JWST Proposal 1207 (Created: Tuesday, June 25, 2019 at 8:00:34 AM Eastern Standard Time) - Overview



1207 - MIRI in the Hubble Ultra-Deep Field

Cycle: 1, Proposal Category: GTO

INVESTIGATORS

HIVESTIGHTONS		
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OBSERVATIONS

Folder	Observation	Label	Observing Template	Science Target
HUDF :	MIRI Imaging Su	urvey		
	1	Mosaic	MIRI Imaging	(1) XDF-OFFCENTER
Observa	ation Folder			
	2	HUDF-NIRSPEC1	NIRSpec MultiObject Spectroscopy	(5) TARGET-OBSERVATION-3
	3	HUDF-NIRSPEC2	NIRSpec MultiObject Spectroscopy	(6) TARGET-OBSERVATION-4
	4	HUDF-NIRSPEC3	NIRSpec MultiObject Spectroscopy	(7) TARGET-OBSERVATION-5

ABSTRACT

We plan a MIRI multi-band survey in the GOODS-S/HUDF region, covering about 30 square arcmin and using all the MIRI imaging bands. Selected galaxies from this survey will be observed with NIRSpec at R = 1000 and from 1 to 5.2 microns. We will model the photometry to separate AGNs from star forming galaxies. We expect to find 30 - 40 AGN of known types and detected at 10:1 signal to noise or higher at 21 microns. We should be able to identify any previously unknown obscured AGNs, and with the deep X-ray, optical and radio data in the same field should obtain a complete sample of these objects. This same survey will provide high quality measurements of some 2000 star forming galaxies. We will compare star formation rates determined from X-rays, hydrogen recombination lines, UV, and mid-infrared to calibrate these indicators at z = 2. We will also

JWST Proposal 1207 (Created: Tuesday, June 25, 2019 at 8:00:34 AM Eastern Standard Time) - Overview use the spectra to estimate metallicities and study the dependence of the aromatic bands and other properties of the galaxies on this parameter. Together these measurements will let us determine accurate SFR densities, luminosity functions, and other parameters relevant to galaxy evolution.

GRIEKE_4001-4011

OBSERVING DESCRIPTION

This program will be executed in two parts. First, MIRI imaging in 8 bands will be obtained in a 3x5 mosaic of the HUDF/GOODS-S region. The mosaic position was chosen to optimize overlap with the NIRCam GTO HUDF/GOODS-S program. AGN can be identified by their SEDs filling in the minimum in star forming galaxy SEDs near rest 4.5 microns. We are developing techniques based on the relative flux densities in the MIRI and some of the NIRCam bands for rapid separation of AGN from star forming galaxies, whose spectra not only have the aforementioned minimum but also are characterized by strong aromatic emission features. Deriving robust star formation rates from 21 micron photometry of highly luminous galaxies (i.e. ULIRGs) will require: 1.) distinguishing nuclear- concentrated star formation from galaxies where the SF is distributed over the disk, an issue for z < 1 (Rujopakarn et et al. 2013; Shipley et al. 2016); and 2.) determining in which cases the 21 microns output is contaminated by AGN. The mid-IR spectral energy distribution can be used as a proxy for the central concentration of star forming galaxies. JWST photometry can help select the appropriate spectral template through comparing the behavior of the output of the aromatic bands, which are dominant for 6 - 13 microns, with that of the dust grains that dominate the emission at 13 - 30 microns. For example, the MIRI photometric bands at 12.8 microns and 21 microns sample these two spectral components separately for 0 < z < 0.6. Their relative behavior can be expressed by the ratio of the flux densities in the two bands. Using the templates suggested by Rujopakarn et al. (2013), the ratio f(12.8)/f(21) for (low central concentration) over (high central concentration) is a factor of 1.5 or more over this redshift range. This approach can be extended to z = 1 by using the 15 micron and 25.5 micron bands. A relatively short integration in the latter band suffices for this application. AGN contamination can be identified with moderately deep observations in the additional MIRI bands, and the results can be tested with the ancillary datasets (e.g., deep X-ray and radio). Finally, the most luminous and embedded star forming galaxies can be identified from previous ultradeep surveys with Spitzer and Herschel, plus the high resolution (0.35 arcsec) and deep radio observations in this region.

This MIRI imaging, together with the NIRCam GTO imaging, will then be used to select targets for NIRSpec MOS follow-up. Three pointings with NIRSpec will obtain R=1000, 1-5um spectra for selected targets. The target selection will prioritize conditional targets identified with MIRI (color-selected AGN, etc) and other interesting sources relevant to this program (radio sources, proto-cluster members, etc). Accurate positions will be determined through NIRCam imaging. As NIRSpec MOS follow-up in this proposal requires both MIRI and NIRCam pre-imaging, it is equired that the NIRSpec follow-up not be scheduled for at least 60 days following the MIRI or NIRCam imaging, whichever is observed last.

	#	Name	Target Coordinates	Targ. Coord. Corrections	Miscellaneous								
	(1)	XDF-OFFCENTER	RA: 03 32 37.3890 (53.1557875d)										
			Dec: -27 48 0.01 (-27.80000d)										
			Equinox: J2000										
	Comments: The program. Whe Category=Un Description=	nen this information is available, identified	n the XDF and the proposed location of NIRCam GTO i this pointing center and PA will be re-assessed.	maging (MRIEKE_0001-0068). Exact positioning wi	ll rely on the launch date and final PA of the NIRCam GTO								
	(2)	NIRSPEC1	RA: 03 32 40.3000 (53.1679167d)										
			Dec: -27 46 17.10 (-27.77142d)										
			Equinox: J2000										
		sponding observation/MSA setup		rily target sources selected via NIRCam and MIRI pre	e-imaging. When this data is available, this pointing center								
		Infrared sources, Radio sources,	X-ray sources]										
	(3)	NIRSPEC2	RA: 03 32 33.6001 (53.1400004d)										
			Dec: -27 47 23.00 (-27.78972d)										
			Equinox: J2000										
ets	and the corres Category=Un	sponding observation/MSA setup	will be re-assessed.	rily target sources selected via NIRCam and MIRI pre	e-imaging. When this data is available, this pointing center								
ľg	(4)	NIRSPEC3	RA: 03 32 39.1000 (53.1629167d)										
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	(5)	Infrared sources, Radio sources,	RA: 03 32 36.3463 (53.1514429d)										
	(5)	TARGET-OBSERVATION-3	,										
			Dec: -27 46 39.39 (-27.77761d)										
	C T		Equinox: J2000										
	Category=Un	nis target was generated automat identified	ically for MSA Observation 3										
	Description=[Infrared sources, Radio sources,	Visible sources, X-ray sources]										
	(6)	TARGET-OBSERVATION-4	RA: 03 32 36.3463 (53.1514429d)										
			Dec: -27 46 39.39 (-27.77761d)										
			Equinox: J2000										
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			7 - Observation	- MIRI i	n the	Hubble	Ultra-Deep	Field
ſ	n	Proposal 1207,	Observation 1: Mosaic					

Tue Jun 25 13:00:34 GMT 2019

Comments: This mosaic is roughly centered on the XDF and the proposed location of NIRCam GTO imaging (MRIEKE_0001-0068). Exact positioning will rely on the launch date and final PA of the NIRCam GTO program. When this information is available, this pointing center and PA will be re-assessed.

(Visit 1:1) Warning (Form): Overheads are provisional until the Visit Planner has been run.

(Visit 1:2) Warning (Form): Overheads are provisional until the Visit Planner has been run.

(Visit 1:3) Warning (Form): Overheads are provisional until the Visit Planner has been run.

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(Visit 1:14) Warning (Form): Overheads are provisional until the Visit Planner has been run.

(Visit 1:15) Warning (Form): Overheads are provisional until the Visit Planner has been run.

'n	# Name	Target Coordinates	Targ. Coord. Corrections	Miscellaneous

Fixed Targets XDF-OFFCENTER RA: 03 32 37.3890 (53.1557875d) Dec: -27 48 0.01 (-27.80000d)

Equinox: J2000

Comments: This pointing is roughly centered on the XDF and the proposed location of NIRCam GTO imaging (MRIEKE_0001-0068). Exact positioning will rely on the launch date and final PA of the NIRCam GTO program. When this information is available, this pointing center and PA will be re-assessed.

Category=Unidentified Description=[Blank field]

Subarray

Template **FULL**

Observatic

Diagnostics

Diagnostic Status: Warning Observing Template: MIRI Imaging

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=	5	2	4-Point-Sets	1	8		5	2	POINT SOURCE	POSITIVE	MEDIUM

Proposal 1207 - Observation 1 - MIRI in the Hubble Ultra-Deep Field

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) 	1	F560W	FAST	59	1	1	Dither 1	4	4	654.909	
	Ĕ	2	F770W	FAST	78	1	1	Dither 1	4	4	865.812	
ľ	: I	3	F1000W	FAST	58	1	1	Dither 1	4	4	643.809	
	- -	4	F1280W	FAST	68	1	1	Dither 1	4	4	754.811	
	5	5	F1500W	FAST	101	1	1	Dither 1	4	4	1121.116	
	je l	6	F1800W	FAST	68	1	1	Dither 1	4	4	754.811	
Ġ	င္က	7	F2100W	FAST	32	3	1	Dither 2	8	24	2131.231	
		8	F2550W	FAST	18	4	1	Dither 1	4	16	799.212	

Sequence Visits within 53.0 Days
Aperture PA Range 41 to 41 Degrees (V3 30 Visits Same PA

2 After 1 by 60 Days to <None specified>
3 After 1 by 60 Days to <None specified>
4 After 1 by 60 Days to <None specified> Sequence Visits within 53.0 Days Aperture PA Range 41 to 41 Degrees (V3 36.550295 to 36.550295) Visits Same PA

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l e	1		G140M/F100LP	Configuration: c0	NRSIRS2	32	1	NONE	3	3	7046.434	
	2		G235M/F170LP	Configuration: c0	NRSIRS2	32	1	NONE	3	3	7046.434	
tral	3		G395M/F290LP	Configuration: c0	NRSIRS2	16	1	NONE	3	3	3545.1	
Spectral Eleme												

Proposal 1207 - Observation 2 - MIRI in the Hubble Ultra-Deep Field

On Hold contingent on NIRCam imaging (MRIEKE_0001-0068) and MIRI imaging (this proposal).

MSA Planned Aperture PA 180.0 to 180.0 Degrees (V3 41.50766 to 41.50766)

2 After 1 by 60 Days to <None specified>
Sequence Observations 2, 3, 4, Non-interruptible

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1 🖁	2		G235M/F170LP	Configuration: c0		32	1	NONE	3	3	7046.434	
ā	3		G395M/F290LP	Configuration: c0	NRSIRS2	16	1	NONE	3	3	3545.1	
Spectral Eleme												

Proposal 1207 - Observation 3 - MIRI in the Hubble Ultra-Deep Field

On Hold contingent on NIRCam imaging (MRIEKE_0001-0068) and MIRI imaging (this proposal).

MSA Planned Aperture PA 180.0 to 180.0 Degrees (V3 41.50766 to 41.50766)

3 After 1 by 60 Days to <None specified>
Sequence Observations 2, 3, 4, Non-interruptible

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_	_		Observation 4: HUD		Tille Hubble	Ollia-Deep	r ieid				Tue Iun 25	13:00:34 GMT 2019
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Observation	1		observation is a plac		Pec pointing with thi	s exposure set-up.	The final NIRSpec p	ointing and MSA sli	it configuration is de	pendent on conditio	onal targets identifie	d through MIRI and
Diagnostics	•	(Visit 4:1) Warnii	ng (Form): Overhead	ls are provisional un	til the Visit Planner	has been run.						
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Fixed Targets		Comments: This t Category=Unider	arget was generated	automatically for M	SA Observation 5							
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fion	#	#	Reference Star Bin	Target	Filter	MSA Configuration	Readout Pattern	Groups/Int	Integrations/Exp	Total Integrations	Total Exposure Time	ETC Wkbk.Calc ID
Acquisition	-	1		SAME	F140X	Auto Acq MSA Config	NRS	3	1	4	558.312	
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۱Ř	1	1	G140M/F100LP	Configuration: c0	NRSIRS2	32	1	NONE	3	3	7046.434	
	2	2	G235M/F170LP	Configuration: c0		32	1	NONE	3	3	7046.434	
ā	3	3	G395M/F290LP	Configuration: c0	NRSIRS2	16	1	NONE	3	3	3545.1	
Spectral Elements	-											

Proposal 1207 - Observation 4 - MIRI in the Hubble Ultra-Deep Field

On Hold contingent on NIRCam imaging (MRIEKE_0001-0068) and MIRI imaging (this proposal).

MSA Planned Aperture PA 180.0 to 180.0 Degrees (V3 41.50766 to 41.50766)

4 After 1 by 60 Days to <None specified>
Sequence Observations 2, 3, 4, Non-interruptible