

Understanding the vector spherical harmonics

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Working out the form of the vector spherical harmonics.

The vector spherical harmonics can be used to decompose a vector field in spherical polar coordinates. When the field is in this form we can more easily calculate the scattering off of spherical particles. In fact each component of the incident field contributes to the scattered field multiplied by the corresponding a_n and b_n of Lorentz-Mie theory [1]. Each textbook uses its own notation so it is hard to compare between the two. Here we'll go through the calculation to compare Jackson [3] to Gouesbet [2].

Jackson begins with the scalar wave equation,

$$\nabla^2 \psi - \frac{1}{c^2} \frac{\partial^2 \psi}{\partial t^2} = 0. \quad (1)$$

Solutions can be written as an expansion in multipole terms,

$$\psi(\mathbf{r}) = \sum_{l,m} [A_{lm} z_l(kr) + B_{lm} z_l(kr)] Y_{lm}(\theta, \phi). \quad (2)$$

We need to make a series of definitions. Y_{lm} are the spherical harmonics, defined by,

$$Y_{lm}(\theta, \phi) = \tilde{P}_l^m(\cos \theta) e^{im\theta} \quad (3)$$

Where $\tilde{P}_l^m(\cos \theta)$ are the normalized associated Legendre polynomials (which are calculated by `dbr_plegendre.pro`), which are defined by,

$$\tilde{P}_l^m(x) = N_p (-1)^m (1-x^2)^{m/2} \frac{d^m}{dx^m} P_l(x) \quad (4)$$

Where $P_l(x)$ are the Legendre Polynomials defined by Rodrigues' formula and the normalization is $N_p = \sqrt{\frac{(2l+1)(l-m)!}{4\pi(l+m)!}}$. And $z_l(kr)$ is one of the spherical Bessel functions, for instance

$$j_l(x) = \sqrt{\frac{\pi}{2x}} J_{l+1/2}(x), \quad (5)$$

Where $J_l(x)$ is a Bessel function of the first kind of order l .

Next Jackson points out that Maxwell's equations can be written in the form of the Helmholtz equation and a divergence condition,

$$(\nabla^2 + k^2)\mathbf{E} = 0 \quad \text{and} \quad \nabla \cdot \mathbf{E} = 0 \quad (6)$$

with \mathbf{H} given by

$$\mathbf{H} = \frac{-i}{kZ_0} \nabla \times \mathbf{E} \quad (7)$$

Where $Z_0 = \frac{1}{c\epsilon_0}$ is the impedance of free space. The hard part is enforcing the divergence condition. We can instead show that $\mathbf{r} \cdot \mathbf{E}$ and $\mathbf{r} \cdot \mathbf{H}$ each satisfy the scalar Helmholtz equation by using the vector identity,

$$\nabla^2(\mathbf{r} \cdot \mathbf{A}) = \mathbf{r} \cdot \nabla^2 \mathbf{A} + 2\nabla \cdot \mathbf{A}. \quad (8)$$

Consequently we can expand $\mathbf{r} \cdot \mathbf{E}$ and $\mathbf{r} \cdot \mathbf{H}$ in terms of Eq. (2) which is why you need the two families of coefficients. It's interesting that you only need two scalar fields to construct a vector field, but this comes from the transverse nature of the fields.

To extract the electric and magnetic fields from the scalar fields we need to use the angular momentum operator, \mathbf{L} , defined by

$$\mathbf{L} = -i\mathbf{r} \times \nabla. \quad (9)$$

Jackson shows that if,

$$\mathbf{r} \cdot \mathbf{H}_{lm}^{(M)} = \frac{l(l+1)}{k} z_l(kr) Y_{lm}(\theta, \phi) \quad (10)$$

Then the electric field corresponding to this magnetic multipole of order l, m is,

$$\mathbf{E}_{lm}^{(M)} = Z_0 z_l(kr) \mathbf{L} Y_{lm}(\theta, \phi). \quad (11)$$

Similarly for an electric multipole field,

$$\mathbf{E}_{lm}^{(E)} = \frac{iZ_0}{k} \nabla \times (z_l(kr) \mathbf{L} Y_{lm}(\theta, \phi)). \quad (12)$$

These two sets of functions are the vector spherical harmonics which are orthogonal functions which allow us to write a general vector field as a sum of them,

$$\mathbf{E} = \sum_{l,m} \left[\frac{i}{k} a_E(l, m) \nabla \times (z_l(kr) \mathbf{L} Y_{lm}(\theta, \phi)) \right. \quad (13)$$

$$\left. + a_M(l, m) \mathbf{L} Y_{lm}(\theta, \phi) \right]. \quad (14)$$

Here we depart slightly from Jackson who defines his vector spherical harmonic function \mathbf{X}_{lm} with a factor of $1/\sqrt{l(l+1)}$ which we omit to match with other formulations.

We can write these functions in spherical coordinates to enable comparison with other works. Considering the magnetic term first,

$$j_l(kr) \mathbf{L} Y_{lm}(\theta, \phi) = -i j_l(kr) \mathbf{r} \times \nabla Y_{lm}(\theta, \phi) \quad (15)$$

$$= -\hat{\phi} i \frac{\psi_l(kr)}{kr} \partial_\theta Y_{lm}(\theta, \phi) \quad (16)$$

$$- \hat{\theta} \frac{\psi_l(kr)}{kr} \frac{m Y_{lm}(\theta, \phi)}{\sin \theta}, \quad (17)$$

Where $\psi_l(kr) = krz_l(kr)$. Next we evaluate the electric term,

$$\frac{-i}{k} \nabla \times j_l(kr) \mathbf{L} Y_{lm}(\theta, \phi) = \frac{-i}{k} \nabla \times \mathbf{L}(j_l(kr) Y_{lm}(\theta, \phi)) \quad (18)$$

$$= \hat{r} \frac{l(l+1)}{(kr)^2} \psi_l(kr) Y_{lm}(\theta, \phi) \quad (19)$$

$$+ \hat{\theta} \frac{1}{kr} \psi'_l(kr) \partial_\theta Y_{lm}(\theta, \phi) \quad (20)$$

$$+ \hat{\phi} \frac{im}{kr} \psi'_l(kr) \frac{Y_{lm}(\theta, \phi)}{\sin \theta}. \quad (21)$$

There are a number of steps in evaluating this term. We used the fact that the angular momentum operator only operates on the angular variables also we used (9.125) in Jackson and Eq.(8) in Gouesbet [2]. This gives the same result as Eq.(1) and Eq.(2) in Gouesbet for the magnetic and the electric VSHs respectively, except for a factor of $(-1)^m$.

For a circularly polarized plane wave travelling along the z axis the coefficients are,

$$a_M(l, m)_\pm = i^l \sqrt{\frac{4\pi(2l+1)}{l(l+1)}} \delta_{m, \pm 1} \quad (22)$$

$$a_E(l, m)_\pm = \pm i a_M(l, m)_\pm \quad (23)$$

These are different than the ones in Bohren and Huffman by the square root.

Numerically it's more stable to write the field in terms of the normalized angular functions $\tilde{\pi}_{mn}$ and $\tilde{\tau}_{mn}$, which are defined by,

$$\tilde{\pi}_{mn}(\cos \theta) = m \frac{\tilde{P}_n^m(\cos \theta)}{\sin \theta} \quad (24)$$

and

$$\tilde{\tau}_{mn}(\cos \theta) = \partial_\theta \tilde{P}_n^m(\cos \theta). \quad (25)$$

This gives us the vector spherical harmonics,

$$j_l(kr) \mathbf{L} Y_{lm}(\theta, \phi) = -\hat{\phi} i \frac{\psi_l(kr)}{kr} \tilde{\tau}_{mn}(\cos \theta) e^{im\phi} \quad (26)$$

$$-\hat{\theta} \frac{\psi_l(kr)}{kr} \tilde{\pi}_{mn}(\cos \theta) e^{im\phi}, \quad (27)$$

and,

$$\frac{-i}{k} \nabla \times j_l(kr) \mathbf{L} Y_{lm}(\theta, \phi) = \hat{r} \frac{l(l+1)}{(kr)^2} \psi_l(kr) Y_{lm}(\theta, \phi) \quad (28)$$

$$+ \hat{\theta} \frac{1}{kr} \psi'_l(kr) \tilde{\tau}_{mn}(\cos \theta) e^{im\phi} \quad (29)$$

$$+ \hat{\phi} \frac{i}{kr} \psi'_l(kr) \tilde{\pi}_{mn}(\cos \theta) e^{im\phi}. \quad (30)$$

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- [1] Bohren, C. F. and D. R. Huffman (2008, September). *Absorption and Scattering of Light by Small Particles*. John Wiley & Sons.
 - [2] Gouesbet, G. (2010, February). T-matrix formulation and generalized Lorenz-Mie theories in spherical coordinates. *Optics Communications* 283(4), 517–521.
 - [3] Jackson, J. D. (1998, August). *Classical Electrodynamics Third Edition* (3 ed.). Wiley.