

# **OPTICAL SATELLITE COMMUNICATIONS**

## GUEST LECTURE – COMMUNICATION THEORY

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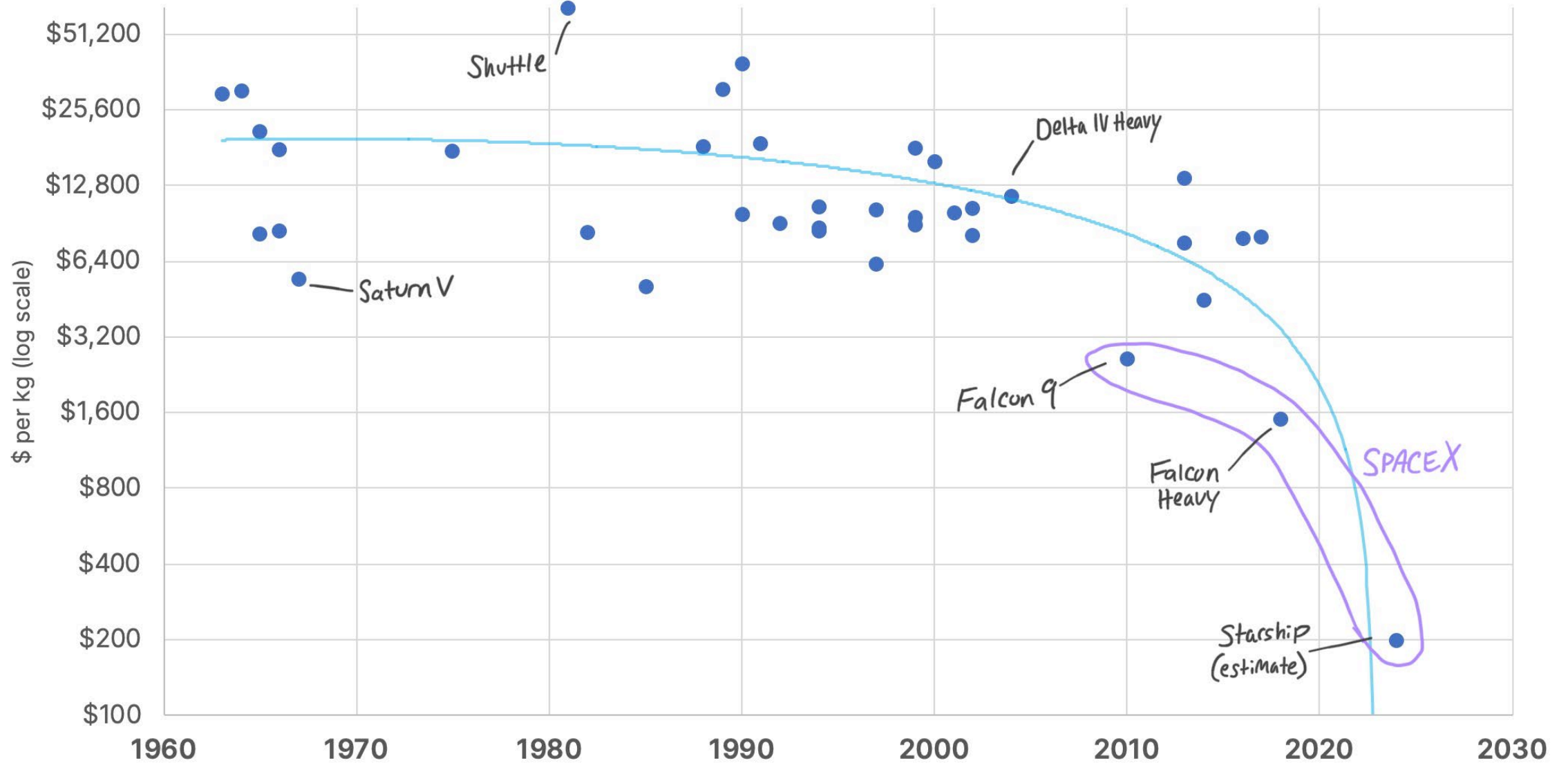
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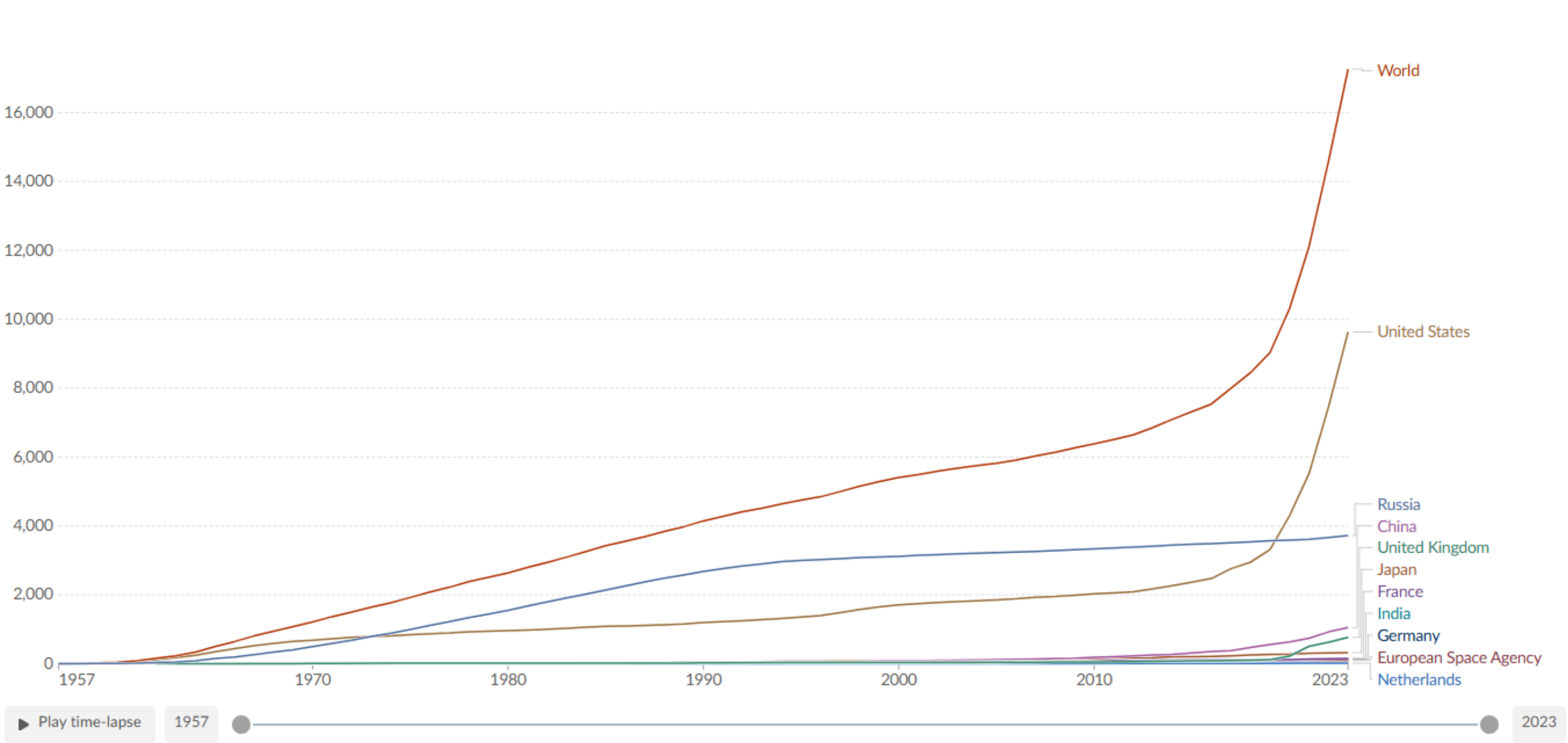






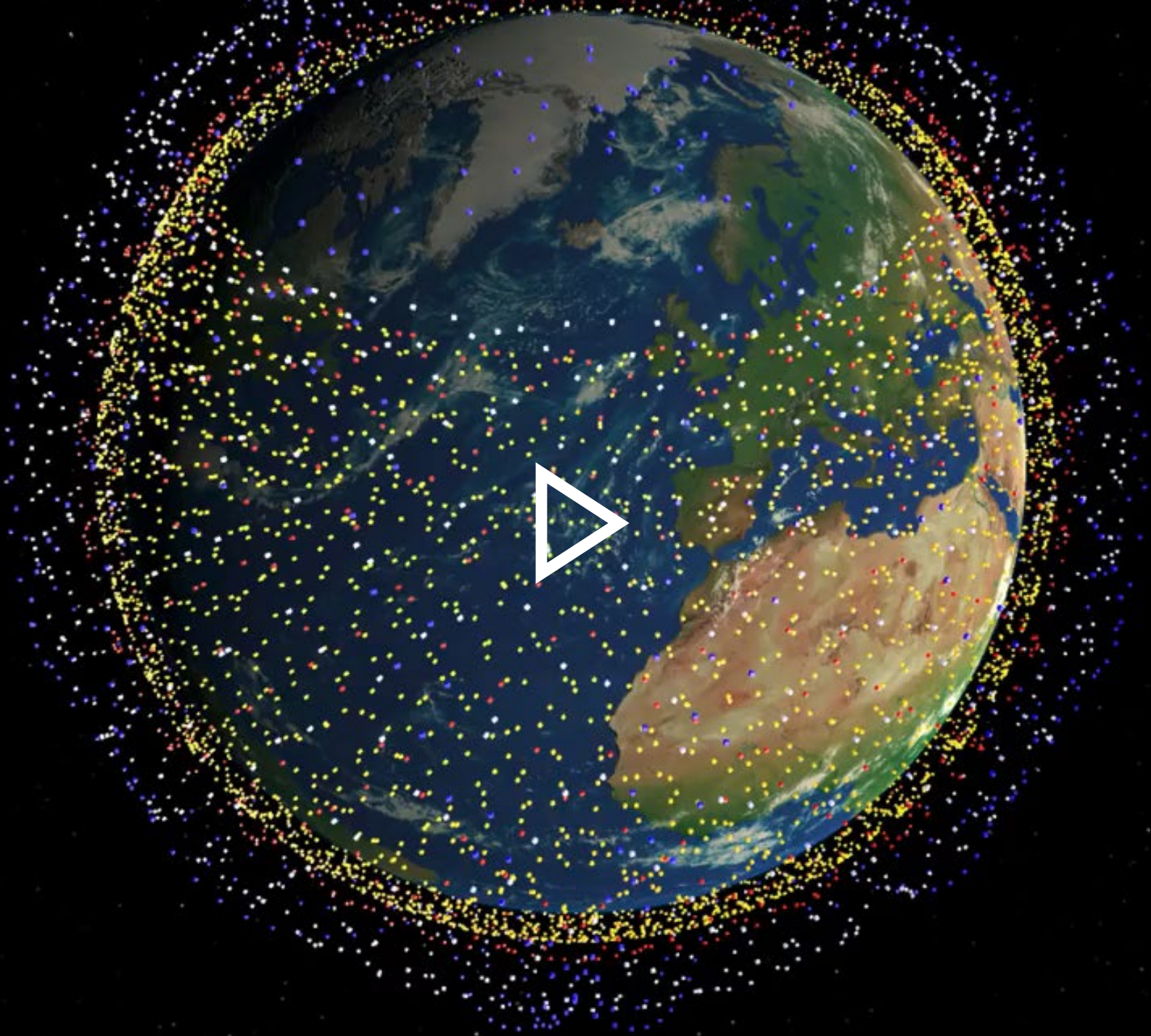
## Launch Cost per Kilogram (medium and heavy)





Data source: United Nations Office for Outer Space Affairs (2024) – [Learn more about this data](#)





# SPECTRUM

FM



5G LOW BAND



5G MID BAND



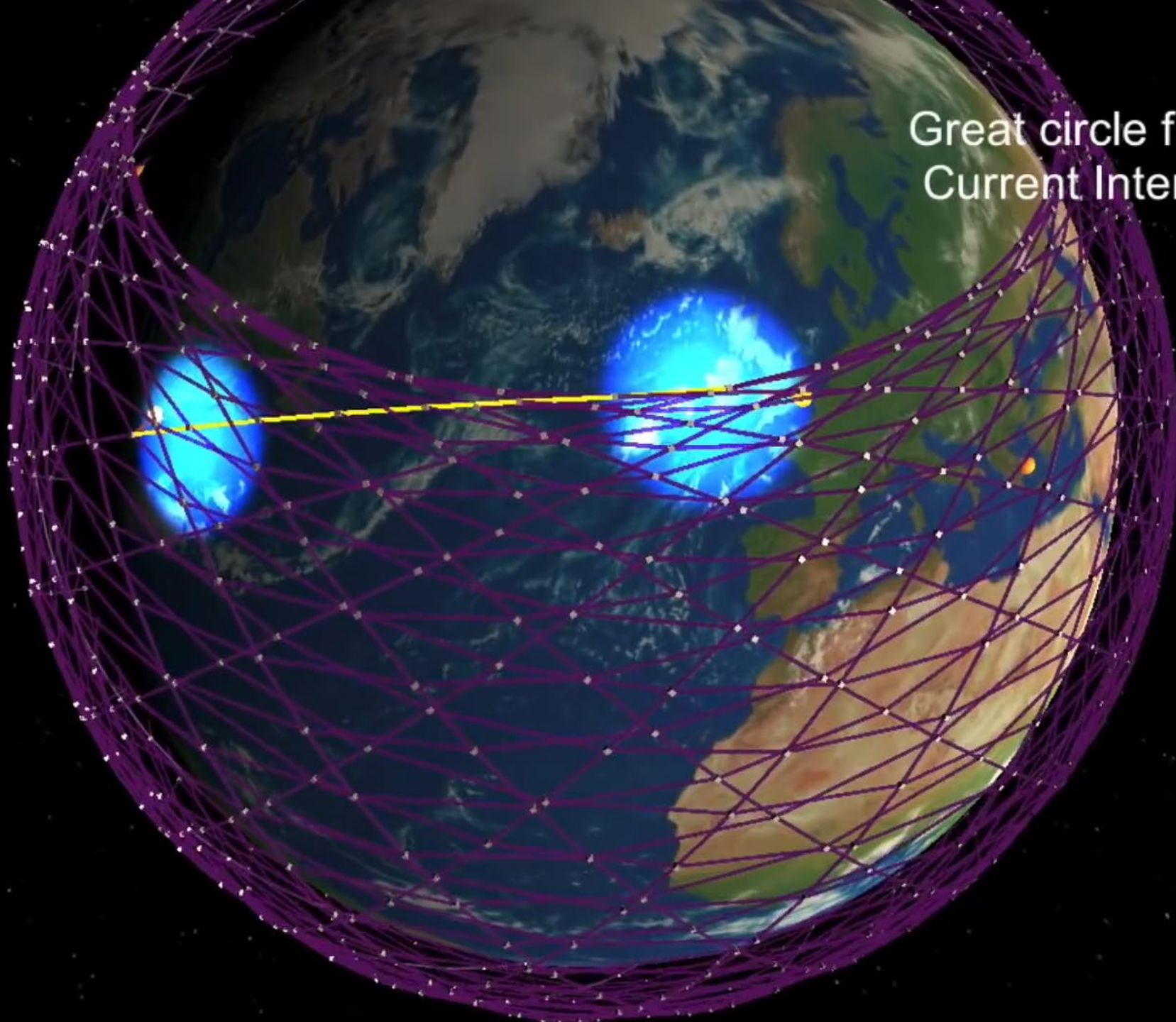
5G MMWAVE BAND



INFRARED C-BAND







Sat dist: 7009  
Sat RTT: 46ms  
Great circle fibre RTT: 55ms  
Current Internet RTT: 76ms



## FSO VERSUS RF – LINK BUDGET

- › Receive power:

$$P_R = \eta P_T \cdot \left( \frac{\pi D_T}{\lambda} \right)^2 \left( \frac{\pi D_R}{\lambda} \right)^2 \left( \frac{\lambda}{4\pi R} \right)^2$$

- › which is proportional to  $\frac{1}{\lambda^2}$  and  $f^2$ .

- › Example: Ka-band versus FSO

- › with RF Ka-band (30 GHz) versus FSO C-band (193 THz)
- › this gives about 75 dB gain in receive power for FSO over RF
- › assuming antenna sizes  $D_T$  and  $D_R$ , transmit power  $P_T$ , distance  $R$  all the same.

### Example #3: 100 Mbps link from Mars to Earth

Parameter	RF Ka-band	FSO C-band	units
Wavelength $\lambda$	$10^7$	1550	nm
Frequency $c/\lambda$	30	193 550	GHz
Transmit power $P_T$	500	50	W
Transmitter diameter $D_T$	5.6	0.5	m
Receiver diameter $D_R$	4x 34	6x 2.5	m

By NASA B. L. Edwards, D. Antsos, A. Biswas, and others, "An envisioned future for space optical communications," 2023 International Conference on Space Optical Systems and Applications (ICSOS), 2023.

## FSO VERSUS RF - RISK OF DETECTION & INTERCEPTION

Thanks to the high transmitter gains and high directivity  
the FSO beams have a low probability of detection (LPD) and interception (LPI)

Diffraction-limited spot size (in far field, assuming perfect optics and antennas):

$$w(R) = w_0 \sqrt{1 + \left( \frac{\lambda R}{\pi w_0^2} \right)^2} \approx \frac{\lambda R}{\pi w_0}$$

Example: **Eavesdropping a LEO satellite on the ground**

with RF Ka-band (30 GHz) versus FSO C-band (193 THz), distance of 500 km

and a transmit beam waist radius  $w_0$  of 10 cm gives +/- 15900 m for Ka-band versus 2.5 m for FSO link.



## › THE PROMISES OF FSO

- › High-throughput
- › High-security
- › Low-interference
- › Low energy-per-bit
- › License-free

## › THE KEY CHALLENGES

### SOME KEY FACTORS

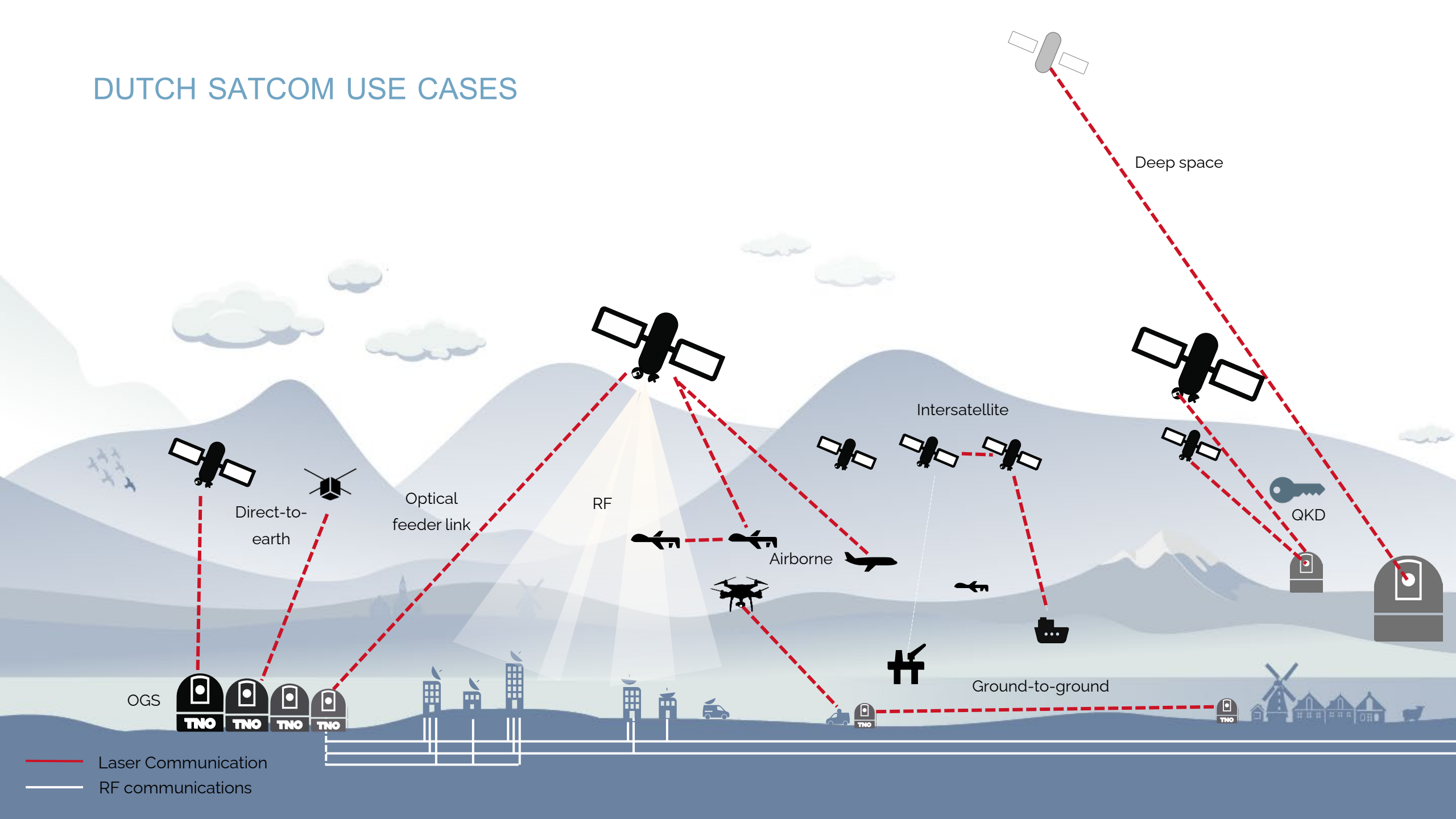
- **Availability** – limited by weather conditions
- **Locations** – direct line of sight
- **Eye safety** – limit on transmit power

### TECHNOLOGICAL CHALLENGES

- Acquisition, tracking & pointing
- Atmospheric turbulence & robustness
- Modem & DSP design for Gbps/Tbps links



# DUTCH SATCOM USE CASES



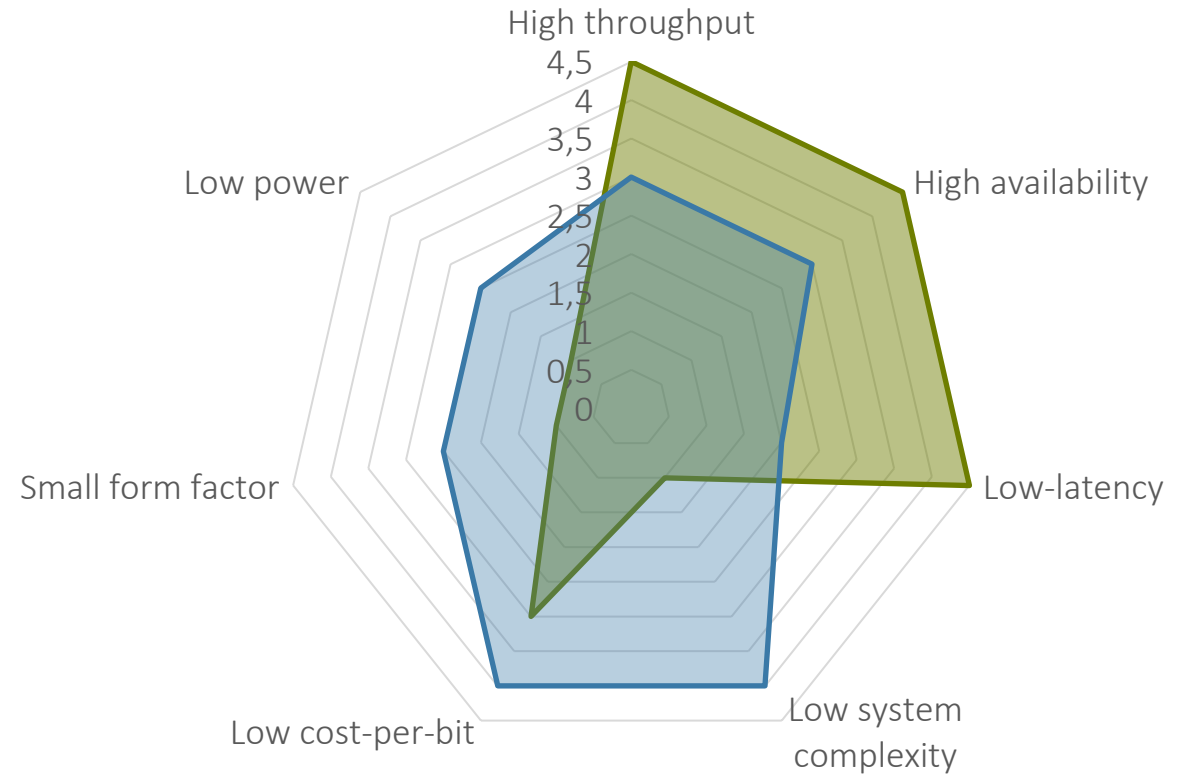
## › DEFINE KEY PERFORMANCE GOALS

### TYPICAL OBJECTIVES:

- 1. THROUGHPUT
- 2. RELIABILITY
- 3. HIGH-AVAILABILITY
- 4. LOW-LATENCY

### WHILE MEETING/MINIMIZING THE REQUIRED:

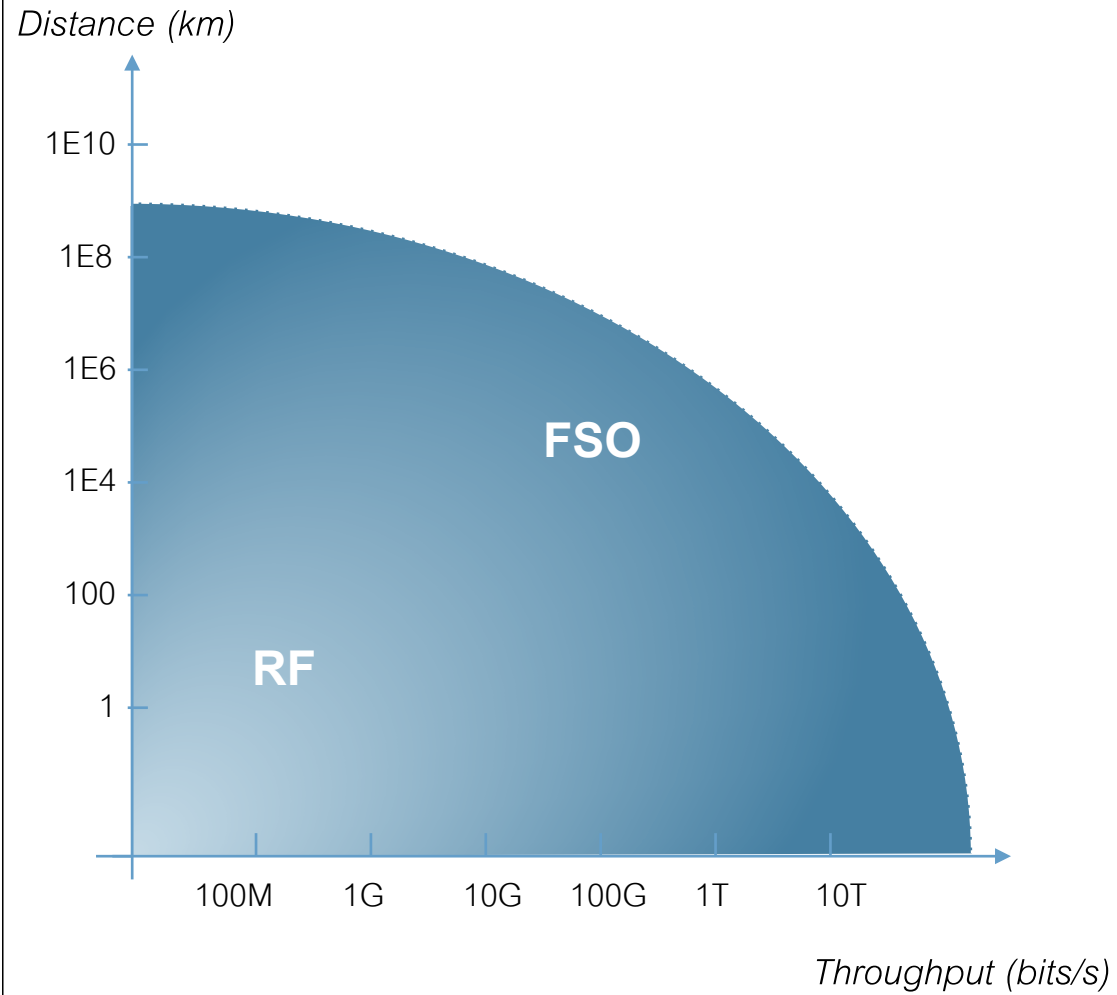
- COST-PER-BIT
- BANDWIDTH / WAVELENGTH
- POWER / EYE-SAFETY
- FORM FACTOR (SIZE & WEIGHT)
- SYSTEM COMPLEXITY
- ....



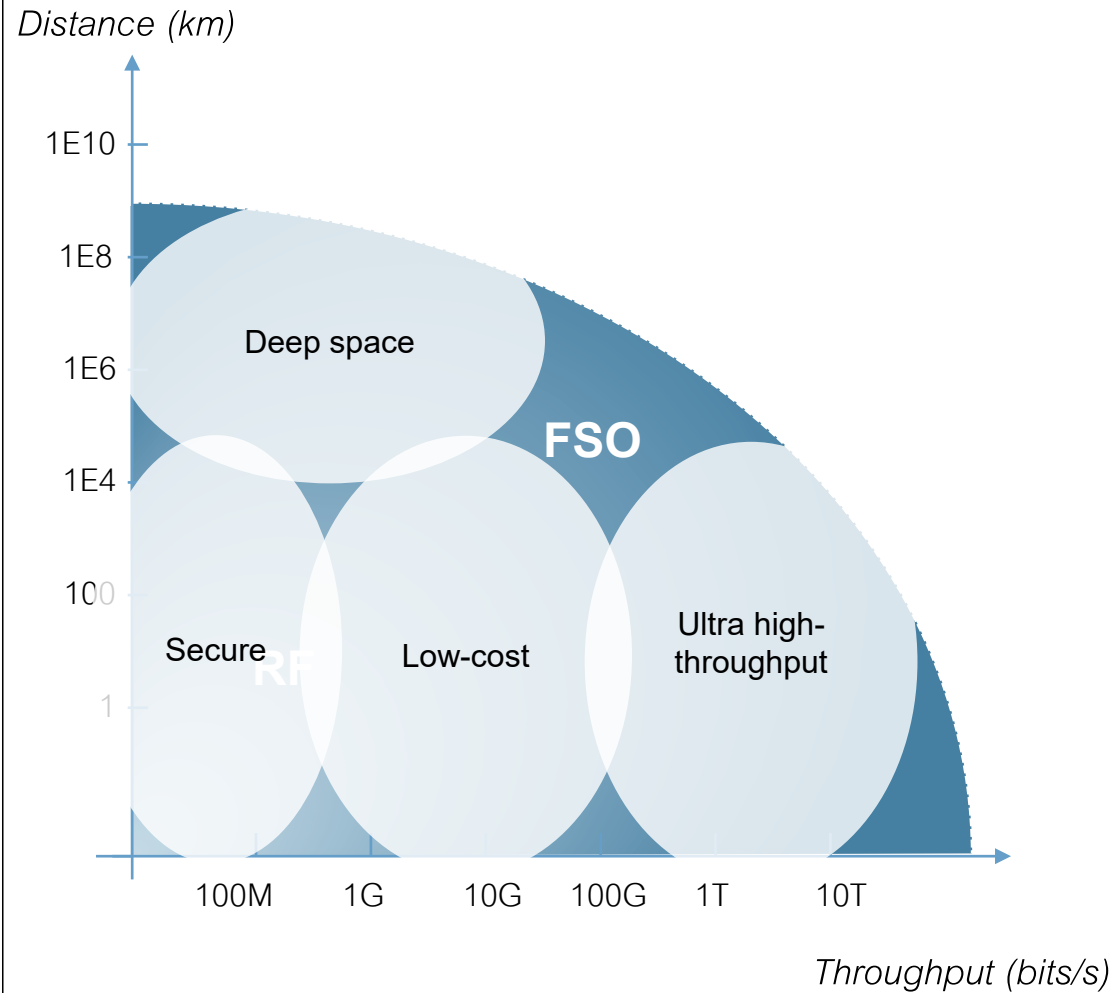
*Example visualization of performance of two different designs*



## › RF & FSO FIGURES OF MERIT



## USE CASES FOR FSO



Use case & USP	Typical datarate	Modulation
1. Low-cost	Gbps class	OOK
2. High-throughput	Tbps class	WDM-QPSK
3. Deep space – Photon-efficient	Mbps - Gbps	PPM
4. Ultra-secure - QKD	bps - kbps	CV-DV QKD

## ACHIEVABLE RATE – A MOONSHOT

Assume you want to set up an FSO laser link from the Moon to the Earth.

- You dimension your system such that 0.01 nW of average power is received.
- You may assume that we use single photon detectors, lossless receiver and no background noise.
- We use the NIR wavelength of 1550 nm.



Now the question, given 256-PPM, what is the achievable rate and photon efficiency?

Example Moon link:

The photon energy is  $h \cdot f$ , such that we have  $P_{TX}/(h \cdot f)$  photons / second.

With M-PPM, you need M timeslots per symbol, to send  $\log_2 (M)$  bits per symbol.

If you only need one photon per symbol, and transmit  $\log_2 (M)$  bits per symbol, you'll need  $1/\log_2(M)$  photons-per-bit.

With  $P_{TX}=0.01$  nW, we have  $7.8E7$  photons / second, with  $M=256$ , we get to 623 Mbps and get down to 0.125 ppb.



## › CALCULATE THROUGHPUT

- A. Number of optical channels, e.g. 96 x 50 GHz channels
- B. Baudrate per channel, e.g. 25 Gbaud per channel
- C. Modulation format, e.g., QPSK, 2 bits per symbol
- D. FEC & coding rate, e.g., coding rate of 0.9 means 90% is used for actual data
- E. Polarizations, e.g., 2 for dual-polarization doubles the throughput
- F. Transmission success factor, e.g., 0.99, on average 1% of retransmissions
- G. Overhead, headers & protocols, e.g., 0.9 to account for 10% overhead

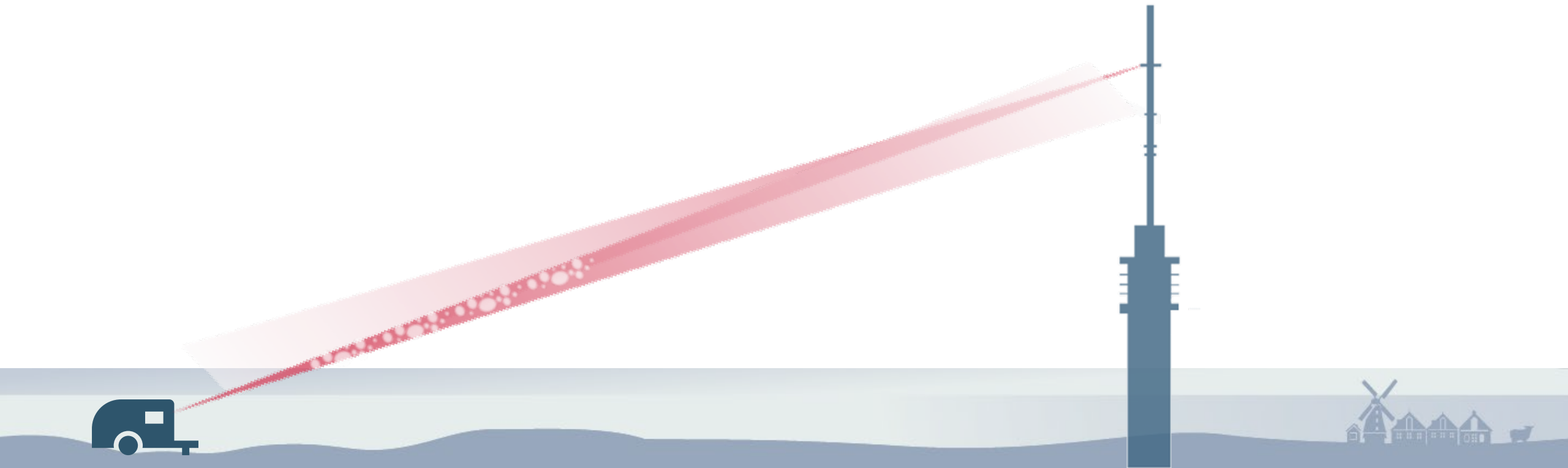
**Estimated throughput is:  $A \times B \times C \times D \times E \times F \times G$**

*(for the example values you get: 7.8 Tbps)*



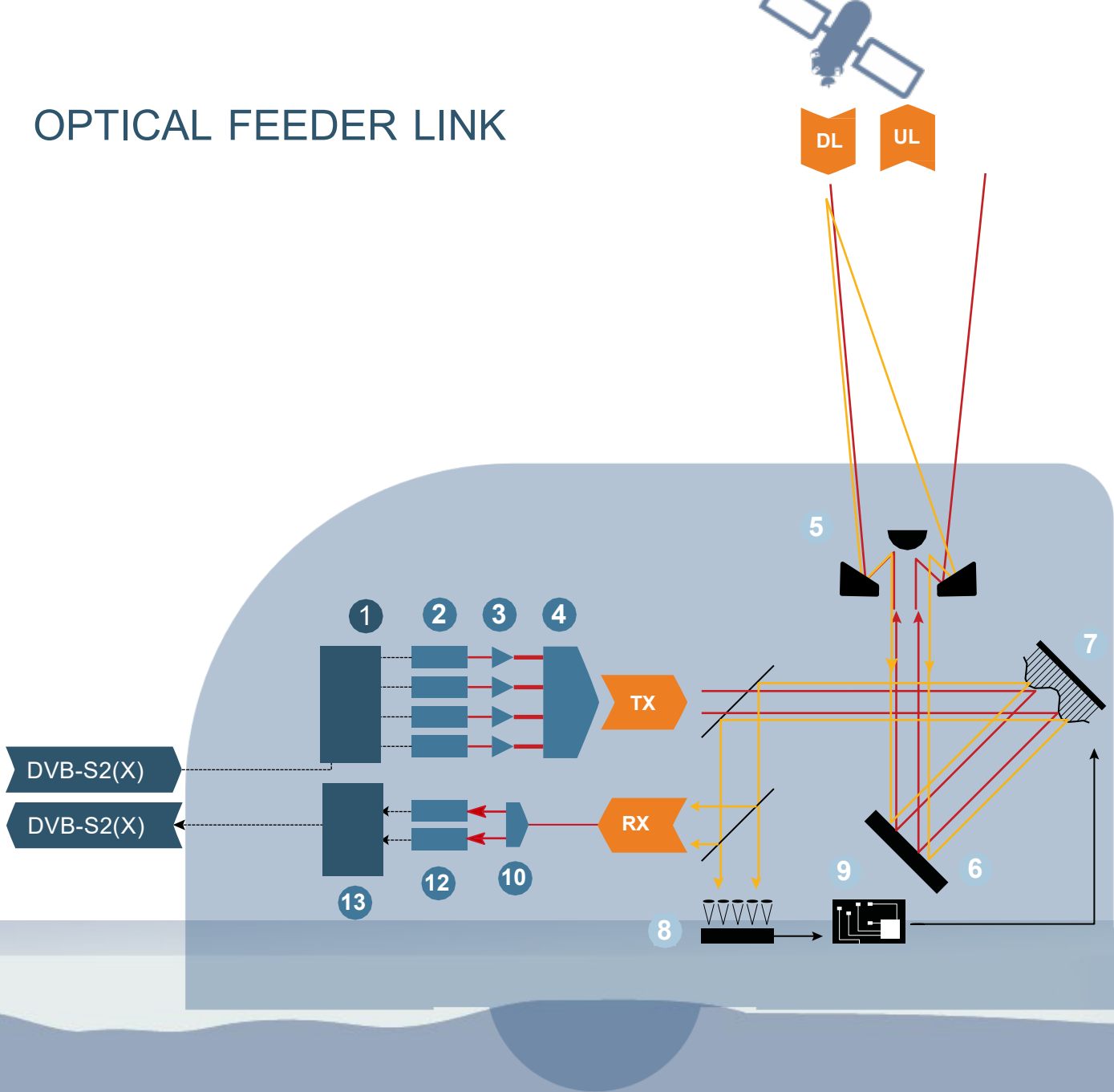
**TNO**

## OPTICAL FEEDER LINK



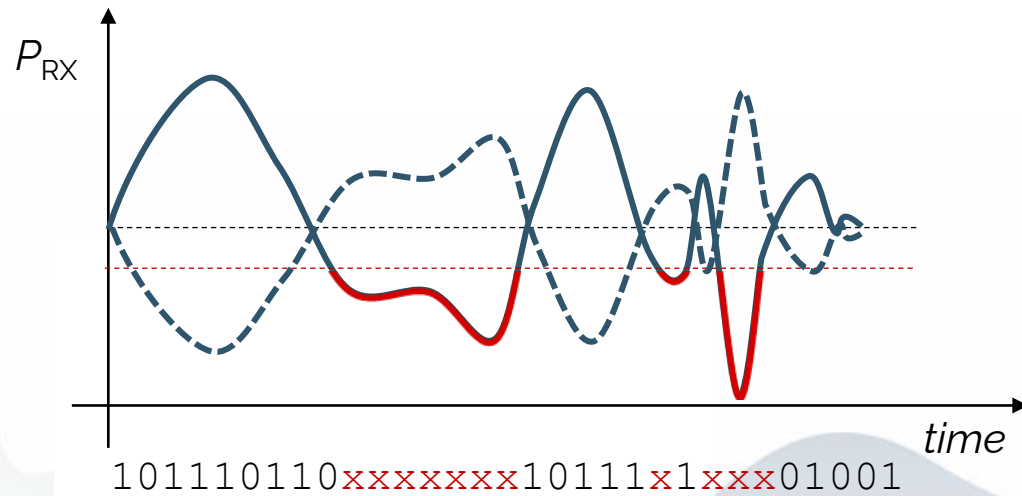


# OPTICAL FEEDER LINK



- 1. Digital Processor
- 2. Optical modulators
- 3. Booster amplifier
- 4. High power multiplexer
- 5. Telescope
- 6. Tip/tilt corrector
- 7. Deformable mirror
- 8. Wavefront sensor
- 9. Control system
- 10. De-multiplexer
- 11. Preamplifier
- 12. Detectors, LIA, TIA & CDR
- 13. Digital processor

# THE PROBLEM OF FADING & MITIGATION METHODS



Retransmissions

Overdimensioning

Power precompensation \*

AO precompensation \*

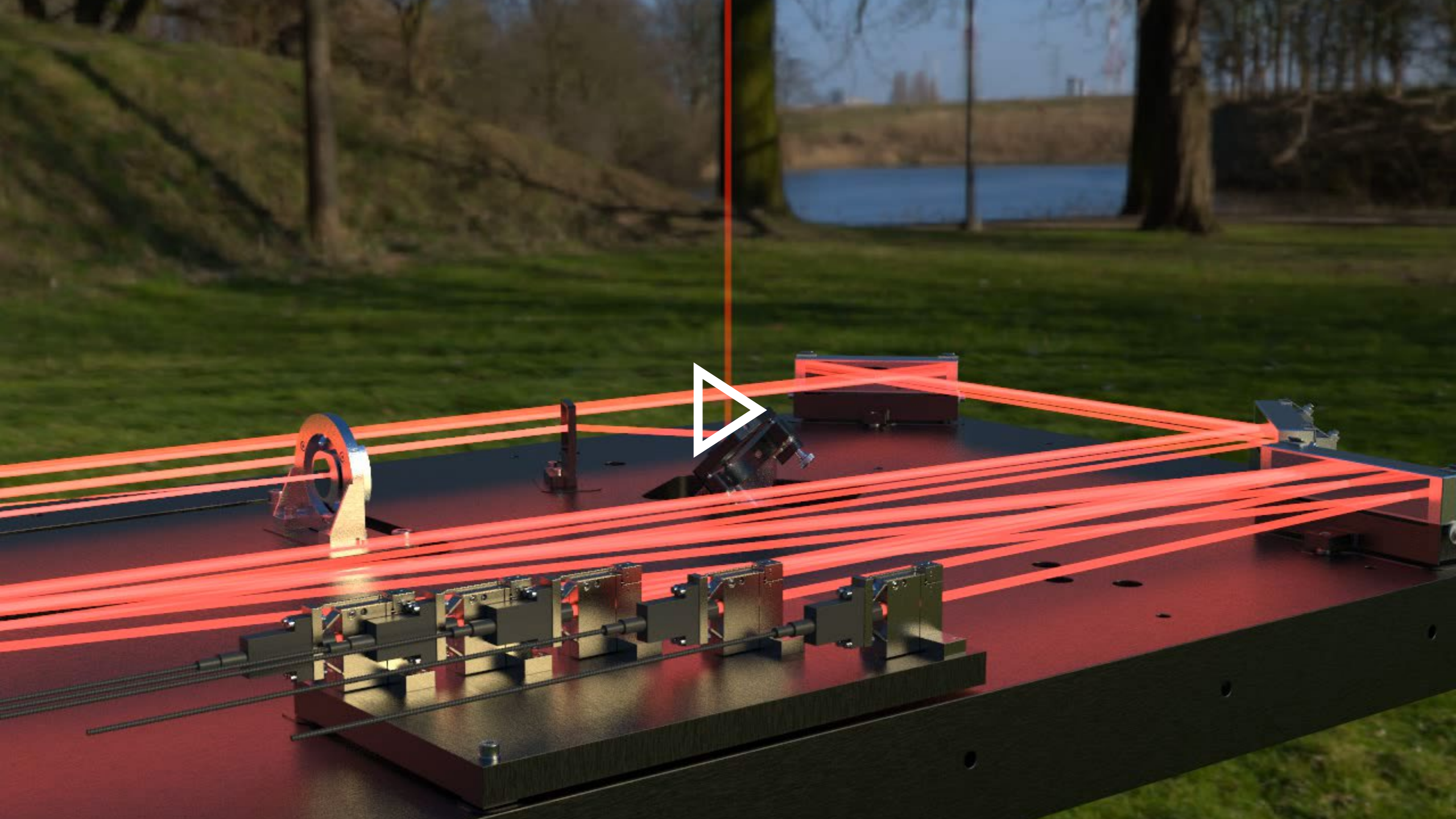
\* Assuming CSI @ TX

# THE PROBLEM OF FADING & DIVERSITY METHODS

- Temporal diversity
- Spatial diversity
- Site diversity
- Frequency diversity



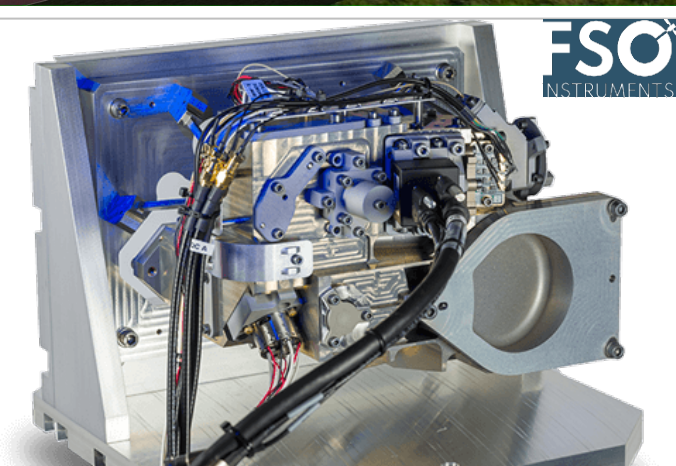
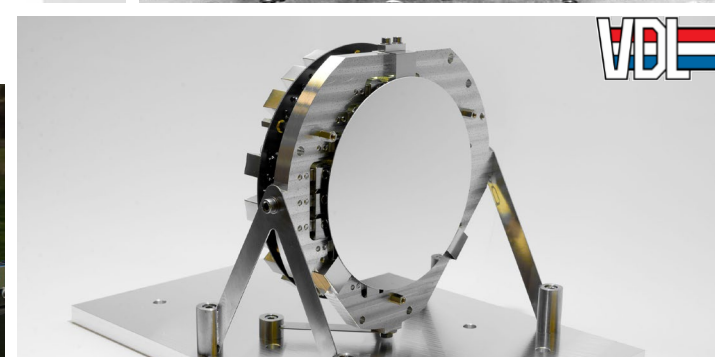
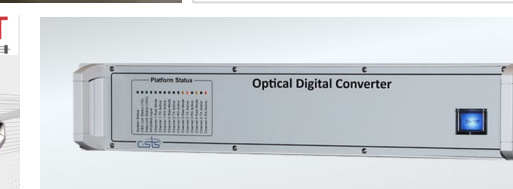
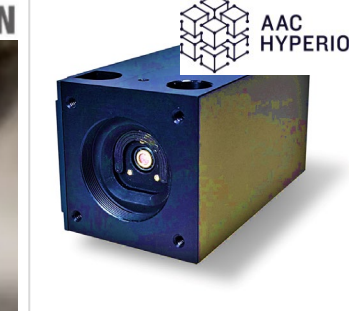
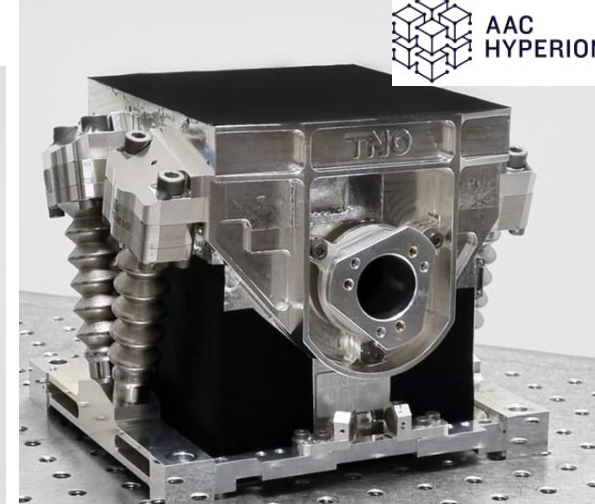


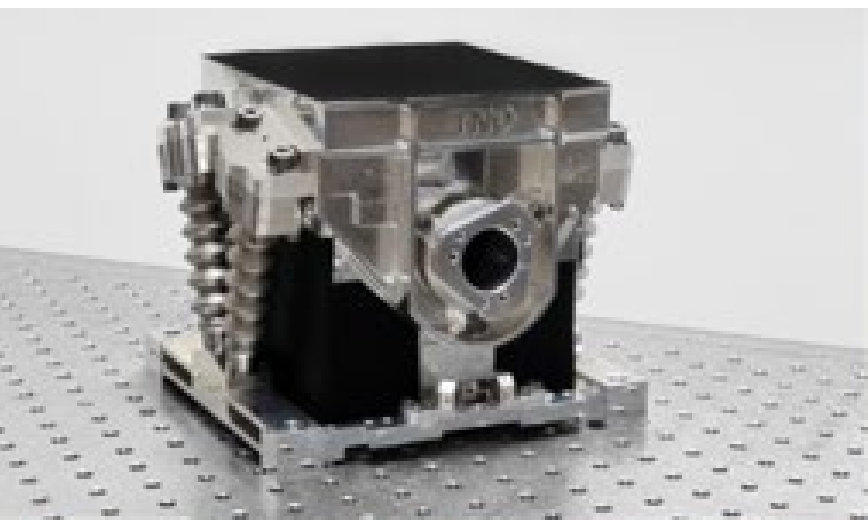


















**THANK YOU** | Q&A

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