

Newton 2:

$\vec{P}$ : momentum (total quantity of motion)

$\vec{F}$ : force

If object has mass  $m$  and velocity  $\vec{v}$

$$\vec{p} = m\vec{v} = m\vec{r}'$$

The rate of change of momentum is force.

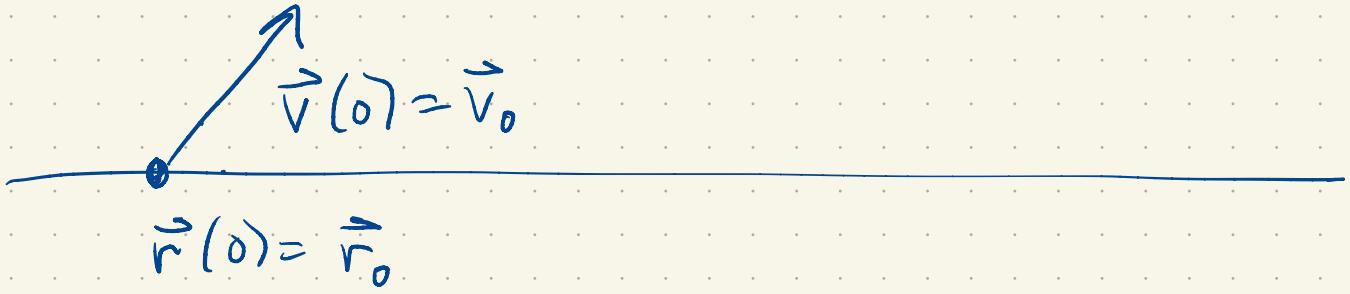
$$\frac{d}{dt} \vec{p} = \vec{F}$$

$$m \text{ constant: } m\vec{r}'' = \vec{F} \quad (\vec{F} = m\vec{a})$$

If you know the force acting on an object,  
you know the acceleration.

$$\vec{a} = \frac{1}{m} \vec{F}$$

And if you know initial position and velocity  
then you can reconstruct the position.



Projectiles close to earth:

$$\vec{F}_g = -9.8 \hat{k} \text{ m/s}^2$$

$$\vec{v}(t) = \int -9.8 \hat{k} dt + \vec{C}_1$$

$$= -9.8t \hat{k} + \vec{C}_1$$

$$\vec{v}(0) = \vec{C}_1$$

$$\vec{v}(t) = -9.8t \hat{k} + \vec{v}_0$$

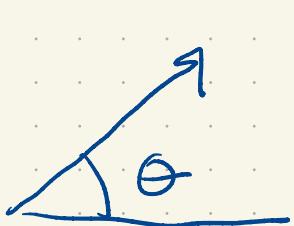
$$\vec{r}(t) = \int \vec{v}(t) dt + \vec{C}_2$$

$$\vec{r}(t) = -\frac{9.8 t^2}{2} \hat{k} + \vec{v}_0 t + \vec{C}_2$$

$$\vec{r}(0) = 0 + 0 + \vec{C}_2$$

$$\vec{r}(t) = \vec{r}_0 + \vec{v}_0(t) - \frac{9.8}{2} t^2 \hat{k}$$

$(g, \theta \rightarrow 0 \Rightarrow \text{linear motion!})$



$$\vec{r}_0 = \vec{0}$$

$$\vec{v}_0 = v_0 \cos(\theta) \hat{i} + v_0 \sin(\theta) \hat{k}$$

$$\vec{r}(t) = v_0 \cos \theta t \hat{i} + \left[ v_0 \sin \theta t - \frac{9.8}{2} t^2 \right] \hat{k}$$

$$x = ct \quad t = x/c$$

$$\alpha = \tan \theta$$

$$z = \alpha x - \beta x^2$$

$$\beta = \frac{9.8}{2} \frac{\sec^2 \theta}{v_0^2}$$

This is a parabolic trajectory.

When is  $z = 0$ ?  $t \left[ v_0 \sin \theta - \frac{9.8}{2} t \right] = 0$

$t = 0$  or

$$t = \frac{2v_0 \sin \theta}{9.8}$$

e.g.  $v_0 = 150 \text{ m/s}$   $\theta = \pi/4 = 45^\circ$

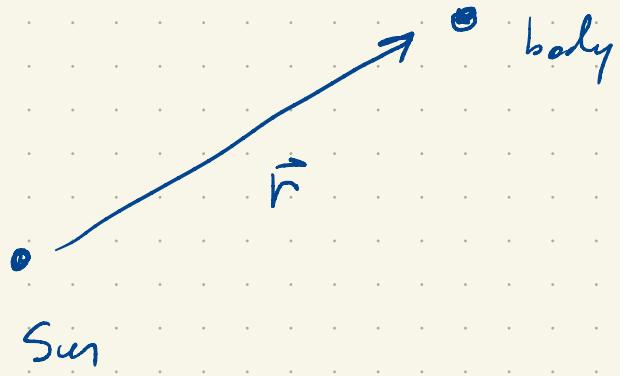
How far when strikes ground?

$$t = \frac{300}{9.8} \cdot \frac{1}{\sqrt{2}} \approx 21.64$$

$$x = v_0 \cos \theta t$$

$$= 150 \cdot \frac{1}{\sqrt{2}} \cdot \frac{300}{9.8} \cdot \frac{1}{\sqrt{2}} = \frac{150 \cdot 150}{9.8} \approx 2295 \text{ m}$$

which is nearly constant.



$$F_G = -\frac{GM_S m_b}{|\vec{r}|^2} \frac{\vec{r}}{|\vec{r}|}$$

Acceleration of body:

$$\frac{d}{dt} \left( m_b \vec{r}' \right) = -\frac{GM_S \vec{r}}{|\vec{r}|^3}$$

$$\vec{r}'' = -\frac{GM_S \vec{r}}{|\vec{r}|^3} \quad M_S \rightarrow M$$

$$\vec{p} = m_b \vec{r}' \rightarrow \text{"linear momentum"}$$

$$\vec{L} = \vec{r} \times \vec{p} \quad \text{angular momentum about } \vec{O}.$$

"total amount of angular motion!"

$\frac{d}{dt} \vec{p} = \vec{F}_g \neq 0$ . The body's momentum is not preserved.

But

$$\frac{d}{dt} \vec{L} = \frac{d}{dt} \vec{r} \times \vec{p} = \vec{r}' \times \vec{p} + \vec{r} \times \frac{d}{dt} \vec{p}$$

$$= \vec{r}' \times (m_b \vec{r}') + \vec{r} \times \vec{F}_g$$

$$= 0 + 0$$

$$(\vec{r}' \times \vec{r}'), (\vec{r} \times \vec{r})$$

Angular momentum ( $\vec{L}$ ) about  $\vec{O}$  is constant.

$\oint \vec{(\vec{r} - \vec{a})} \times \vec{p} = -\vec{a} \times \vec{F_g} \neq 0$  in general,

Note  $\vec{r} \perp \vec{L}$  so  $\vec{r}$  is always contained  
in the plane  $\uparrow$  perp to  $\vec{L}$ .  
thru origin

$$\vec{L} = L \hat{k} \quad L > 0 \quad \text{WLOG}$$

$$\vec{r} = r \cos\theta \hat{i} + r \sin\theta \hat{j}$$

$$\vec{r}' = r' \frac{\vec{r}}{r} + r (-\sin\theta \hat{i} + \cos\theta \hat{j}) \theta'$$

$$\vec{r} \times \vec{r}' = \vec{r} \times r (-\sin\theta \hat{i} + \cos\theta \hat{j}) \theta'$$

$$= r^2 (c \hat{i} + s \hat{j})(-s \hat{i} + c \hat{j}) \theta'$$

$$= r^2 (c^2 \hat{i} \times \hat{j} - s^2 \hat{j} \times \hat{i}) \theta'$$

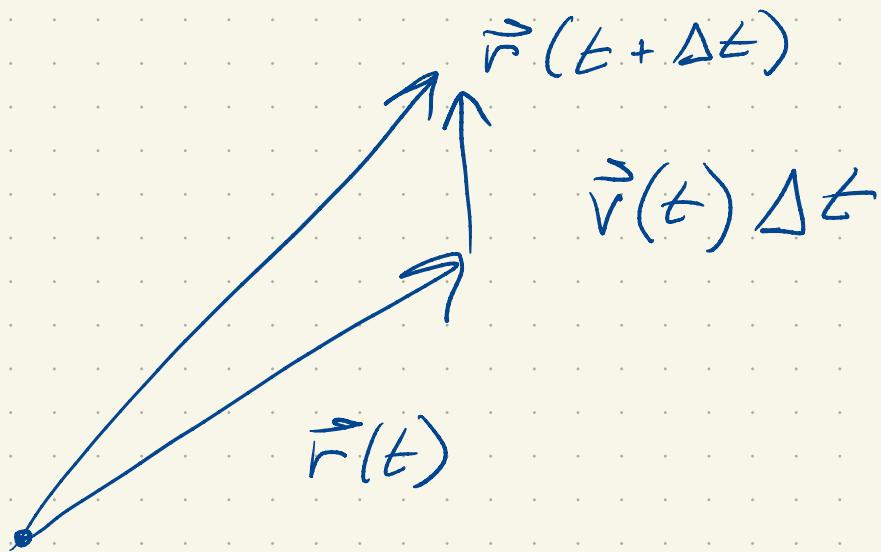
$$= r^2 (\cos \theta \hat{i} + \sin \theta \hat{j}) \hat{k} \times \hat{j} = \theta \hat{k}$$

$$= r^2 \theta' \hat{k}$$

$$L = |\vec{r} \times p| = m r^2 \theta'$$

$$\theta' = \frac{1}{m} \frac{L}{r^2}$$

Angle changes slowly as  $r \rightarrow 0$ .

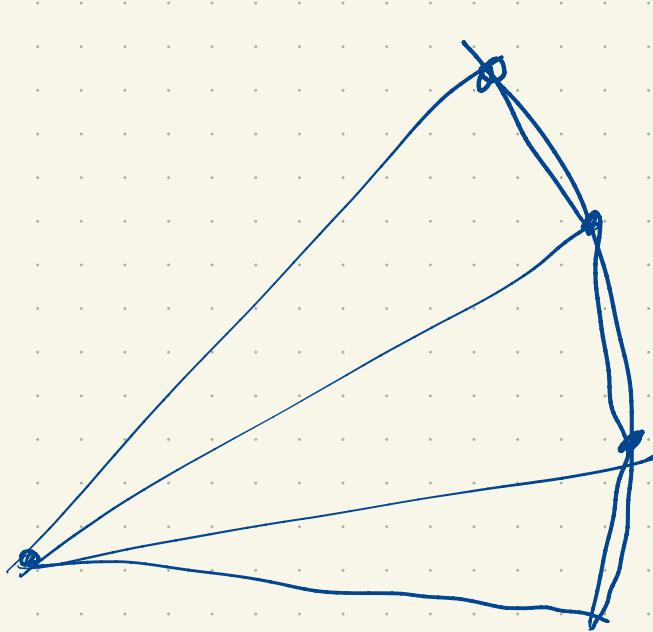


$$\frac{1}{2} |\vec{r} \times \vec{v} \Delta t| = \frac{1}{2} |\vec{r} \times \vec{p}| \frac{\Delta t}{m} = \frac{1}{2} \frac{L}{m} \Delta t$$

is the area of the triangle.

That is, in time  $\Delta t$ , the position vector sweeps out an area approximately  $\frac{L}{2m} \Delta t$

Top: equal  $\Delta t \leftrightarrow$  equal area.



$$\frac{dA}{dt} = \frac{1}{2} \frac{L}{m}$$

Kepler 2: equal times  $\Rightarrow$  equal area swept out.

is exactly conservation of angular momentum