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#####
# Global parameters
#####
# Virial velocity of the galaxy [km/s]
v200          150.0
# Virial mass of the galaxy [1e10 Msol]
# Overrides the v200 parameter
m200          107.4
# Halo spin parameter
lambda         0.04
# Refinement level of the potential grid
level_coarse   7
# Refinement level of the plane plane density grid
level_grid_mid_dens 7
# Refinement level of the turbulence grid
level_grid_turb 7
# Refinement level of the gaussian field grid
level_grid_dens_fluct 7
# Size of the potential grid [kpc]
boxsize1       100.0
boxsize2       14.0
boxsize3       5.0
# Dispersion for the Gaussian field fluctuations
dens_fluct_sigma 0.50
# Physical injection scale of the random field fluctuations [kpc]
dens_fluct_scale_inj 2.00
# Physical dissipation scale of the random field fluctuations [kpc]
dens_fluct_scale_diss 0.25
dens_fluct_seed 1212
# Seed for the random number generator
seed           1246
# Switch to MCMC ntry algorithm to position particles for a value > 1
mcmc_ntry     1
# Number of iterations to reach hydrostatic equilibrium (zero to deactivate)
hydro_eq_niter 3

#####
# Components parameters
#####

#####
# Component 1: Halo
#####
# Fraction of the virial mass in the component 1
mass_frac1    0.9636819104
# Number of particles for the component 1
npart1        100000
# Target mass of individual particles [Msol]
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part_mass1           3.564e8
# Number of particles for the potential computation
#npart_pot1         200000
# Target mass of individual particles for potential computation [Msol]
#part_mass_pot1     1e6
# Component 1 density model
# Available models:
# 1 = Exponential disk + sech-z profile
# 2 = Myamoto-Nagai profile
# 3 = Exponential disk + exponential-z profile
# 4 = Hernquist profile
# 5 = Plummer profile
# 6 = Jaffe profile
# 7 = Isothermal profile
# 8 = NFW profile
# 9 = Burkert
# 10 = Einasto profile
# 11 = Mestel profile
# 12 = Kalnajs profile
# 13 = Sersic profile
# 14 = Toomre-Kuzmin profile
# 15 = Uniform profile
# 16 = Pseudo-isothermal profile
model1              8
# Scale length of the density profile [kpc]
scale_length1        1.
# Core radius in the density profile [kpc]
rcore1               0.
# Concentration parameter
# if positive, the scale length is recomputed to match the concentration
# the concentration parameter of an NFW halo with a mass M200*mass_frac
concentration1       10.0
# Gravitational softening for the poisson solver [kpc]
softening1           0.6
# Component 1 radial density cut [kpc]
cut1                 0.
# Component 1 thickness parameters
flatx1               1.00
flaty1               1.00
flatz1               1.00
# Gaussian step for the MCMC Metropolis-Hasting particle positionning algorithm,
# expressed in units of the component scale length. Default value is 0.5.
mcmc_step1            0.50
# Maximum velocity for the component 1 particles in expressed in units of escape velocity
vmax_esc1             5.0
# Particles type (GADGET format -- 0=Gas,1=Halo,2=Disk,3=Bulge,4=Stars)
type1                 1
# Streaming fraction of the component 1
stream_fraction1      0.00

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# Radius at which the density profile should have the exact same
# value as the NFW profile with the previously defined concentration
# and with a mass M200*mass_frac [kpc]
radius_nfw1      1.0
# Cut the density function in the Jeans equation integration
jeans_mass_cut1      1
# Compute component velocity (debug option)
compute_vel1      1
# Number of integral of motion for the Jeans equations
# 0 = Do not use Jeans equation for sigma_r and sigma_z
# 1 = Spherically symmetric Jeans equation
# 2 = Jeans equations with 2 integrals of motion
# 3 = Jeans equations with 3 integrals of motion (solved on a 2D grid in the r-z plane)
jeans_dim1      1
# Method for streaming velocity computation
# 0 =      - User defined fixed fraction of the circular velocity profile (stream_fraction keyword)
# 1 = Bullock 2001 - Streaming velocity profile following the cumulative mass profile
# 2 = Springel 1999 - Streaming velocity profile following is a fixed fraction of the rotation curve
# 3 =      - Solid Body rotation
stream_method1      1
# Alpha coefficient for the generalized normal distribution [Gaussian=2.0]
# for the random generation of velocities
ggd_beta1      2.0
# Minimum acceptance for the MCMC chain
accept_min1      0.80
# Maximum acceptance for the MCMC chain
accept_max1      0.95

#####
# Component 2: Thin stellar disk
#####
mass_frac2      0.02641287
npart2      100000
#npart_pot2      200000
part_mass2      3.437e5
model2      3
# If the value is zero and the particle type is not 1 the size is determined using the spin conservation
# (Fitting formula from Mo, Mao & White 1998)
scale_length2      3.432
cut2      4.5
flatz2      0.15
mcmc_step2      0.3
type2      2
stream_fraction2      1.00
# Epicycle approximation in the Jeans equations
epicycle2      1
# Minimal value for the Toomre parameter
Q_lim2      1.25
# Fixed value for the Toomre parameter

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Q_fixed2          0.0
# Additional term for the Toomre parameter
Q_boost2          0.0
# Past constant Star Formation Rate [Msol/yr]
# Negative value recomptes the SFR according to Bouché et al. 2010
SFR2              -1.
# Minimum age of the stars [Myr]
min_age2          0.
# Metallicity of the stars [Zsolar]
metal2             0.01
# Gaussian fluctuations in the density field
dens_fluct2        0
# Use the density cut during the Jeans equation integration
jeans_mass_cut2   1
compute_vel2       1

#####
# Component 3: Gaseous disk
#####
mass_frac3         0.0066036034
npart3             100000
#npart_pot3        200000
part_mass3         8.593e4
model3              3
scale_length3      3.432
cut3                4.5
flatz3              0.25
mcmc_step3         0.15
type3               0
stream_fraction3   1.00
metal3              0.01
# Metallicity follows density gradient
metal_gradient3    1
# Temperature of the gas particles [K]
t_init3            1e5
# Turbulent velocity dispersion [km/s]
turb_sigma3         0.0
# Turbulence injection scale [kpc]
turb_scale_inj3    1.0
# Turbulence dissipation scale [kpc]
turb_scale_diss3   0.01
# Seed for the turbulent gaussian field
turb_seed3          1234
# Compute hydrostatic equilibrium
hydro_eq3           1
# Gaussian step for the MCMC Metropolis-Hasting particle postionning algorithm
# within the hydro equilibrium algorithm
mcmc_step_hydro3   0.3
compute_vel3         1

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# Polytropic index for the gas (1 = isothermal)
gamma_poly3          1.0

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# Component 4: Stellar bulge
#####
mass_frac4           0.0033016095
npart4               12500
#npart_pot4          25000
part_mass4            3.437e5
model4                4
scale_length4          2.0
cut4                  4.0
flatz4                0.80
mcmc_step4             0.3
vmax_esc4              2.0
type4                  3
stream_fraction4        0.5
min_age4               50.
metal4                 0.001
# Structural parameter specific to the Einasto density profile
alpha_struct5          0.80
compute_vel4             1

#####
# Component 5: Stellar spheroid
#####
mass_frac5             0.0
npart5                 0
npart_pot5              0
model5                  4
scale_length5            2.0
cut5                   6.0
flatz5                 0.40
mcmc_step5              0.3
vmax_esc5                2.0
type5                   2
stream_fraction5         1.00
min_age5                50.
metal5                  0.001
Q_lim5                  1.5
compute_vel5              1

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