



STACUP-2023



Hands-On  
Session-3

# Python Based Statistics and its Application in Astrophysics & Cosmology

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# Project Allotment

# Project-1.

Constraint Cosmological Parameters of Flat- $\Lambda$ CDM Model using Hubble Parameter Measurements.

$$H(z; H_0, \Omega_{m0}) = H_0 \sqrt{\Omega_{m0}(1+z)^3 + 1 - \Omega_{m0}}$$

Parameters:  $\mathbb{P} = \{H_0, \Omega_{m0}\}$

## Project-2.

Constraint Cosmological Parameters of Non-Flat- $\Lambda$ CDM Model using Hubble Parameter Measurements.

$$H(z; H_0, \Omega_{m0}, \Omega_{\Lambda 0}) = H_0 \sqrt{\Omega_{m0}(1+z)^3 + (1 - \Omega_{m0} - \Omega_{\Lambda 0})(1+z)^2 + \Omega_{\Lambda 0}}$$

Parameters:  $\mathbb{P} = \{H_0, \Omega_{m0}, \Omega_{\Lambda 0}\}$

## Project-3.

Constraint Cosmological Parameters of Flat- $\omega$ CDM Model using Hubble Parameter Measurements.

$$H(z; H_0, \Omega_{m0}, \omega) = H_0 \sqrt{\Omega_{m0}(1+z)^3 + (1 - \Omega_{m0})^{3(1+\omega)}}$$

Parameters:  $\mathbb{P} = \{H_0, \Omega_{m0}, \omega\}$

## Project-4.

Constraint Cosmological Parameters of Flat- $\Lambda$ CDM Model using Hubble Parameter Measurements.

$$H(z; H_0, z_t) = H_0 \sqrt{\frac{(1+z)^3}{\frac{1}{2}(1+z_t)^3 + 1} + \frac{(1+z_t)^3}{(1+z_t)^3 + 2}}$$

Parameters:  $\mathbb{P} = \{H_0, z_t\}$

## Project-5.

Constraint Cosmological Parameters of Flat- $q$ CDM Model using Hubble Parameter Measurements.

$$H(z; H_0, q) = H_0 (1 + z)^{1+q}$$

Parameters:  $\mathbb{P} = \{H_0, q\}$

## Project-6.

Constraint Cosmological Parameters of Flat- $\Lambda$ CDM Model using Type Ia Supernovae Observations.

$$\text{Luminosity distance: } d_L = c(1+z) \int_0^z \frac{dx}{H(x)}$$

$$H(z; H_0, \Omega_{m0}) = H_0 \sqrt{\Omega_{m0}(1+z)^3 + 1 - \Omega_{m0}}$$

$$\text{Parameters: } \mathbb{P} = \{H_0, \Omega_{m0}\}$$



## Project-7.

### Constraint Cosmological Parameters of Flat- $\omega$ CDM Model using Type Ia Supernovae Observations

$$\text{Luminosity distance: } d_L = c(1+z) \int_0^z \frac{dx}{H(x)}$$

$$H(z; H_0, \Omega_{m0}, \omega) = H_0 \sqrt{\Omega_{m0}(1+z)^3 + (1-\Omega_{m0})^{3(1+\omega)}}$$

$$\text{Parameters: } \mathbb{P} = \{H_0, \Omega_{m0}, \omega\}$$

## Project-8.

### Constraint Cosmological Parameters of Flat- $\Lambda$ CDM Model using Type Ia Supernovae Observations

$$\text{Luminosity distance: } d_L = c(1+z) \int_0^z \frac{dx}{H(x)}$$

$$H(z; H_0, z_t) = H_0 \sqrt{\frac{(1+z)^3}{\frac{1}{2}(1+z_t)^3 + 1} + \frac{(1+z_t)^3}{(1+z_t)^3 + 2}}$$

$$\text{Parameters: } \mathbb{P} = \{H_0, z_t\}$$

## Project-9.

### Constraint Cosmological Parameters of Flat-qCDM Model using Type Ia Supernovae Observations

$$\text{Luminosity distance: } d_L = c(1+z) \int_0^z \frac{dx}{H(x)}$$

$$H(z; H_0, q) = H_0 (1+z)^{1+q}$$

$$\text{Parameters: } \mathbb{P} = \{H_0, q\}$$

## Project-10.

### Determine Mars Orbit using Kepler's Triangulation Technique

The position of object moving under the influence of a gravitational force:

$$r(\theta; \alpha, \epsilon, \omega) = \frac{\alpha}{1 + \epsilon \cos \theta - \omega}$$

Parameters:  $\mathbb{P} = \{\alpha, \epsilon, \omega\}$