

So far we have limited our discussion of ①
 3D QM to angular solutions for which
 we forgo modeling the interactions as
 they feature in the radial eqn.

We posited solutions that we
 separable $\Psi(r, \theta, \phi) = R(r) \Theta(\theta) \Phi(\phi)$

and we found that the spherical harmonics
 could fully describe our angular results,

$$Y_l^m(\theta, \phi) = \Theta_l^m(\theta) \Phi_m(\phi)$$

We also found that the separation
 constant, A , that we introduced was
 equal to $l(l+1)$. All of this results
 in a radial equation given by,

$$\left[\frac{-\hbar^2}{2\mu r^2} \frac{d}{dr} \left(r^2 \frac{d}{dr} \right) + V(r) + l(l+1) \frac{\hbar^2}{2\mu r^2} \right] R(r) = E R(r)$$

b/c the last two terms depend only on ② r , it's common to refer to their sum as the "effective potential" (like in classical)

$$V_{\text{eff}}(r) = V(r) + \ell(\ell+1) \frac{\hbar^2}{2mr^2}$$

But to develop a solution we need a particular $V(r)$. In this case, we want to work with Hydrogenic atoms, so

$$V(r) = -\frac{ze^2}{4\pi\epsilon_0 r} \quad \text{Coulomb Potential}$$

We can rewrite the DiffyQ,

$$\frac{d^2R}{dr^2} + \frac{2}{r} \frac{dR}{dr} + \frac{2\mu}{\hbar^2} \left[E + \frac{ze^2}{4\pi\epsilon_0 r} - \frac{\hbar^2\ell(\ell+1)}{2mr^2} \right] R = 0$$

$V(r \rightarrow \infty) \rightarrow 0$ so that we cannot "get rid" of $V(r)$ and we have $E < 0$ bound states & $E > 0$ unbound states.

"Non-dimensionalizing" a Drift Q

(3)

It is common practice in theoretical physics to remove the dimensionality in analysis. This leads to find characteristic length, mass, time, energy, etc. scales, but also parameterizes our results in terms of these characteristic scales.

We will do this partially for $R(r)$ by recasting our analysis using a dimensionless

variable $\rho = r/a \leftarrow$ as of yet unknown length scale

So that

$$R(r) \rightarrow R(\rho)$$

this is a relatively straight forward process, which we can do via "replacement"

(4)

Replace!

$$\overline{p} = r/a \quad \text{thus} \quad r = pa$$

$$\text{and } \frac{d}{dr} = \frac{dp}{dr} \frac{d}{dp} = \frac{1}{a} \frac{d}{dp}$$

$$\text{and } \frac{d^2}{dr^2} = \frac{d}{dr} \left(\frac{1}{a} \frac{d}{dp} \right) = \frac{dp}{dr} \left(\frac{1}{a} \frac{d^2}{dp^2} \right) = \frac{1}{a^2} \frac{d^2}{dp^2}$$

This leads to,

$$\frac{1}{a^2} \frac{d^2 R}{dp^2} + \frac{1}{a^2} \frac{2}{p} \frac{dR}{dp} + \frac{2\mu}{\hbar^2} \left[E + \frac{ze^2}{4\pi\epsilon_0 a p} - \frac{\hbar^2 l(l+1)}{2\mu a^2 p^2} \right] R = 0$$

or,

$$\frac{d^2 R}{dp^2} + \frac{2}{p} \frac{dR}{dp} + \left[\frac{2\mu a^2}{\hbar^2} E + \left(\frac{\mu ze^2}{4\pi\epsilon_0 \hbar^2} \right) \frac{2a}{p} - \frac{l(l+1)}{p^2} \right] R = 0$$

p is dimensionless, so is $\frac{2a}{p}$ so, the units

of $\frac{\mu ze^2}{4\pi\epsilon_0 \hbar^2}$ are $1/\text{length}$

We identify this as our characteristic length

$$a = \frac{4\pi\epsilon_0 \hbar^2}{\mu ze^2}$$

In addition $\frac{2\mu a^2}{\hbar^2}$ has units of $1/\text{energy}$ (5)

So we identify $\frac{\hbar^2}{2\mu a^2}$ as a characteristic energy scale and take the ratio,

$E / \left(\frac{\hbar^2}{2\mu a^2} \right)$ as a negative quantity b/c $E < 0$ gives bound states

So,

$$-\gamma^2 \equiv \frac{E}{\hbar^2/2\mu a^2}$$

where

$$\gamma^2 > 0$$

Thus,

$$\frac{d^2R}{dp^2} + \frac{2}{p} \frac{dR}{dp} + \left[-\gamma^2 + \frac{2}{p} - \frac{l(l+1)}{p^2} \right] R = 0$$

is our eigen value eqn.

Solving for $R(p)$ (or $R(r)$)

(6)

We will bring a new approach to solving this differential equation

⇒ Matching asymptotic solutions ($p \rightarrow 0$ & $p \rightarrow \infty$)

This is done in 3 steps,

① Find Approx Diffy Q for $p \rightarrow \infty$

② Find Approx Diffy Q for $p \rightarrow 0$

③ Match asymptotic solutions with full Diffy Q.

①

Let $p \rightarrow \infty$,

$$\frac{d^2R}{dp^2} + \frac{2}{p} \frac{dR}{dp} + \left[-\gamma^2 + \frac{2}{p} - \frac{l(l+1)}{p^2} \right] R = 0$$

$\underbrace{}_{\rightarrow 0} \quad \underbrace{_{\rightarrow 0}}_{\rightarrow 0} \quad \underbrace{_{\rightarrow 0}}_{\rightarrow 0}$

as $p \rightarrow \infty$

$$\frac{d^2R}{dp^2} - \gamma^2 R \approx 0 \quad \text{Approx Diffy Q}$$

for $p \rightarrow \infty$

(7)

We thus expect $R(p) \sim e^{\pm \gamma p}$

But $e^{\pm \gamma p}$ blows up as $p \rightarrow \infty$ so,

$R(p) \sim e^{-\gamma p}$ is our asymptotic solution.
for $p \rightarrow \infty$

② Let $p \rightarrow 0$

$$\frac{d^2R}{dp^2} + \frac{2}{p} \frac{dR}{dp} + \left[-\gamma^2 + \frac{2}{p} - \frac{\ell(\ell+1)}{p^2} \right] R = 0$$

nominal big really big!

$$\frac{d^2R}{dp^2} + \frac{2}{p} \frac{dR}{dp} - \frac{\ell(\ell+1)}{p^2} R \approx 0 \quad \text{Approx Diffy Q}$$

for $p \rightarrow 0$

it looks like a polynomial $R(p) = p^2$
works as all the terms give $p^{\ell-2}$

so lets pop that in, (note we could have \$Cp^g\$ but the Cs) (8)

$$\frac{dR}{dp} = g p^{g-1} \quad \frac{d^2R}{dg^2} = \cancel{g(g-1)} p^{g-2}$$

$$g(g-1)p^{g-2} + 2g p^{g-2} - \ell(\ell+1)p^{g-2} = 0$$

$$g^2 - g + 2g - \ell(\ell+1) = 0$$

$$g(g+1) - \ell(\ell+1) = 0$$

thus $g = \ell$ or $-\ell-1$

so $R = p^\ell$ or $R = p^{-\ell-1}$ blows up for $p \rightarrow 0$

so

$R(p) \sim p^\ell$ for our asymptotic solution
as $p \rightarrow 0$

so we get

$$R(p) \sim p^\ell e^{-\gamma p} \quad \text{as } p \rightarrow 0 \text{ & } p \rightarrow \infty$$

This behaves fine

(10)

③ Intermediate ρ ?

Assume so function $f(\rho)$ as of yet determined. and find the Diffy Q it satisfies,

$$R(\rho) = \rho^l e^{-\gamma\rho} f(\rho)$$

$$\begin{aligned}\frac{dR}{d\rho} &= l\rho^{l-1} e^{-\gamma\rho} f(\rho) + \rho^l (-\gamma e^{-\gamma\rho}) f(\rho) + \rho^l e^{-\gamma\rho} f'(\rho) \\ &= \rho^{l-1} e^{-\gamma\rho} [lf(\rho) - \gamma\rho f(\rho) + pf'(\rho)]\end{aligned}$$

$$f'(\rho) = \frac{df}{d\rho} \quad \underline{\text{BTW}}$$

$$\begin{aligned}\frac{d^2R}{d\rho^2} &= \rho^{l-1} e^{-\gamma\rho} [(2-2\gamma-2\gamma l)f(\rho) + (2+2l-2\gamma\rho)f'(\rho) \\ &\quad + pf''(\rho)]\end{aligned}$$

$$f''(\rho) = \frac{d^2f}{d\rho^2} \quad \underline{\text{BTW}}$$

Substitution gives,

(11)

$$P \frac{d^2 f}{dp^2} + 2(\lambda+1-\gamma p) \frac{df}{dp} + 2(1-\gamma-\gamma\lambda) f(p) = 0$$

looks like a mess, but let's try a series
solution,

$$f(p) = \sum_{j=0}^{\infty} c_j p^j$$

$$\begin{aligned} \frac{df}{dp} &= \sum_{j=0}^{\infty} j c_j p^{j-1} && \text{index shift} && \text{note } j=-1 \text{ term} \\ &= \sum_{j=-1}^{\infty} (j+1) c_{j+1} p^j && && = \sum_{j=0}^{\infty} (j+1) c_{j+1} p^j \\ \frac{d^2 f}{dp^2} &= \sum_{j=0}^{\infty} j(j+1) c_{j+1} p^{j-1} && && \text{maintain } p^j \text{ order} \\ P \frac{d^2 f}{dp^2} + 2(\lambda+1) \frac{df}{dp} - 2\gamma p \frac{df}{dp} + 2(1-\gamma-\gamma\lambda) f &= 0 && \text{maintain } p^j \text{ order} && \end{aligned}$$

$$\begin{aligned} \sum_{j=0}^{\infty} j(j+1) c_{j+1} p^j + 2(\lambda+1) \sum_{j=0}^{\infty} (j+1) c_{j+1} p^j \\ - 2\gamma \sum_{j=0}^{\infty} j c_j p^j + 2(1-\gamma-\gamma\lambda) \sum_{n=0}^{\infty} c_n p^n = 0 \end{aligned}$$

OK,

(12)

$$\sum_{j=0}^{\infty} \left(j(j+1) c_{j+1} + 2(\ell+1)(j+1) c_{j+1} - 2\gamma_j c_j + 2(1-\gamma-\gamma\ell) c_j \right) \varphi^j = 0$$

holds for each j and any φ so sum vanishes for each j !

$$j(j+1) c_{j+1} + 2(\ell+1)(j+1) c_{j+1} - 2\gamma_j c_j + 2(1-\gamma-\gamma\ell) c_j = 0$$

Thus,

$$c_{j+1} \left(j(j+1) + 2(\ell+1)(j+1) \right) - (2\gamma_j - 2(1-\gamma-\gamma\ell)) c_j = 0$$

$$c_{j+1} = \frac{2\gamma_j - 2 + 2\gamma + 2\gamma\ell}{(j+1)(j+2\ell+2)} c_j$$

$$c_{j+1} = \frac{2\gamma(j+\ell+1) - 2}{(j+1)(j+2\ell+2)} c_j$$

Recurrence relation
 c_0 determines all coeffs
get c_0 from $\langle \Psi | \Psi \rangle = 1$

$$f(p) = \sum_{j=0}^{\infty} c_j p^j$$

do we have ∞ terms? (13)

let $j \rightarrow \infty$,

$$c_{j+1} \approx \frac{2\alpha j}{j^2} c_j = \frac{2\alpha}{j} c_j$$

Note

$$e^{\alpha x} = 1 + \frac{1}{1!} x + \frac{\alpha^2}{2!} x^2 + \frac{\alpha^3}{3!} x^3 + \dots$$

hence,

$$c_j = \frac{\alpha}{j+1} c_j \quad \text{which is } c_{j+1} \approx \frac{\alpha}{j} c_j$$

for $j \rightarrow \infty$

In the large j limit,

$$f(p) \approx e^{2\alpha p}$$

like this exponential

So,

$$R(p) \approx p^l e^{-\alpha p} e^{2\alpha p} = p^l e^{\alpha p}$$

Oh no!

To get a well behaved

that grows as $p \rightarrow \infty$!

$R(p)$, j must terminate

(like w/
Legendre Polys)

Assume a j_{\max} such that,

(14)

$$2\gamma(j_{\max} + l + 1) - 2 = 0 \quad \left(\begin{array}{l} \text{(numerator of} \\ \text{recurrence} \\ \text{relationship} \end{array} \right)$$

j_{\max}, l are integers

so

$j_{\max} + l + 1$ is an integer, n

$$\boxed{n = j_{\max} + l + 1}$$

Principal Quantum Number, n

j and l start @ 0 so

$$\boxed{n = 1, 2, 3, \dots \infty}$$

$$2\gamma n - 2 = 0 \quad \text{so} \quad \boxed{\gamma = \frac{1}{n}}$$

energy is quantized (by $n!$) ↗

Energy Quantization

(15)

With $\gamma = 1/a$ we get,

$$-\gamma^2 = -\frac{1}{h^2} = \frac{E}{\frac{h^2}{2ma^2}} = \frac{E}{\frac{h^2}{2m}} \left(\frac{4\pi\epsilon_0 e^2}{mZ} \right)^2$$

So that,

$$E_n = -\frac{1}{2n^2} \left(\frac{Ze^2}{4\pi\epsilon_0} \right)^2 \frac{\mu}{h^2} \quad n=1, 2, 3, \dots$$

For a given n ,

$$l = n - j_{\max} - 1$$

And thus we have 3 quantum numbers

$$n = 1, 2, 3, \dots, \infty$$

"shell / "orbital" number

$$l = 0, 1, 2, \dots, n-1$$

ang. mom.

$$m = -l, -l+1, \dots, 0, l-1, l$$

mag. quantum