# Code Repositories: What they are, why you might care, and How to Use Them

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# Where can you do this?

Atlassian Bitbucket

github SOCIAL CODING

source forge


GitLab

PERFORCE

# Where can you do this?

github
SOCIAL CODING

git

# As a "case study"...

# What is the philosophy behind the code?

The Astropy Project is a community effort to develop a single core package for Astronomy in Python and foster interoperability between Python astronomy packages.

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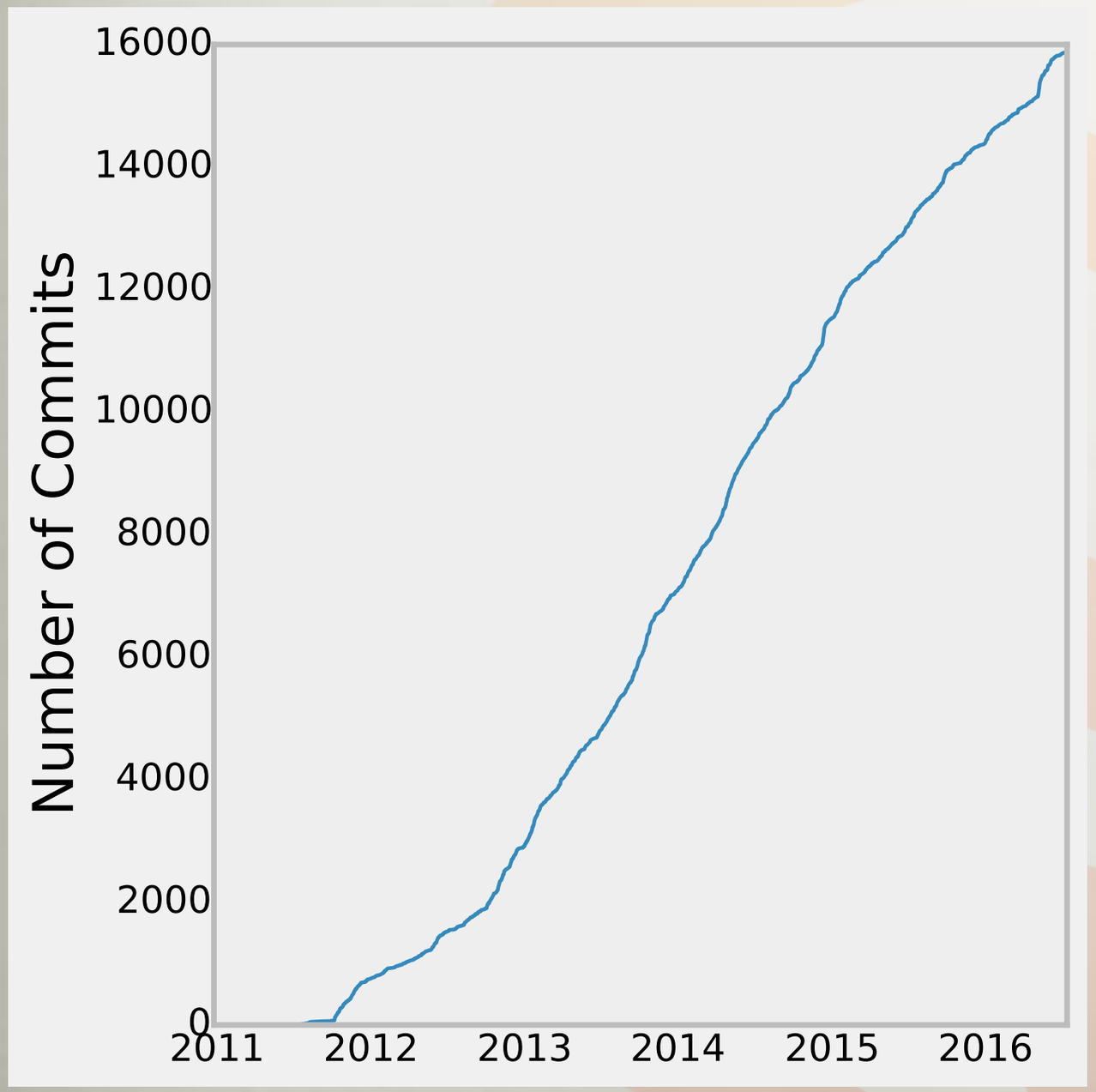
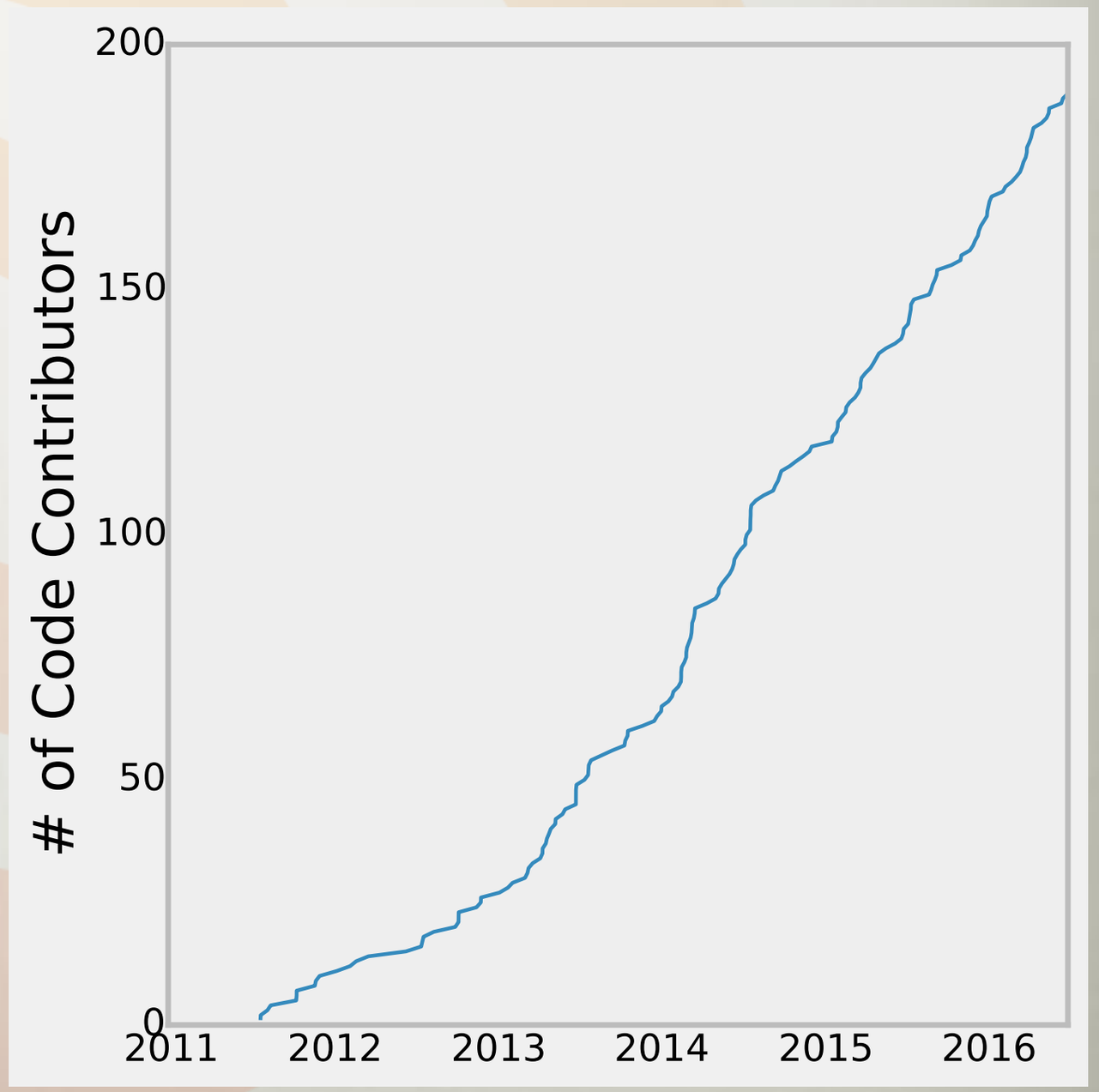
The Astropy Project is a community effort to develop a single core package for Astronomy in Python and foster interoperability between Python astronomy packages.

This means both *by* and *for* the community

(Professional) Astronomers help write it

# Contributions are Growing



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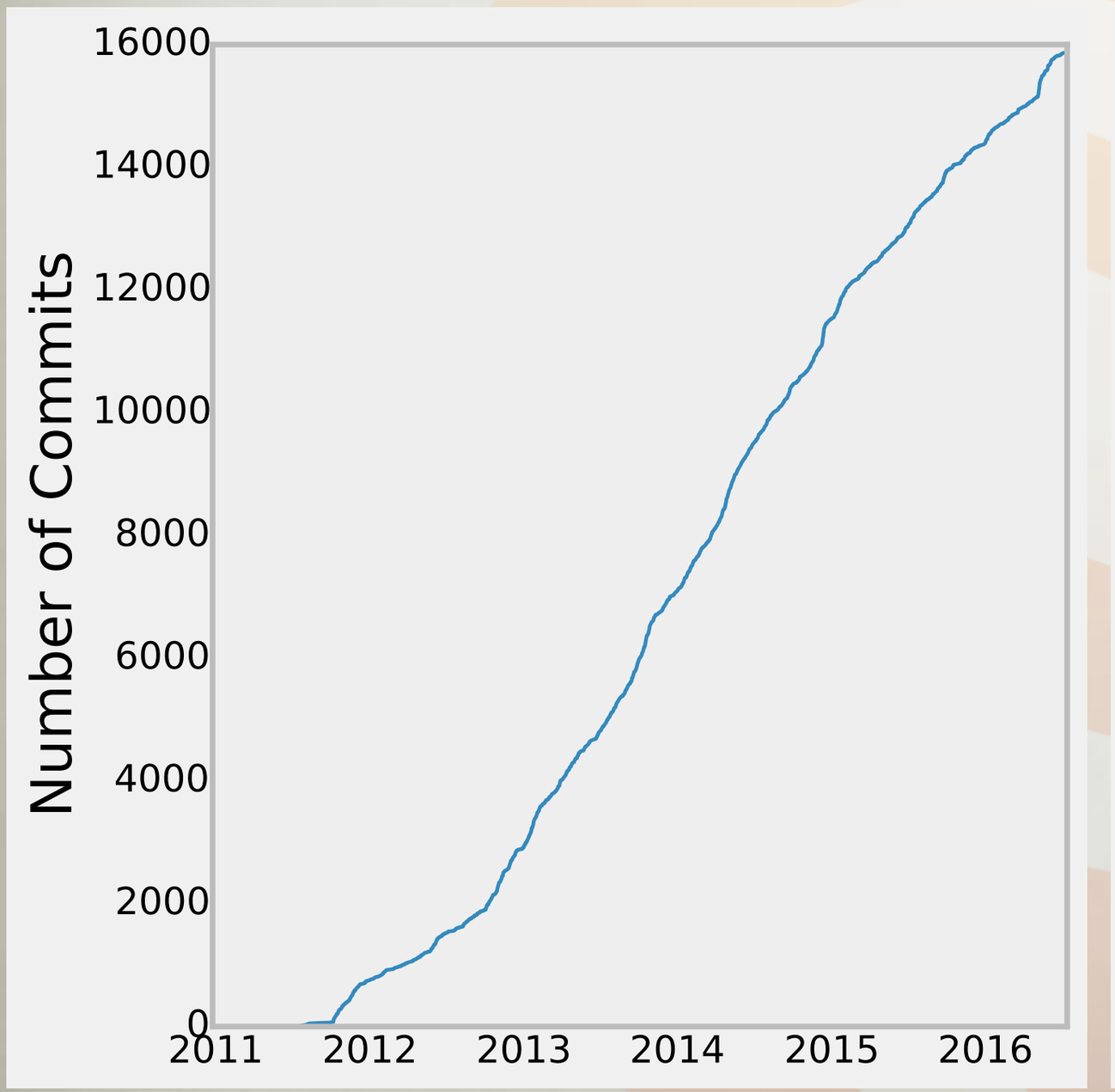
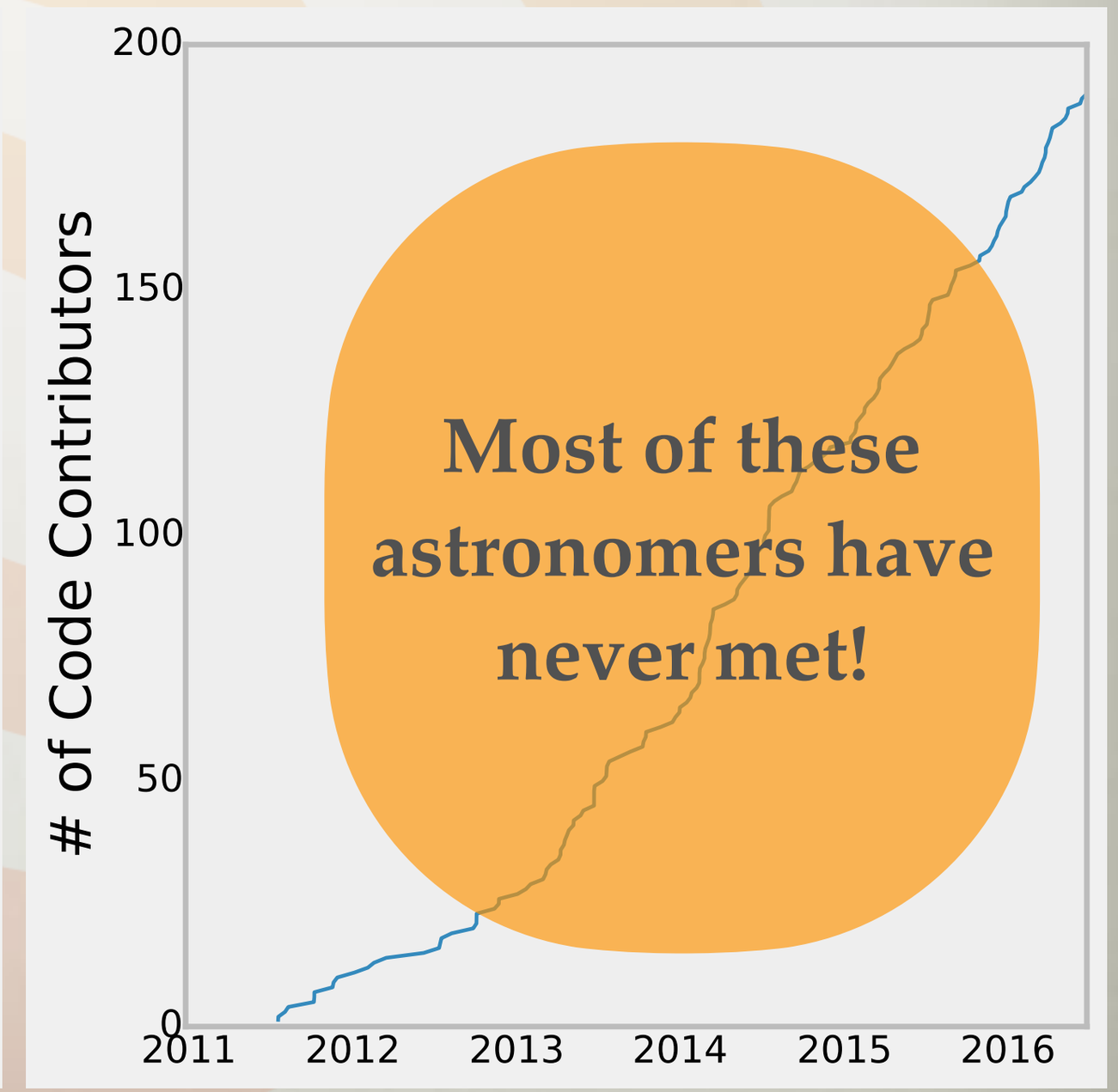
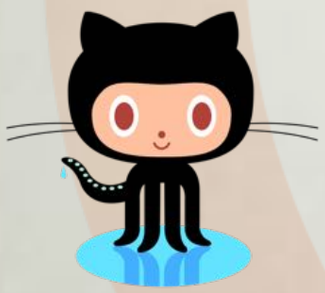


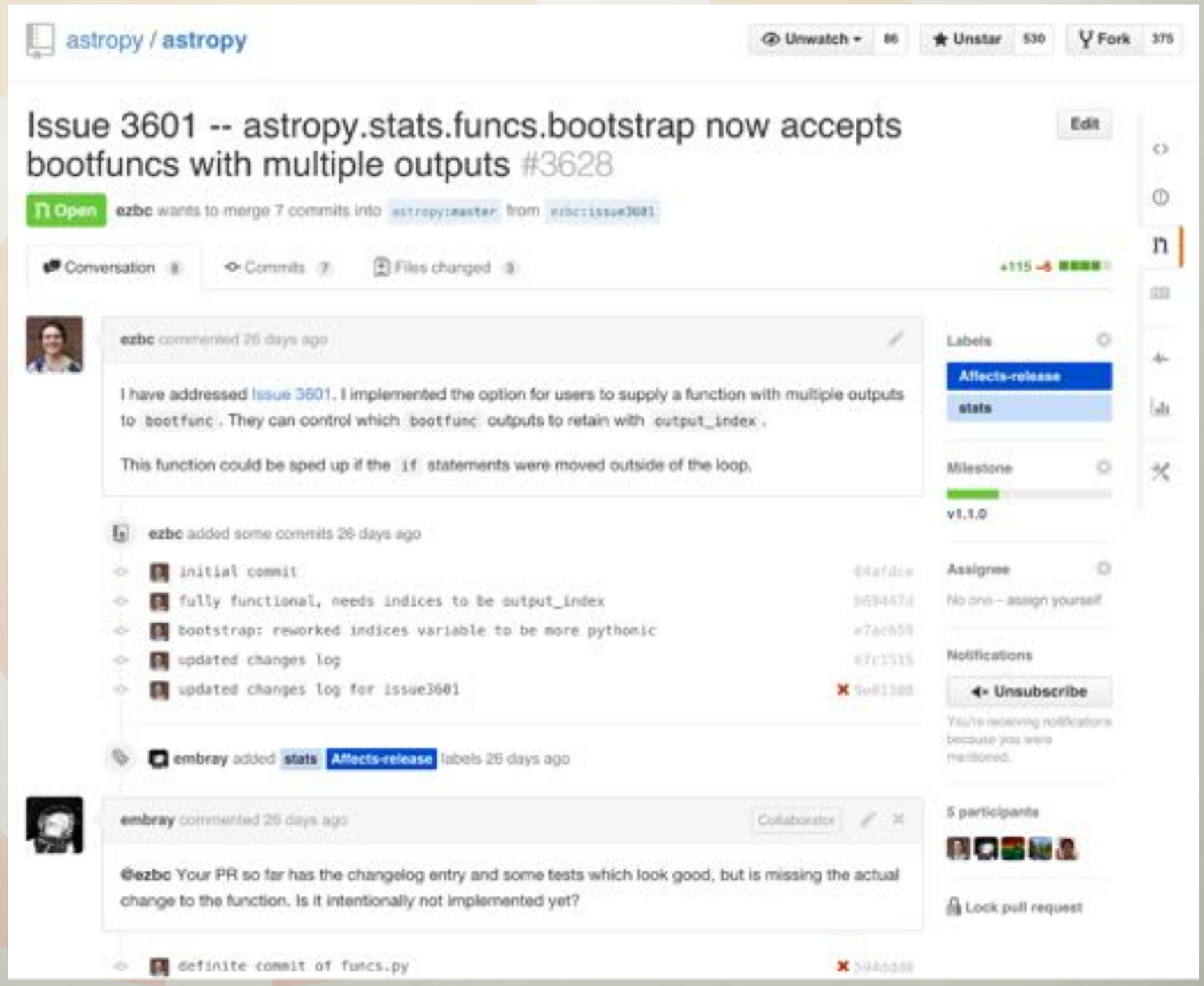



# Keys To Distributed Development

# Keys To Distributed Development: Github

github
SOCIAL CODING





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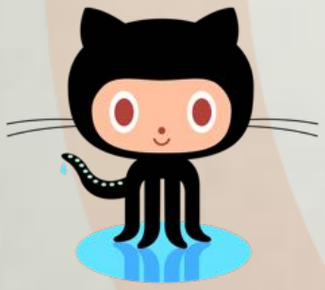

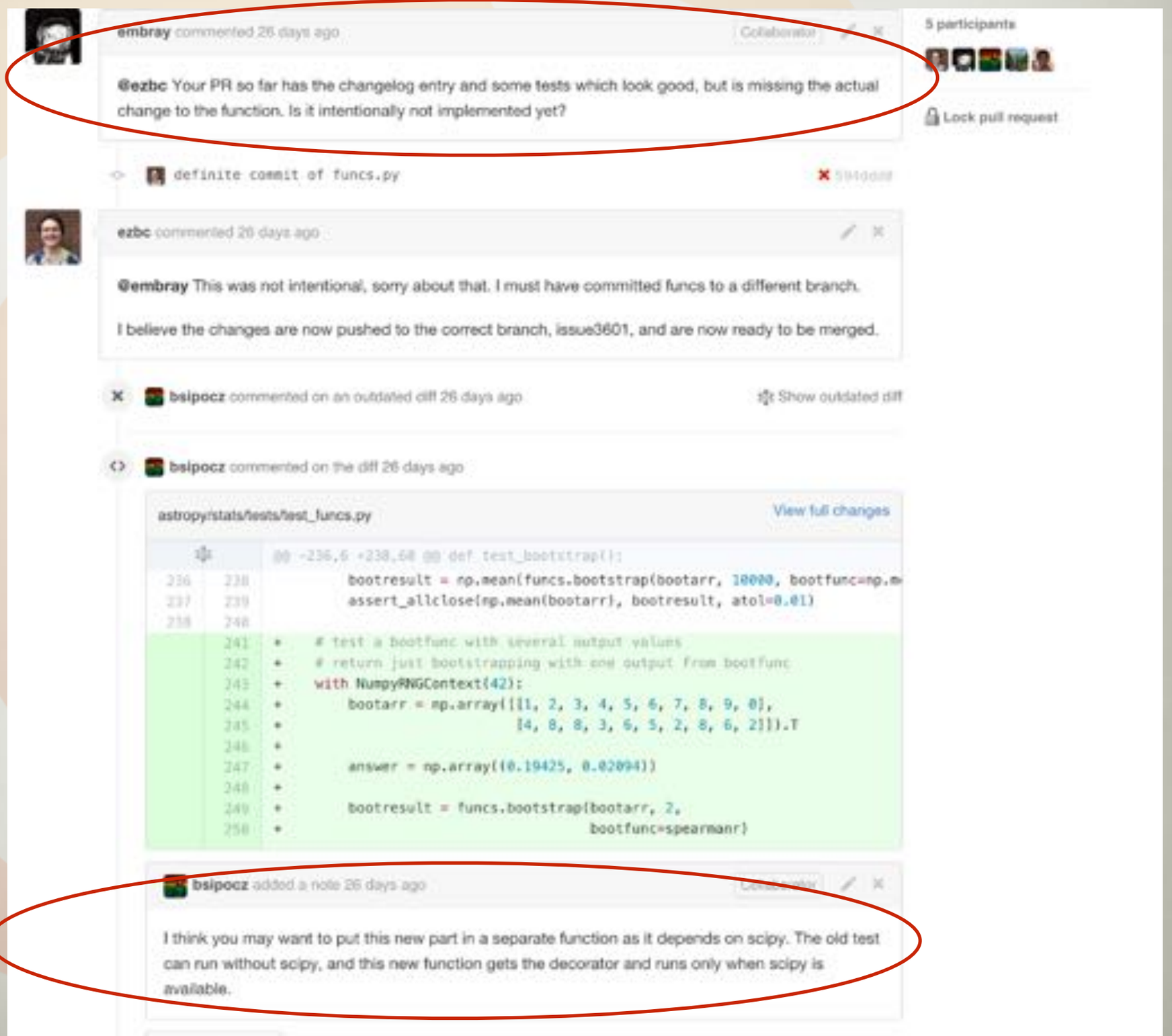

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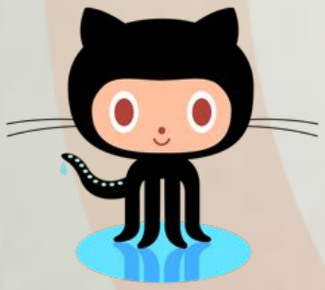

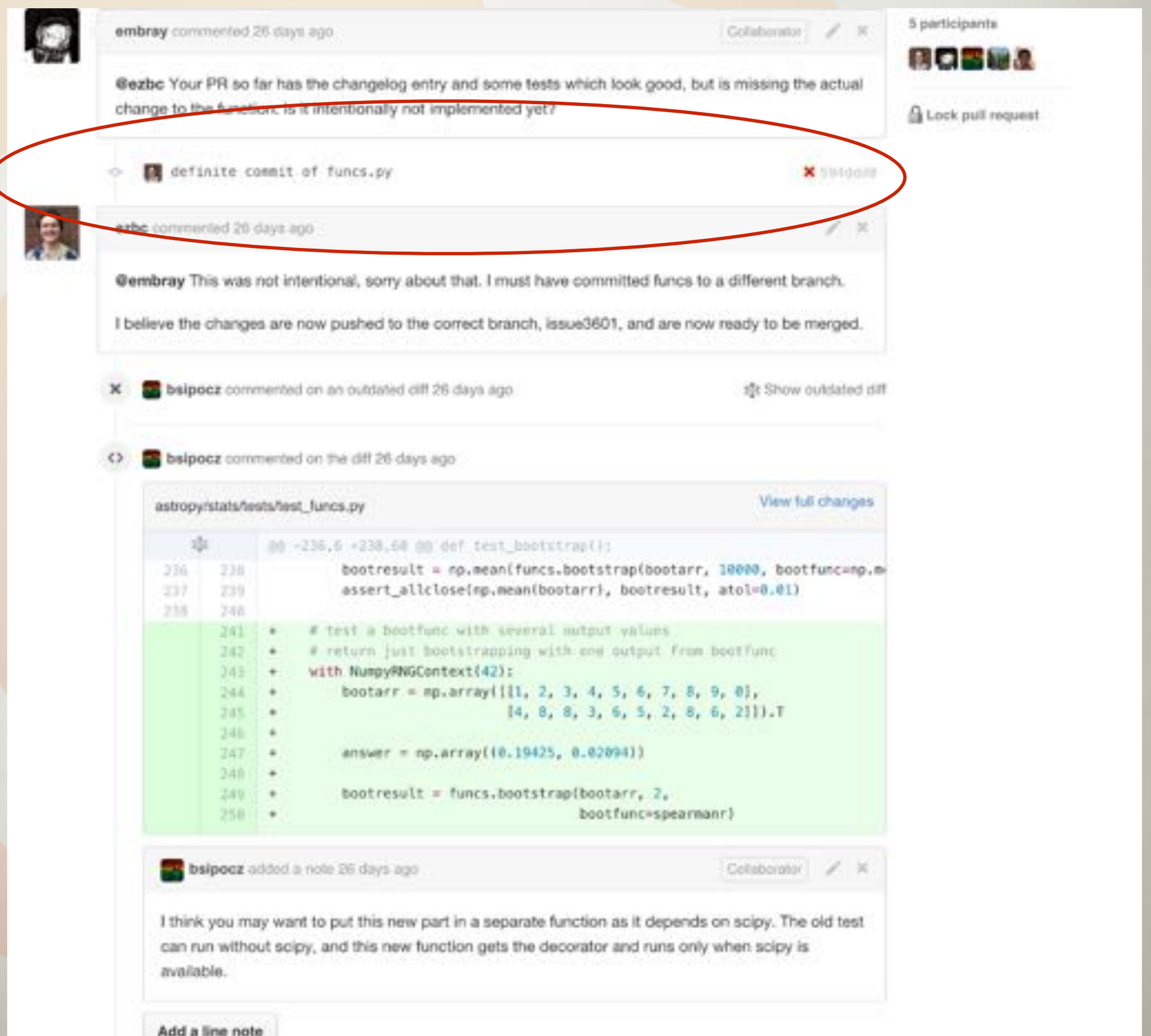

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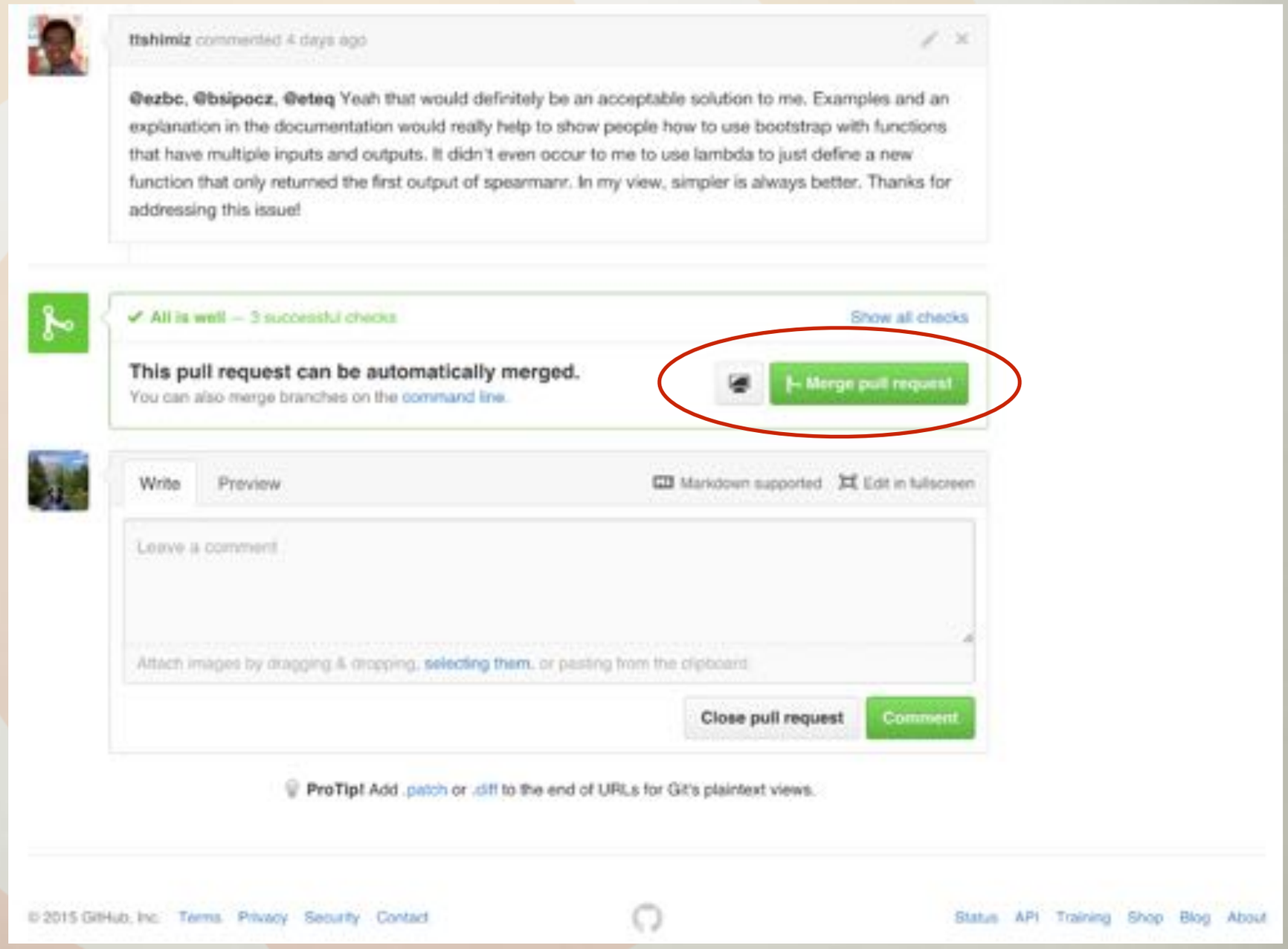
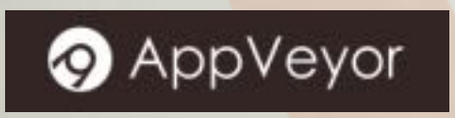
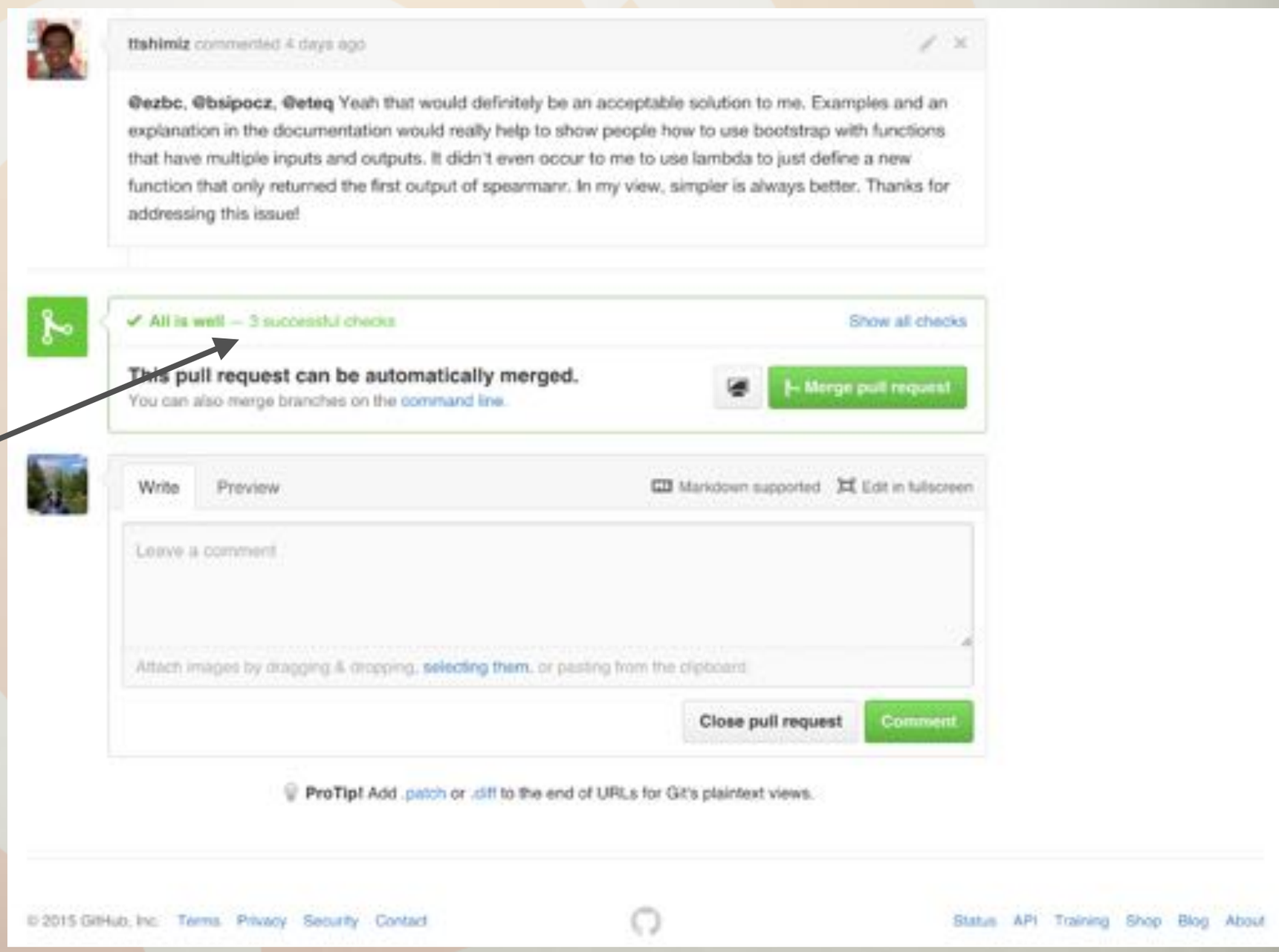


# Keys To Distributed Development: Continuous Testing

py.test

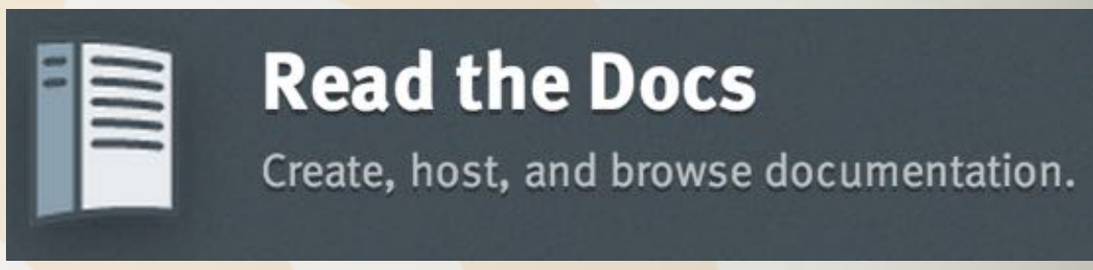



# Keys To Distributed Development: Docs

Sphinx

Python documentation generator




```
class FLRW(Cosmology):
    """ A class describing an isotropic and homogeneous
    (Friedmann-Lemaitre-Robertson-Walker) cosmology.

    This is an abstract base class -- you can't instantiate
    examples of this class, but must work with one of its
    subclasses such as `LambdaCDM` or `wCDM`.

    Parameters
    ----------

    H0 : float or scalar `~astropy.units.Quantity`
        Hubble constant at z = 0.  If a float, must be in [km/sec/Mpc]

    Om0 : float
        Omega matter: density of non-relativistic matter in units of the
        critical density at z=0.  Note that this does not include
        massive neutrinos.

    Ode0 : float
        Omega dark energy: density of dark energy in units of the critical
        density at z=0.

    Tcmb0 : float or scalar `~astropy.units.Quantity`
        Temperature of the CMB z=0. If a float, must be in [K]. Default: 2.725.
        Setting this to zero will turn off both photons and neutrinos (even
        massive ones)

    Neff : float
        Effective number of Neutrino species. Default 3.04.

    m_nu : `~astropy.units.Quantity`
        Mass of each neutrino species. If this is a scalar Quantity, then all
        neutrino species are assumed to have that mass. Otherwise, the mass of
        each species. The actual number of neutrino species (and hence the
        number of elements of m_nu if it is not scalar) must be the floor of
        Neff. Usually this means you must provide three neutrino masses unless
        you are considering something like a sterile neutrino.

    name : str
        Optional name for this cosmological object.

    Ob0 : float
        Omega baryons: density of baryonic matter in units of the critical
        density at z=0.

    Notes
    -----
    Class instances are static -- you can't change the values
    of the parameters.  That is, all of the attributes above are
    read only.
    """
    def __init__(self, H0, Om0, Ode0, Tcmb0=2.725, Neff=3.04,
                 m_nu=u.Quantity(0.0, u.eV), name=None, Ob0=None):

        # all densities are in units of the critical density
        self._Om0 = float(Om0)
        if self._Om0 < 0.0:
            raise ValueError("Matter density can not be negative")
        self._Ode0 = float(Ode0)
        if Ob0 is not None:
            self._Ob0 = float(Ob0)
            if self._Ob0 < 0.0:
```

class astropy.cosmology. **FLRW**(*H0, Om0, Ode0, Tcmb0=2.725, Neff=3.04, m_nu=<Quantity 0.0 eV>, name=None, Ob0=None*) [edit on github][source]

Bases: `astropy.cosmology.core.Cosmology`

A class describing an isotropic and homogeneous (Friedmann-Lemaitre-Robertson-Walker) cosmology.

This is an abstract base class – you can't instantiate examples of this class, but must work with one of its subclasses such as `LambdaCDM` or `wCDM`.

**Parameters:**

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Effective number of Neutrino species. Default 3.04.

**m_nu** : `Quantity`

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**name** : str

Optional name for this cosmological object.

**Ob0** : float

Omega baryons: density of baryonic matter in units of the critical density at z=0.

**Notes**

Class instances are static – you can't change the values of the parameters. That is, all of the attributes above are read only.

**Attributes Summary**

| | |
|---|---|
| `H0` | Return the Hubble constant as an `Quantity` at z=0 |

# Keys To Distributed Development In a Shared Repo

- *git* + Github for sharing
- Test *everything* (automatically)
- *Easy* documentation (⇒ thorough)

# Lets dig down on how you do Shared Development with Public Repos

# Lets dig down...

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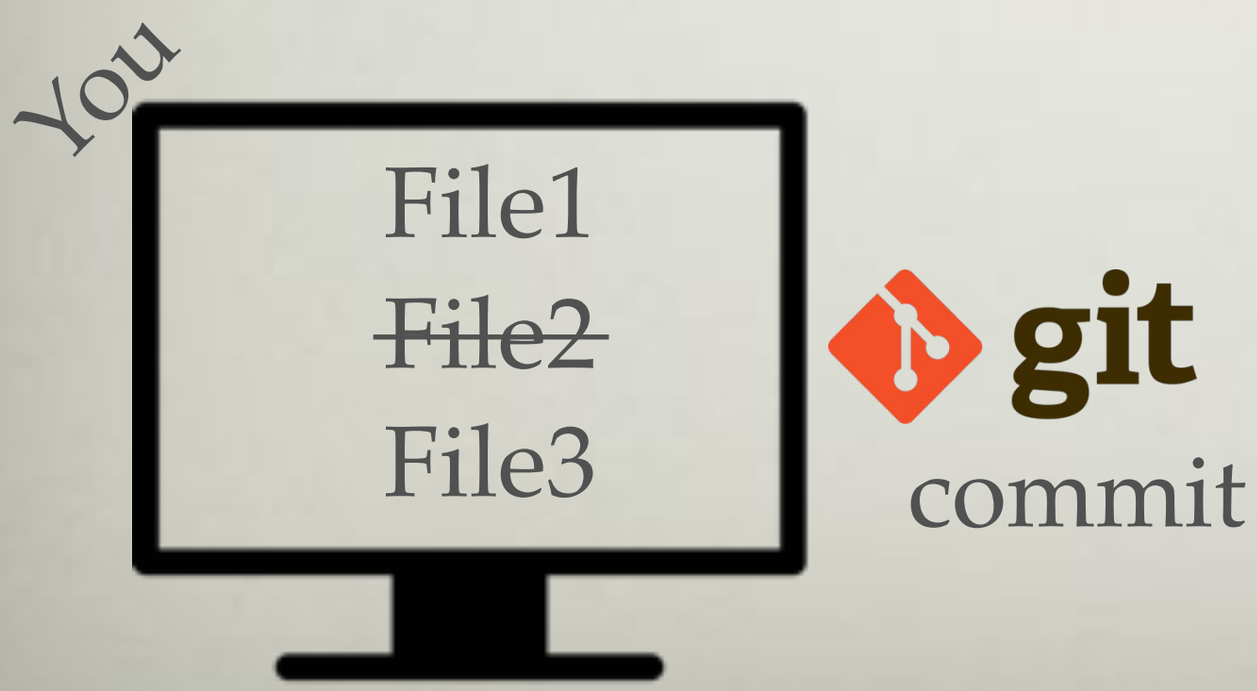

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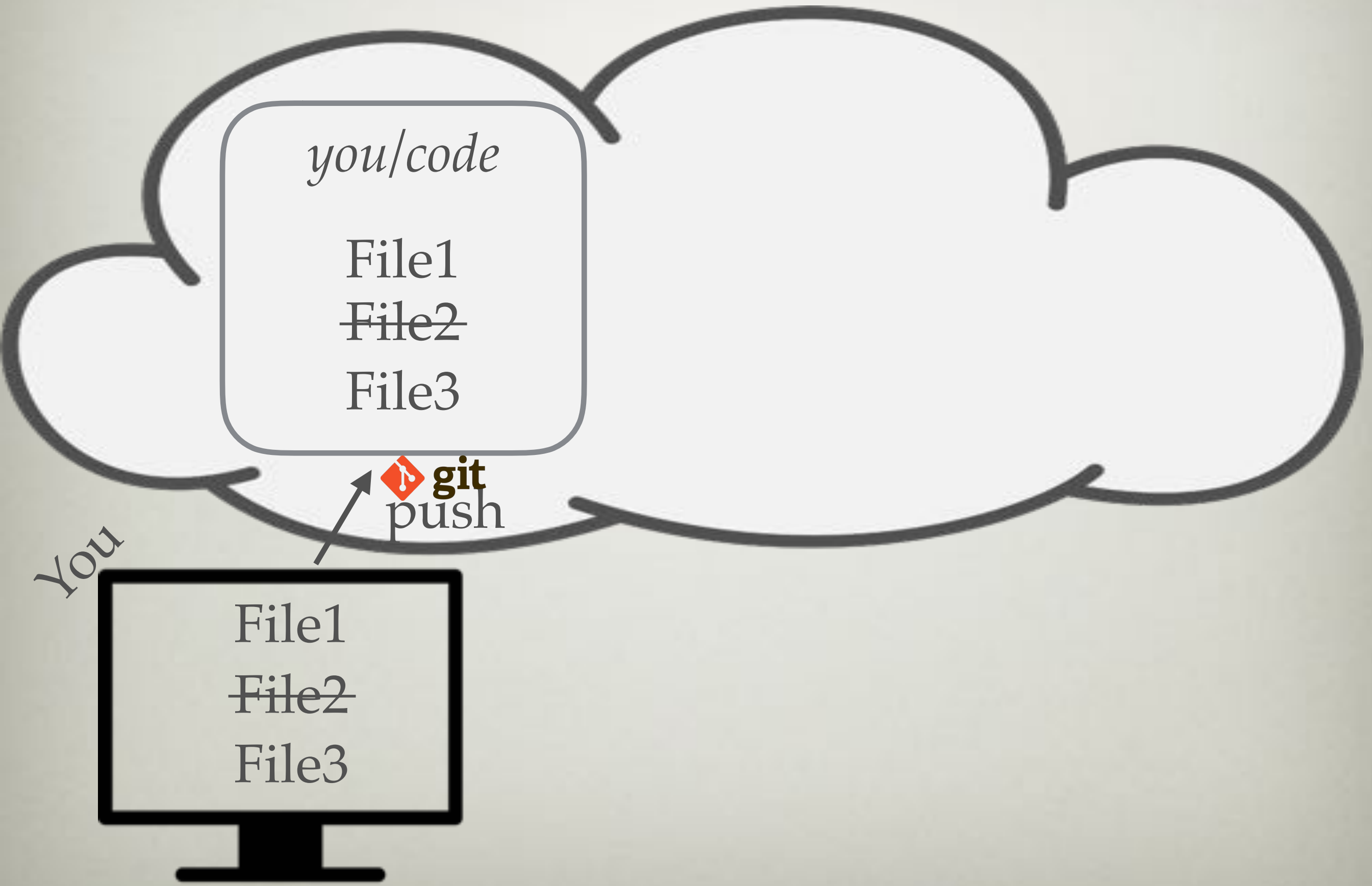

# Lets dig down...

*you/code*

File1
~~File2~~
File3

fork

*me/code*

File1
~~File2~~
File3

You

File1
~~File2~~
File3

# Lets dig down...

*you/code*

File1
~~File2~~
File3

*me/code*

File1
~~File2~~
File3

git clone

You

File1
~~File2~~
File3

Me

File1
~~File2~~
File3

# Lets dig down...

*you/code*

File1
~~File2~~
File3

*me/code*

File1
~~File2~~
File3

You

File1
~~File2~~
File3

git
commit

Me

File1
~~File2~~
File3*

# Lets dig down...

*you/code*

File1
~~File2~~
File3

*me/code*

File1
~~File2~~
File3*

git
push

You

File1
~~File2~~
File3

Me

File1
~~File2~~
File3*

# Lets dig down...

*you/code*

File1
~~File2~~
File3

Pull Request

*me/code*

File1
~~File2~~
File3*

You

File1
~~File2~~
File3

Me

File1
~~File2~~
File3*

# Lets dig down...

*you/code*

File1
~~File2~~
File3*

Merge Button

*me/code*

File1
~~File2~~
File3*

You

File1
~~File2~~
File3

Me

File1
~~File2~~
File3*

# Lets dig down...

*you/code*

File1
~~File2~~
File3*

*me/code*

File1
~~File2~~
File3*

git pull

You

File1
~~File2~~
File3*

Me

File1
~~File2~~
File3*

# Lets dig down...

*you/code*

File1*
~~File2~~
File3*

*me/code*

File1
~~File2~~
File3*

git push

You

File1*
~~File2~~
File3*

git commit

Me

File1
~~File2~~
File3*

# Lets dig down...

*you/code*

File1*
~~File2~~
File3*

*me/code*

File1
~~File2~~
File3*

git
pull

You

File1*
~~File2~~
File3*

Me

File1*
~~File2~~
File3*

# What About Actually Laying out the Code?

- There's not an easy answer for science code - it tends to develop "organically".
- Often it's best just to split files when they get too big.
- Always keep the novice user (or future you) in mind… Use descriptive names.
- ***Think modular!***

# What About Packaging Code?

- Deliver your code in some form that others can install without thinking too hard about where anything goes.
  - Makefiles, ruby gems, python packages, etc.
  - (Includes sensible versioning!)

# What About Packaging Python Code?

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- "**source directory/folder**": the directory/folder with all of a codes "stuff"
- "**repository**"/"**repo**": the source directory *in version control*
- "**submodule**": a git repo embedded in *another* git repo
  - "**astropy-helpers**": an example seen in Astropy packages

# Sample Package Layout

README
LICENSE
setup.py
mypackage/__init__.py
mypackage/mymodule.py
mypackage/secondmodule.py
mypackage/subpackage/__init__.py
mypackage/subpackage/anothermodule.py

```
import mypackage
from mypackage import my module
from mypackage import secondmodule
from mypackage import subpackage
from mypackage.subpackage import anothermodule
```

The goal of packaging and installing is basically to make that work anywhere

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- Anything x.y.z<something else> is a pre-release
  - E.g., 1.2.3beta, 2.1.6rc2
- But don't get too worked up. 0.1 -> 0.2 -> 0.3 is better than nothing.

# Licensing Your Code

- Rule #1: Have a license!
- Rule #2: There is no rule #2.

(see problem sets for more)

# Now go do it yourself!

# Astropy's Development

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