

1219 - NIRSpec and MIRI spectroscopy of QSOs - part #3

Cycle: 1, Proposal Category: GTO

INVESTIGATORS

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OBSERVATIONS

Folder	Observation	Label	Observing Template	Science Target							
NIRSpec MSA of J1342											
	6	Quasar in S200A1 and S200A2	NIRSpec MultiObject Spectroscopy	(4) TARGET-OBSERVATION-6							
NIRSpe	ec IFU of J1342										
	7	NIRSpec IFU	NIRSpec IFU Spectroscopy	(2) ULASJ1342+0928							
MIRI o	MIRI observations of J1342										
	(2) ULASJ1342+0928										
	5	J1342-Imager	MIRI Imaging	(5) ULASJ1342+0928-IMAGER							

JWST Proposal 1219 (Created: Wednesday, June 26, 2019 at 1:01:05 PM Eastern Standard Time) - Overview

ABSTRACT

Update June 2019:

We have removed target j2348-3054 and added ULASJ1342+0928

NIRSpec-IFU observations now also include R100

Note that the 'charged time' is larger than the 'allocated time'

because we have internally redistributed the time among the NIRSpec-IFU GTO programmes

MIRI, with its spectral coverage from 5 to 28 um and sensitivity, is the only instrument onboard JWST able to explore the optical and near-infrared spectrum and light distribution of galaxies and QSOs at redshifts above 6.7. A complete 5 to 28 spectrum (~0.6 to 3.5 microns rest-frame) of the highest redshift (z=7.54) QSO J1342+0928, will be obtained, together with F560W MIRI imaging of the host and surrounding field. In addition simultaneous MIRI imaging of a nearby field will be taken with the F560W and F1000W filters.

The same APT file includes the NIRSpec observation of the same target with the IFU with the G395H grating (aimed primarily at mapping the main optical nebular lines Hbeta, [OIII], Halpha, [NII]) and with the fixed slit with the G140H and G235H gratings (aimed primarily at detecting IGM metal absorption systems). Simultaneously with the fixed slit observation (centered onto the quasar) the MSA will be used to observe galaxies imaged by the HST in the field of view, therefore these observations have been set up in MOS mode.

OBSERVING DESCRIPTION

NIRSpec MSA OBSERVATION

This corresponds to NIRSpec Proposal IDs: FERRUIT_4004 and FERRUIT_4104

(NIRSpec Contact Person: Chris Willott, chriswillott1@gmail.com)

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The quasar will be placed in the fixed slits S200A1 and S200A2 whilst simultaneously configuring MSA shutters to target other galaxies identified in HST imaging. We use NIRSpec team software to design the MSA configurations and ensure they agree with positions in the APT MPT software.

We use the G140H/F070LP grating and filter combination to do spectroscopy at 0.7 to 1.8 microns. We realise there will be some spectral overlap at >1.4 microns but our prime targets have almost zero flux below 0.9 microns so this will not strongly affect the spectra. We also use the G235H/F170LP grating and filter combination to do spectroscopy at 1.7 to 3.1 microns.

NIRSpec IFU OBSERVATION

This corresponds to NIRSpec Proposal ID: FERRUIT_3054

(NIRSpec Contact Persons: Roberto Maiolino, r.maiolino@mrao.cam.ac.uk

The NIRSpec IFU observation is done with the G395H grating and it is aimed aimed primarily at mapping the strongest optical nebular lines (Hbeta, [OIII], Halpha, [NII]).

We also include R100 observation (+ 1 leakcal) for the continuum (and the fluxes of other lines)

We are using no target acquisition (i.e. point-and-shoot).

At any of the constrained PA range there are Gaia GS that can be selected for guiding and which will ensure the proper location of the target within the IFU aperture, with the required accuracy.

We are using NRSIRS2RAPID for a better identification and rejection of cosmic rays.

MIRI OBSERVING DESCRIPTIONS:

JWST Proposal 1219 (Created: Wednesday, June 26, 2019 at 1:01:05 PM Eastern Standard Time) - Overview

This corresponds to MIRI Proposal ID: WRIGHT_0601 and WRIGHT_0602

(MIRI Contact Person: Javier Alvarez-Marquez, javier.alvarez@cab.inta-csic.es)

The purpose of the program is to get a full 5 - 30 um spectrum of J2348 using the 3 MRS configurations with simultaneous Imager observations (see additional note 1). In addition, we request the imaging of the target in two filters: F560W, F770W.

The dithering strategies (4-pt, point source) were selected to optimize the PSF and detector effects in all MRS channels, and IMAGER filters. These strategies could be subject to change without modifying the total time.

Proposal 1219 - Targets - NIRSpec and MIRI spectroscopy of QSOs - part #3

	#	Name	Target Coordinates	Targ. Coord. Corrections	Miscellaneous
	(2)	ULASJ1342+0928	RA: 13 42 8.0970 (205.5337375d)		
			Dec: +09 28 38.28 (9.47730d)		
			Equinox: J2000		
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argets	(4)	TARGET-OBSERVATION-6	RA: 13 42 11.2177 (205.5467404d)		
JE			Dec: +09 27 14.36 (9.45399d)		
1 -			Equinox: J2000		
ixed	Comments:	This target was generated automa	tically for MSA Observation 6		
Ê	Category=0 Description Extended=1	=[High-redshift galaxies, Primora	lial galaxies, Quasars]		
1	(5)	ULASJ1342+0928-IMAGER	RA: 13 42 7.1000 (205.5295833d)		
			Dec: +09 28 48.00 (9.48000d)		
			Equinox: J2000		
	Comments:	7.1			
	Category=0 Description				
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	3	LONG(C)	MRSSHORT		SLOW	36	1	1	Dither 1	4	4	3440.148	

<u>Pro</u>	pposal 1219 - Observation 4 - NIRSpec and MIRI spectroscopy of QSOs - part #3
ents	Aperture PA Range 90 to 150 Degrees (V3 90.0 to 150.0) Background Limited. Background no more than 40% above minimum
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