# Chapter 1

# **Basis Sets**

## 1.1 Methods

#### 1.1.1 Basis set size optimization

In keeping with the theme of "computers should work, people should think", we decided to use a somewhat novel method of generating our basis sets. We begin by choosing an arbitrarily large basis set size, such as s(1:40)p(1:40)d(1:40)f(1:40) (here i(#:#) refers to the starting and ending index of  $\zeta$ 's used for symmetry i). We then find the optimal  $\alpha$ ,  $\beta$ ,  $\delta$ , and  $\gamma$  wtbs parameters for this basis set. We then use the following steps to find the optimal basis set size.

Step 1. Begin by finding the fewest number of f functions necessary. This can be done by generating .inps files that range in size from s(1:40)p(1:40)d(1:40)f(1:1) to s(1:40)p(1:40)d(1:40)f(1:40).

- Step 2. Optimize the basis sets and select the smallest basis set that is still below some minimum accuracy threshold (we chose a relative error of no greater than  $5.0 \times 10^{-9}$  to numerical calculations). The size of this basis set is  $s(1:40)p(1:40)d(1:40)f(1:x_f)$ .
- Step 3. Replace the wtbs parameters with those from the newly optimized set and generate a list of .inp files that range in size from  $s(1:40)p(1:40)d(1:x_f)f(1:x_f)$  to  $s(1:40)p(1:40)d(1:40)f(1:x_f)$ .
- Step 4. Optimize these new sets and select the smallest that is still below the accuracy threshold. The size of this basis set is  $s(1:40)p(1:40)d(1:x_d)f(1:x_f)$ .
- Step 5. Repeat steps 3 and 4 for the remaining symmetries. The size of the basis set at the end of this step will be of size  $s(1:x_s)p(1:x_p)d(1:x_d)f(1:x_f)$ .
- Step 6. Replace the wtbs parameters with those from the newly optimized set and generate a list of .inp files that range in size from  $s(1:x_s)p(1:x_p)d(1:x_d)f(1:x_f)$  to  $s(1:x_s)p(1:x_p)d(1:x_d)f(x_f:x_f)$ .
- Step 7. Optimize and select from the basis sets. The new set will be of size  $s(1:x_s)p(1:x_p)d(1:x_d)f(y_f:x_f)$ .
- Step 8. Repeat steps 6 and 7 for the other symmetries except s. The final basis set will be of size  $s(1:x_s)p(y_p:x_p)d(y_d:x_d)f(y_f:x_f)$ .

This process has a major drawback in that it requires a lot of computer power to run efficiently. But this is not really a problem if access to large computer clusters is available. The advantages of this are it finds a very small 1.2. DISCUSSION 3

basis set that is still accurate, and it is also completely automatable. If there are no problems with individual calculations not converging, the output files never even need to be manually examined!

## 1.2 Discussion

The optimized wtbs parameters and basis set sizes for elements 2 to 86 are shown in Table 1.1.

Table 1.1: Basis sets optimized using rwtbs

Element	Configuration	$\alpha$	$\beta$	$\delta$	$\gamma$	$\mathbf{S}$	p	d	f
02he	1s	$8.1402 \times 10^{-02}$	1.9536	4.5046	1.5158				
03li	2s	$1.5961 \times 10^{-02}$	1.9334	5.7019	1.5733				
04be	1s	$2.6475 \times 10^{-02}$	1.9382	5.8417	1.5941				
05b	2p	$3.2387 \times 10^{-02}$	1.9480	5.5736	1.5233				
06c	3p	$4.6132 \times 10^{-02}$	1.9415	4.7179	1.3177				
07n	4s	$5.9765 \times 10^{-02}$	1.9239	5.1835	1.4424				
08o	3p	$6.9671 \times 10^{-02}$	1.9366	5.1701	1.4261				
09f	2p	$8.3219 \times 10^{-02}$	1.9426	5.0845	1.4080				
10ne	1s	$9.9432 \times 10^{-02}$	1.9457	4.9889	1.3924				
11na	2s	$1.6960 \times 10^{-02}$	1.9502	6.0567	1.4695				
12mg	1s	$2.3756 \times 10^{-02}$	1.9320	5.6865	1.4306				
13al	2p	$2.2391 \times 10^{-02}$	1.8983	5.5553	1.4135				
14si	3p	$3.3523 \times 10^{-02}$	1.9218	5.4190	1.4103				
15p	4s	$4.4108 \times 10^{-02}$	1.9072	5.1968	1.3909				
16s	3p	$5.0047 \times 10^{-02}$	1.8967	4.9626	1.3629				
17cl	2p	$5.8312 \times 10^{-02}$	1.8869	4.7841	1.3443				
18ar	1s	$6.8345 \times 10^{-02}$	1.8787	4.6546	1.3314				
19k	2s	$1.4020 \times 10^{-02}$	1.8907	6.3886	1.5274				
20ca	1s	$1.8614 \times 10^{-02}$	1.8758	6.1617	1.5117				
21sc	2d	$2.0595 \times 10^{-02}$	1.8799	6.1968	1.5113				
22ti	3f	$2.2312\times10^{-02}$	1.8870	6.2747	1.5134				

23v	4f	$2.3706 \times 10^{-02}$	1.8876	6.6055	1.5128
24cr	$7\mathrm{s}$	$2.3579 \times 10^{-02}$	1.8655	5.3555	1.4093
$25 \mathrm{mn}$	6s	$2.5779 \times 10^{-02}$	1.8659	5.3494	1.4107
26 fe	5d	$2.7386 \times 10^{-02}$	1.8685	5.3547	1.4103
27co	4f	$2.8470 \times 10^{-02}$	1.8696	5.9737	1.4317
28ni	3f	$3.0030 \times 10^{-02}$	1.8738	5.9737	1.4250
29cu	2s	$2.6735 \times 10^{-02}$	1.8610	5.7993	1.4774
$30\mathrm{zn}$	1s	$3.2991 \times 10^{-02}$	1.8812	5.8188	1.3393
31ga	2p	$2.7458 \times 10^{-02}$	1.8637	5.7090	1.4648
32ge	3p	$3.6139 \times 10^{-02}$	1.8533	5.6198	1.4547
33as	4s	$4.7438 \times 10^{-02}$	1.8695	5.3427	1.3448
34se	3p	$5.0979 \times 10^{-02}$	1.8471	4.9562	1.5367
$35 \mathrm{br}$	2p	$5.9414 \times 10^{-02}$	1.8621	5.2129	1.3388
36 kr	1s	$6.8233 \times 10^{-02}$	1.8600	5.5404	1.3764
$37 \mathrm{rb}$	2s	$1.4319 \times 10^{-02}$	1.8574	7.1985	1.6287
38sr	1s	$1.8315 \times 10^{-02}$	1.8459	6.6780	1.5548
39y	2d	$2.0484 \times 10^{-02}$	1.8397	6.5392	1.5414
40zr	5f	$2.2027 \times 10^{-02}$	1.8373	6.4805	1.5354
41nb	6d	$2.4370 \times 10^{-02}$	1.8321	6.3626	1.5234
42 mo	$7\mathrm{s}$	$2.5726 \times 10^{-02}$	1.8317	6.3487	1.5218
43tc	6s	$2.6038 \times 10^{-02}$	1.8362	6.4554	1.5349
$44 \mathrm{ru}$	5f	$2.6910 \times 10^{-02}$	1.8329	6.7863	1.6058
$45\mathrm{rh}$	4f	$2.7554 \times 10^{-02}$	1.8358	6.8855	1.6196
46pd	1s	$5.9550 \times 10^{-02}$	1.8279	5.9798	1.5036
47ag	2s	$2.5958 \times 10^{-02}$	1.7951	5.4672	1.4209
48cd	1s	$3.0102 \times 10^{-02}$	1.7848	5.3332	1.4114
49in	2p	$2.7173 \times 10^{-02}$	1.7993	5.5382	1.4313
$50\mathrm{sn}$	3p	$3.4062 \times 10^{-02}$	1.7793	5.2471	1.4039
$51 \mathrm{sb}$	4s	$4.5051 \times 10^{-02}$	1.8149	6.5596	1.6032
52te	3p	$4.8683 \times 10^{-02}$	1.8168	6.2438	1.5279
53i	2p	$5.3928 \times 10^{-02}$	1.8104	6.0171	1.5066
54xe	1s	$6.0103 \times 10^{-02}$	1.8026	5.8608	1.4907
55cs	2s	$1.2855 \times 10^{-02}$	1.7828	6.2757	1.4895
56ba	1s	$1.5853 \times 10^{-02}$	1.7692	6.1433	1.5188
57la	2d	$1.7344 \times 10^{-02}$	1.7639	6.0521	1.5123
58ce	3h	$1.6781 \times 10^{-02}$	1.7686	6.0001	1.4644
59 pr	4i	$1.7185 \times 10^{-02}$	1.7758	6.2865	1.5337
60nd	5i	$1.7564 \times 10^{-02}$	1.7775	6.5512	1.5794

61pm	6h	$1.8016 \times 10^{-02}$	1.7797	6.6017	1.5841
$62\mathrm{sm}$	7f	$1.8709 \times 10^{-02}$	1.7777	5.8441	1.4672
63eu	8s	$1.8909 \times 10^{-02}$	1.7840	6.7077	1.5952
64 gd	7f	$1.9561 \times 10^{-02}$	1.7816	6.1306	1.5217
$65 \mathrm{tb}$	6h	$1.9924 \times 10^{-02}$	1.7878	6.5207	1.5707
66 dy	5i	$2.0346 \times 10^{-02}$	1.7876	6.5033	1.5289
67ho	4i	$2.0740 \times 10^{-02}$	1.7904	6.9684	1.6317
$68\mathrm{er}$	3h	$2.1199 \times 10^{-02}$	1.7921	7.0221	1.6383
$69 \mathrm{tm}$	2f	$2.1652 \times 10^{-02}$	1.7938	7.0776	1.6442
70yb	1s	$2.2102 \times 10^{-02}$	1.7952	7.1405	1.6542
71lu	2d	$2.4224 \times 10^{-02}$	1.7889	6.9809	1.6379
72hf	3f	$2.5724 \times 10^{-02}$	1.7860	6.9232	1.6343
73ta	4f	$2.7058 \times 10^{-02}$	1.7842	6.8929	1.6324
74w	5d	$2.8296 \times 10^{-02}$	1.7830	6.8797	1.6327
75re	6s	$2.9400 \times 10^{-02}$	1.7826	6.8837	1.6343
76os	5d	$3.0756 \times 10^{-02}$	1.7813	6.8979	1.6378
77ir	4f	$3.2004 \times 10^{-02}$	1.7807	6.8690	1.6364
$78 \mathrm{pt}$	3d	$3.1806 \times 10^{-02}$	1.7857	7.0167	1.6534
79au	2s	$3.2248 \times 10^{-02}$	1.7899	6.5803	1.5341
80 hg	1s	$3.5498 \times 10^{-02}$	1.7819	6.6918	1.6049
81tl	2p	$2.9054 \times 10^{-02}$	1.7445	5.0229	1.3663
82 pb	3p	$3.8974 \times 10^{-02}$	1.7778	6.9364	1.6561
83bi	4s	$4.5333 \times 10^{-02}$	1.7662	6.1580	1.5460
84po	3p	$4.8209 \times 10^{-02}$	1.7631	5.9691	1.5162
85at	2p	$5.2512 \times 10^{-02}$	1.7565	5.8363	1.5019
$86\mathrm{rn}$	1s	$5.7612 \times 10^{-02}$	1.7488	5.6871	1.4861

# Chapter 2

# **CUDAProphet**

### 2.1 GPU Architecture

While general purpose computing on graphics processing units (GPGPU) (GPUs) has been adopted by the high performance computing (HPC) community for quite some time, it can seem quite complex to the uninitiated. While most quantum chemists who decide to dip their toes into computational waters can get away with having little to no understanding of what is actually going on under the hood when programming for a central processing uint (CPU), the same can not at all be said for GPUs. Therefore, in this section I will explain what makes a GPU tick in order to help understand the terms and techniques used in the following chapter. As CUDAProphet was optimized for (and is currently hardcoded for :() the Tesla C2050, I will be referring to its specs for examples when needed. I will begin will the most

granularity possible, and zoom out so that by the end of this section, the reader should come away prepared for the rest of this chapter.

#### 2.1.1 Threads, Blocks, and Grids

The most granular element of computation on a GPU is the thread. When a CUDA function or subroutine (hereafter called a kernel) is called, it is a thread that actually executes the code. What makes GPGPU so powerful is that while a thread in a kernel is always executing the same code, the data a thread works with can be completely different between threads. This method of parallelism is called single instruction, multiple data (SIMD).

32 threads are organized into a structure called a warp. Within a warp, all threads execute code in lock-step. If a condition arises where some threads in warp must execute some code, and other threads in the same warp execute some other code (for instance in a **if eles** statement), this leads to serialization of code execution, a process called warp divergence. In a worst case scenario, this could cause all 32 threads to execute serially which could lead to a massive hit to performance. It can also cause warps to fall behind other warps and has the potential to cause race conditions to appear. If one warp need data generated by another warp that is several steps behind, this can cause all kinds of confusing errors to appear. Therefore, it is generally advisable to avoid warp divergence whenever possible. But, for programs of any considerable usefulness, warp divergence will be inevitable. Thankfully, the good people of Nvidia have included the **syncthreads** command. This

provides a barrier that all threads within a block must meet before any can continue.

The next highest structure is the thread block, or block for short. Simply put, a block is a collection of 1 or more threads. When a kernel is called, the block scheduler assigns each block to a streaming multiprocessor (SM). How many blocks a SM can run at once depends on the resources each block uses (more on this in section ??), but no more than 8 block per SM can ever be run at once on our C2050. Blocks also add an element of dimensionality to threads through the thread index (threadIdx) variable. This variable has three parts, threadIdx.x, threadIdx.y, and threadIdx.z. These variables make dealing with matrixes and arrays much simpler and can help with paralyzing large nested loops. The size of dimensions of the grid can be accessed through the blockDim variable. blockDim also has x, y, and z parts. These variables are all integer type, and the threadIdx variables range from 1 in Fortran (or 0 in C/++) to the relevant part of blockDim in Fortran (or blockDim -1 in C/++). The threadIdx set of variables are all reset at the boundary of each block.

Finally there is the grid. The grid is the entire collection of blocks that are launched for a kernel. They also give a dimensionality to blocks through the blockIdx variable which has similar parts to the threadIdx variable. It also adds in the gridDim variable which is analogous to blockDim.

## 2.2 Program flow

In this section, the control flow of a typical calculation is given.

- Step 1. The input file is read by the *intin* subroutine. If needed, the basis set and open-shell configurations are calculated by *formbs* and *find\_bin\_configurations* respectively. The options set in the input file are rewritten to stdout.
- Step 2. The calculation of small arrays and other constants is performed by calc\_parameters and bsnorm. calc\_parameters calls bcoef and setvc.
- Step 3. The mapping of threads to the unrolled one and two-electron integral matrixes are calculated by *lmpqrsa* on the GPU. All other data from step 1 and 2 is then uploaded to the GPU.
- Step 4. The one and two electron integrals are calculated on the GPU by eint1gpu and eint2gpu respectively.
- Step 5. The initial guess of the density matrix is calculated by **guess** on the GPU. This is done by the diagonalization of the one-electron Hamiltonian matrix.
- Step 6. scfiter then performs the SCF until convergence of the density matrix has been reached, or the maximum number of iterations has been reached. SCF is performed with cuSOLVER functions, as well as a few custom helper functions. The converged eigenvectors, values, and energies are downloaded from the GPU and then written to stdout.

Step 7 (optional). If jobtyp='bsopt', then an optional basis set optimization is then carried out. This starts by assigning pointers to variables on the CPU and GPU through **hookup\_cpu** and **hookup\_gpu**. Then, the four wtbs parameters are optimized by **newuoa**. **newuoa** calls **calc\_energy** which essentially reconstructs the basis set from new parameters from **newuoa**, then repeats steps 2 - 6 and feeds the energy back into **newuoa**. This repeats until optimal wtbs parameters have been found.

## 2.3 Alterations for CUDA

While almost all of the code from the original DFRATOM was modified in someway, the most extreme changes were the integral evaluation and the formation of the P and Q matrixes. Therefore, we will discuss these changes in detail in this section.

## 2.3.1 Two-electron Integrals

The original DFRATOM calculates the two electron integrals with a long series of nested for loops. Working from the outside in, the indices of the loops are L, P, Q, M, R, and S where L and M are spinor symmetries, P and Q are basis functions of symmetry L, and R and S are basis functions of symmetry M. The pseudocode for these is shown in Algorithm 1. Where nsym is the total number of symmetries, and nbs(i) is the number of basis functions for symmetry species i. Thus, each set of J and K integrals are

uniquely defined by their L, M, P, Q, R, and S values. In order to have an effective CUDA implementation of this algorithm, we to find a way to map these six numbers to CUDA threads. The simplest approach would be to map the values of L, P, and Q onto the threadIdx.x, threadIdx.y, and threadIdx.z variables for each thread, launch the needed number of blocks, and then have each thread loop over the remaining indices. The pseudocode for this can be seen in Algorithm 2. A similar method is employed by many other programs, and for larger systems it works perfectly well. But because this program is for only single atoms, problems begin to appear.

The first is the problem of warp divergence. Because the maximum values of M, R, and S depend on L, P, and Q, different threads will have a different number of loops to complete than others. Because a streaming multiprocessor (SM) must finish the block is it currently working on before it can grab another, there is the possibility that most of the threads in a block are idling while waiting for others in the same block to finish. The second problem is that with this method, there will always be several blocks that have threads that remain idle throughout the block's runtime, no matter what. This can be seen more clearly in Figure whatever. The third problem with this is that with the maximum values for L, P, and Q available for a single atom, there might not even be enough combinations to completely fill the GPU. While all of these issues begin to disappear once the number of integrals to evaluate becomes large enough, we are still very much in the range where they are in play. Therefore a smarter algorithm had to be used.

### Algorithm 1 The original

```
for L = 1 to nsym do
 for P = 1 to nbs(L) do
   for Q = 1 to P do
      for M = 1 to L do
        if L = M then
          maxr = P
        else
          maxr = nbs(M)
        end if
        for R = 1 to maxr do
          if (L = M) and (P = R) then
            maxs = Q
          else
            maxs = R
          end if
          for S = 1 to maxs do
            compute the J and K integrals of L, M, P, Q, R, and S
          end for
        end for
      end for
    end for
  end for
end for
```

### Algorithm 2 Easy Code

```
L = threadIdx.x + (blockIdx.x - 1) * blockDim.x
P = threadIdx.y + (blockIdx.y - 1) * blockDim.y
Q = threadIdx.z + (blockIdx.z - 1) * blockDim.z
if (L \leq nsym) and (P \leq nbs(L)) and (Q \leq P) then
  for M = 1 to L do
    if L = M then
      maxr = P
    else
      maxr = nbs(M)
    end if
    for R = 1 to maxr do
      if (L = M) and (P = R) then
        maxs = Q
      else
        maxs = R
      end if
      for S = 1 to maxs do
        compute the J and K integrals of L, M, P, Q, R, and S
      end for
    end for
  end for
end if
```

The second problem can be solved by opting for a one dimensional solution instead of the three dimensional one of Algorithm 2. By restricting ourselves to only using threadIdx.x, we ensure that only the last block to run will have the possibility of having threads remaining idle throughout the block's runtime. The third problem can be solved by having each thread calculate one and only one set of J and K integrals. If we start with a valid combination of L, M, P, Q, R, and S values, we can very easily figure out which thread will calculate that set of integrals by using the following equations.

$$n'(j) = \frac{j^2 + j}{2} \tag{2.1}$$

$$y = n'(nbs(L-1)) + n'(P-1) + Q$$
(2.2)

$$x = n'(nbs(M-1)) + n'(R-1) + S$$
(2.3)

$$i = n'(x) + y \tag{2.4}$$

$$i_{max} = n'(\sum_{j=1}^{nsym} n'(nbs(j)))$$
 (2.5)

nbs(0) = 0 and i would be equal to threadIdx.x + (blockIdx.x - 1) \* blockDim.x. Starting with a value of i and working our way back though is a much more challenging task. It becomes easier if we reframe it in the

following way.

Consider Figure whatever. It shows the top half of the symmetric twoelectron integral matrix for a problem with nsym = 2 and 3 basis function for symmetry one, and two for symmetry two. In each element of this matrix, there is a set of seven numbers. The top two are L and M, then P and Q, then R and S, and the last number is the value of i for the thread calculating that integral. With this, it can be seen that each element in the same row have the same M, R, and S values and the elements in the same column have the same L, P, and Q values. Therefore, finding out which column the element belongs to gives us our value for y, and finding the row give us the value of x. This can be done with the following binary search algorithm.

Where variables with the  $s_-$  prefix refer to those in the shared memory, and lownum and highnum refer to the minimum and values the i could be for the current guess (mid) of y. From here, L can be found with Algorithm 4, and P and Q can be found with Algorithm 5. The same set of algorithms can then be used to get M, R, and S by substituting the relevant variables. If there is sufficient global memory available, these values can be stored and referred to later as needed. Otherwise, they could be calculated on the fly as needed. Because binary search scales as  $\mathcal{O}(n \log n)$ , this should ensure that this remains a fast method of mapping threads to integrals for large problems as well. (CHECK WITH MARIUSZ TO BE SURE THE NEXT IS TRUE!!!!!!) With some alterations, this method could also apply to molecular symmetries other than a single atom. For instance in C1, all

#### **Algorithm 3** Binary Search for x and y

```
if theadIdx.x \leq nsym then
  s\_nsym = nsym
  s\_nbs(threadIdx.x) = nbs(threadIdx.x)
  s\_nprime(threadIdx.x) = n'(s\_nbs(threadIdx.x))
end if
call syncthreads
i = threadIdx.x + (blockIdx.x - 1) * blockDim.x
if i \leq i_{max} then
  low = 1
  high = \mathbf{sum}(s\_nprime(1:s\_nsym))
  while low \le high do mid = \frac{(low + high)}{2}
    lownum = \frac{2mid-1)(mid-2)}{2} + mid
    highnum = lownum - 1 + mid
    if (i \leq highnum) and (i \geq lownum) then
       y = mid
       exit
    else if i > highnum then
       low = mid + 1
    else if i < lownum then
       high = mid - 1
    end if
  end while
  x = i - lownum + 1
end if
```

### $\overline{\textbf{Algorithm 4}}$ Binary Search for L

```
i = threadIdx.x + (blockIdx.x - 1) * blockDim.x
if i \leq i_{max} then
  low = 1
  high = s nsym
  while low \le high \text{ do}

mid = \frac{(low + high)}{2}
    lownum = 1 + \mathbf{sum}(s\_nprime(1:mid-1))
    highnum = lownum - 1 + s\_nprime(mid)
    if (i \leq highnum) and (i \geq lownum) then
       L = mid
       exit
    else if y > highnum then
       low=mid+1
    else if y < lownum then
       high = mid - 1
    end if
  end while
end if
```

#### **Algorithm 5** Binary Search for P and Q

```
i = threadIdx.x + (blockIdx.x - 1) * blockDim.x
if i \leq i_{max} then
  low = 1
  high = s\_nbs(L)
  while low \le high do
mid = \frac{(low + high)}{2}
lownum = \frac{(mid-1)(mid-2)}{2} + mid + \mathbf{sum}(s\_nprime(1:L-1))
     highnum = lownum + mid - 1
    if (i \leq highnum) and (i \geq lownum) then
       P = mid
       exit
     else if y > highnum then
       low = mid + 1
     else if y < lownum then
       high = mid - 1
     end if
  end while
  Q = y - lownum + 1
end if
```

possible combinations of four basis functions must be used (ignoring those that appear on the bottom triangle of the two electron integral matrix of course). We could simply remove the search for L and M, have the initial value of high in Algorithm 5 be the total number of basis functions, and remove the  $\mathbf{sum}(s\_nprime(1:L-1))$  term from lownum.

From here, the code for actually evaluating the integrals remains largely the same as the original code, except for some minor changes to allow for more efficient global or shared memory access. We also use a process referred to as "grid-stride looping" where all these binary search algorithms have their if  $i \leq i_{max}$  then removed, and then are placed within the following loop: for i = threadIdx.x + (blockIdx.x - 1) \* blockDim.x to  $i_{max}$ , i += blockDim.x \* gridDim.x do. If we know the occupancy of the algorithm on the GPU beforehand, we can launch exactly the number of blocks that will fill the GPU. This reduces the overhead of block swapping and lets us further eke out some performance.

## 2.3.2 P Q Matrix Formation and SCF

In the previous subsection we discussed the calculation of the two-electron integrals, now we will discuss how they are actually used.

## 2.4 Input Description

The program can be excited on Unix-like systems in the following way

#### 2.4. INPUT DESCRIPTION

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\$ path\_to\_executable input\_file > output\_file

The input file must end in ".inp" or an error will be given. Redirection of stdout to an output file is optional, but is recommended to save the results of a calculation. The input file is read using the namelist functionality of Fortran. A description of what must appear on each line of the input file is given below. Sample input files are also given at the end of this document.

#### 1. A title of no more than 200 characters.

#### 2. \$contrl

jobtype The type of calculation to be performed.

= 'energy' Will do a single point energy calculation.

= 'bsopt' Will optimize the basis set. Can only be used if bastype equals 'wtbs'.

c The speed of light. If not given, the default is set to 137.03599976 au.

#### 3. \$nuc

znuc The charge of the nucleus.

nucmdl The nuclear model to use.

= 1 Point nucleus (default).

= 2 Finite sphere (not yet supported).

= 3 Gaussian.

rnuc The radius of the finite sphere nucleus.

alpha The exponent for the gaussian nucleus.

Defaults are given in litdata.f90.

#### 4. \$bas

nsym The number of symmetries to be used.

bastype The type of basis set given.

= 'wtbs' Use a wtbs.

= 'rdin' Read in the basis set from the input file.

ngroup The number of different groups to use in

the wtbs scheme (default 1).

The next line will depend on what bastype was set to. If bastype equals 'rdin' then the following lines must be the number of functions for the S+ symmetry, followed by the exponents to use, each on a new line. The pattern repeats for each new symmetry. See the sample input files for further clarification. Otherwise, if bastype equals 'wtbs' the \$wtbs group is read next.

#### 5. \$wtbs has to be given if bastype='wtbs'.

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wtbspara	The $\alpha$ ,	β,	$\delta$ .	and $\gamma$	wtbs	parameters.	If

there is more than one group, the order

would be  $\alpha_1$ ,  $\beta_1$ ,  $\delta_1$ ,  $\gamma_1$ ,  $\alpha_2$ ,  $\beta_2$ ,  $\delta_2$ ,  $\gamma_2$  and

so on.

nbs The number of functions used in each sym-

metry.

start Where in the  $\zeta$  pool each symme-

try starts taking exponents from (de-

fault=1,1,1,1,1,1,1).

groups What group each symmetry belongs to

(default=1,1,1,1,1,1,1).

The next line depends on what the jobtype was set to. If jobtype equals 'energy' \$newuoa is skipped and \$econfig will be read next. If jobtype equals 'bsopt', then the \$newuoa group will be needed.

6. \$newuoa has to be given if jobtype='bsopt'. Refer to the newuoa documentation for more information if needed.

rhobeg			The initial value of the trust region used
			by newuoa (default=0.1).
rhoend			The final value of the trust region used
			by newuoa. Must be smaller than rhobeg
			(default= $1.0 \times 10^{-4}$ ).
iprint			The print level for newuoa.
	=	0	No printing from newuoa (default).
	=	1	Print only when newuoa has finished.
	=	2	Print only when the trust region has de-
			creased by an order of magnitude.
	=	3	Print every iteration of newuoa.
maxfun			The maximum number of calls to calfun
			newuoa will make before terminating (de-
			fault=500).

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nclose The number of closed spinors for each

symmetry.

nopen The number of open spinors for each sym-

metry. There is a limit to one open orbital

per symmetry.

freeel The number of electrons available in the

open spinors.

autogen Automatically generates all possible com-

binations of spinor occupancies (de-

fault=.false.).

nconf The number of configurations to be read

in (needed if autogen is false).

If autogen is false, then the next nonf lines will be the spinor occupancies. They will be given as real numbers with one configuration per line. maxitr The maximum number of SCF iterations

(default=50).

ixtrp The method of extrapolation.

= 0 No extrapolation (default).

= 1 Extrapolate the Fock matrix.

dfctr Damping factor for Fock maxtrix (de-

fault=0.3).

thdll Convergence limit for the large-large

components of the density matrix

 $(default=1.0 \times 10^{-5})$ 

thdsl Convergence limit for the small-large

components of the density matrix

 $(default=1.0 \times 10^{-7})$ 

thdss Convergence limit for the small-small

components of the density matrix

 $(default=1.0 \times 10^{-9})$