

## Homework 2: Classification and Bias-Variance Trade-offs

### Introduction

This homework is about classification, bias-variance trade-offs, and uncertainty quantification.

The datasets that we will be working with relate to astronomical observations. The first dataset, found at `data/planet-obs.csv`, contains information on whether a planet was observed (as a binary variable) at given points in time. This will be used in Problem 1. The second dataset, available at `data/hr.csv`, details different kinds of stars and their measured magnitude and temperature. You will work with this data in Problem 3.

As a general note, for classification problems we imagine that we have the input matrix  $\mathbf{X} \in \mathbb{R}^{n \times d}$  (or perhaps they have been mapped to some basis  $\Phi$ , without loss of generality) with outputs now “one-hot encoded.” This means that if there are  $K$  output classes, rather than representing the output label  $y$  as an integer  $1, 2, \dots, K$ , we represent  $\mathbf{y}$  as a “one-hot” vector of length  $K$ . A “one-hot” vector is defined as having every component equal to 0 except for a single component which has value equal to 1. For example, if there are  $K = 7$  classes and a particular data point belongs to class 3, then the target vector for this data point would be  $\mathbf{y} = [0, 0, 1, 0, 0, 0, 0]$ . We will define  $C_1$  to be the one-hot vector for the 1st class,  $C_2$  for the 2nd class, etc. Thus, in the previous example  $\mathbf{y} = C_3$ . If there are  $K$  total classes, then the set of possible labels is  $\{C_1 \dots C_K\} = \{C_k\}_{k=1}^K$ . Throughout the assignment we will assume that each label  $\mathbf{y} \in \{C_k\}_{k=1}^K$  unless otherwise specified. The most common exception is the case of binary classification ( $K = 2$ ), in which case labels are the typical integers  $y \in \{0, 1\}$ .

### Resources and Submission Instructions

We encourage you to read CS181 Textbook’s Chapter 3 for more information on linear classification, gradient descent, and classification in the discriminative setting. Read Chapter 2.8 for more information on the trade-offs between bias and variance.

In problems 1 and 3, you may use `numpy` or `scipy`, but not `scipy.optimize` or `sklearn`. Example code is given in the provided notebook.

Please type your solutions after the corresponding problems using this L<sup>A</sup>T<sub>E</sub>X template, and start each problem on a new page.

Please submit the **writeup PDF to the Gradescope assignment ‘HW2’**. Remember to assign pages for each question. Please submit your **L<sup>A</sup>T<sub>E</sub>X file and code files to the Gradescope assignment ‘HW2 - Supplemental’**. **You must include your plots in your writeup PDF**. The supplemental files will only be checked in special cases, e.g. honor code issues, etc.

**Problem 1** (Exploring Bias-Variance and Uncertainty)

In this problem, we will explore the bias and variance of a few different model classes when it comes to logistic regression and investigate two sources of predictive uncertainty in a synthetic (made-up) scenario.

We are using a powerful telescope in the northern hemisphere to gather measurements of some planet of interest. At certain times however, our telescope is unable to detect the planet due to its positioning around its star. The data in `data/planet-obs.csv` records the observation time in the “Time” column and whether the planet was detected in the “Observed” column (with the value 1 representing that it was observed). These observations were taken over a dark, clear week, which is representative of the region. Since telescope time is expensive, we would like to build a model to help us schedule and find times when we are likely to detect the planet.

1. Split the data into 10 mini-datasets of size  $N = 30$  (i.e. dataset 1 consists of the first 30 observations, dataset 2 consists of the next 30, etc. This has already been done for you). Consider the three bases  $\phi_1(t) = [1, t]$ ,  $\phi_2(t) = [1, t, t^2]$ , and  $\phi_3(t) = [1, t, t^2, t^3, t^4, t^5]$ . For each of these bases, fit a logistic regression model using  $\text{sigmoid}(\mathbf{w}^\top \phi(t))$  to each dataset by using gradient descent to minimize the negative log-likelihood. This means you will be running gradient descent 10 times for each basis, once for each dataset.

Use the given starting values of  $\mathbf{w}$  and a learning rate of  $\eta = 0.001$ , take 10,000 update steps for each gradient descent run, and make sure to average the gradient over the data points at each step. These parameters, while not perfect, will ensure your code runs reasonably quickly.

2. After consulting with a domain expert, we find that the probability of observing the planet is periodic as the planet revolves around its star—we are more likely to observe the planet when it is in front of its star than when it is behind it. In fact, the expert determines that observation follows the generating process  $y \sim \text{Bern}(f(t))$ , where  $f(t) = 0.4 \times \cos(1.1t + 1) + 0.5$  for  $t \in [0, 6]$  and  $y \in \{0, 1\}$ . Note that we, the modelers, do not usually see the true data distribution. Knowledge of the true  $f(t)$  is only exposed in this problem to allow for verification of the true bias.

Use the given code to plot the true process versus your learned models. Include your plots in your solution PDF.

**In no more than 5 sentences**, explain how bias and variance reflected in the 3 types of curves on the graphs. How do the fits of the individual and mean prediction functions change? Keeping in mind that none of the model classes match the true generating process exactly, discuss the extent to which each of the bases approximates the true process.

### Problem 1 (cont.)

3. If we were to increase the size of each dataset drawn from  $N = 30$  to a larger number, how would the bias and variance change for each basis? Why might this be the case? You may experiment with generating your own data that follows the true process and plotting the results, but this is **not** necessary. **Your response should not be longer than 5 sentences.**
4. Consider the test point  $t = 0.1$ . Using your models trained on basis  $\phi_3$ , report the predicted probability of observation of the *first* model (the model trained on the first 30 data points). How can we interpret this probability as a measure of uncertainty? Then, compute the variance of the classification probability over your 10 models at the same point  $t = 0.1$ . How does this measurement capture another source of uncertainty, and how does this differ from the uncertainty represented by the classification probability? Repeat this process (reporting the first model's classification probability and the variance over the 10 models) for the point  $t = 3.2$ .

Compare the uncertainties and their sources at times  $t = 0.1$  and  $t = 3.2$ .

5. We now need to make some decisions about when to request time on the telescope. The justifications of your decisions will be sent to your funding agency, which will determine whether you will be allocated funds to use the telescope for your project.
  - To identify the ideal time, which model(s) would you use and why?
  - What time would you request, and why?
  - Your funding agency suggests using a different telescope in a humid area near the equator. Can you still use your model to determine when the planet is likely to be visible? Why? Are there adaptations that may be necessary?
  - You seek out a team that has used the alternative telescope for observing this planet, and they provide you their observation file `data/planet-obs-alternate.csv`. Compare the observations from your telescope to theirs. What seems to be happening? What might be an appropriate model for this? Your funding agency asks you to refit your models on these new data. Do you think this is a reasonable ask, and if so, how will it help you make better decisions about when to request viewing time? If not, why do you think the additional modeling will not help? You do *not* need to do any modeling for this question!
  - The team who shared the data tells you that the effect you're seeing is due to the weather around the telescope. Weather forecasts for the area are usually accurate a few days in advance. Does that affect your strategy for requesting time? If so, how? If not, why not?

In these questions, we are looking for your reasoning; there may be more than one valid answer.

1. I'm not entirely sure what the submittable portion of this problem is, because it is simply saying to set up the way the model will be run, so to avoid not submitting anything, I decided to submit my code in this section, which passed the tests that the question talks about. For this part of the problem, I changed the Code in Jupyter Notebook to the following. For Basis2 and Basis3 I wrote:

```
def basis2(t):
    return np.stack([np.ones(len(t)), t, t**2], axis=1)

def basis3(t):
    return np.stack([np.ones(len(t)), t, t**2, t**3, t**4, t**5], axis=1)
```

Then, for the fit and predict functions, I wrote the following code:

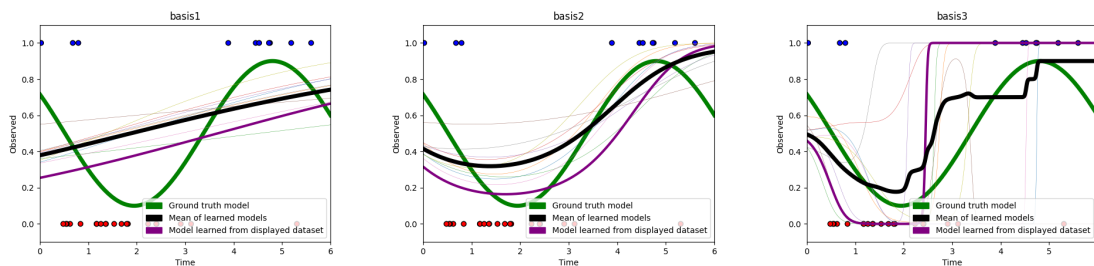
```
def fit(self, x, y, w_init):
    # Keep this for the autograder
    self.W = w_init
    # TODO
    n = x.shape[0]

    for run in range(self.runs):
        y_pred = self.predict(x)
        gradient = np.dot(x.T, (y_pred - y)) / n
        self.W -= self.eta * gradient

    def predict(self, x):
        z = np.dot(x, self.W)
        probs = sigmoid(z)
        return probs
```

With these functions, I was able to make all three bases pass the test.

2. The plots I got are as given:



From looking at these plots, it is apparent the relationship between variance and bias. In the first basis, which is the simplest, there is a high level of bias, as can be seen through the distance between the mean of the learned model and the ground truth of the dataset. Of the three graphs, the first basis has the most bias, but it also has a very low variance, because the line does not fit the data well. That means that outlying

points will do comparatively little to sway the mean. This trend is continued with bases 2 and 3, with the trend being that as the complexity increases, the bias decreases, with the model fitting the data better, but also the curve gets increasingly more susceptible to variation from new points. The third basis has the best approximation of the true data in terms of bias, but also is prone to overfitting as seen in the irregular behaviors of the graph.

3. If we increased the size of each dataset to a larger number, that would lead to a decrease in bias across all three bases. This is because this increased  $N$  would allow each model more information to train on and would lead to a more informed approximation, which will have a lower bias than with less data points. The more data points there are though the more potential there is for variance. As data continues to be collected, that data picks up more on variability within what is being measured, so the models are likely to have a higher variance.

4. I found that at the test point  $t = 0.1$ , Model 1, which is the first 30 data points, gives a predicted probability of 0.3482, which can be interpreted by saying that the first model is .3482 certain that the planet is visible at that time. This means that because this value is closer to 0 than it is to 1 that this model is relatively more certain that the planet is not visible at this time than that it is visible. Then I calculated the variance of the answers that all ten models gave, and found that the variance between all ten models was 0.0049, indicating that there is not much variation at all between individual models. This also helps to quantify the uncertainty because it is clear to see that all ten of the models are very similar to each other, so all the models roughly agree that the probability is closer to 0 than 1, which make the assertion that the planet is not usually seen at time  $t = 1$  even more certain. Then, when done again for  $t = 3.2$ , I found that for model 1, the predicted probability at 3.2 was 0.6074, which again means that for model 1, there is a roughly 60% chance that the planet can be seen at time 3.2, which means it is more likely to be present than not, but does not have a certainty or guarantee to be seen. When calculating the variance over the 10 model at time 3.2 I got Variance = 0.0056, which once again shows that all ten models agree pretty closely and increases the certainty of model 1's findings.

5. To identify the ideal time, I would use the second model; the one based off of basis 2. That is because that model is in a nice sweet spot of not being overfit and having a reasonably small bias and also a reasonably small variance. Of the three bases, it is the most stable and accurate, so it is the best to use for analyzing the information. I would then request the time to be between time 4 and time 5, since those are times when the planet was very frequently spotted through the telescope, while also having a minimal mean squared error across that interval. If this project were to be moved to a more humid climate, the model found above will most likely not be applicable. These models are specific to a cloudless environment where there is no interference from weather, and in a different climate that is unlikely to also be the case, so different data must be collected to account for weather patterns and cloud patterns in the new area. Looking at the data provided in planet-obs-alternate, it appears that the alternate location has a very different graph for planet observation. The initial location has a much more sinusoidal nature to the curve of planet visibility, but it appears that the alternate dataset has much more consistent visibility with no clear sinusoidal function to model it. Refitting the models to match this data makes sense because of the difference in distribution between the locations. This being a result of the weather in the area makes sense and it does effect your request for time, because you want to make sure you avoid a time when there is no probability of seeing the planet just due to clouds being in the way. It would also be helpful in this case to prove that the clouds being there does not have a spurious connection to the planet being visible which is skewing the data, but it is still best to request time according to the weather patterns.

**Problem 2** (Maximum likelihood in classification, 15pts)

Consider now a generative  $K$ -class model. We adopt class prior  $p(\mathbf{y} = C_k; \boldsymbol{\pi}) = \pi_k$  for all  $k \in \{1, \dots, K\}$  (where  $\pi_k$  is a parameter of the prior). Let  $p(\mathbf{x}|\mathbf{y} = C_k)$  denote the class-conditional density of features  $\mathbf{x}$  (in this case for class  $C_k$ ). Consider the data set  $D = \{(\mathbf{x}_i, \mathbf{y}_i)\}_{i=1}^n$  where as above  $\mathbf{y}_i \in \{C_k\}_{k=1}^K$  is encoded as a one-hot target vector and the data are independent.

1. Write out the log-likelihood of the data set,  $\ln p(D; \boldsymbol{\pi})$ .
2. Since the prior forms a distribution, it has the constraint that  $\sum_k \pi_k - 1 = 0$ . Using the hint on Lagrange multipliers below, give the expression for the maximum-likelihood estimator for the prior class-membership probabilities, i.e.  $\hat{\pi}_k$ . Make sure to write out the intermediary equation you need to solve to obtain this estimator. Briefly state why your final answer is intuitive.

For the remaining questions, let the class-conditional probabilities be Gaussian distributions with the same covariance matrix

$$p(\mathbf{x}|\mathbf{y} = C_k) = \mathcal{N}(\mathbf{x}|\boldsymbol{\mu}_k, \boldsymbol{\Sigma}), \text{ for } k \in \{1, \dots, K\}$$

and different means  $\boldsymbol{\mu}_k$  for each class.

3. Derive the gradient of the log-likelihood with respect to vector  $\boldsymbol{\mu}_k$ . Write the expression in matrix form as a function of the variables defined throughout this exercise. Simplify as much as possible for full credit.
4. Derive the maximum-likelihood estimator  $\hat{\boldsymbol{\mu}}_k$  for vector  $\boldsymbol{\mu}_k$ . Briefly state why your final answer is intuitive.
5. Derive the gradient for the log-likelihood with respect to the covariance matrix  $\boldsymbol{\Sigma}$  (i.e., looking to find an MLE for the covariance). Since you are differentiating with respect to a *matrix*, the resulting expression should be a matrix!
6. Derive the maximum likelihood estimator  $\hat{\boldsymbol{\Sigma}}$  of the covariance matrix.

**Hint: Lagrange Multipliers.** Lagrange Multipliers are a method for optimizing a function  $f$  with respect to an equality constraint, i.e.

$$\min_{\mathbf{x}} f(\mathbf{x}) \text{ s.t. } g(\mathbf{x}) = 0.$$

This can be turned into an unconstrained problem by introducing a Lagrange multiplier  $\lambda$  and constructing the Lagrangian function,

$$L(\mathbf{x}, \lambda) = f(\mathbf{x}) + \lambda g(\mathbf{x}).$$

It can be shown that it is a necessary condition that the optimum is a critical point of this new function. We can find this point by solving two equations:

$$\frac{\partial L(\mathbf{x}, \lambda)}{\partial \mathbf{x}} = 0 \quad \text{and} \quad \frac{\partial L(\mathbf{x}, \lambda)}{\partial \lambda} = 0$$

**Cookbook formulas.** Here are some formulas you might want to consider using to compute difficult gradients. You can use them in the homework without proof. If you are looking to hone your matrix calculus skills, try to find different ways to prove these formulas yourself (will not be part of the evaluation of this homework). In general, you can use any formula from the matrix cookbook, as long as you cite it. We opt for the following common notation:  $\mathbf{X}^{-\top} := (\mathbf{X}^{\top})^{-1}$

$$\frac{\partial \mathbf{a}^{\top} \mathbf{X}^{-1} \mathbf{b}}{\partial \mathbf{X}} = -\mathbf{X}^{-\top} \mathbf{a} \mathbf{b}^{\top} \mathbf{X}^{-\top}$$

$$\frac{\partial \ln |\det(\mathbf{X})|}{\partial \mathbf{X}} = \mathbf{X}^{-\top}$$

1. Since we want to say something about every data point  $y_i$ , we can make  $n_k$  be equal to the amount of times in  $1, 2, \dots, K$  where  $y_i = C_k$ . To find the log-likelihood of  $P(D|\pi)$ , we first must calculate the likelihood. Which is  $\prod_{i=1}^n [P(x_i, y_i) | \pi] = \prod_{i=1}^n P(x_i | y_i, \pi) P(y_i | \pi)$ . That gives us the probability of the data given  $\pi$ . But more important than  $\pi$  is the idea that only when  $y_i = C_k$  will the value be non-zero, just by definition of one-hot encoding. This means we can rewrite the probability as  $\prod_{i=1}^n [P(x_i | y_i = C_k) \pi_k]$ . Then the next step is to take this equation and take the log of it, then using log rules:

$$\ln[\prod_{i=1}^n [P(x_i | y_i = C_k) \pi_k]] = \sum_{i=1}^n \ln(P(x_i | y_i = C_k) \pi_k) = \sum_{i=1}^n [\ln(P(x_i | y_i = C_k)) + \ln(\pi_k)]$$

Now from this we can complete the calculation using an indicator function, so that we only include values where  $y_i = C_k$ . We can define this as  $\mathbb{I}(y_i = C_k)$  where  $\mathbb{I}$  is 1 where  $y_i = C_k$  and 0 otherwise. Then we can make two sums, one over all data points and the other which is to sum over all the non-zero data points, or points where  $y_i \in n_k$ . With that defined we can now write it out fully as:

$$\ln(P(D|\pi)) = \sum_{k=1}^K \sum_{i=1}^n \mathbb{I}(y_i = C_k) (\ln(P(x_i | y_i = C_k)) + \ln(\pi_k))$$

2. Using the hint given in the problem, we want to optimize the log likelihood, but use the constraint  $\sum_k \pi_k - 1 = 0$ . Then also using the hint, we know we can set this up as:

$$\begin{aligned} L(\pi, \lambda) &= \ln(P(D|\pi)) + \lambda(\sum_k \pi_k - 1) \text{ Then as per Part 1} \\ &= \sum_{k=1}^K \sum_{i=1}^n \mathbb{I}(y_i = C_k) (\ln(P(x_i | y_i = C_k)) + \ln(\pi_k)) + \lambda(\sum_k \pi_k - 1) \end{aligned}$$

Now, we want to solve for both  $\frac{\partial L(\pi, \lambda)}{\partial \pi} = 0$  and  $\frac{\partial L(\pi, \lambda)}{\partial \lambda} = 0$ . Focusing on the first one with respect to  $\pi$ , we can see that in order to solve this equation, we can think about it in terms of the sum from 1 to  $K$  of  $\pi_k$ , and so we can sum over all the data points to get that  $\sum_{i=1}^n \frac{y_i}{\pi_k} + \lambda = 0$ . Rearranging for  $\pi_k$ , we get

$$\text{that } \pi_k = -\frac{\sum_{i=1}^n y_i}{\lambda}$$

Next, we can solve for  $\frac{\partial L(\pi, \lambda)}{\partial \lambda}$ . This is simpler, because the first probability that we found in Part 1 above does not include any lambda terms, so we only need to take into account  $\lambda(\sum_{k=1}^K \pi_k - 1)$ . Then since this equation has lambda on the outside, the derivative is simply  $\sum_{k=1}^K \pi_k - 1$ . This leaves us with the equation  $\sum_{k=1}^K \pi_k - 1 = 0$ , then we can substitute the  $\pi_k$  found above and get  $\sum_{k=1}^K -\frac{\sum_{i=1}^n y_i}{\lambda} - 1 = 0$ . Then, since we defined  $n_k$  to be the number of values such that  $y_i = C_k$ , and those are the only nonzero additions, we know that  $\sum_{k=1}^K \sum_{i=1}^n y_i = \sum_{i=1}^n y_i = n_k$  just by nature of definition. This leaves us with the expression  $\frac{n_k}{\lambda} = -1$ , so  $\lambda = -n_k$ . Then using this, we can go back to the equation for  $\pi_k$  and we find that  $\pi_k = \frac{n_k}{n}$ . This makes intuitive sense because we want to maximize the size of  $n_k$ , so it makes sense to have the maximum parameter be the true proportion of those points in order to have the best representation.

3. Again, in part one we found:  $\ln(P(D|\pi)) = \sum_{k=1}^K \sum_{i=1}^n y_{ik} (\ln(P(x_i | y_i = C_k)) + \ln(\pi_k))$ , and now we can rewrite this given the class-conditioned probability as  $\sum_{k=1}^K \sum_{i=1}^n y_{ik} (\ln(N(x_i | \mu_k, \Sigma))) + \ln(\pi_k)$ . Now we want to take the gradient of this with respect to  $\mu_k$  and we get:

$$\begin{aligned} \frac{\partial}{\partial \mu_k} (\ln(P(D|\pi))) &= \frac{\partial}{\partial \mu_k} [\sum_{k=1}^K \sum_{i=1}^n y_{ik} (\ln(N(x_i | \mu_k, \Sigma)))] \\ &= \frac{\partial}{\partial \mu_k} [\sum_{k=1}^K \sum_{i=1}^n y_{ik} \ln(\frac{1}{\sqrt{2\pi\Sigma}} \exp(-\frac{(x_i - \mu_k)^T \Sigma^{-1} (x_i - \mu_k)}{2}))] \\ &= \frac{\partial}{\partial \mu_k} [\sum_{i=1}^n -y_{ik} (\frac{(x_i - \mu_k)^T \Sigma^{-1} (x_i - \mu_k)}{2})]. \end{aligned}$$

Then we can use the Matrix Cookbook provided in the question to see that this is all equal to:

$$\sum_{i=1}^n -y_{ik}\Sigma^{-1}(\mu_k - x_i) = \sum_{i=1}^n (\mathbb{I}(y_i = C_k))(-\Sigma^{-1}(\mu_k - x_i))$$

4. For the maximum likelihood, we can take the equation we solved for in part 3 and just set it equal to 0:  $\sum_{i=1}^n -y_{ik}\Sigma^{-1}(\mu_k - x_i)$ , then we can distribute the  $y_{ik}$  since it is impacted by the sum and keep the  $\Sigma^{-1}$  outside of the sum which leaves us with  $\Sigma^{-1} \sum_{i=1}^n (-y_{ik}\mu_k + x_i y_{ik}) = \Sigma^{-1} [\mu_k \sum_{i=1}^n (-y_{ik}) + \sum_{i=1}^n x_i y_{ik}] = 0$ . Then, since we want to solve for  $\mu_k$ , we don't care about  $\Sigma^{-1}$  because it has no relevance to  $\mu_k$ , and we can rearrange to get that:

$$\hat{\mu}_k = \frac{\sum_{i=1}^n x_i y_{ik}}{\sum_{i=1}^n y_{ik}}.$$

Then, just as we did in the last part, we can rewrite this as the Indicator function to get that

$$\hat{\mu}_k = \frac{\sum_{i=1}^n x_i \mathbb{I}(y_i = C_k)}{\sum_{i=1}^n \mathbb{I}(y_i = C_k)}$$

Once again this answer is intuitive because it shows that the maximum likelihood approximates the sum of all  $x$  values of a class divided by all the points in the class, effectively giving an average of that class, which is the maximum likelihood.

5. This question wants  $\frac{\partial}{\partial \Sigma} [\ln(P(d|\pi))]$ .

Once again, we can build off the work we've done so far in this problem to write this more concretely as:

$$\begin{aligned} \frac{\partial}{\partial \Sigma} [\ln(P(d|\pi))] &= \frac{\partial}{\partial \Sigma} [\sum_{k=1}^K \sum_{i=1}^n y_{ik} (\ln(N(x_i|\mu_k, \Sigma))) + \ln(\pi_k)] = \frac{\partial}{\partial \Sigma} [\sum_{k=1}^K \sum_{i=1}^n y_{ik} (\ln(N(x_i|\mu_k, \Sigma)))] \\ &= \sum_{k=1}^K \sum_{i=1}^n y_{ik} \frac{\partial}{\partial \Sigma} \ln\left(\frac{1}{\sqrt{2\pi\Sigma}}\right) + \ln\left(\exp\left(-\frac{(x_i - \mu_k)^T \Sigma^{-1} (x_i - \mu_k)}{2}\right)\right) \\ &= \sum_{k=1}^K \sum_{i=1}^n y_{ik} \frac{\partial}{\partial \Sigma} \left[ \ln\left(\frac{1}{\sqrt{2\pi}}\right) + \ln(\Sigma^{-0.5}) - \frac{(x_i - \mu_k)^T \Sigma^{-1} (x_i - \mu_k)}{2} \right] \end{aligned}$$

From there, we can once again use the cookbook formulas and simplify this as:

$$\sum_{k=1}^K \sum_{i=1}^n y_{ik} \left( -\frac{1}{2} \Sigma^{-T} + \frac{\Sigma^{-T} (x_i - \mu_k) (x_i - \mu_k)^T \Sigma^{-T}}{2} \right)$$

And we finish once again by putting it in terms of the indicator function to get

$$\sum_{k=1}^K \sum_{i=1}^n \mathbb{I}(y_i = C_k) \left( -\frac{1}{2} \Sigma^{-T} + \frac{\Sigma^{-T} (x_i - \mu_k) (x_i - \mu_k)^T \Sigma^{-T}}{2} \right)$$

6. Very similar to Problem 4, we can find the Maximum Likelihood just by setting our answer in part 5 equal to 0 and solving. First, we can pull the term  $\frac{1}{2} \Sigma^{-T}$  from the Sums because it does not depend on  $k$  or  $i$ . As we pull it out though, we have to acknowledge that it will be multiplied by  $n_k$ , or the number of times that  $y_i = C_k$ , so we can rewrite this as  $-\frac{n}{2} \Sigma^{-T} + \sum_{k=1}^K \sum_{i=1}^n y_{ik} \left( \frac{\Sigma^{-T} (x_i - \mu_k) (x_i - \mu_k)^T \Sigma^{-T}}{2} \right)$ , then we can subtract that over to get  $\sum_{k=1}^K \sum_{i=1}^n y_{ik} \left( \frac{\Sigma^{-T} (x_i - \mu_k) (x_i - \mu_k)^T \Sigma^{-T}}{2} \right) = \frac{n}{2} \Sigma^{-T}$ .

Now we can multiply  $\Sigma^T$  on the left and right of both sides to eliminate the  $\Sigma^{-T}$ , which results in:

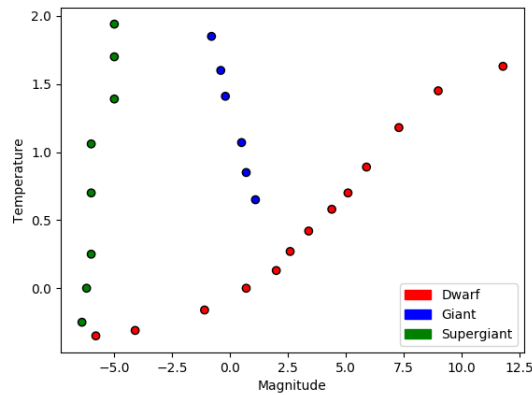
$$\begin{aligned} \sum_{k=1}^K \sum_{i=1}^n y_{ik} \left( \frac{(x_i - \mu_k) (x_i - \mu_k)^T}{2} \right) &= \frac{n}{2} \Sigma^T, \text{ so we have} \\ \Sigma^T &= \sum_{k=1}^K \sum_{i=1}^n \mathbb{I}(y_i = C_k) \left( \frac{(x_i - \mu_k) (x_i - \mu_k)^T}{n} \right) \end{aligned}$$





### Problem 3 (Classifying Stars)

In this problem, you will code up three different classifiers to classify different types of stars. The file `data/hr.csv` contains data on magnitude and temperature. The data can be plotted on these two axes:



Please implement the following classifiers in the `SoftmaxRegression` and `KNNClassifier` classes:

- A generative classifier with Gaussian class-conditional densities with a *shared covariance matrix*** across all classes. Feel free to re-use your Problem 2 results.
- Another generative classifier with Gaussian class-conditional densities, but now with a *separate covariance matrix*** learned for each class. (Note: The staff implementation can switch between the two Gaussian generative classifiers with just a few lines of code.)
- A multi-class logistic regression classifier** using the softmax activation function. In your implementation of gradient descent, **make sure to include a bias term and use L2 regularization** with regularization parameter  $\lambda = 0.001$ . Limit the number of iterations of gradient descent to 200,000, and set the learning rate to be  $\eta = 0.001$ .
- Another multi-class logistic regression classifier** with feature map  $\phi(\mathbf{x}) = [\ln(x_1 + 10), x_2^\top]^\top$ , where  $x_1$  and  $x_2$  represent the values for magnitude and temperature, respectively.
- A kNN classifier** in which you classify based on the  $k = 1$  and  $k = 5$  nearest neighbors and the following distance function:

$$\text{dist}(\text{star}_1, \text{star}_2) = (\text{mag}_1 - \text{mag}_2)^2/9 + (\text{temp}_1 - \text{temp}_2)^2$$

where nearest neighbors are those with the smallest distances from a given point.

Note 1: When there are more than two labels, no label may have the majority of neighbors. Use the label that has the most votes among the neighbors as the choice of label.

Note 2: The grid of points for which you are making predictions should be interpreted as our test space. Thus, it is not necessary to make a test point that happens to be on top of a training point ignore itself when selecting neighbors.

After implementing the above classifiers, complete the following exercises:

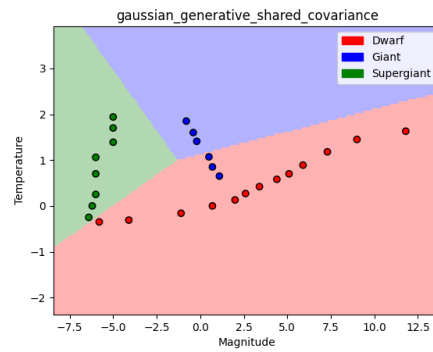
- Plot the decision boundaries generated by each classifier for the dataset. Include them in your PDF. Identify the similarities and differences among the classifiers. What explains the differences—in particular, which aspects or properties of each model dictate the shape of its decision boundary?
- Consider a star with Magnitude 3 and Temperature -2. To which class does each classifier assign this star? Report the classification probabilities of this star for each model.

Interpret how each model makes its classification decision. What are the pros and cons of each interpretation? What else should we, the modelers, be aware of when making predictions on a test point “far” from our training data? **Your response should no be longer than 5 sentences.**

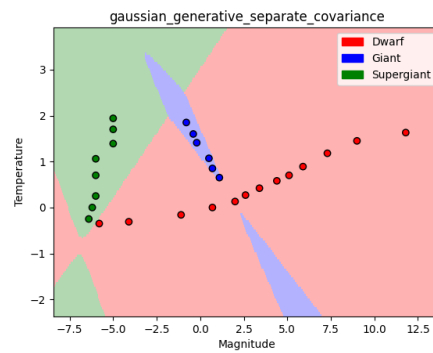
**Solution:**

1. I was able to make the following decision boundaries:

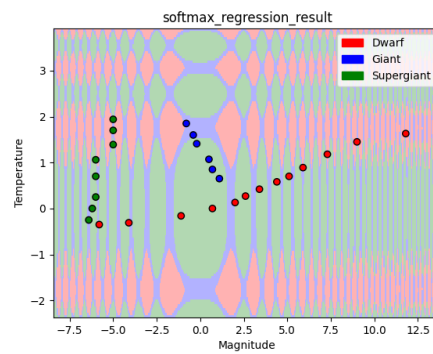
Gaussian with Shared Covariance:



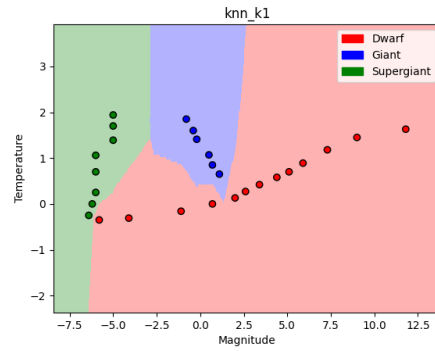
Gaussian with Separate Covariance:



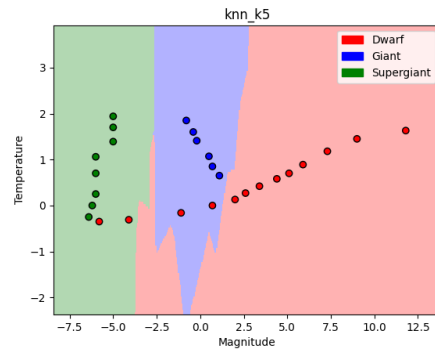
Logistic Regression Classifier:



kNN with  $k = 1$ :



kNN with  $k = 5$ :



**Name**

**Dan Tracy**

**Collaborators and Resources**

Whom did you work with, and did you use any resources beyond cs181-textbook and your notes? I got assistance from Students Teresa Lu-Remeo and Christine Lee, but most of the work was done independently.