Point source approximation



We can approximate an earthquake as being a point source when looking at

- wavelengths 2 >> & fault surface, or - period T >> to suppose time

Moment tensor: For paint sources, we can introduce 2 symmetrie moment tensor M such that

 $M = \int_{0}^{t_{final}} \int_{t} \int_{\Sigma} dx dt = \int_{\Sigma} \int_{t_{inal}} d^{3}x$ $= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} dx^2$

is related to the final stress glut still and surface mement-density tensor mind For planar faults and isotropic media, the moment tensor M scales with the final slip offset on the fault: double couples

 $M = \int u \Delta u^{\text{final}} \left(\frac{\lambda \hat{b} + \hat{b} \hat{v}}{\lambda \hat{b} + \hat{b} \hat{v}} \right) d^{2}x$

 $M_o = \int_{\mathcal{L}} u \, du \, final \, dx \qquad scalar \\ = \sum_{moment} u \, dx \qquad moment$

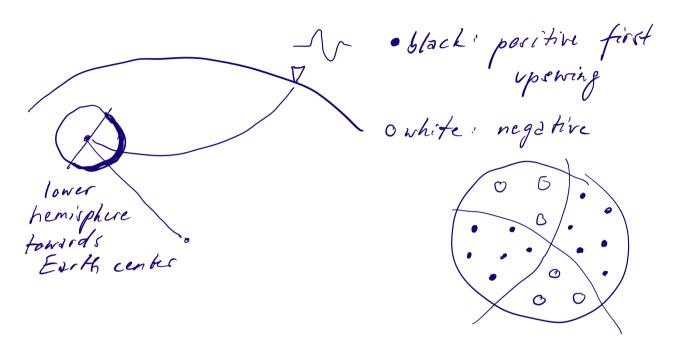
The scalar moment $M_0 = \frac{1}{\sqrt{2}} (M:M)^2$ has become a standard measure of the size of an earthquake.

For point sauces, we can now write the earthquake in terms of equivalent body force $f = -M \cdot \nabla_x \delta(x - x_s) H(t - t_s)$ moment tensor

Heariside skp function

To separate volumetrie and deviatorie effects: $\underline{\underline{M}} = \frac{1}{3}(\underline{+}\underline{\underline{M}}) \underline{\underline{I}} + \underline{\underline{M}}^{\text{dev}}$ =0 for hydrostatic M = Miso + M devisotropic deviaterie where $\underline{M}^{dev} = \underline{M}^{DC} + \underline{M}^{CLVD}$ double compensated couple linear vector dipole rearthquake anagma dyke 1 magma dyke intrusion

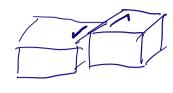
Beach ball representation: Depicts the radiation pattern of an earthquake



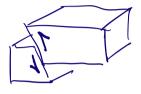
We can always find 2 planes as solution to double couples (true & auxiliary), which leads to the fault plane ambiguity.



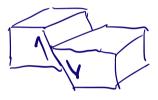












Volterra's theorem: Writes displacement as convolution

in time of moment tensor Mand Green's functions G eonvolvtion $u_n(x,t) = \int u_i(S,t) c_{ijkl} \hat{v}_i * \partial_k G_{kn} d^2S$ $= \int m_{kl} * \partial_k G_{ln} d^2S$ $= M_{kl} * G_{kn}, k$ $M_{kl} * G_{kn}, k$

Source-time function: What is the time dependence?

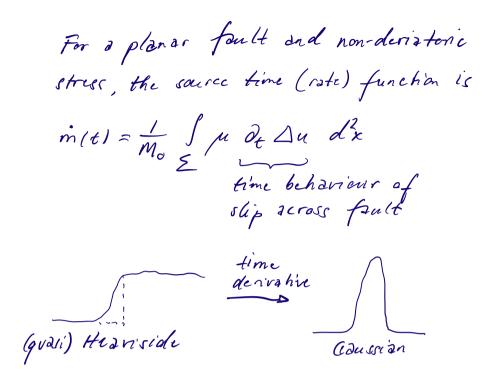
stress glute rak $S = M \delta(x-x_s)$ moment-rate tensor $M = \int_{t}^{t} S dx$ E

assuming synchronuous in all components

of M, we find normalized source time function $M(t) = \sqrt{2} M_0 M m(t)$ sealer moment directions

a size a radiation

pottern



Analythe solution: For an elastic, homogeneous, whole space,
the far-field compressional displacement
writes as [Akib Richards, 2002]

$$u_r(x,t) = \frac{1}{4\pi g \, v_p^3 \, r} \qquad \dot{m}(t - \frac{r}{v_p}) \quad \sin 2\theta \cos \theta$$
radial
$$zmplitude \qquad temperal \qquad radiation \\ pulse \qquad pattern$$

with position x = (r, 8, 4) spherical coord. the shear displacement writes as $u_{09} = \frac{1}{478 \, v_{s}^{3} \, r} \, \dot{M} \left(t - \frac{r}{v_{s}} \right) \cos 2 \vartheta \cos 4$ $u_{4} = \frac{1}{478 \, v_{c}^{3} \, r} \, \dot{M} \left(t - \frac{r}{v_{s}} \right) \left(-\cos 2 \vartheta \sin 4 \right)$ Note that for these far-field displacements:

- · + facter due to geometrical spreading
- · shear displacement has no nodal planes, therefore P-waves are used for beachballs
- higher amplitudes in slow media (sediments) and S-waves due to $\frac{1}{v_p^3} \operatorname{resp.} \frac{1}{v_s^3} \operatorname{factor}$
- $\left(\frac{Vp}{Vs}\right)^3 \approx 5$ since for a Parison

 solid $Vp = \sqrt{3} V_S$ -> S-wsves are about $5 \times stronger$