a copy with a different name, but the same extension. (Otherwise one can simply use **BIGSTICK** to generate the density matrices.) The output will be a `.dres` file.

You will be given an option to combine the two sets of wave functions into a single list:

Do you want to combine initial/final levels into a single list? (y/n)

If you choose ‘n’, then the output will be a `.dres` file, which lists both the initial and final sets of states, e.g.,

```
# Wavefunctions -- initial      4      4
State      E      Ex      J      T      par
  1      -87.10445    0.00000    -0.000    0.000    1
  2      -85.60215    1.50230     2.000    0.000    1
  3      -82.98830    4.11615     2.000    0.000    1
  4      -82.73202    4.37243     4.000    0.000    1
  5      -82.03408    5.07037     3.000    0.000    1
#
# Wavefunctions -- final      4      4
  1      -92.77905    0.00000    -0.000    0.000    1
  2      -91.11964    1.65940     2.000    0.000    1
  3      -88.47786    4.30119     2.000    0.000    1
  4      -87.97810    4.80094     4.000    0.000    1
  5      -87.43483    5.34422     3.000    0.000    1
  6      -86.54257    6.23647     4.000    0.000    1
  7      -84.79132    7.98772    -0.000    0.000    1
  8      -84.54140    8.23765     2.000    0.000    1
```

This is slightly different from the standard BIGSTICK output `.dres` file, which provides a single lists of levels. If you choose ‘y’, that is, a single combined list, then the output `.dres` file will generally be compatible with various postprocessing codes.

The output `.dres` file also includes a convenient list of single-particle orbitals.

```

Single particle state quantum numbers
ORBIT      N      L      2 x J
    1      0      2      3
    2      0      2      5
    3      1      0      1

```

The output densities are in the same format as from BIGSTICK, either in isospin format

```

Initial state #    1 E = -87.10445 2xJ, 2xT =    0    0
Final state  #    1 E = -92.77905 2xJ, 2xT =    0    0
Jt =    0, Tt = 0      1
    1    1    0.3876706    0.0000000
    2    2    1.7379967    0.0000000
    3    3    0.4368686    0.0000000

```

or in explicit proton neutron format,

```

Initial state #    1 E = -87.10445 2xJ, 2xT =    0    0
Final state  #    1 E = -92.77905 2xJ, 2xT =    0    0
Jt =    0, proton      neutron
    1    1    0.2741245    0.2741245
    2    2    1.2289493    1.2289492
    3    3    0.3089128    0.3089128

```

In both these cases, the initial wave vectors are drawn exclusively from the list of initial levels, and final from final.

Here is a sample script for obtaining cross-parity (i.e., between positive and negative parity levels). Note that one needs basis files for both parities, as well as wave function files for both parities.

```

n24pos.m0    ! INITIAL BASIS FILENAME
n24neg.m0    ! FINAL BASIS FILENAME
dn1xpn      ! MENU CHOICE: proton-neutron 1-body densities
y           ! combine positive, negative levels into one list
n24pos2neg   ! OUTPUT FILE NAME
n24pos      ! INITIAL WFN FILENAME
1 500       ! selected initial levels
n24neg      ! FINAL WFN FILENAME
1 500       ! selected initial levels

```

Note time-reversed densities, that is, both positive-to-negative and negative-to-positive, are automatically generated.



### 5.3.1 Generating ‘missing’ density matrix elements

In computing reduced density matrices, one must divide by a Clebsch-Gordan coefficient or an equivalent Wigner 3- $j$  symbol. However in some cases these vanish due to symmetries, as explained above. This means one is missing one-body density matrix elements  $\langle J_f || [\hat{c}^\dagger \otimes \hat{c}]_K || J_i \rangle$  when  $(-1)^{J_i+J_f+K} = -1$ . (Note that here we use  $K$  for the angular momentum rank of the operator to be consistent with the definition of reduced density matrices in Eq. (4.5); however in `.dres` file this is written as `Jt`).

RHODIUM allows you to retrieve these ‘missing’ matrix elements without re-running BIGSTICK. By applying the angular momentum raising operator,  $\hat{J}_+$ , RHODIUM can break the troublesome symmetry.

Recall this only happens if  $M = 0$  (and hence an even number of particles). If the initial basis and the final basis differ by exactly one unit of  $M$ , then you will see three additional options:

```
(jraise) Apply J+ raising operator to wave functions and write out
(dn1jrm) Compute missing density matrix elements by applying J+ raising
(dn1mis) Compute missing density matrix elements
```

When seeking missing density matrix elements, there are two cases: first, when the list of initial and final states are the same, and second, when they are different. In practice, these two scenarios happen if the initial and final states have the same parity, or the opposite parities, respectively.

If the initial and final sets of states (levels) are the same, i.e., are drawn from the same parity, then use option ‘`(dn1jrm)`.’ In this case, the original levels have been generated with  $M = 0$ . You will need to create basis files with extension `.bas` with both  $M = 0$  and  $M = 1$ . With option ‘`(dn1jrm)`’, you will enter the filename for the wavefunctions computed with  $M = 0$ ; RHODIUM will then automatically apply the angular momentum raising operator  $\hat{J}_+$ , converting internally from  $M = 0$  to  $M = 1$  (levels with  $J = 0$  which, after application of  $\hat{J}_\pm$ , have zero norm and will be excluded; they don’t contribute to missing matrix elements at any rate), and then compute a `.dres` file with only the ‘missing’ matrix elements, namely those that satisfy  $(-1)^{J_i+J_f+K} = -1$ .

Here is a sample script for obtaining ‘missing’ density matrix elements when working from the same set of states (i.e., same parity):

```
n24pos.m0    ! INITIAL (INPUT) BASIS FILENAME
n24pos.m1    ! FINAL (TEMPORARY) BASIS FILENAME
dn1jrm       ! MENU OPTION
n24pos       ! INPUT FILENAME this is M = 0
1 500        ! selected levels
n24posmis    ! OUTPUT FILENAME
```

Note that one needs the basis files for  $M = 0, 1$ , but one only reads in the  $M = 0$  wave function file. The  $M = 1$  basis file is needed as a reference, but the  $M = 1$

wave functions are automatically generated (and written temporarily to the file `TEMP.wfn`) and then discarded.

If the initial and final sets of states (levels) are different, for example if they are drawn from different parities, then there is a two-step process: you must run `RHODIUM` twice. This is because there are *three* bases involved.

Let's assume you start with two `.wfn` files, one with basis  $+$ ,  $M = 0$  and the other with basis  $-$ ,  $M = 0$ . In order to compute the missing density matrix elements, we have to put one of those sets of wave functions into a basis with  $M = 1$  (one could select instead  $M = -1$ ; the choice does not matter), by applying the angular momentum raising operator  $\hat{J}_+$ , as

$$\hat{J}_+|M\rangle \propto |M+1\rangle. \quad (5.1)$$

Let's assume we choose to raise the  $-$  parity state; the choice does not matter.

To begin the process, one has to generate three `.bas` files:  $+$ ,  $M = 0$ ;  $-$ ,  $M = 0$ ; and  $-$ ,  $M = 1$ . In the first step, in `RHODIUM` take  $-$ ,  $M = 0$  as the initial basis and  $-$ ,  $M = 1$  as the final basis, and select option '`(jraise)`'. This option will request the input `.wfn` file of wave vectors in the  $-$ ,  $M = 0$  basis and then write out a new file those same wave vectors in the  $-$ ,  $M = 1$  basis. The states will be normalized. Any state with  $J = 0$  has zero norm after raising (or lowering) and is ignored.

In the second step, use the option '`(dn1mis)`' to compute the missing density matrix elements between the states in the  $+$ ,  $M = 0$  basis and the newly constructed states in the  $-$ ,  $M = 1$  basis; the order of initial and final does not matter. When asked

Do you want to combine initial/final levels into a single list? (y/n)

answer 'y'. This will produce a `.dres` file in the standard format, but only containing the troublesome missing density matrix elements.

Here are sample scripts. For the first step, raising the negative parity states from  $M = 0$  to  $M = 1$ . (Note: at the time of this writing, unfixed bugs require that the negative parity states get raised.)

```
negbasis.M0
negbasis.M1
jraise          ! MENU CHOICE
inputwfn        ! must be negative parity, M = 0
1, 500
outputwfn       ! will be negative parity, M = 1
```

The second step is to find the 'missing' matrix elements, generated between positive parity,  $M = 0$  states and negative parity,  $M = 1$  states constructed in the previous step:

```

n24pos.m0
n24neg.m1
dn1mis          ! MENU CHOICE
y               ! COMBINE BOTH SETS OF LEVELS INTO ONE LIST
n24neg2posmis   ! OUTPUT FILENAME
n24pos          ! M = 0 WFN FILE
1 500
n24negm1        ! M = 1 WFN FILE (from previous step)
1 500

```

Again, time-reversed densities are automatically constructed. **As stated before, it's important to remember that we must have  $J_t \geq |M| = |M_i - M_f|$** , where  $J_t$  is the angular momentum rank of the transition matrix elements. RHODIUM will automatically skip such density matrix elements.

While these steps can seem complicated, they ultimately save compute time by reducing the costly Lanczos runs in BIGSTICK to generate the wave functions. These procedures also eliminate ambiguities of phase and convergence, the latter of which can happen when a large number of states are generated.

## 5.4 Applying a (non-charge-changing) one-body operator

This option is very similar to the option ‘(o)’ in BIGSTICK: the code reads in a wave function `.wfn` file, and a `.opme` file of reduced matrix elements specifying a one-body operator  $\hat{O}$ , then applies  $\hat{O}$  to each of the wave vectors, and finally write all the so modified wave functions to a new `.wfn` file.

The difference here is that the final states can have different basis quantum numbers from the initial states. For example, an E1 operator could be applied to + parity states to generate - parity states. As another example, a raising operator could take states from  $M = 0$  to  $M = 1$  (though RHODIUM already has an option, `(jraise)`, which does this automatically without having to read in the raising operator matrix elements).

## 5.5 Charge-changing one-body density matrices

If the total number of nucleons is conserved, but  $N$  and  $Z$  both change by 1, RHODIUM provides these options:

```

(dn1bcc) Charge-changing 1-body densities (pn)
(d1iscc) Charge-changing 1-body densities (iso)

```

These allow you to compute charge-changing one-body density matrix elements, for example for beta decay, without assuming isospin is a good quantum number. In this case, the first option is preferred.

Note in this case, especially using the `pn` option, to compute beta-decay matrix elements one cannot use the usual code `gtstrength` but rather the `pntrans` code.

## 5.6 Spectroscopic amplitudes

Spectroscopic amplitudes are reduced matrix elements of one-nucleon creation/annihilation operators, e.g.,

$$\mathcal{A}_j^- = -\frac{\langle A-1 || \hat{a}_j || A \rangle}{\sqrt{2J_{A-1} + 1}}, \quad (5.2)$$

where  $|A\rangle$  is the initial wave function with  $A$  nucleons,  $|A-1\rangle$  is the final wave function with  $A-1$  nucleons (and angular momentum  $J_{A-1}$ , and  $\hat{a}_j$  is the nucleon annihilation operator for the  $j$ th orbital. Similarly,

$$\mathcal{A}_j^+ = -\frac{\langle A+1 || \hat{a}_j^\dagger || A \rangle}{\sqrt{2J_{A+1} + 1}}. \quad (5.3)$$

The reduced matrix elements are computed according to the conventions of Edmonds [1996]. In the literature, one more commonly uses the *spectroscopic factor*,

$$S_j = (\mathcal{A}_j^+)^2. \quad (5.4)$$

To compute the spectroscopic amplitude in `RHODIUM`, you need to first create the following files using `BIGSTICK`:

- The `.wfn` wave function file for the initial wave function (typically with option ‘`(n)`’, though other options can also create it);
- The `.bas` basis file for the initial model space using the option ‘`(b)`’;
- And the `.wfn` and `.bas` files for the final wave function. These must have one more or one less nucleon than the initial wave function.

As an example, let’s suppose we’ve created files for  $^{47,48}\text{Cr}$  in the  $pf$  valence space using the `GX1A` interaction. Starting `RHODIUM`, we enter the names of the basis files, `cr47.bas` and `cr48.bas`, leaving off the mandatory extension:

```

Enter name of INITIAL .bas file
cr47
Basis file successfully opened
about to read in basis
dimbasischeck=          483887
.....
DONE READING IN
Enter name of FINAL .bas file
cr48
Basis file successfully opened
```

```

dimbasischeck=          1963461
....
Choose one of the following menu options

Enter the six-character code for your choice::
(spamp1) 1-particle spectroscopic amplitude (change in number of particles)
(appsp1) Add/remove a single particle (linear combination)

Because RHODIUM detected these two bases differ by exactly one nucleon, it
offered these two menu choices. In the next section we will consider the option
(appsp1). Here we consider the first option, (spamp1).

You are calculating spectroscopic amplitudes with changes in the number of particles
Enter output name (enter "none" if none)
Cr47toCr48
....
finished setting up local vectors...
Initial state wave function file
Enter input name of .wfn file
cr47
dimbasischeck=          483887
.bas and .wfn headers agree
There are          10  initial wave functions
Enter start, stop to compute spectroscopic factors
(Enter 0,0 to take all )
0 0
   1      -83.81071      0.00000      2.500      0.500
   2      -83.78111      0.02959      1.500      0.500
   3      -83.69141      0.11930      3.500      0.500
   4      -82.56263      1.24808      5.500      0.500
...
Final state wave function file
Important: cannot be the same file as initial
Enter input name of .wfn file
cr48
dimbasischeck=          1963461
.bas and .wfn headers agree
There are           5  final wave functions
Enter start, stop to compute spectroscopic factors
(Enter 0,0 to take all )
0 0
   1      -99.57887      0.00000      0.000      0.500
   2      -98.79021      0.78866      2.000      0.500
...

All done with spectroscopic factors

```

The output is written to a `.spres` file, here `Cr47toCr48.spres`. That file has information on the initial and final wave functions:

```

RHODIUM version 0.8.9 May 2024
# Wavefunctions -- initial      4    3
  1      -83.81071      0.00000      2.500    0.500
  2      -83.78111      0.02959      1.500    0.500
....
#
# Wavefunctions -- final        4    4
  1      -99.57887      0.00000      0.000    0.500
  2      -98.79021      0.78866      2.000    0.500
...

```

and on the single particle orbitals

```

Single particle state quantum numbers
ORBIT      N      L      2 x J
  1         0       3       7
  2         1       1       3
  3         0       3       5
  4         1       1       1

```

and finally on the single particle amplitudes themselves:

```

Initial state #    1 E = -83.81071 2xJ, 2xT =    5    0
Final state   #    1 E = -99.57887 2xJ, 2xT =    0    0
orbit  amp
  3 -0.353712

Initial state #    1 E = -83.81071 2xJ, 2xT =    5    0
Final state   #    2 E = -98.79021 2xJ, 2xT =    4    0
orbit  amp
  1  0.697354
  2  0.227033
  3 -0.083912
  4 -0.061296
....

```

The code will automatically output the amplitude for proton or neutron; it's up to the user to remember which species it is.

At the end of the `.spres` file, the code also writes out sum rules (to be finished).

### 5.6.1 Sum rules

A check on the spectroscopic amplitudes are sum rules.

We define the spectroscopic amplitude for adding a particle (there is actually a different definition) as

$$A_j^{fi,+} = \frac{\langle J_f || \hat{a}_j^\dagger || J_i \rangle}{[J_f]}. \quad (5.5)$$

One can show Rowe and Wood [2010] that

$$A_j^{fi,+} = -\langle J_f M_f | \left[ \hat{a}_j^\dagger \otimes |J_i\rangle \right]_{J_f M_f}, \quad (5.6)$$

independent of  $M_f$ . With a bit of work, one can show that

$$\sum_i \left( A_j^{fi,+} \right)^2 = \langle f | \hat{n}_j | f \rangle, \quad (5.7)$$

where  $\hat{n}_j$  is the number operator for the orbital  $j$  (it should be apparent I am suppressing all other quantum numbers on the orbital). We can sum the other way: by taking particular care with the sum over Clebsch-Gordan coefficients, one can show that

$$\sum_f \left( A_j^{fi,+} \right)^2 \frac{2J_f + 1}{2J_i + 1} = 2j + 1 - \langle i | \hat{n}_j | i \rangle. \quad (5.8)$$

As  $2j + 1$  is the maximum number of particles in this orbital, this sum rule counts how many holes there are in the orbital  $j$ .

These are sum rules for *adding* a particle to an initial state. What about removing a particle? By using the definition of reduced matrix elements, one can show

$$\langle J_f || \hat{a}_j^\dagger || J_i \rangle = (-1)^{J_f + j - J_i} \langle J_i || \tilde{a}_j || J_f \rangle. \quad (5.9)$$

If we define

$$A_j^{fi,-} = \frac{\langle J_f || \tilde{a}_j || J_i \rangle}{[J_f]}, \quad (5.10)$$

then

$$A_j^{fi,-} = (-1)^{J_i + j - J_f} \frac{[J_i]}{[J_f]} A_j^{if,+}. \quad (5.11)$$

With this we then get the sum rules

$$\sum_i \left( A_j^{fi,-} \right)^2 = 2j + 1 - \langle f | \hat{n}_j | f \rangle \quad (5.12)$$

and

$$\sum_f \left( A_j^{fi,-} \right)^2 \frac{2J_f + 1}{2J_i + 1} = \langle i | \hat{n}_j | i \rangle. \quad (5.13)$$

## 5.7 Adding or removing a nucleon

Rather than choose the ‘(spamp1)’ option to compute the spectroscopic amplitude, one can choose the option

(appsp1) Add/remove a single particle (linear combination)

which constructs new wave function(s) by adding or removing a single nucleon, or a linear combination of nucleons—for example, a plane wave, and writing that to a .wfn file.

To do that you need a file with the single-particle amplitudes, with extension .spa. The RHODIUM distribution comes with an example, SAMPLE.spa:

```
#SAMPLE SPA FILE
3  ! number of orbitals
1  0 2 1.5 ! index  n 1  j
2  0 2 2.5
3  1 0 0.5
2  ! species = 1 protons 2 neutrons
1 -0.1 ! orbital amplitude
2  1.3
3  0.05
```

The file includes:

1. An optional header, started by # or !
2. The number of orbitals (assumed to be the same for protons and neutrons)
3. A list of the orbitals, including (a) the index of the orbital (b) the integer radial nodal number (c) the integer orbital angular momentum  $l$  (d) the real total angular momentum  $j$
4. The species, = 1 for protons, = 2 for neutrons (this must be specified, as I assume that protons could be affected by Coulomb)
5. Finally, a list of the amplitude  $c_i$  for each orbital index  $i$

This will create a wave function, e.g., for creating a nucleon:

$$|\Psi_{A+1}\rangle = \left( \sum_i c_i \hat{a}_i^\dagger \right) |\Psi_A\rangle. \quad (5.14)$$

RHODIUM automatically takes care of  $m$ -values, and uses the input file the same for creating or annihilating a nucleon.

After reading in the .spa amplitude file, you will be asked if you want to normalize the wave function or not:

Do you want to normalize final wfns? (y/n)



The next step is to read in the initial `.wfn` file previously generated (either by BIGSTICK or by a previous application of RHODIUM).

```
There are          N  initial wave functions
Enter start, stop to apply one particle creation/removal
(Enter 0,0 to take all )
```

Finally you will be asked the name of the output file (with extension `.wfn`, but do not include the extension)

```
Enter name of output .wfn file
test
```

The option `appsp1` produces only a single final state in the output `.wfn` file. This can be inconvenient. For example, in generating an optical potential using Green's functions, one may want a nucleon added/removed for each of  $0d_{5/2}, 0d_{3/2}, 1s_{1/2}$ . Using only option `appsp1`, to add/remove a nucleon systematically from each available orbital, one would have to run the code multiple times, once for each orbital. In the *sd*-shell, this would require three files, one for each of  $0d_{5/2}, 0d_{3/2}, 1s_{1/2}$ .

## Chapter 6

# Acknowledgements

This material is based upon work supported by the U.S. Department of Energy, Office of Science, Office of Nuclear Physics, under Award Number DE-FG02-03ER41272. I thank Mark Caprio for pointing out some inconsistencies and infelicities in the text.

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