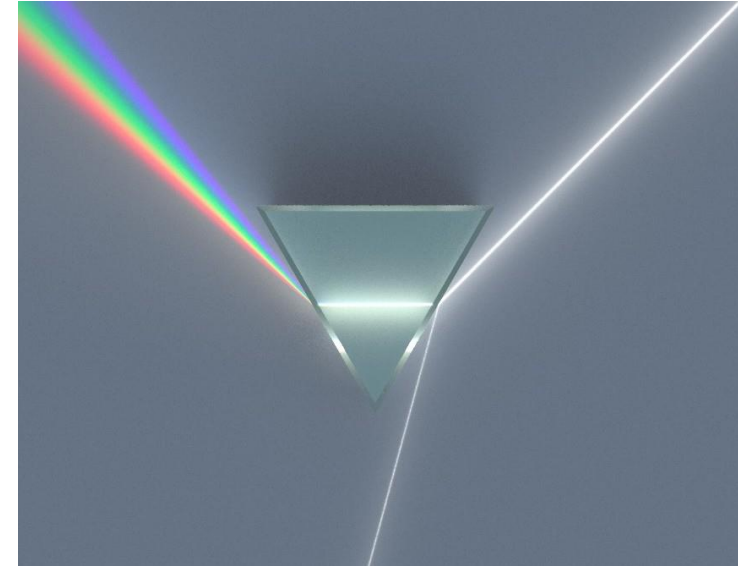
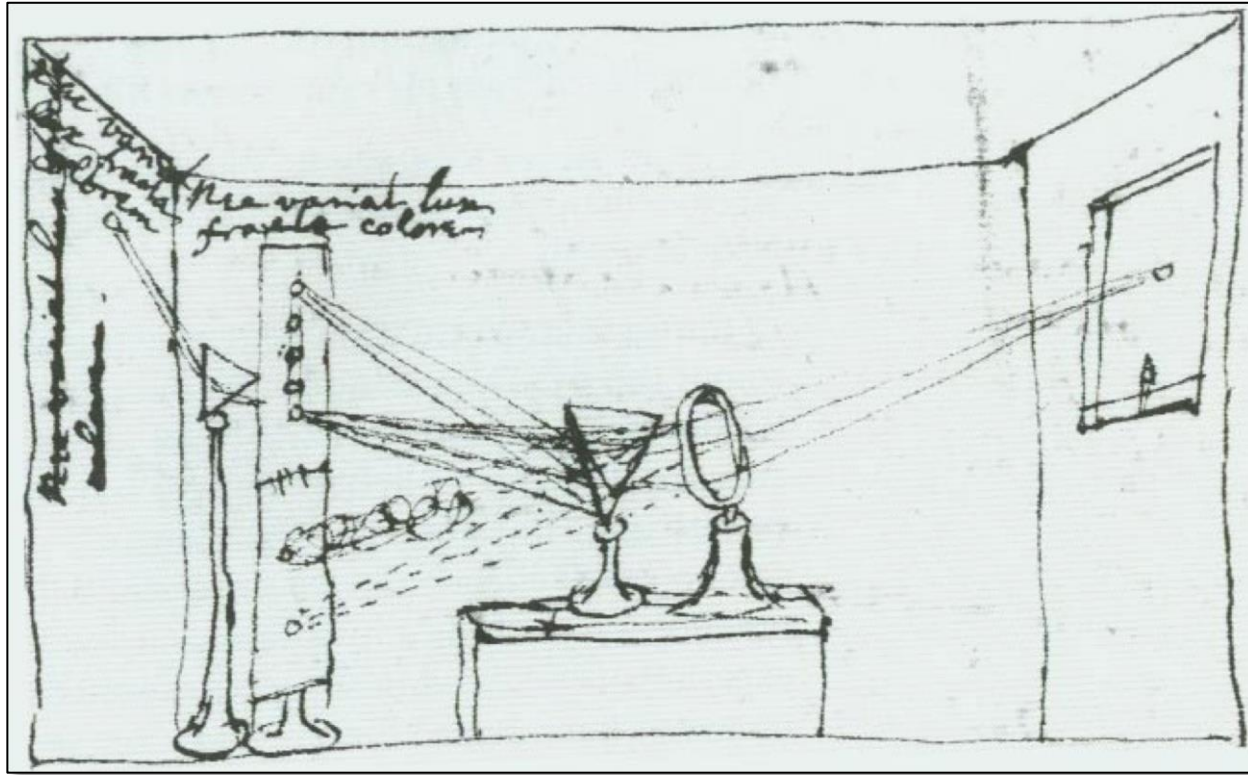


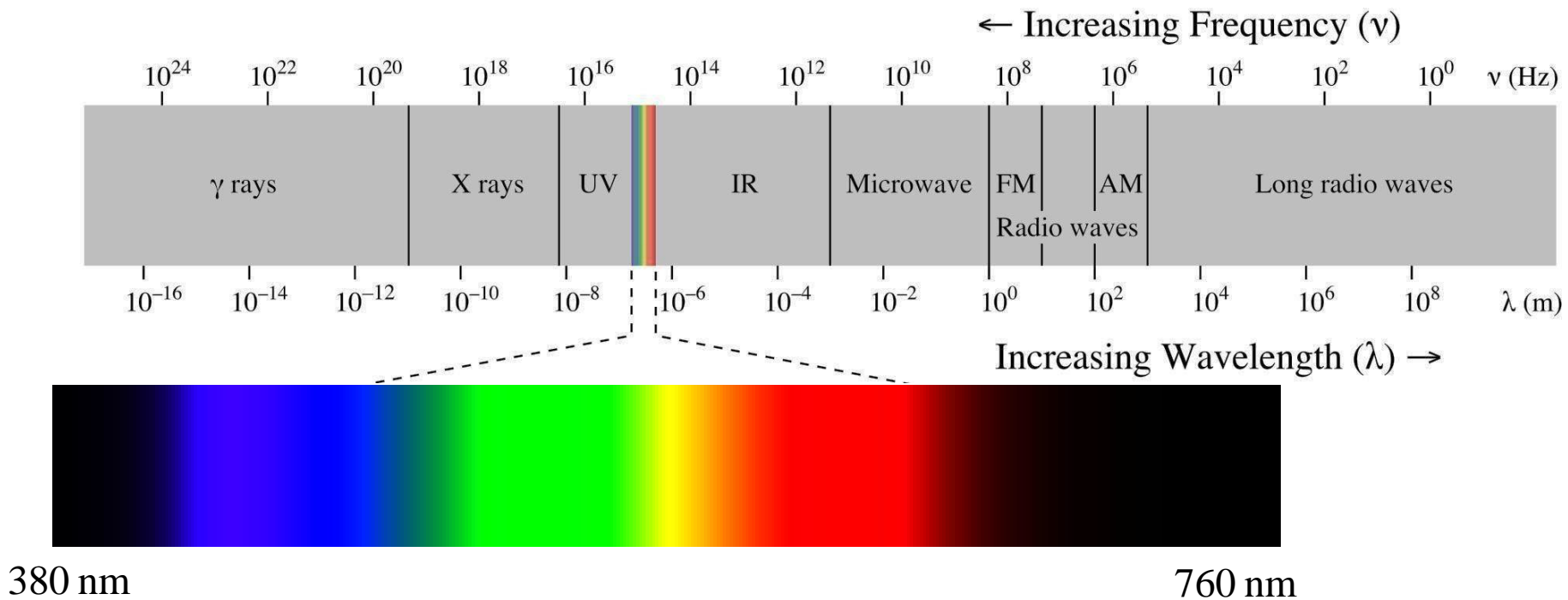
Introduction to color science

- Trichromacy
- Spectral matching functions
- CIE XYZ color system
- xy-chromaticity diagram
- Color gamut
- Color temperature
- Color balancing algorithms

Newton's Prism Experiment - 1666



Color: visible range of the electromagnetic spectrum



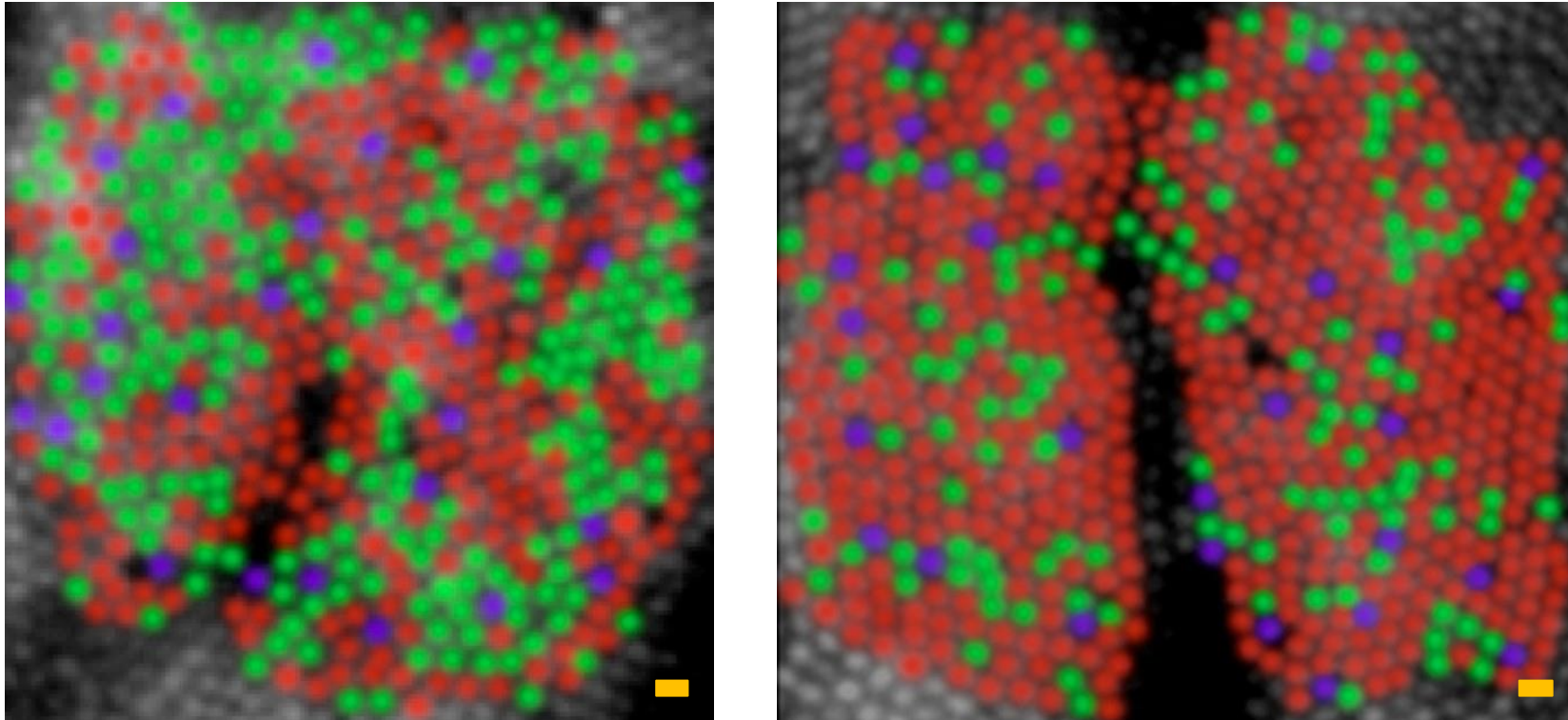
Radiometric quantities

Quantity		Unit			Dimension	Notes					
Name	Symbol ^[nb 1]	Name	Symbol	Symbol							
Radiant energy		Radiance	$L_{\text{e},\Omega}$ ^[nb 5]	watt per steradian per square metre	$\text{W}\cdot\text{sr}^{-1}\cdot\text{m}^{-2}$	$\text{M}\cdot\text{T}^{-3}$	Radiant flux emitted, reflected, transmitted or received by a <i>surface</i> , per unit solid angle per unit projected area. This is a <i>directional</i> quantity. This is sometimes also confusingly called "intensity".				
Radiant energy dens											
Radiant flux											
Spectral flux											
Radiant intensity											
Spectral intensity											
Radiance		Spectral radiance	$L_{\text{e},\Omega,\nu}$ ^[nb 3] or $L_{\text{e},\Omega,\lambda}$ ^[nb 4]	watt per steradian per square metre per hertz <i>or</i> watt per steradian per square metre, per metre	$\text{W}\cdot\text{sr}^{-1}\cdot\text{m}^{-2}\cdot\text{Hz}^{-1}$ <i>or</i> $\text{W}\cdot\text{sr}^{-1}\cdot\text{m}^{-3}$	$\text{M}\cdot\text{T}^{-2}$ <i>or</i> $\text{M}\cdot\text{L}^{-1}\cdot\text{T}^{-3}$	Radiance of a <i>surface</i> per unit frequency or wavelength. The latter is commonly measured in $\text{W}\cdot\text{sr}^{-1}\cdot\text{m}^{-2}\cdot\text{nm}^{-1}$. This is a <i>directional</i> quantity. This is sometimes also confusingly called "spectral intensity".				
Spectral radiance											
Irradiance											
Spectral irradiance											
Radiosity											
Spectral radiosity											
Radiant exitance		Irradiance	E_{e} ^[nb 2]	watt per square metre	W/m^2	$\text{M}\cdot\text{T}^{-3}$	Radiant flux <i>received</i> by a <i>surface</i> per unit area. This is sometimes also confusingly called "intensity".				
Spectral exitance	<i>or</i> $M_{\text{e},\lambda}$ ^[nb 4]							<i>or</i> watt per square metre, per metre	<i>or</i> W/m^3	<i>or</i> $\text{M}\cdot\text{L}^{-1}\cdot\text{T}^{-3}$	"Spectral emittance" is an old term for this quantity. This is sometimes also confusingly called "spectral intensity".
Radiant exposure	H_{e}							joule per square metre	J/m^2	$\text{M}\cdot\text{T}^{-2}$	Radiant energy received by a <i>surface</i> per unit area, or equivalently irradiance of a <i>surface</i> integrated over time of irradiation. This is sometimes also called "radiant fluence".
Spectral exposure	$H_{\text{e},\nu}$ ^[nb 3] <i>or</i> $H_{\text{e},\lambda}$ ^[nb 4]	joule per square metre per hertz <i>or</i> joule per square metre, per metre	$\text{J}\cdot\text{m}^{-2}\cdot\text{Hz}^{-1}$ <i>or</i> J/m^3	$\text{M}\cdot\text{T}^{-1}$ <i>or</i> $\text{M}\cdot\text{L}^{-1}\cdot\text{T}^{-2}$	Radiant exposure of a <i>surface</i> per unit frequency or wavelength. The latter is commonly measured in $\text{J}\cdot\text{m}^{-2}\cdot\text{nm}^{-1}$. This is sometimes also called "spectral fluence".						

Photometric quantities

Quantity		Unit		Dimension	Notes
Name	Symbol ^[nb 1]	Name	Symbol	Symbol	
Luminous energy	Q_v ^[nb 2]	lumen second	$\text{lm} \cdot \text{s}$	$\text{T} \cdot \text{J}$ ^[nb 3]	Units are sometimes called <i>talbots</i> .
Luminous flux / Luminous power	ϕ_v ^[nb 2]	lumen (= cd · sr)	lm	J ^[nb 3]	Luminous energy per unit time.
Luminous intensity	I_v	candela (= lm/sr)	cd	J ^[nb 3]	Luminous power per unit solid angle.
Luminance	L_v	candela per square metre	cd/m^2	$\text{L}^{-2} \cdot \text{J}$	Luminous power per unit solid angle per unit <i>projected</i> source area. Units are sometimes called <i>nits</i> .
Illuminance	E_v	lux (= lm/m ²)	lx	$\text{L}^{-2} \cdot \text{J}$	Luminous power <i>incident</i> on a surface.
Luminous exitance / Luminous emittance	M_v	lux	lx	$\text{L}^{-2} \cdot \text{J}$	Luminous power <i>emitted</i> from a surface.
Luminous exposure	H_v	lux second	$\text{lx} \cdot \text{s}$	$\text{L}^{-2} \cdot \text{T} \cdot \text{J}$	
Luminous energy density	ω_v	lumen second per cubic metre	$\text{lm} \cdot \text{s} \cdot \text{m}^{-3}$	$\text{L}^{-3} \cdot \text{T} \cdot \text{J}$	
Luminous efficacy	η ^[nb 2]	lumen per watt	lm/W	$\text{M}^{-1} \cdot \text{L}^{-2} \cdot \text{T}^3 \cdot \text{J}$	Ratio of luminous flux to radiant flux.
Luminous efficiency / Luminous coefficient	V			1	

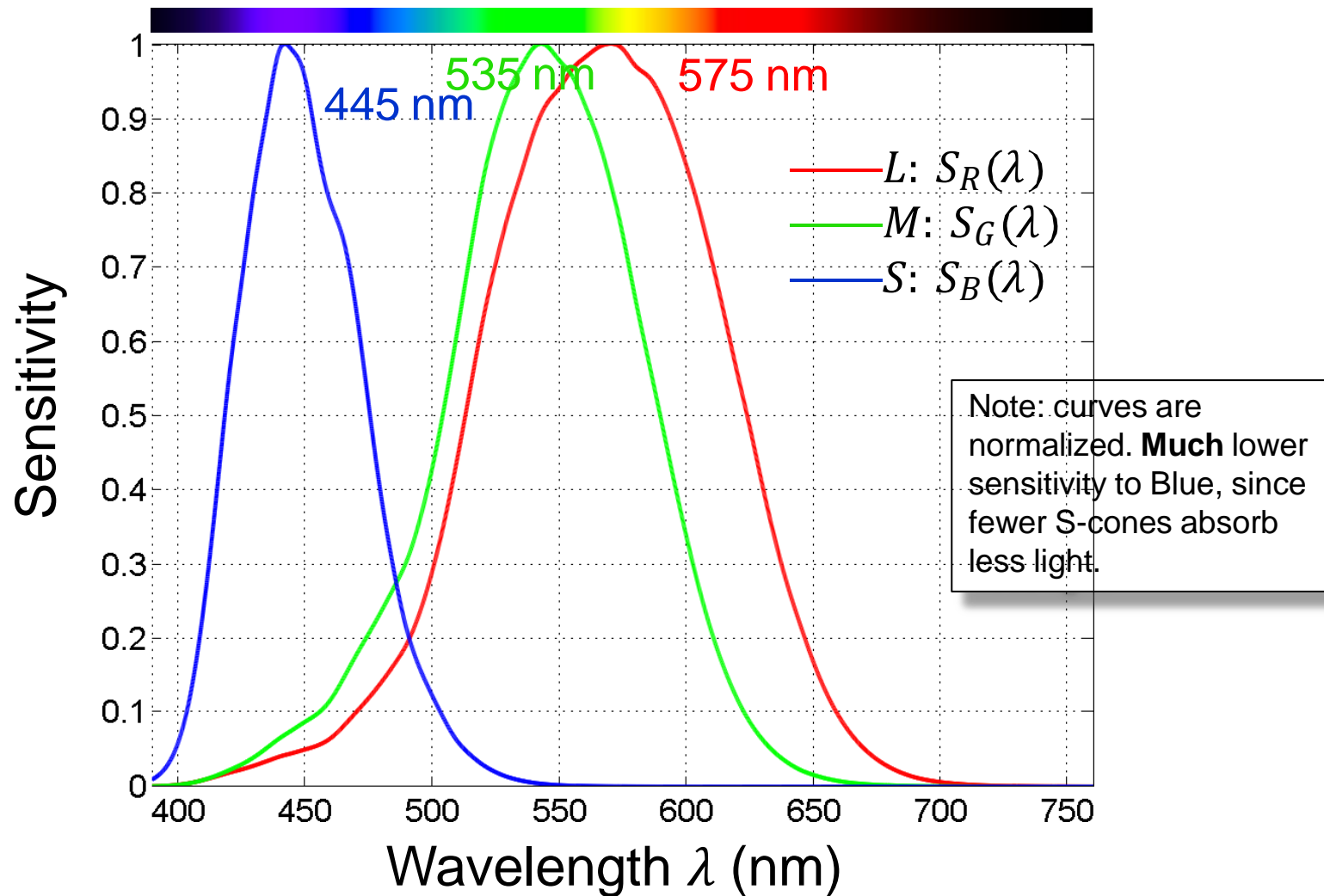
Human retina



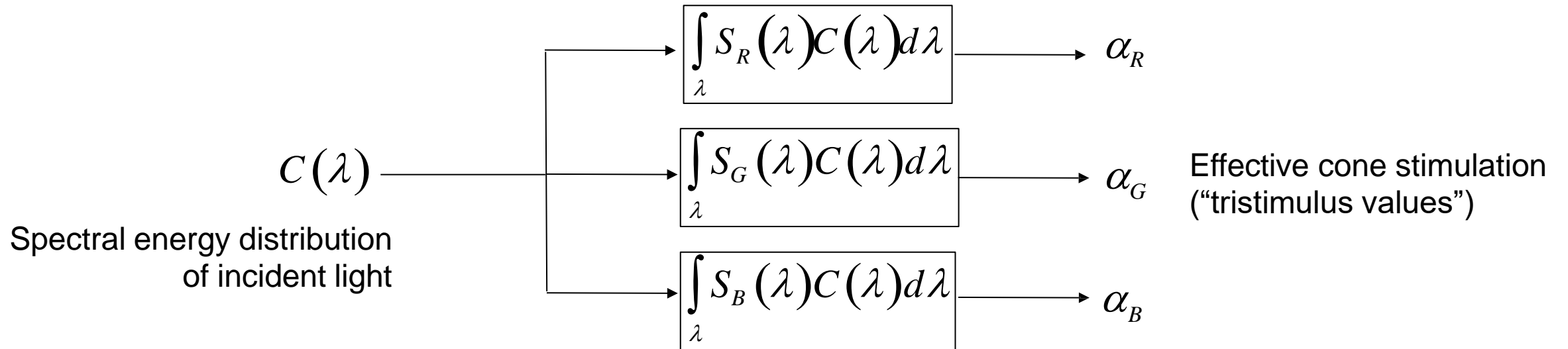
[Roorda, Williams, 1999]

Pseudo-color image of nasal retina,
1 degree eccentricity, in two male subjects, scale bar 5 micron

Absorption of light in the cones of the human retina



Three-receptor model of color perception

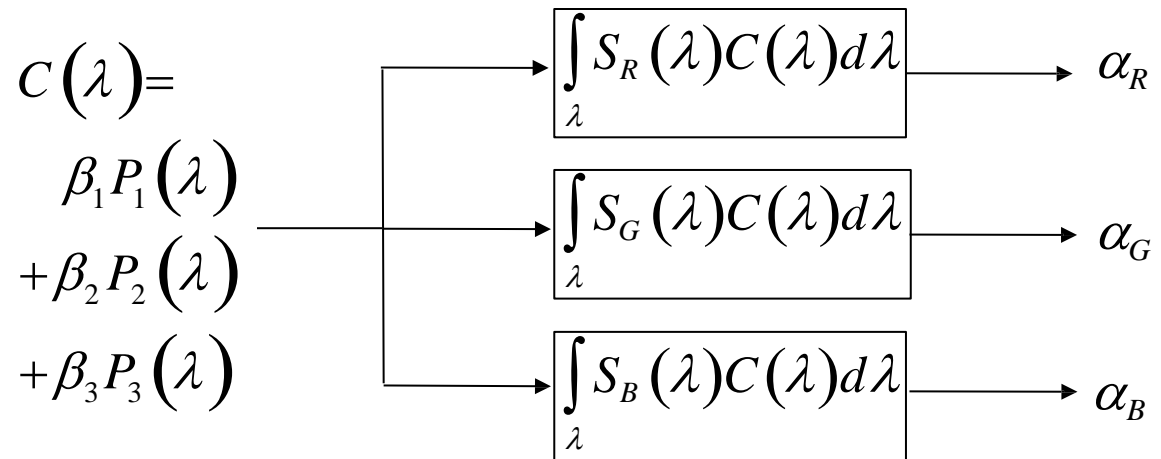


[T. Young, 1802] [J.C. Maxwell, 1890]

- Different spectra can map into the same tristimulus values and hence look identical ("metamers")
- Three numbers suffice to represent any color

Color matching

- Suppose 3 primary light sources with spectra $P_k(\lambda)$, $k=1,2,3$
- Intensity of each light source can be adjusted by factor β_k
- How to choose β_k , $k=1,2,3$, such that desired tristimulus values $(\alpha_R, \alpha_G, \alpha_B)$ result ?



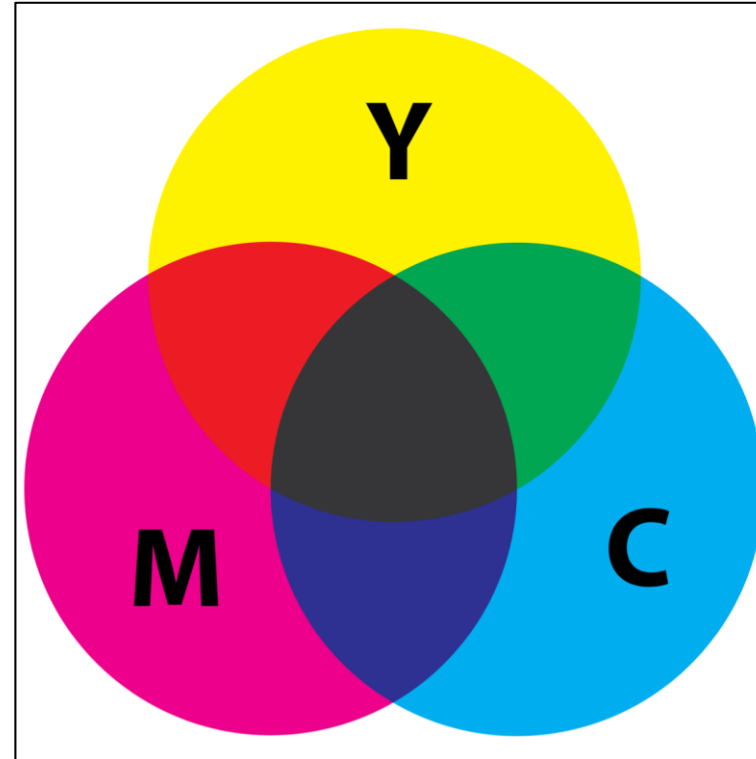
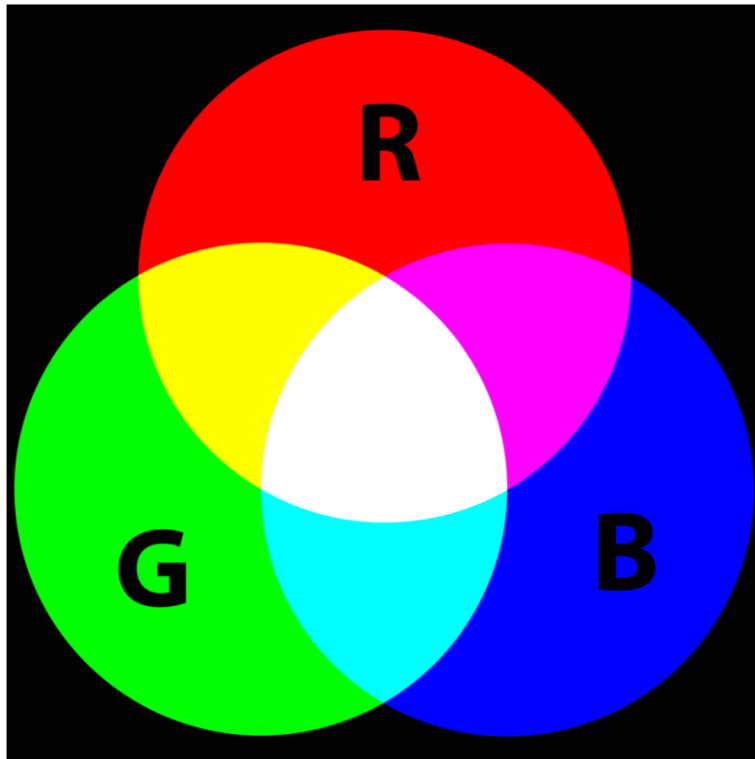
$$\alpha_i = \int_{\lambda} S_i(\lambda) [\beta_1 P_1(\lambda) + \beta_2 P_2(\lambda) + \beta_3 P_3(\lambda)] d\lambda$$

$$= \beta_1 \cdot K_{i,1} + \beta_2 \cdot K_{i,2} + \beta_3 \cdot K_{i,3}$$

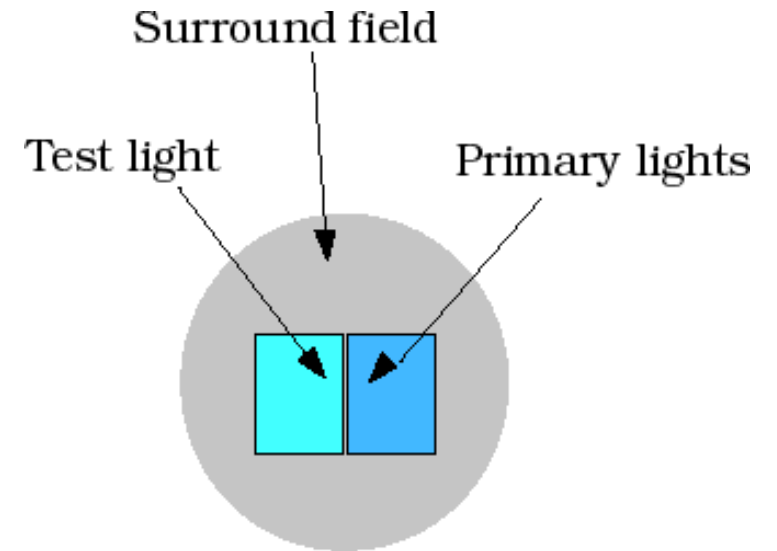
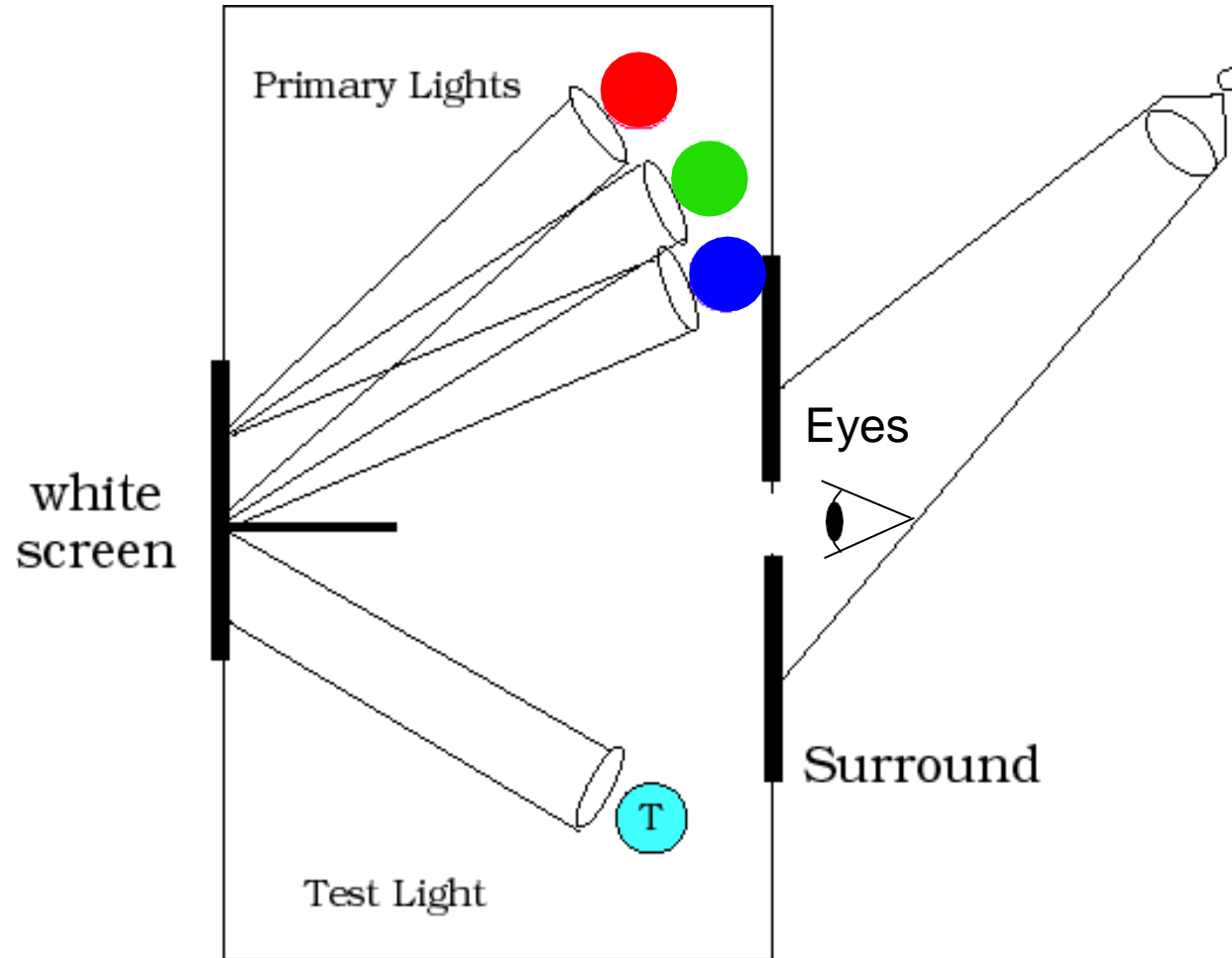
$$\text{with } K_{i,j} = \int_{\lambda} S_i(\lambda) P_j(\lambda) d\lambda$$

Color matching is linear!

Additive vs. subtractive color mixing

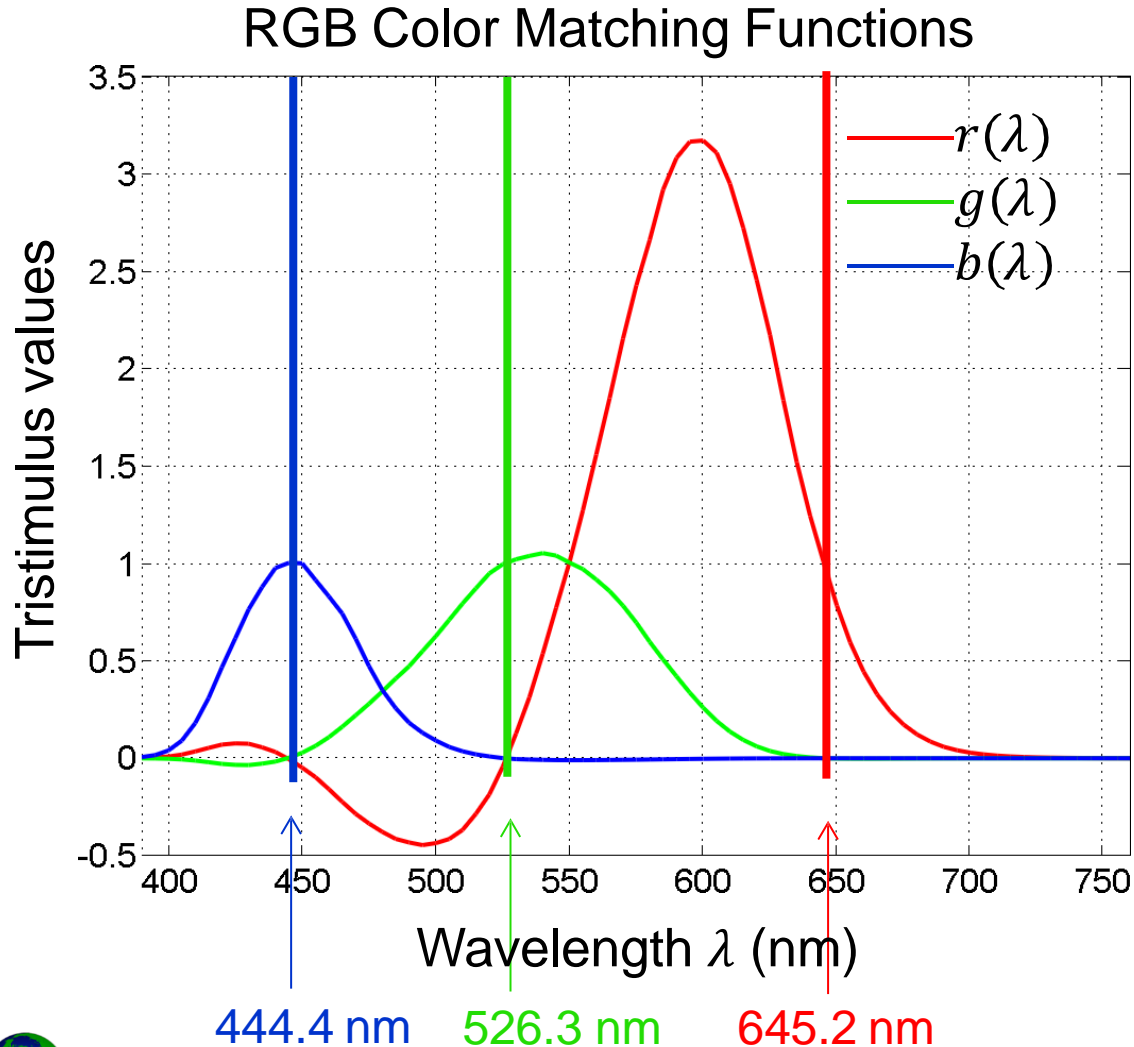


Color matching experiment



Courtesy B. Wandell, from [Foundations of Vision, 1996]

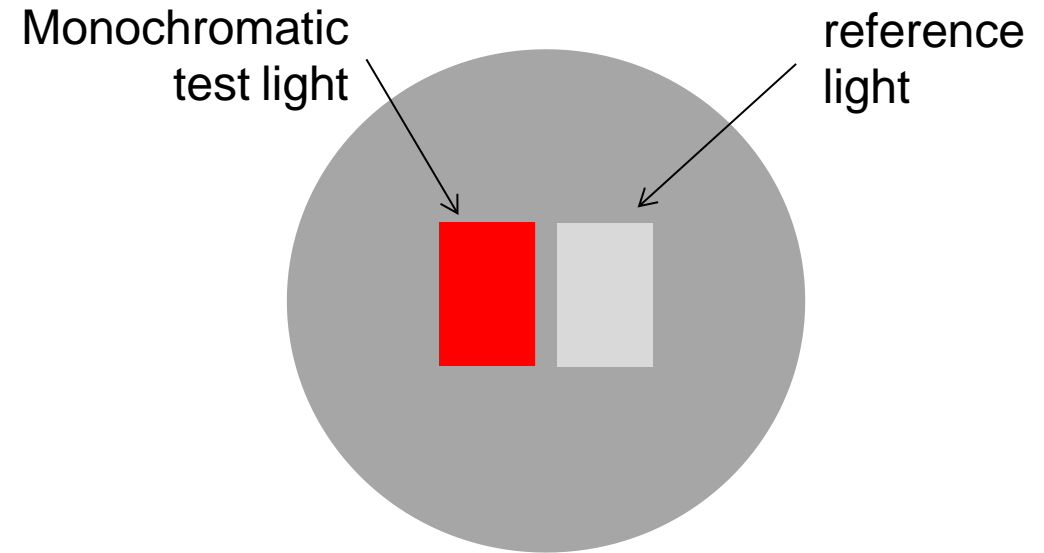
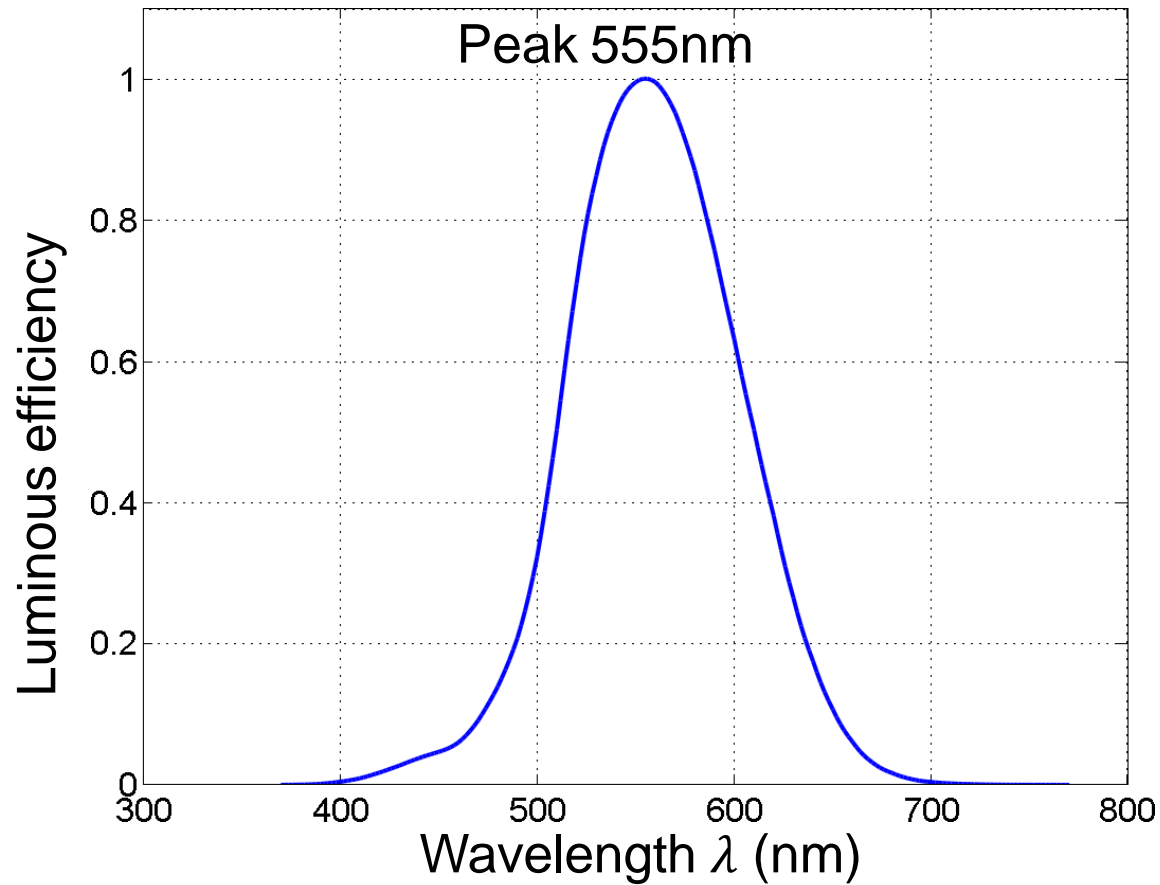
Spectral matching functions



- Color matching experiment: Monochromatic test light vs. mixture of 3 monochromatic primaries
- “Negative intensity”: color is added to test color
- CIE (Commision Internationale de L’Eclairage), 1931: Spectral RGB primaries (scaled, such that $R_\lambda=G_\lambda=B_\lambda$ matches spectrally flat white).



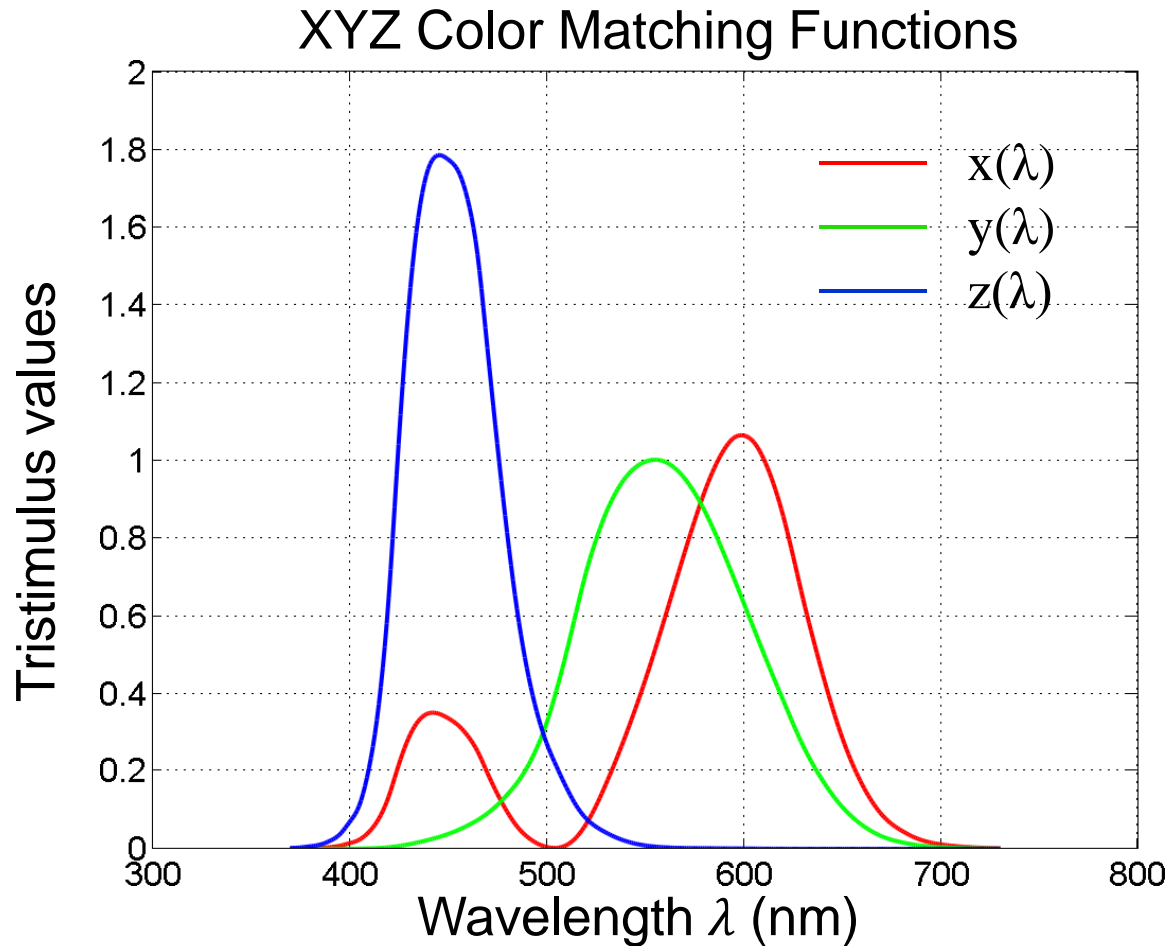
Luminosity function



- Experiment:
Match the brightness of a white reference light and a monochromatic test light of wavelength λ
- Links photometric to radiometric quantities



CIE 1931 XYZ color system



Properties:

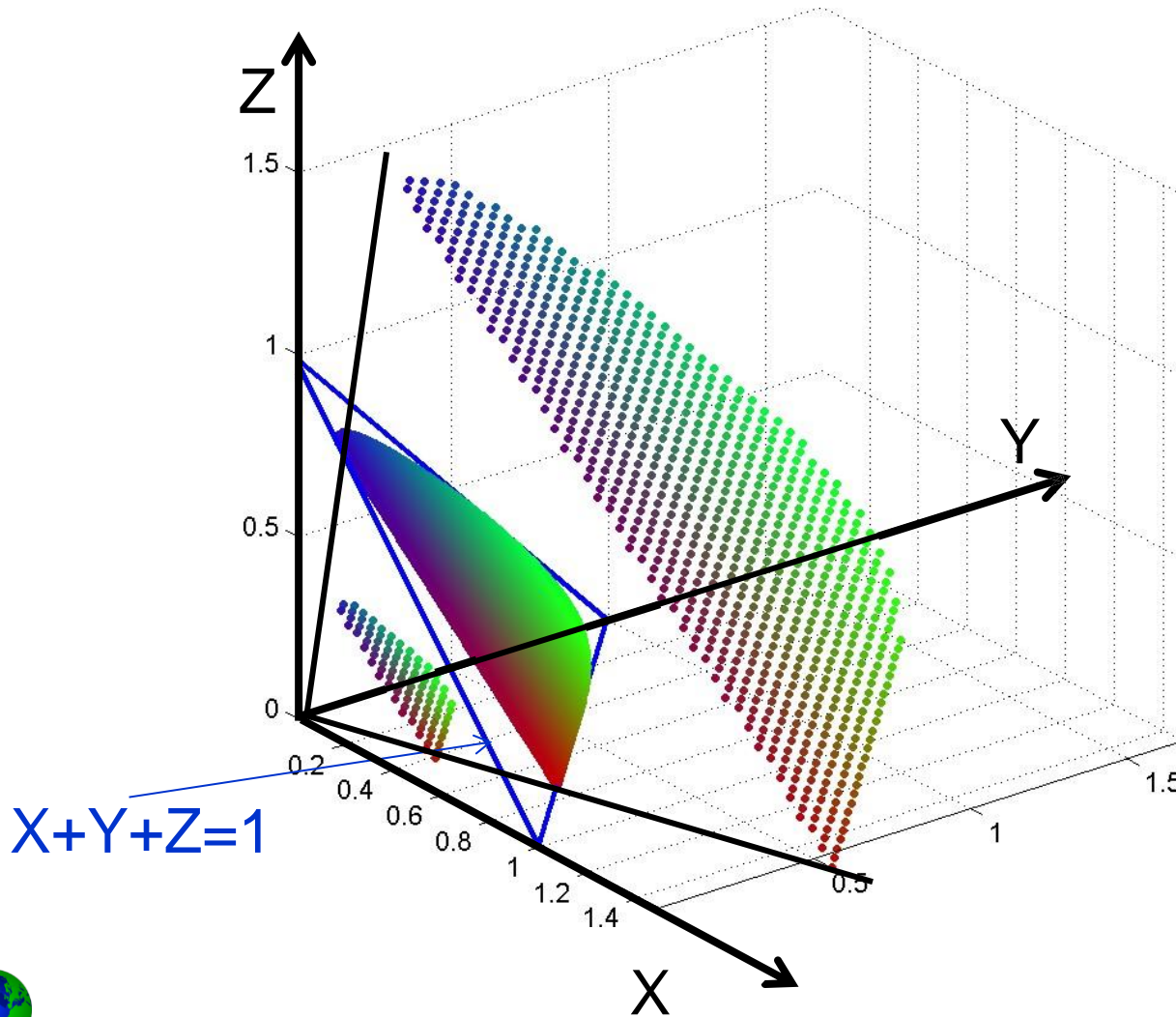
- All positive spectral matching functions

$$\begin{pmatrix} X \\ Y \\ Z \end{pmatrix} = \begin{pmatrix} .490 & .310 & .200 \\ .177 & .813 & .011 \\ .000 & .010 & .990 \end{pmatrix} \begin{pmatrix} R_\lambda \\ G_\lambda \\ B_\lambda \end{pmatrix}$$

- Y corresponds to luminance
- Equal energy white: $X=Y=Z$
- Virtual primaries



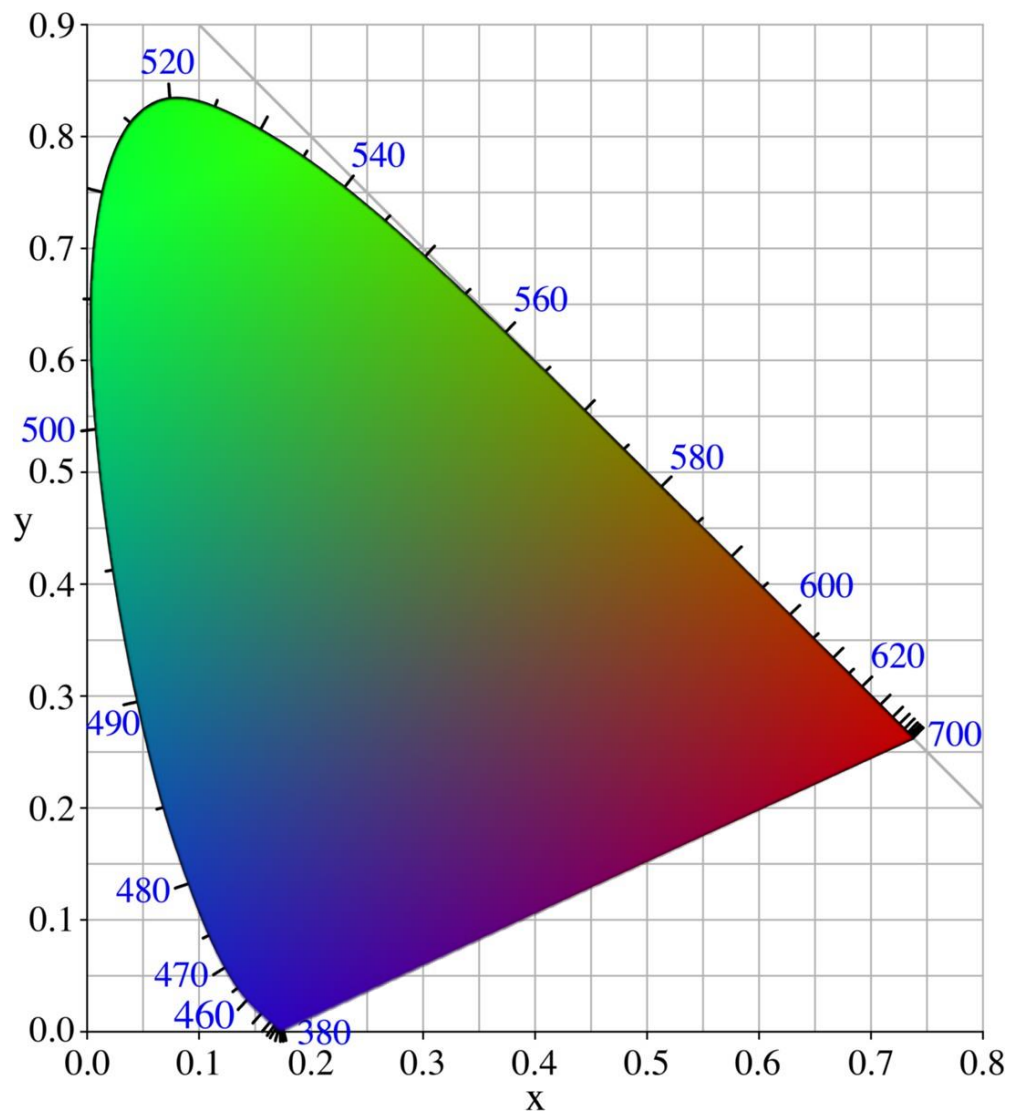
Color gamut and chromaticity



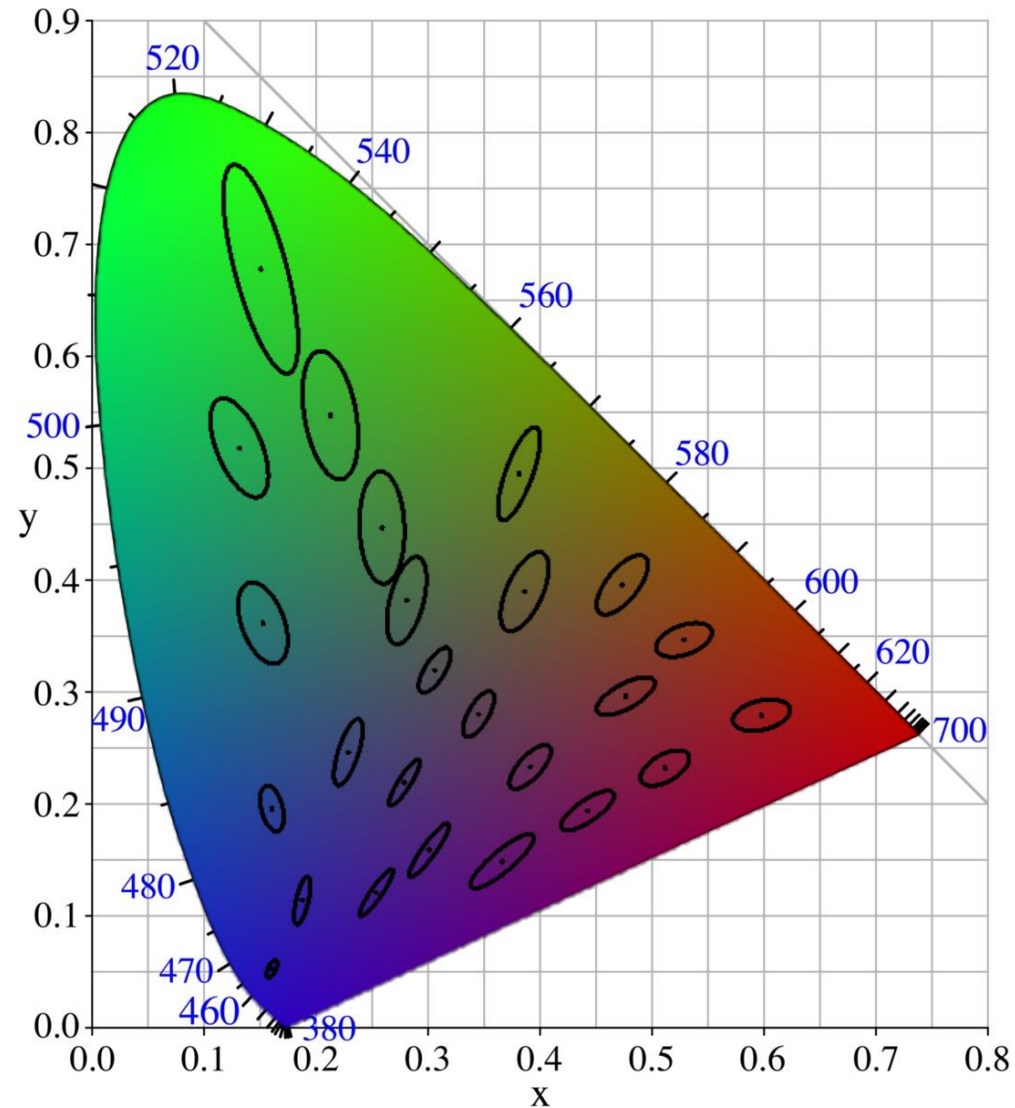
$$x = \frac{X}{X + Y + Z}$$
$$y = \frac{Y}{X + Y + Z}$$



CIE chromaticity diagram



Perceptual non-uniformity of xy chromaticity

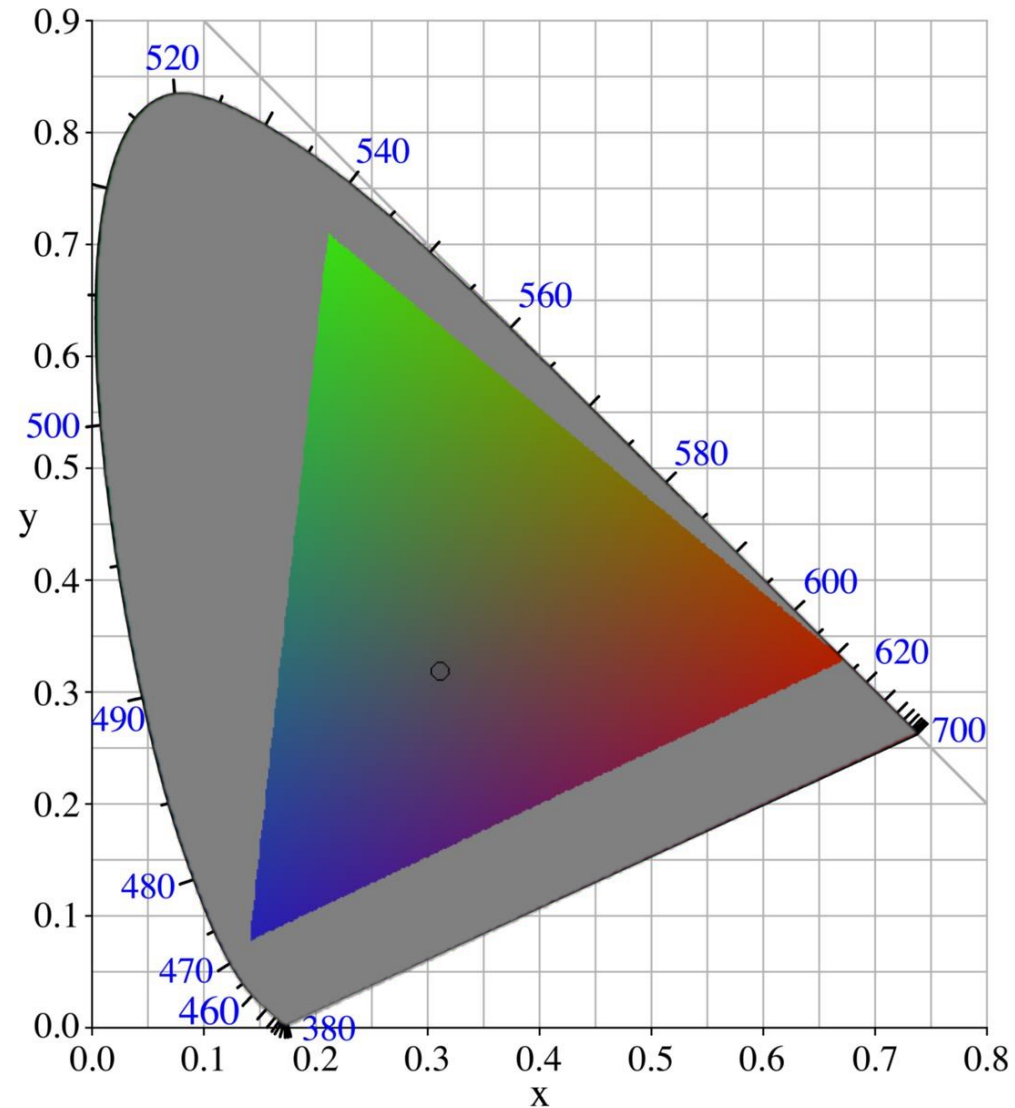


Just noticeable chromaticity differences (10X enlarged)

[MacAdam, 1942]



Color gamut



NTSC phosphors

R: $x=0.67$, $y=0.33$

G: $x=0.21$, $y=0.71$

B: $x=0.14$, $y=0.08$

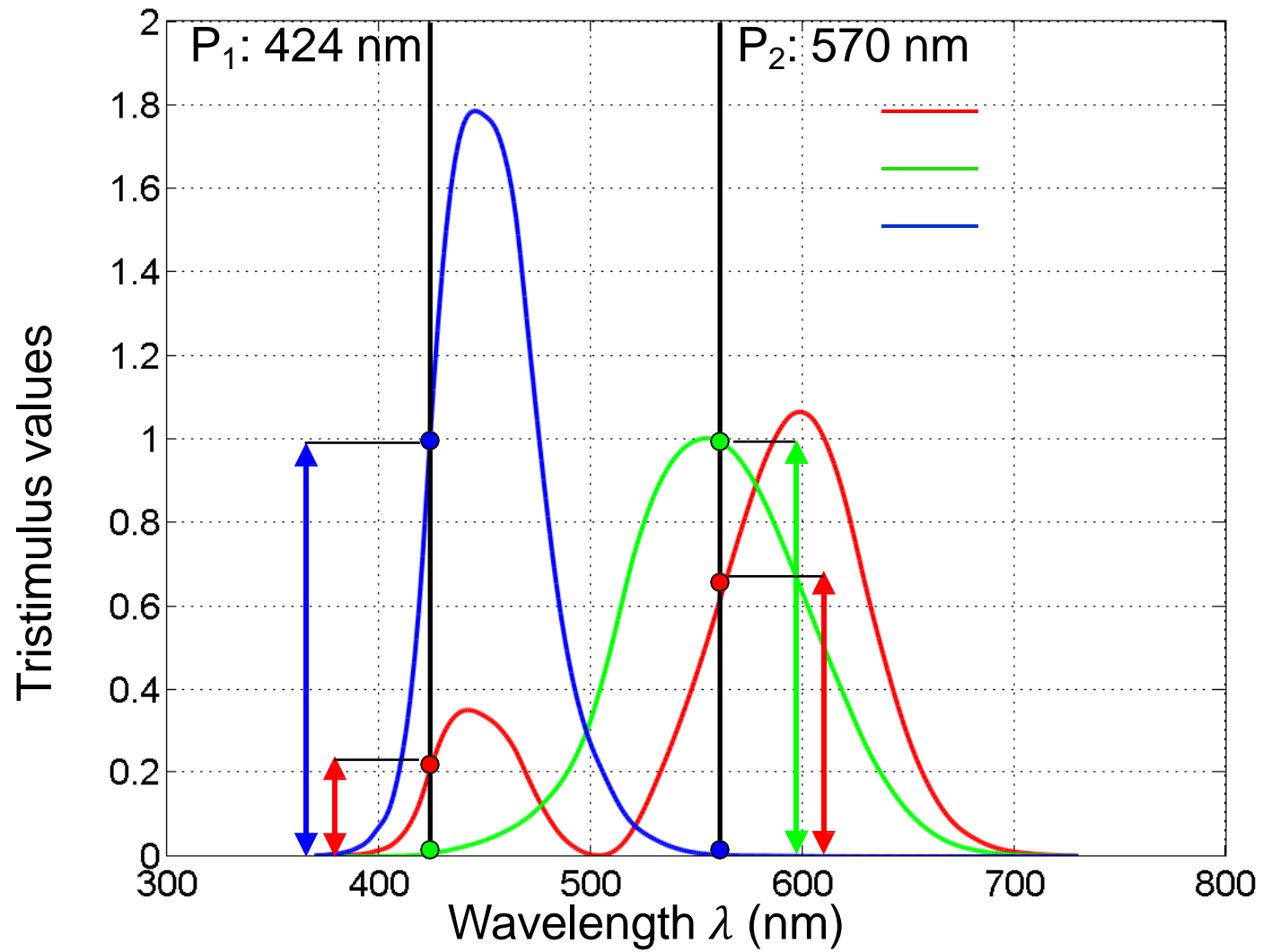
Reference white:

$x=0.31$, $y=0.32$

Illuminant C



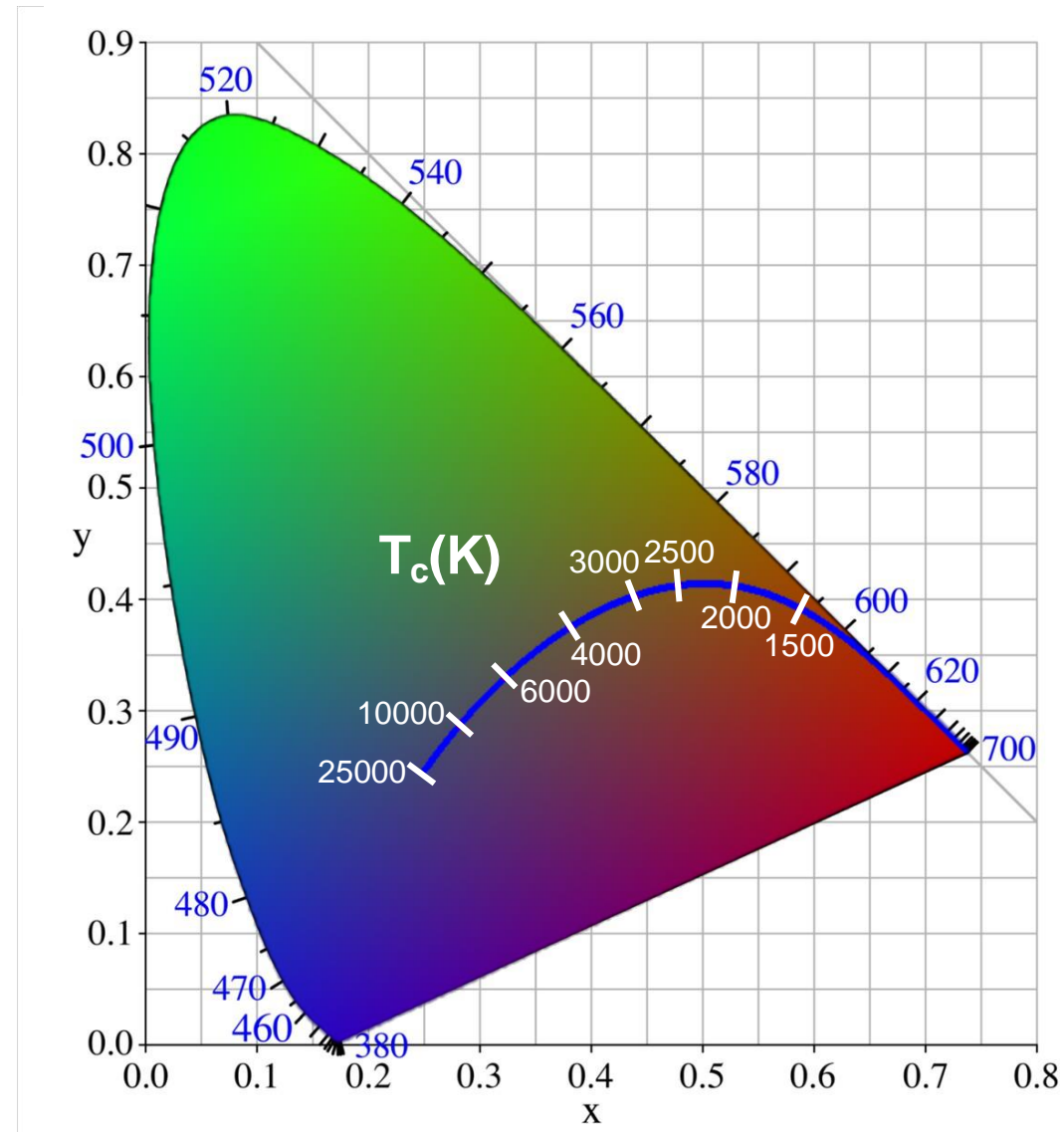
XYZ Color Matching Functions



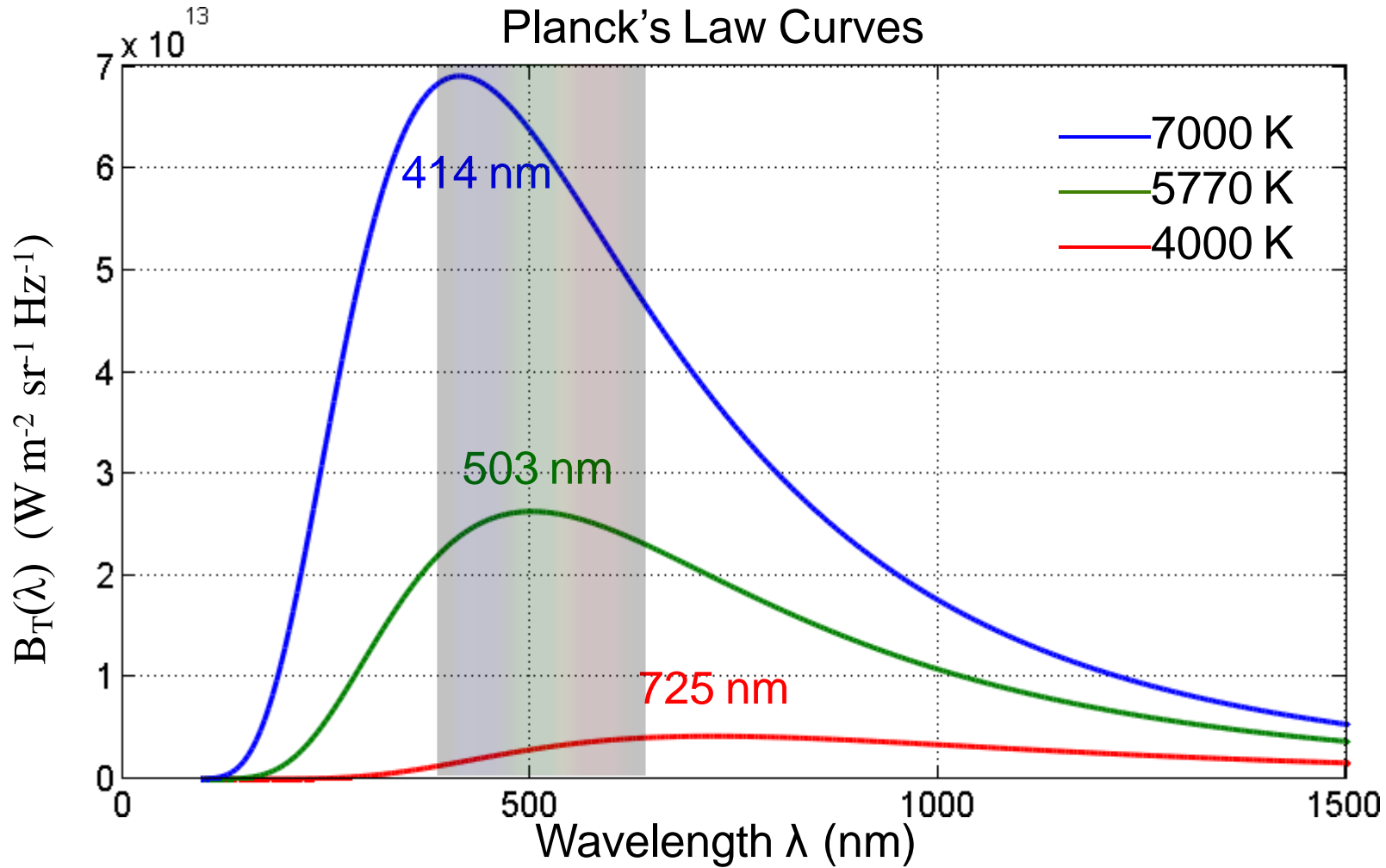
$Z_1 + Z_2$ $Y_1 + Y_2$ $X_1 + X_2$

Blue arrow (up and down) Green arrow (up and down) Red arrow (up and down)

White at different color temperatures



Blackbody radiation



Planck's Law, 1900

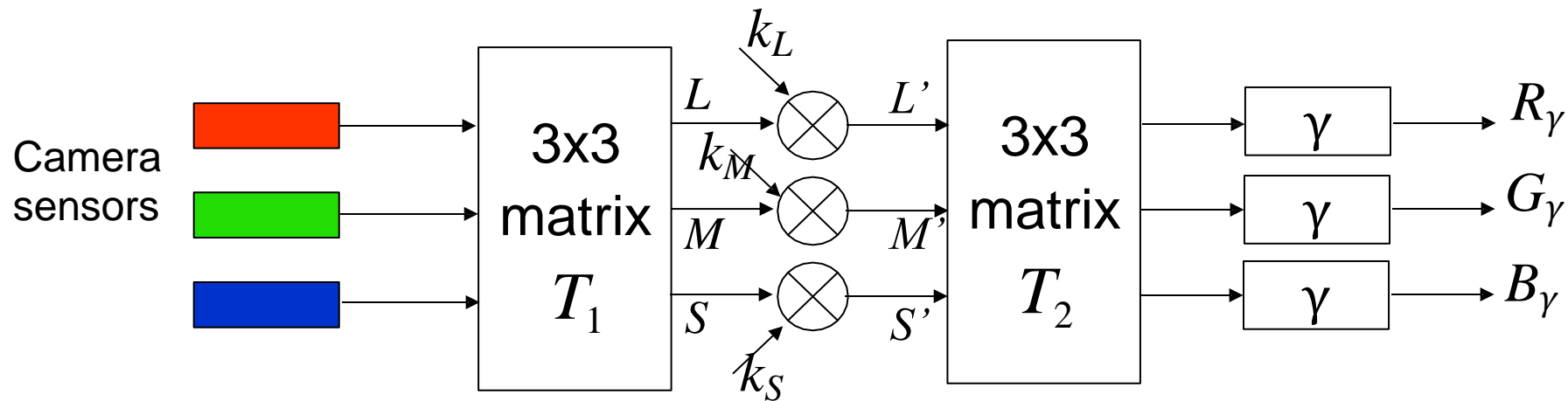
$$B_T(\lambda) = \frac{2hc^2 / \lambda^5}{e^{hc/\lambda kT} - 1}$$

Wien's Law

$$\lambda_{peak} [nm] = \frac{2,900,000}{T[K]}$$

Color balancing

- Effect of different illuminants can be cancelled only in the spectral domain (impractical)
- Color balancing in 3-d color space is practical approximation
- Color constancy in human visual system: gain control in cone space LMS [[von Kries, 1902](#)]
- Von Kries hypothesis applied to image acquisition devices (cameras, scanners)



- How to determine k_L , k_M , k_S automatically?

Color balancing (cont.)

- Von Kries hypothesis

$$\begin{pmatrix} L' \\ M' \\ S' \end{pmatrix} = \begin{pmatrix} k_L & 0 & 0 \\ 0 & k_M & 0 \\ 0 & 0 & k_S \end{pmatrix} \begin{pmatrix} L \\ M \\ S \end{pmatrix}$$

- If illumination (or a patch of white in the scene) is known, calculate

$$k_L = \frac{L_{desired}}{L_{actual}}; \quad k_M = \frac{M_{desired}}{M_{actual}}; \quad k_S = \frac{S_{desired}}{S_{actual}}$$

Color balancing with unknown illumination

- Gray-world

$$k_L \sum_{x,y} L[x, y] = k_M \sum_{x,y} M[x, y] = k_S \sum_{x,y} S[x, y]$$

- Scale-by-max

$$k_L \max_{x,y} L[x, y] = k_M \max_{x,y} M[x, y] = k_S \max_{x,y} S[x, y]$$

- Shades-of-gray

[Finlayson, Trezzi, 2004]

$$k_L \left(\sum_{x,y} L^p[x, y] \right)^{\frac{1}{p}} = k_M \left(\sum_{x,y} M^p[x, y] \right)^{\frac{1}{p}} = k_S \left(\sum_{x,y} S^p[x, y] \right)^{\frac{1}{p}}$$

- » Special cases: gray-world ($p = 1$), scale-by-max ($p = \infty$)
- » Best performance for $p \approx 6$

- Refinements:

smooth image, exclude saturated color/dark pixels,
use spatial derivatives instead (“gray-edge,” “max-edge”)

[van de Weijer, 2007]

Color balancing example



Original



Gray-world



Scale-by-max



Gray-edge



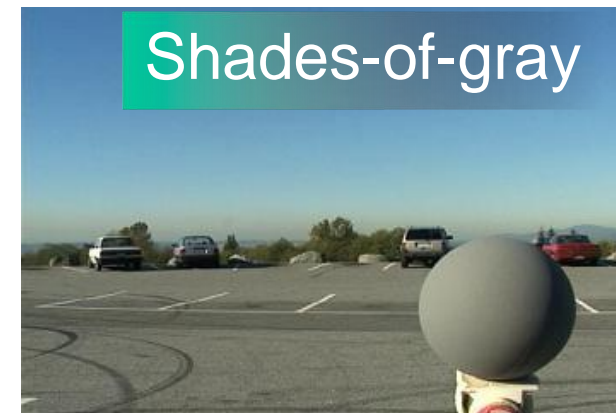
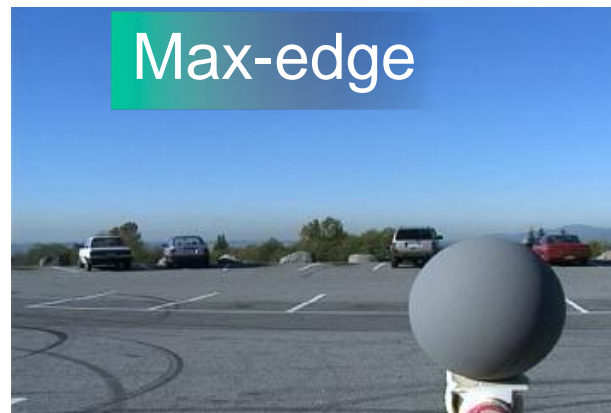
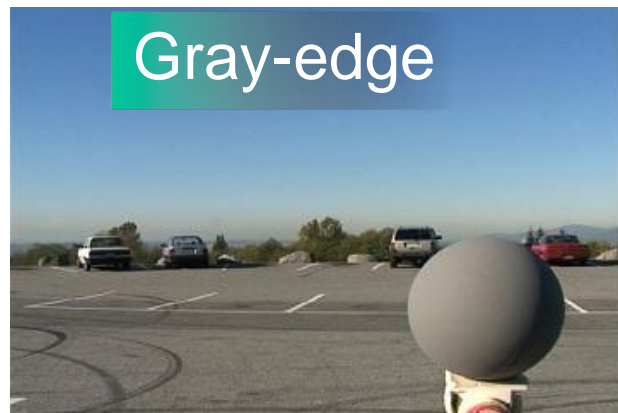
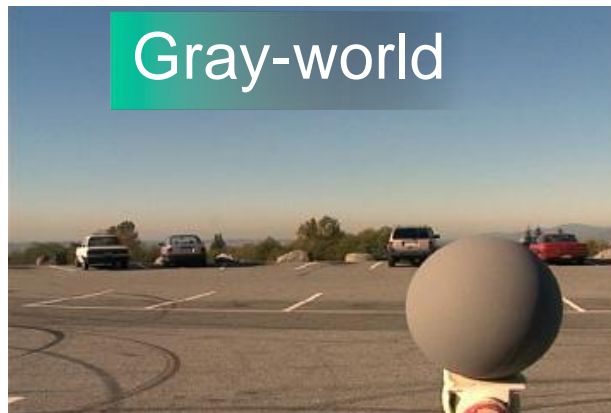
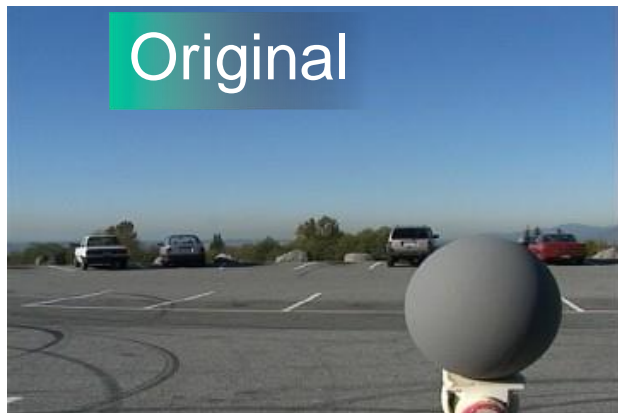
Max-edge



Shades-of-gray

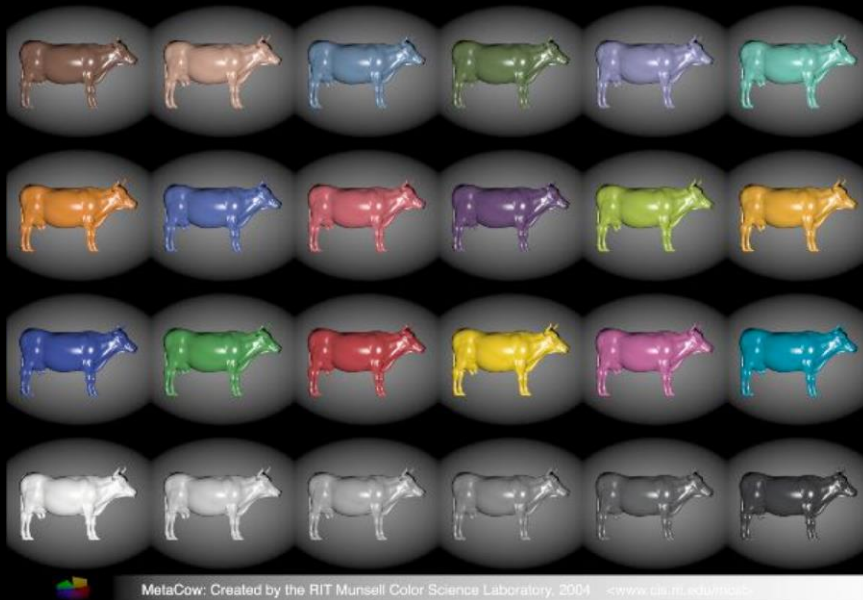


Color balancing example



Original image courtesy Ciurea and Funt

Daylight D65
CIE observer



Daylight D65
cheap camera



Illuminant A
CIE observer



Color conversion cheat sheet (e.g., for HW2)

- Great website for insights, every possible color conversion scheme, and much more:
www.brucelindbloom.com

- Spectrum to CIE XYZ:
(no illuminant)

$$X = \int_{\lambda} \bar{x}(\lambda) P(\lambda) d\lambda$$

$$Y = \int_{\lambda} \bar{y}(\lambda) P(\lambda) d\lambda$$

$$Z = \int_{\lambda} \bar{z}(\lambda) P(\lambda) d\lambda$$

CIE XYZ to CIE xyY:

$$x = \frac{X}{X + Y + Z}$$

$$y = \frac{Y}{X + Y + Z}$$

$$Y = Y$$

- CIE XYZ to CIE RGB:

$$\begin{bmatrix} R_{linear} \\ G_{linear} \\ B_{linear} \end{bmatrix} = M^{-1} \begin{bmatrix} X \\ Y \\ Z \end{bmatrix}$$

approximation of CIE gamma:

$$\{R, G, B\} = \left\{ R, G, B \right\}_{linear}^{1/\gamma}$$

- CIE RGB to CIE XYZ:

$$\begin{bmatrix} X \\ Y \\ Z \end{bmatrix} = M \begin{bmatrix} R_{linear} \\ G_{linear} \\ B_{linear} \end{bmatrix}$$

$$M = \begin{bmatrix} .490 & .310 & .200 \\ .177 & .813 & .011 \\ .000 & .010 & .990 \end{bmatrix}$$