Comparing Classical and Nonclassical Symmetries of Nonlinear Partial Differential Equations

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Project Goal

Our research objective for this project was to calculate the classical and nonclassical symmetry groups for the reduced Gibbons-Tsarev equation and the Born-Infeld equation and compare them.

Definition

A symmetry is a transformation that leaves an object invariant.

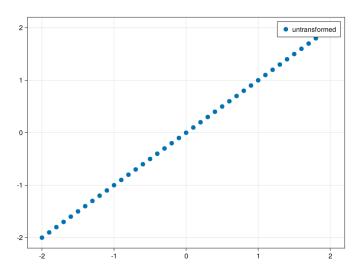
Definition

A symmetry is a transformation that leaves an object invariant.

Definition

A symmetry is a change that doesn't change anything.

Let's see this in action using the simple linear equation x - y = 0.



Example (A Non-Example)

- For our first transformation, let's define new variables $\bar{x} = x + 1$ and $\bar{y} = y$.
- Now we rewrite our equation using these new variables.

$$ar x-ar y=0$$
 by definition $x+1-y=0$ by substitution $y=x+1$ by rewriting in slope-intercept form

► This transformation is not a symmetry:

$$x - y + 1 \neq x - y$$

A Transformation that is a Symmetry

Example (2)

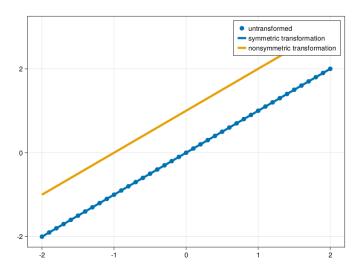
- Let's define some new variables again $\bar{x} = x + 1$ and $\bar{y} = y + 1$.
- Now we rewrite our equation using these new variables.

$$\begin{array}{c} \bar{x}-\bar{y}=0 & \text{by definition} \\ (x+1)-(y+1)=0 & \text{by substitution} \\ (x-y)+(1-1)=0 & \text{by algebra} \\ x-y=0 & \text{by algebra} \\ y=x & \text{by rewriting in slope-intercept form} \end{array}$$

► This transformation is a symmetry:

$$x - y = x - y$$

The graphs of our three equations.



Who cares?

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- Symmetries encode physically meaningful aspects of equations, like conservation laws in physics.

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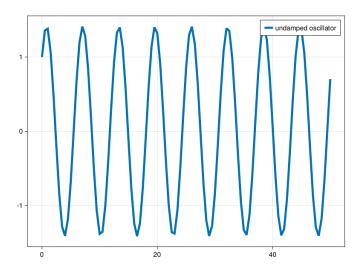
- ▶ A differential equation is an equation that contains both an unknown function and information about how that function relates to its rates of change.
- Differential equations show up everywhere we model something using information about how that thing changes. This includes everything from population dynamics to planetary orbits.
- Differential equations are different from algebraic equations, and can't be solved in the same ways.

As an example, we'll use the equation for an undamped spring.

$$m\ddot{x} = -\kappa x$$

- ▶ Here m is the mass, \ddot{x} is the acceleration, κ is the spring constant, and x is the position.
- ▶ If we set both m and κ to 1, we get $\ddot{x} = -x$.
- This is a simple differential equation with the algebraic solution of $x(t) = c_1 \cos(t) + c_2 \sin(t)$.

The graph of our spring system $x(t) = \cos(t) + \sin(t)$, where $c_1 = c_2 = 1$.



The History of the Born-Infeld and the reduced Gibbons-Tsarev Equations

- The goal of our project was to find symmetry transformations for these two PDEs, both classical and nonclassical, and compare them.
- It turns out that there's a standard method for calculating both kinds of symmetry transformations for differential equations.
 - Input Equation \rightarrow Calculate Lie's invariance condition
 - \rightarrow Solve linear system of PDEs
 - \rightarrow Symmetries!

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- ► For a differential equation, this means that the function and any derivatives of that function that are present in the equation must not change under the transformation.
- Lie's invariance condition gives us a way to generate transformations that meet all these requirements.
- ► If you can solve Lie's invariance condition for your equation, you can find its symmetries.

The general form of Lie's invariance condition is the falsely unassuming

$$\Gamma^{(n)}\Delta|_{\Delta=0}=0$$

Here Γ is the infinitesimal operator, n is the order of the equation, and Δ is the equation itself set to equal 0 (this is standard form).

For our equations, which are both second order PDEs in two independent variables, Lie's invariance condition looks like the following:

. . .

► The result of putting our PDEs through Lie's invariance condition was a system of ... linear PDEs for the Born-Infeld equation and ... for the reduced Gibbons-Tsarev equation.

- ► The result of putting our PDEs through Lie's invariance condition was a system of ... linear PDEs for the Born-Infeld equation and ... for the reduced Gibbons-Tsarev equation.
- ➤ Though difficult, these systems were easier to solve than the original single equations because of their linearity. This resulted in the following symmetries.

Classical Symmetries of the Born-Infeld Equation

- **.**

Classical Symmetries of the reduced Gibbons-Tsarev Equation

- **.**

Future Work: Does Integrability Imply Equivalence of Classical and Nonclassical Symmetries?

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