

Clifford spinors and the relativistic point particle

Kaare Borchsenius *

Abstract

We examine the structure of the Clifford algebra associated with a Hermitian bilinear form and apply the result to a dynamical model of the relativistic point particle. The dynamics of the particle is described by a Dirac spinor with components in a Clifford algebra. This spinor determines, through the Clifford algebra, both the space-time coordinates and their conjugate momenta and satisfies a first order equation of motion which leads to the usual space-time canonical equations of motion. The constraints appear as the equations of motion for the einbein and spin connection which are needed to ensure the reparametrization invariance and local Lorentz invariance of the action.

*Bollerisvej 8, 3782 Klemensker, Denmark, e-mail: bdge@post5.tele.dk

1 Introduction

Spinors are usually associated with field theories for Fermions. Since, however, they are simply the ‘building blocks’ of the Lorentz group, there is no reason why other field theories could not, in principle, result from the dynamics of underlying spinors. In [1] we demonstrated an example of this kind in the simplest possible case, that of the relativistic point particle. We examined a dynamical model in which the quantized canonical space-time coordinates X of the particle were resolved into two-component spinors ψ with components in a Clifford algebra

$$X_{ab}^{A\dot{B}} = \{C_a^A, C_b^{*\dot{B}}\}, \quad \{C_a^A, C_b^B\} = 0, \quad X_{ab}^{A\dot{B}} \stackrel{\text{def}}{=} \sigma_\mu^{A\dot{B}} X_{ab}^\mu \quad (1)$$

where X_{ab}^μ transform under a Lorentz transformation in the index μ and under a unitary change of basis in Hilbert space in the indices a and b . The model also addressed the more general question as to whether space-time has a substructure which relates the complex properties of the Lorentz group and quantum mechanics. Spinor substructures of space-time have been discussed in Schwartz and Van Nieuwenhuizen [2] and in Borchsenius [3, 4, 5]

In this paper we examine the structure of the Clifford algebra which is associated with a Hermitian bilinear form, and apply the result to a dynamical model of the relativistic point particle. We find that the structure of the generating space, taken together with the need to accommodate parity, points to a Dirac spinor with Clifford components as the primary dynamical variable. This Dirac spinor determines through the Clifford algebra both the space-time coordinates and conjugate momenta of the particle and satisfies a first order equation of motion which generates the usual space-time canonical equations of motion. The fact that the momenta are no longer defined in terms of the coordinates and their derivatives means that the problem in constrained dynamics of eliminating the derivatives of the coordinates in favour of the momenta does not arise. The constraints appear instead as the equations of motion for the einbein and spin connection which are needed to ensure the reparametrization invariance and local Lorentz invariance of the action.

2 Clifford algebras and Hermitian bilinear forms

We recall that a real Clifford algebra arises naturally as the ‘square root’ of a real quadratic form Q on a real linear space V

$$v^2 = Q(v), \quad v \in V \quad (2)$$

where Q can have any signature s . In case Q is degenerate, the algebra contains Grassmann elements. On expanding v on an orthogonal basis e_i of V , it follows that (2) is satisfied if e_i satisfies the generating algebra

$$\frac{1}{2}\{e_i, e_j\} = \delta_{ij}Q(e_i) \quad (3)$$

To relate Hermitian bilinear forms on a complex linear space V to real Clifford algebras, we need to consider vectors in $V \oplus V$ which we shall write as

$\begin{pmatrix} u \\ v \end{pmatrix}, u, v \in V$. Then the following proposition holds:

Let B be a Hermitian bilinear form of signature s on an n -dimensional complex linear space V with complex involution $*$. Then there exists a $2n$ -dimensional Clifford algebra of signature $2s$ for $V \oplus V$ so that

$$B(u, v) = \left\{ \begin{pmatrix} u \\ 0 \end{pmatrix}, \begin{pmatrix} 0 \\ v^* \end{pmatrix} \right\}, \quad \left\{ \begin{pmatrix} u \\ 0 \end{pmatrix}, \begin{pmatrix} v \\ 0 \end{pmatrix} \right\} = 0 \quad (4)$$

Proof The involution $*$ of V induces the complex involution

$$\begin{pmatrix} u \\ v \end{pmatrix}^* \stackrel{\text{def}}{=} \begin{pmatrix} v^* \\ u^* \end{pmatrix} \quad (5)$$

of $V \oplus V$. Let f_i be an orthogonal normalized basis for V

$$B(f_i, f_j) = \delta_{ij} g_i, \quad g_i \in \{-1, 0, 1\} \quad (6)$$

Then $\begin{pmatrix} f_i \\ 0 \end{pmatrix}, \begin{pmatrix} 0 \\ f_i^* \end{pmatrix}$ is a basis for $V \oplus V$. It follows that the $2n$ -dimensional real subspace W of vectors $w = w^*, w \in V \oplus V$ has the basis

$$a_j = \frac{1}{2} \begin{pmatrix} f_j \\ 0 \end{pmatrix} + \frac{1}{2} \begin{pmatrix} 0 \\ f_j^* \end{pmatrix}, \quad b_j = -\frac{i}{2} \begin{pmatrix} f_j \\ 0 \end{pmatrix} + \frac{i}{2} \begin{pmatrix} 0 \\ f_j^* \end{pmatrix} \quad (7)$$

For this space there exists a $2n$ -dimensional real Clifford algebra of signature $2s$ with the generating algebra

$$\{a_i, a_j\} = \{b_i, b_j\} = \frac{1}{2} \delta_{ij} g_j, \quad \{a_i, b_j\} = 0 \quad (8)$$

Since a_i, b_i also form a basis for the whole of $V \oplus V$, the Clifford product can be extended from the real subspace W to the whole of $V \oplus V$. This permits us to write the generating algebra (8) in the equivalent form

$$\left\{ \begin{pmatrix} f_i \\ 0 \end{pmatrix}, \begin{pmatrix} 0 \\ f_j^* \end{pmatrix} \right\} = \delta_{ij} g_j, \quad \left\{ \begin{pmatrix} f_i \\ 0 \end{pmatrix}, \begin{pmatrix} f_j \\ 0 \end{pmatrix} \right\} = 0 \quad (9)$$

and (4) follows by expanding u and v on f_i, \square

In matrix language, this proposition implies that any $n \times n$ Hermitian matrix H_{ij} of signature s can be expressed in terms of the elements v_i of a $2n$ -dimensional Clifford algebra

$$H_{ij} = \{v_i, v_j^*\}, \quad \{v_i, v_j\} = 0 \quad (10)$$

Applying this to the Hermitian spinor components of the space-time coordinates

$$x^{A\dot{B}} \stackrel{\text{def}}{=} \sigma_\mu^{A\dot{B}} x^\mu \quad (11)$$

where σ^μ are the four Hermitian Pauli matrices, we obtain

$$x^{A\dot{B}} = \left\{ \begin{pmatrix} c^A \\ 0 \end{pmatrix}, \begin{pmatrix} 0 \\ c^{*\dot{B}} \end{pmatrix} \right\}, \quad \left\{ \begin{pmatrix} c^A \\ 0 \end{pmatrix}, \begin{pmatrix} c^B \\ o \end{pmatrix} \right\} = 0 \quad (12)$$

which was the basis for the discussion in [1].

3 Dirac spinors with Clifford components

In [1] we started out with the single Weyl spinor ψ^A in (12) as the primary dynamical variable of the point particle. There are two objections to this. First, the Clifford product is defined in $\mathbb{V} \oplus \mathbb{V}$; so there is no formal reason to restrict the components of the primary variables to \mathbb{V} . Secondly, there is no simple way of representing space and time inversion with only a single Weyl spinor. Both objections are met by taking instead as basic dynamical variable the Dirac spinor

$$\psi = \begin{pmatrix} \psi_l^A \\ \psi_{r\dot{B}} \end{pmatrix} \quad (13)$$

with left- and right-handed Weyl components ψ_l and ψ_r . For each value of (A, B) , ψ is an element of $\mathbb{V} \oplus \mathbb{V}$. Whereas the Weyl spinor in (12) determined only the space-time coordinates of the particle, the Dirac spinor (13) will, as we shall see, determine both the coordinates and their conjugate momenta.

To construct an action we must be able to form Lorentz invariant commutators between Dirac spinors. From the Clifford algebra it follows that

$$\left\{ \begin{pmatrix} u \\ 0 \end{pmatrix}, \begin{pmatrix} v \\ 0 \end{pmatrix} \right\} = \left\{ \begin{pmatrix} 0 \\ u \end{pmatrix}, \begin{pmatrix} 0 \\ v \end{pmatrix} \right\} = 0 \quad (14)$$

Hence, for the Dirac spinors $\psi = \begin{pmatrix} \psi_l \\ \psi_r \end{pmatrix}$ and $\chi = \begin{pmatrix} \chi_l \\ \chi_r \end{pmatrix}$ we obtain the commutator

$$\{\psi, \chi\} = \left\{ \begin{pmatrix} \psi_l \\ 0 \end{pmatrix}, \begin{pmatrix} 0 \\ \chi_r \end{pmatrix} \right\} + \left\{ \begin{pmatrix} 0 \\ \psi_r \end{pmatrix}, \begin{pmatrix} \chi_l \\ 0 \end{pmatrix} \right\} \quad (15)$$

We observe that the Clifford product couples the Weyl spinors in the opposite way of an ordinary contraction. It is therefore convenient to define the bracket

$$\{\psi, \chi\}_\sim \stackrel{\text{def}}{=} \{\psi, \tilde{\chi}\}, \quad \begin{pmatrix} \tilde{\chi}_l \\ \tilde{\chi}_r \end{pmatrix} \stackrel{\text{def}}{=} \begin{pmatrix} \chi_r \\ \chi_l \end{pmatrix} \quad (16)$$

The Dirac conjugate is defined in the usual way but with \sim taking the place of ordinary complex conjugation of the Weyl spinors

$$\bar{\psi} \stackrel{\text{def}}{=} (\psi_l^*, \psi_r^*) \gamma^0 \quad (17)$$

In the Weyl representation the Lorentz scalar corresponding to $\bar{\psi}\chi$ is therefore

$$\{\bar{\psi}, \chi\}_\sim = \left\{ \begin{pmatrix} \psi_{rB}^* \\ 0 \end{pmatrix}, \begin{pmatrix} 0 \\ \chi_l^B \end{pmatrix} \right\} + \left\{ \begin{pmatrix} 0 \\ \psi_l^{*A} \end{pmatrix}, \begin{pmatrix} \chi_{r\dot{A}} \\ 0 \end{pmatrix} \right\} \quad (18)$$

The basic Dirac spinor ψ determines the complex four-vector

$$a^\mu = \{\bar{\psi}_c, \gamma^\mu \psi\}_\sim, \quad \psi_c \stackrel{\text{def}}{=} C \bar{\psi}^T \quad (19)$$

which, since both ψ and its charge conjugate ψ_c transform like Dirac spinors, will transform like a vector under space and time inversion. So will its real and imaginary parts:

$$a^\mu = x^\mu + i p^\mu \quad (20)$$

$$x^\mu \stackrel{\text{def}}{=} \frac{1}{2} \{\bar{\psi}_c, \gamma^\mu \psi\}_\sim + \frac{1}{2} \{\bar{\psi}, \gamma^\mu \psi_c\}_\sim \quad (21)$$

$$p^\mu \stackrel{\text{def}}{=} -\frac{i}{2} \{\bar{\psi}_c, \gamma^\mu \psi\}_\sim + \frac{i}{2} \{\bar{\psi}, \gamma^\mu \psi_c\}_\sim \quad (22)$$

To show that x^μ and p^μ are independent vectors we express them in terms of the Weyl components of ψ using (19)

$$\sigma_\mu^{A\dot{B}}(x^\mu + ip^\mu) = 4i \left\{ \begin{pmatrix} \psi_l^A \\ 0 \end{pmatrix}, \begin{pmatrix} 0 \\ \psi_r^{\dot{B}} \end{pmatrix} \right\} \quad (23)$$

In [1] we showed that different eigenvalues of the quantized space-time coordinates \mathbf{X} corresponded to mutually anticommuting Weyl spinors. Since \mathbf{X} has a continuous spectrum it follows that any two four-vectors can be expressed in the form (12) with mutually anticommuting Weyl spinors. Using this property it follows immediately that (23) can be solved with respect to ψ_l and ψ_r for arbitrary values of \mathbf{x} and \mathbf{p} .

4 Equations of motion

The dynamics of the relativistic point particle will be described by the Dirac spinor ψ as a function of a parameter-time τ . We consider the action

$$\int \left\{ \bar{\psi}, i \left(\frac{d}{d\tau} + \Gamma(\tau) \right) \psi \right\}_\sim + c.c. - e(\tau) H(x, p) d\tau \quad (24)$$

which is invariant under a reparametrization $\tau \rightarrow \tau'$ of the world line with the einbein $e(\tau)$ and spin connection $\Gamma(\tau)$ transforming as

$$e'(\tau') = e(\tau) \frac{d\tau}{d\tau'}, \quad \Gamma'(\tau') = \Gamma(\tau) \frac{d\tau}{d\tau'} \quad (25)$$

and under a local Lorentz transformation $\psi \rightarrow S(\tau)\psi$ with the spin connection $\Gamma(\tau)$ transforming as

$$\Gamma(\tau) \rightarrow S\Gamma(\tau)S^{-1} - \frac{dS}{d\tau}S^{-1}, \quad \Gamma(\tau) = i\Gamma_{\mu\nu}(\tau)\sigma^{\mu\nu}, \quad \sigma^{\mu\nu} \stackrel{\text{def}}{=} \frac{i}{2}[\gamma^\mu, \gamma^\nu] \quad (26)$$

The equations of motion for $e(\tau)$ and $\Gamma(\tau)$ give the constraints

$$H(x, p) = 0 \quad (27)$$

$$\{\bar{\psi}, \sigma^{\mu\nu}\psi\}_\sim = 0 \quad (28)$$

Because the parameter-space is one-dimensional we can choose the gauge

$$e(\tau) = 1, \quad \Gamma(\tau) = 0 \quad (29)$$

The equations of motion for ψ are obtained by varying the action (26) with respect to $\bar{\psi}$. The contributions from \mathbf{x} and \mathbf{p} come from ψ and ψ_c in the second terms in (21) and (22), and we obtain

$$\frac{d\psi}{d\tau} = \frac{1}{2} \left(\frac{\partial H}{\partial p^\mu} - i \frac{\partial H}{\partial x^\mu} \right) \gamma^\mu \psi_c \quad (30)$$

When applied to (19) this gives the equations of motion for a^μ

$$\frac{da^\mu}{d\tau} = \left(\frac{\partial H}{\partial p^\nu} - i \frac{\partial H}{\partial x^\nu} \right) \{\bar{\psi}, \gamma^\nu \gamma^\mu \psi\}_\sim \quad (31)$$

which reduce to

$$\frac{da^\mu}{d\tau} = \left(\frac{\partial H}{\partial p_\mu} - i \frac{\partial H}{\partial x_\mu} \right) \{\bar{\psi}, \psi\}_\sim \quad (32)$$

by use of the constraint (28). For non-vanishing $\{\bar{\psi}, \psi\}_\sim$ we can reparametrize these equations with the new parameter $\bar{\tau}$

$$\frac{d\bar{\tau}}{d\tau} = \{\bar{\psi}, \psi\}_\sim \quad (33)$$

and split them into real and imaginary parts to obtain the canonical equations of motion

$$\frac{dx^\mu}{d\bar{\tau}} = \frac{\partial H}{\partial p_\mu}, \quad \frac{dp^\mu}{d\bar{\tau}} = -\frac{\partial H}{\partial x_\mu} \quad (34)$$

This leads to the interpretation of \bar{x} , \bar{p} and $\bar{\tau}$ as the coordinates, conjugate momenta and proper-time of the particle.

5 The free particle

The free particle is described by the Hamiltonian

$$H(x, p) = \frac{1}{2m} (p_\mu p^\mu - m^2) \quad (35)$$

with the constraint

$$p_\mu p^\mu = m^2 \quad (36)$$

The equations of motion (30) become

$$\frac{d\psi}{d\tau} = \frac{p_\mu}{2m} \gamma^\mu \psi_c \quad (37)$$

From (32) it follows that \bar{p} is constant. Using the constraint (36) we obtain the complete solution to (37)

$$\psi = \alpha e^{\frac{1}{2}\bar{\tau}} + \beta e^{-\frac{1}{2}\bar{\tau}} \quad (38)$$

where the Dirac spinors α and β are constants of integration which satisfy

$$\alpha = \frac{p_\mu}{m} \gamma^\mu \alpha_c, \quad \beta = -\frac{p_\mu}{m} \gamma^\mu \beta_c \quad (39)$$

From (39) it follows that

$$\{\bar{\alpha}, \beta\}_\sim = -\{\bar{\beta}, \alpha\}_\sim \quad (40)$$

so that $\{\bar{\alpha}, \beta\}_\sim$ is imaginary. The constraint (28) gives

$$\{\bar{\alpha}, \sigma^{\mu\nu} \alpha\}_\sim = \{\bar{\beta}, \sigma^{\mu\nu} \beta\}_\sim = 0 \quad (41)$$

$$\{\bar{\alpha}, \sigma^{\mu\nu} \beta\}_\sim + \{\bar{\beta}, \sigma^{\mu\nu} \alpha\}_\sim = 0 \quad (42)$$

a^μ is obtained by inserting the expression (38) for ψ into (19)

$$\begin{aligned} a^\mu = & \frac{p_\nu}{m} (\{\bar{\alpha}, \gamma^\nu \gamma^\mu \alpha\}_\sim e^\tau - \{\bar{\beta}, \gamma^\nu \gamma^\mu \beta\}_\sim e^{-\tau}) \\ & + \frac{p_\nu}{m} (\{\bar{\alpha}, \gamma^\nu \gamma^\mu \beta\}_\sim - \{\bar{\beta}, \gamma^\nu \gamma^\mu \alpha\}_\sim) \end{aligned} \quad (43)$$

To obtain the coordinates and momenta we split a^μ into its real and imaginary parts by applying conditions (40), (41) and (42)

$$p^\mu = -\frac{2i}{m}\{\bar{\alpha}, \beta\}_\sim p^\mu \quad (44)$$

$$x^\mu = \frac{p^\mu}{m} (\{\bar{\alpha}, \alpha\}_\sim e^\tau - \{\bar{\beta}, \beta\}_\sim e^{-\tau}) - 2i\{\bar{\alpha}, \sigma^{\nu\mu}\beta\}_\sim p_\nu \quad (45)$$

(44) simply adjusts the value of $\{\bar{\alpha}, \beta\}_\sim$ to be consistent with the constant value of p^μ . The proper-time is found by integrating (33)

$$\frac{d\bar{\tau}}{d\tau} = \{\bar{\psi}, \psi\}_\sim = \{\bar{\alpha}, \alpha\}_\sim e^\tau + \{\bar{\beta}, \beta\}_\sim e^{-\tau} \quad (46)$$

to give

$$\bar{\tau} = \{\bar{\alpha}, \alpha\}_\sim e^\tau - \{\bar{\beta}, \beta\}_\sim e^{-\tau} + \tau_0 \quad (47)$$

The expression for τ in terms of proper-time is obtained from (45) and (47):

$$x^\mu(\bar{\tau}) = \frac{p^\mu}{m}\bar{\tau} + x^\mu(0), \quad x^\mu(0) \stackrel{\text{def}}{=} -\frac{p^\mu}{m}\tau_0 - 2i\{\bar{\alpha}, \sigma^{\nu\mu}\beta\}_\sim p_\nu \quad (48)$$

The solutions fall into two classes depending on the relative signs of $\{\bar{\alpha}, \alpha\}_\sim$ and $\{\bar{\beta}, \beta\}_\sim$. When they have the same signs a complete solution to the equations of motion for ψ generates a complete solution to the space-time equations of motion. When they have opposite signs the space-time solution has an endpoint in proper-time and the complete solutions to the equations of motion for ψ are double coverings of incomplete solutions to the space-time equations. Before rejecting this last type of solution as unphysical the possibility should be kept in mind that in a realistic theory a common starting point in proper-time could have a cosmological interpretation. The double covering property also provides a simple local interpretation of the quantum interference between alternative space-time paths [1].

6 Conclusion

The model of the relativistic point particle discussed in the foregoing suggests that the concept of canonically conjugate variables has an algebraic origin in the relationship between Hermitian bilinear forms and Clifford algebras. The $V \oplus V$ structure of the generating space of the Clifford algebra lends itself naturally to a representation of parity in terms of a Dirac spinor with Clifford components. From this spinor the space-time coordinates and momenta emerge in a unified manner, and the problem in constrained dynamics of eliminating the derivatives of the coordinates in favour of the momenta does not arise.

It is customary to obtain a many-particle theory by applying the coordinate representation to the constraints. However, since the space-time coordinates are no longer primary variables, such a procedure would only be an approximation. Another option which naturally suggests itself in the present context is to increase the dimension of the parameter-space by adding a world-sheet spinor index to the world spinor index of ψ . It remains to be seen which of the features of the point particle model discussed in the foregoing can be carried over into such a higher dimensional theory.

References

- [1] Borchsenius, K. (2000). *Math. Phys. EJ*, Vol 6, No.4.
- [2] Schwartz, J.H. and Van Nieuwenhuizen, P. (1982). *Lettere al Nuovo Cimento*, 34, 21.
- [3] Borchsenius, K. (1987). *General Relativity and Gravitation*, 19, 643.
- [4] Borchsenius, K. (1989). *General Relativity and Gravitation*, 21, 959.
- [5] Borchsenius, K. (1995). *International Journal of Theoretical Physics*,r 34, 1863.