

Benchmarking Substellar Evolutionary Models Using New Age Estimates for HD 4747 B and HD 19467 B

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Benchmarking Substellar Evolutionary Models Using New Age Estimates for HD 4747 B and HD 19467 B

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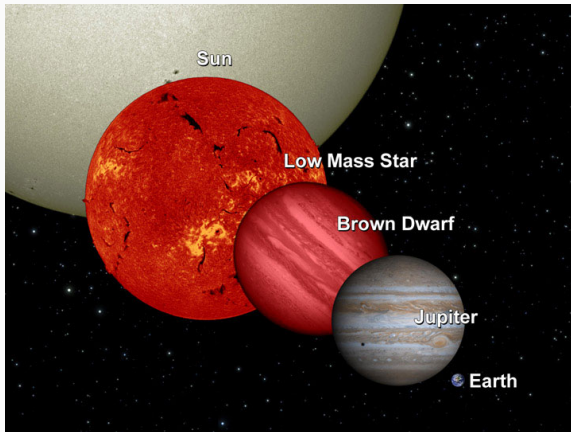
2. Observations and Results

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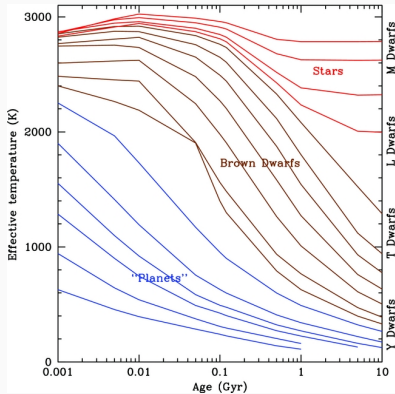
Introduction and Motivation

Not a Star and Not a Planet

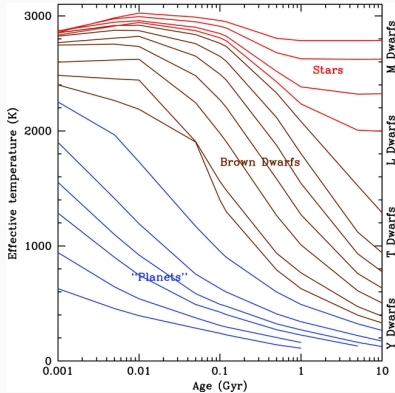
- A brown dwarf is an object that does not sustain nuclear fusion
- Generally considered “failed stars”
- Mass $\sim 13 - 72 M_{\text{Jup}}$



Brown Dwarfs Continuously Cool

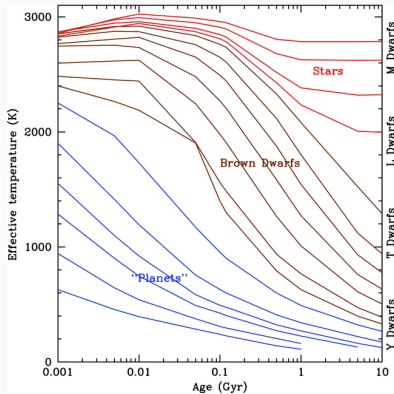


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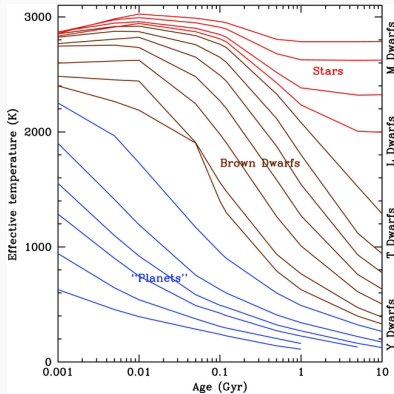
- Larger mass \rightarrow hotter \rightarrow higher luminosity

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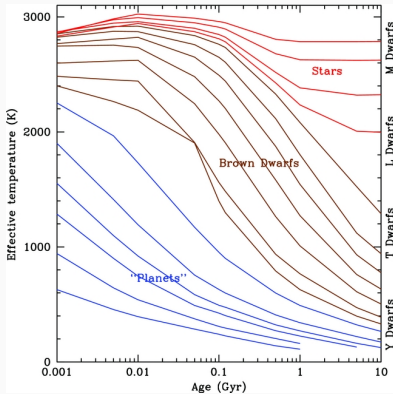
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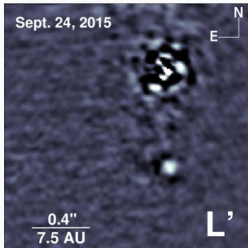
Brown Dwarfs Continuously Cool



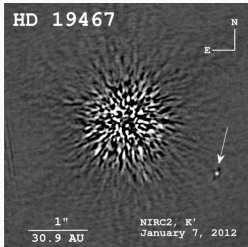
- Larger mass \rightarrow hotter \rightarrow higher luminosity
- Older age \rightarrow lower luminosity
- Need three parameters: mass, age, and luminosity
- How do you determine if a brown dwarf is massive but old or light but young?

Are the models actually any good?

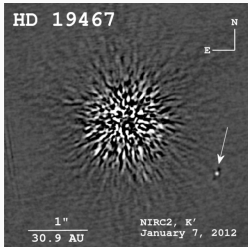
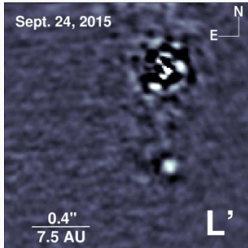
Finding “Benchmark” Objects



- Directly imaged brown dwarfs in binary (or higher) systems have model-independent masses, ages, and luminosities



Finding “Benchmark” Objects

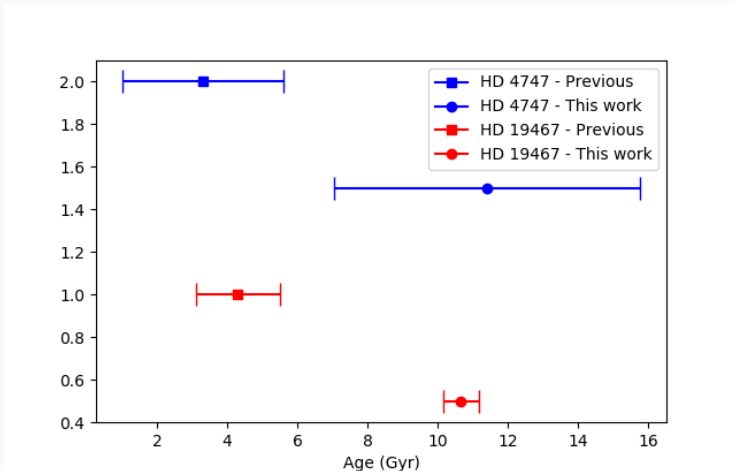


- Directly imaged brown dwarfs in binary (or higher) systems have model-independent masses, ages, and luminosities
- Use these “benchmark” brown dwarfs to test and calibrate models

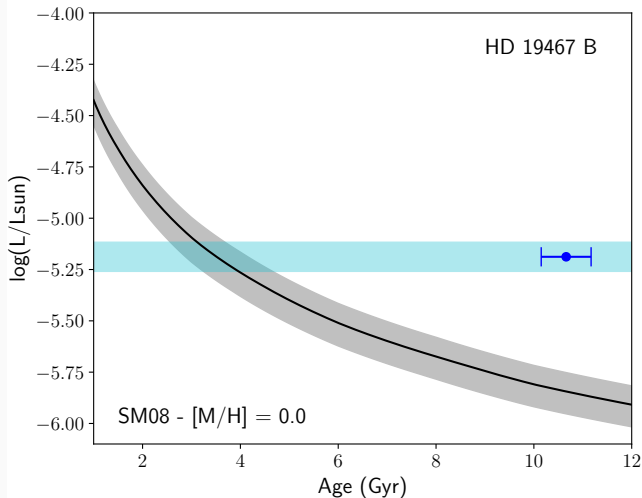
Observations and Results

Study the Host Star Instead

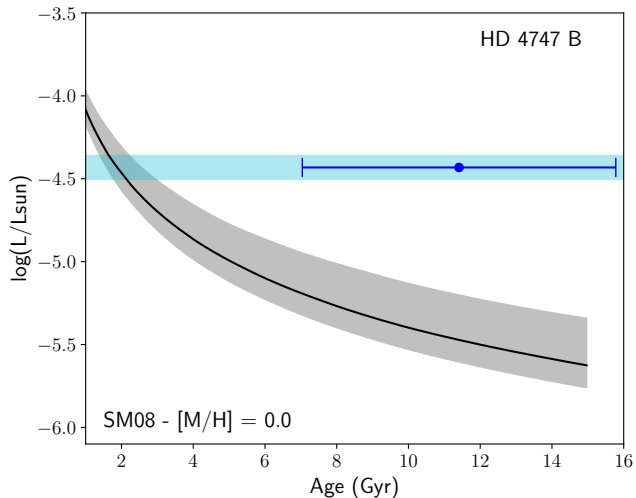
Stellar models are more rigorously tested than brown dwarf models, but age results depend on method used.



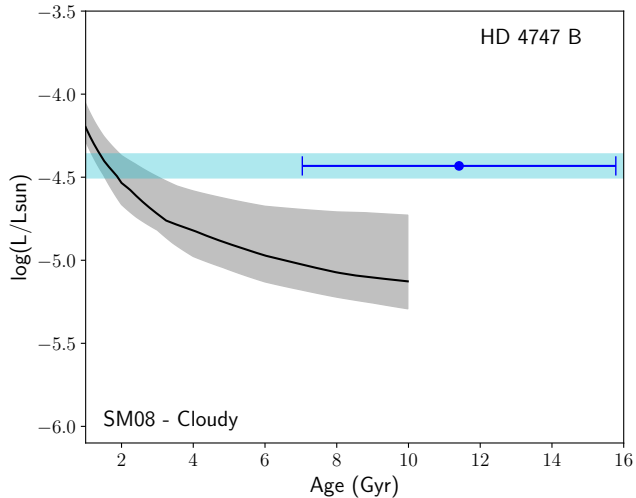
Comparing the Luminosities



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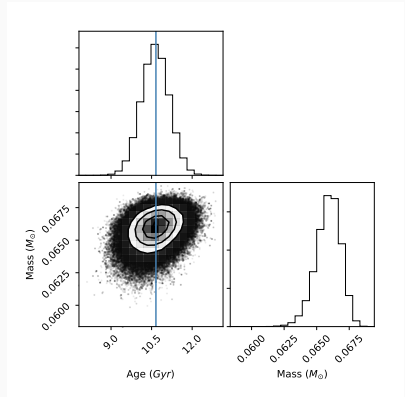


HD 4747 B Might Have Clouds

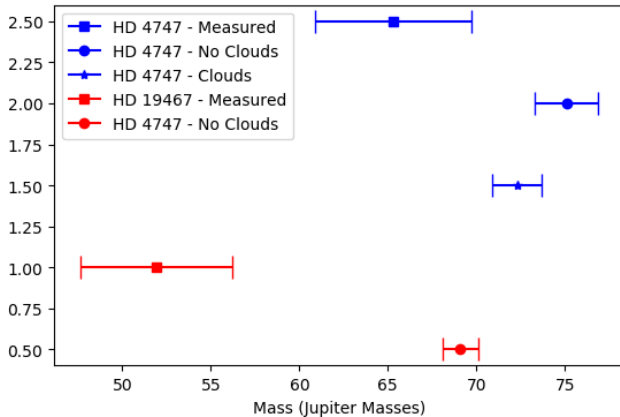


Comparing the Masses

- Models come in form of a grid
 - each (mass, age) pair has a corresponding luminosity
- Explore the parameter space in an unbiased way using a Markov Chain Monte Carlo simulation
- Returns a probability distribution for the explored parameter



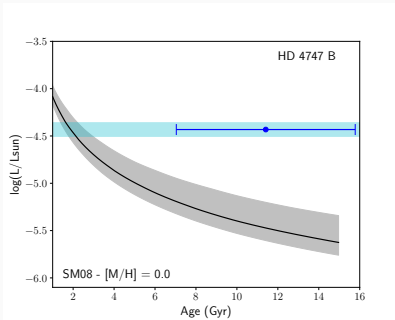
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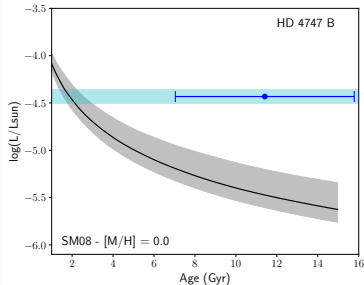
Conclusions

Conclusions & Future Work

- Current brown dwarf models under-predict the luminosities of HD 4747 B and HD 19467 B by at least a factor of 3 and over-predict the masses by at least 15%

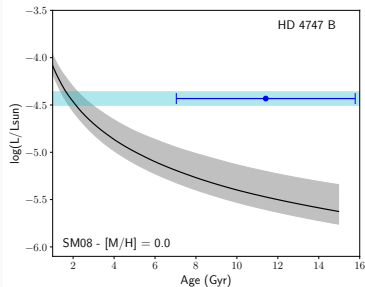


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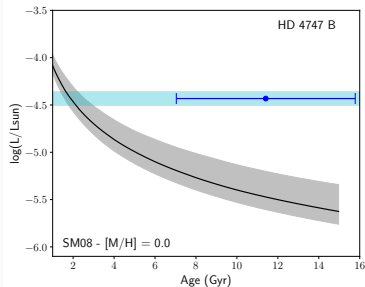
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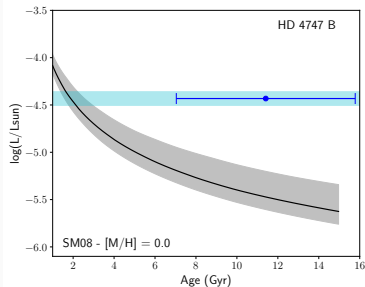
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Questions?