HW 10 - ASTR501

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a)
|a_1|_{1=1} = s\theta = \theta \rightarrow ArcTan[H/r] /.H \rightarrow .04r (*Angle with respect to mid plane in radians*)
Out[1]= \Theta \rightarrow 0.0399787
ln[2]:= ArcTan[D[nr^{(9/7)}, r]] - ArcTan[H/r] //. {H <math>\rightarrow .04 r, r -> 1 au, n \rightarrow .04 r/r^(9/7)}
         (*Angle with surface of disk in radians*)
Out[2]= 0.0114046
       b)
ln[3] =  peak wavelength 5000K [[2]] < 2\pi 0.1 microns
Out[3]= True
ln[4]:= mdust = \left(4/3 \text{ Pi a}^3 3 \text{ g/cm}^3\right);
       s\Sigma = Solve[{1 == Pi a^2 Q Ndust, Ndust == .01 \Sigma gas / mdust}, {\Sigma gas, Ndust}][[1, 1]] //.
           \left\{\text{a}\rightarrow\text{0.1}\,\mu\text{m}\,,\,Q\rightarrow\text{If}\left[\lambda<2\,\text{Pia,1,2}\,\text{Pia}/\lambda\right],\right.
            \lambda \rightarrow \Box peak wavelength 5000K [[2]] // Echo
   \lambda \rightarrow 5.795545800000000 \times 10^{-7} \text{ m}
Out[4]= \Sigmagas \rightarrow 0.04 kg/m<sup>2</sup>
       I don't like to be really precise, since this is Astronomy where everything is order of magnitude.
       c)
ln[5] = Solve[1000 g/cm^2 .01/mdust/H \pi a^2 z 1 au/Sin@\theta == 1, z][[1, 1, 2]] //.
         \{H \rightarrow .04\,r,\,r \rightarrow 1\,au\,,\,s\theta,\,a \rightarrow 0.1\,\mu\text{m}\,\}
Out[5]= 6.39489 \times 10^{-9}
       Ratio with respect to column in b)
ln[6] = \Sigma gas / 1000 g/cm^2 /. s\Sigma
Out[6]= 4. \times 10^{-6}
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