Force on a spherical particle in an optical conveyor

David Ruffner and David G. Grier

Department of Physics and Center for Soft Matter Research, New York University, New York, NY 10003

(Dated: March 5, 2014)

Derivation of the calculation of the force on a spherical particle in an optical conveyor using Lorenz-Mie theory.

In order to calculate the force we must write down the incident field in terms of the vector spherical harmonics. This allows us to calculate the scattered field by matching the boundary conditions at the scattering sphere [1]. Furthermore given the scattered field we can calculate the force and torque on the spherical particle [2].

The field of an optical conveyor is a sum of two Bessel beams, and each Bessel beam can be thought of as a cone of plan waves [8]. The electric field of the Bessel beam as a function of position, **r**, can be written as,

$$\mathbf{E}(\mathbf{r}) = E_0 \int_0^{2\pi} \mathbf{e}_0(\theta_0, \phi') e^{i\mathbf{k}(\theta_0, \phi') \cdot \mathbf{r}} d\phi', \qquad (1)$$

Where E_0 is the amplitude of each component plane wave, $\mathbf{e}_0(\theta_0, \phi')$ is the polarization, and $\mathbf{k}(\theta_0, \phi')$ is the wavevector. The cone angle, θ_0 , gives us the wavevector of each component plane wave,

$$\mathbf{k}(\theta_0, \phi') = k\hat{r}(\theta_0, \phi') \tag{2}$$

$$= k(\sin \theta_0 \cos \phi' \hat{x} + \sin \theta_0 \sin \phi' \hat{y} + \cos \theta_0) \quad (3)$$

in terms of the wavenumber, $k = 2\pi n_{ext}/\lambda$, where n_{ext} is the refractive index of the medium and λ is the wavelength of the light. If we assume that the Bessel beam was formed from light polarized in the \hat{x} direction, then we can get the polarization of each plane wave component decomposing it in terms of $\hat{\theta}$ and $\hat{\phi}'$ since these are the two directions transverse to $\mathbf{k}(\theta_0, \phi')$. Consequently the polarization of each plane wave component becomes,

$$\mathbf{e}_0(\theta_0, \phi') = \cos \phi' \,\hat{\theta}(\theta_0, \phi') + \sin \phi' \,\hat{\phi}'(\theta_0, \phi'). \tag{4}$$

The electric field can always expressed in terms of the vector spherical harmonics $\mathbf{M}_{mn}^{(1)}$ and $\mathbf{N}_{mn}^{(1)}$ [3], with the following expansion,

$$\mathbf{E} = -i\sqrt{\frac{2n+1}{4\pi} \frac{(n-m)!}{(n+m)!}} \sum_{n=1}^{\infty} \sum_{m=-n}^{n} \left(p_{mn} \mathbf{N}_{mn}^{(1)} + q_{mn} \mathbf{M}_{mn}^{(1)} \right)$$

Where we are neglecting the position dependence for clarity. The coefficients p_{mn} and q_{mn} are called the beam shape coefficients (BSCs) and they can be determined by the following integral,

$$p_{mn} = i\sqrt{\frac{4\pi}{2n+1}} \frac{(n+m)!}{(n-m)!} \frac{\int_0^{2\pi} \int_0^{\pi} \mathbf{E} \cdot \mathbf{N}_{mn}^* \sin\theta d\theta d\phi'}{\int_0^{2\pi} \int_0^{\pi} |\mathbf{N}_{mn}|^2 \sin\theta d\theta d\phi'},$$
(6)

And a similar one for q_{mn} with \mathbf{M}_{mn} instead of \mathbf{N}_{mn} . This integral is difficult in general however Taylor and Love solved it analytically in the case of a Bessel beam [7]. Essentially they wrote the Bessel beam in terms of a cone of plane waves like we have above. The found the BSCs for each plane wave as was shown in [5], and then added them all together and were able to compute the integral. This enables us to write down the BSCs for a Bessel beam,

$$p_{mn}\left(\mathbf{r}\right) = E_0 U_n e^{ikz\cos\theta_0} \tag{7}$$

$$\left[\tilde{\tau}_{mn}(\cos\theta_0)\,I^+(\rho,\phi) + \right. \tag{8}$$

$$\tilde{\pi}_{mn}(\cos\theta_0) I^-(\rho,\phi)$$
 (9)

where $\rho = \sqrt{x^2 + y^2}$ is the distance from the beam axis, and $\phi = \arctan(-y/x) - \pi/2$ is the azimuthal angle. To calculate q_{mn} just exchange the $\tilde{\tau}_{mn}$ with the $\tilde{\pi}_{mn}$. There is the factor $U_n = \frac{4\pi i^n}{n(n+1)}$ and then functions in terms of the modified Legendre Polynomials, P_n^m ,

$$\tilde{\pi}_{mn}(\cos \theta_0) = \sqrt{\frac{2n+1}{4\pi} \frac{(n-m)!}{(n+m)!}} \frac{m}{\sin \theta_0} P_n^m(\cos \theta_0),$$
(10)

$$\tilde{\tau}_{mn}(\cos\theta_0) = \sqrt{\frac{2n+1}{4\pi} \frac{(n-m)!}{(n+m)!}} \frac{d}{d\theta_0} P_n^m(\cos\theta_0).$$
 (11)

Finally there are the results from solving the integrals,

$$I^{\pm}(\rho,\phi) = \pi \left(e^{i(m-1)\phi} J_{1-m}(k\sin\theta_0 \rho) \pm \right)$$
 (12)

$$e^{i(m+1)\phi}J_{-1-m}(k\sin\theta_0\rho)\right). \tag{13}$$

The BSCs for the Bessel beam are still complicated, however this form is much more useful than the integral in Eq. (6). All one has to do to calculate p_{mn} is to evaluate Bessel functions and associated Legendre polynomials.

The next step is to determine the scattered field from the BSCs. The BSCs depend on position because they refer to a spherical polar coordinate system centered on the particle. This allows us to easily match the boundary conditions which determines the coefficients of the scattered field [1]. One notational issue is Barton uses a different convention for the BSCs but they are related according to,

$$A_{mn} = \frac{ip_{mn}}{2\pi ka}, \quad B_{mn} = \frac{n_{ext}q_{mn}}{2\pi ka}.$$
 (14)

The coefficients of the scattered field are,

$$a_{mn} = \frac{\psi'_{n}(k_{int}a)\psi_{n}(k_{ext}a) - \bar{n}\psi_{n}(k_{int}a)\psi'_{n}(k_{ext}a)}{\bar{n}\psi_{n}(k_{int}a)\xi_{n}^{(1)'}(k_{ext}a) - \psi'_{n}(k_{int}a)\xi_{n}^{(1)}(k_{ext}a)} A_{mn},$$
(15)

$$b_{mn} = \frac{\bar{n}\psi'_{n}(k_{int}a)\psi(k_{ext}a) - \psi_{n}(k_{int}a)\psi'_{n}(k_{ext}a)}{\psi_{n}(k_{int}a)\xi_{n}^{(1)'}(k_{ext}a) - \bar{n}\psi'_{n}(k_{int}a)\xi_{n}^{(1)}(k_{ext}a)} B_{mn},$$
(16)

Where $\xi_n^{(1)} = \psi_n - i\chi_n$ and ψ_n and χ_n are the Ricatti-Bessel functions. Additionally, k_{int} is the wavenumber inside the sphere, k_{ext} is the wavenumber outside the sphere, and a is the radius of the spherical particle. Finally the complex index of refraction is given by $\bar{n} = \sqrt{\bar{\epsilon}_{int}/\epsilon_{ext}}$ where $\bar{\epsilon}_{int} = \epsilon_{int} + i4\pi\sigma/\omega$ and σ is the electrical conductivity of the particle. This can be simplified using the a_n and b_n in Bohren and Huffman,

$$a_{mn} = -a_n A_{mn}, (17)$$

$$b_{mn} = -b_n B_{mn}, (18)$$

(CHECK).

Given these BSCs for the incident and the scattered fields we can now calculate the force on a spherical particle using the formalism derived by Barton et. al. [2]. Integrating the Maxwell's stress tensor, **T**, over a surface enclosing the particle determines the optical force,

$$\langle \mathbf{F}(\mathbf{r}) \rangle = \langle \oint_{S} \hat{n} \cdot \mathbf{T} dS \rangle.$$
 (19)

Barton substitutes the series expression for the fields, Eq. (5), into the surface integral and then evaluates in the limit large radius. It's possible then to directly integrate and express the force on the particle in terms of the BSCs of the incident, A_{mn} and B_{mn} and scattered fields, a_{mn} and b_{mn} . The expression is huge and nasty:

$$\frac{\langle F_x \rangle + i \langle F_y \rangle}{a^2 E_0^2} = \frac{\alpha^2}{16\pi} i \sum_{n=1}^{\infty} \sum_{m=-n}^{n} \left[\sqrt{\frac{(n+m+2)(n+m+1)}{(2n+1)(2n+3)}} n(n+2) \times \right]$$
(20)

$$(2\epsilon_{ext}a_{mn}a_{m+1,n+1}^* + \epsilon_{ext}a_{mn}A_{m+1,n+1}^* + \epsilon_{ext}A_{mn}a_{m+1,n+1}^* + 2b_{mn}b_{m+1,n+1}^* + b_{mn}B_{m+1,n+1}^* + (21)$$

$$B_{mn}b_{m+1,n+1}^{*}) + \sqrt{\frac{(n-m+1)(n-m+2)}{(2n+1)(2n+3)}}n(n+2)(2\epsilon_{ext}a_{m-1,n+1}a_{mn}^{*} + \epsilon_{ext}a_{m-1,n+1}A_{mn}^{*} + (2n+1)(2n+3))$$
(22)

$$\epsilon_{ext} A_{m-1,n+1} a_{mn}^* + 2b_{m-1,n+1} b_{mn}^* + b_{m-1,n+1} B_{mn}^* + B_{m-1,n+1} b_{mn}^*$$
(23)

$$-\sqrt{(n+m+1)(n-m)}\sqrt{\epsilon_{ext}}(-2a_{mn}b_{n,m+1}^* + 2b_{mn}a_{n,m+1}^* + -a_{mn}B_{n,m+1}^* +$$
(24)

$$b_{mn}A_{n,m+1}^* + B_{mn}a_{n,m+1}^* + -A_{mn}b_{n,m+1}^*)$$
(25)

and,

$$\frac{\langle F_z \rangle}{a^2 E_0^2} = -\frac{\alpha^2}{8\pi} \sum_{n=1}^{\infty} \sum_{m=-n}^{n} \Im\left\{ \sqrt{\frac{(n-m+1)(n+m+1)}{(2n+1)(2n+3)}} n(n+2) \times \right.$$
 (26)

$$(2\epsilon_{ext}a_{m,n+1}a_{mn}^* + \epsilon_{ext}a_{m,n+1}A_{mn}^* + \epsilon_{ext}A_{m,n+1}a_{mn}^* + 2b_{m,n+1}b_{mn}^* + b_{m,n+1}B_{mn}^* + (27)$$

$$B_{m,n+1}b_{mn}^*) + \sqrt{\epsilon_{ext}}m(2a_{mn}b_{mn}^* + a_{mn}B_{mn}^* + A_{mn}b_{mn}^*)\}.$$
(28)

There are also similar expressions in Farsund and Felder-hof [4] and in Mazolli et. al. [6]. So we should be able to catch any mistakes in these equations.

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