

# 1219 - NIRSpec and MIRI spectroscopy of QSOs - part #3

Cycle: 1, Proposal Category: GTO

## **INVESTIGATORS**

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## **OBSERVATIONS**

Folder	Observation	Label	Observing Template	Science Target						
NIRSpec MSA of J2348										
	6	Quasar in S200A1 and S200A2 v2	NIRSpec MultiObject Spectroscopy	(4) TARGET-OBSERVATION-6						
NIRSpe	ec IFU of J2348 1	new								
	3	J2348 with NIRSpec IF U	NIRSpec IFU Spectroscopy	(2) J2348-3054						
MIRI o	bservations of J2	348 new								
	4	J2348	MIRI Medium Resolution Spectroscopy	(2) J2348-3054						
	5	J2348-Imager	MIRI Imaging	(3) J2348-3054-IMAGER						

JWST Proposal 1219 (Created: Tuesday, February 20, 2018 6:19:00 PM EST) - Overview

### **ABSTRACT**

This proposal is associated with:

NIRSpec Proposal IDs: FERRUIT\_4004 and FERRUIT\_4104

NIRSpec Proposal ID: FERRUIT\_3054

MIRI Proposal ID: WRIGHT\_0601 and WRIGHT\_0602

MIRI, with its spectral coverage from 5 to 28 um and sensitivity, is the only instrument onboard JWST able to explore the optical and near-infrared spectrum and light distribution of galaxies and QSOs at redshifts above 6.7. A complete 5 to 28 spectrum (~0.6 to 3.5 microns rest-frame) of the IR-luminous QSO J2348-3054 (z=6.9018), will be obtained, together with, MIRI imaging to map the rest-frame near-IR light distribution of the host galaxy at ~0.2-0.3 arcsec angular resolution.

The same APT file includes the NIRSpec observation of the same target with the IFU with the G395H grating (aimed primarily at mapping the primary optical nebular lines Hbeta, [OIII], Halpha, [NII]) and with the fixed slit with the G140H grating (aimed primarily at detecting IGM metal absorption systems). Simultaneously with the fixed slit observation (centered onto the quasar) the MSA will be used to observe galaxies imaged by the HST in the field of view.

#### **OBSERVING DESCRIPTION**

NIRSpec MSA OBSERVATION

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This corresponds to NIRSpec Proposal IDs: FERRUIT\_4004 and FERRUIT\_4104

(NIRSpec Contact Person: Chris Willott, chriswillott1@gmail.com)

The quasar will be placed in the fixed slits S200A1 and S200A2 whilst simultaneously configuring MSA shutters to target other galaxies identified in HST imaging. We use NIRSpec team software to design the MSA configurations and ensure they agree with positions in the APT MPT software.

We use the G140H/F070LP grating and filter combination to do spectroscopy at 0.7 to 1.8 microns. We realise there will be some spectral overlap at >1.4 microns but our prime targets have almost zero flux below 0.9 microns so this will not strongly affect the spectra.

JWST P	roposal '	1219 (	Created:	Tuesday,	February	20,	2018	6:19:00	PM ES	T) - Overvie	W
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NIRSpec IFU OBSERVATION

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This corresponds to NIRSpec Proposal ID: FERRUIT\_3054

(NIRSpec Contact Person: Roberto Maiolino, r.maiolino@mrao.cam.ac.uk)

The NIRSpec IFU observation is done with the G395H grating and it is aimed aimed primarily at mapping the strongest optical nebular lines (Hbeta, [OIII], Halpha, [NII]).

We are using no target acquisition (i.e. point-and-shoot).

At any of the constrained PA range there are Gaia GS that can be selected for guiding and which will ensure the proper location of the target within the IFU aperture, with the required accuracy.

We are using NRSIRS2RAPID for a better identification and rejection of cosmic rays.

#### MIRI OBSERVING DESCRIPTIONS:

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This corresponds to MIRI Proposal ID: WRIGHT\_0601 and WRIGHT\_0602

(MIRI Contact Person: Javier Alvarez-Marquez, javier.alvarez@cab.inta-csic.es)

The purpose of the program is to get a full 5 - 30 um spectrum of J2348 using the 3 MRS configurations with simultaneous Imager observations (see

JWST Proposal 1219 (Created: Tuesday, February 20, 2018 6:19:00 PM EST) - Overview additional note 1). In addition, we request the imaging of the target in two filters: F560W, F770W.

The dithering strategies (4-pt, point source) were selected to optimize the PSF and detector effects in all MRS channels, and IMAGER filters. These strategies could be subject to change without modifying the total time.

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The PA\_V3 constraints (30 < PA\_V3 (deg) < 43) for MIRI and NIRSpec IFU observations were done not to obtain:

- bright stars located in the IMAGER of MIRI that can saturate
- bright stars located in the MOS area that can contaminate the NIRSpec IFU observation

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