



1178 - Ram Pressure Stripping in ESO 137-001

Cycle: 1, Proposal Category: GTO

INVESTIGATORS

<i>Name</i>	<i>Institution</i>	<i>E-Mail</i>
Dr. Scott D. Friedman (PI)	Space Telescope Science Institute	friedman@stsci.edu
Dr. Stacey Alberts (CoI) (Contact)	University of Arizona	salberts@email.arizona.edu
Dr. George Rieke (CoI)	University of Arizona	ghrieke@gmail.com

OBSERVATIONS

<i>Folder</i>	<i>Observation</i>	<i>Label</i>	<i>Observing Template</i>	<i>Science Target</i>
Observation Folder				
	1	Inner Tail [MRS+Imaging]	MIRI Medium Resolution Spectroscopy	(18) Group INNER-TAIL
	2	Outer Tail [MRS+Imaging]	MIRI Medium Resolution Spectroscopy	(21) Group OUTER-TAIL-BACKGROUND

ABSTRACT

Once thought to be rare, the number of known ram pressure stripping (RPS) events has been steadily rising, observed as truncated or disturbed gaseous disks or one-sided tails in the X-ray through the radio. These events hold key information regarding the relation between galaxy transformation and environment. We propose MIRI MRS observations of ESO 137-001, a well studied local galaxy ($z=0.01625$) with a spectacular double ram pressure stripped tail. At both high spectral ($R\sim 2700$) and spatial (~ 0.1 arcsec) resolution over 5-28.8m, we will detect multiple transitions of rotational H₂ lines as well as a suite of fine structure lines at high significance. From these observables, we will deduce the kinematics and the temperature/density structure of warm and hot gas components in the tail on sub-kpc scales as well as the excitation mechanism(s) responsible. This information will reveal how the (star forming) interstellar medium of the host galaxy responds to strong RPS and how the stripped gas subsequently interacts with the intra-cluster medium. Notably, the detailed state of H₂ will identify the spatial extent of shocked gas and constrain the mechanisms and timescales for the cooling of molecular gas, revealing whether star-forming regions in the tail were formed in situ or

from molecular gas stripped directly from the galactic disk. Additionally, high resolution MIRI 7.7m imaging obtained simultaneously with MRS pointings in the far-tail (~40 kpc from the main galaxy) will fall back on the main galaxy and near- to mid-tail regions, providing a measure of the aromatic features and a SFR indicator.

FRIEDMAN_0001

OBSERVING DESCRIPTION

This program targets multiple regions at the galaxy-tail interface and along the stripped tail of ESO 137-001 with the MIRI MRS. The full MRS wavelength range with all channels will be observed. The 16 science targets have been grouped into the inner tail region (14 targets in target group INNER-TAIL) and the outer tail region (2 targets in a 2x1 mosaic). Our science goal is to detect and characterize warm molecular hydrogen emission and fine structure lines in key regions, selected HII regions with evidence for star formation and/or cold molecular gas, at the galaxy-tail interface and along the RPS tail. Our primary objective is the H2 S(1) emission line at 17.035um (rest). As this line is in the MRS channel 3 at $z=0.01625$, our pointings and dither strategy are optimized for this channel. Secondary objectives include the H2 S(2)-S(7) lines, fine structure lines such as [NeII]-[NeVI], [SIII], [OIV], [FeII], etc. and PAH features. The H2 S(0) emission line, which falls at the low sensitivity edge of channel 4 long, will be recovered if possible using spatial and spectral binning.

One dedicated background pointing, with the same exposure setup as our science pointings, is included in the target group OUTER-TAIL for calibration purposes in order to ensure proper subtraction of the sky/telescope background. This background pointing must be taken during the same epoch as our observations and therefore this entire program is non-interruptible and needs to be executed at the same PA.

MIRI imaging, taken simultaneously with MRS mosaic in the outer tail, will provide a F770W image of the galaxy ESO 137-001 and its RPS tail out to a distance of ~25 kpc. This imaging requires a position angle of ~334-346 deg (V3). For this APT file, we have fixed the PA to 334 deg (V3) to satisfy the special requirement that target groups have a fix, zero-width PA range. Since this imaging can improve spatial registration of the MRS exposures, we include it in the inner tail region as well at a minimum increase to program duration.

Given that our targets are extended, the absolute pointing accuracy of JWST is adequate for our needs and target acquisition is not required for this program.

Proposal 1178 - Targets - Ram Pressure Stripping in ESO 137-001

Fixed Targets	#	Name	Target Coordinates	Targ. Coord. Corrections	Miscellaneous
	(2)	KNOT-2	RA: 16 13 24.9680 (243.3540333d) Dec: -60 45 43.51 (-60.76209d) Equinox: J2000 <i>Comments: Coordinates from Sivanandam+10, Table 4.</i> <i>We are observing one background (fixed target 10, observation 2) for all MRS pointings.</i> <i>Category=ISM</i> <i>Description=[H II regions, Molecular gas]</i> <i>Extended=YES</i>		
	(3)	KNOT-3	RA: 16 13 26.4800 (243.3603333d) Dec: -60 45 32.85 (-60.75912d) Equinox: J2000 <i>Comments: Coordinates from Sivanandam+10, Table 4.</i> <i>We are observing one background (fixed target 10, observation 2) for all MRS pointings.</i> <i>Category=ISM</i> <i>Description=[H II regions]</i> <i>Extended=YES</i>		
	(4)	KNOT-4	RA: 16 13 25.4500 (243.3560417d) Dec: -60 45 35.19 (-60.75977d) Equinox: J2000 <i>Comments: Coordinates from Sivanandam+10, Table 4.</i> <i>We are observing one background (fixed target 10, observation 2) for all MRS pointings.</i> <i>Category=ISM</i> <i>Description=[H II regions, Molecular gas]</i> <i>Extended=YES</i>		
	(5)	KNOT-5	RA: 16 13 23.9300 (243.3497083d) Dec: -60 45 53.60 (-60.76489d) Equinox: J2000 <i>Comments: Coordinates from Sivanandam+10, Table 4.</i> <i>We are observing one background (fixed target 10, observation 2) for all MRS pointings.</i> <i>Category=ISM</i> <i>Description=[H II regions]</i> <i>Extended=YES</i>		
	(6)	KNOT-6	RA: 16 13 24.5400 (243.3522500d) Dec: -60 46 5.92 (-60.76831d) Equinox: J2000 <i>Comments: Coordinates from Sivanandam+10, Table 4.</i> <i>We are observing one background (fixed target 10, observation 2) for all MRS pointings.</i> <i>Category=ISM</i> <i>Description=[H II regions, Molecular gas]</i> <i>Extended=YES</i>		
	(7)	KNOT-7	RA: 16 13 24.1300 (243.3505417d) Dec: -60 45 33.46 (-60.75929d) Equinox: J2000 <i>Comments: Coordinates from Sivanandam+10, Table 4.</i> <i>We are observing one background (fixed target 10, observation 2) for all MRS pointings.</i> <i>Category=ISM</i> <i>Description=[H II regions, Molecular gas]</i> <i>Extended=YES</i>		

Proposal 1178 - Targets - Ram Pressure Stripping in ESO 137-001

(8)	KNOT-11	RA: 16 13 22.8300 (243.3451250d) Dec: -60 45 20.55 (-60.75571d) Equinox: J2000 <i>Comments: Coordinates from Sivanandam+10, Table 4.</i> <i>We are observing one background (fixed target 10, observation 2) for all MRS pointings.</i> <i>Category=ISM</i> <i>Description=[H II regions, Molecular gas]</i> <i>Extended=YES</i>
(10)	BACKGROUND	RA: 16 13 11.5000 (243.2979167d) Dec: -60 45 0.00 (-60.75000d) Equinox: J2000 <i>Comments: Chosen to have no 8um emission.</i> <i>Category=Calibration</i> <i>Description=[Telescope/sky background]</i> <i>Extended=YES</i>
(11)	ESO137-001-TILE-2	RA: 16 13 26.1575 (243.3589896d) Dec: -60 45 42.12 (-60.76170d) Equinox: J2000 <i>Comments: Tile in mosaic of galaxy-tail interface covering the base of the ram pressure stripped tail of ESO137-001. Mosaic was designed using the APT mosaic tool with 15% overlap, but then the tiles were split into individual targets in order to include them in Target Group INNER-TAIL. This allows us to observe the mosaic and Knots 2, 3, 4, 5, 6, 7, 11 in the same visit as they are within the visit splitting distance.</i> <i>We are observing one background (fixed target 10, observation 2) for all MRS pointings.</i> <i>Category=Galaxy</i> <i>Description=[Spiral galaxies, Tidal tails]</i> <i>Extended=YES</i>
(12)	ESO137-001-TILE-4	RA: 16 13 25.7573 (243.3573221d) Dec: -60 45 49.48 (-60.76374d) Equinox: J2000 <i>Comments: Tile in mosaic of galaxy-tail interface covering the base of the ram pressure stripped tail of ESO137-001. Mosaic was designed using the APT mosaic tool with 15% overlap, but then the tiles were split into individual targets in order to include them in Target Group INNER-TAIL. This allows us to observe the mosaic and Knots 2, 3, 4, 5, 6, 7, 11 in the same visit as they are within the visit splitting distance.</i> <i>We are observing one background (fixed target 10, observation 2) for all MRS pointings.</i> <i>Category=Galaxy</i> <i>Description=[Active galactic nuclei, Spiral galaxies, Tidal tails]</i> <i>Extended=YES</i>
(13)	ESO137-001-TILE-5	RA: 16 13 26.4339 (243.3601412d) Dec: -60 45 47.68 (-60.76324d) Equinox: J2000 <i>Comments: Tile in mosaic of galaxy-tail interface covering the base of the ram pressure stripped tail of ESO137-001. Mosaic was designed using the APT mosaic tool with 15% overlap, but then the tiles were split into individual targets in order to include them in Target Group INNER-TAIL. This allows us to observe the mosaic and Knots 2, 3, 4, 5, 6, 7, 11 in the same visit as they are within the visit splitting distance.</i> <i>We are observing one background (fixed target 10, observation 2) for all MRS pointings.</i> <i>Category=Galaxy</i> <i>Description=[Spiral galaxies, Tidal tails]</i> <i>Extended=YES</i>

Proposal 1178 - Targets - Ram Pressure Stripping in ESO 137-001

(14)	ESO137-001-TILE-6	RA: 16 13 27.1105 (243.3629604d) Dec: -60 45 45.88 (-60.76274d) Equinox: J2000
<p><i>Comments: Tile in mosaic of galaxy-tail interface covering the base of the ram pressure stripped tail of ESO137-001. Mosaic was designed using the APT mosaic tool with 15% overlap, but then the tiles were split into individual targets in order to include them in Target Group INNER-TAIL. This allows us to observe the mosaic and Knots 2, 3, 4, 5, 6, 7, 11 in the same visit as they are within the visit splitting distance.</i></p> <p><i>We are observing one background (fixed target 10, observation 2) for all MRS pointings.</i></p> <p>Category=Galaxy Description=[Spiral galaxies, Tidal tails] Extended=YES</p>		
(15)	ESO137-001-TILE-7	RA: 16 13 26.0337 (243.3584737d) Dec: -60 45 55.05 (-60.76529d) Equinox: J2000
<p><i>Comments: Tile in mosaic of galaxy-tail interface covering the base of the ram pressure stripped tail of ESO137-001. Mosaic was designed using the APT mosaic tool with 15% overlap, but then the tiles were split into individual targets in order to include them in Target Group INNER-TAIL. This allows us to observe the mosaic and Knots 2, 3, 4, 5, 6, 7, 11 in the same visit as they are within the visit splitting distance.</i></p> <p><i>We are observing one background (fixed target 10, observation 2) for all MRS pointings.</i></p> <p>Category=Galaxy Description=[Spiral galaxies, Tidal tails] Extended=YES</p>		
(16)	ESO137-001-TILE-8	RA: 16 13 26.7103 (243.3612929d) Dec: -60 45 53.24 (-60.76479d) Equinox: J2000
<p><i>Comments: Tile in mosaic of galaxy-tail interface covering the base of the ram pressure stripped tail of ESO137-001. Mosaic was designed using the APT mosaic tool with 15% overlap, but then the tiles were split into individual targets in order to include them in Target Group INNER-TAIL. This allows us to observe the mosaic and Knots 2, 3, 4, 5, 6, 7, 11 in the same visit as they are within the visit splitting distance.</i></p> <p><i>We are observing one background (fixed target 10, observation 2) for all MRS pointings.</i></p> <p>Category=Galaxy Description=[Spiral galaxies, Tidal tails] Extended=YES</p>		
(17)	ESO137-001-TILE-9	RA: 16 13 27.3869 (243.3641121d) Dec: -60 45 51.44 (-60.76429d) Equinox: J2000
<p><i>Comments: Tile in mosaic of galaxy-tail interface covering the base of the ram pressure stripped tail of ESO137-001. Mosaic was designed using the APT mosaic tool with 15% overlap, but then the tiles were split into individual targets in order to include them in Target Group INNER-TAIL. This allows us to observe the mosaic and Knots 2, 3, 4, 5, 6, 7, 11 in the same visit as they are within the visit splitting distance.</i></p> <p><i>We are observing one background (fixed target 10, observation 2) for all MRS pointings.</i></p> <p>Category=Galaxy Description=[Spiral galaxies, Tidal tails] Extended=YES</p>		
(18)	Group INNER-TAIL	
<p><i>Comments:</i> <i>Target Selection=[2 KNOT-2, 3 KNOT-3, 4 KNOT-4, 5 KNOT-5, 6 KNOT-6, 7 KNOT-7, 8 KNOT-11, 11 ESO137-001-TILE-2, 12 ESO137-001-TILE-4, 13 ESO137-001-TILE-5, 14 ESO137-001-TILE-6, 15 ESO137-001-TILE-7, 16 ESO137-001-TILE-8, 17 ESO137-001-TILE-9]</i></p>		
(19)	OUTER-TAIL-1	RA: 16 13 14.6400 (243.3110000d) Dec: -60 44 45.30 (-60.74592d) Equinox: J2000
<p><i>Comments: We are observing one background (fixed target 10, observation 2) for all MRS pointings.</i></p> <p>Category=ISM Description=[H II regions, Molecular gas]</p>		

Proposal 1178 - Targets - Ram Pressure Stripping in ESO 137-001

(20)	OUTER-TAIL-2	RA: 16 13 14.3800 (243.3099167d) Dec: -60 44 39.40 (-60.74428d) Equinox: J2000 <i>Comments: We are observing one background (fixed target 10, observation 2) for all MRS pointings.</i> <i>Category=ISM</i> <i>Description=[H II regions, Molecular gas]</i>
(21)	Group OUTER-TAIL- BACKGROUND	<i>Comments:</i> <i>Target Selection=[10 BACKGROUND, 19 OUTER-TAIL-1, 20 OUTER-TAIL-2]</i>

Proposal 1178 - Observation 1 - Ram Pressure Stripping in ESO 137-001

Observation	Proposal 1178, Observation 1: Inner Tail [MRS+Imaging]												Tue Jun 25 20:00:26 GMT 2019		
	Diagnostic Status: Warning														
	Observing Template: MIRI Medium Resolution Spectroscopy														
	Comments: PA is set to 334 because target groups require a zero-width PA range as a special requirement. However, as long as all pointings have the same PA, any PA between 334 and 346 deg (V3) would be suitable for this program.														
We are observing one background (fixed target 10, observation 2) for all MRS pointings.															
Diagnostics	(Visit 1:1) Warning (Form): Data volume for this visit 45858.19 MB exceeds half the maximum allowed of 58000.0 MB.														
	(Visit 1:1) Warning (Form): Overheads are provisional until the Visit Planner has been run.														
Fixed Targets	#	Name				Target Coordinates				Targ. Coord. Corrections				Miscellaneous	
	(18)	Group INNER-TAIL													
Comments:															
Target Selection=[2 KNOT-2, 3 KNOT-3, 4 KNOT-4, 5 KNOT-5, 6 KNOT-6, 7 KNOT-7, 8 KNOT-11, 11 ESO137-001-TILE-2, 12 ESO137-001-TILE-4, 13 ESO137-001-TILE-5, 14 ESO137-001-TILE-6, 15 ESO137-001-TILE-7, 16 ESO137-001-TILE-8, 17 ESO137-001-TILE-9]															
Acquisition	#											Target			
	1											NONE			
Template	AcqFilter		Primary Channel				Simultaneous Imaging				Imager Subarray				
	F560W		CHANNEL3				YES				FULL				
Dithers	#	Dither Type				Optimized For				Direction					
	1	2-Point				EXTENDED SOURCE				NEGATIVE					
Spectral Elements	#	Wavelength Range	Detector	Filter	Readout Pattern	Groups/Int	Integrations/Exp	Exposures/Dith	Dither	Total Dithers	Total Integrations	Total Exposure Time	ETC Wkbk.Calc ID		
	1		IMAGER	F770W	FAST	10	1	1	Dither 1	2	2	55.501			
	1	SHORT(A)	MRSLONG		FAST	95	1	1	Dither 1	2	2	527.258	25204		
	1	SHORT(A)	MRSSHORT		FAST	95	1	1	Dither 1	2	2	527.258	25204		
	2		IMAGER	F770W	FAST	10	1	1	Dither 1	2	2	55.501			
	2	MEDIUM(B)	MRSLONG		FAST	95	1	1	Dither 1	2	2	527.258	25204		
	2	MEDIUM(B)	MRSSHORT		FAST	95	1	1	Dither 1	2	2	527.258	25204		
	3		IMAGER	F770W	FAST	10	1	1	Dither 1	2	2	55.501			
	3	LONG(C)	MRSLONG		FAST	95	1	1	Dither 1	2	2	527.258	25204		
	3	LONG(C)	MRSSHORT		FAST	95	1	1	Dither 1	2	2	527.258	25204		

Proposal 1178 - Observation 1 - Ram Pressure Stripping in ESO 137-001

Special Requirements	Aperture PA Range 334 to 334 Degrees (V3 334.0 to 334.0) Sequence Observations 1, 2, Non-interruptible
----------------------	---

Proposal 1178 - Observation 2 - Ram Pressure Stripping in ESO 137-001

Observation	Proposal 1178, Observation 2: Outer Tail [MRS+Imaging]												Tue Jun 25 20:00:26 GMT 2019	
	Diagnostic Status: Warning													
	Observing Template: MIRI Medium Resolution Spectroscopy													
	Comments: PA is set to 334 because target groups require a zero-width PA range as a special requirement. However, as long as all pointings have the same PA, any PA between 334 and 346 deg (V3) would be suitable for this program.													
We are observing one background (fixed target 10, observation 2) for all MRS pointings.														
Diagnostics	(Visit 2:1) Warning (Form): Overheads are provisional until the Visit Planner has been run.													
Fixed Targets	#	Name			Target Coordinates			Targ. Coord. Corrections			Miscellaneous			
	(21)	Group OUTER-TAIL-BACKGROUND												
Comments: Target Selection=[10 BACKGROUND, 19 OUTER-TAIL-1, 20 OUTER-TAIL-2]														
Acquisition	#	Target												
	1	NONE												
Template	AcqFilter	Primary Channel			Simultaneous Imaging			Imager Subarray						
	F560W	CHANNEL3			YES			FULL						
Dithers	#	Dither Type			Optimized For			Direction						
	1	2-Point			EXTENDED SOURCE			NEGATIVE						
Spectral Elements	#	Wavelength Range	Detector	Filter	Readout Pattern	Groups/Int	Integrations/Exp	Exposures/Dith	Dither	Total Dithers	Total Integrations	Total Exposure Time	ETC Wkbk.Calc ID	
	1		IMAGER	F770W	FAST	95	1	1	Dither 1	2	2	527.258	25204	
	1	SHORT(A)	MRSLONG		FAST	95	1	1	Dither 1	2	2	527.258	25204	
	1	SHORT(A)	MRSSHORT		FAST	95	1	1	Dither 1	2	2	527.258	25204	
	2		IMAGER	F770W	FAST	95	1	1	Dither 1	2	2	527.258	25204	
	2	MEDIUM(B)	MRSLONG		FAST	95	1	1	Dither 1	2	2	527.258	25204	
	2	MEDIUM(B)	MRSSHORT		FAST	95	1	1	Dither 1	2	2	527.258	25204	
	3		IMAGER	F770W	FAST	95	1	1	Dither 1	2	2	527.258	25204	
	3	LONG(C)	MRSLONG		FAST	95	1	1	Dither 1	2	2	527.258	25204	
	3	LONG(C)	MRSSHORT		FAST	95	1	1	Dither 1	2	2	527.258	25204	

Proposal 1178 - Observation 2 - Ram Pressure Stripping in ESO 137-001

Special Requirements	<p>Aperture PA Range 334 to 334 Degrees (V3 334.0 to 334.0)</p> <p>Sequence Observations 1, 2, Non-interruptible</p>
-----------------------------	--