# **HW 4 - ASTR503**

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1)

# Q1)

#### Frame width

$$FrameWidth = \frac{DetectorWidth}{FocalLength}$$

#### Numerical value

```
In[2]:= SFW = FW -> UnitConvert[w/f rad //. {f \rightarrow 6 m , w \rightarrow 1024. 18 \mum }, "ArcMinutes"] Out[2]= FW \rightarrow 10.56076'
```

#### Pixel scale

#### For n x n pixels:

$$PixelScale = \frac{FrameWidth}{n}$$

#### Numerical value

```
\label{eq:ln[3]} $$ \mbox{PS} \to \mbox{UnitConvert[FW / n /. {sFW, n} \to \mbox{1024}}, \mbox{"ArcSeconds"]} $$ \mbox{Out[3]= PS} \to \mbox{0.6187944"}
```

# Q2)

#### Formula for resolution at diffraction limit

```
ln[4]:= f\theta[\lambda_{,} d_{,}] := \theta_{min} \rightarrow UnitConvert[1.22 \lambda/d rad, "ArcSeconds"]
```

### Hubble space telescope

#### At 300nm

$$In[5]:=$$
 f $\theta$ [ 300 nm , 2.4 m]

Out[5]=  $\theta_{min} \rightarrow 0.03145538$ "

#### At $2\mu m$

$$In[6]:=$$
  $f\theta$  [  $2 \mu m$  ,  $2.4 m$  ]

Out[6]=  $\theta_{min} \rightarrow 0.2097026$ "

### 8m space telescope

#### At 2 $\mu$ m

$$\label{eq:normalized_loss} \text{Out}[7]\text{:=} \ \ \textbf{f}\theta \left[ \ \textbf{2} \ \mu \textbf{m} \ \textbf{,} \ \textbf{8} \ \textbf{m} \ \right]$$
 
$$\text{Out}[7]\text{:=} \ \ \theta_{\text{min}} \ \rightarrow \ \textbf{0.06291077"}$$

# Q3)

```
\label{eq:ln[8]:=} $$ \textbf{FormulaData["FNumber"]}$$ $$ Out[8]=$$ N == $\frac{f}{D}$$ $$
```

Here N is the f-number, f is the focal length and D is the aperture.

# a) Refractor (f/10)

```
In[9]:= With [\{f = 10 \text{ m}\}, TubeLength \rightarrow f]
Out[9]= TubeLength \rightarrow 10 \text{ m}
```

# b) Schmidt (f/10)

```
In[10]:= With [\{f = 10 \text{ m}\}, TubeLength \rightarrow 2 \text{ f}]
Out[10]:= TubeLength \rightarrow 20 \text{ m}
```

```
ln[11]:= With [\{f = 2.5 m\}, TubeLength \rightarrow 2 f]
Out[11]= TubeLength \rightarrow 5. m
```

d) Cassegrain (f/10, f/3 primary, final focus at 20cm)

#### **Equations**

$$ln[12]:=$$
 eqd = {f1 == 3 m , f == 10 m ,  $\beta$  f1 == 0.2 m ,  
 m == f / f1, 1 - d / f1 == (1 +  $\beta$ ) / (1 + m) , TubeLength == d +  $\beta$  f1};

#### Solving

```
In[13]:= NSolve[eqd, {TubeLength, \beta, f1, d, f, m}] [1, -1]
Out[13]:= TubeLength \rightarrow 2.461538 m
```

Therefore the TubeLength is 2.462 meters.

e) Cassegrain (f/10, f/2 primary, final focus at 20cm)

### **Equations**

$$ln[14]:=$$
 eqe = {f1 == 2 m , f == 10 m ,  $\beta$  f1 == 0.2 m ,  
 m == f / f1, 1 - d / f1 == (1 +  $\beta$ ) / (1 + m) , TubeLength == d +  $\beta$  f1};

### Solving

```
In [37]:= NSolve [eqe, {TubeLength, \beta, f1, d, f, m}] [1, -1]]
Out[37]= TubeLength \rightarrow 1.8333333 m
```

Therefore the TubeLength is 1.833 meters.

# Q4)

### **Equations**

$$\begin{array}{lll} \text{m1} & -\text{ m2} & = & -2.5 \text{ Log}_{10} \text{ (F1 / F2)} \\ \text{mb1} & -\text{ mb2} & = & -2.5 \text{ Log}_{10} \text{ (b1 / b2)} & = & -1 \\ \\ \frac{\text{F1}}{\text{F2}} & = & \sqrt{\frac{\text{b1}}{\text{b2}}} & \frac{\text{D2}}{\text{D1}} \\ \text{m1} & = & 26 \end{array}$$

### Solving for m2

```
In[16]:= Solve [26 - m2 == -2.5 Log10 [Sqrt@rb 6.5 / 2.4] /. Solve [1 == -2.5 Log10@rb, rb] [1], m2] [1, 1] Out[16]= m2 \rightarrow 26.58176
```

# Q2) Observing Andromeda

#### Fetching values from WolframAlpha:

```
Grid Prepend Thread@Prepend (ev={"Rise","Transit","Set"}) @

{| 1t = (ob = M31 ) @EntityProperty [ob@"ObjectType",#,

{"Location"→ (ps = 40.1105° N, 88.2284° W), "Date"→ sep 30 }] &/@

(#<>"Time"&/@ev), st=SiderealTime[ps,#]&/@lt,st-ob@"RightAscension"},

{| ob@"Name", "Local Time", "Local Sidereal Time", "Hour Angle"}], Frame→All
```

	M31	Local Time	Local Sidereal Time	Hour Angle
	Rise	Fri 30 Sep 2016 15:40 GMT-5.	15 <sup>h</sup> 25 <sup>m</sup> 26.23772 <sup>s</sup>	14 <sup>h</sup> 42 <sup>m</sup> 41.93772 <sup>s</sup>
Out[17]=	Transit	Fri 30 Sep 2016 00:59 GMT-5.	0 <sup>h</sup> 42 <sup>m</sup> 1.511847 <sup>s</sup>	-42.78815 <sup>s</sup>
	Set	Fri 30 Sep 2016 10:14 GMT-5.	9 <sup>h</sup> 58 <sup>m</sup> 32.68422 <sup>s</sup>	9 <sup>h</sup> 15 <sup>m</sup> 48.38422 <sup>s</sup>

Q3)

#### SPT

```
In[18]:= \{\sigma \rightarrow 8 \times 10^-5, SNR_{Max} \rightarrow 8.57\} // Column

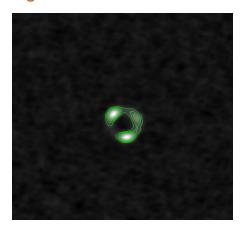
Out[18]:= \sigma \rightarrow \frac{1}{12500}

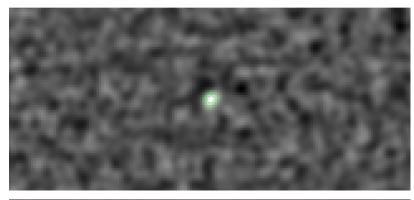
SNR_{Max} \rightarrow 8.57
```

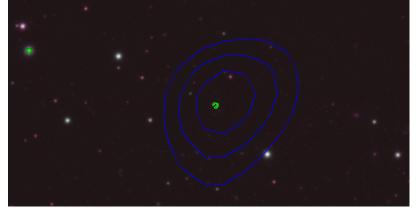
#### **ALMA**

```
In[19]:= \{\sigma \to 1.4 \times 10^{-3}, SNR_{Max} \to 18.99\} // Column
Out[19]:= \begin{cases} \sigma \to 0.0014 \\ SNR_{Max} \to 18.99 \end{cases}
```

# Images







Q4)

### Selection criteria

- SNR of either 2mm or 1.4mm >  $4.5\sigma$
- Does not appear in IRAC or SUMSS catalog
- Posterior probability  $P(\alpha>1.66) > 0.5$

# Importing SPT catalog

```
In[36]:= catSPT = SemanticImport[NotebookDirectory[] <> "PointSourceSPT.dat"];
```

#### Selecting lensed sources

```
ln[21]:= lensed = catSPT[Select[(#["S/N_2.0mm"] > 4.5 || #["S/N_1.4mm"] > 4.5) &&
            #["P(alpha>1)"] > .5 && #["dR_SUMSS[arcsec]"] > 60 && #["dR_IRAS[arcsec]"] > 60 & ]];
```

#### Number of lensed sources

```
In[22]:= Length@lensed
```

 $\mathsf{Out}[22] = \ 33$ 

#### **Plot**

 $\label{listLogLogPlot[Normal/@(#[All, {"S_raw_2.0mm[mJy]", "S_raw_1.4mm[mJy]"}) & /@ {catSPT, lensed}), }$ PlotMarkers → {".", "x"}, Frame → True, PlotLegends → Placed[{"All", "Lensed"}, {Left, Bottom}], FrameLabel → {"S\_raw\_2.0mm[mJy]", "S\_raw\_1.4mm[mJy]"}, ImageSize → 500]

