

COMPUTATIONAL PHYSICS (PHY241)

HOMEWORK ASSIGNMENT 7

Due Date: **Thursday, March 31, 2022**

The Planck theory of thermal radiation tells us that in the (angular) frequency interval ω to $\omega + d\omega$, a black body of unit area radiates electromagnetically an amount of thermal energy per second equal to $I(\omega) d\omega$, where

$$I(\omega) = \frac{\hbar}{4\pi^2 c^2} \frac{\omega^3}{(e^{\hbar\omega/k_B T} - 1)}.$$

Here \hbar is Planck's constant over 2π , c is the speed of light, and k_B is Boltzmann's constant.

1. (10 points) Show that the total energy per unit area radiated by a black body is

$$W = \frac{k_B^4 T^4}{4\pi^2 c^2 \hbar^3} \int_0^\infty \frac{x^3}{e^x - 1} dx.$$

2. (10 points) Briefly explain the integration algorithm you would like to choose. At least, the corresponding finite difference expressions should be given and explained.
3. (40 points) Write a program to evaluate the integral in this expression.
4. (10 points) Discuss how accurate your answer is. Plot W as a function of the number of your chosen grid points (N) at a fixed temperature (say 300K).
5. (10 points) Make a plot showing the variation of W with respect to temperature T .
6. (20 points) Even before Planck gave his theory of thermal radiation around the turn of the 20th century, it was known that the total energy W given off by a black body per unit area per second followed Stefan's law: $W = \sigma T^4$, where σ is the Stefan-Boltzmann constant. Use your value for the integral above to compute a value for the Stefan-Boltzmann constant (in SI units) to three significant figures. Check your result against the known value, which you can find in books or on-line. You should get good agreement.