

■ Scientific Justification

Exploring the Red Sequence of AGN with the *James Webb Space Telescope*

Things We Know About AGN/Quasars::

- Blue Quasars are in enhanced distrubed systems at $z \sim 0.7$ (Villforth et al.)
- “Red” quasars generally merger at $z \sim 2.5$ (Glikman et al. 2016), but “red” here is a somewhat red+radio definition...
- Peak of optical QLF at $z \sim 2 - 3$ (Richards et al. 2006; Ross et al. 2014)
- There is a trend of radio fraction in QSOs with $(g - i)$ colour; the redder the colour, the larger the radio fraction (Klindt et al. 2018)

Things We *don't* Know About AGN/Quasars::

- The host properties of SDSS/BOSS $z = 2 - 3$ QSOs
- Is there a *range* in red quasar host properities??
- Is there a “*transition colour*” above which mergers are enhanced?
- Is there a *transitional Radio Loudness* above which mergers are enhanced?

General Idea::

NIRCam Imaging, and/or NIRSpec spectroscopy (Long Slit? IFU?) of a sample of “red” to “extremely red” quasars.

- What are the host galaxy morphologies of Red Quasars?
- Are “Red” quasars more distrubed than “Extremely Red” quasars?
- Are red radio-loud quasars in different hosts than red radio-quiet quasars??
- Are the narrow lines offset from the broadlines

General Sample::

X-Shooter Red Quasar Sample (Radio Loud? Radio Quiet? TBD...)

“Core” ERQs from Hamann et al. (2017). *i*-W3 selected, with C ivEW selection too.

Select the subset of “core” ERQs that are still *r*-W4 objects...??

“Hot DOGs” (aka W1W2-drops)

Questions to answer/things to address::

- Why not *HST*?? Want to go redder than e.g. F160W (*H*-Short at $1.545\mu\text{m}$, FWHM= $0.29\mu\text{m}$)
- Why not *ALMA*?? Will/can use ALMA for e.g. SFRs instead of MIRI.

“Cool Ideas....”

- Hopkins (2008) Figure 1, for real, for the Red objects, at $z \approx 2.5$.

■ Technical Justification

Sample is given in Figure 2. There are 11 QSOs with confirmed spectroscopic redshifts, with redshift range $5.0 < z < 6.7$ and strong, $\text{SNR} \geq 5$ WISE detections in the W3 $12\mu\text{m}$ band.

Things to think about::

NIRSpec vs. NIRISS?

NIRSpec since it has the higher resolution modes

Things to think about::

NIRSpec IFU vs. NIRSpec fixed slits (FS) ??

Both have \approx the same wavelength coverage. Need to run ETC.

Our targets are well spaced in R.A. and Decl.

■ Special Requirements (if any)

■ Justify Coordinated Parallel Observations (if any)

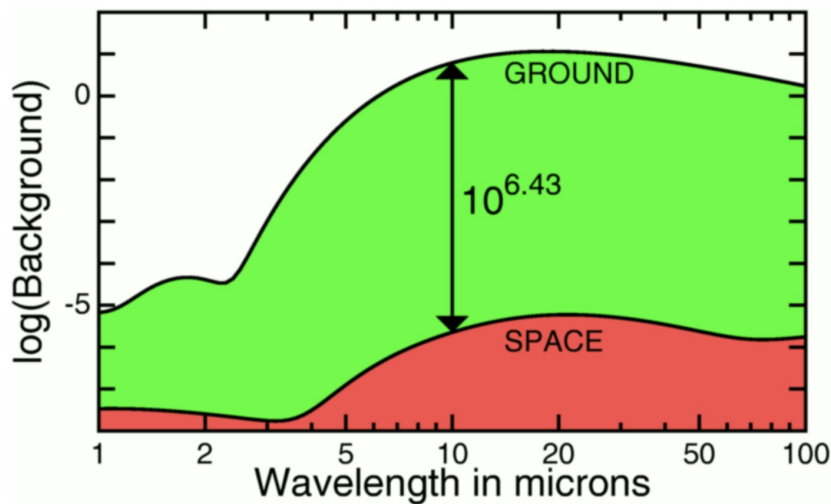
■ Justify Duplications (if any)

■ Data Processing & Analysis Plan (AR only)

■ Management Plan (AR only)

Why Space?

“Ground-based infrared astronomy is like observing stars in broad daylight with a telescope made out of fluorescent lights” — George Rieke.



40 cm WISE
telescope in
space equals
six thousand
8-meter
telescopes on
the ground!

Figure 1: Ned Wright's talk; <https://www.ipac.caltech.edu/exgal2011/sched.shtml>

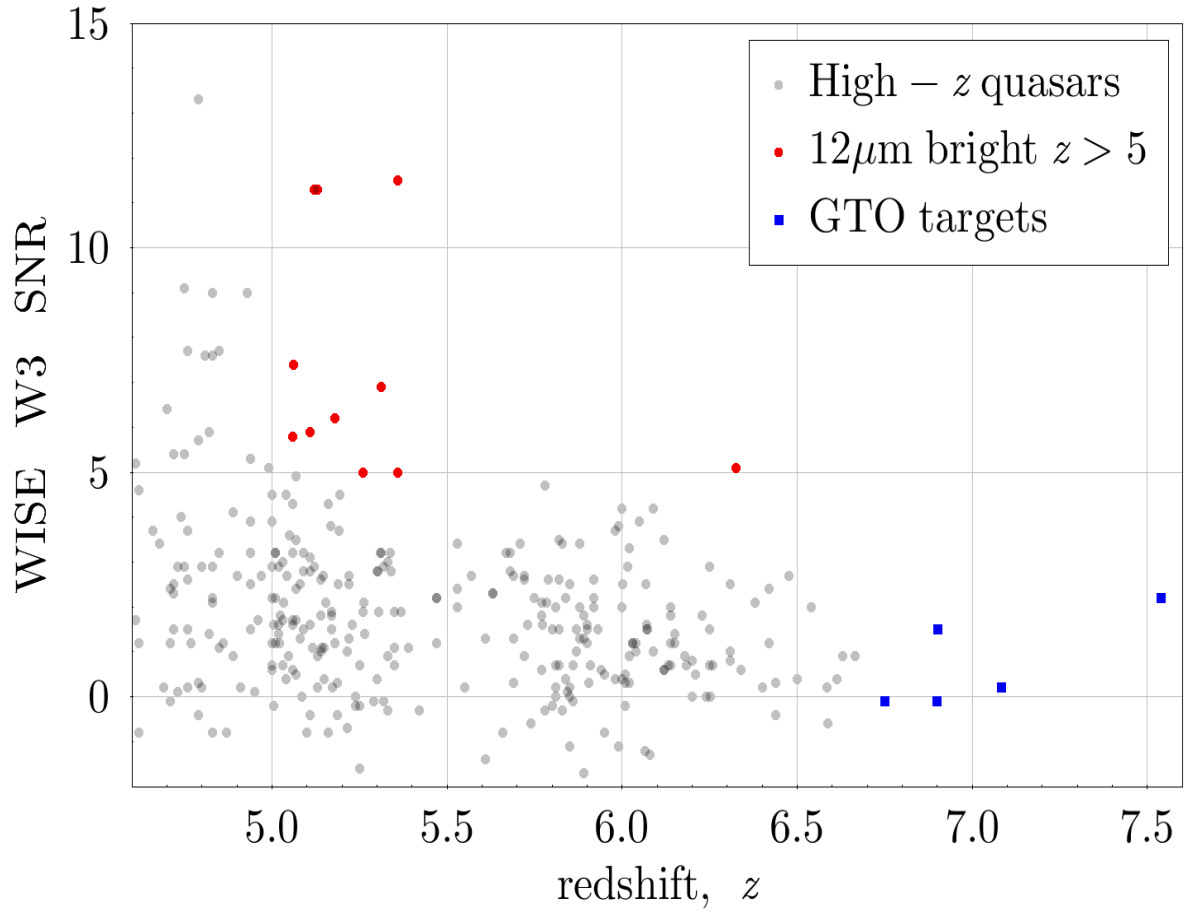


Figure 2: The WISE W3 SNR values for all spectroscopically confirmed $z \geq 5$ quasars. Our sample are the red circles; the GTO reserved targets are the blue squares.