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A Representative Democracy to reduce interdependency in a

multi-model ensemble

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ABSTRACT

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The collection of Earth System Models available in the CMIP5 archive represents, at least to some degree, a sample of uncertainty of future climate evolution. The presence of duplicated code as well as shared forcing and validation data in the multiple models in the archive 9 raises at least three potential problems; biases in the mean and variance, the overestimation 10 of sample size and the potential for spurious correlations to emerge in the archive due to 11 model replication. Analytical evidence is presented to demonstrate that the distribution 12 of models in the CMIP5 archive is not consistent with a random sample, and a weighting 13 scheme is proposed to reduce some aspects of model co-dependency in the ensemble. A 14 method is proposed for selecting diverse and skillful subsets of models in the archive which could be used for impact studies in cases where physically consistent joint projections of multiple variables (and their temporal and spatial characteristics) are required.

1. Introduction

Today's Earth System Models (ESMs) are great testament to collaborative scientific 19 thinking. Millions of lines of computer code represent the pinnacle of understanding of the 20 intricate coupled interactions of the Earth's land, oceanic, cryospheric, and atmospheric 21 systems. Unlike the more simple atmospheric models of the past, few people (if any) now 22 understand the models in their entirety and so the models themselves have become vehicles 23 of a scientific consensus which we use to project future climates which cannot directly be 24 validated for decades to come. For some parts, such as the representation of the equations 25 of fluid flow, understanding is mature and thus (relatively) uncontentious. But other com-26 ponents, such as the effect of a changing climate on ecosystem dynamics, are sufficiently 27 complex that any computational code must inevitably make significant approximations in 28 order to even represent the bulk behavior of the system in any tractable fashion. 29

A given model is thus more than a computer program, it is a collection of axioms and 30 beliefs about which processes might be important for evaluating how our environment might 31 change, and how those processes should be represented, and as such, each model is a self-32 consistent entity. The challenge arises, however, when one wishes to combine the results of 33 many models to attain some more comprehensive understanding of the uncertainties present 34 in their individual implementation. Given a set of models of the climate system, assessing 35 the value of adding another model clearly requires a consideration of whether the model is 36 fit for purpose (e.g. the validity of its axioms, forcing data and tuning protocols). We would 37 argue also that it is important to assess if the model provides new information; to measure 38 how independent is the new model from those in the original set. In an extreme case, adding 39 an exact duplicate of a model already in the set would not add value, rather it would bias any combination of model results towards the results of the duplicated model (Caldwell et al. 2014). 42

The latest Coupled Model Inter-comparison Project (CMIP5, Taylor et al. 2012) is the largest archive of climate data the world has seen to date. Such Multi Model Ensembles

(MMEs) have often been referred to as 'ensembles of opportunity' (Tebaldi and Knutti 2007), because the range of models represent some sample of the systematic choices which developers face in the course of representing the climate system in the form of computer code. But, as has been noted before (Knutti 2010), this sample is far from perfect.

Firstly, the models available may vary in their ability to resolve certain processes which
might be observed in the Earth System. For any given process, a researcher may find relevant
observations to rank models for their purposes but the output of the ESMs is sufficiently
high dimensional that any ranking is unlikely to be universal (Santer et al. 2009). In contrast
to weather forecast models, ESMs can also rarely be validated out of sample and so there
remains a risk that empirical components of ESMs can be calibrated using the only available
observations, and although this might be a pragmatic approach it leaves little opportunity
for assessing and contrasting model performance (Sanderson and Knutti 2012).

A second problem lies in the lack of independence of models, where independence is not 57 meant in a statistical sense but in a more loose sense of models sharing ideas for parame-58 terizations and simplifications or sharing actual computer code, and therefore being biased 59 in similar ways relative to reality. At the time of writing, 61 models are listed in the Earth System Grid database. This doesn't necessarily mean that each of these models provides an 61 independent estimate of future climate change. Indeed, some of these co-dependencies are trivial and can be accounted for by considering models submitted with different resolutions (for example, MPI-ESM-MR and MPI-ESM-LR, see Knutti et al. 2013). Most institutions also produce model variants with a range of different configurations, with options for inter-65 active atmospheric chemistry or carbon cycle (CMCC-CESM and CMCC-CM, for example). Finally, different institutions can share model components, for example the FIO-ESM model 67 shares its atmosphere, ocean, sea ice and land surface code with CCSM4, but adds a surface 68 ocean wave parameterization. Submodel replication is common throughout the ensemble, 69 for example in the models considered for this study over 25 percent use some variant of 70 the Community Atmosphere Model (CAM3, CAM3.5, CAM4 or CAM5) to represent atmospheric processes. The GFDL MOM ocean model is similarly popular (MOM2.2, MOM4.0 and MOM4.1). Table 2 shows a broad illustration of shared model components in the CMIP5 models considered for this study.

This extensive model replication in the CMIP5 and its predecessors is not a problem 75 per se, in fact it seems natural to copy successful parts and build on the work of others, and it requires enormous effort to develop entirely new model components. Hence, each institution understandably focuses on certain aspects but copies other components. But model replication presents a number of issues for model ensemble analysis. The first is simply a matter of representation: the Assessment Reports of the Intergovernmental Panel on Climate Change (IPCC) have often used the multi-model mean of the CMIP ensembles to represent a consensus view of model projections of future climate, but clearly this mean will be biased towards models which are highly replicated within the ensemble. Similarly, 83 model agreement on the sign or magnitude of a change in future climate is often taken to 84 imply confidence in a result (Tebaldi et al. 2011, Knutti and Sedáček 2013), but if models 85 are highly replicated within the ensemble, such agreement becomes less significant. 86

Another issue lies in the possible effect of replicated models in studies which attempt to constrain aspects of future climate change. If a researcher discovers a correlation between an observable quantity and some unknown climate parameter in a multi-model ensemble (such as in Fasullo and Trenberth 2012 or Qu and Hall 2013), the statistical significance of that correlation would be inflated if some points are repeated. This argument is developed in Caldwell et al. (2014) who show that although a data-mining approach will yield more strong correlations between Climate Sensitivity and potentially observable fields than one would expect to see by chance in CMIP5, this may be attributable in part to model codependencies.

This is the second in a series of papers examining interdependency in the CMIP ensembles. In Sanderson et al. (submitted), we developed a distance metric which enabled both models and observations to be represented as points in a multi-dimensional space. We then

showed that model properties could be interpolated within this space, allowing a resampling of model properties in a manner which was less sensitive to model replication and could take 100 into account a measure of performance in reproducing observations. However, the approach 101 of Sanderson et al. (submitted) is also unable to provide full spatial and temporal variations 102 in quantities. For example, a farmer may not want an estimate of the change in average 103 rainfall, but a set of representative summers with full spatial and temporal information, and 104 the corresponding temperature, sunshine and wind data. For such cases it may be better to use the raw or bias corrected model output directly, but that requires selecting a set of 106 models to use. 107

It has been proposed before that subsets of larger ensembles may produce more sta-108 tistically robust results, Evans et al. (2013) investigated this concept using subsets of a 109 multi-physics ensemble of weather forecasting models. Perhaps the simplest approach to 110 achieve this might be to take a single model from each institution, but there are numerous 111 issues with this. Firstly, although there are often similarities between models published by 112 single institutions, such a crude approach would eliminate cases where significantly different 113 models were produced by the same group. There are several examples of the latter case, 114 the GISS-E2 model, for example is published with two structurally different oceans. Fur-115 thermore, several groups (CESM, GFDL, UKMO amongst others) publish both a 'bleeding 116 edge' model and a legacy model to the archive, where there might be significant structural 117 changes between the releases. Finally, an institution-based pruning approach would not help identify models from different institutions which share a large fraction of their code.

It could be argued that one could account for many of these problems through careful consideration of model lineages, by documenting the basic parameterizations shared by different
models or by assessing the fraction of common code between different models. This, however, would be a considerable undertaking - and the results would require a comprehensive
understanding of each model's code. Firstly, although some models document and publish
their code-base in full before submitting simulations to the CMIP archive, this practice is far

from universal. A model could in theory be defined by summarizing the parameterizations,
their values and other structural assumptions which have been employed in that model, but
assessing the relative importance of each of those parameterizations in terms of model climatology or response to external forcing would require good prior intuition of the relationships
between the parameterizations and the process to be studied, which might be possible in
some but not necessarily all cases. Such an approach would clearly be worthwhile, and could
greatly aid in the interpretation of differences in climate change projections, but it would be
a monumental undertaking.

An alternative approach is to utilize output from the models themselves to establish 134 codependencies. This approach has been demonstrated with some promise by Masson and 135 Knutti (2011) and Masson and Knutti (2013), who used inter-model distances derived from 136 spatial patterns of climatological temperature and precipitation to establish a hierarchical 137 clustering of models which resembles a tree showing structural relationships one might expect 138 from considering model lineages. As noted in Masson and Knutti (2011) and Sanderson et al. 139 (submitted), the distribution of inter-model distances shows recognizable structure, with 140 models from the same institution and models with common heritage generally exhibiting 141 similar patterns of mean state bias. However, the aforementioned studies did not establish 142 any quantitative assessment of inter-model distance, which we attempt to address here. 143

To this end, we formalize an approach to use model similarity information to select models 144 based on their skill and independence. This does not eliminate model inter-dependency, but allows us to select a subset of models where the most glaring examples of model replication 146 are no longer present. In Section 2.a, we establish a method for identifying near-neighbors 147 in a climate model ensemble, in Section 2.d, we use model similarity information to produce 148 a weighting scheme which accounts for both model skill and model interdependence. Section 149 2.e shows how this framework can be used to select a subset of models from an archive of 150 climate models. Finally, Section 3.b demonstrates this method using the CMIP5 multi-model 151 archive. 152

$_{153}$ 2. Method

2.a. Processing model output

In this study, as in our accompanying paper Sanderson et al. (submitted), we produce a 155 matrix of inter-model distances in an EOF space derived from 30 year mean climatological 156 output from each model's historical simulation conducted for CMIP5. The details of the 157 construction of the distance matrix are identical to that of Sanderson et al. (submitted). We use the 'historical' and 'rcp85' experiments, and the 'r1i1p1' simulations in each case. 159 In the special case of CCSM4, we also consider the sensitivity of the technique to internal 160 variability by repeating the analysis with all available simulations in the CMIP5 archive 161 (r1i1p1, r1i2p1, r1i2p2, r2i1p1, r3i1p1, r4i1p1, r5i1p1 and r6i1p1 for the historical runs and 162 r1i1p1, r2i1p1, r3i1p1, r4i1p1, r5i1p1 and r6i1p1 for the RCP8.5 simulations). 163 The input data for this study is both processed, and used to conduct an EOF analysis 164 in a similar fashion to Sanderson et al. (submitted). Minor differences in the inter-model 165 distances occur because the former study considers both CMIP3 and CMIP5 models, which 166 slightly changes the exact form of the EOFs. For each model, a number of monthly, gridded 167 diagnostic variables are considered to represent the climatology of the model. For each avail-168 able model in the CMIP3 and CMIP5 ensembles, monthly climatologies are obtained from a 169 single historical simulation by averaging monthly mean fields for the time period 1970-2000. 170 Data is obtained for five 2 dimensional fields (surface air temperature (TAS), total precipita-171 tion (PR), outgoing top-of-atmosphere shortwave radiative flux (RSUT), outgoing longwave 172 top-of-atmosphere flux (RLUT), sea level pressure (PSL)) and two three-dimensional fields 173 (atmospheric temperature (T) and relative humidity (RH)). Three dimensional fields are 174 zonally averaged. Corresponding observational monthly mean climatologies are obtained by 175 averaging available years for each field type, as shown in Table 1. 176

Data from each model and dataset are regridded onto a 2.5 by 3.75 degree latitude longitude grid, and zonal vertical fields are regridded onto a 2.5 degree latitude grid at 17

pressure levels. For each variable, values are area weighted. Vertically resolved fields are 179 also weighted by the pressure difference between the top and bottom of the corresponding 180 level. In order to usefully concatenate the multivariate field for EOF analysis, the variables 181 must be normalized for each to represent a similar amount of variance in the multi-model 182 ensemble. We normalize each observable field using values obtained from the observations. 183 For 2 dimensional fields, we calculate the inter-monthly variance of tropical grid-cells and 184 take the average over the tropics to obtain a single normalization factor for each variable. For 3 dimensional fields, we take the inter-monthly variance of zonally averaged fields in the 186 tropics between 700 and 400 hPa, and then average the variances over the spatial domain to 187 obtain the normalization factor. Normalization factors are calculated from the observations 188 only, and the fields from each model are divided by the same factor (shown in Table 1). 189 Each field is then reformulated into a single vector. If any elements of the vector in any 190 single model or in the observations are missing, those particular elements are removed from 191 all models. Each field vector is then normalized by the number of remaining elements, and 192 the 2d and 3d fields are concatenated into a single vector length n (where n=358,248 when 193 all fields are utilized). Each of the m vectors are combined to form a matrix X^{20c} (size m 194 by n, where m is 36, comprising 36 CMIP5 model vectors). The ensemble mean value is 195 calculated by averaging the m rows of the matrix, and this is subtracted from each row to 196 yield the anomaly matrix ΔX^{20c} , such that 197

$$\Delta X^{20c} = X^{20c} - \overline{X^{20c}}. (1)$$

The analysis is also repeated with a number of different subsets of the entire set of variables. In these cases, the matrix ΔX^{20c} is formed using only that subset, and the analysis continues in the same fashion.

The process is repeated to produce a similar matrix to represent the climate change between the historical simulation (1970-2000) and the RCP8.5 simulation (2070-2100). In this second analysis, the anomaly between the two 30 year periods is taken to form the

matrix ΔX^{21c} . The future analysis is also repeated with a number of different subsets of the entire set of variables. In these cases, the matrix ΔX^{21c} is formed using only that subset, and the analysis continues in the same fashion.

207 2.b. Principal Component Analysis

We conduct a principal component analysis on the resulting matrix formed by combining 208 the climatology vectors from each participating model, such that the EOF loadings define a 209 t-dimensional space (where t is the truncation length of the Principal Component Analysis) 210 in which inter-model and observation-model Euclidean distances may be defined. The use 211 of the EOF pre-filter combines fields which are trivially correlated (such as adjacent grid-212 cells) into a single mode. The results of the analysis do change in a subtle fashion with 213 truncation length, and we discuss this sensitivity further in Section 3.c.1, but for the initial 214 analysis we use a truncation length of t=9. This truncation length effectively provides 215 enough degrees of freedom to represent some subtle differences between related models in the resulting distance metric, but not so many as to introduce excessive random noise into 217 the calculation. 218

The PCA analysis on any ΔX can be performed by singular value decomposition and truncated to t modes, such that:

$$\Delta X^{20c} = U^{20c} \lambda^{20c} V^{20c^T}, \tag{2}$$

for the present day case (20c), and

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$$\Delta X^{21c} = U^{21c} \lambda^{21c} V^{21cT}, \tag{3}$$

for the future case (21c). U^{20c} and U^{21c} (sized m by t) are matrices of model loadings, V^{20c} and V^{21c} (sized n by t) are spatial patterns of ensemble variability while λ^{20c} and λ^{21c} (sized t by t) are diagonal matrices representing the variances associated with each mode. The inter-model distances can then be measured in a Euclidean sense in the loadings matrices U^{20c} and U^{21c} , such that the distances between 2 models i and j can be expressed as:

$$\delta_{ij}^{20c} = \left(\sum_{l=1}^{t} \left(U^{20c}(i,l) - U^{20c}(j,l)\right)^2\right)^{1/2},\tag{4}$$

for the present day and

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$$\delta_{ij}^{21c} = \left(\sum_{l=1}^{t} \left(U^{21c}(i,l) - U^{21c}(j,l)\right)^2\right)^{1/2},\tag{5}$$

for the future. Model-observation distances $\delta_{i(obs)}^{20c}$ which can obviously only be calculated for the present day case are created using a climatological vector from an observational dataset X^{obs} prepared in the same fashion as X^{20c} :

$$\Delta X_{(obs)n}^{20c} = X_{(obs)} - \overline{X^{20c}} \tag{6}$$

where $\overline{X^{20c}}$ is the multi-model mean of X^{20c} , length n. This observational anomaly vector can be projected onto V^{20c} to form an observational loading vector $U_{(obs)}$ (length t). The distance between each model and the observations can be then calculated in a similar fashion:

$$\delta_{i(obs)}^{20c}(i) = \left(\sum_{l=1}^{t} \left(U^{20c}(i,l) - U_{(obs)}^{20c}(l)\right)^{2}\right)^{1/2},\tag{7}$$

Finally, we calculate the variability expected in an initial condition ensemble by taking $n_{ic} = 8$ (historical) or $n_{ic} = 6$ (future) member CCSM4 ensemble for both the historical simulation and RCP8.5. In each case, the data is processed in the same fashion as for the multi-model case to create an n_{ic} by n matrix, X_{ic}^{20c} and X_{ic}^{21c} . We then take anomalies from the CMIP5 ensemble mean:

$$\Delta X_{ic}^{20c} = X_{ic}^{20c} - \overline{X^{20c}},\tag{8}$$

and

$$\Delta X_{ic}^{21c} = X_{ic}^{21c} - \overline{X^{21c}}. (9)$$

These can also be projected onto V^{20c} and V^{21c} to form loading vectors $U^{20c}_{(ic)}$ and $U^{21c}_{(ic)}$ (size n_{ic} by t). The distance between initial condition ensemble members can be then calculated as before for the multi-model case.

244 2.c. Forming Random ensembles

In order to compare inter-model distances in the CMIP5 archive with distances expected by chance, we create a set of 10^5 matrices of random data with the same dimensions as U^{20c} and U^{21c} (where m is 36). Each random distribution represents inter-point distances for all possible pair-wise combinations m points (703 distances, in this case). Our results are not sensitive to further increasing the number of random cases.

Each row of one of these random matrices is populated with draws from a Gaussian PDF with variance equal to that from the rows of U^{20c} and U^{21c} (all of the rows have equal variance in each case). As a result, data in these random matrices is independent in directions corresponding to both the EOF number and the model number. We desire matrices with an independent model dimension in order to test the likelihood that CMIP5 output was drawn from a set of independent models. Having independence in the field direction is appropriate because the columns of U^{20c} and U^{21c} are independent by construction.

Our assumption that the t dimensional normal distribution is representative of an independent ensemble of climate projections is subject to some caveats; we are making the
effective assumption that a normal distribution of models in the space defined by U^{20c} or U^{21c} is plausible, and that there are no parts of that space which might represent an unphysical
climate state. There are some justifications for this assumption; the random distributions
are compared with the loading matrices U^{20c} and U^{21c} , which are themselves orthogonal
basis sets defined by multi-model variability. As such, we are making the assumption that

if there are physical relationships between variables in the model output data (say between 264 adjacent grid-cells or between surface temperature and outgoing longwave radiation for ex-265 ample), then any correlation between these would be represented as a single mode in the 266 EOF analysis. Thus, any linear relationships which exist in the original data are effectively 267 preserved in the random ensemble also. However, a strong nonlinear relationship between 268 two variables in the CMIP5 archive could not be represented in a single EOF mode, and 269 might be represented in two or more modes. In this case, then there would be some of 270 the space which should physically off-limits. Hence, by using normally distributed data to define the random ensembles and their associated length scale for inter-point distances, we 272 make the assumption that multi-model variability can be appropriately described by a linear basis set. Although one could potentially consider designing a random sample which fitted a 274 high-dimensional distribution to the existing ensemble to account for nonlinear relationships 275 between modes, the increase in complexity, the lack of samples in the original ensemble and 276 the necessary subjective parameterization of such a distribution means this is impractical 277 for the present study. 278

279 2.d. Weighting for Uniqueness

In this section, we seek to use the relationships derived in the Section 2.b to define a weighting scheme which would effectively down-weight closely related model pairs the ensemble, which we can assess using the expectation values for near neighbor distances in the random ensembles proposed in Section 2.c. Our scheme should also provide the capability to down-weight models which exhibit low fidelity in a desirable metric.

The limiting cases of such a scheme are easy to define. We consider the models, as before, to be represented as points in a space defined by the loadings of the model in an ensemble-wide EOF analysis. In the extreme case, if the distance between two models is exactly zero then the models are considered identical and each member of the pair should be given half the weight that they would otherwise have (equivalently, a statement that adding

290 an identical model to an existing ensemble member should not change the results).

We propose a simple functional form for model similarity which satisfies the requirements for a given model pair [i, j], separated by a distance δ_{ij}^{20c} or δ_{ij}^{21c} :

$$S(\delta_{ij}^{20c}) = e^{-\left(\frac{\delta_{ij}^{20c}}{D_u}\right)^2}$$
 (10)

$$S(\delta_{ij}^{21c}) = e^{-\left(\frac{\delta_{ij}^{21c}}{D_u}\right)^2},$$
 (11)

where D_u is a free parameter, a 'radius of similarity', such that model pairs separated by 293 less than this value are considered similar. The distance is squared so that the metric tends 294 to unity for values $\ll D_u$. The smallest reasonable value for D_u would be the expected 295 distance between two identical models exhibiting different realizations of internal model 296 variability, given this represents a case where the model structure is identical. As D_u is 297 increased from this value, increasingly distant pairs of models are considered similar. In the 298 extreme case, as D_u approaches the largest inter-point distances (i.e. the largest values of 299 δ_{ij}^{20c} or δ_{ij}^{21c}) in the ensemble, then only the models with the largest biases would exhibit a 300 value of S of close to unity and all other members would be down-weighted. 301

In Section 2.c, we derived D_u empirically by considering the nearest neighbors one would expect to find by chance in a t dimensional normal distribution of equal population, variance and dimensionality as U. This is achieved in practice by considering the randomly generated distributions from the Section 2.a. We define D_u to be the 50th percentile of nearest-neighbor distances in the 10^5 randomly generated ensembles.

One can thus obtain a value for the effective repetition of model i in the ensemble:

$$R_u(i)^{20c} = 1 + \sum_{j \neq i}^m S(\delta_{ij}^{20c})$$
 (12)

$$R_u(i)^{21c} = 1 + \sum_{j \neq i}^m S(\delta_{ij}^{21c}),$$
 (13)

for the the past and future cases respectively, where m is the total number of models.

We then propose a uniqueness weighting for model i by taking the inverse of the number of models similar to i:

$$w_u(i)^{20c} = \left(R_u(i)^{20c}\right)^{-1} \tag{14}$$

$$w_u(i)^{21c} = (R_u(i)^{21c})^{-1}.$$
 (15)

for the the past and future cases respectively. If desired, a weighting scheme could also consider model quality, a model should be given increasingly less weight the further that model lies from the point representing the observations in the EOF space. In the limiting case, the model weight should tend to zero as the distance of the model to the observations tends to infinity. These attributes are satisfied by the following construction for w_q , the model quality weighting:

$$w_q(i) = e^{-\left(\frac{\delta_{i(obs)}^{20c}}{D_q}\right)^2}, \tag{16}$$

where $\delta_{i(obs)}^{20c}$ is the Euclidean distance between the EOF loading for model i and the 317 loading of the observed climatology projected onto the same EOF basis set. This is only 318 calculated for the historical data where observations are available. D_q is a 'radius of model 319 quality', and is a free parameter in the weighting scheme. As $D_q \to +\infty$, then $w_q \to 1$ 320 for all models, and the quality weighting has no distinguishing effect. As the value of D_q 321 is reduced, models closer to the observations are increasingly up-weighted. The smallest 322 reasonable value for D_q would be the smallest observational bias seen in the ensemble (i.e. 323 $min(\delta_{i(obs)})$). In the extreme case as $D_q \to 0$, the majority of the weight is placed on the 324 single best performing model. 325

To explore the sensitivity to this parameter, we consider two values for D_q : a 'wide' choice where D_q is equal to the mean inter-model distance in the CMIP5 ensemble and a

'narrow' choice which is half of this value. Expressing D_q in terms of the CMIP variance 328 has the disadvantage that the variance itself can be influenced by both model quality and 329 reproduction, but this decision is a matter of practicality. We present the values of D_q as 330 subjective, effectively as a statement that relative skill, rather than any absolute measure, 331 should define whether we accept or reject a model. In effect, the 'wide' case describes a 332 situation where only the models with the largest biases in the ensemble are down-weighted, while in the narrow case a distinction is made between the 'average' and 'best' performers. It might be desirable to let internal or natural variability define D_q , but as we show in Section 335 3.a, this would lead to a situation where $\delta_{i(obs)}^{20c} >> D_q$ for all i, which given Equation 16, 336 would place the majority of the weight on the model with the lowest value of $\delta_{i(obs)}^{20c}$. 337

338 2.e. Eliminating interdependent models

If the researcher's goal is simply to produce a multi-model average which is less susceptible to bias by model replication, then simply weighting each model by the appropriate value of w_u would suffice. This approach could be used directly for calculating a central estimate of combined multi-model projections.

However, some issues associated with model co-dependence cannot be solved by weighting alone. For example, the potential bias associated with regression-based predictions of
unknown climate parameters can only be addressed by removing the interdependent models. This can be achieved in a pure statistical fashion (see Caldwell et al. 2014) but the
interpretation of such constructions is not always intuitive.

We propose here a less formal approach which should be readily reproducible for a variety
of purposes where it is desired to remove the most blatant model codependencies. Our
method is a step-wise model elimination, where the models with the highest co-dependencies
are removed first.

The simplest approach here would be to recursively remove a member of the closest nearneighbor pair until the remaining ensemble conforms to a plausible random distribution in the *n* dimensional EOF space. Since better models are replicated more, however, such an approach preferentially eliminates the models clustering closer to observations while models with large biases would be preserved. This has a significant detrimental effect on the mean performance of the remaining ensemble. Instead, we propose a strategy which considers both model performance and model independence when creating an ensemble subset.

Firstly, we introduce a bulk quantity which describes the ensemble characteristics, the 'independent ensemble quality score':

$$S_m^{20c} = \sum_{i}^{m} w_u^{20c}(i) w_q(i)$$
 (17)

$$S_m^{21c} = \sum_{i}^{m} w_u^{21c}(i) w_q(i), \tag{18}$$

for historical and future cases, where w_u^{20c} , w_u^{21c} and w_q are described in Section 2.d as the individual model weights corresponding to model i. Using the product of the two weights is a subjective decision, and other functional forms could potentially be explored. However, as we now demonstrate, this simple combination of the uniqueness and quality weights addresses our goals to remove the influence of exactly replicated models and of very poor models.

This can be illustrated as follows for the historical simulation: If an independent model is added to the ensemble, $w_u^{20c}(i)$ equals 1 for model i, and so S_m will increase by the model quality score, $w_q(i)$. The increase is large for a high performing model, and approaches zero for a very poor model. However, if two identical models i and j are added to the ensemble together $w_u^{20c}(i)$ and $w_u^{20c}(j)$ each equal 0.5, and so S_N will still only increase by $w_q(i)$.

If we start with an N member ensemble, we eliminate a single member by considering the maximum possible ensemble quality score for each combination of N-1 members. The excluded model j is removed from the ensemble and the process is repeated until an appropriate stopping criterion has been reached. We can assess the effective number of models remaining at any point by considering the 'number of effective models', for both 377 historical and future cases:

$$n_{eff}^{20c} = \sum_{i}^{m} w_u^{20c}(i) \tag{19}$$

$$n_{eff}^{21c} = \sum_{i}^{m} w_{u}^{21c}(i), (20)$$

each representing the sum of the uniqueness weights for the remaining models in the ensemble.

The approach outlined here is quantitative but subjective, with a number of free parameters. In order to demonstrate its utility, we consider a case study of the CMIP5 ensemble, where we can objectively demonstrate that we can use the algorithm to produce a subset of CMIP5 models which provides comparable model diversity, improved mean model performance and reduced model replication in comparison to the original model sample.

$_{55}$ 3. Results

3.a. CMIP5 Ensemble Properties

The initial dataset from which we draw our conclusion is the matrix of pairwise distances between models in the CMIP-5 archive, δ^{20c} and δ^{21c} which are calculated from U^{20c} and U^{21c} matrices. This matrix is represented graphically in Figure 1 for the all-variable case using both present day climatological fields calculated from 1970 to 2000 in historical simulations, and the anomalies from those fields in the RCP8.5 simulation between 2070 and 2100. In both cases, recognizable structure relating to model genealogy is visible in the inter-model distance field.

We can compare, in a bulk sense, the distribution of distances in the matrices to that one
might expect from a purely random distribution. The distributions for the CMIP5 derived
matrix and the random distributions are plotted in Figures 2(a) and (b) for a number of
different variable choices.

The random distributions have the same variance as the original CMIP5 distributions by
design because each dimension of the random psuedo-ensembles is normally distributed with
the same variance as the original CMIP5 case in each dimension of U_{20c} and U_{21c} . Because we
consider a large number of pseudo-random normally distributed ensembles, we can produce
best estimates and confidence intervals for the distribution of inter-model distances one
would expect if the models were normally distributed in the space defined by U_{20c} and U_{21c} .

If the CMIP5 distribution falls outside of this range, this implies that the models in CMIP5
are distributed in a non-normal fashion in the space.

We find there are some significant deviations in the CMIP5 distribution from what one 406 would expect in a purely random case. Firstly, there are a number of model pairs which lie 407 closer to each other in the EOF space than ever occurs by chance in the random samples (less 408 than 50 percent of the expected mean inter-point distance for the random case). However, 409 there is also an absence of models at intermediate distances (between 50 and 90 percent of 410 the mean inter-point distance), relative to the random distributions. This indicates that 411 the distribution of CMIP5 models in the EOF space has a rather heterogeneous, clustered 412 distribution - with families of closely related models lying close together but with significant 413 voids in-between model clusters. These features are especially clear in the future case, where 414 the distances are measured in terms of (2070-2100) anomalies from the (1970-2000) climate 415 mean state. We also show the histogram of inter-model distances in initial condition CCSM4 416 ensemble, demonstrating that inter-model distances due to internal model variability alone are an order of magnitude smaller than the mean inter-model distances seen in the CMIP5 418 archive. 419

The responsible model pairs can be explicitly plotted. Figure 3(a) shows model pairs which are closer together than the expected nearest-neighbor distances in the random distributions, using all variables. Many of these samples correspond to identical models from the same institution submitted at a different resolution (IPSL-CM5A-MR/LR, MPI-ESM-LR/MR for example). Other model pairs relate to changes in model configuration which

do not influence the set of atmospheric diagnostics considered here (HadGEM2-AO and 425 HadGEM2-ES for example share the same atmospheric, ocean and ice models, but the for-426 mer lacks treatment of the carbon cycle which has little effect in these concentration driven 427 simulations). Finally, there are some cases where models from two institutions share a large 428 fraction of code-base, and this is reflected in their proximity in EOF space (HadGEM2-AO 429 and ACCESS1-0 or FIO-ESM and BNU-ESM, for example). Several other model pairs are 430 plotted with dotted lines. These, to a lesser degree, still occur closer together than one might 431 expect by chance (for the models joined by a black line, one such pair would be expected by 432 chance in a 36 member ensemble). These connections can also be related to common model 433 components (for example, NorESM and CCSM4 share atmosphere and land surface, MPI-434 ESM and CMCC-CSM5 share atmospheric code). We also include the observational point 435 in the same analysis in Figure 3(a), which shows that none of the models in the CMIP5 436 archive are considered closer to the observations than would be expected by chance. In 437 the later part of the study, where we prune similar models from the archive, this give us 438 some confidence that similar models are not being removed because they are all converging 439 on the 'true' climate. We can repeat the analysis for future changes in the same variables 440 (Figure 3(b)), which show a similar close relationships to present day case. Using specific 441 fields produces similar (but non-identical) relationships (Figure 3(c-e)). The all-variables 442 case shows that all close relationships would be expected from a genealogical perspective. 443 However, when one uses single variables (PR especially), there are some unexpected results (e.g. MIROC and CAM5 are considered close). We attribute this to the difficulty of representing inter-model precipitation variability in a low dimensional basis set (although models from different centers may in some cases share parameterizations). 447

3.b. Stepwise model elimination

There are various arguments to support the hypothesis that the CMIP5 ensemble is biased by the inclusion of common components, some of which are featured more frequently than others. One can make this argument from a consideration of the models themselves (see Introduction and Table 2), or by examining the spatial distribution of models in orthogonal dimensions derived from model output. We have proposed a method of model removal which maximizes a metric reflecting both model diversity and fidelity. The iterative model elimination process is illustrated for the CMIP5 ensemble in Figure 4.

The plot shows the consecutive removal of models from the set of 36 considered in this study until a single model remains. The process is demonstrated by eliminating interdependent models as judged by the simulation of present day climatology. The model quality
weights w_q are obtained using the mean state climatology from the models as compared to
the observations. Model uniqueness is calculated as in Section 2.e after each iteration.

We demonstrate the sequence of model removal in Figure 4 (for present day similarities, all variables and a 'wide' quality radius). The figures show the order in which models are removed from the archive to achieve the maximum independent ensemble quality. If the removed model is closer than D_u (a function of the number of models remaining) to any other remaining model, then that model is shown to merge with its nearest neighbor. However, if the model is further than D_u from any other model, the model branch is shown as terminating in the diagram.

We have not yet fully discussed an appropriate point to stop trimming models. This
question is ultimately subjective, and the conclusion is somewhat dependent on the specific
needs of the researcher. However, Figure 5 shows some changing characteristics of the
remaining ensemble as the ensemble size is decreased, and these can be used to recommend
ensemble subsets for different scenarios. In essence, a first phase of eliminating models just
removes redundant data, a second improves the characteristics of the ensemble by removing
poor models and partly redundant ones. Going beyond that potentially worsens the ensemble
mean bias representation.

Figure 5(a) shows how n_{eff} varies as models are removed from the archive as described in Section 2.e. The actual number is dependent on the choice of D_u , the radius of similarity.

Two choices of D_u are illustrated, using either the 50th percentile of nearest-neighbor distances in the set of 10^5 random ensembles (as was used in Section 2.d) or, for comparison, the 90th percentile. Using all the models in the archive, n_{eff} is 15.5 using the larger value for D_u , or 22.5 using the smaller value (using present day climatology metrics of similarity). The removal of the first 10 models has little effect on n_{eff} (especially using the larger value of D_u). The removal of the remaining models results in a monotonic decrease in n_{eff} .

As was indicated by Figure 5(a), most of the early model eliminations have little effect on 484 n_{eff} . Figure 4 shows that many of the initial removals represent models (CCSM4 to CESM1-485 BGC, HadGEM2-ES to HadGEM2-AO, GFDL-ESM2M to GFDL-ESM2G) which are largely 486 structurally identical, at least in terms of their long term atmospheric climatology - differing 487 only in the presence of an active carbon cycle which would not influence the diagnostics 488 used in this study. It is thus largely random which member of the pair is eliminated. In 489 this regime, there is a strong inverse relationship between model quality weights (w_q) and 490 uniqueness weights (w_u) , as shown in Figure 6(a). 491

The second broad class of eliminations is models with strong connections, often from the same institutions but with some differing components. In these cases, the model with the higher value quality weighting (w_q) is generally preserved (for instance, GISS-E2-H and GISS-E2-R which differ in their ocean components). In this regime, the inverse relationship between the model quality weight and uniqueness weights is weaker (Figure 6(b)), as the clear duplicates have already been removed. Note that the uniqueness weights now refer to uniqueness within the remaining subset, and not within the full CMIP5 archive.

The final stages of removal (approximately the final 20 models) do result in a reduction in the number of effective models, illustrated by the termination of the model path. As shown in Figure 5(b), in this regime - the distribution of inter-model distances are now consistent with what one might expect from a purely random sample. Each family of closely related models is now represented, to a large extent, by its own 'champion'. Figure 6(c) shows that when only 10 models remain, the relationship between w_u and w_q is rather weak, with all remaining models having comparable uniqueness weights.

Our value judgment for an appropriate stopping criterion is thus dependent on the application. If one wishes to only remove near-identical models, one should stop trimming when the number of effective models n_{eff} begins to significantly decrease. However, if one wishes to produce the best-performing ensemble mean simulation of the mean state, it is more logical to also remove the worst performing models such that the RMSE error of the sub-ensemble mean is minimized.

$_{512}$ 3.c. Sensitivity to initial choices

The algorithm as described in Section 2.e requires several assumptions and we explore 513 the sensitivity of the results to those choices in this section. Figure 7 shows the models 514 which are retained in the analysis with a range of different initial variable and parameter 515 choices. In each case, the analysis is repeated and there is a stepwise removal of models 516 based on maximising the ensemble quality score. On each line of the plot, we show which models remain when the smallest inter-point distance in the remaining archive is first greater 518 than 50% (unfilled symbols) or 10% (filled symbols) of purely random distributions of the 519 same population, variance and dimensionality (regions marked by mid grey and dark dray 520 shading in Figure 4). Thus, we can explore the sensitivity of the retained models to our 521 initial assumptions. 522

Firstly, there is the choice of which variables are used to derive the inter-model distance matrix. To address this, we repeat the analysis with a variety of individual fields, as well as the multivariate example discussed in the previous section. The analysis is repeated for zonal mean temperature and humidity (TQ), gridded precipitation (PR), gridded Top of Atmosphere shortwave and longwave fluxes (TOA), Gridded surface air temperature (TAS) and all variables combined (ALL). Secondly, we explore the 'radius of model quality' D_q introduced in Equation 16. The analysis is repeated for two values, a 'wide' value where D_q is equal to the mean inter-model distance in the CMIP5 ensemble and a 'narrow' choice which is half of this value. The latter 'narrow' case effectively increases the role of the model quality metric, such that models with a low quality score are removed earlier in the algorithm, unlike in the 'wide' case, where highly interdependent models are removed first. Finally, we construct the model uniqueness weightings w_u using the inter-model distances derived from the 30 year mean 1970-2000 present day data in the 'present' case, but use the anomaly between 2070-2100 and 1970-2000 for the 'future' case.

We find that variable choice has little impact on the final choice of model subsets. Al-537 though in some cases, the choice of model from a given institution can change, the overall 538 number of models retained is similar for each of the variable choices. The use of the 'narrow' radius of model quality, however, significantly decreases the number of retained models with 540 respect to the 'wide' value. This can be explained by considering that the narrow setting 541 increases the ratio of the model quality weighting for models lying close to the observations, 542 and those far away. In the 'narrow' regime, the ensemble quality score is best maximised 543 by removing the poorly performing models earlier in the analysis, and thus after the inter-544 dependent remaining models have been removed, the number of remaining unique models is 545 smaller than in the 'wide' case. 546

3.C.1) EOF TRUNCATION CHOICES

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Some subjective decisions are required in the interpretation and subsequent usage of the PCA conducted in Section 2.a, and we discuss these at greater length here. In previous studies like Masson and Knutti (2011), the inter-model distances were calculated without the PCA stage, simply calculating distances in the space defined by the anomaly matrices, ΔX^{20c} and ΔX^{21c} . For the purposes of this study, and its companion studies (Sanderson et al. submitted), it is neccessary to decrease the dimensionality (and co-dependence) of the data in order to establish prior expectations of near-neighbor distances.

In this study, as in Sanderson et al. (submitted), the inter-model distances are calculated with the truncated set of 9 modes. The resulting inter-model distance matrix calculated with

 U^{20c} truncated to 9 modes has a 0.93 correlation with the matrix one would calculate using the full-field matrix ΔX^{20c} , but using the orthogonal basis set allows us to form random matrices with which to compare the results (Figure 2).

For smaller values of t, only the leading patterns of model difference are retained, which 560 results in large inter-model distances between different model families (e.g. CESM and GFDL 561 models) and very small distances between models in the same family (e.g. CESM-CAM5 and CESM-CAM4). With such few degrees of freedom, very small intermodel distances cannot be ruled out by chance in the random ensembles, and so no models can be excluded from the 564 ensemble (see Figure 8 for truncation values of 3 or less. The analysis produces very similar results, and the minimum number of retained models, for values of t between 8 and 12 (see 566 Figure 8), with relatively little sensitivity to variable choice (not shown). For values of t of 15 567 or greater, the higher order modes increasingly represent subtle and often noisy differences 568 between models in the archive, which inflates the distance between the near-neighbors in the 569 ensemble. Hence, once again we see fewer models ruled out. 570

To test the sensitivity of the inter-model distance matrix to variable choice, we also repeat 571 the EOF analysis with a number of different subsets of diagnostic variables. The resulting 572 correlation depends significantly on which exact variable is retained. The inter-model dis-573 tances calculated using gridded surface temperature only ('TAS') are highly correlated with 574 the multi-variate case (R=0.95, untruncated). Top of atmosphere radiative fluxes (RAD, 575 R=0.85 untruncated), Total Precipitation (PR, R=0.66 untruncated), and zonally averaged vertical temperature and humidity (QT, R=0.42 untruncated) are increasingly poorly cor-577 related with the full field multi-variate case. This implies that some fields, such as surface 578 temperature have sufficient information to render a multi-variate approach unnecessary. 579

With a truncation length of 9, which we used for the bulk of this study, the resulting
distance matrix remains highly correlated to the full field distance matrix, but the influence
of covariant fields and models is reduced (see Caldwell et al. 2014 for an extensive discussion
of these issues).

84 3.d. Ensemble Mean Performance

The results of Section 3.b suggest that eliminating the strongest interdependent models to leave a plausibly random distribution would leave between 10 and 25 of the 36 CMIP5 models considered here (depending on variable and parameter choices). Trimming the ensemble to its more independent subset does not worsen the fidelity of the climatological mean result, and removing the poorer performing outliers (models with large biases) can actually improve it, as we show in this section.

We can first examine how the multi-model mean of present day climatology compares against observations. Figure 5(c) considers the Root Mean Square Errors (RMSE) of various weighted and unweighted multi-model means calculated using the same multi-variate climate state vectors described in Section 2.a and the observations listed in Table 1. We illustrate this using the 'ALL' variable case, with the 'wide' radius of model quality and present day derived inter-model distances. We also compare with the average RMSE seen when a completely random sample (without replacement) of the same size is taken, as compared to the detailed technique outlined in Section 3.b.

If one considers only the far left of the plot, where all 36 models are retained, weighting
the models by uniqueness actually increases the RMSE. This is largely to be expected - as we
have seen in Figure 6(a) that the best performing models have the lowest uniqueness weights.

It also suggests that a mean of the CMIP5 ensemble is already weakly weighted towards the
better performing models. If we explicitly weight the model mean towards models which lie
closer to the observations in the EOF space, the RMSE can be reduced significantly.

As the first 10 (highly interdependent) models are removed from the archive, the simple mean RMSE increases slightly while the random draw RMSE remains constant, likely because the high-performing models have less representation when the duplicates have been pruned. The uniqueness weighted mean also becomes more similar to the simple mean case (u_w) is now more consistent across the ensemble). Between 28 and 12 models remaining, the simple RMSE decreases significantly and when 20 models remain, the subset outperforms the

RMSE of the random sample. The lowest RMSE values occur with between 12 and 5 models remaining. Removing any further models increases the RMSE of the simple multi-model mean. With 5 or fewer models remaining, all models have a high value of both w_u and w_q , so weighting by uniqueness or quality has little effect. In all cases, any further removal of models (below 5) significantly increases the RMSE, a fact which is likely attributable to the Cauchy-Schwartz inequality (Annan and Hargreaves 2011).

₇ 4. Discussion and Conclusions

The present study considers how one might remove potential biases which might arise from shared components in the CMIP5 archive of climate models, and its predecessors. We also propose some simple diagnostics which might be used to identify interdependent models using model diagnostic output, and a possible strategy to choose a model subset to maintain model diversity without replication and to incorporate model quality information into this decision.

This study represents a proof of concept; the choice of diagnostics used in this study are 624 of course arbitrary, to some degree, though the results of which models are interdependent do seem to be relatively resilient to changes in variable and time period (see Figure 3, Pennell and Reichler 2011 and Knutti et al. 2013). However, we do assume that a model's mean state 627 climatology can be used to assess both its skill and independence. Clearly, if our final goal 628 is to assess the plausibility of a model's future simulations then the mean state simulation 629 is not a perfect assessment of model skill, although it could be argued that it is a neccessary 630 condition and as such a weighting strategy based on present day climatology can be justified 631 in the absence of any additional information. 632

Certainly, which model exhibits the highest quality score is very much dependent on the specific metrics in which the researcher might be interested (Santer et al. 2009), and it is far beyond the scope of this study to conduct an exhaustive comparison of possible model metrics. In this study, we have focussed primarily on diagnostic output from the atmospheric model, and our results are thus liable to be most sensitive to common component in that model. As such, the results of this study should be interpreted as illustrative of a potential method for reducing the effects model interdependency, and not as a prescriptive list of models which should be used for future studies. Most studies based on CMIP5 could easily use such a framework, but the value judgements of future researchers should be embedded into the choice of metric used to assess model similarity and quality.

We assess the likelihood of near-neighbor models occurring by chance using a large num-643 ber of random distributions of the same dimensionality as the truncated orthogonal set of EOF loadings we derive from the original ensemble. The random sample is not a proxy for 645 the space which might be attainable by the real climate, rather it is a proxy for the distri-646 bution of models represented in an orthogonal basis set defined by multi-model variability. 647 As such, we are making the assumption that if there are physical relationships between vari-648 ables in the model output data (say between surface temperature and outgoing longwave 649 radiation), then any correlation between these would be represented as a single mode in the 650 EOF analysis. However, if there exists a strong nonlinear relationship between two variables 651 in the CMIP5 archive then this relationship could not be represented in a single EOF mode, 652 and might be represented in two or more modes. In this case, then the distribution of models 653 in the space could be more complex than a simple Gaussian. One could imagine designing 654 a random sample which fitted a high-dimensional distribution to the CMIP5 ensemble to account for such nonlinearities, but the increase in complexity, the lack of samples in the 656 original ensemble and the neccessary parameterization of such a distribution means this is 657 impractical. 658

We also assume, by drawing random samples using the variance defined by the original ensemble, that none of the CMIP5 members can be ruled out *a priori*. One could imagine a situation where an arbitrarily poor model was included in the ensemble which would increase the variance represented in each mode such that any realistic models would look self similar

and would be down weighted by the uniqueness weighting. Therefore, the method only
makes sense if there is some level of base confidence that none of the models in the archive
are completely unrepresentative of the true system. But, we would argue that this is true
of any analysis which uses the CMIP5 archive and that even a simple mutli-model mean is
subject to a sanity check of the participating models.

Caveats aside, this study illustrates some interesting characteristics of the CMIP5 archive and potential issues which might arise from treating this archive as a random sample of possible climate models. There is extensive replication of model code in the archive, primarily 670 within institutions but also in some cases between institutions (see Table 2). This should 671 come as little surprise, a quick examination of AOGCM makeup in the CMIP5 models 672 indicates that some individual components are used by over 25 percent of the archive. But, 673 we show in this study (like in Masson and Knutti 2011 and others) that many of those 674 similarities can be identified also through a simple analysis of model output. A more detailed 675 discussion of shared model components is given in the supplementary material of Knutti et al. 676 (2013).677

Similarities in diagnostic output are not always predictable from a consideration of model 678 construction alone. One can find examples of cases with significant changes in code-base, 679 but with minor changes in diagnostic similarity. For example, CCSM4 and CESM1-CAM5 680 have significantly different aerosol schemes, dynamics, cloud microphysics and yet our results 681 show the two models as very strongly related when considering the distribution of inter-model 682 distances. This indicates that tuning strategies and non-atmospheric components may play a significant role in diagnostic model similarity, even when primarily atmospheric output is 684 used to assess inter-model distance. This implies that although the diagnostic output is a 685 useful indicator of model similarities, those similarities may not be a function of shared code 686 alone. The climateprediction.net (Stainforth et al. 2005) and QUMP (Murphy et al. 2007) 687 experiments, for example show that considerable diversity in model behavior is achievable 688 through parameter perturbation alone with an identical codebase. 689

There are several possible additional factors which might influence diagnostic similarity. 690 Firstly, the tendency for various generations of models from a single institution to exhibit 691 strong similarities in spite of extensive model component changes (see Figure 2 in Sanderson 692 and Knutti 2012 with reference to CESM, GFDL or Hadley Centre models) indicates that 693 some elements of model calibration tend to cluster models from a given modeling center. The 694 reasons for this clustering have multiple possible candidates which could lie in institutional policy or regional focus (institutions might be more concerned with their model's performance in the region's climate). Standard metrics used to judge model performance during the model 697 development process or preferred observational datasets may also vary from institution to 698 institution. Secondly, models rarely change all components at the same time, so we would 699 posit that evaluating when a model is 'new' is a subjective matter. Finally, the CMIP5 700 protocol allows for some flexibility in the way that models implement external forcings - so 701 different groups, even with identical models, can choose to represent the historical and future 702 boundary conditions in different ways to produce differences in the simulated climate. Knutti 703 et al. (2013) see similar relationships in control simulations, but one cannot exclude the 704 possibility that the control simulations themselves might also include common assumptions 705 on boundary conditions. 706

In summary, we confirm earlier arguments that models are not independent, some are 707 essentially duplicates, and the effective number of independent models based on this method 708 is less than half of the actual number of models, consistent with earlier studies (Jun et al. 2008, Annan and Hargreaves 2011, Sanderson and Knutti 2012). Some models are closer to observations than others (Gleckler et al. 2008, Knutti and Sedáček 2013). We believe that our 711 method, and results do not strongly hinge on the way in which one interprets the ensemble 712 as 'truth centered' (Knutti 2010), 'indistinguishable from truth' (Annan and Hargreaves 713 2011, Rougier et al. 2013) or neither (Sanderson and Knutti 2012, Bishop and Abramowitz 714 2013). One could imagine a hypothetical ensemble following any of these frameworks, and 715 by duplicating some of its members, bias would be introduced in the ensemble distribution. 716

By evaluating our ensemble subset performance in terms of ensemble mean performance, we do not necessarily advocate a truth centered ensemble, as the ensemble mean would also be the best estimate of future change in the indistinguishable case.

There are of course different ways to account for model performance and interdependence. 720 In the companion paper (Sanderson et al. submitted), we proposed a method to produce 721 probabilistic estimates that are largely insensitive to model duplicates and can consider 722 model performance. However, when high dimensional data and/or spatially and temporally 723 consistent fields are required (e.g., for impact models), a fully probabilistic method becomes 724 unwieldy and might even hinder the development of tractable impact analyses (Dessai and 725 Hulme 2004). Bishop and Abramowitz (2013) also proposes an alternative technique where models in the archive are subject to a linear transformation, where the weighted mean of 727 transformed models is calculated to be optimally close to an observed climate. This transfor-728 mation and weighting can then be extrapolated for future projections. This method has the 729 advantage that the resulting transformed models have independent errors, and weight future 730 projections by climatological skill. However, the transformed models are not, themselves 731 physically self-consistent and there is a potential for simulations to be over-fitted to histor-732 ical data in a manner which could potentially result in overconfident future projections. In 733 comparison, the method we present here preserves a subset of self-consistent physical models 734 (for both present day and future projections), and although they might not be independent 735 in the strict sense of orthogonality, this subset can be simply used for almost any application 736 or analysis. 737

We thus propose that there is significant utility in spanning the potential uncertainty in future climate by representing spread with an appropriate subset of models. This study introduces weights which assess model uniqueness and model climatology fidelity. We find that the two were inversely related such that the models with the best simulations of the present day climate were also least unique. A part of this is possibly due to the fact that models have been calibrated by the observations, and will thus appear to cluster around

those observations (and each other). But, a closer examination reveals that a large fraction
of the high-scoring models' lack of uniqueness can be explained by other models which have
duplicated some, or all of their code. When these duplicates are removed, this strong inverse
relationship is weakened (but not entirely eliminated).

This property of the ensemble is clearly to some extent contingent on the choice of 748 metrics used, but it does raise a potentially interesting property of the ensemble; that the best performing models might also be the most promiscuous. This situation implies that 750 the ensemble as a whole is already strongly weighted towards the better performing models. We show that if the models are weighted to reward their uniqueness, then the RMSE of 752 the ensemble mean is increased. Thus, through a mechanism of quasi natural selection, the 753 climate community has created an ensemble of models which has already up-weighted its 754 climatologically best performing members. In other words, relying on model democracy is 755 to some degree upweighting skilled model structures without deliberately thinking about it 756 or discussing it, by the mechanism of duplication of well-proven code. 757

This could be seen as an argument in support of keeping the entire ensemble when 758 performing an analysis, and at least some justification that the multi-model mean result is 759 a defensible best estimate. But, it is at best an accidental property that is not guaranteed 760 to remain in future ensembles, and may not at all be visible for more specific questions or 761 metrics. Whether a model is extensively duplicated is not a pure function of its quality or 762 fidelity. A sub-model with open source code and few restrictions on its use is more likely to be utilized by another group than another model with a closed-source policy. However, a 764 model which is jointly used by a large number of groups also has a large development pool 765 invested in improving that model. Duplication within institutions depends also on funding 766 and the available computing resources. One could make the argument that the CMIP5 767 ensemble distribution and the social and intellectual landscape of the climate community 768 are surely related, but certainly not in any simple fashion. 769

A question also remains of whether the original CMIP5 ensemble is sufficient to assess

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systematic uncertainty in future climate change. This question could easily form a study in itself, but our results are somewhat informative in this matter. Firstly, the number of truly independent models in the archive is significantly less than the number of submitted models, when gauged by model output. Hence, adding another model to the existing archive has most value if the developers introduce novel components and assumptions. It is true that exploring different configurations of existing components through sub-model exchange or parameter perturbation can certainly modify model behavior, and we would argue that such experiments should continue in order to fully explore the inherent uncertainties in the existing model set.

However, this uncertainty is conditional on the number of independent models available 780 to us, and establishing whether the current set is sufficient is a question which might not 781 be a useful, because there is not a convenient space in which systematic model assumptions 782 can be defined. For example, the current CMIP5 ensemble might have n fundamentally 783 different convection schemes, each with its own advantages and biases, but nobody would 784 argue that this constituted a "full set". Where there is approximation and parameterization, 785 there are potentially limitless ways to address this. And because nobody can know the 786 behavior of the $n+1^{th}$ model, the question of ensemble adequacy cannot be answered in 787 a strict sense. Within the ensemble we have, we can tractably experiment with subsetting 788 to assess how many models are required to have confidence in the distribution of future 789 climate change formed by the full set, but we can never know if the $n+1^{th}$ model will adopt different assumptions or resolve a new process to place its projection outside of the existing distribution. 792

We argue that a joint consideration of model similarity and quality metrics allows the researcher to make use of a more quantitatively defensible sample of simulations available in the CMIP archives, either through weighting or by model elimination (in itself, an extreme form of weighting) to produce a best estimate of combined model projections. Our approach for achieving this can be controlled with a small number of subjective but clearly defined

parameters, which can potentially mitigate some of the arbitrary sampling issues which arise from relying on model democracy, and can be tailored to specific questions by choosing appropriate metrics and datasets.

It should be noted in this discussion that the CMIP5 archive is not a full representation 801 of the uncertainty space for GCM projections. Rather, it is a collection of intended 'best 802 possible models', the final iterations of their respective tuning processes as model developers 803 calibrate their parameterization choices to best represent the observed climate properties which they find most important, although there may be other acceptable configurations (Mauritsen et al. 2012). Clearly, these choices and targets will vary from model to model, but 806 the fact that there are implicitly a near-infinite number of rejected parameter configurations for each model must be remembered when trying to interpret the significance of the spread 808 of simulations in the archive. In a practical sense, we ignore these rejected configurations 809 because we do not have access to them. In addition, there is some evidence to suggest 810 that the model diversity one can attain by structural changes significantly exceeds that of 811 parameter changes in currently available Perturbed Parameter ensembles (Yokohata et al. 812 2013). Nevertheless, it should be remembered that both the CMIP5 ensemble (and by 813 definition our subsets of that ensemble) is already a subset of all possible model configurations 814 which have been chosen by model developers. 815

There are some cases where we would argue it is essential to eliminate interdependent 816 models, such as when a correlation found in the multi-model ensemble is used as a constraint on a climate parameter (such as for climate sensitivity in Fasullo and Trenberth 2012, or for 818 high latitude surface albedo feedbacks in Hall and Qu 2006). The presence of closely related, 819 or even identical models in the archive would tend to artificially inflate the significance 820 of any correlation simply because identical models would exhibit similar values for both 821 the predictor and for the unknown quantity (Caldwell et al. 2014). Removing the obvious 822 interdependent models as shown in this study would certainly be better than assessing a 823 correlation based on the entire archive, but a method for achieving this in a strict statistical 824

sense is presented in Caldwell et al. (2014).

There is a danger that as models improve, the better models have the potential to con-826 verge on the 'true' climate state, which might lead to their elimination if interdependent 827 models are removed. We show in Figure 3 that this is unlikely to be the case for CMIP5, 828 given none of the models lie close enough to the observations to be influenced by the unique-829 ness weighting. However, one could imagine if a small group of models make a real advance 830 which removes a long-standing systematic bias (for example, as some models begin to explicitly resolve convection), then it would be necessary to accept a higher level of similarity 832 among the better performing models (i.e. the uniqueness weighting u_w could no longer be independent of the skill weighting u_s). 834

Proposing a subset of models to consider for a less biased analysis could be seen as overly 835 prescriptive, but our aim is not to focus on the exact set of models which should be used 836 for future studies, rather to establish a framework in which researchers could make their 837 selection based upon metrics which are most relevant to their question. We would argue 838 that although the collection of models which arise from the 'ensemble of opportunity' is 839 often seen as sacrosanct, the democratic policy of one-model, one-vote is no longer a logical 840 one in the increasingly complex family tree of models available to the researcher. A subset 841 of 10-20 models that are reasonably independent and perform well for the criteria that are 842 judged to be relevant is very likely to be more skillful than the full ensemble. Giving equal 843 weight to all models which have completed a simulation of interest is, albeit implicitly, adopting a weighting scheme which rewards model components which are highly replicated. 845 This weighting scheme might fortuitously have the property of rewarding the most skilled 846 components but, we would argue, this property should be demonstrated and the decision 847 how to incorporate it should be made consciously. 848

5. Acknowledgments

We acknowledge the World Climate Research Programme's Working Group on Coupled 850 Modeling, which is responsible for CMIP, and we thank the climate modeling groups for 851 producing and making available their model output. For CMIP the U.S. Department of 852 Energy's Program for Climate Model Diagnosis and Inter-comparison provides coordinating 853 support and led development of software infrastructure in partnership with the Global Or-854 ganization for Earth System Science Portals. Portions of this study were supported by the 855 Office of Science (BER), US Department of Energy, Cooperative Agreement No DE-FC02-856 97ER62402. The National Center for Atmospheric Research is sponsored by the National 857 Science Foundation. We would also like to thank our anonymous reviewers for their extensive 858 and insightful comments. 859

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REFERENCES

- Adler, R., et al., 2003: The version 2 global precipitation climatology project (GPCP)
 monthly precipitation analysis (1979-present). Journal of Hydrometeorology, 4 (6), 1147–
 1167.
- Annan, J. and J. Hargreaves, 2011: Understanding the CMIP3 Multimodel Ensemble. *Jour-nal of Climate*, **24** (**16**), 4529–4538.
- Aumann, H. H., et al., 2003: Airs/amsu/hsb on the aqua mission: Design, science objectives,
 data products, and processing systems. *IEEE TRANSACTIONS ON GEOSCIENCE AND*REMOTE SENSING, 41 (2), 253.
- Bishop, C. H. and G. Abramowitz, 2013: Climate model dependence and the replicate earth paradigm. *Climate Dynamics*, **41** (3-4), 885–900.

- Brohan, P., J. Kennedy, I. Harris, S. Tett, and P. Jones, 2006: Uncertainty estimates in
- regional and global observed temperature changes: a new dataset from 1850. J. Geophys.
- 874 Res., **111** (**D12**), D12 106.
- Caldwell, P. M., C. S. Bretherton, M. D. Zelinka, S. A. Klein, B. D. Santer, and B. M.
- Sanderson, 2014: Statistical significance of climate sensitivity predictors obtained by data
- mining. Geophysical Research Letters, 41 (5), 1803–1808.
- Dessai, S. and M. Hulme, 2004: Does climate adaptation policy need probabilities? *Climate policy*, **4** (2), 107–128.
- Evans, J. P., F. Ji, G. Abramowitz, and M. Ekström, 2013: Optimally choosing small
- ensemble members to produce robust climate simulations. Environmental Research Letters,
- 8 **(4)**, 044 050.
- Fasullo, J. T. and K. E. Trenberth, 2012: A less cloudy future: The role of subtropical subsidence in climate sensitivity. *Science*, **338** (6108), 792–794.
- Gleckler, P., K. Taylor, and C. Doutriaux, 2008: Performance metrics for climate models. *J. Geophys. Res.*, **113**, D06 104.
- Hall, A. and X. Qu, 2006: Using the current seasonal cycle to constrain snow albedo feedback in future climate change. *Geophys. Res. Lett.*, **33**, L03 502, doi:10.1029/2005GL025127.
- Jun, M., R. Knutti, and D. W. Nychka, 2008: Local eigenvalue analysis of CMIP3 climate model errors. *Tellus A*, **60** (5), 992–1000.
- Knutti, R., 2010: The end of model democracy? Climatic Change, 102 (3-4), 395–404.
- Knutti, R., D. Masson, and A. Gettelman, 2013: Climate model genealogy: Generation

 CMIP5 and how we got there. *Geophysical Research Letters*, **40** (6), 1194–1199.
- Knutti, R. and J. Sedáček, 2013: Robustness and uncertainties in the new CMIP5 climate model projections. *Nature Climate Change*, **3 (4)**, 369–373.

- Masson, D. and R. Knutti, 2011: Climate model genealogy. Geophys. Res. Lett, 38 (8), L08 703.
- Masson, D. and R. Knutti, 2013: Predictor screening, calibration, and observational con-
- straints in climate model ensembles: An illustration using climate sensitivity. Journal of
- 900 Climate, **26** (3), 887–898.
- Mauritsen, T., et al., 2012: Tuning the climate of a global model. Journal of Advances in
- Modeling Earth Systems, 4 (3), M00A01.
- 903 Murphy, J., B. Booth, M. Collins, G. Harris, D. Sexton, and M. Webb, 2007: A methodology
- for probabilistic predictions of regional climate change from perturbed physics ensembles.
- Philosophical Transactions A, 365 (1857), 1993.
- NASA, 2011: CERES EBAF Data Sets. Available online, URL http://eosweb.larc.nasa.
- gov/PRODOCS/ceres/level4_ebaf_table.html.
- Pennell, C. and T. Reichler, 2011: On the effective number of climate models. Journal of
- 909 Climate, **24** (9), 2358–2367.
- Qu, X. and A. Hall, 2013: On the persistent spread in snow-albedo feedback. Climate Dy-
- 911 namics, 1–13.
- Rougier, J., M. Goldstein, and L. House, 2013: Second-order exchangeability analysis for
- multimodel ensembles. Journal of the American Statistical Association, 108 (503), 852–
- 914 863.
- 915 Sanderson, B., R. Knutti, and P. Caldwell, submitted: Addressing interdependency in a
- multi-model ensemble by interpolation of model properties. Journal of Climate.
- Sanderson, B. M. and R. Knutti, 2012: On the interpretation of constrained climate
- model ensembles. Geophysical Research Letters, 39 (16), L16708, doi:DOI:10.1029/
- 919 2012GL052665.

- Santer, B., et al., 2009: Incorporating model quality information in climate change detec-
- tion and attribution studies. Proceedings of the National Academy of Sciences, 106 (35),
- 922 14778–14783.
- 923 Stainforth, D. A., et al., 2005: Uncertainty in predictions of the climate response to rising
- levels of greenhouse gases. *Nature*, **433** (**7024**), 403–406.
- Taylor, K. E., R. J. Stouffer, and G. A. Meehl, 2012: An overview of cmip5 and the experi-
- ment design. Bulletin of the American Meteorological Society, 93 (4), 485–498.
- Tebaldi, C., J. M. Arblaster, and R. Knutti, 2011: Mapping model agreement on future
- cliamte projections. Geophys. Res. Lett., 38, L23 701, doi:10.1029/2011GL049863.
- Tebaldi, C. and R. Knutti, 2007: The use of the multi-model ensemble in probabilistic climate
- projections. Philosophical Transactions of the Royal Society A: Mathematical, Physical and
- Engineering Sciences, **365** (1857), 2053–2075.
- Yokohata, T., et al., 2013: Reliability and importance of structural diversity of climate model
- ensembles. Climate Dynamics, 41 (9-10), 2745–2763.

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936		in this effort were acquired as part of the activities of NASA's Science Mission	
937		Directorate, and are archived and distributed by the Goddard Earth Sciences	
938		(GES) Data and Information Services Center (DISC)."	40
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Table 1. Observational Datasets used as 'observations throughout. * "The data used in this effort were acquired as part of the activities of NASA's Science Mission Directorate, and are archived and distributed by the Goddard Earth Sciences (GES) Data and Information Services Center (DISC)."

Field	Source	Reference	Years	Global normalization
TS	HadCRUT3	Brohan et al. (2006)	1970-2000	2.09 K
PR	GPCP	Adler et al. (2003)	1979-2001	$30.1 \ Wm^{-2}$
RSUT	CERES-EBAF	NASA (2011)	2000-2005	$25.8 \ Wm^{-2}$
RLUT	CERES-EBAF	NASA (2011)	2000-2005	$3.32 \ mm/day$
T	AIRS*	Aumann et al. (2003)	2002-2010	$0.28 \ K$
RH	AIRS*	Aumann et al. (2003)	2002-2010	12.12 %

TABLE 2. Submodel components for the 38 CMIP5 models considered in this study.

Model		TABLE	z. Dubino	ter components		TABLE 2. Submodel components for the 30 Civil 9 models considered in this study.
CAM4 CLM4 MICOM-HAMOCC CICE MRL-AGCM3 CLM4 MICOM-HAMOCC CICE MRL-AGCM3 HAL MRI-COM3 CICE ECHAM6 JSBACH MPIOM Bitz/Lipscomb FECGC-AGCM MATSIRO CCSR-COCO Bitz/Lipscomb FRCGC-AGCM MATSIRO CCSR-COCO Bitz/Lipscomb FRCGC-AGCM MATSIRO CCSR-COCO Bitz/Lipscomb FRCGC-AGCM MATSIRO CCSR-COCO Bitz/Lipscomb LMDZ ORCHIDEE NEMO-OPA NEMO-LIM LMDZ TRIFFID HadGOM2 NEMO-LIM HAGAMIL 20 CCLM3 HAGGOM2 HAGCOM2 GISS HAGCOM2 HAGCOM2 SIS GFDL-AM2 TRIFFID HAGCOM2 SIS	Model	Atmosphere	Land	Ocean	Ice	Source
CAM4 CLM4 MICOM-HAMOCC CICE MRL-AGCM3 HACL MRLCOM3 CICE ECHAM6 JSBACH MPIOM Bitz/Lipscomb FRCGC-AGCM MATSIRO CCSR-COCO Bitz/Lipscomb I MDZ ORCHIDEE NEMO-OPA NEMO-LIM I LMDZ ORCHIDEE NEMO-OPA NEMO-LIM I INMCM INMCM INMCM INMCM I INMCM INMCM INMCM INMCM I RAGGAMI TRIFFID HadGOM2 ARGCMI HAdGAM2(N96L38) TRIFFID HadGOM2 ARGCMI HAGGAM2(N96L38) TRIFFID HadGOM2 ARGCMI GISS GISS <td< td=""><td>NorESM1-ME</td><td>CAM4</td><td>CLM4</td><td>MICOM-HAMOCC</td><td>CICE</td><td>https://verc.enes.org/ISENES2/models/earthsystem-models/ncc/noresm</td></td<>	NorESM1-ME	CAM4	CLM4	MICOM-HAMOCC	CICE	https://verc.enes.org/ISENES2/models/earthsystem-models/ncc/noresm
MRI-AGCM3	NorESM1-M	CAM4	CLM4	MICOM-HAMOCC	CICE	https://verc.enes.org/ISENES2/models/earthsystem-models/ncc/noresm
ECHAM6	MRI-CGCM3	MRI-AGCM3	HAL	MRI.COM3		http://www.mri-jma.go.jp/Publish/Technical/DATA/VOL_64/index_en.html
Dechambe JSBACH MPIOM Bitz/Lipscomb FRCGC-AGCM MATSIRO CCSR-COCO Bitz/Lipscomb ERCGC-AGCM MATSIRO CCSR-COCO Bitz/Lipscomb Implementary MATSIRO CCSR-COCO Bitz/Lipscomb MATSIRO CCMAIDEE NEMO-OPA NEMO-LIM Implementary NEMO-OPA NEMO-LIM Implementary	MPI-ESM-MR	ECH AM6	JSBACH	MPIOM		http://www.mpimet.mpg.de/en/science/models/mpi-esm.html
FRCGC-AGCM	MPI-ESM-LR	ECHAM6	JSBACH	MPIOM		https://www.enes.org/models/system-models/mpi-m/mpi-esm
FRCGC-AGCM	MIROC5	FRCGC-AGCM	MATSIRO	CCSR-COCO	Bitz/Lipscomb	http://journals.ametsoc.org/doi/full/10.1175/2010JCLI3679.1
FRCGC-AGCM MATSIRO CCSR-COCO Bitz/Lipscomb	MIROC-ESM-CHEM	FRCGC-AGCM	MATSIRO	CCSR-COCO	Bitz/Lipscomb	http://www.wcrp-climate.org/wgcm/WGCM15/presentations/21Oct/KIMOTO_Japan.pdf
LMDZ (CM4) ORCHIDEE NEMO-OPA NEMO-LIM LMDZ ORCHIDEE NEMO-OPA NEMO-LIM LMDZ ORCHIDEE NEMO-OPA NEMO-LIM INMCM INMCM INMCM INMCM INMCM INMCM INMCM INMCM GAMIL 2.0 CLM3 LICOM2 CICE4-LASG HadGAM2 (N96L38) TRIFFID HadGOM2 CICE4-LASG HadGAM2 (N96L38) TRIFFID HadGOM2 CICE4-LASG GISS GISS Russell HYCOM GISS GISS HAGGOM2 SIS GISS GISS HAGGOM2 SIS GFDL-AM2.1 LM3 MOM4.1 SIS GFDL-AM2.1 LM3 MOM4.1 SIS GCDLAM3 POP2 CICE4 AGCM4 CLASS NOAR SIS GCM4 POP2 CICE4 GCAM4 CLM4 POP2 CICE4 CAM4 CLM4 POP2 CICE4 C	MIROC-ESM	FRCGC-AGCM	MATSIRO	CCSR-COCO	Bitz/Lipscomb	http://www.wcrp-climate.org/wgcm/WGCM15/presentations/21Oct/KIMOTO-Japan.pdf
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LMDZ	IPSL-CM5A-MR	LMDZ	ORCHIDEE	NEMO-OPA	NEMO-LIM	http://icmc.ipsl.fr/index.php/icmc-models/icmc-ipsl-cm5
INMCM	IPSL-CM5A-LR	LMDZ	ORCHIDEE	NEMO-OPA	NEMO-LIM	http://icmc.ipsl.fr/index.php/icmc-models/icmc-ipsl-cm5
HadGAM2 (N96L38) TRIFFID HadGOM2	INMCM4	INMCM	INMCM	INMCM	INMCM	http://link.springer.com/article/10.1134%2FS000143381004002X
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HadGAM2 (N96L60) TRIFFID	HadGEM2-ES	HadGAM2 (N96L38)	TRIFFID	HadGOM2		http://cms.ncas.ac.uk/wiki/UM/Configurations/HadGEM2
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GISS	GISS-E2-R	GISS	GISS	Russell	Russell	$\mathrm{http:}//\mathrm{data.giss.nasa.gov/modelE/ar5}/$
GFDL-AM2.1 LM3 MOM4.1 SIS	GISS-E2-H	GISS	GISS	HYCOM	HYCOM	http://data.giss.nasa.gov/modelE/ar5/
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CAM3.5 CLM3 POPP2 CICE4	GFDL-CM3	GFDL-AM3	LM3	MOM4.1	SIS	http://www.gfdl.noaa.gov/earth-system-model
AGCM4 CLASS NCAR	FIO-ESM	CAM3.5	CLM3	POP2	CICE4	http://www.wcrp-climate.org/wgcm/WGCM15/presentations/21Oct/WANG-WGCM.pdi
Gordon CABLE MOM2.2 SIS	CanESM2	AGCM4	CLASS	NCAR		http://journals.ametsoc.org/doi/pdf/10.1175/JCLI-D-11-00715.1
ARPEGE-Climate ISBA NEMO-OPA GELATO	CSIRO-Mk3-6-0	Gordon	CABLE	MOM2.2	SIS	http://www.bom.gov.au/amoj/docs/2013/jeffrey_hres.pdf
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ECHAM5 SIIVA OPA8.2 LIM	CMCC-CM	ECHAM5	SILVA	OPA8.2	$_{ m TIM}$	http://www.cmcc.it/models/cmcc-cm
CAM5	CMCC-CESM	ECHAM5	SILVA	OPA8.2	LIM	http://www.cmcc.it/models/cmcc-cm
CAM4	CESM1-CAM5	CAM5	CLM4	POP2	CICE4	https://www2.cesm.ucar.edu/models
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CAM3.5 CLM/BNU MOM4.1 CICE4.1 BCC_AGCM 2.1 CLM3 MOM4 SIS BCC_AGGM 2.1 CLM3 MOM4 GFDL SIS UKMO GA1.0 CABLE v1.8 MOM4.1 CICE4.1 HadGEM2 r1.1 MOSES MOM4.1 CICE4.1	CCSM4	CAM4	CLM4	POP2	CICE4	${\rm https://www2.cesm.ucar.edu/models}$
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	ACCESS1-0	HadGEM2 r1.1	MOSES	MOM4.1	CICE4.1	http://www.cawcr.gov.au/publications/technicalreports/CTR_059.pdf

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 climate model (or observation). Each box represents a pair-wise combination,
 where warm colors indicate a greater distance. Distances are measured as a
 fraction of the mean inter-model distance in CMIP5.
- 2 Histograms of CMIP5 inter-model euclidean distances in the EOF loading 948 space derived from (a) 1970-2000 monthly mean climatological fields as de-949 fined in Table 1 and (b) changes in the aforementioned fields between (1970-950 2000) and (2070-2100), as compared to a sample of 10⁵ histograms calculated 951 from randomly sampled distributions. Gray bars show the histogram of inter-952 model distances in the CMIP5 ensemble in an EOF space constructed with all 953 available variables, while other colors show distances constructed with only a 954 subset of variables; Surface Temperature (TAS), Top of Atmosphere Short-955 wave and Longwave fluxes (TOA), Total Precipitation (PR) and zonal mean 956 temperature and humidity (TQ). The yellow bars indicate the distribution 957 using all variables from the CCSM4 initial condition ensemble. The box and 958 whisker plots show the range of bin values observed in the random distribu-959 tions showing the 10th, 50th and 90th percentiles of the distribution. 960

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An illustration of inter-model and observation-model distances in an EOF space defined by (a) 1970-2000 simulated climatology for 'ALL' variables and (b) the anomaly between 1970-2000 and 2070-2100 under the RCP8.5 scenario for 'ALL' variables. Plots are repeated for individual variables, Top of Atmosphere shortwave and longwave fluxes (c), Precipitation (d) and Surface Air Temperature (e). Inter-model lines illustrate where the inter-model distance is less than 50% (dotted) or 90% (solid) of nearest inter-point distances in a randomly generated distribution of with the same dimensionality, variance and population.

An illustration of the stepwise model elimination procedure outlined in Section 2.e as applied to the 36 models from the CMIP5 ensemble, using model similarity information from the present day (1970-2000) climatology for 'ALL' variables and the 'wide' quality radius. The full set of models are shown on the left of each plot, and the order of model removal is shown on the bottom axis with the left-most model removed first. If the number of effective models n_{eff} decreases by less than 0.5, then the removed model is shown merging with its nearest neighbor in EOF space. If the number of effective models decreases by more than 0.5, the line is shown as ending - indicating the removal of that model family from the ensemble. Background shading indicates whether the smallest inter-point distance in EOF space using the remaining archive is less than 90% (light grey), 50% (mid grey) or 10% (dark grey) of purely random distributions of the same population, variance and dimensionality.

- 5 Plots illustrating the stepwise model elimination following the procedure in 983 Section 2.e. Calculations are conducted using model similarity metrics derived 984 from both present day climatology and from future climate change under 985 RCP8.5. (a) The number of effective models as a function of the number 986 of actual models remaining in the ensemble. The percentile cutoff is the 987 fraction of nearest neighbor distances seen in purely random ensembles used 988 to define the radius of similarity D_u in Equation 10. (b) The nearest-neighbor 989 distance as a function of the number of models remaining. For comparison, the 990 10th, 50th and 90th percentile of nearest neighbor distances in purely random 991 ensembles of the same dimensionality and variance are shown. (c) RMSE 992 of weighted and unweighted multi-model means as a function of remaining 993 models. 994
- A plot demonstrating how model uniqueness weights and model quality weights change as models are eliminated in the sequence shown in Figure 4, for (a) 36, (b) 20 and (c) 10 models remaining.

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7 A plot showing suggested subsets of CMIP5 given model quality scores and co-dependencies derived in a number of ways. Each line in the figure repeats the analysis leading to figure 4 with different assumptions. Plotted are the remaining models where the smallest inter-point distance in EOF space using the remaining archive is greater than 10% (unfilled symbols) or 50% (filled symbols) of purely random distributions of the same population, variance and dimensionality (regions marked by mid grey and dark dray shading in Figure 4). The analysis is conducted with zonal mean temperature and humidity (TQ), gridded precipitation (PR), gridded Top of Atmosphere shortwave and longwave fluxes (TOA), Gridded surface air temperature (TAS) and all variables combined (ALL). D_q , the radius of model quality is set to 'wide' or 'narrow' (the latter increasing the role of model quality metrics in model elimination). w_u , the model uniqueness weighting is shown calculated with the future RCP8.5 data, or the present day data. Numbers at the bottom of the plot indicate the number of retained models for the two conditions where the minimum remaining intermodel distance is greater than the 10th or 50th percentile of random smallest inter-model distances.

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A plot as in figure 7 showing suggested subsets of CMIP5 with different truncation lengths for the EOF analysis. Plotted are the remaining models where the smallest inter-point distance in EOF space using the remaining archive is greater than 50% (unfilled symbols) or 10% (filled symbols) of purely random distributions of the same population, variance and dimensionality (regions marked by mid grey and dark dray shading in Figure 4).

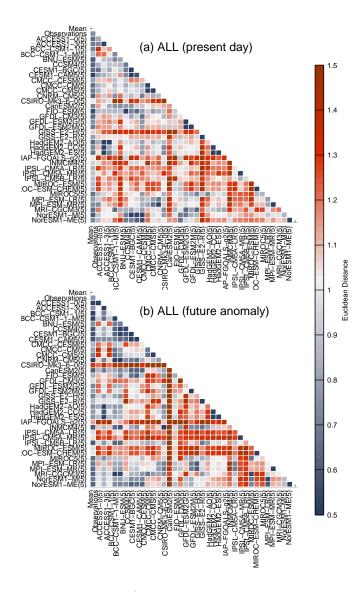


FIG. 1. A graphical representation of the inter-model distance matrix for CMIP5 calculated for ALL variables using (a) 1970-2000 monthly mean climatological fields as defined in Table 1 and (b) changes in the aforementioned fields between (1970-2000) and (2070-2100). Each row and column represents a single climate model (or observation). Each box represents a pairwise combination, where warm colors indicate a greater distance. Distances are measured as a fraction of the mean inter-model distance in CMIP5.

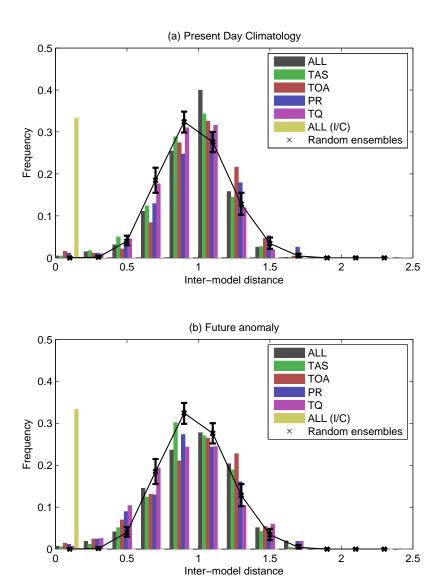


FIG. 2. Histograms of CMIP5 inter-model euclidean distances in the EOF loading space derived from (a) 1970-2000 monthly mean climatological fields as defined in Table 1 and (b) changes in the aforementioned fields between (1970-2000) and (2070-2100), as compared to a sample of 10⁵ histograms calculated from randomly sampled distributions. Gray bars show the histogram of inter-model distances in the CMIP5 ensemble in an EOF space constructed with all available variables, while other colors show distances constructed with only a subset of variables; Surface Temperature (TAS), Top of Atmosphere Shortwave and Longwave fluxes (TOA), Total Precipitation (PR) and zonal mean temperature and humidity (TQ). The yellow bars indicate the distribution using all variables from the CCSM4 initial condition ensemble. The box and whisker plots show the range of bin values observed in the random distributions showing the 10th, 50th and 90th percentiles of the distribution.

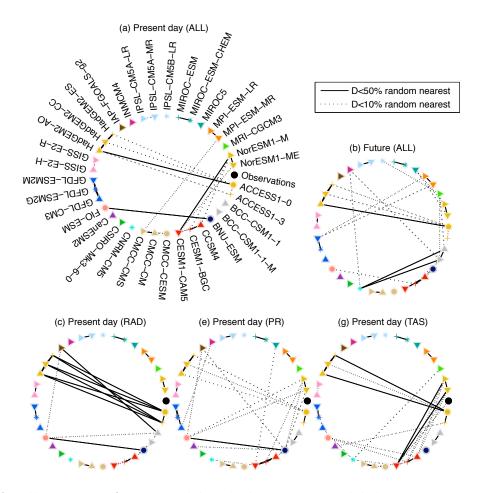


FIG. 3. An illustration of inter-model and observation-model distances in an EOF space defined by (a) 1970-2000 simulated climatology for 'ALL' variables and (b) the anomaly between 1970-2000 and 2070-2100 under the RCP8.5 scenario for 'ALL' variables. Plots are repeated for individual variables, Top of Atmosphere shortwave and longwave fluxes (c), Precipitation (d) and Surface Air Temperature (e). Inter-model lines illustrate where the inter-model distance is less than 50% (dotted) or 90% (solid) of nearest inter-point distances in a randomly generated distribution of with the same dimensionality, variance and population.

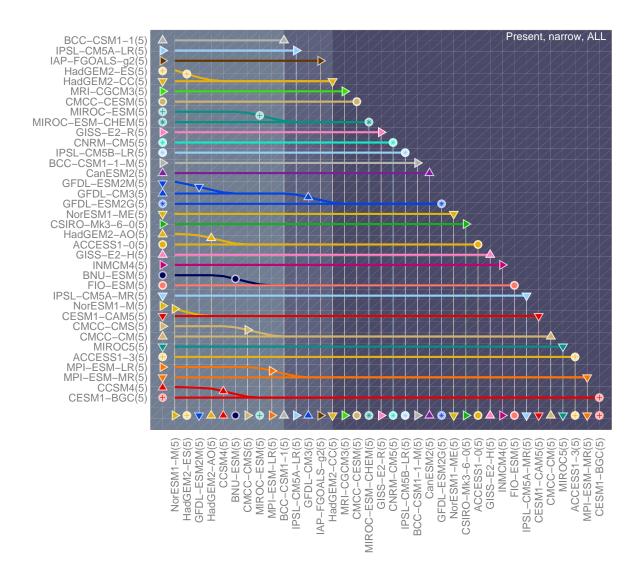


Fig. 4. An illustration of the stepwise model elimination procedure outlined in Section 2.e as applied to the 36 models from the CMIP5 ensemble, using model similarity information from the present day (1970-2000) climatology for 'ALL' variables and the 'wide' quality radius. The full set of models are shown on the left of each plot, and the order of model removal is shown on the bottom axis with the left-most model removed first. If the number of effective models n_{eff} decreases by less than 0.5, then the removed model is shown merging with its nearest neighbor in EOF space. If the number of effective models decreases by more than 0.5, the line is shown as ending - indicating the removal of that model family from the ensemble. Background shading indicates whether the smallest inter-point distance in EOF space using the remaining archive is less than 90% (light grey), 50% (mid grey) or 10% (dark grey) of purely random distributions of the same population, variance and dimensionality.

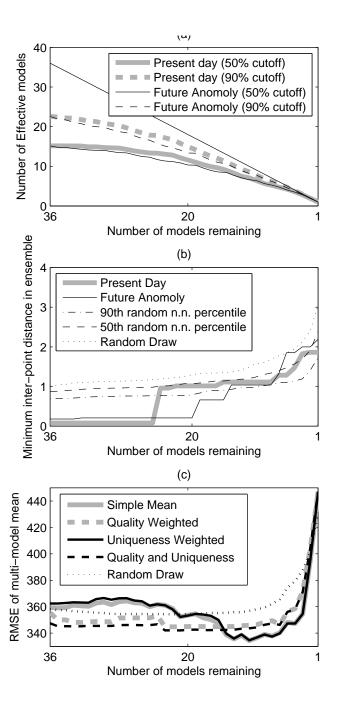


Fig. 5. Plots illustrating the stepwise model elimination following the procedure in Section 2.e. Calculations are conducted using model similarity metrics derived from both present day climatology and from future climate change under RCP8.5. (a) The number of effective models as a function of the number of actual models remaining in the ensemble. The percentile cutoff is the fraction of nearest neighbor distances seen in purely random ensembles used to define the radius of similarity D_u in Equation 10. (b) The nearest-neighbor distance as a function of the number of models remaining. For comparison, the 10th, 50th and 90th percentile of nearest neighbor distances in purely random ensembles of the same dimensionality and variance are shown. (c) RMSE of weighted and unweighted multi-model means as a function of remaining models.

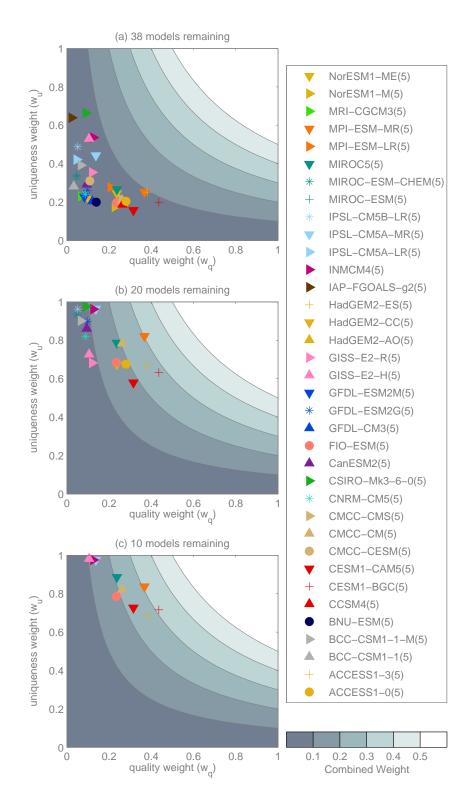


FIG. 6. A plot demonstrating how model uniqueness weights and model quality weights change as models are eliminated in the sequence shown in Figure 4, for (a) 36, (b) 20 and (c) 10 models remaining.

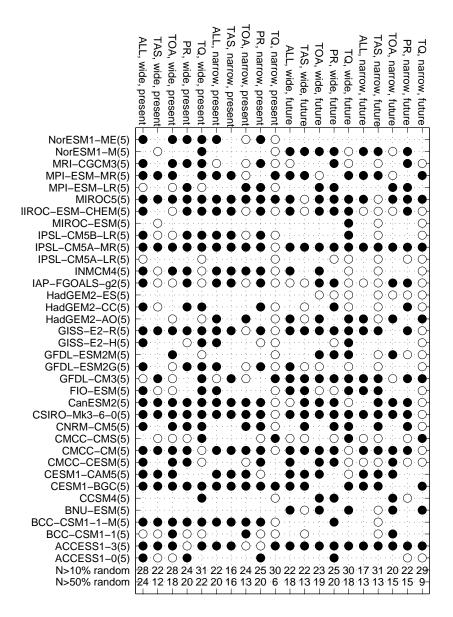


Fig. 7. A plot showing suggested subsets of CMIP5 given model quality scores and codependencies derived in a number of ways. Each line in the figure repeats the analysis leading to figure 4 with different assumptions. Plotted are the remaining models where the smallest inter-point distance in EOF space using the remaining archive is greater than 10% (unfilled symbols) or 50% (filled symbols) of purely random distributions of the same population, variance and dimensionality (regions marked by mid grey and dark dray shading in Figure 4). The analysis is conducted with zonal mean temperature and humidity (TQ), gridded precipitation (PR), gridded Top of Atmosphere shortwave and longwave fluxes (TOA), Gridded surface air temperature (TAS) and all variables combined (ALL). D_q , the radius of model quality is set to 'wide' or 'narrow' (the latter increasing the role of model quality metrics in model elimination). w_u , the model uniqueness weighting is shown calculated with the future RCP8.5 data, or the present day data. Numbers at the bottom of the plot indicate the number of retained models for the two conditions where the minimum remaining intermodel distance is greater than the 10th or $\frac{500}{52}$ h percentile of random smallest inter-model distances.

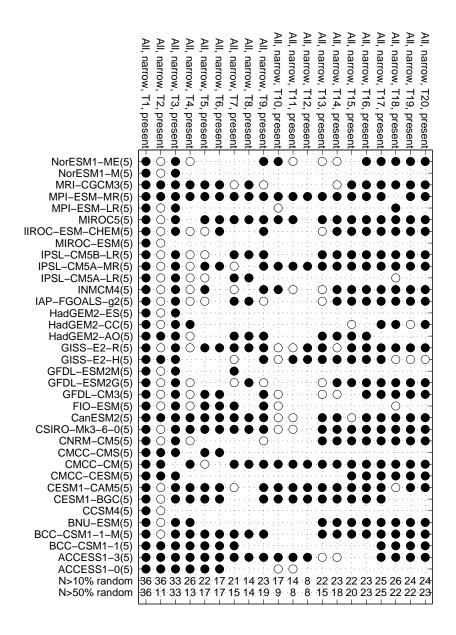


FIG. 8. A plot as in figure 7 showing suggested subsets of CMIP5 with different truncation lengths for the EOF analysis. Plotted are the remaining models where the smallest inter-point distance in EOF space using the remaining archive is greater than 50% (unfilled symbols) or 10% (filled symbols) of purely random distributions of the same population, variance and dimensionality (regions marked by mid grey and dark dray shading in Figure 4).