

1263 - NIRSpec and MIRI Spectroscopy of QSOs - Part 2

Cycle: 1, Proposal Category: GTO

INVESTIGATORS

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OBSERVATIONS

CECE	11110110					
Folder	Observation	Label	Observing Template	Science Target		
MIRI	J1120					
	1	J1120 [WRIGHT_0501	MIRI Medium Resolution Spectroscopy	(1) MIRI-J1120		
	2	J1120-Imager [WRIGH T_0502]	MIRI Imaging	(2) MIRI-J1120-IMAGER		
NIRSPI	EC - J1120					
	3	NIRSpec IFU observati on of J1120 [FERRUIT _3053]	NIRSpec IFU Spectroscopy	(3) NIRSPEC-J1120+0641		

ABSTRACT

This APT is for IFU Observations of the high-z QSO J1120 which is in common with the MIRI programme, hence the programmes are merged to save slew.

JWST Proposal 1263 (Created: Tuesday, February 20, 2018 7:04:39 PM EST) - Overview

The goals of both observations are to map the optical and near-IR nebular lines in the host galaxy circumgalactic region of this quasar at z=7.08.

MIRI, with its spectral coverage from 5 to 28 m and sensitivity, is the only instrument onboard JWST able to explore the optical and near-infrared spectrum and light distribution of galaxies and QSOs at redshifts above 6.7. A complete 5 to 28 spectrum (~0.6 to 3.5 microns rest-frame) of the currently known highest redshift (z of 7.0842) QSO J1120+0641 will be obtained together with MIRI imaging to map the first rest-frame near-IR light distribution of the host galaxy at ~0.2-0.3 arcsec angular resolution.

NIRISpec will observe this quasar with the IFS and with R2700 - band III grating, with the goal of tracing the distribution and kinematics of the main nebular emission lines and, in

particular, [OII], Hbeta, [OIII], [OI], Halpha. The ultimate goal is to trace the presence and properties of a quasar driven outflow and the dynamics and star formration of the host galaxy and

its close environment. This information will be precious to constrain the evolutionary processes of early massive galaxies hosting supermassive accreting black holes.

OBSERVING DESCRIPTION

MIRI MRS and Imager:

The purpose of the program is to get a full 5 - 30 um spectrum of J1120 using the 3 MRS configurations with simultaneous Imager observations (see additional note 1). In addition, we request the imaging of the target in two filters: F560W, F770W. To save slew and maneuvering overheads, we propose all observations to be "non-interruptible". The epoch (January 2020) was selected to guarantee low background conditions for MIRI.

The dithering strategies (4-pt, point source) were selected to optimize the PSF and detector effects in all MRS channels and Imager filters. These strategies could be subject to change without modifying the total charged time.

Constraints:

MIRI: the target should be observed in January to have low background.

JWST Proposal 1263 (Created: Tuesday, February 20, 2018 7:04:39 PM EST) - Overview

NIRSpec: due to the bright stars in the field and MSA contamination issues, the V3 range between 55 to 130 should be avoided.

NIRSpec IFU OBSERVATION:

This part corresponds to NIRSpec IFU Proposal ID: FERRUIT_3053

(NIRSpec contact person: Roberto Maiolino)

Proposal title "NIRSpec and MIRI spectroscopy of QSOs - part #2"

The NIRSpec IFU observation is done with the G395H grating and it is aimed aimed primarily at mapping the strongest optical nebular lines (Hbeta, [OIII], Halpha, [NII]).

PA constraints are driven by the MIRI observation and by the requirement to avoid some bright stars to be in the NIRSpec MSA footprint.

In this specific case, if the target is positioned in the center of the IFU field of view then at this specific redshift the brightest [OIII]5007 is located in the detectors gap over about half of the field of view. Hence, we have offset the centering by -0.7 arcsec in the X-direction (we hope this is in the instrument coordinates, not on sky), which will enable to properly map [OIII]5007 in the central +-0.7 arcsec of the source.

We have adopted a "small" dither pattern, given the reduced effective field as a consequence of the offset discussed above.

We are using no target acquisition (i.e. point-and-shoot).

JWST Proposal 1263 (Created: Tuesday, February 20, 2018 7:04:39 PM EST) - Overview At any of the constrained PA range there are Gaia GS that can be selected for guiding and which will ensure the proper location of the target within the IFU aperture, with the required accuracy.

We are using NRSIRS2RAPID for a better identification and rejection of cosmic rays.

Pro	posal 126	3 - Obser	vation 1 -	NIRSpec	and MIRI S	Spectrosc	opy of QSO	s - Part 2					
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