

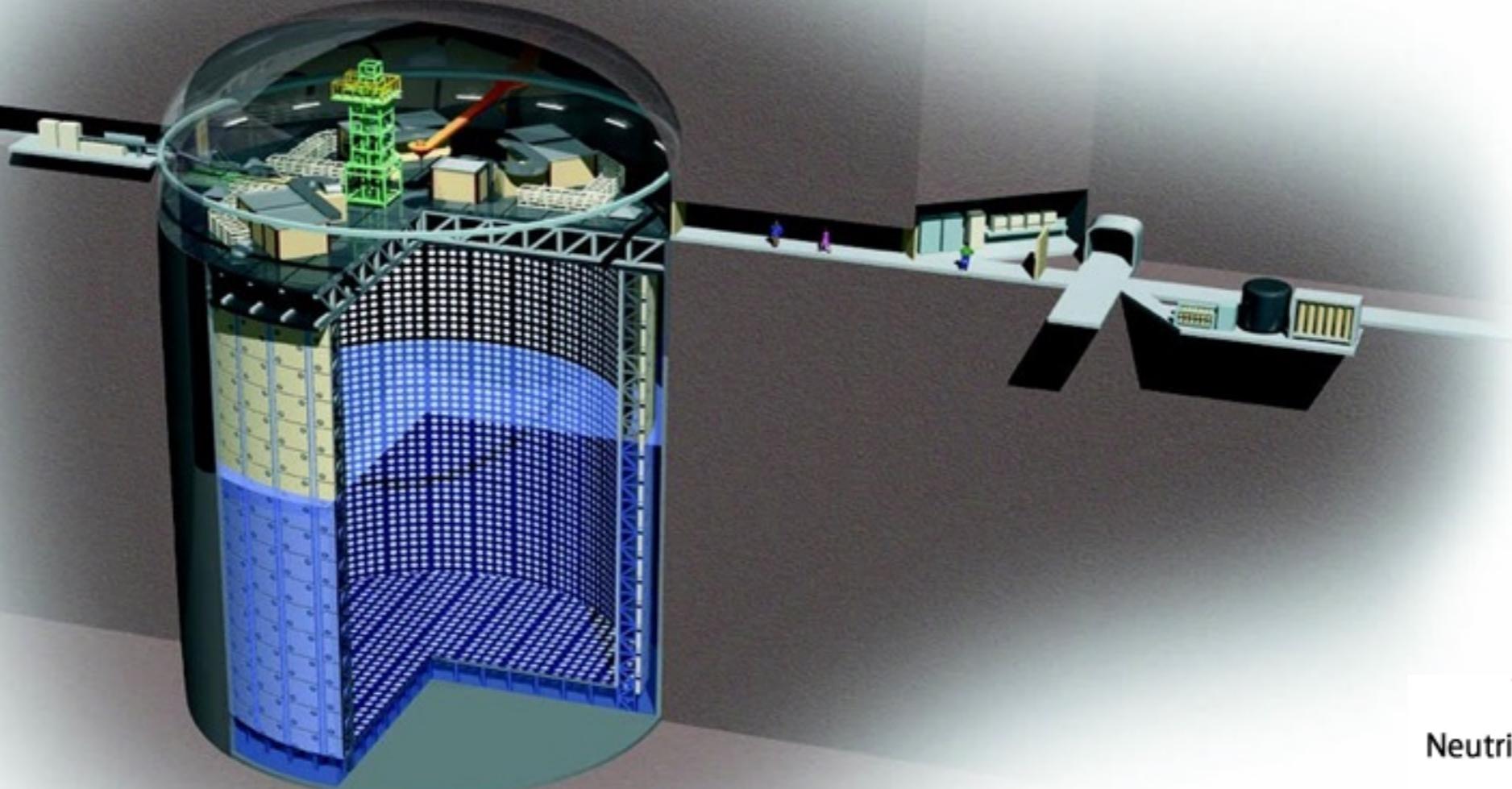
Gadolinium Doped Water Cherenkov Detectors

David Hadley
University of Warwick

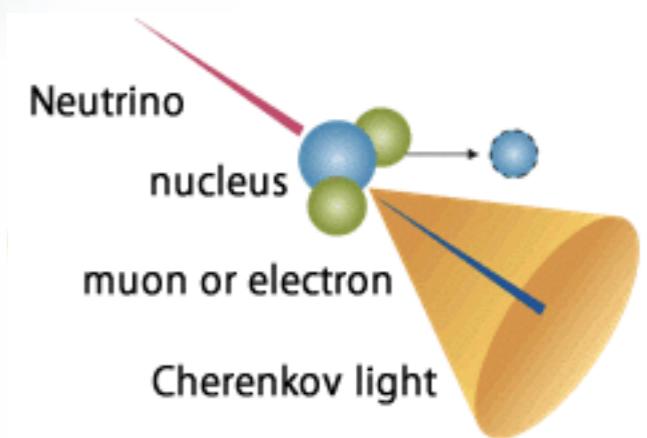
NuInt-UK Workshop

20th July 2015

Water Cherenkov Detector

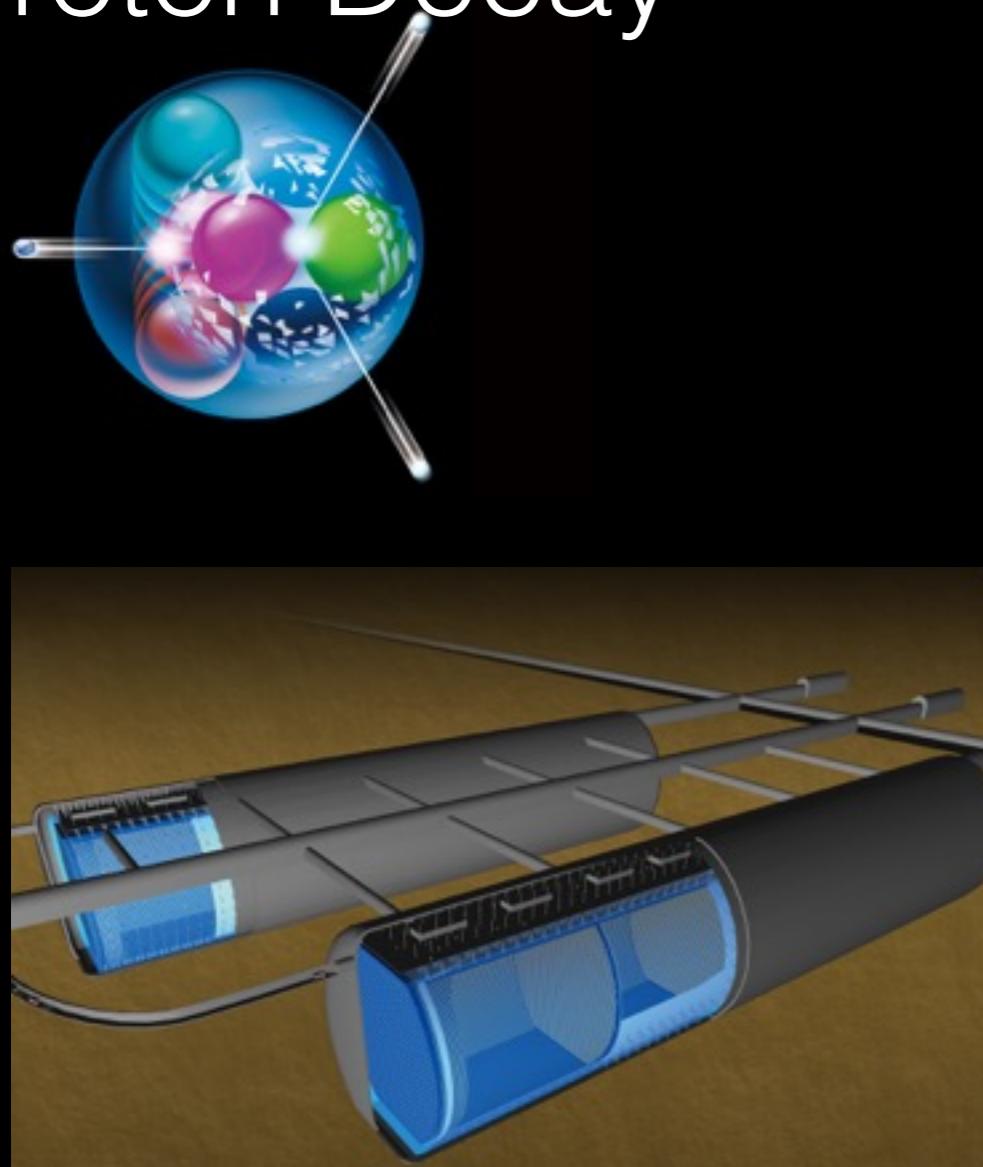


Super-Kamiokande
22.5 kt fiducial mass



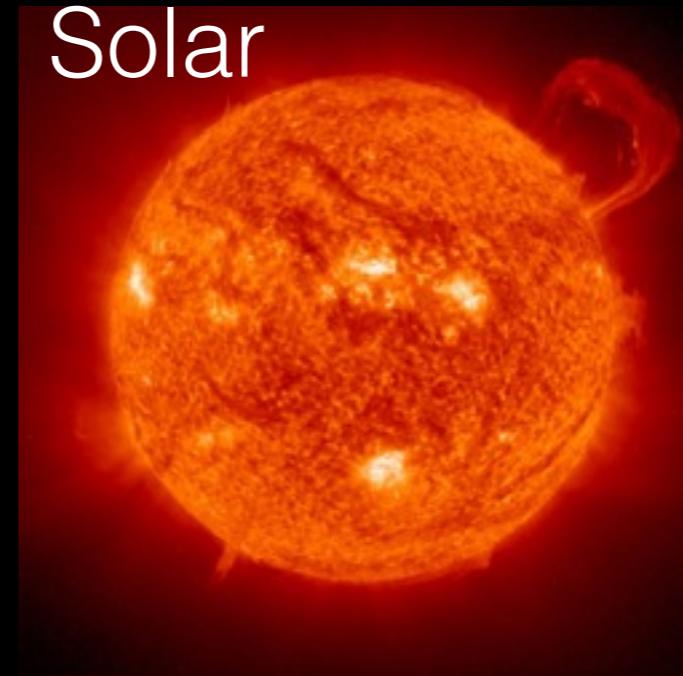
Physics with Large Scale WC

Proton Decay

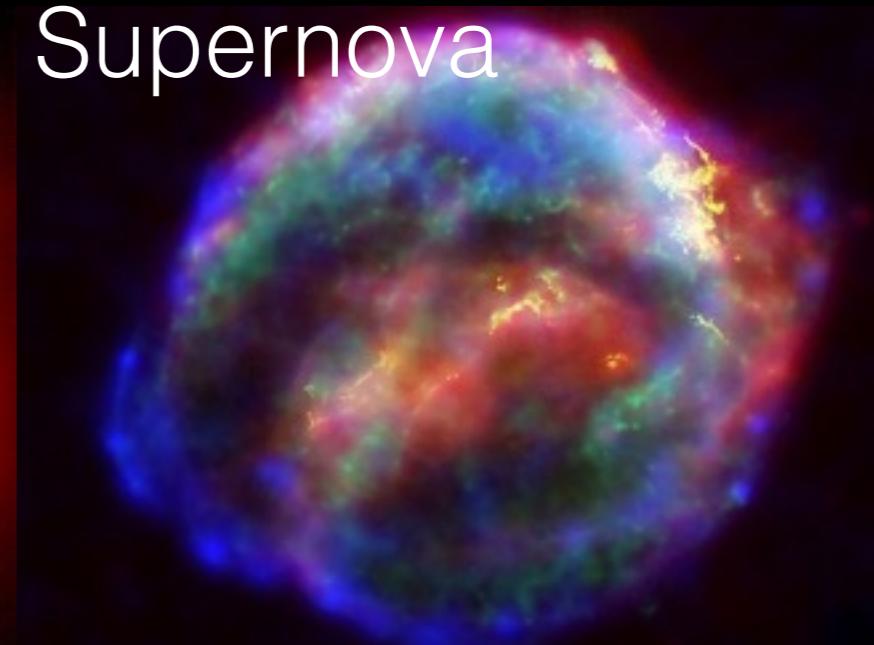


Neutrinos

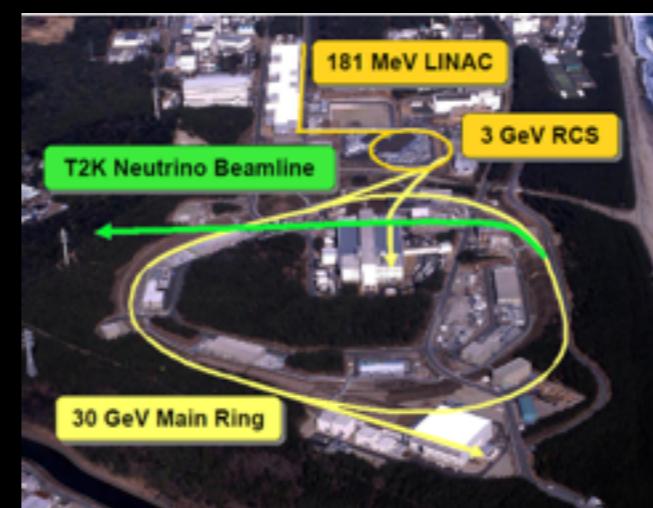
Solar



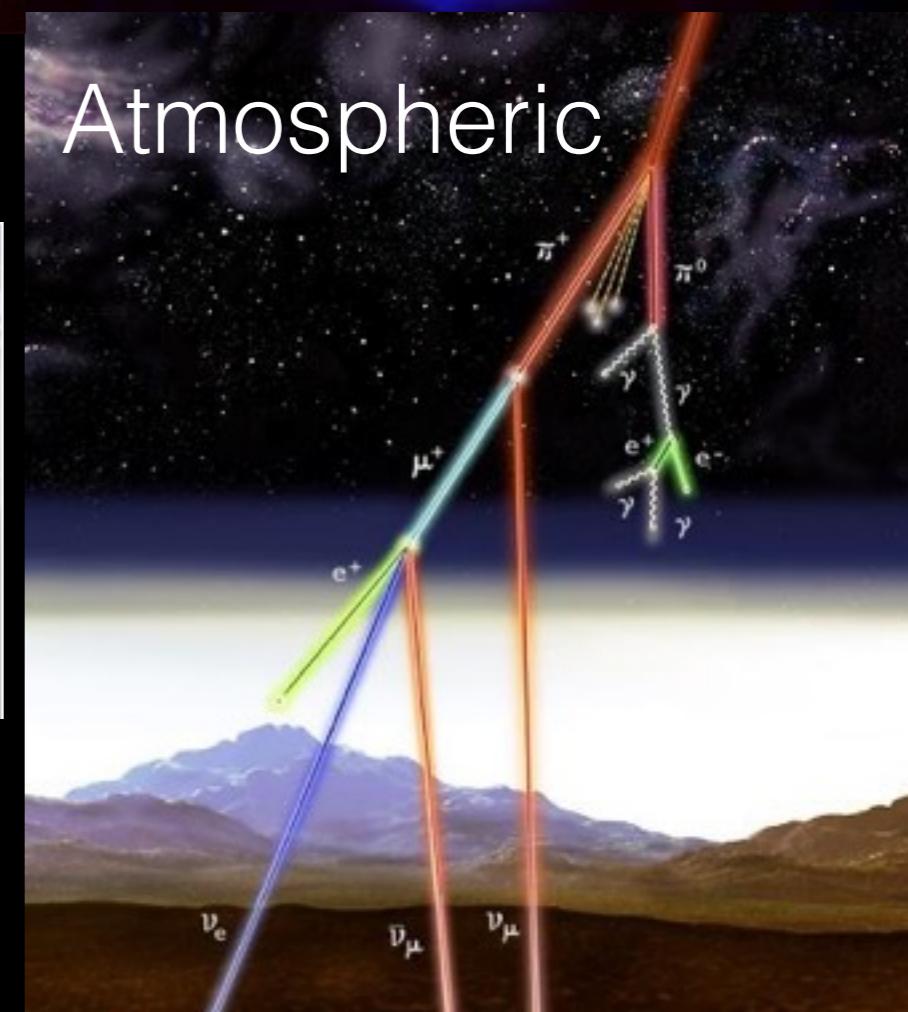
Supernova



Accelerator

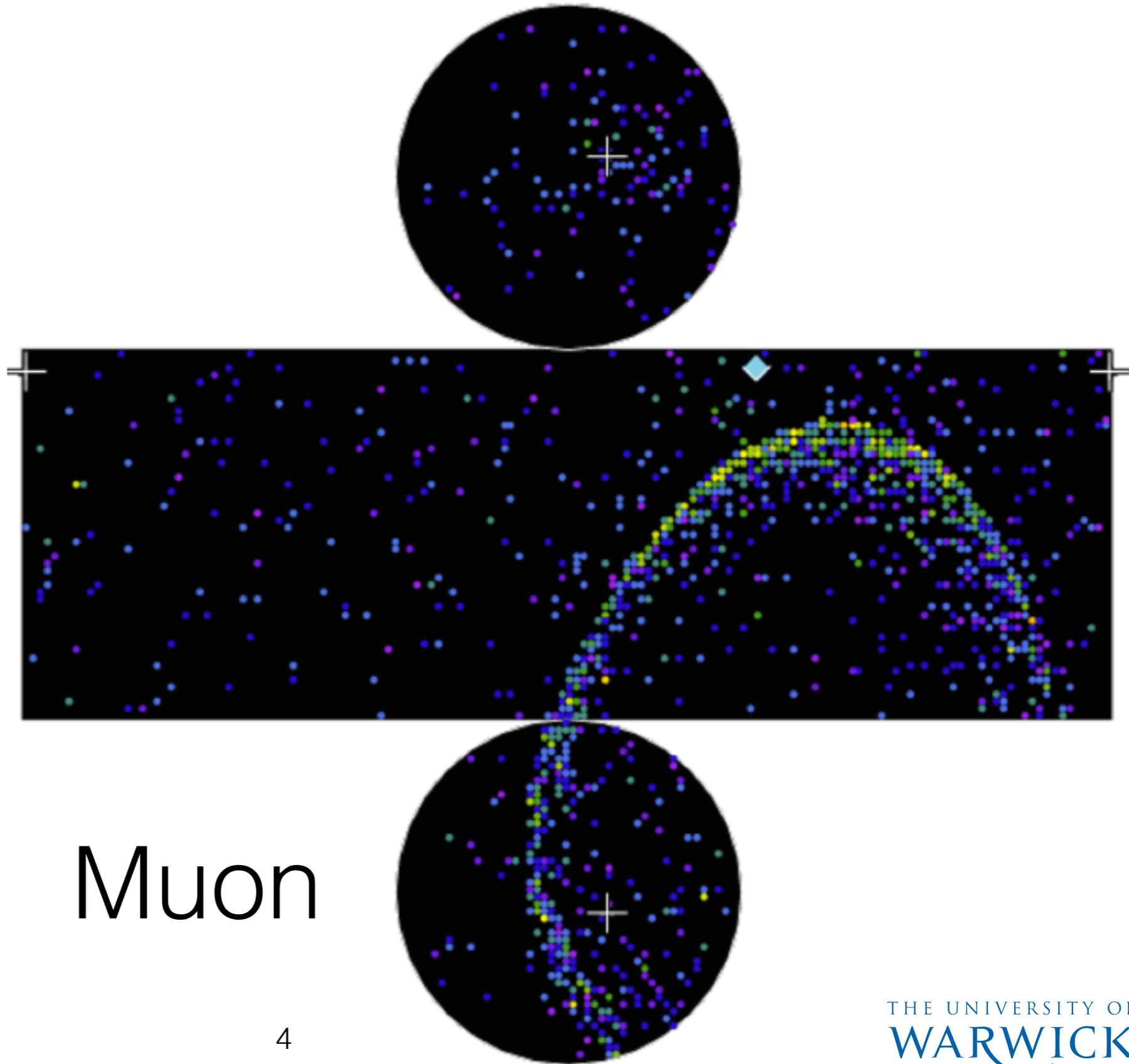


Atmospheric

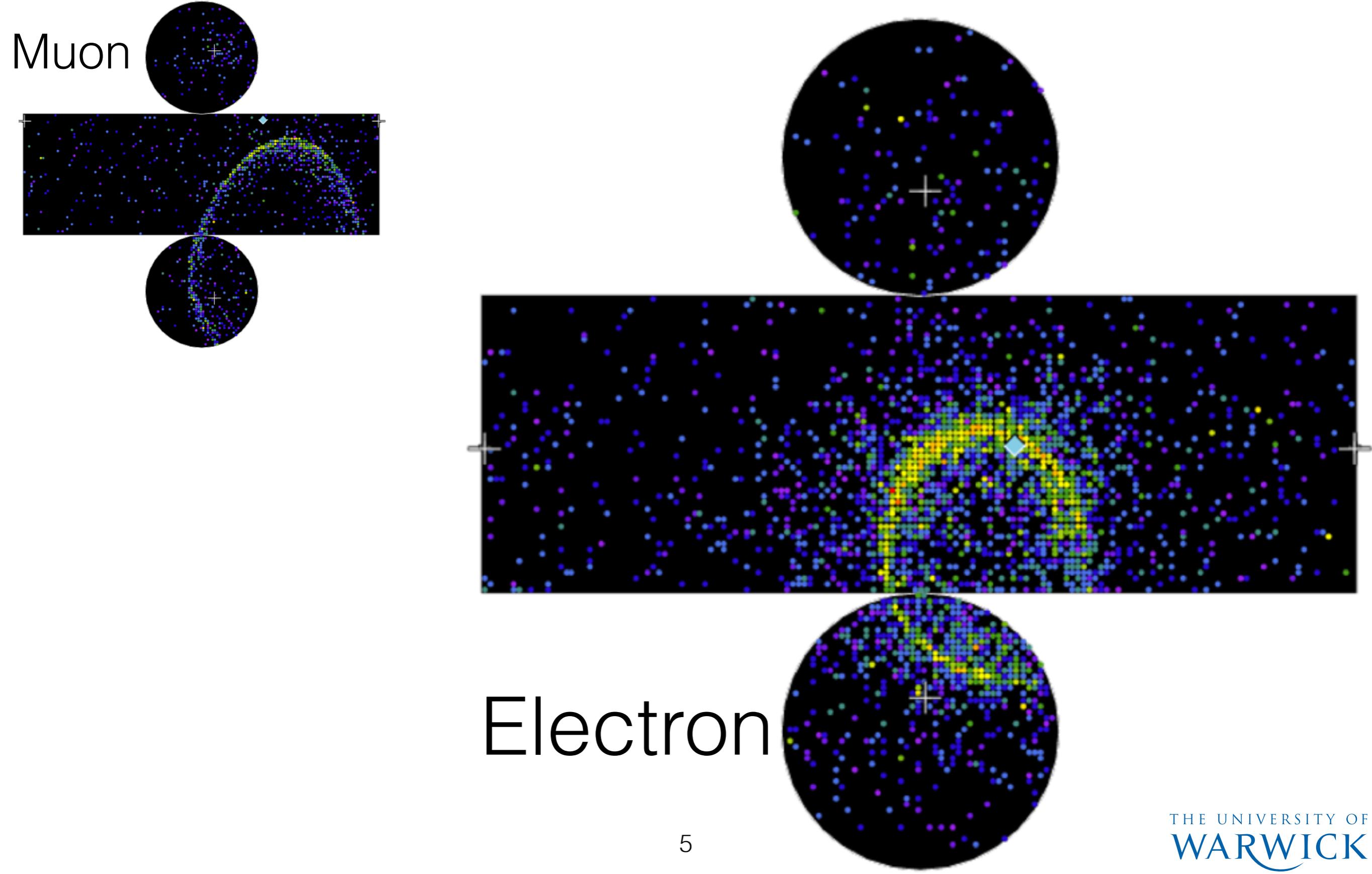


Broad physics topics,
wide energy range

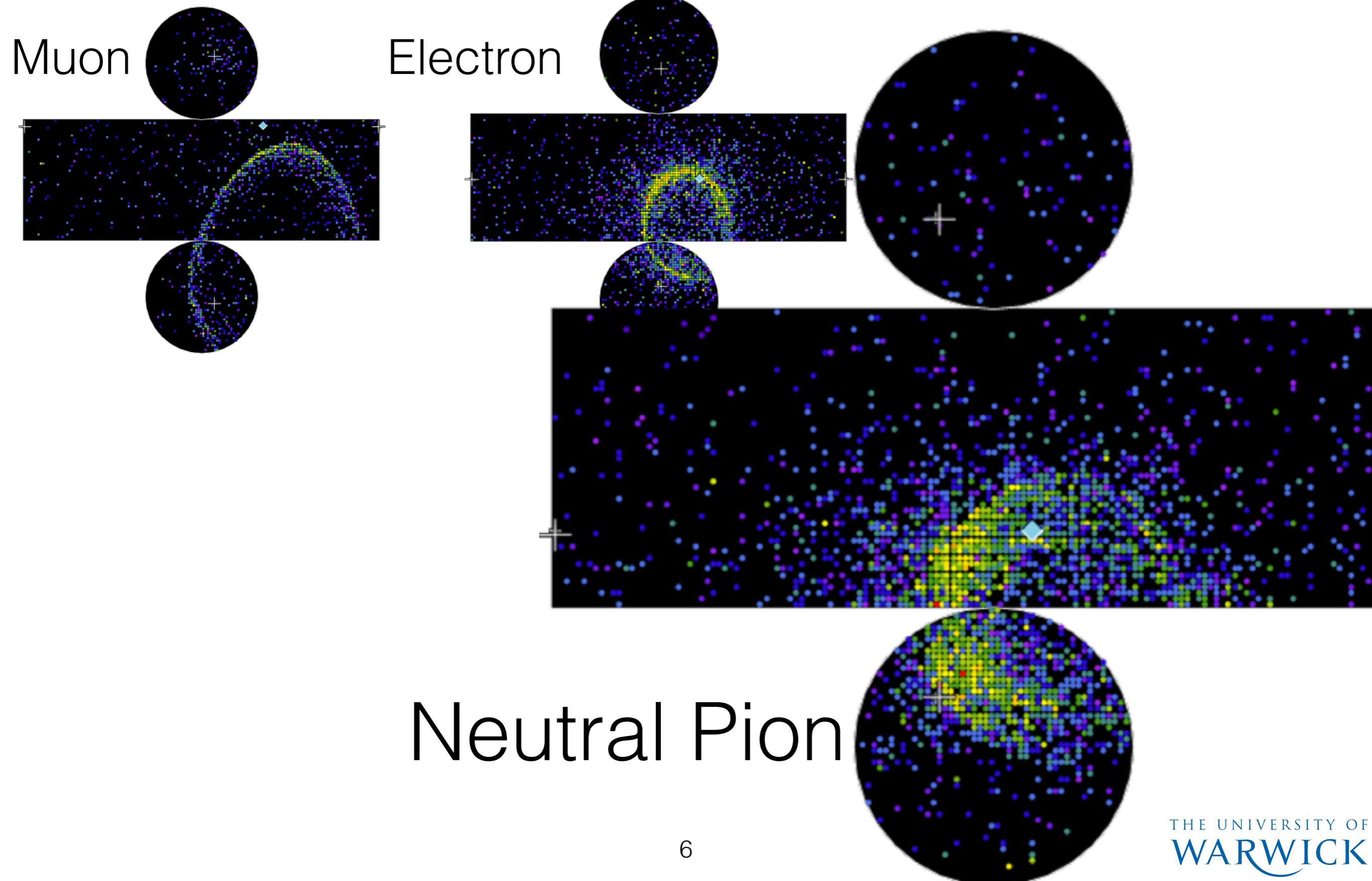
Water Cherenkov Technique



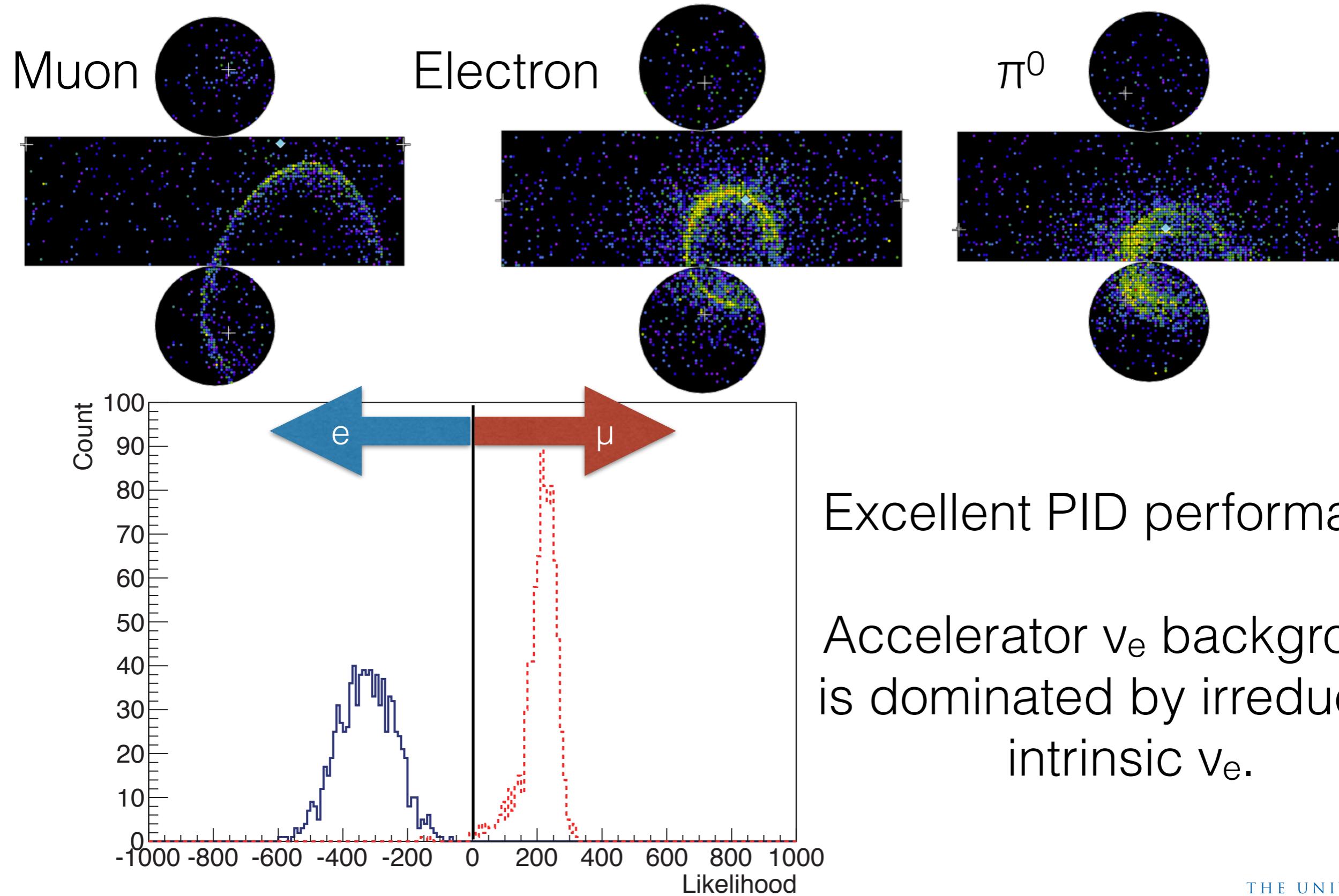
Water Cherenkov Technique



Water Cherenkov Technique



Water Cherenkov Technique



Why Water Cherenkov?

Scalability

Water is cheap, non-toxic, liquid at room temperature
long attenuation length achievable in pure water
(SK > 100m at 400nm)

Proven technology

many years of experience (eg Super-K 1996 to date)
low risk

Excellent performance

for charged particles above Cherenkov threshold

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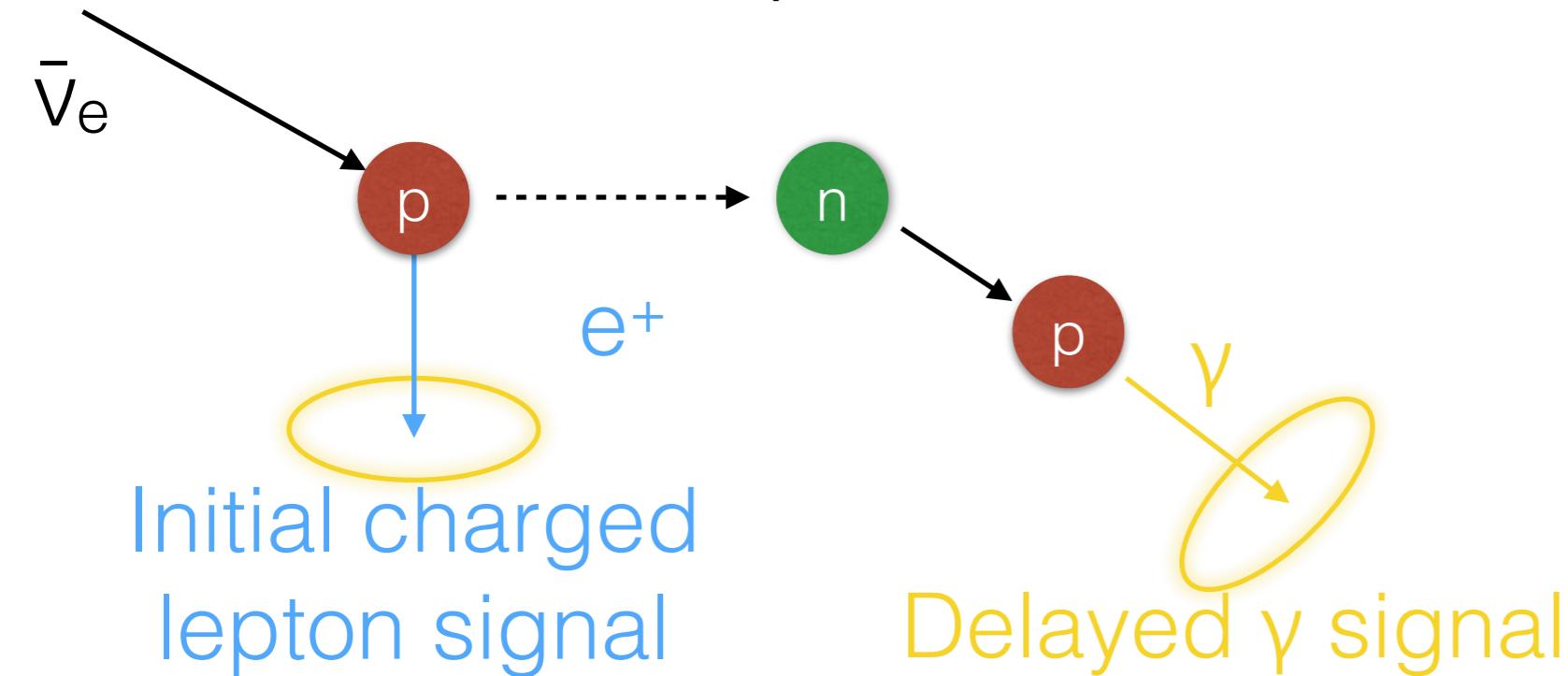
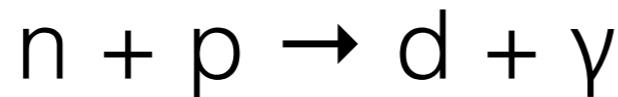
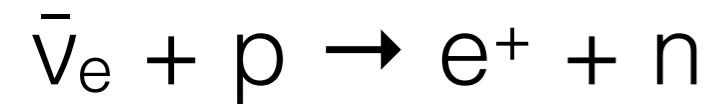
for charged particles above Cherenkov threshold

Why **not** Water Cherenkov?

Blind to particles below Cherenkov threshold

for protons $> 1.1 \text{ GeV}/c.$

Neutron Capture on Hydrogen



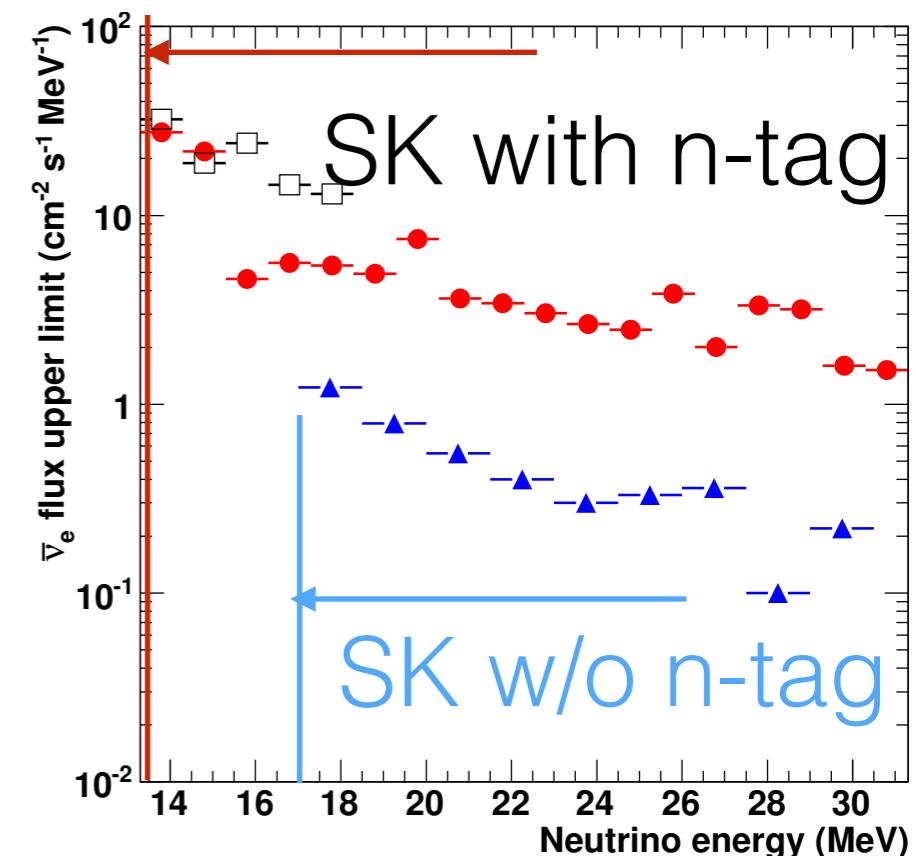
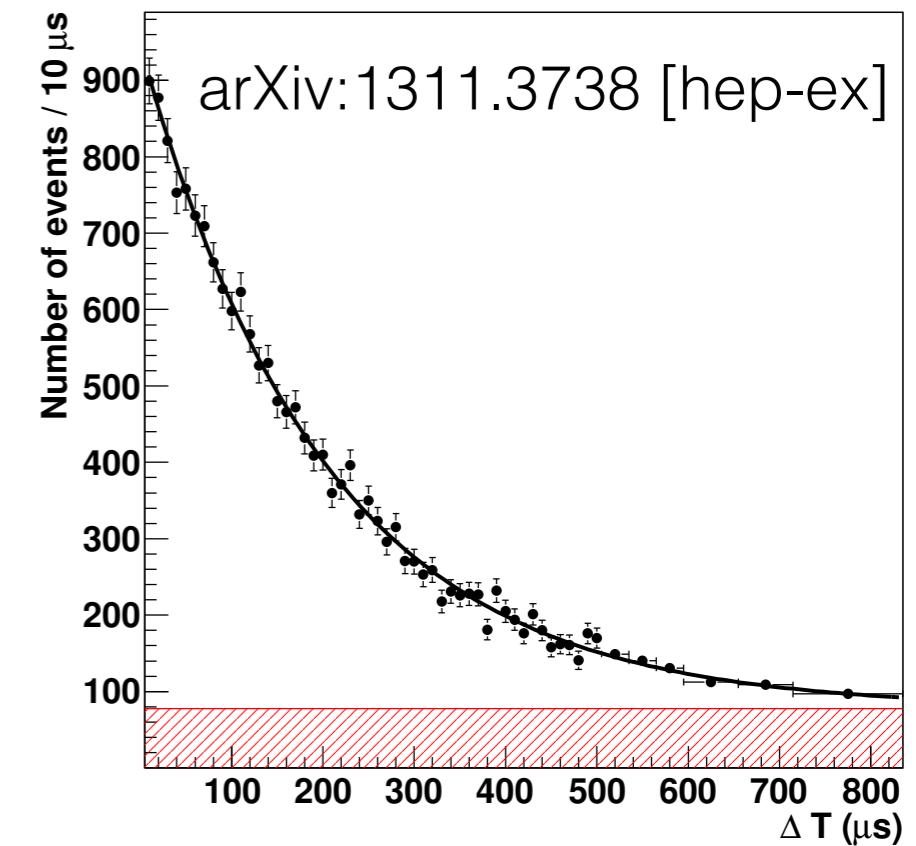
200 μs capture time

$$E_\gamma = 2.2 \text{ MeV}$$

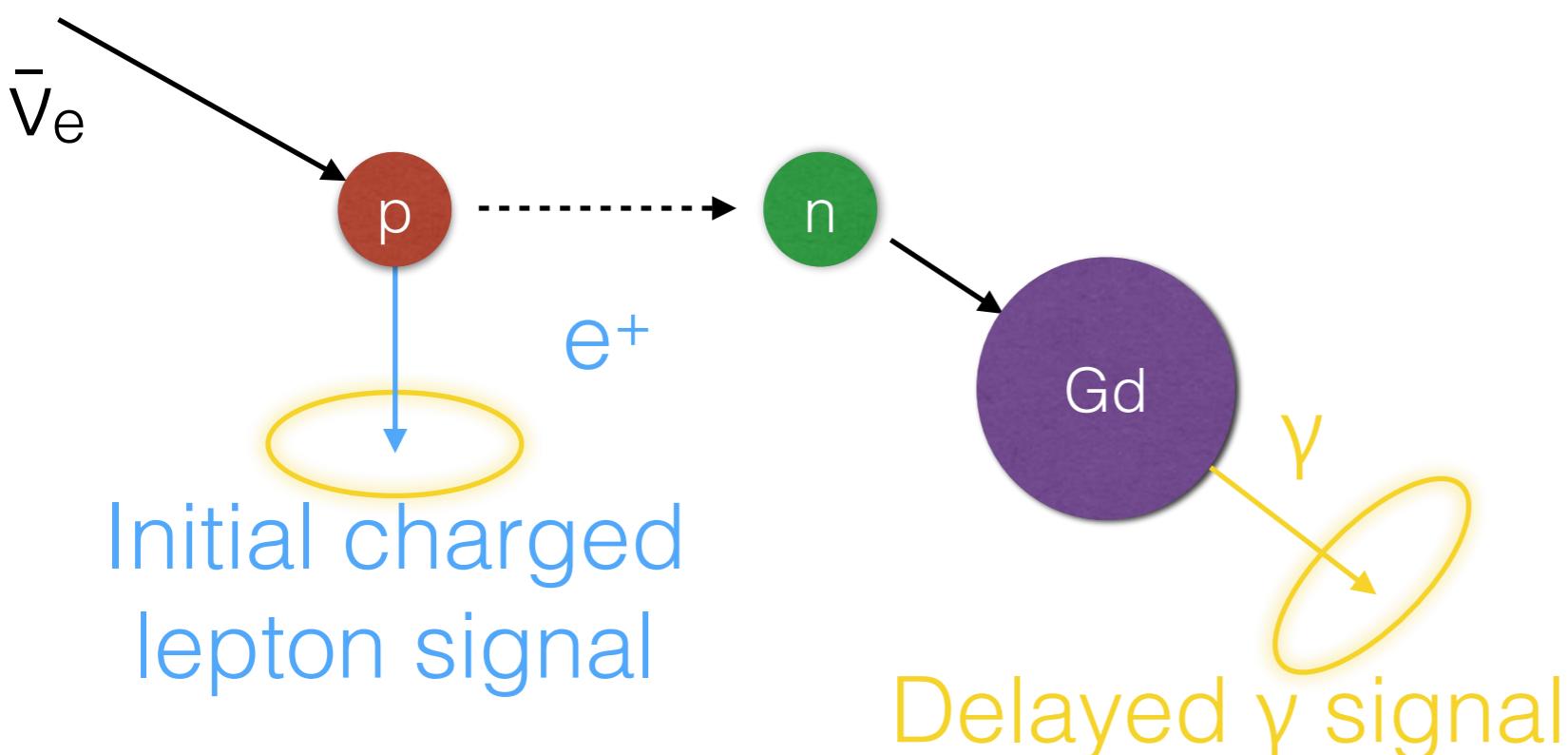
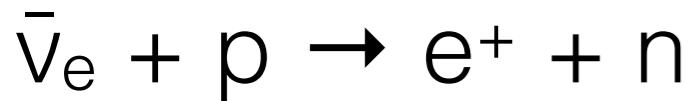
Low light yield

Close to or below trigger threshold

Low detection efficiency ($\sim 18\%$)



Neutron Capture on Gadolinium

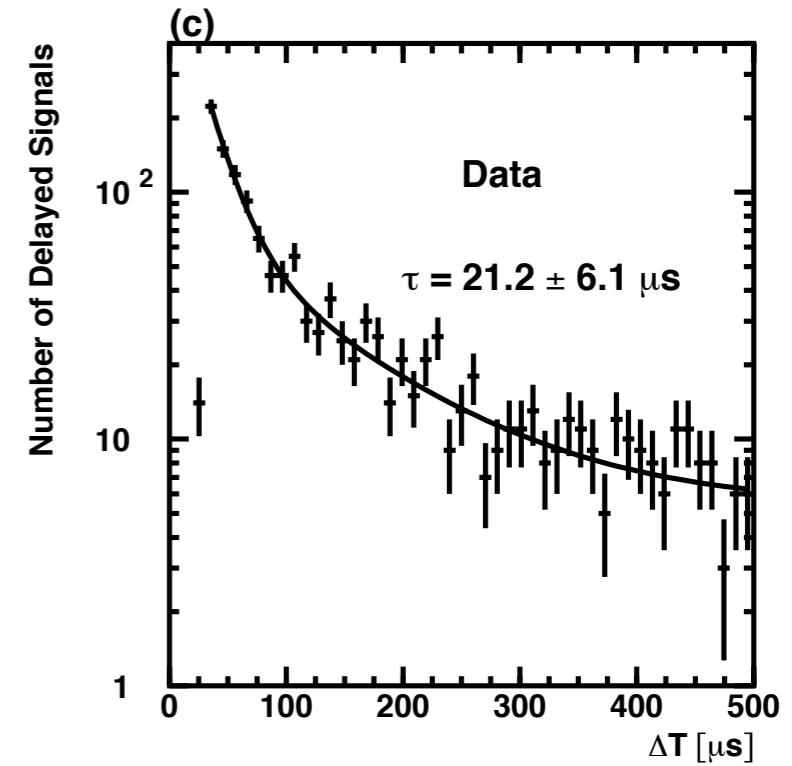
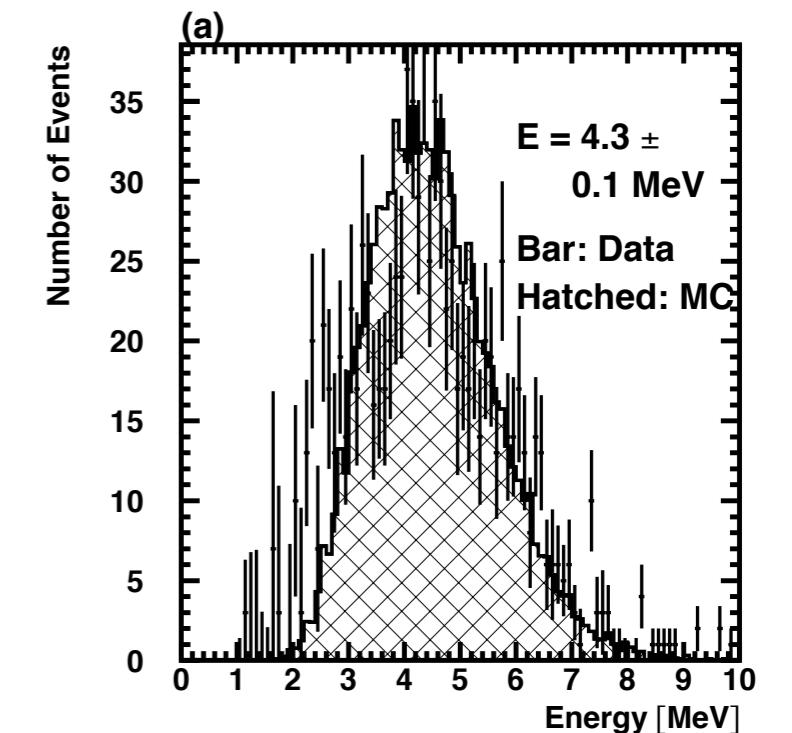


20 μ s capture time

$E_\gamma \sim 8$ MeV cascade (~ 4 MeV visible)

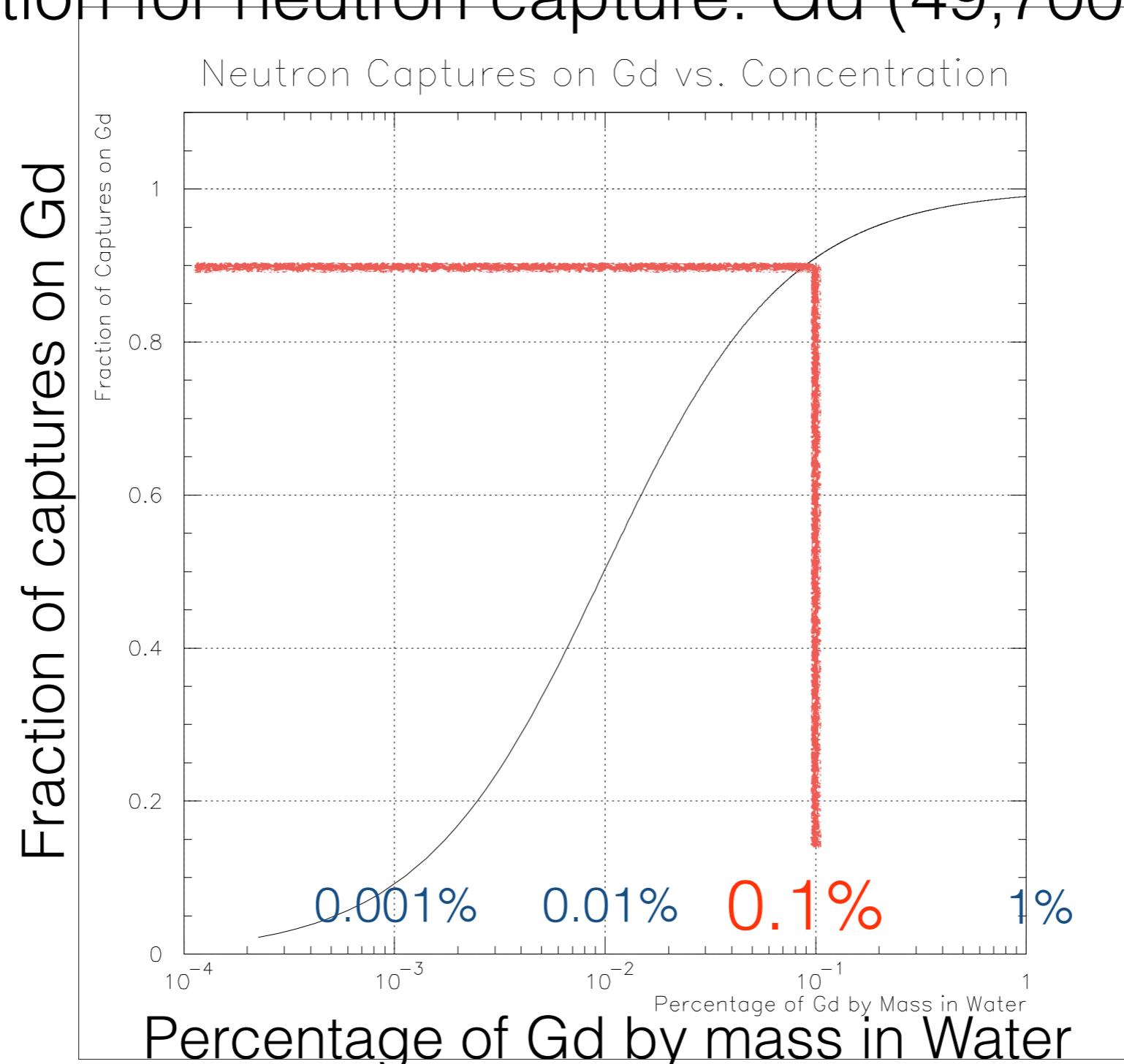
Fast capture time (small ΔT window)
Higher energy γ signal

arXiv:0811.0735 [hep-ex]



Neutron Capture on Gadolinium

Cross section for neutron capture: Gd (49,700 b), H (0.3 b)



0.1% Gd fraction gives 90% neutrons captured on Gd.

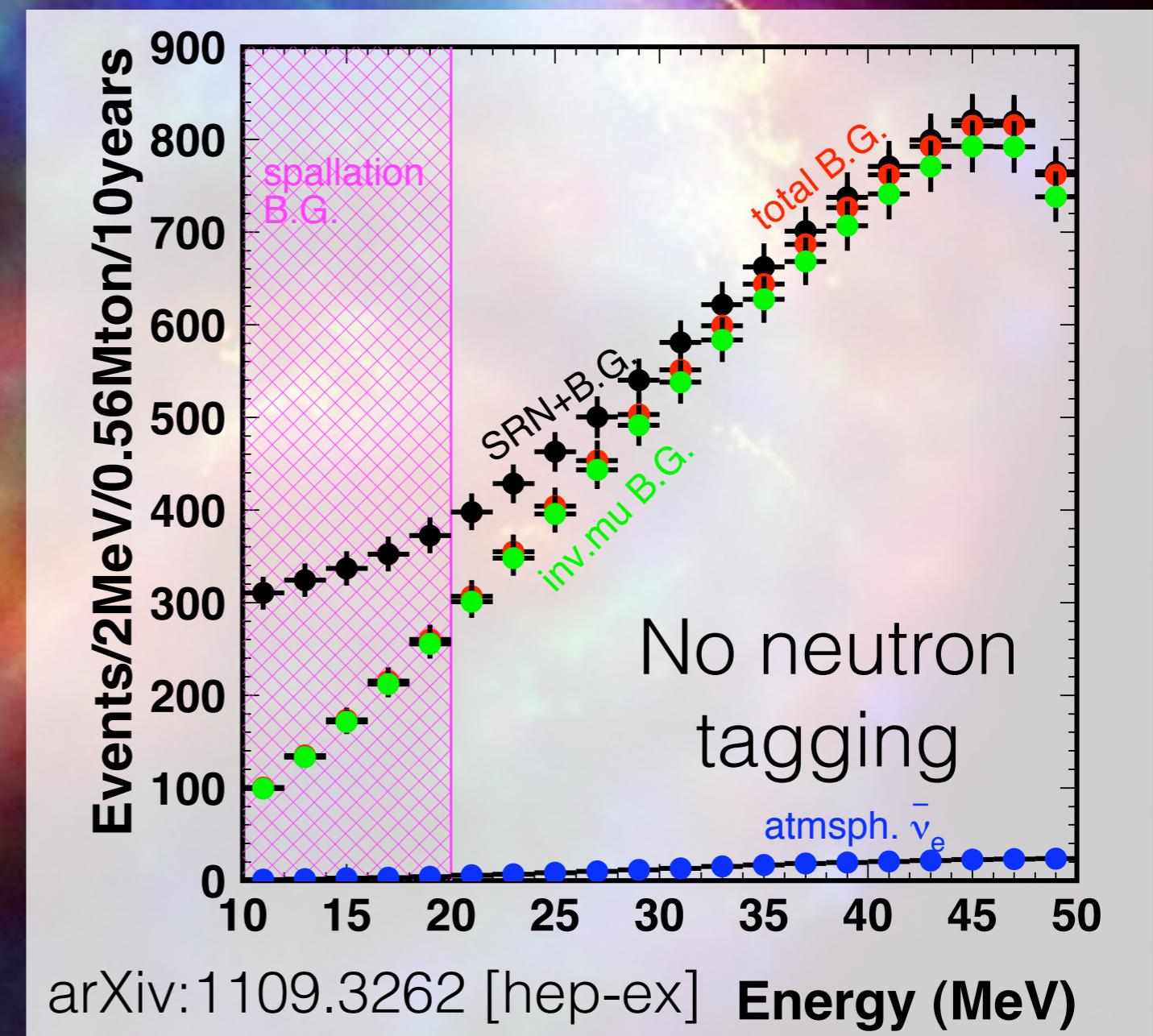
Applications: Supernova Relic Neutrinos

A low energy example

Directly observable local supernova are all too rare

Alternative is to measure diffuse supernova background DSNB/SRN

Very low rate
Large backgrounds



Applications: Supernova Relic Neutrinos

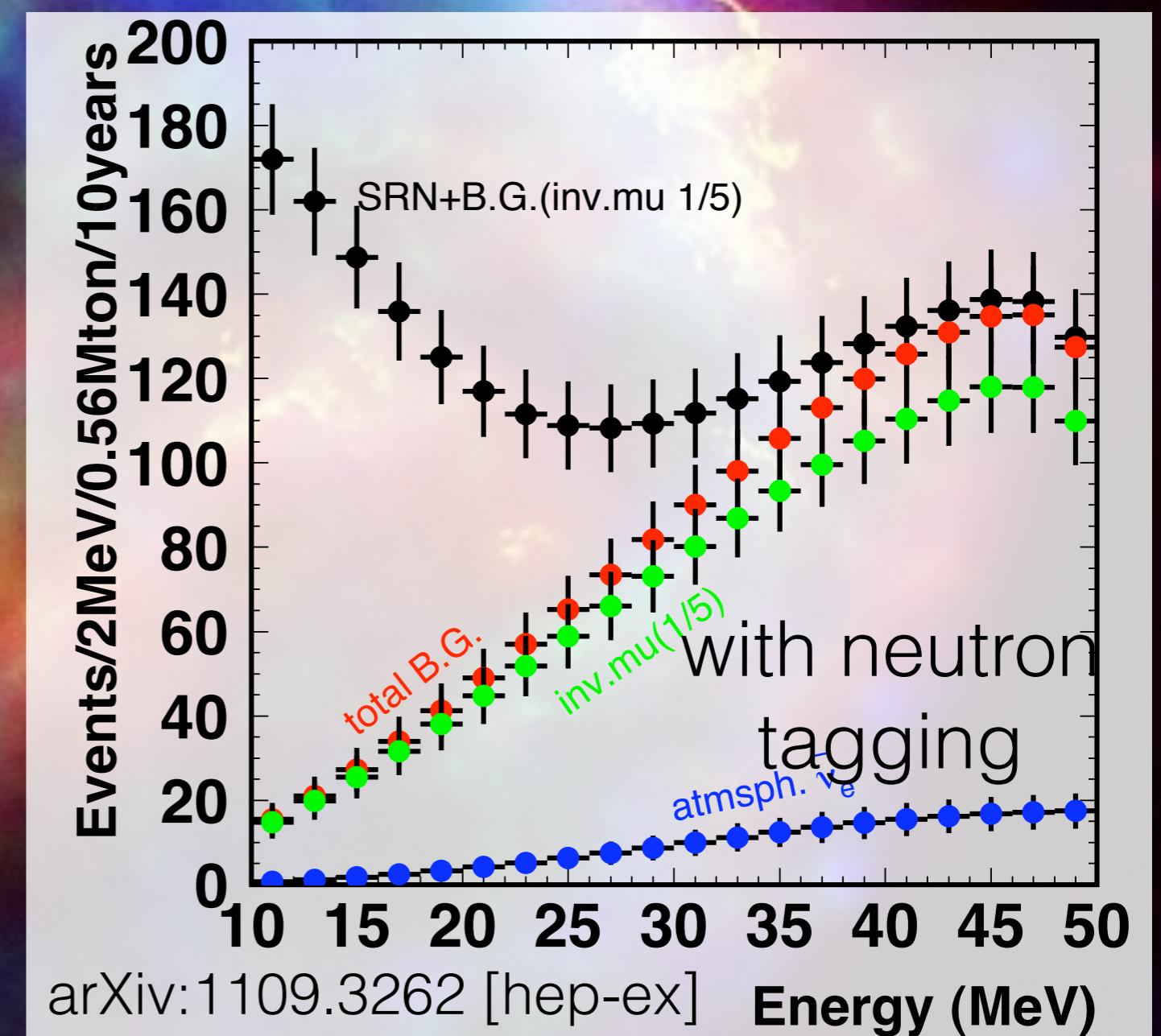
A low energy example

Directly observable local supernova are all too rare

Alternative is to measure diffuse supernova background DSNB/SRN

Very low rate
Large backgrounds

Removed by requiring coincidence with neutron



A few clean events per year in SK
~100s per year in HK

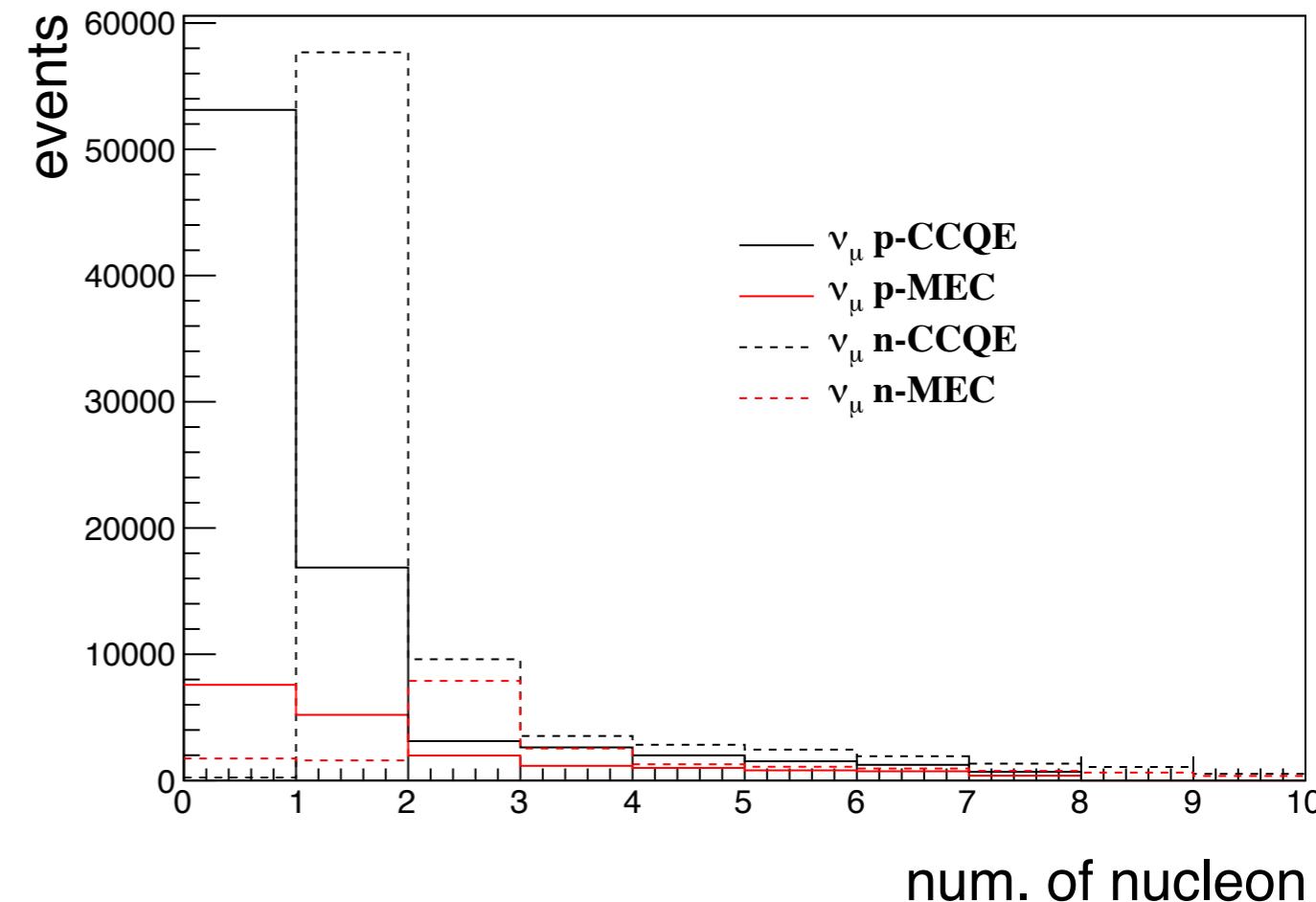
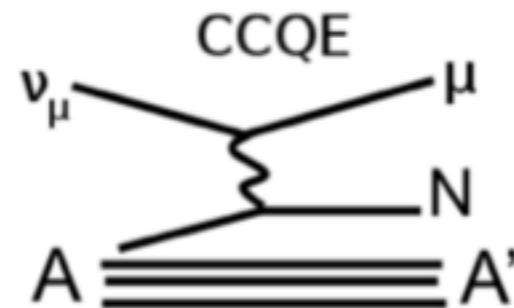
Applications: Accelerator based long baseline neutrino oscillations

A high energy example

T2K / T2HK neutrino beam energy ~ 0.6 GeV

Signal: ν CCQE: $\nu + n \rightarrow l^- + p$

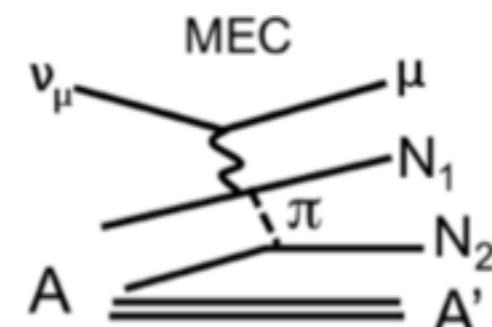
$\bar{\nu}$ CCQE: $\bar{\nu} + p \rightarrow l^+ + n$



Multi-nucleon:

$\nu + (nn) \rightarrow l^- + p + n$

$\nu + (p\ p/n) \rightarrow l^+ + n + p/n$

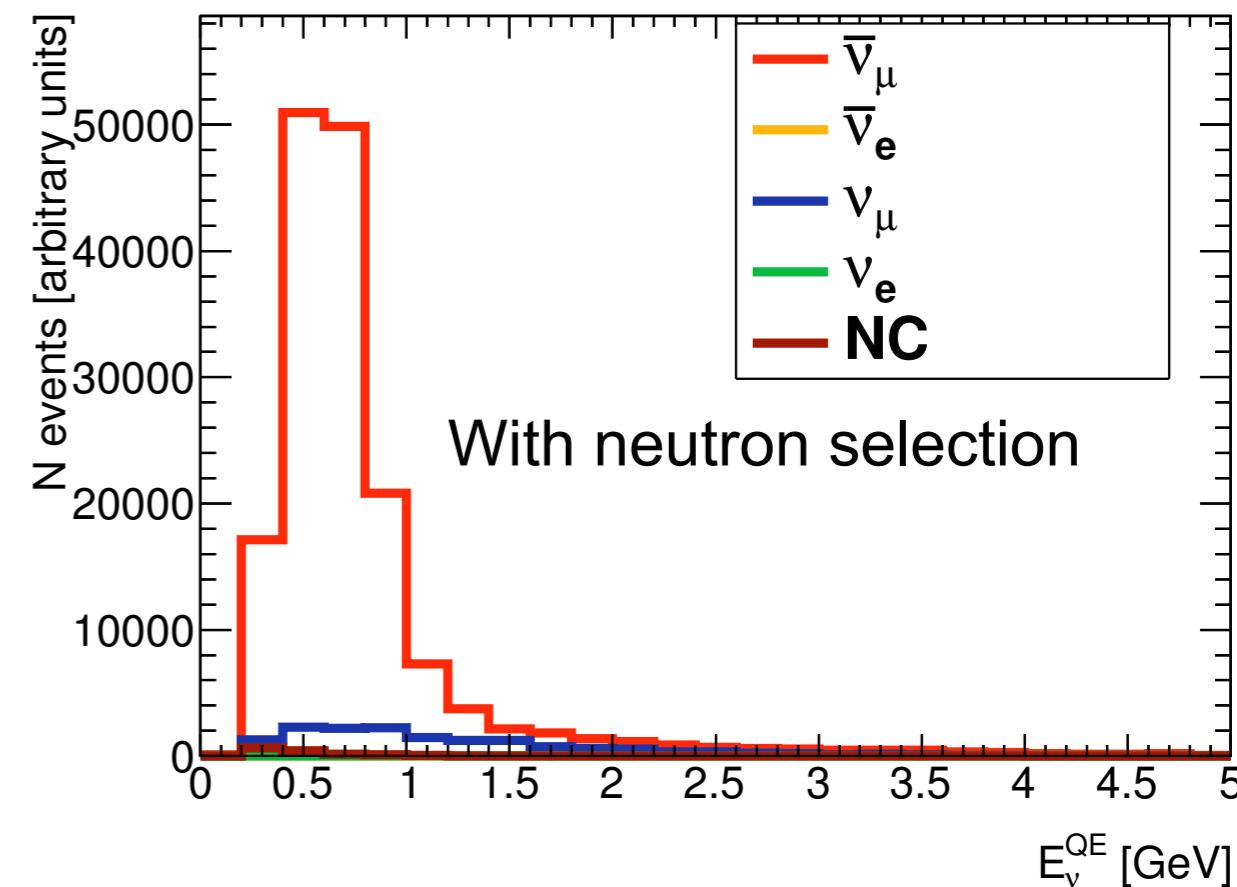
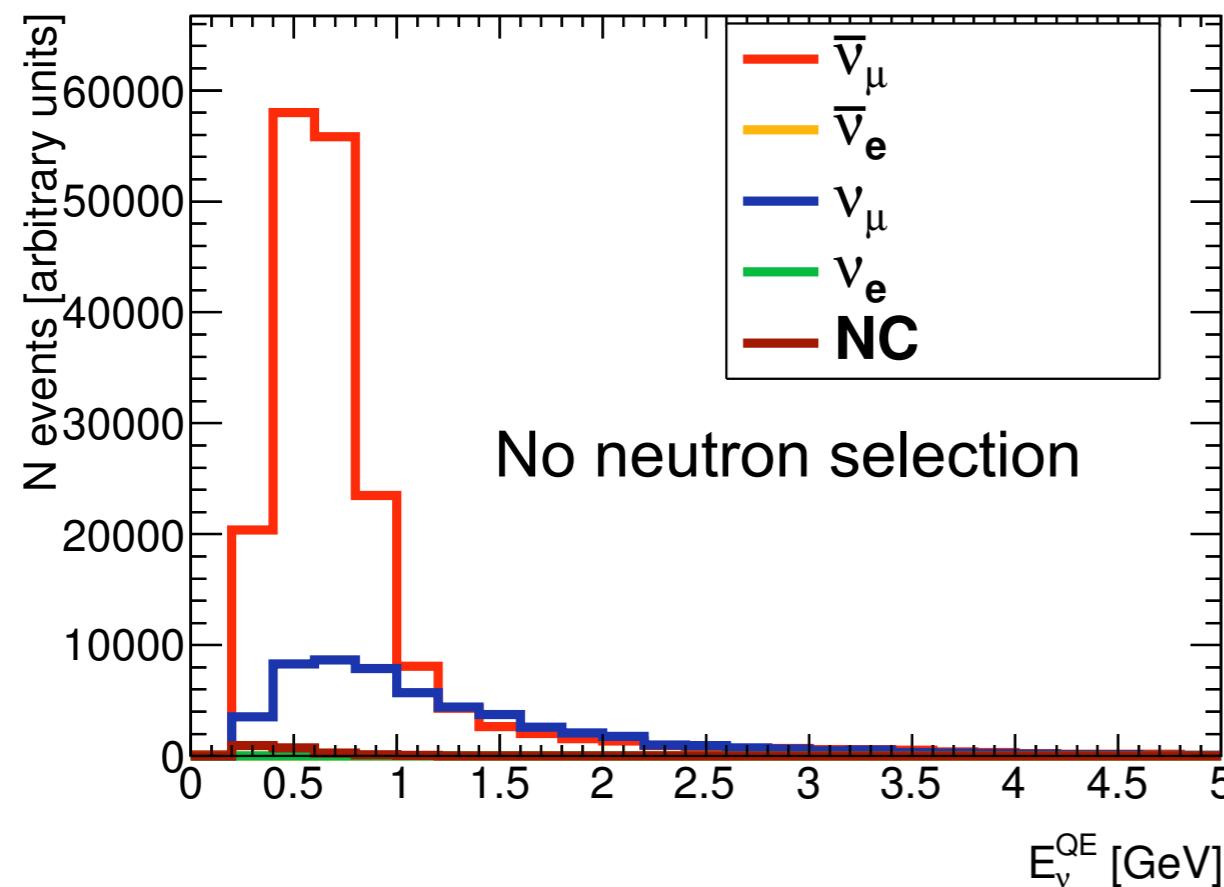


Neutron multiplicity gives an additional observable with which to isolate interaction modes.

Complimentary to LAr proton measurements

Applications: Accelerator based long baseline neutrino oscillations

Tagging neutron reduces wrong-sign background in anti-neutrino mode

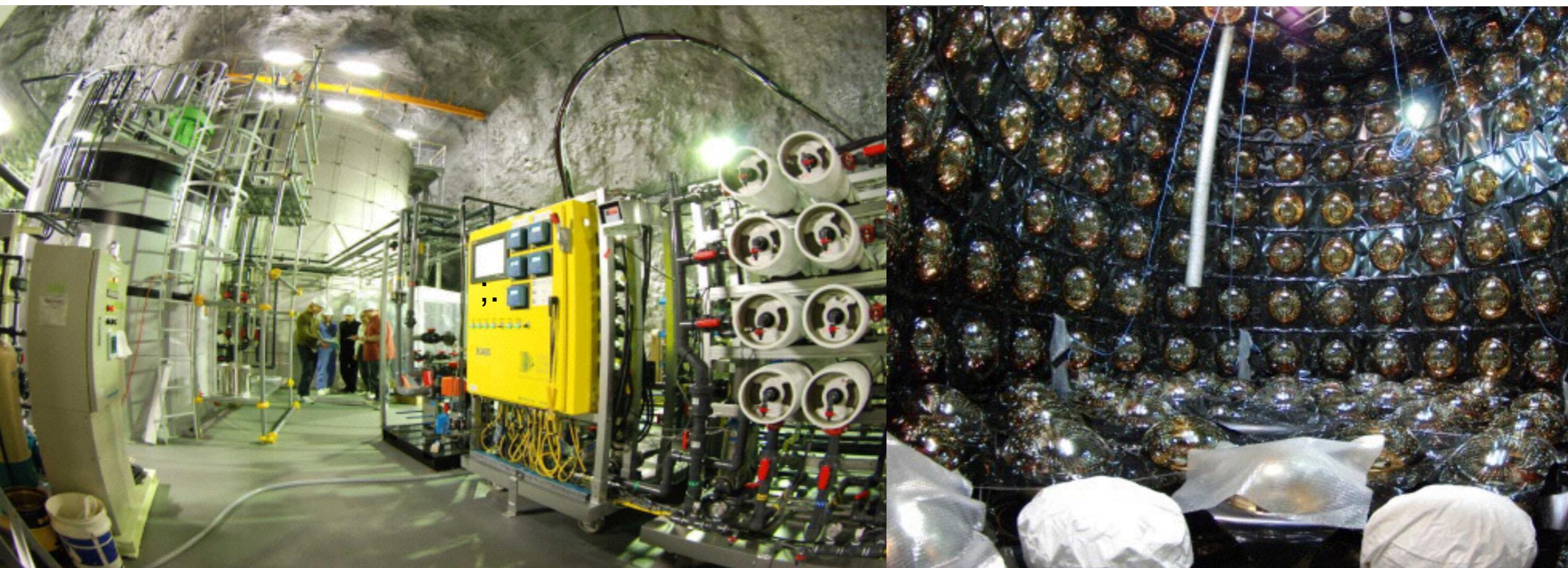


Impact on sensitivity being evaluated by
Hyper-K Gd-doped Near Detector (TITUS)
working group

EGADs

(Evaluating Gadolinium's Action on Detector Systems)

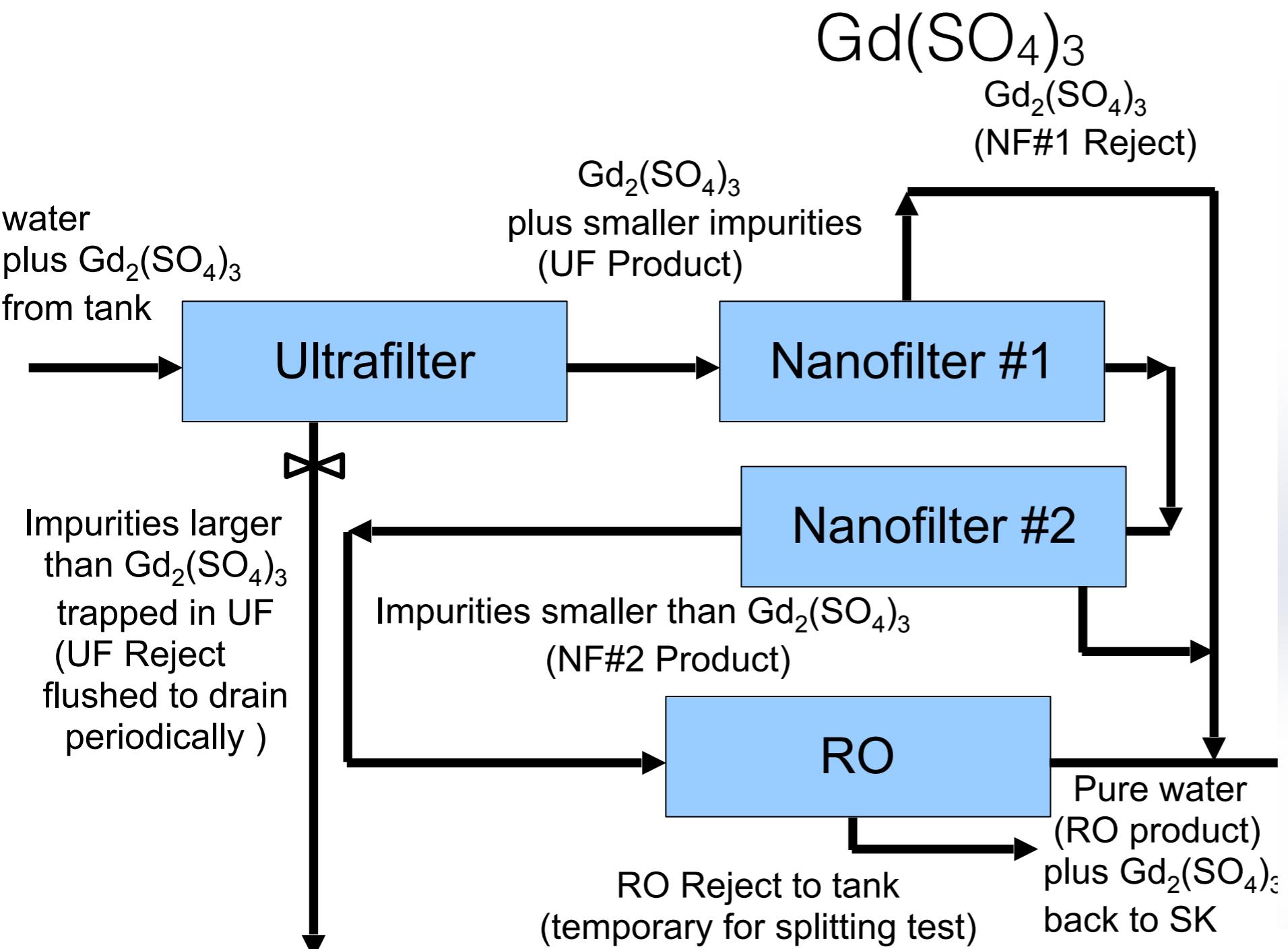
200 t instrumented Water Cherenkov detector to test introduction of a water soluble Gadolinium in a $\text{Gd}(\text{SO}_4)_3$



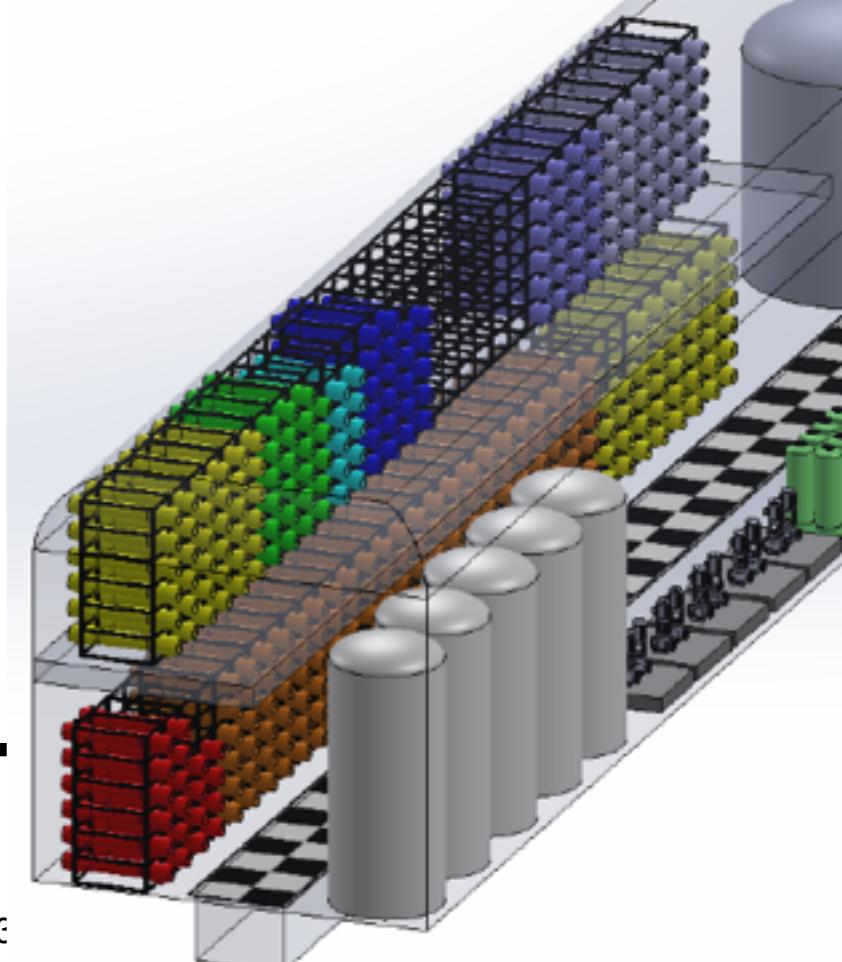
EGADs

(Evaluating Gadolinium's Action on Detector Systems)

Need a water filtration system that removes impurities but not



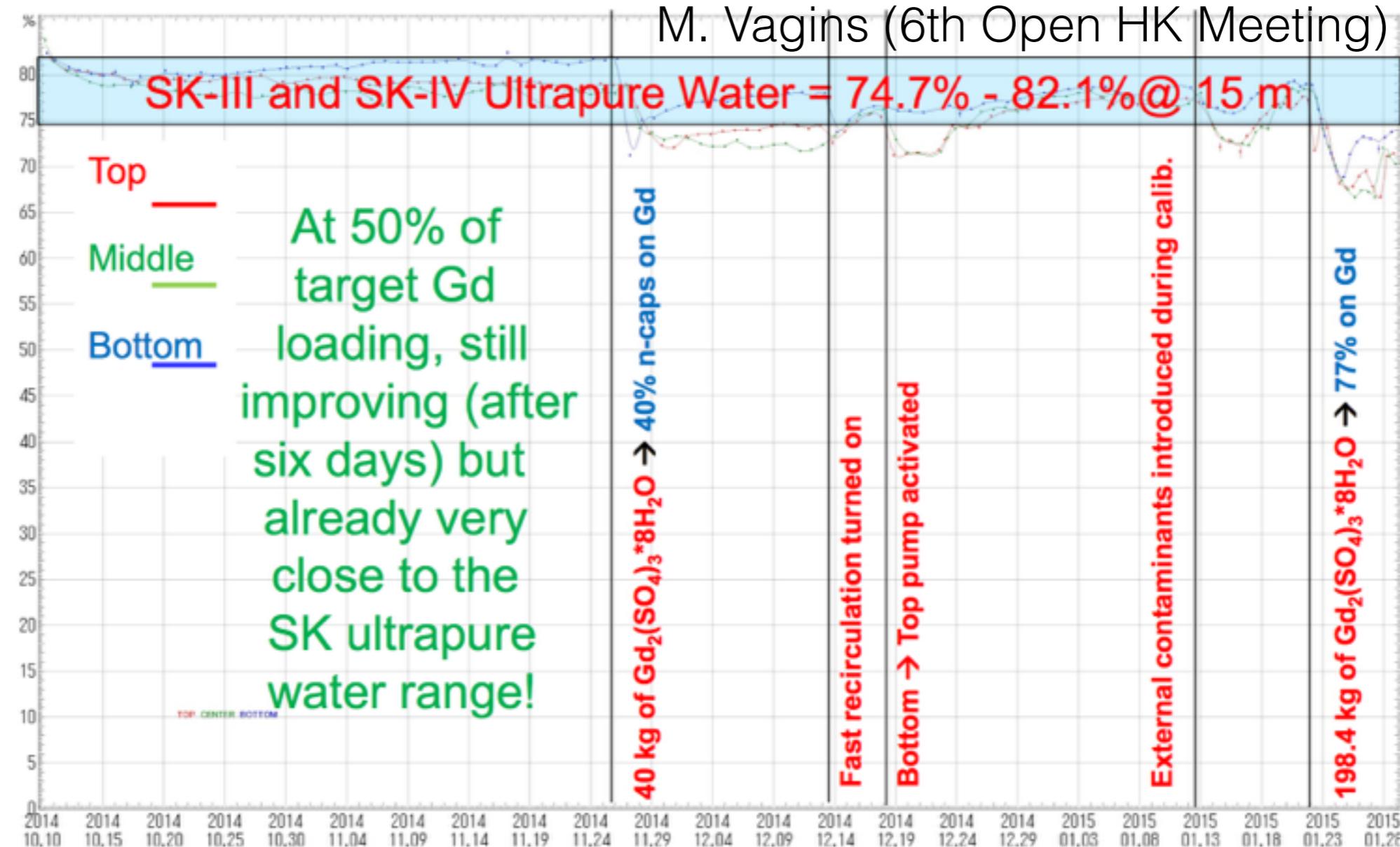
Scaled up design
for SK sized tank



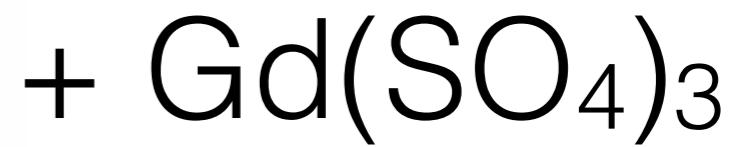
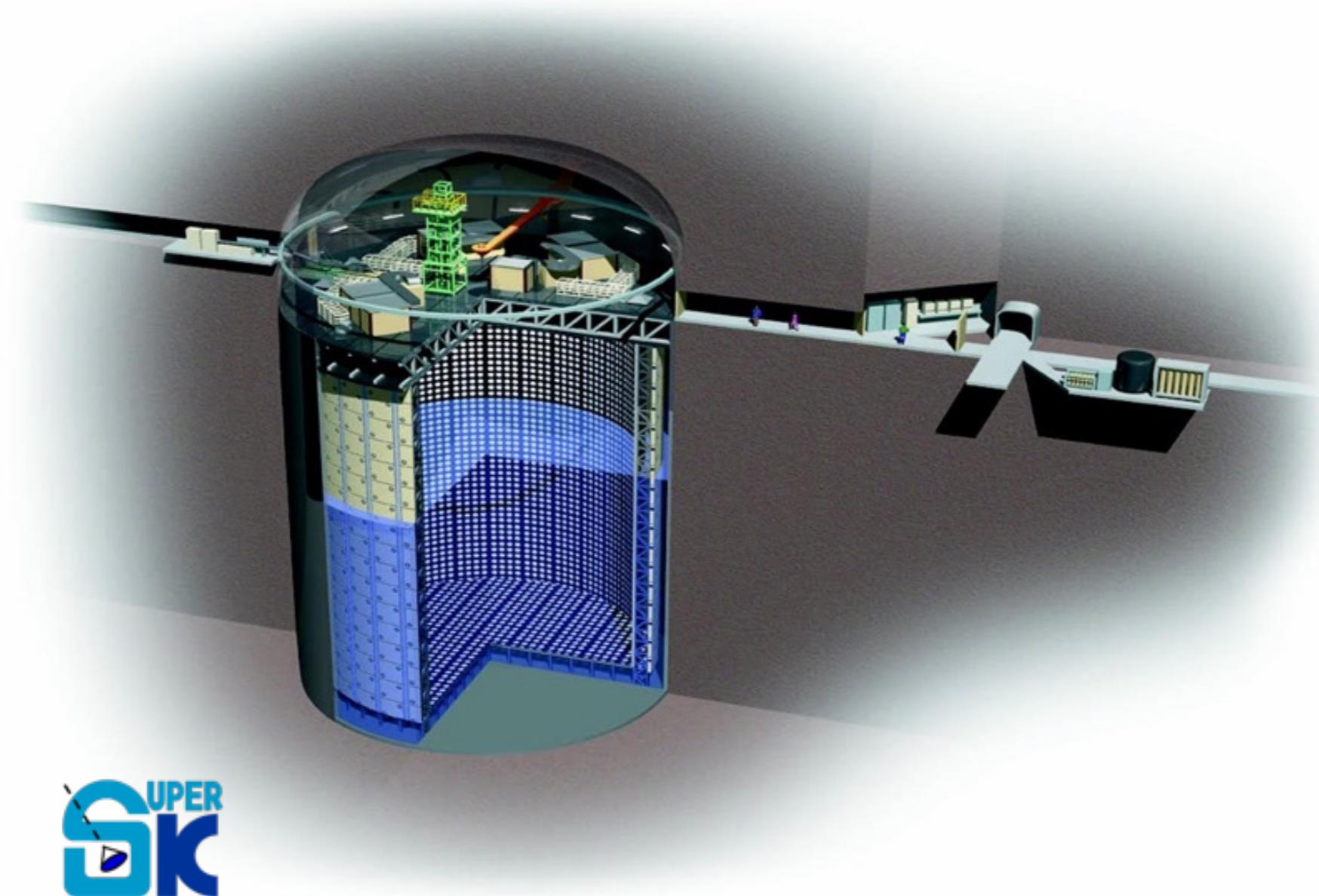
EGADs

(Evaluating Gadolinium's Action on Detector Systems)

Light @ 15 meters in the 200-ton tank (Gd water with PMT's)



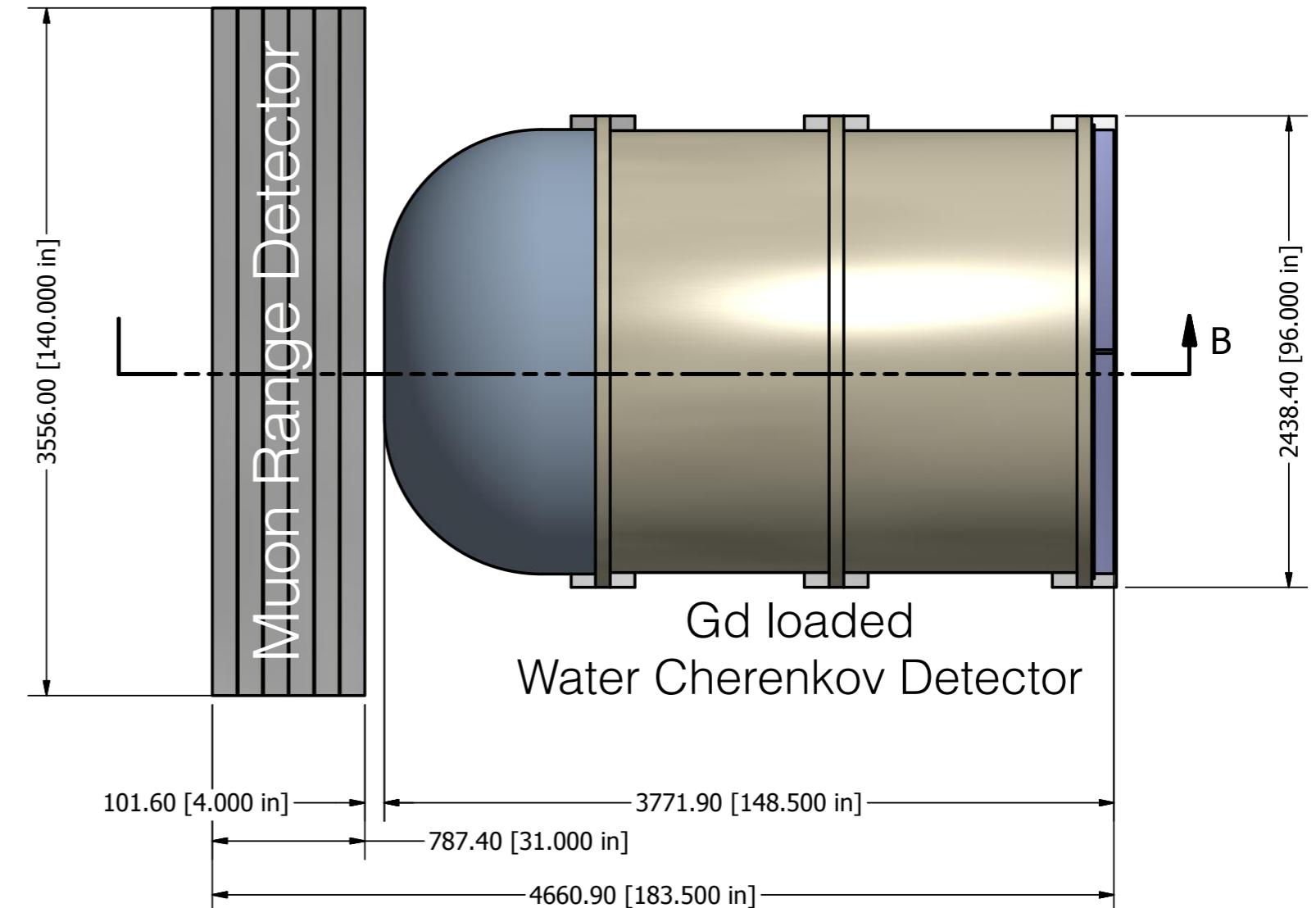
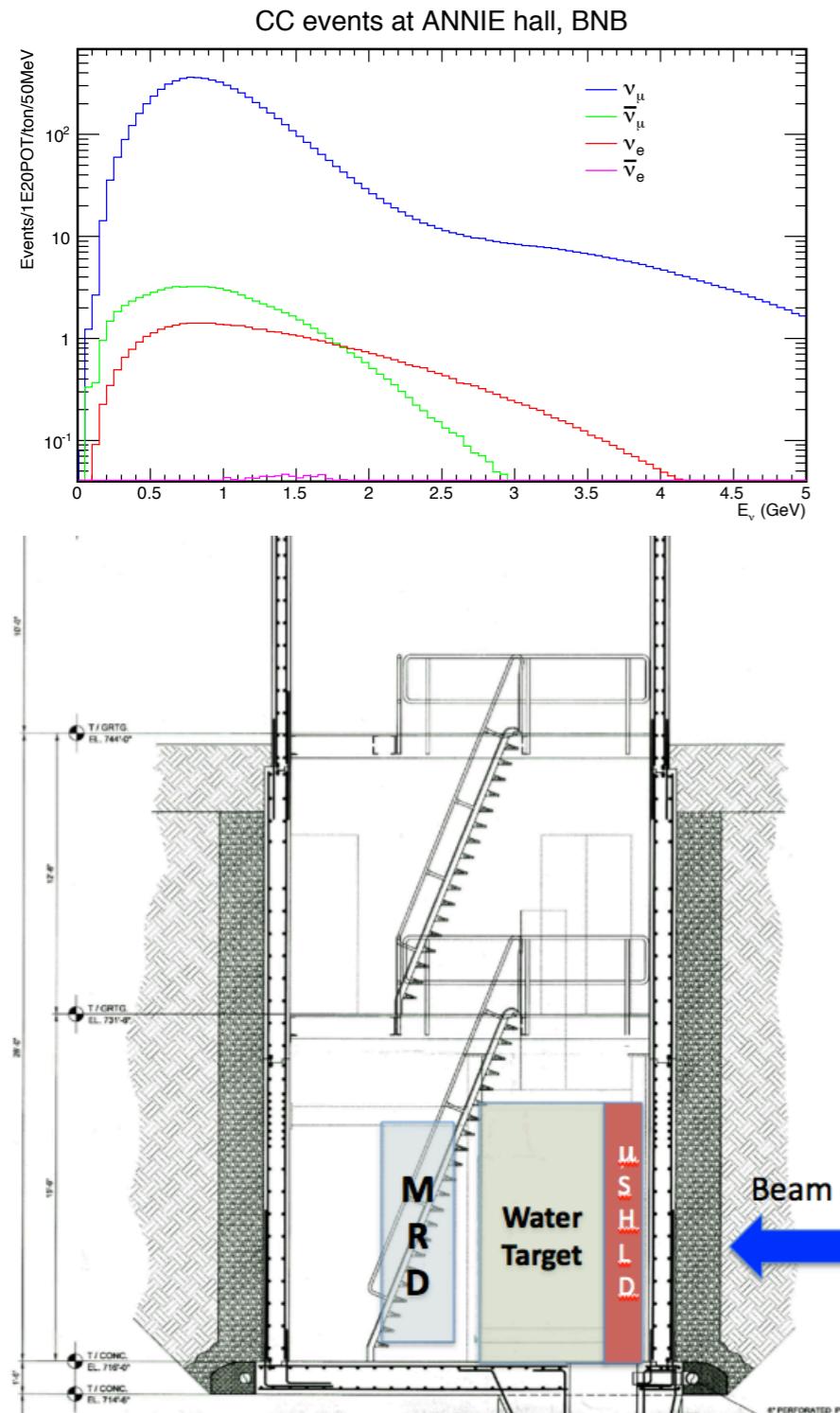
Super Kamiokande



In June 2015 the Super-K collaboration approved Gd-loading.
Gd is also an option for Hyper-K.

ANNIE

(Accelerator Neutrino Neutron Interaction Experiment)



Aim to measure neutron multiplicities for neutrino interactions on Oxygen in the few GeV range

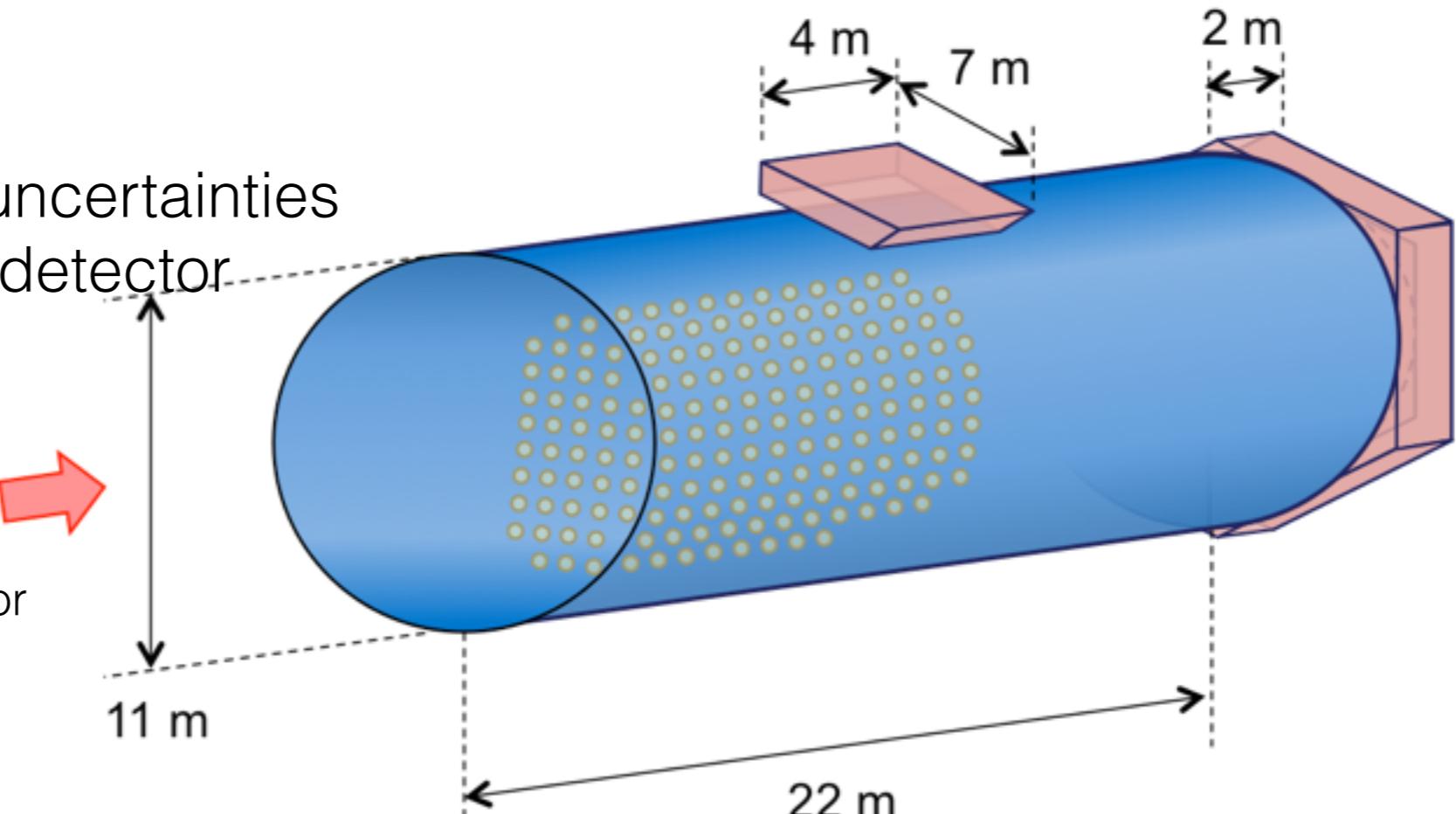
TITUS

Proposed Intermediate Water Cherenkov Detector for T2HK TITUS Detector

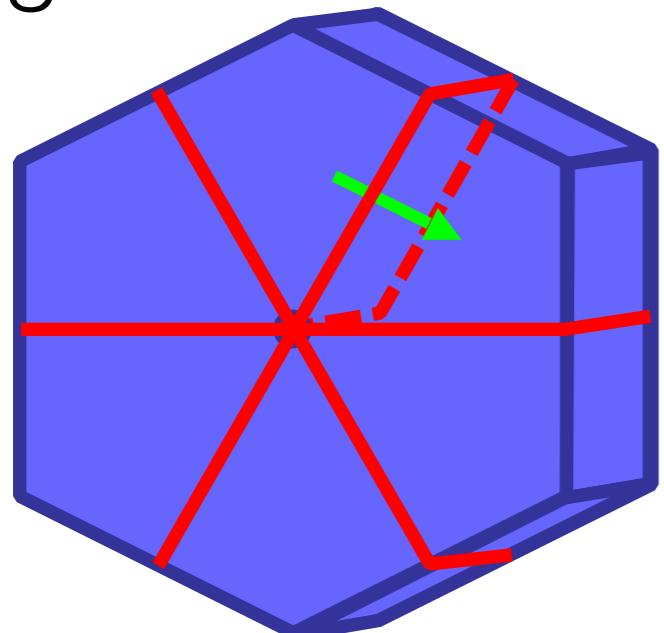
Maximise cancellation of uncertainties
between near and far detector

Identical target nucleus and
detector technologies

~2 km from beam source
match the flux at the far detector

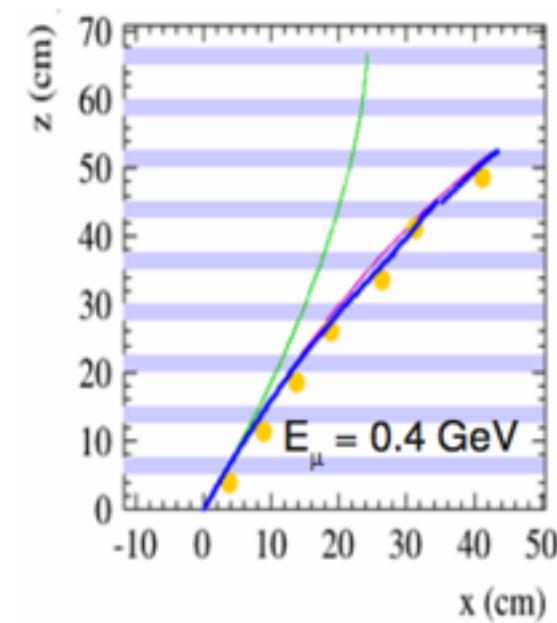
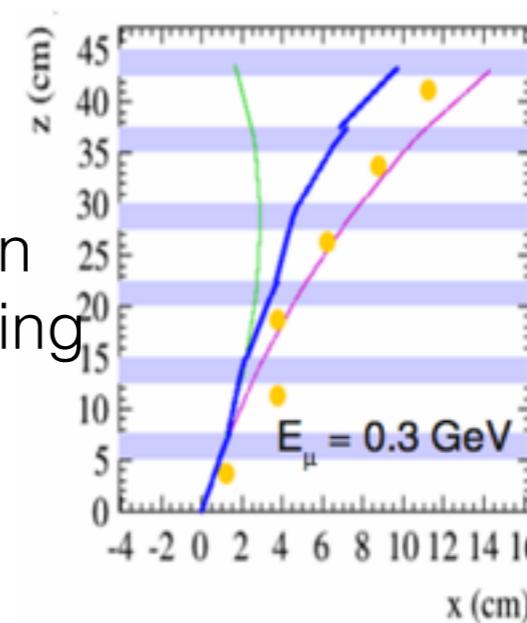


Magnetised Muon Range Detector



Measure momentum of
escaping muons.

In-situ cross-check of sign
selection with neutron tagging
method.



Gadolinium Doped Water Cherenkov Detectors

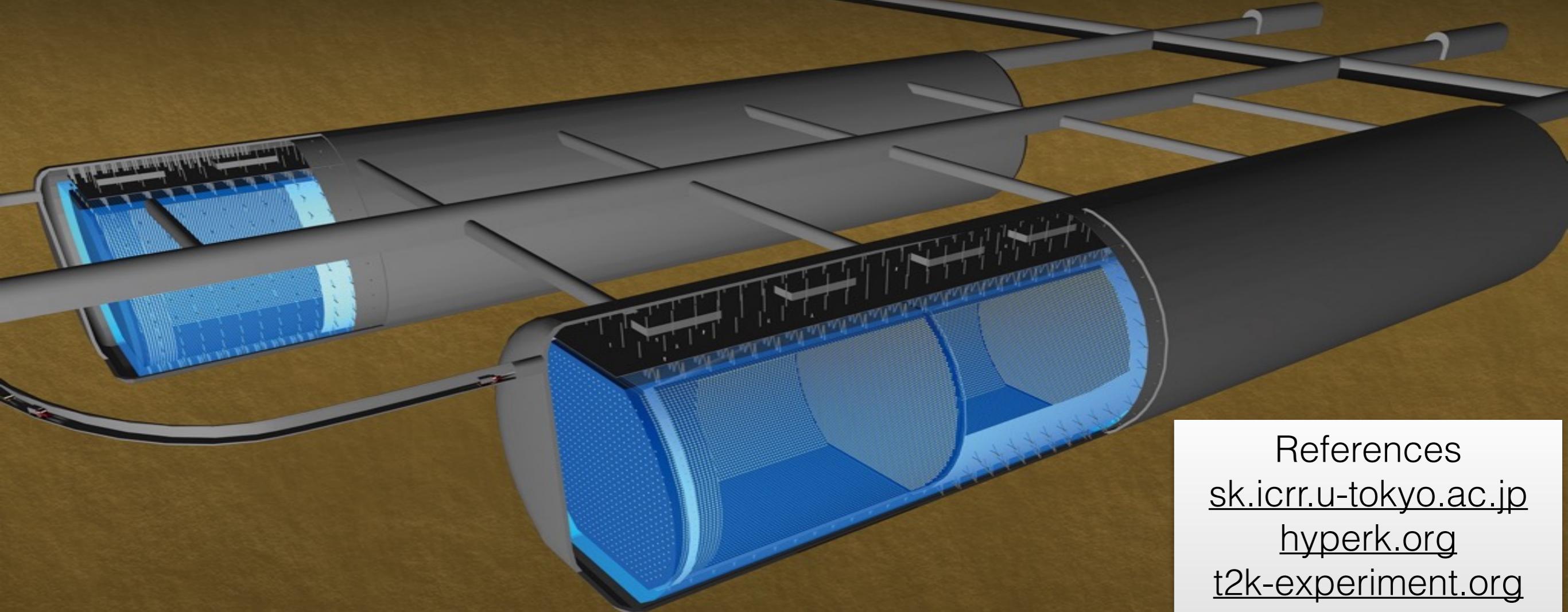
Neutron tagging with Gd-doped WC significantly extends the physics reach of large scale Water Cherenkov detectors.

Technical implementation has been successfully demonstrated (EGADs etc).

Gd-doping is the future for Super-K (and Hyper-K?).

To fully exploit this new technology, we need to make measurements of neutron multiplicity for ν -Oxygen interactions and build models that reproduce them.

Thank you for listening



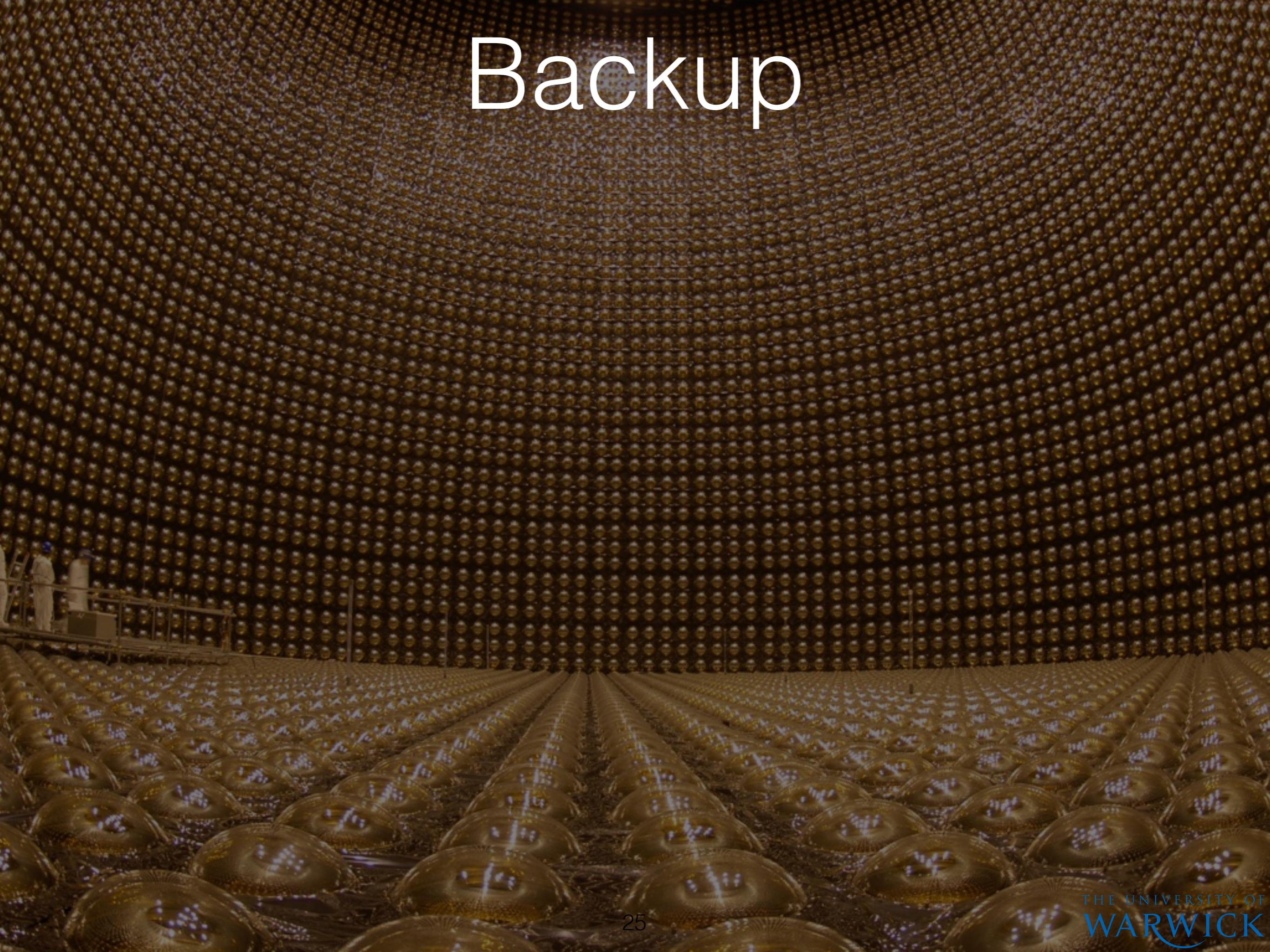
David Hadley
University of Warwick
29th May 2015

References

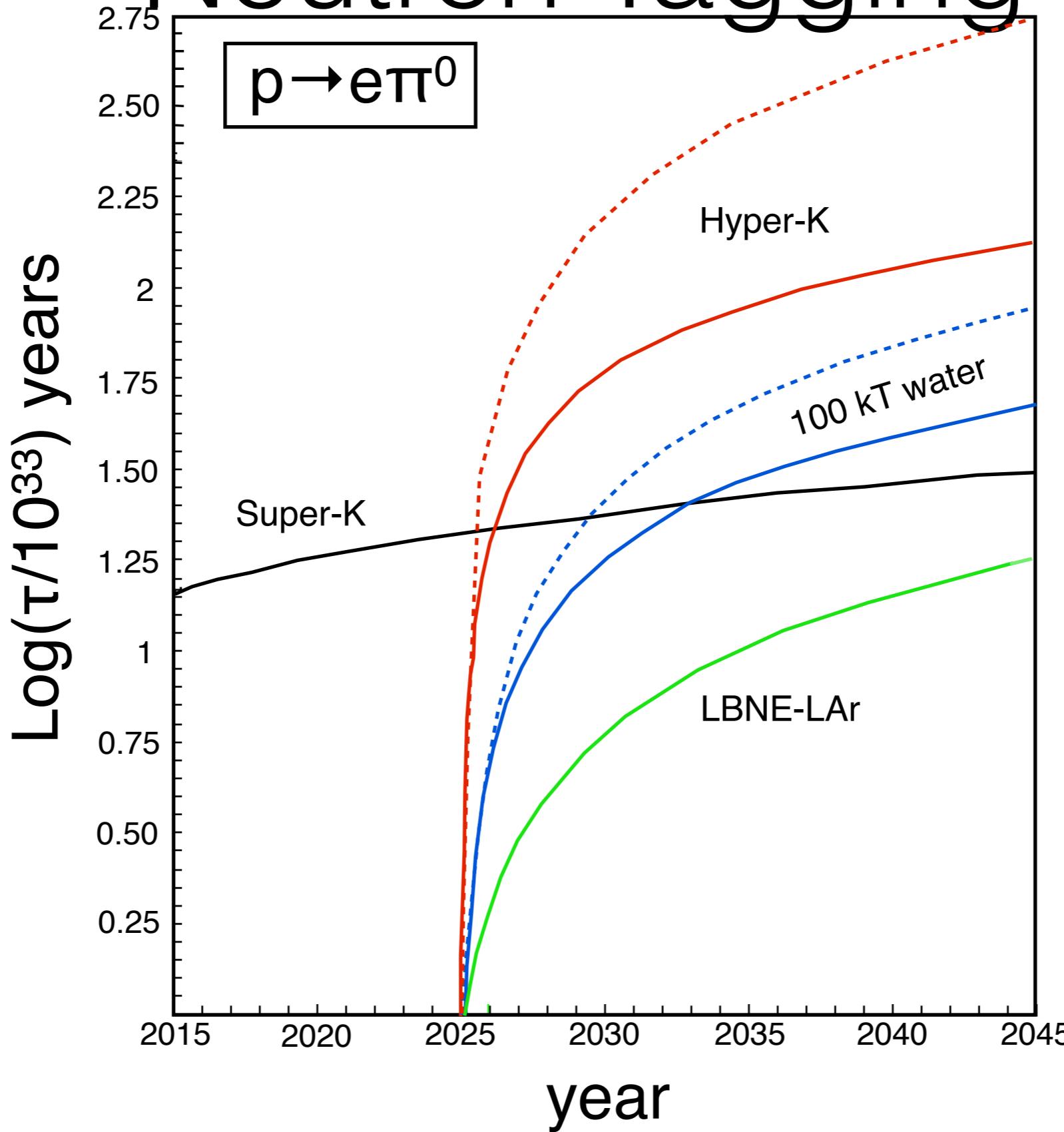
sk.icrr.u-tokyo.ac.jp
hyperk.org
t2k-experiment.org

arXiv:hep-ph/0309300
arXiv:1311.3738
arXiv:0811.0735
arXiv:1109.3262
arXiv:1201.1017
arXiv:1504.01480

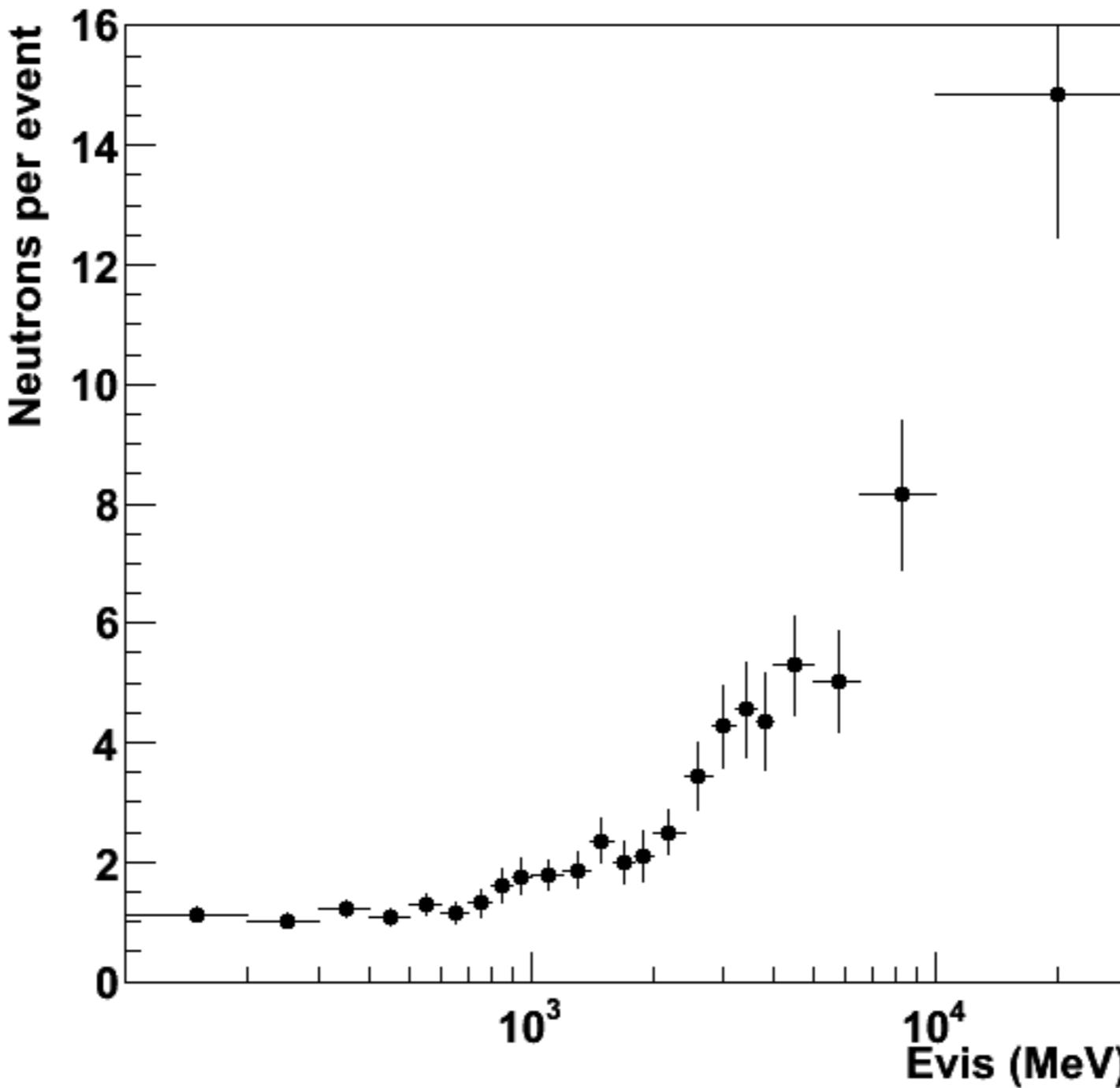
Backup



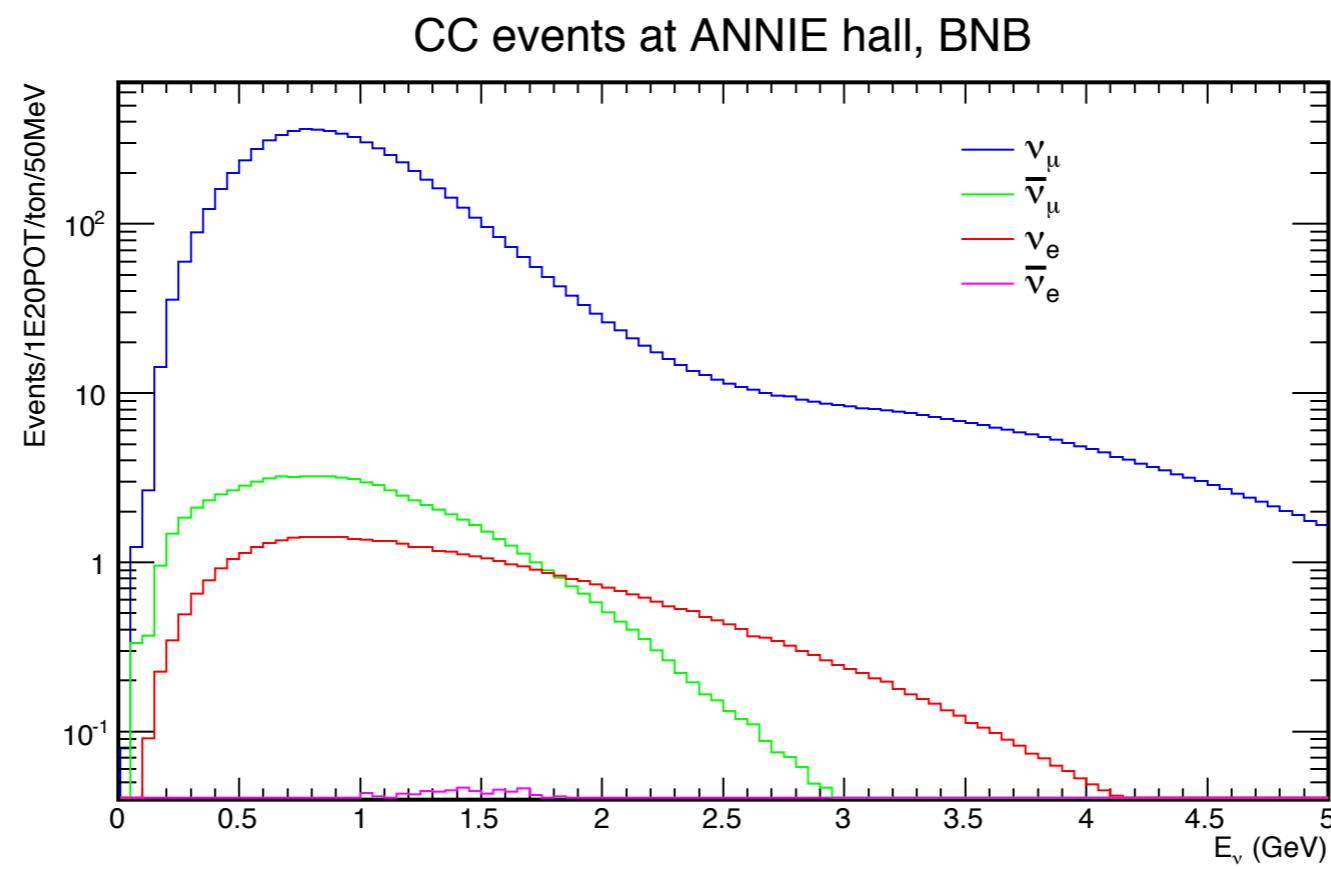
Proton Decay Limits with Neutron Tagging



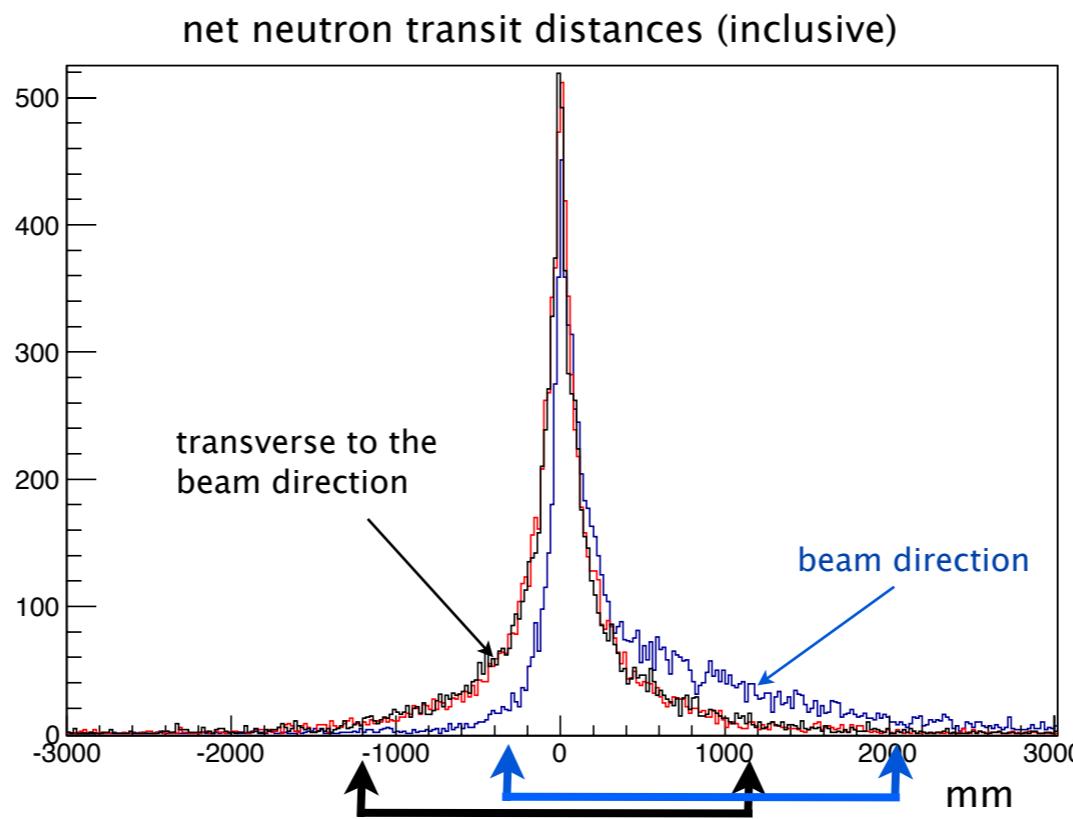
Super-K Measurements of Neutron Multiplicity



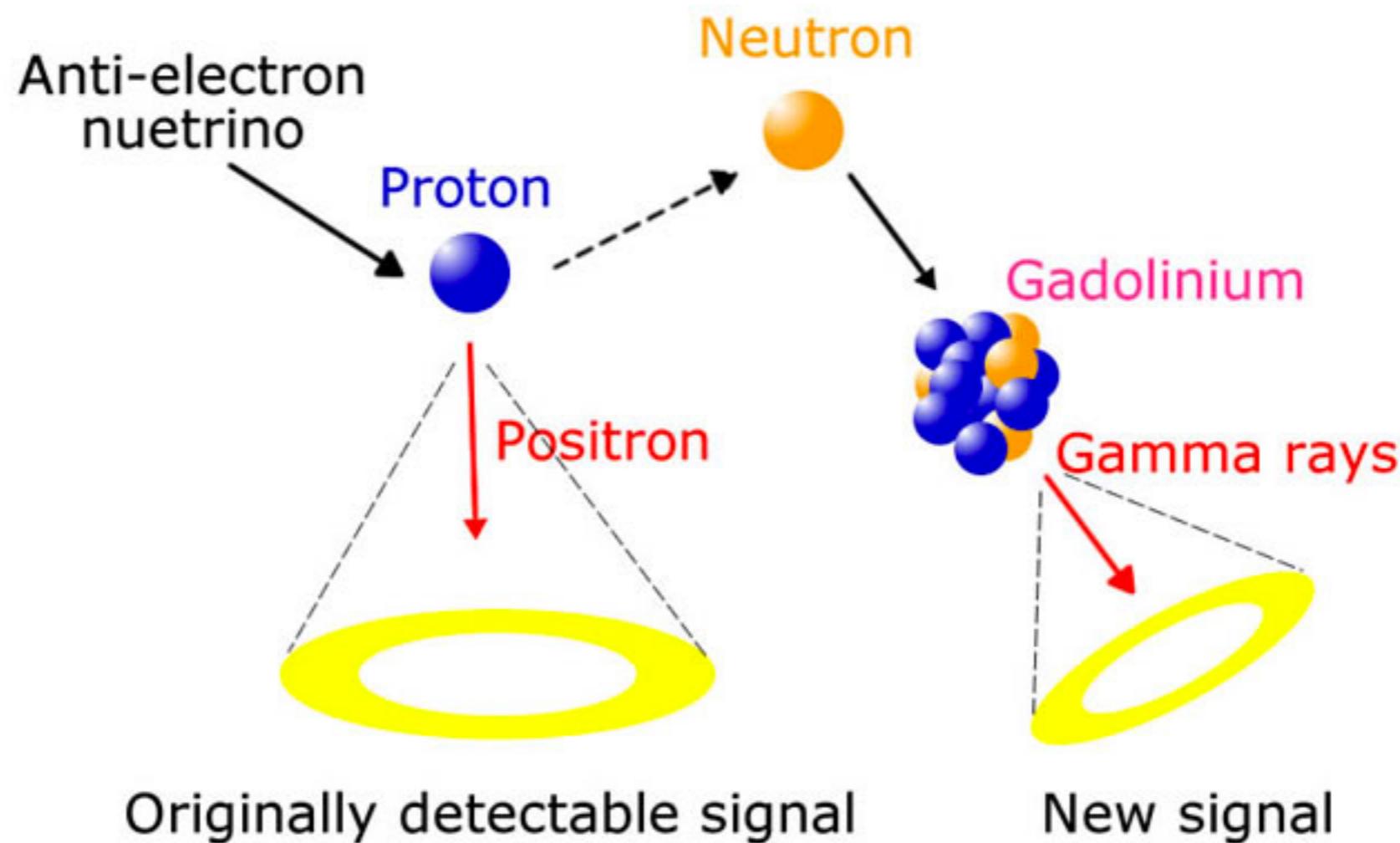
ANNIE Events



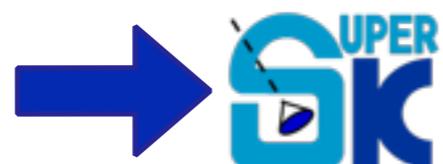
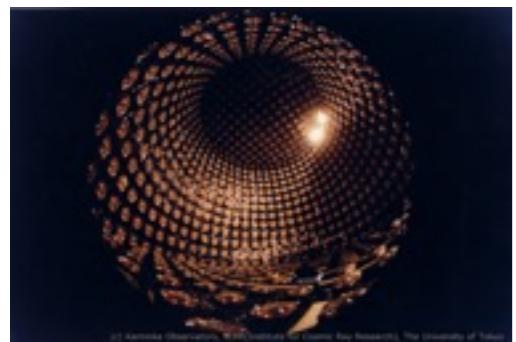
ANNIE Neutron Transit



Neutron Capture on Gd

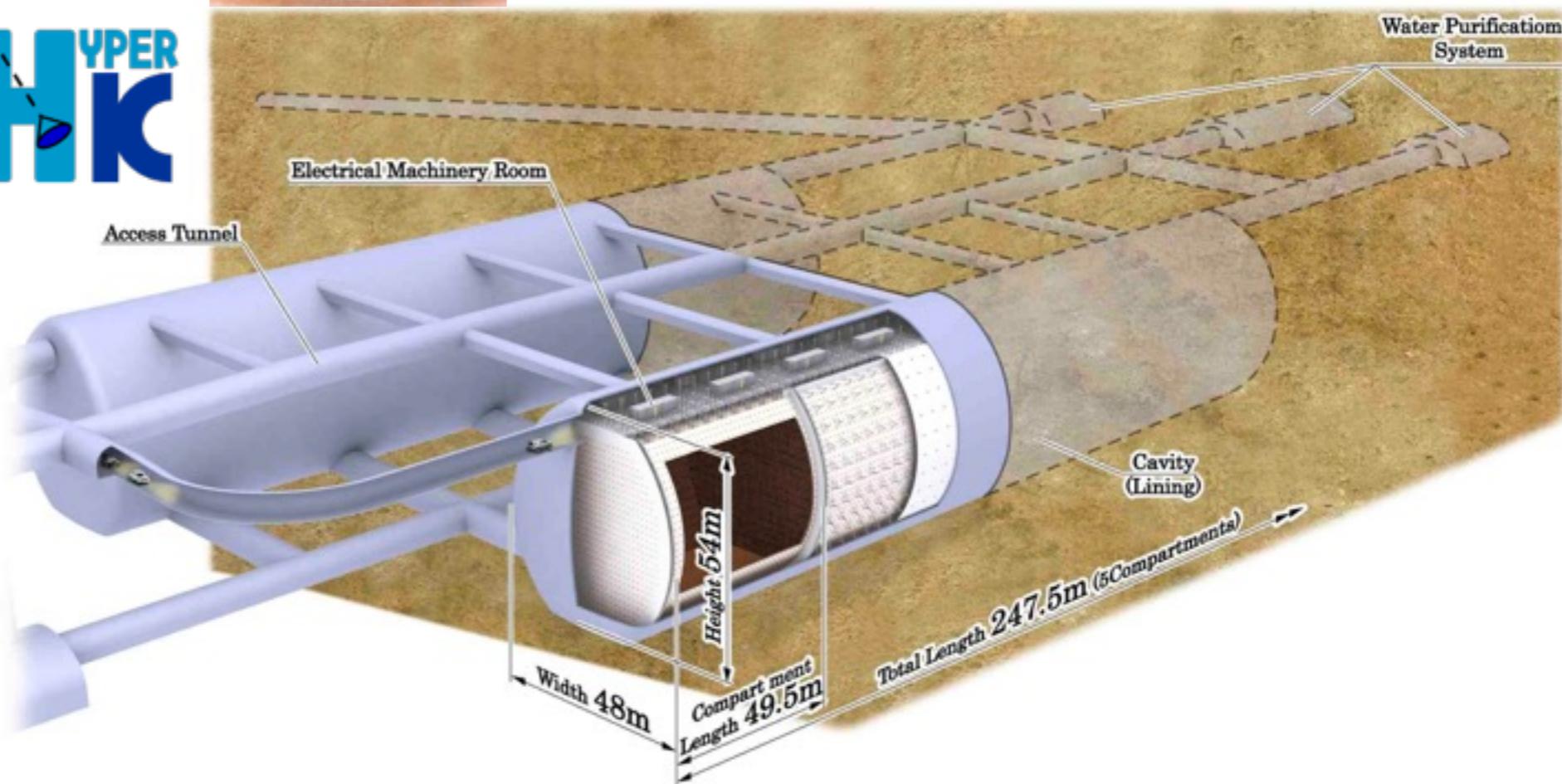


Kamiokande Detectors



Super-Kamiokande
22.5kt fiducial mass
(33x Kamiokande)

Kamiokande
680 tonne
fiducial mass
(1983)



Megaton scale Water Cherenkov detector
x25 larger fiducial volume than Super-K.
(202X)

Physics with Large Scale WC

Proton Decay

$$p \rightarrow e^+ + \pi^0$$

$>1.3 \times 10^{35}$ years 90% CL

$$p \rightarrow \bar{v} + K^+$$

$>3.2 \times 10^{34}$ years 90% CL



Broad physics programme.

Neutrinos

Solar



200 solar ν per day

Indirect dark matter search

Supernova

SN ~200,000 @ 10kPC

SN ~30-50 @ M31

Accelerator



Θ_{23} octant determination
3 σ for $\sin^2 \Theta_{23} > 0.56$ or $\sin^2 \Theta_{23} < 0.46$



Near Detector Development

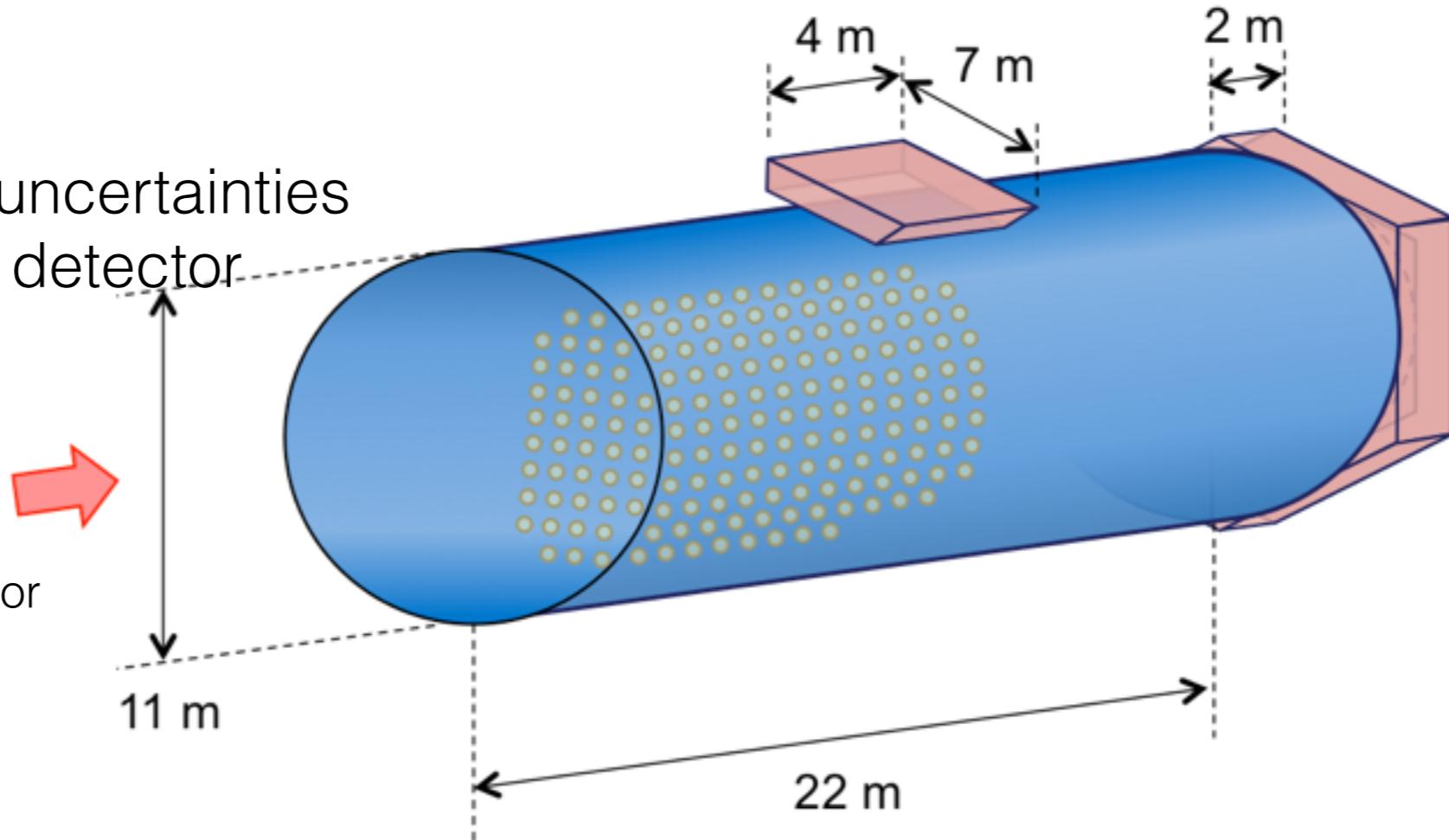
New Intermediate Water Cherenkov Detectors

TITUS Detector

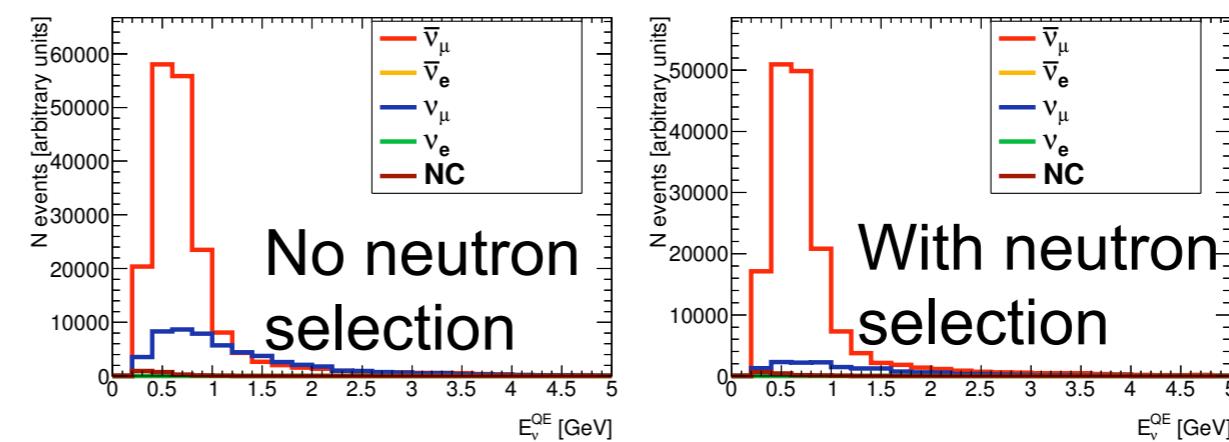
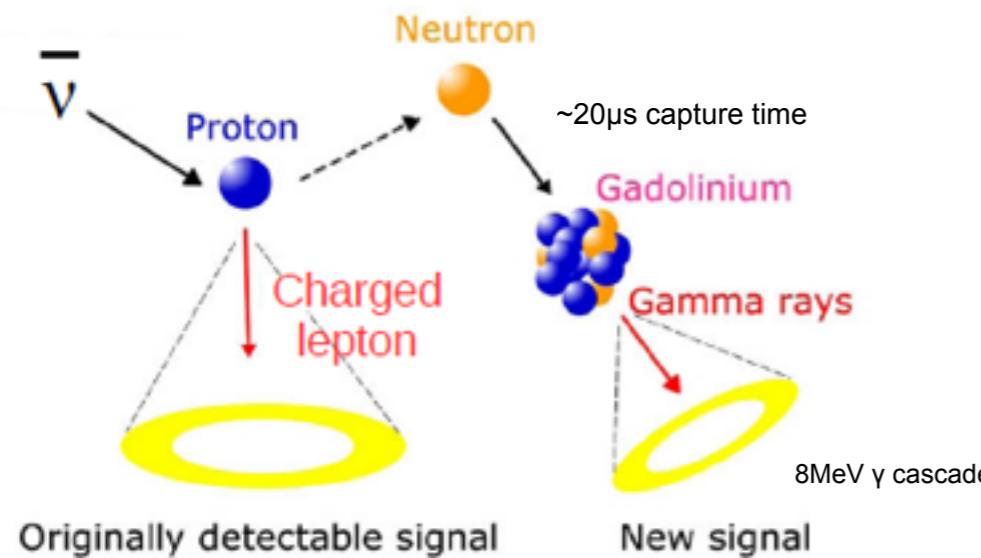
Maximise cancellation of uncertainties
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Neutron Capture on Gd



DSNB at GADZOOKS

