The dynamical timescale:

$$t_{\rm dyn} = \sqrt{\frac{R^3}{GM}} \tag{1}$$

where R is radius, G is the Gravitational Constant, M is mass.

$$x = \frac{R}{R_s} \tag{2}$$

where R_s is the Schwarzschild radius.

$$R = R_s.x \tag{3}$$

$$R = R_s.x$$

$$= \frac{2GMx}{c^2}$$
(3)

 \Rightarrow

$$t_{\rm dyn}^2 = \frac{8x^3G^3M^3}{GMc^6}$$
 (5)
= $\frac{8x^3G^2M^2}{c^6}$ (6)

$$= \frac{8x^3G^2M^2}{c^6} \tag{6}$$

 \Rightarrow

$$t_{\rm dyn} = \frac{\sqrt{8}G}{c^3} x^{3/2} \left(\frac{M}{M_8}\right) \tag{7}$$

where M_8 is the mass of the SMBH is units of $10^8 M_{\odot}$.

Thus,

$$t_{\rm dyn} = 1400 \left(\frac{R}{R_s}\right)^{\frac{3}{2}} \left(\frac{M}{M_8}\right) \text{ seconds.}$$
 (8)