



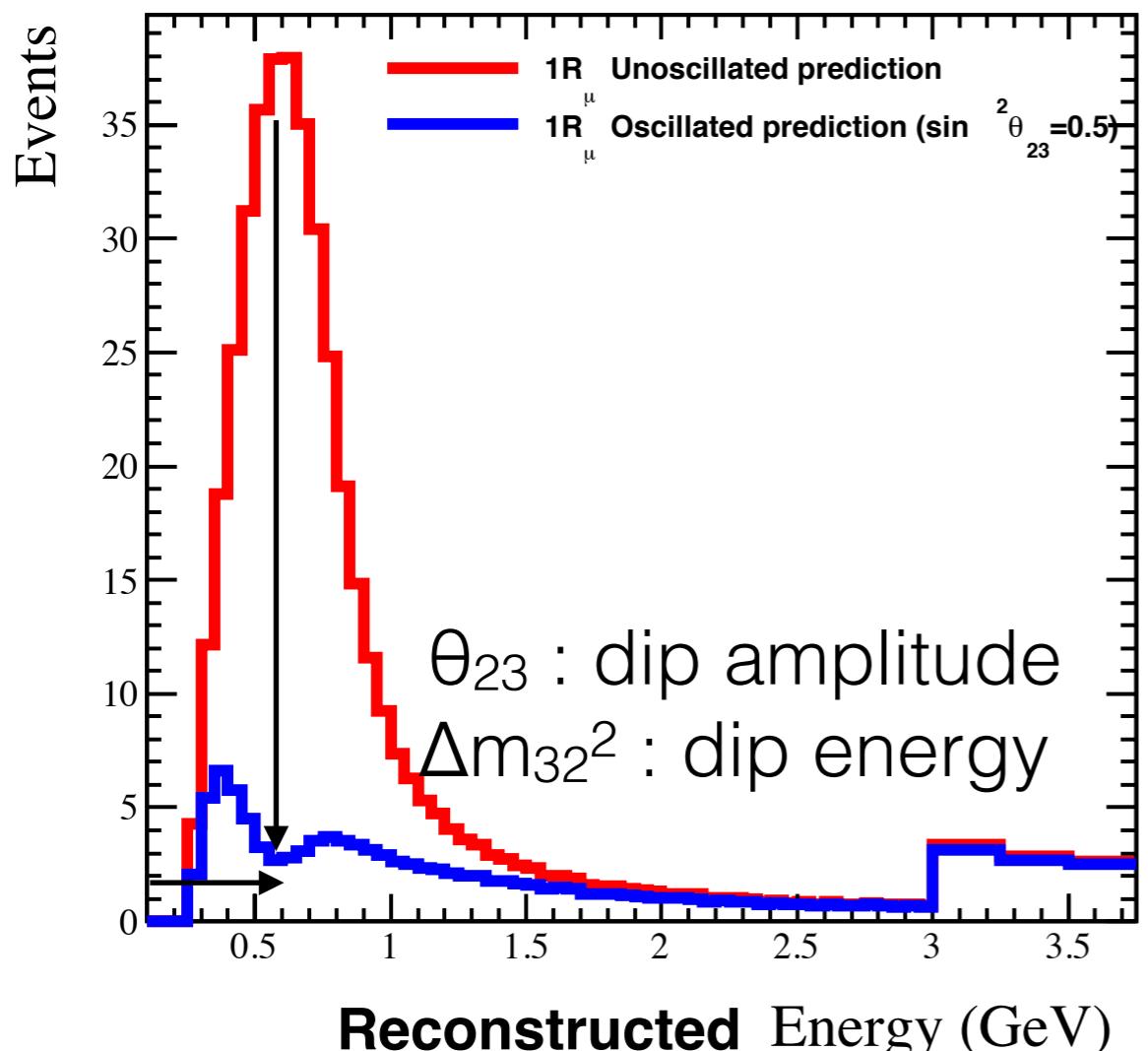
Systematic Uncertainties in Long Baseline Neutrino Oscillation Measurements



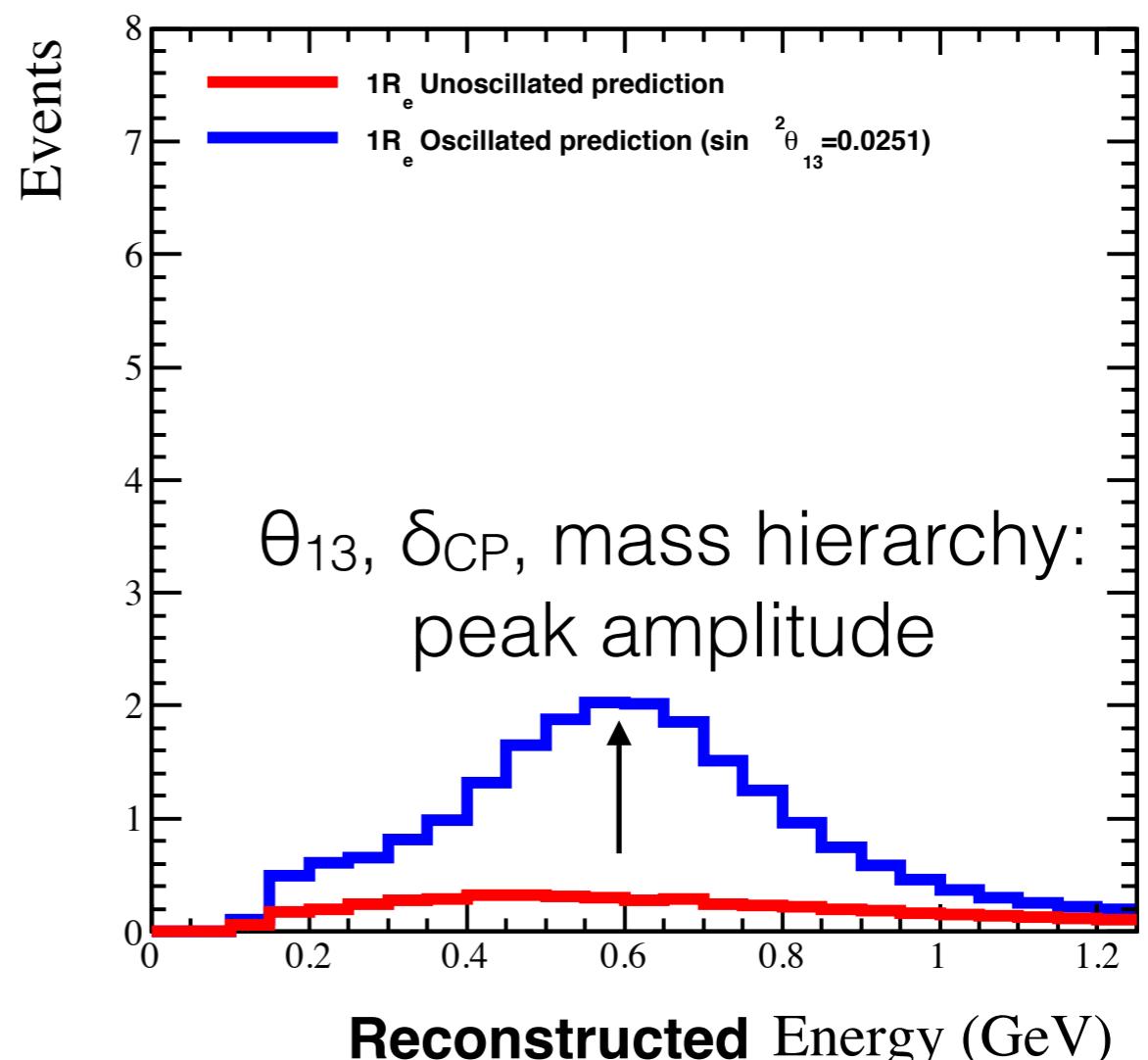
David Hadley
21st December 2017
Prospects in Neutrino Physics, NuPhys2017

What is measured?

ν_μ disappearance



ν_e appearance



Measurement precision limited by:

- Statistics
- Neutrino energy reconstruction
- Knowledge of unoscillated spectrum and background contamination

Where do systematic uncertainties enter?



$$R(\vec{x}_{\text{reco}}) = \int \Phi(E_\nu) \times \sigma(E_\nu, \vec{x}_{\text{true}}) \times \varepsilon(\vec{x}_{\text{true}}, \vec{x}_{\text{reco}}) \times P(E_\nu, \vec{\theta}) dE_\nu d\vec{x}_{\text{true}}$$

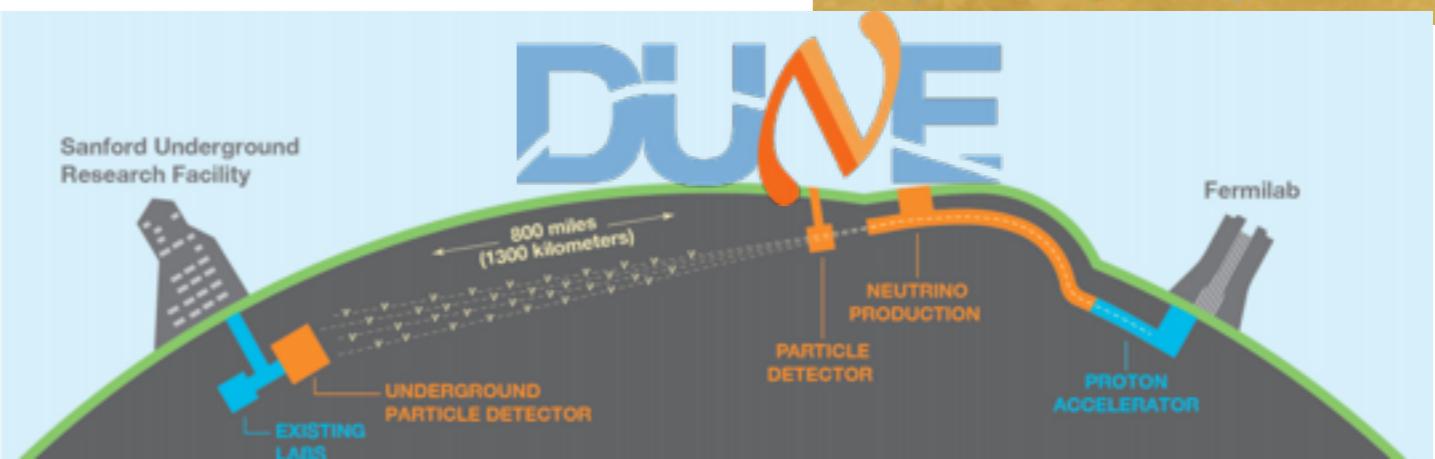
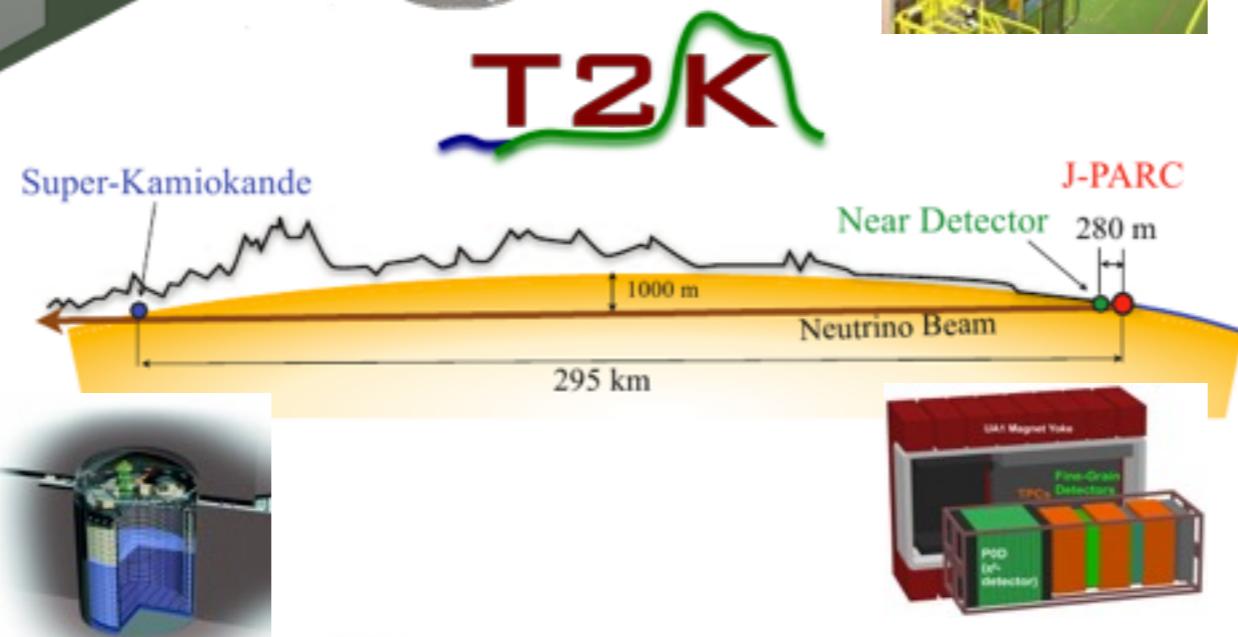
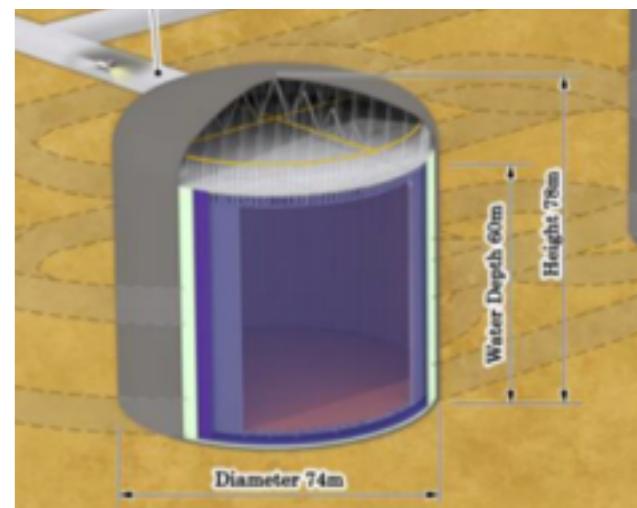
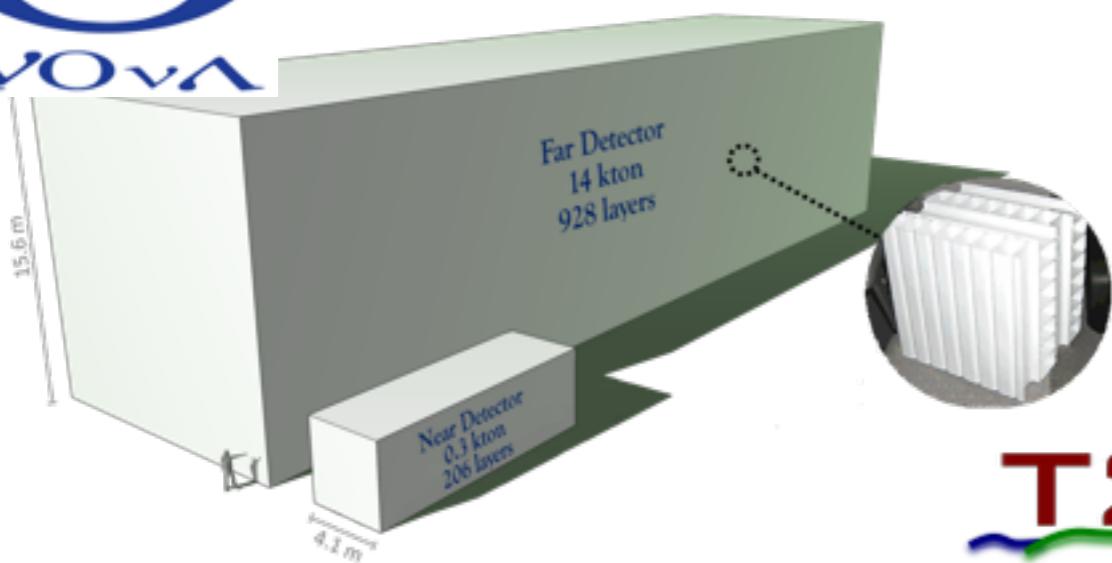
Neutrino Flux Neutrino nucleus interaction model Detector efficiency and resolution Oscillation Probability

Four arrows point from the text labels below the equation to the terms in the equation above them: 'Neutrino Flux' points to $\Phi(E_\nu)$, 'Neutrino nucleus interaction model' points to $\sigma(E_\nu, \vec{x}_{\text{true}})$, 'Detector efficiency and resolution' points to $\varepsilon(\vec{x}_{\text{true}}, \vec{x}_{\text{reco}})$, and 'Oscillation Probability' points to $P(E_\nu, \vec{\theta})$.

Use measurements at near detector to constrain $(\Phi \times \sigma \times \varepsilon)$

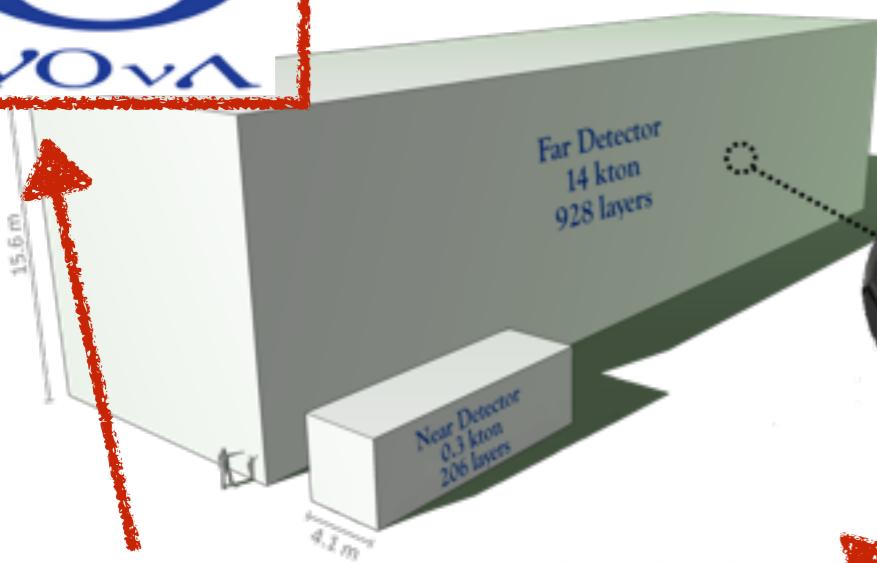
Cancellation of uncertainties is not perfect as no oscillation at the near detector and Φ and ε may differ

Accelerator based Neutrino Oscillation Experiments

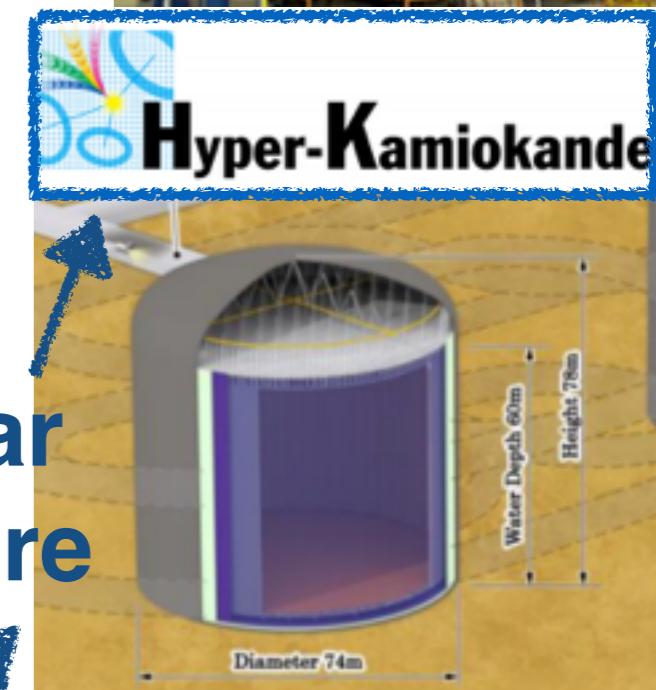
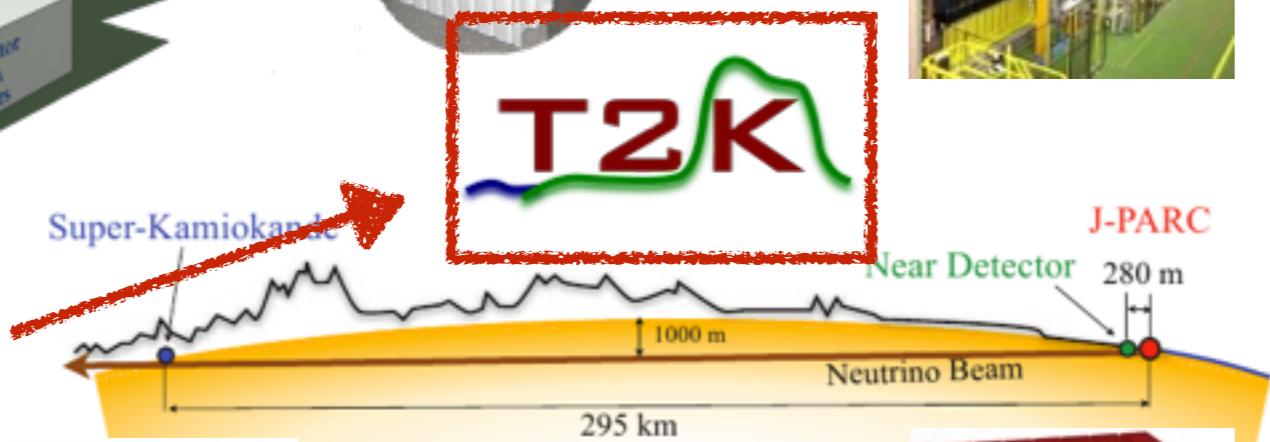


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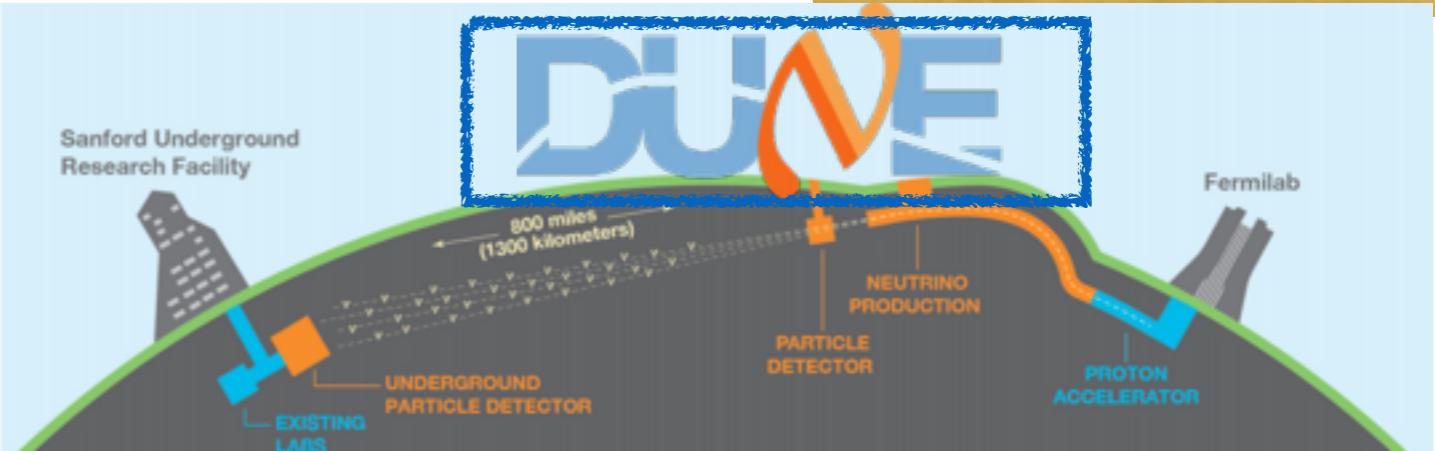
Accelerator based Neutrino Oscillation Experiments



**Currently
running long
baseline
experiments**



**Near
future**



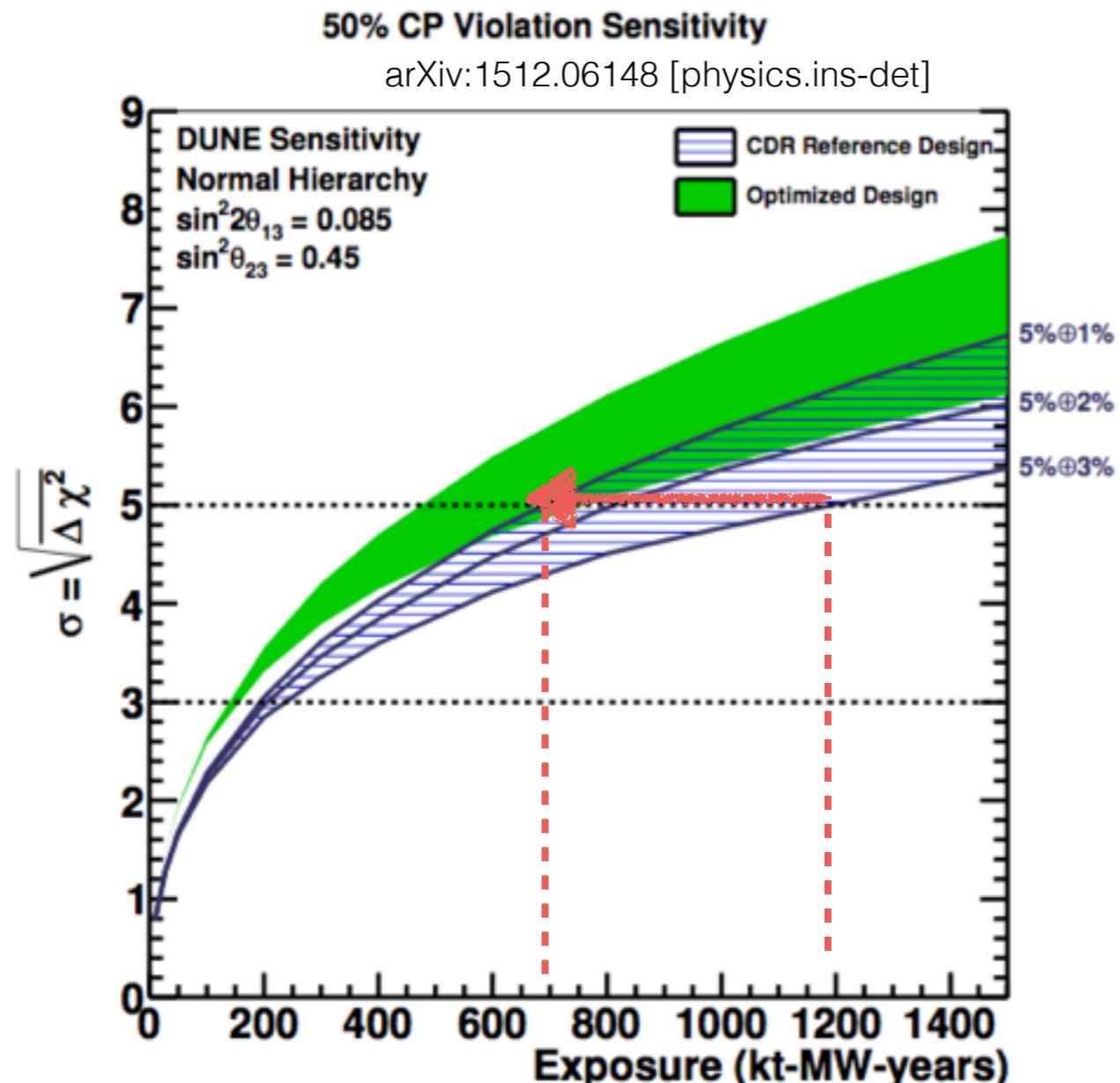
Statistics



Experiment	$\nu_e + \bar{\nu}_e$	$1/\sqrt{N}$	Ref.
T2K (current)	74 + 7	12% + 40%	2.2×10^{21} POT
NOvA (current)	33	17%	FERMILAB-PUB-17-065-ND
NOvA (projected)	110 + 50	10% + 14%	arXiv:1409.7469 [hep-ex]
T2K-I (projected)	150 + 50	8% + 14%	7.8×10^{21} POT, arXiv:1409.7469 [hep-ex]
T2K-II	470 + 130	5% + 9%	20×10^{21} POT, arXiv:1607.08004 [hep-ex]
Hyper-K	2900 + 2700	2% + 2%	10 yrs 2-tank staged KEK Preprint 2016-21
DUNE	1200 + 350	3% + 5%	$3.5+3.5$ yrs x 40kt @ 1.07 MW arXiv:1512.06148 [physics.ins-det]

Current appearance measurements stats dominate
 $O(10^3) \nu_e$ at future experiments \rightarrow demands $\sim 2\%$ systematics
 $O(10^4) \nu_\mu \rightarrow$ need systematics as good as we can get!

Systematic Uncertainties



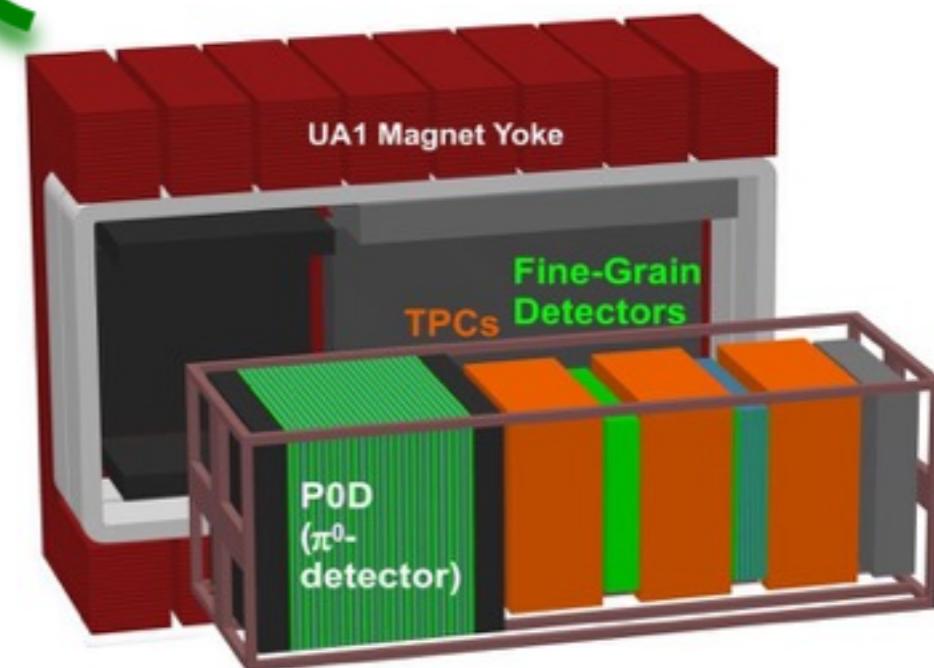
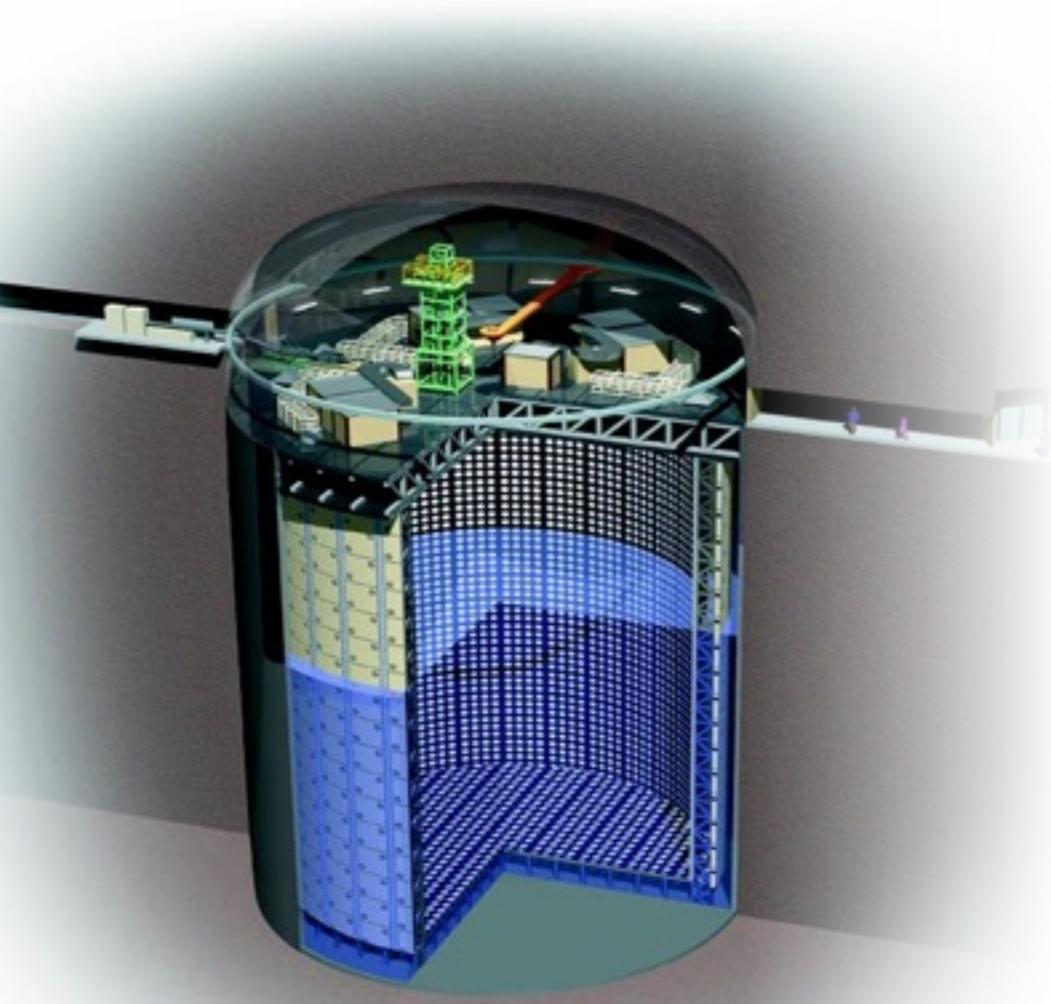
Reduction in systematic uncertainties
can be equivalent to significant boost
in exposure



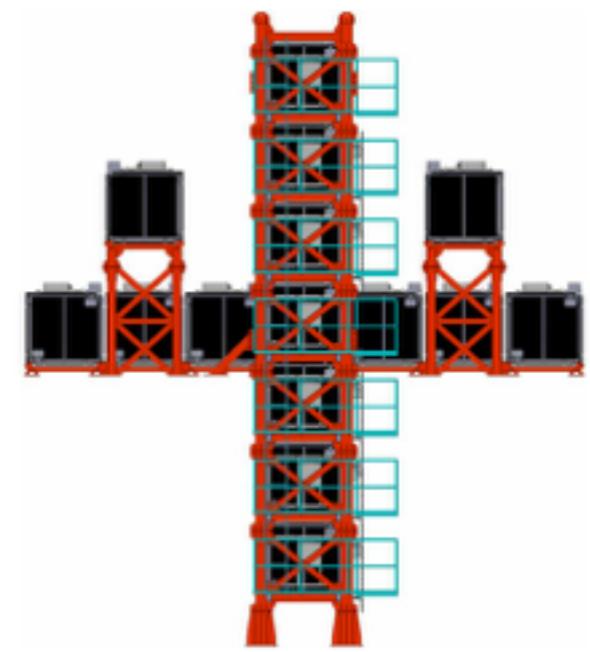
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Far Detector
(Super-K)



Near Detectors
(ND280+INGRID)

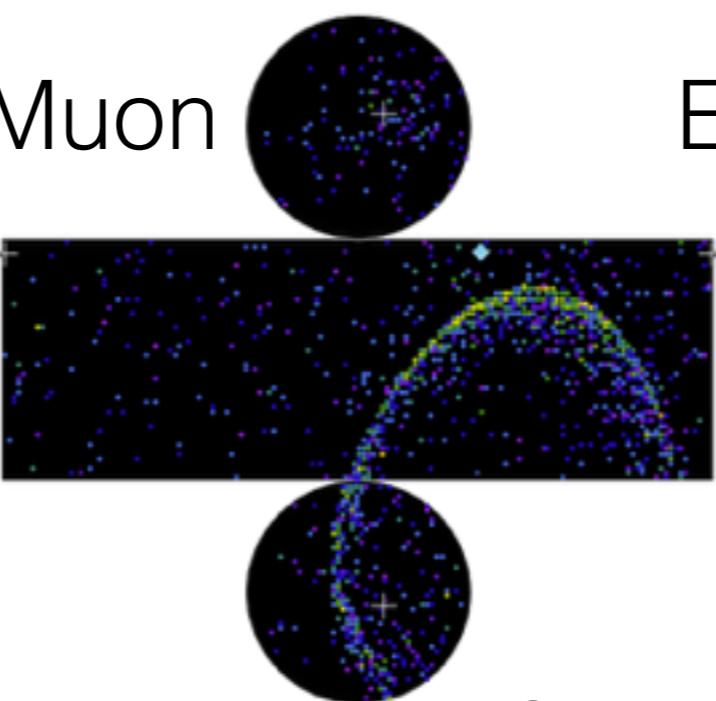




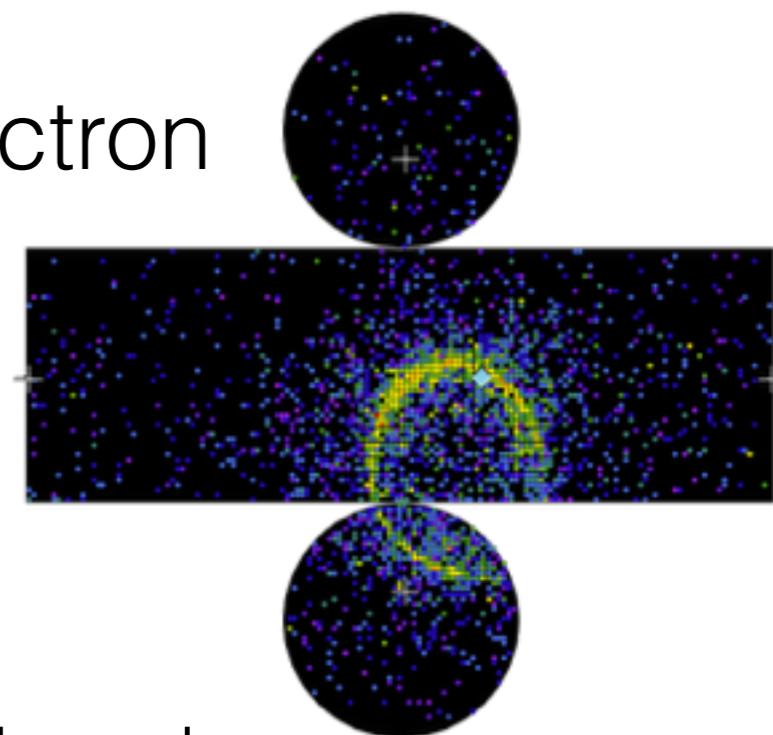
Water Cherenkov Far Detector
>22.5 kt fiducial mass



Muon



Electron



Oxygen target
 4π acceptance

Energy reconstruction from lepton kinematics

Blind to particles below Cherenkov threshold
for protons $< 1.1 \text{ GeV}/c$.
(neutron counting possible with SK-Gd)

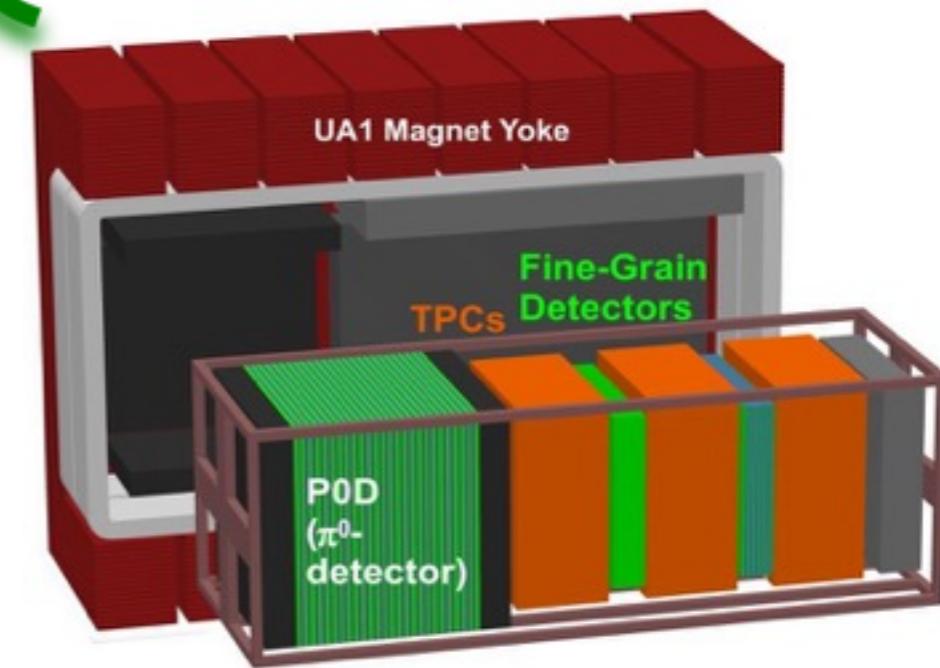


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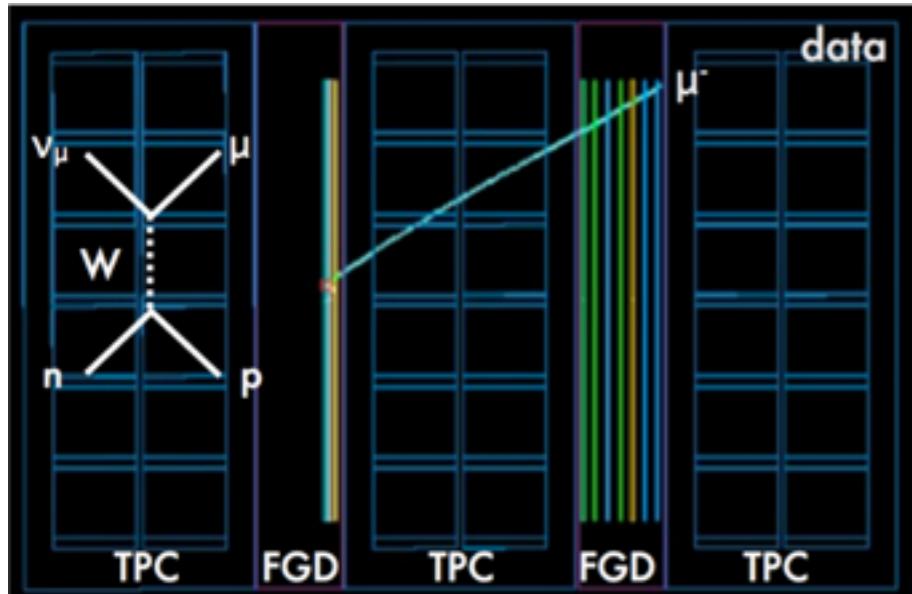
Carbon and Oxygen target materials

Acceptance differs from far detector

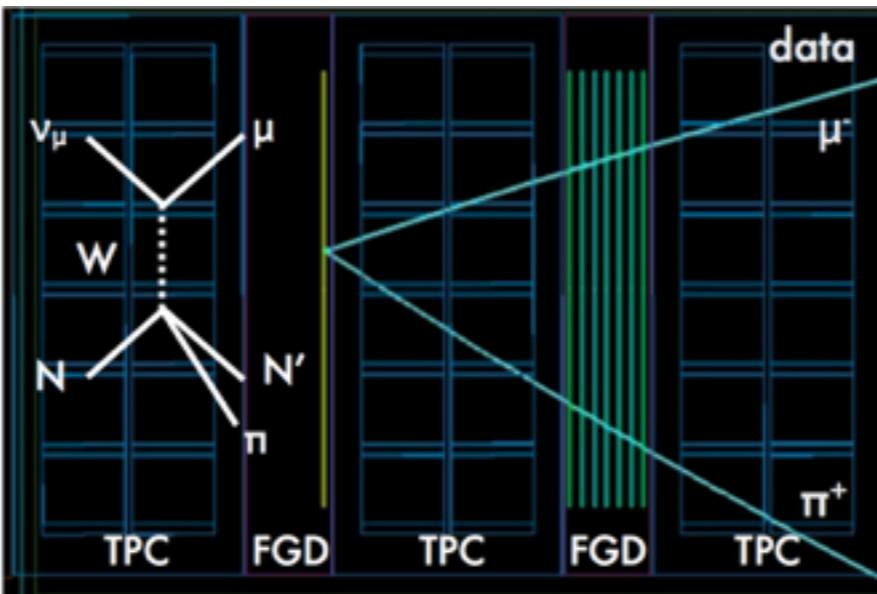
Magnetic field for sign selection



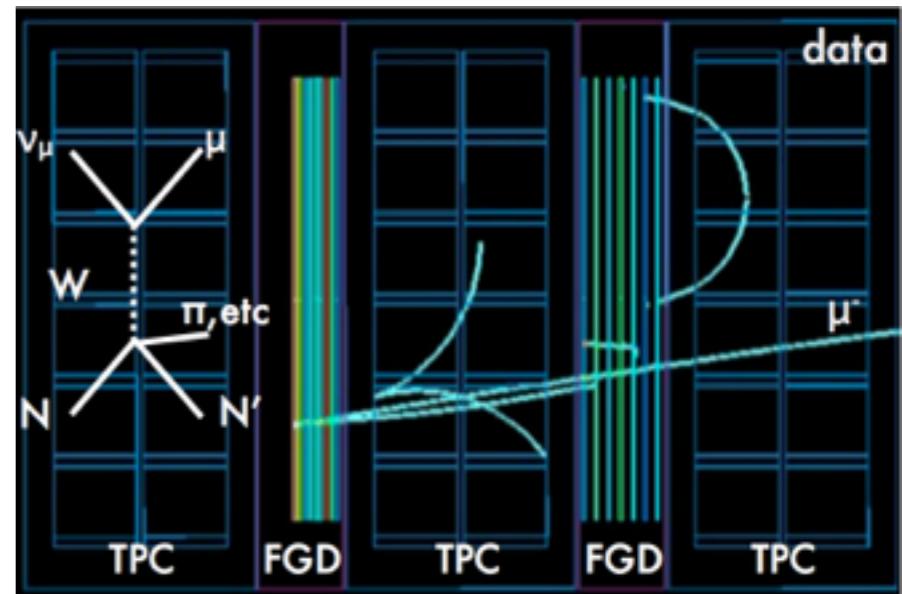
Near Detector (ND280)



CC 1 μ + 0 π + X

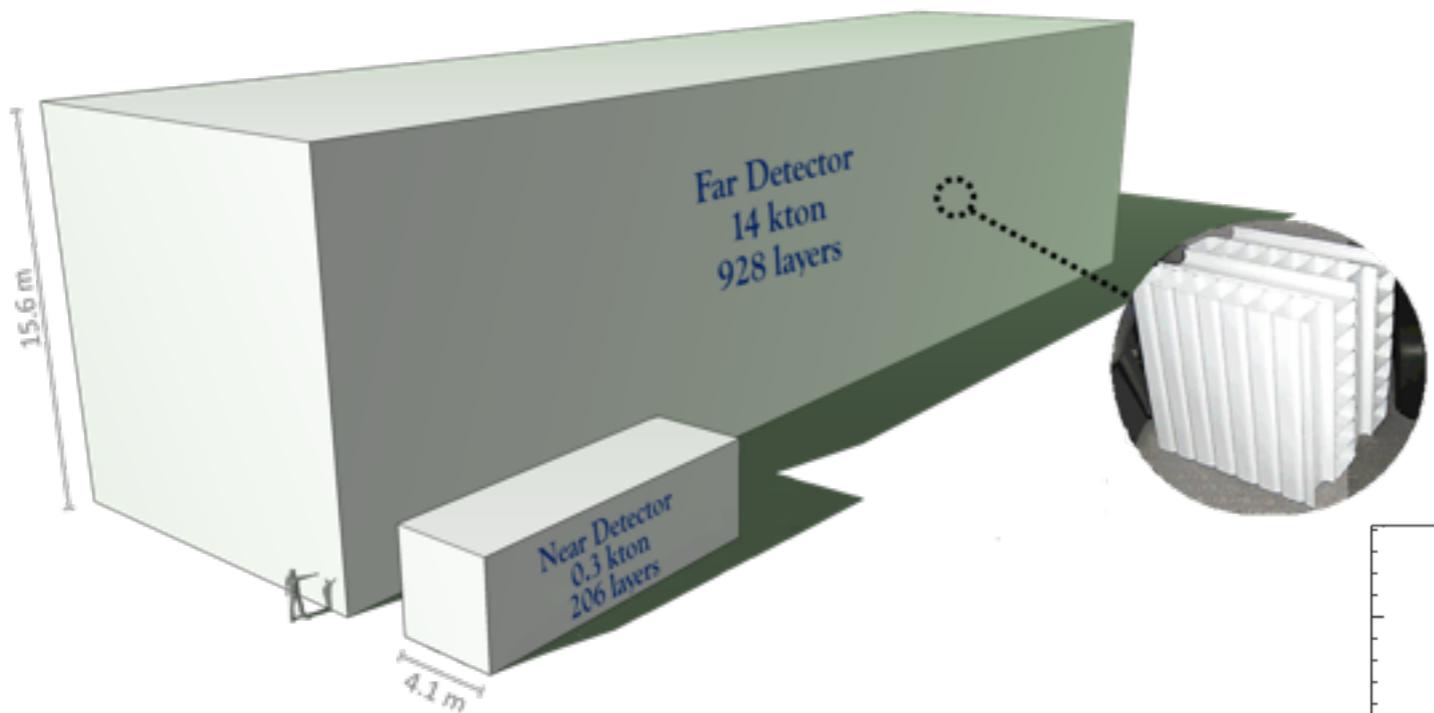


CC 1 μ + 1 π^+ + X



CC other

NOvA Experiment

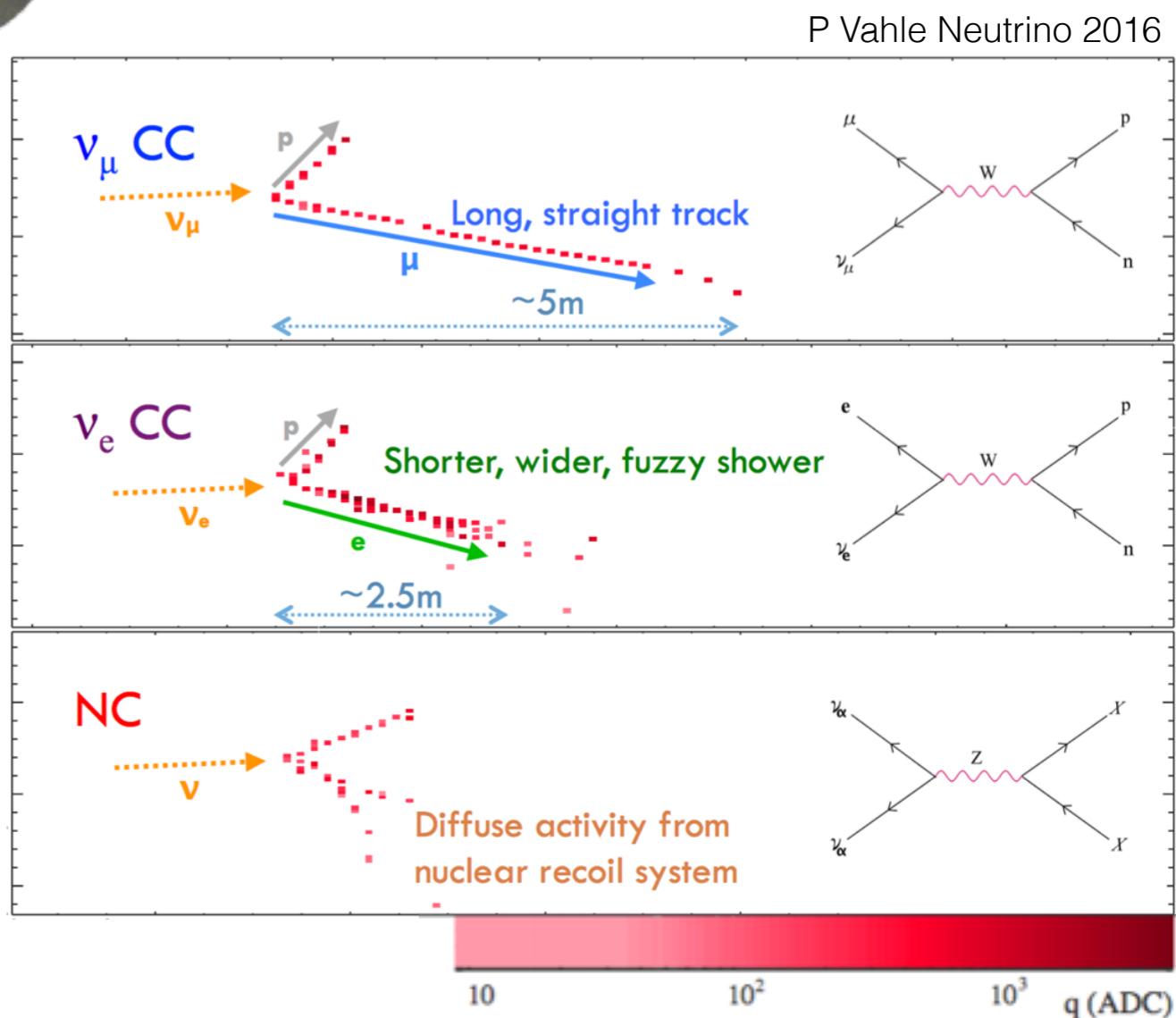


Liquid Scintillator tracking calorimeter

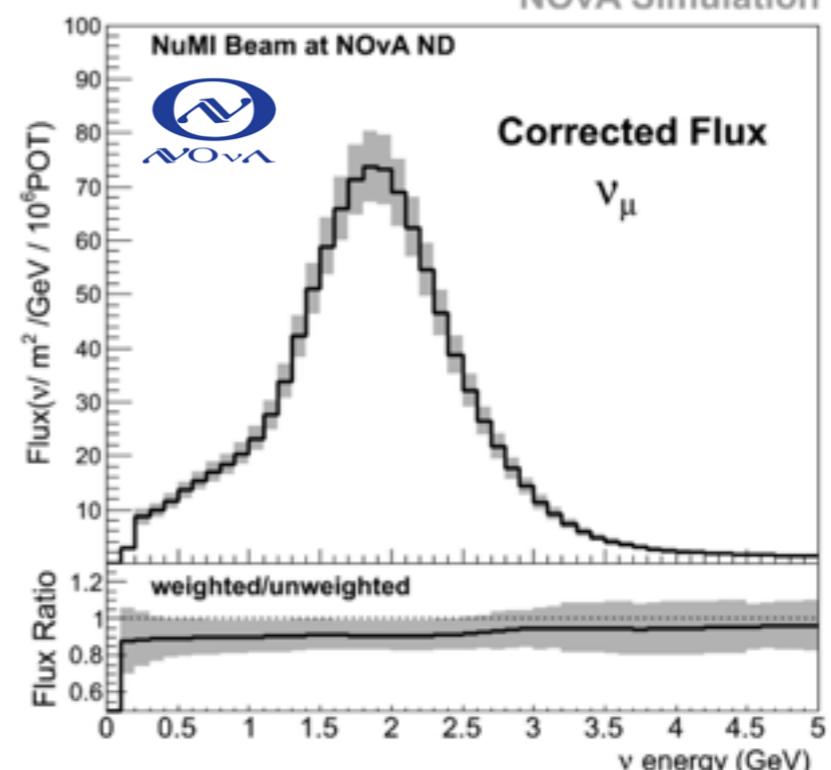
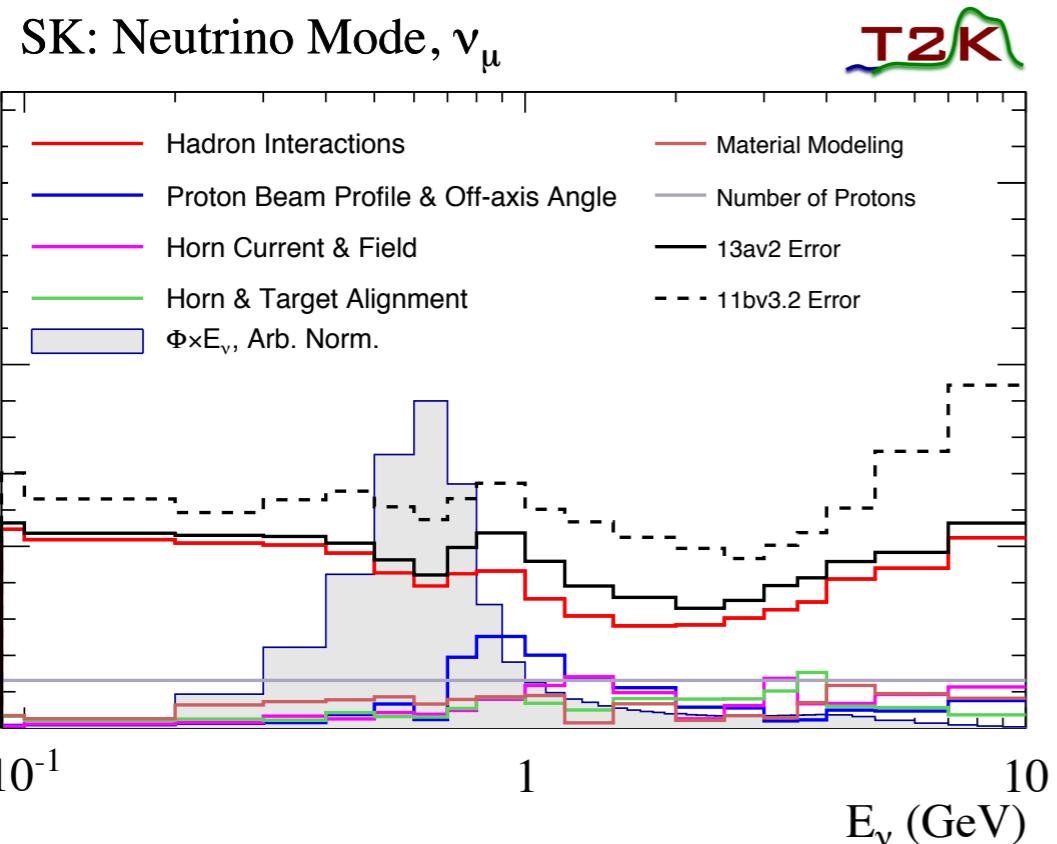
Almost identical near far detectors

Calorimetric Energy Reconstruction

$$E_{\nu \text{ reco}} = E_{\text{elep}} + E_{\text{had}}$$



Flux Uncertainties



T2K ~ 8-12% (based on thin target tuning)

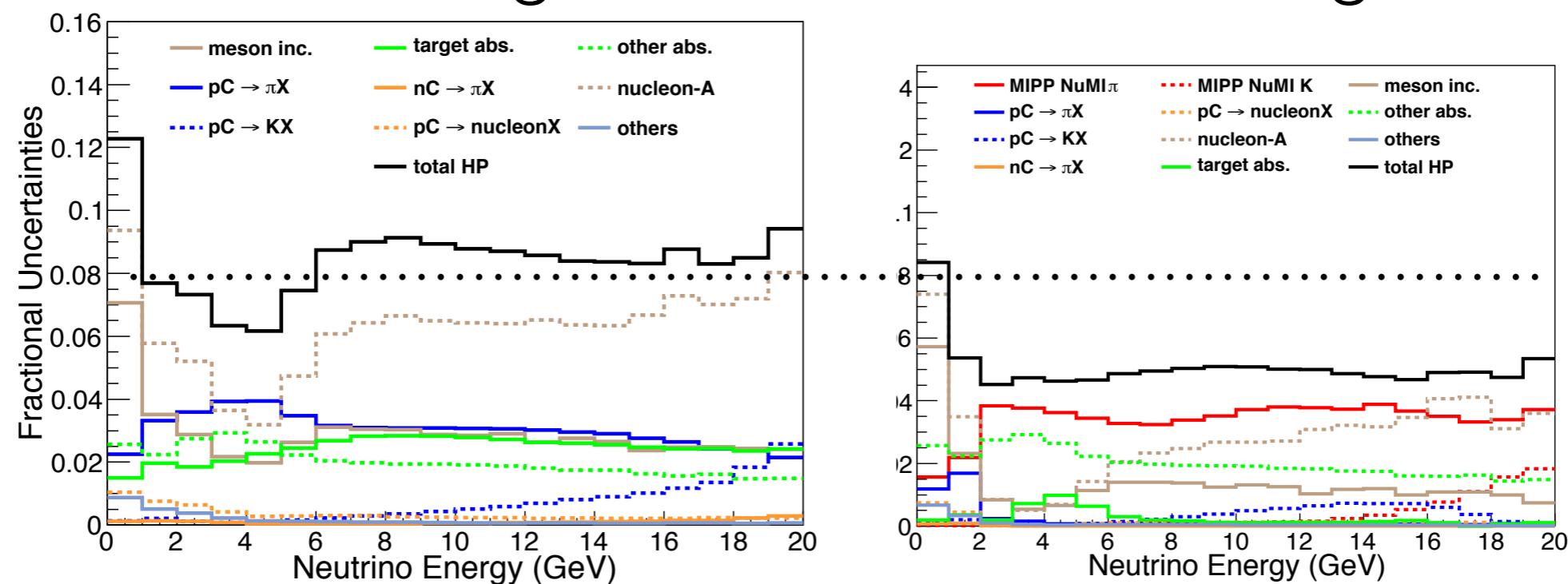
Dominated by hadron interaction modelling

Alignment/focussing uncertainties are also important
(especially for near to far extrapolation)

Flux Uncertainties

Thin Target

Thick Target



MINERvA Low E NuMI Flux Uncertainties, Phys. Rev. D 95, 039903 (2017)

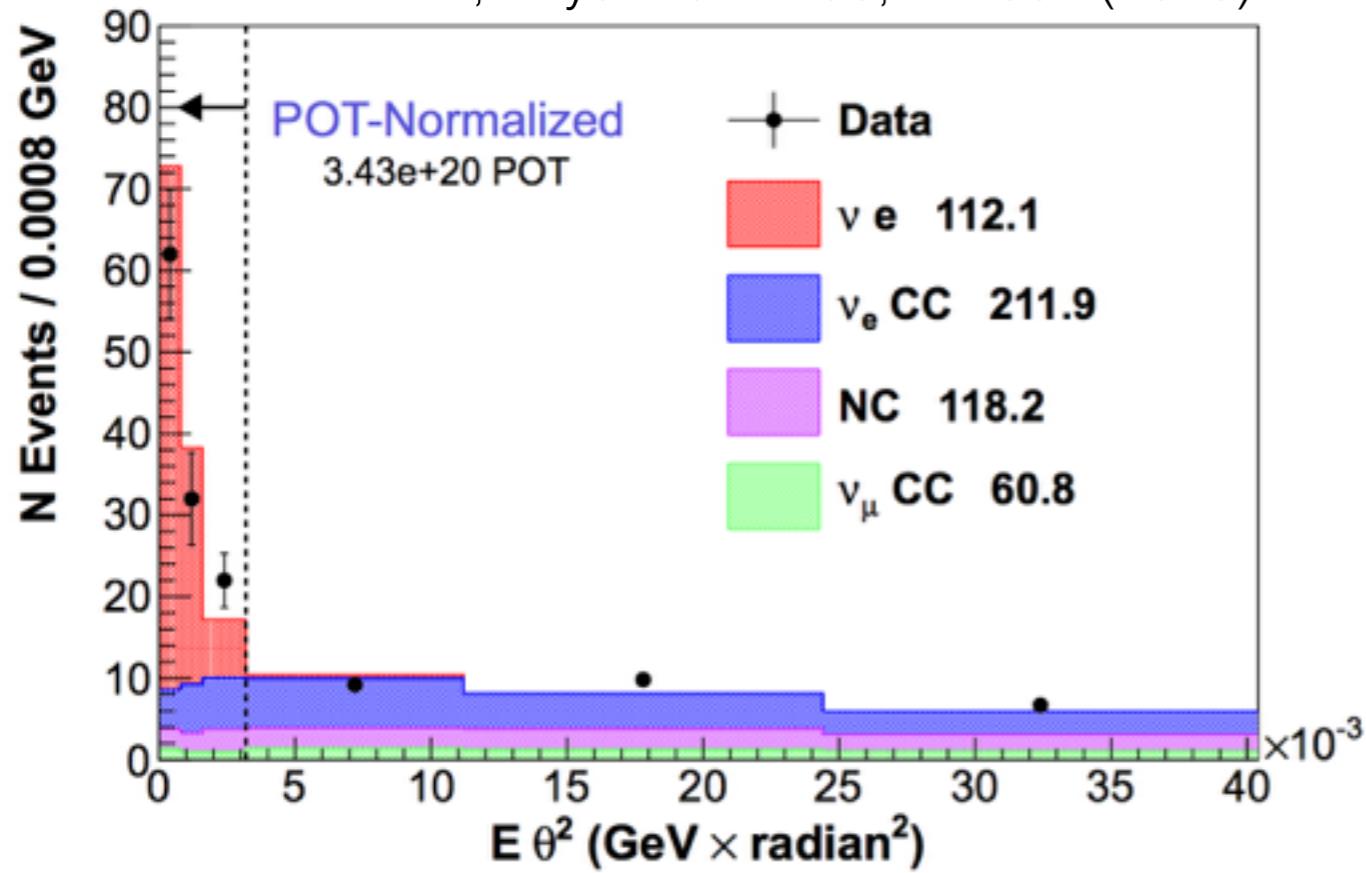
Significant reductions from thick/replica target
 (See Tomislav Vladisavljevic poster for latest T2K tuning)

Future high beam power experiments may have different target material/geometry requiring dedicated hadron production measurements

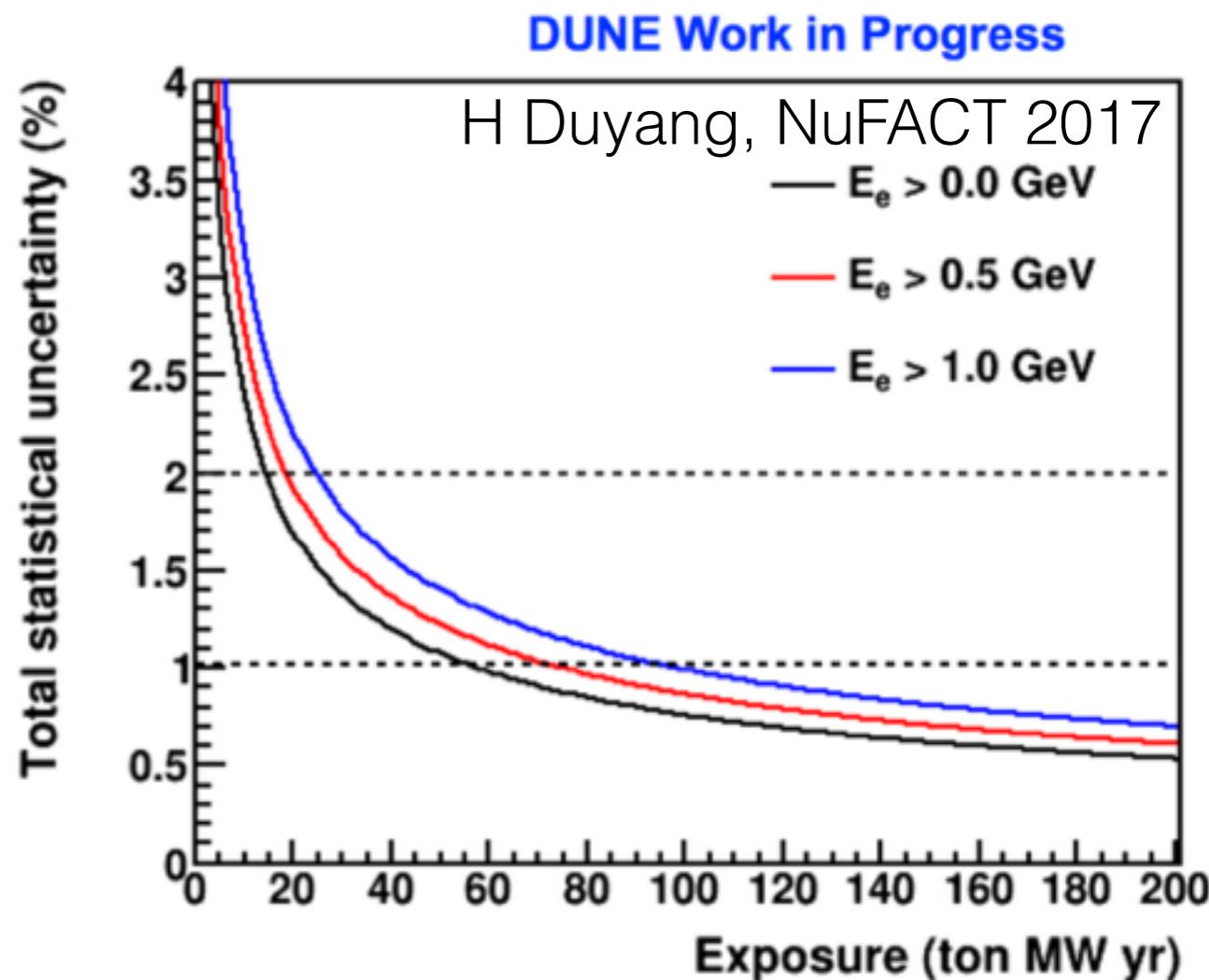
Flux Uncertainties

In situ flux measurements possible
eg neutrino-electron elastic scattering

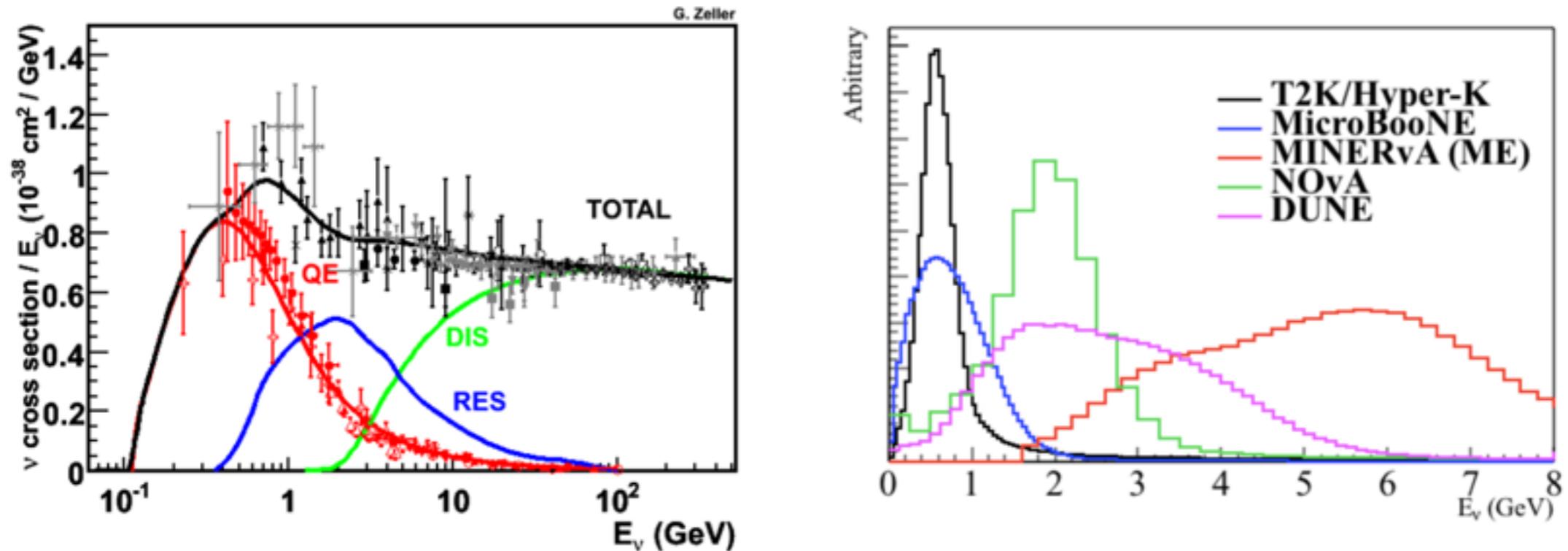
MINERvA, Phys. Rev. D 93, 112007 (2016)



DUNE Work in Progress



Neutrino Interaction Model Uncertainties



Wide range of processes need to be simulated
Require both lepton and hadronic side of the interaction
Nuclear effects important in the relevant energy regime
Experiments rely on MC generators
for $E_{\text{visible}} \rightarrow E_\nu$ extrapolation

Model parameter uncertainties from fits to external datasets

Sometimes parameter error must be inflated or ad-hoc parameters to account
for discrepancies between model and data or known flaws in the model



T2K Cross-Section Model

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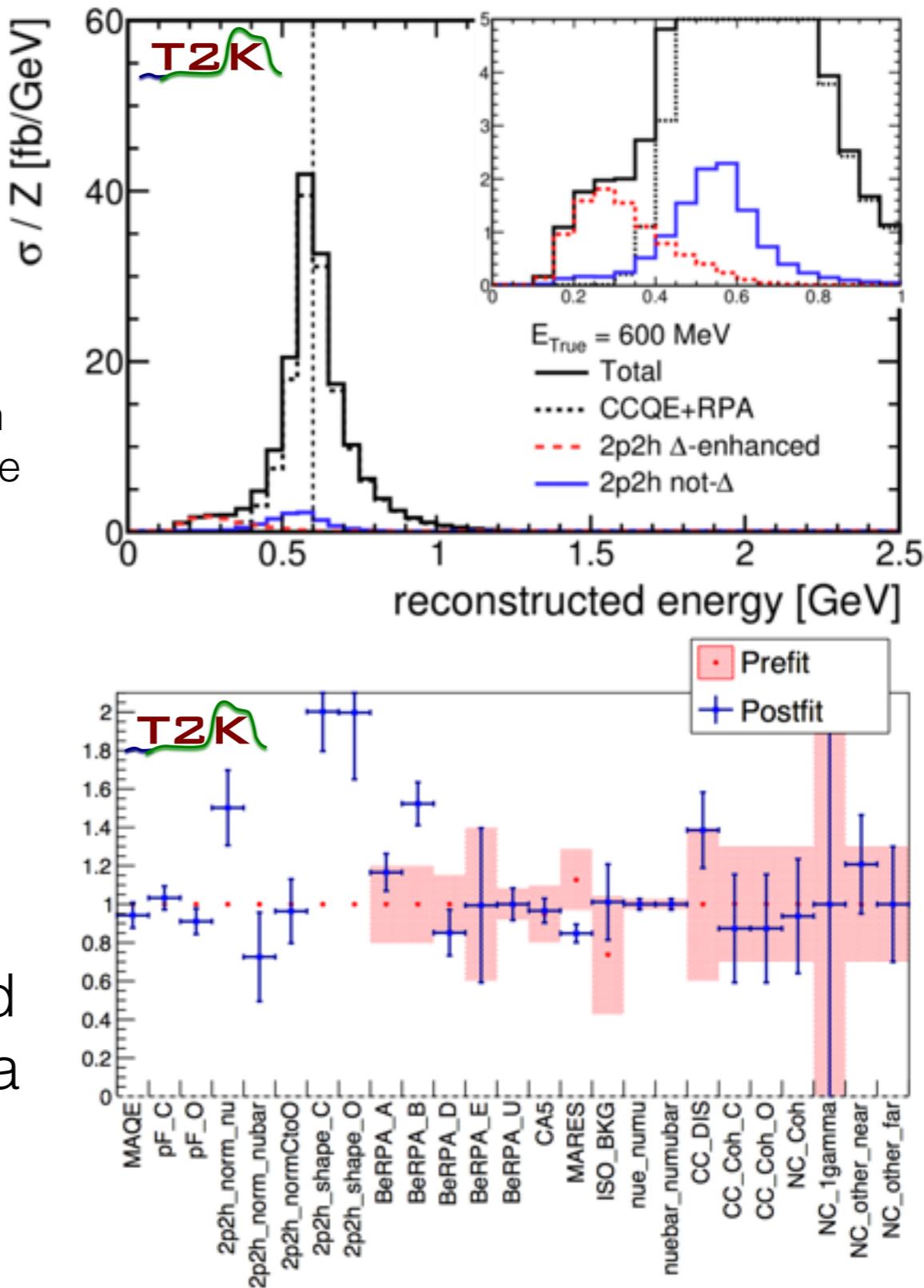
Implemented in NEUT MC generator

Quasi-elastic scattering most important process at T2K energies

- Valencia 2p-2h model Phys. Rev. C83 (2011) 045501
- Long-range effects with Random Phase Approximation
- Parameters introduced to vary normalisation and shape
- Relativistic Fermi Gas (RFG) nuclear model
- Uncertainties from RFG \leftrightarrow Local Fermi Gas
- Final state interactions with cascade model

No priors on most CCQE parameters
Constraint from near detector

Impact of alternative models not implemented in oscillation analysis evaluated with fake data studies





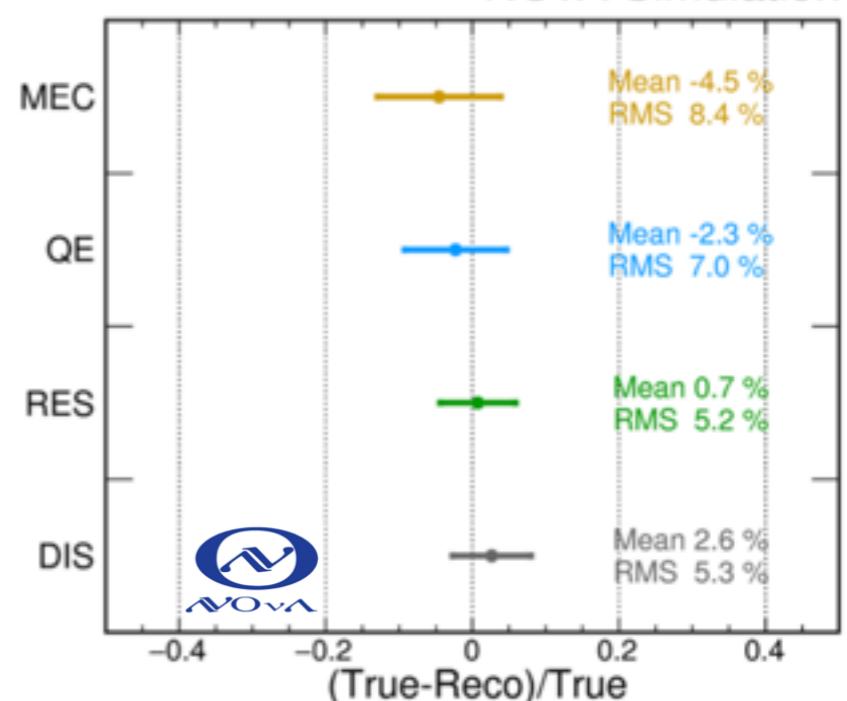
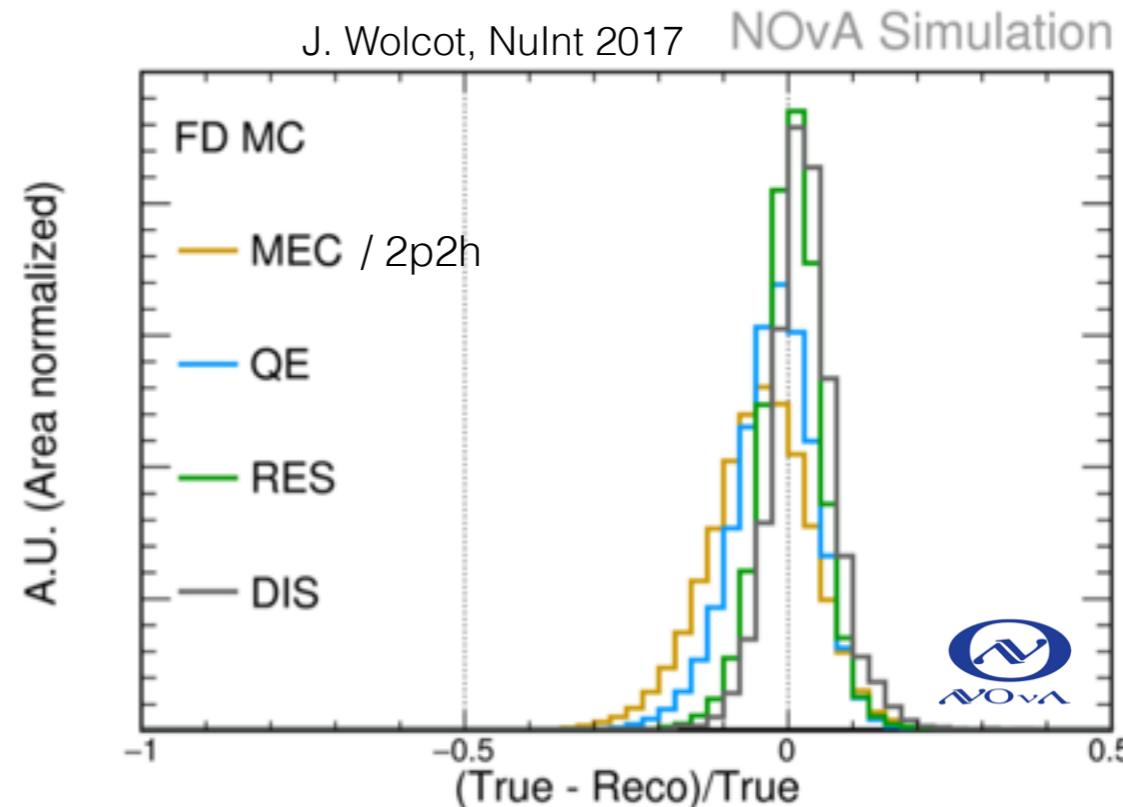
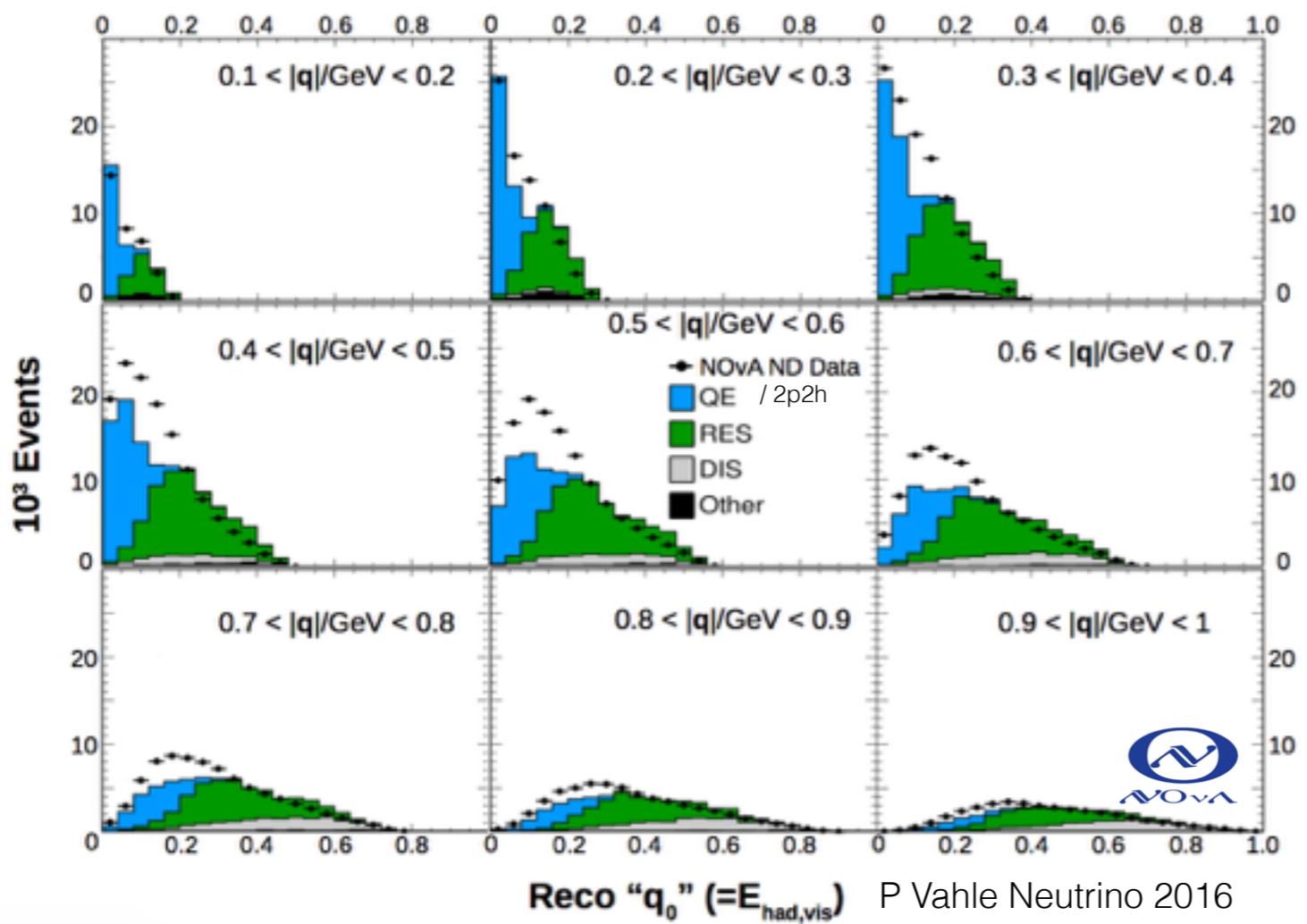
NOvA Cross-section Model

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Use GENIE MC generator and uncertainties

Some additions/modifications

- Empirical 2p2h model, tuned to match ND data
- Parameters to cover RPA uncertainties
- Alternative tuning of CC1 π model [Eur. Phys. J. C 76, 474 (2016)]





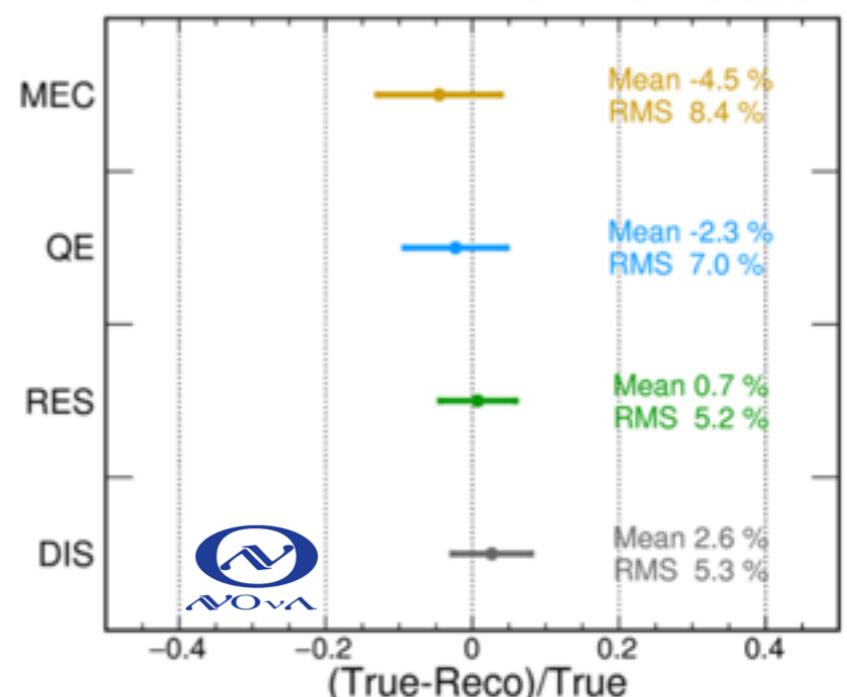
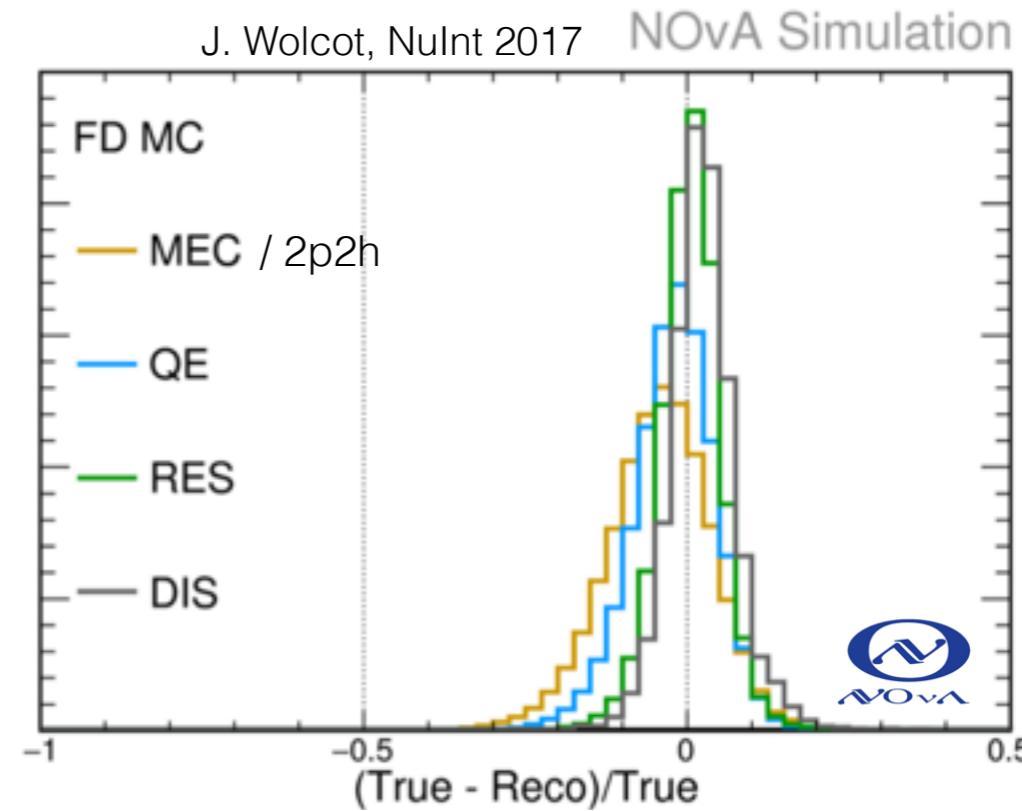
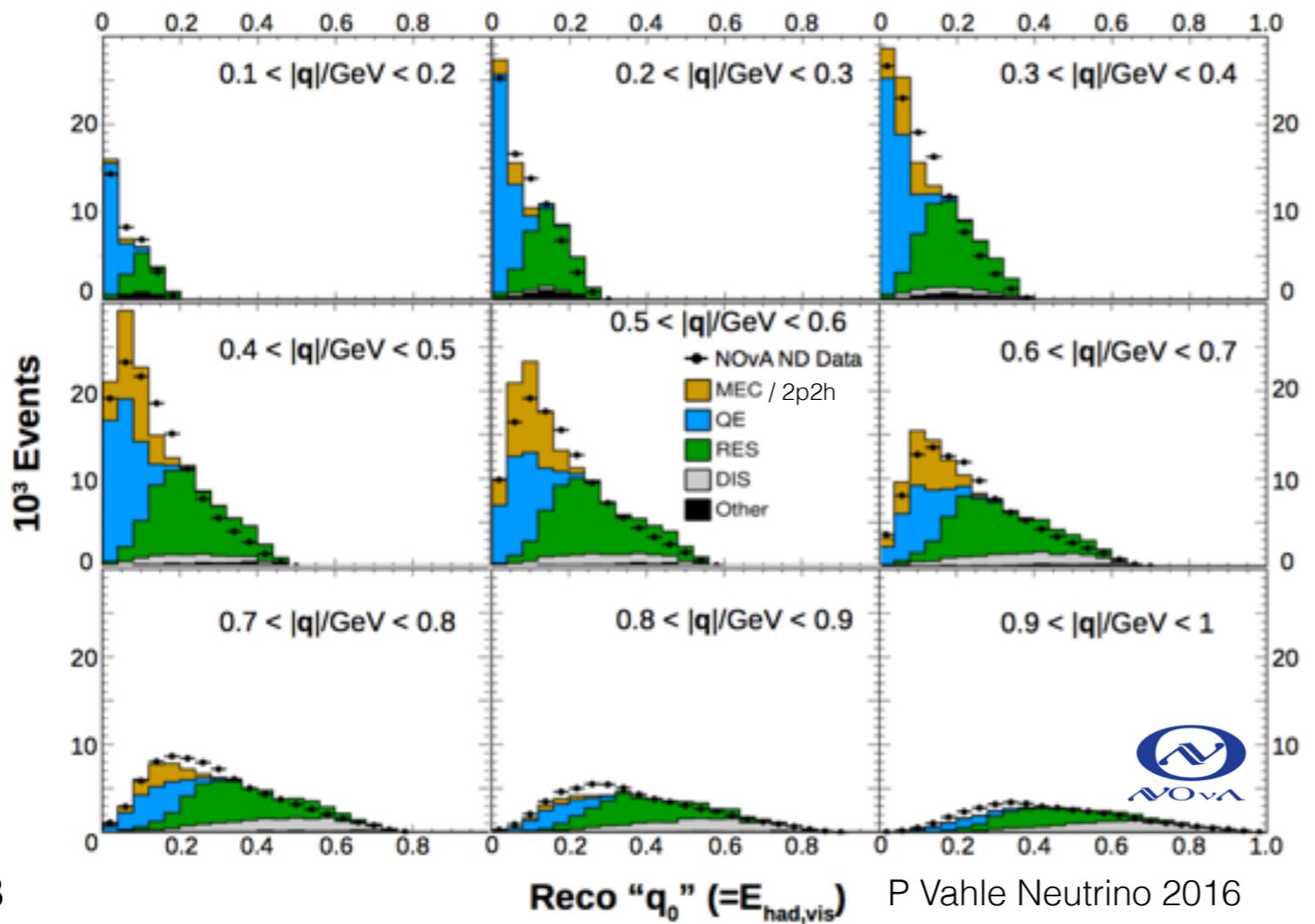
NOvA Cross-section Model

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Use GENIE MC generator and uncertainties

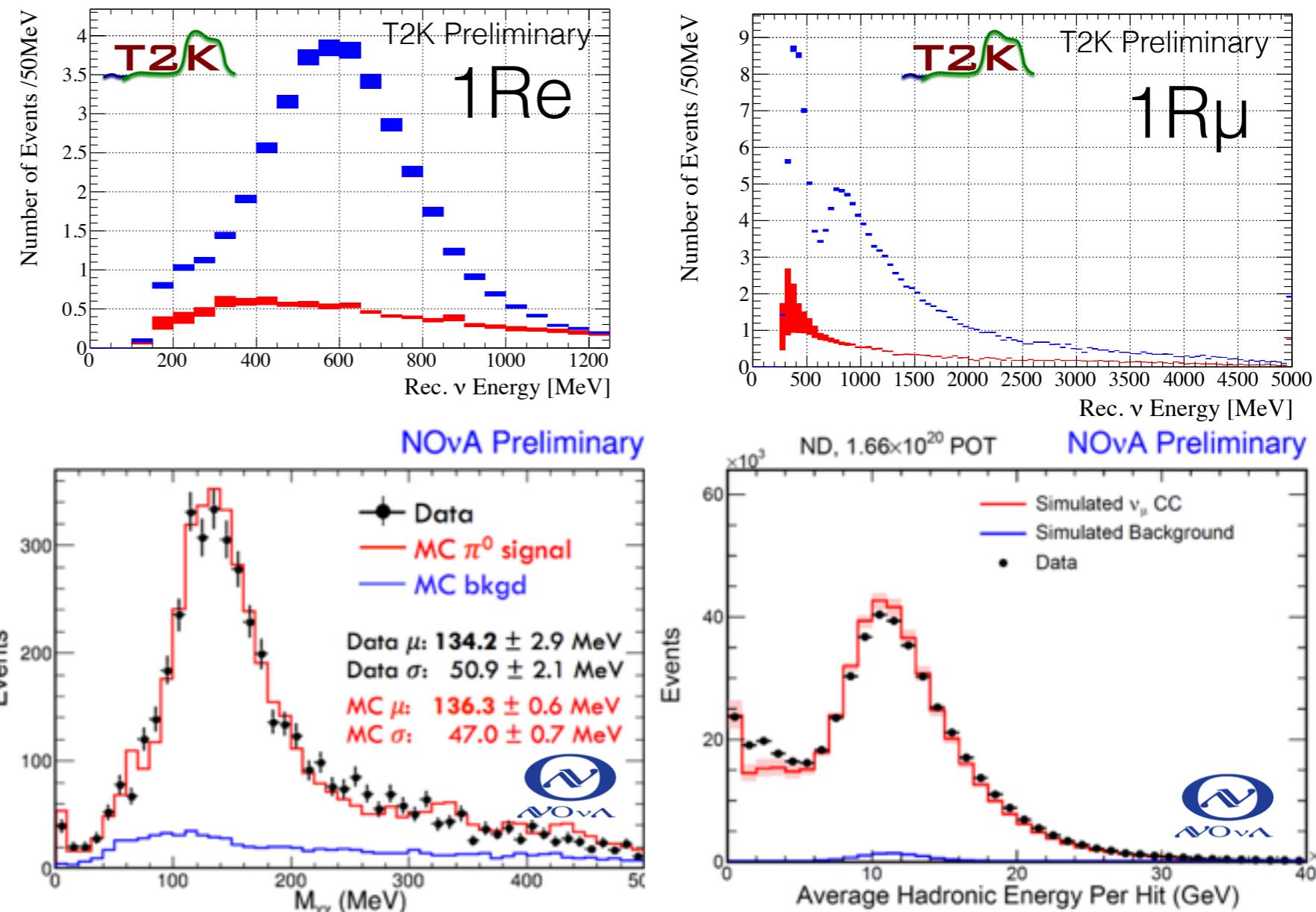
Some additions/modifications

- Empirical 2p2h model, tuned to match ND data
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Detector Modelling Uncertainties

SK detector response evaluated with atmospheric sample



Detector modelling uncertainties typically from data MC comparisons in control samples
 May be limited by control sample statistics

T2K Systematic Uncertainties

ND280 constraint
13% → 3%

Error Source	μ sample [%]		e sample [%]	
	ν	$\bar{\nu}$	ν	$\bar{\nu}$
SK Detector	1.9	1.6	3.0	4.2
SK FSI+SI+PN	2.2	2.0	2.9	2.5
ND280 Constraint (Flux + Cross Section)	3.3	2.7	3.2	2.9
$\sigma(\nu_e)/\sigma(\nu_\mu)$	-	-	2.6	1.5
NC 1γ	-	-	1.1	2.6
NC other	0.3	0.1	0.1	0.3
Total Systematic	4.4	3.8	6.3	6.4
Statistical	6.5	12	12	40

T2K preliminary (final systematics pending)

Total systematic uncertainty
~4 - 6%

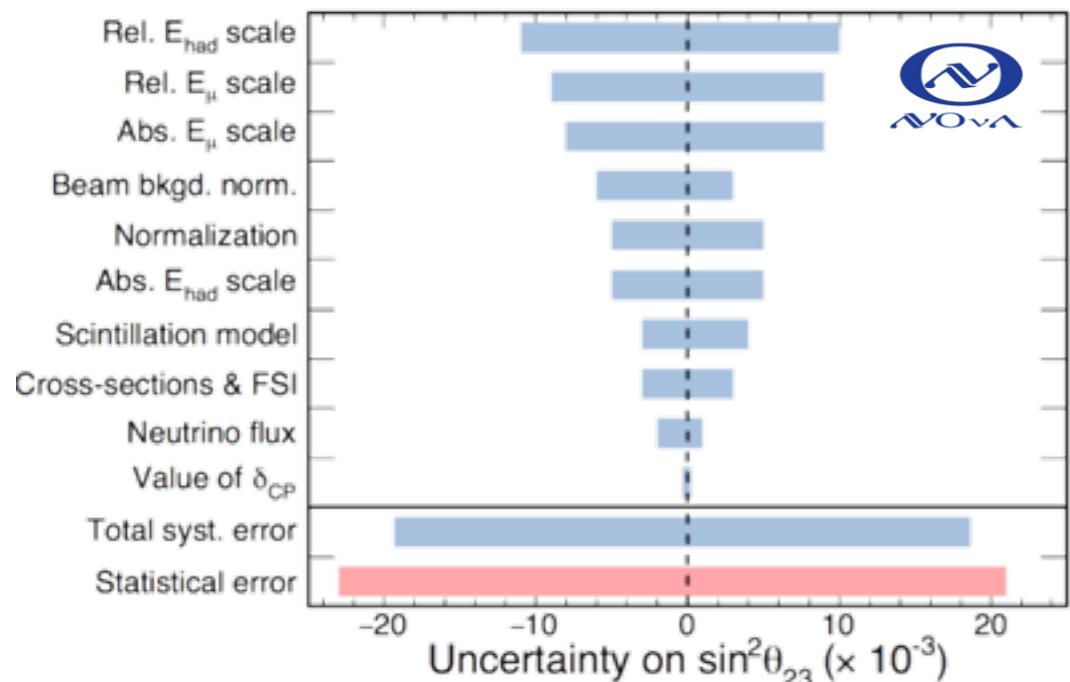
Smaller than stats. uncertainty
(for now!)

Pion Final State
Interactions (FSI) and
Secondary Interactions
(SI) modelling important

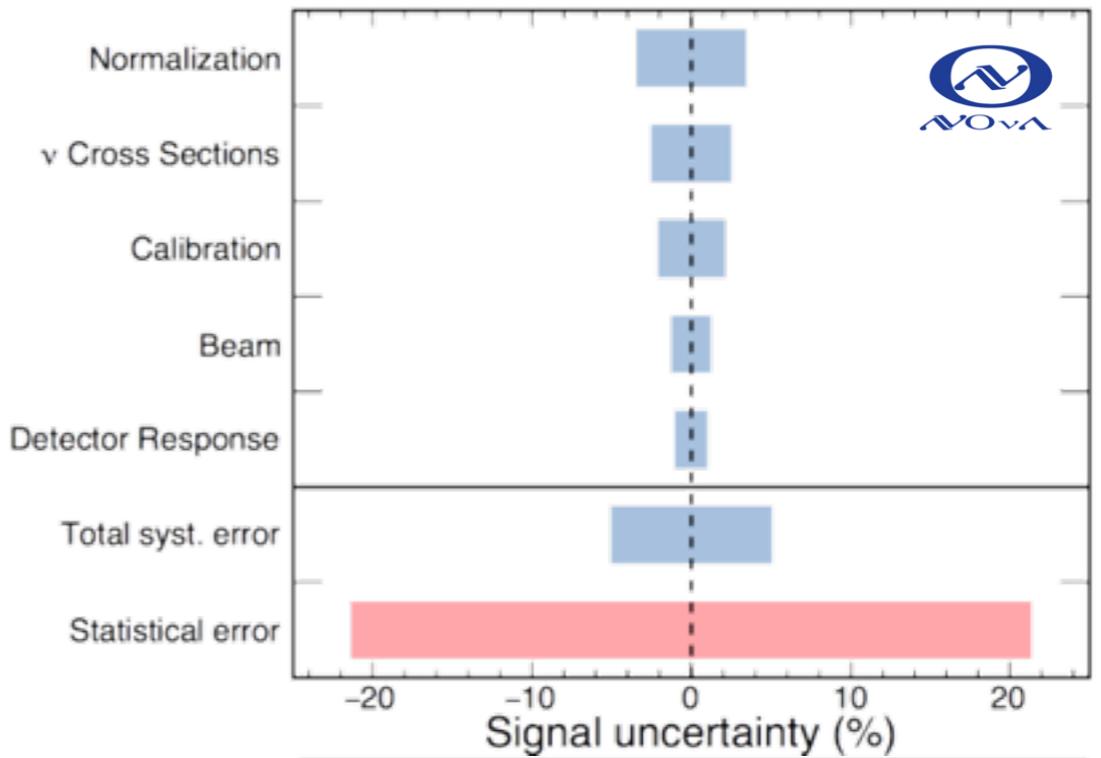
Theoretical uncertainty
 ν_e to ν_μ
Difficult to constrain with
near detector

NOvA Systematic Uncertainties

v_μ



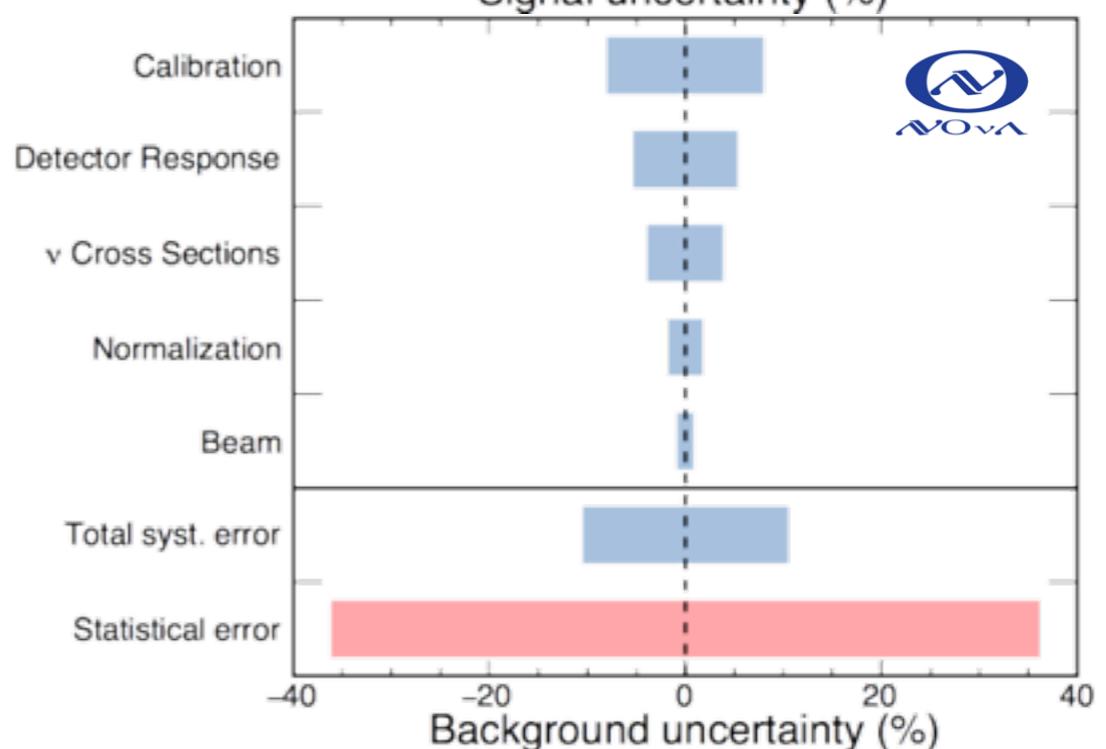
v_e



$v_e \sim 5\text{-}10\%$

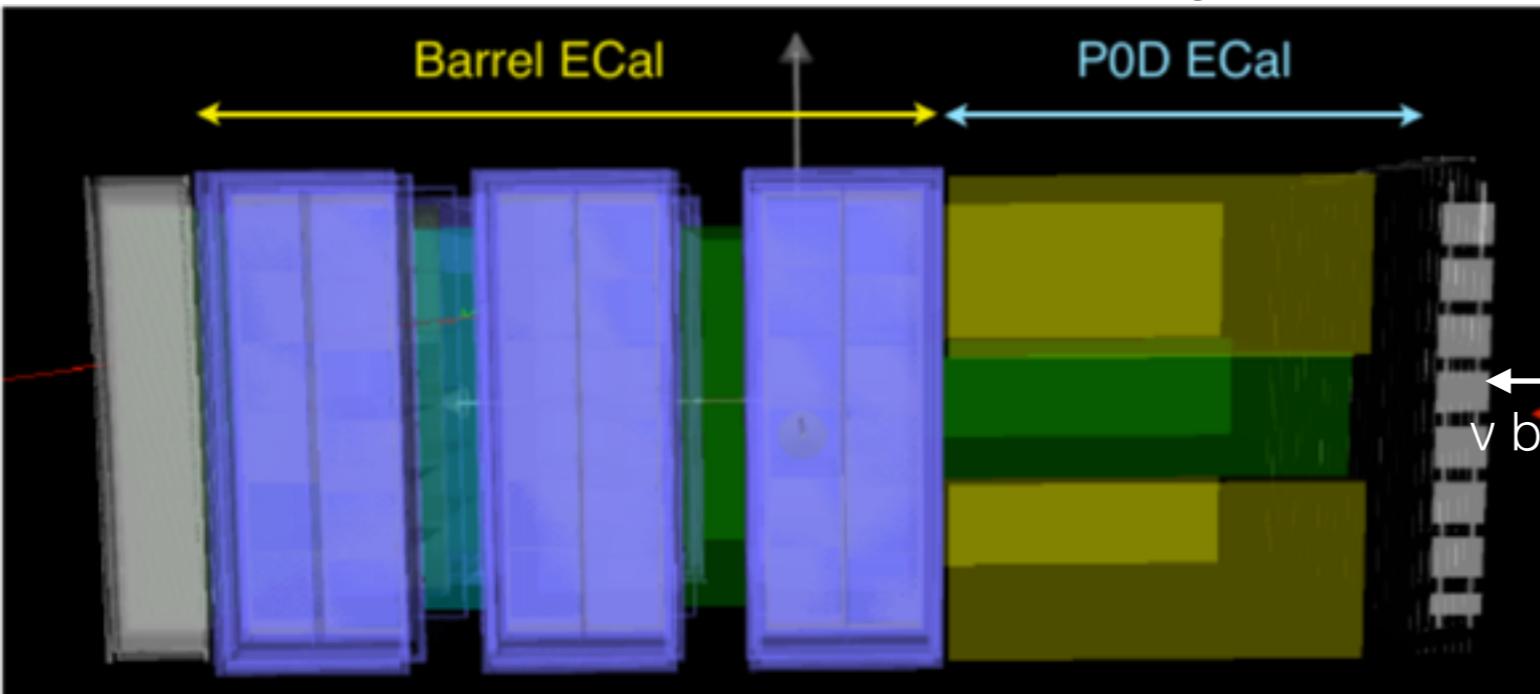
$v_\mu \sim 3\text{-}4\%$

Energy scale
uncertainties dominate v_μ

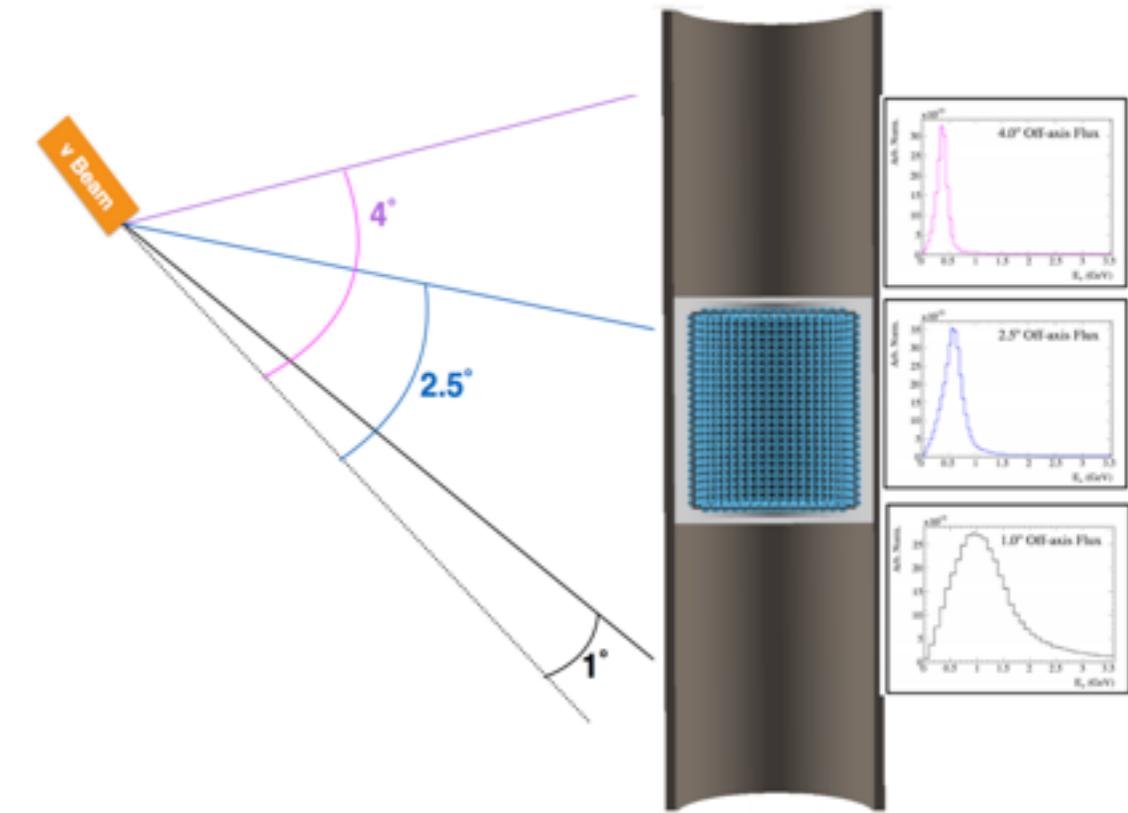


Near Detector Development

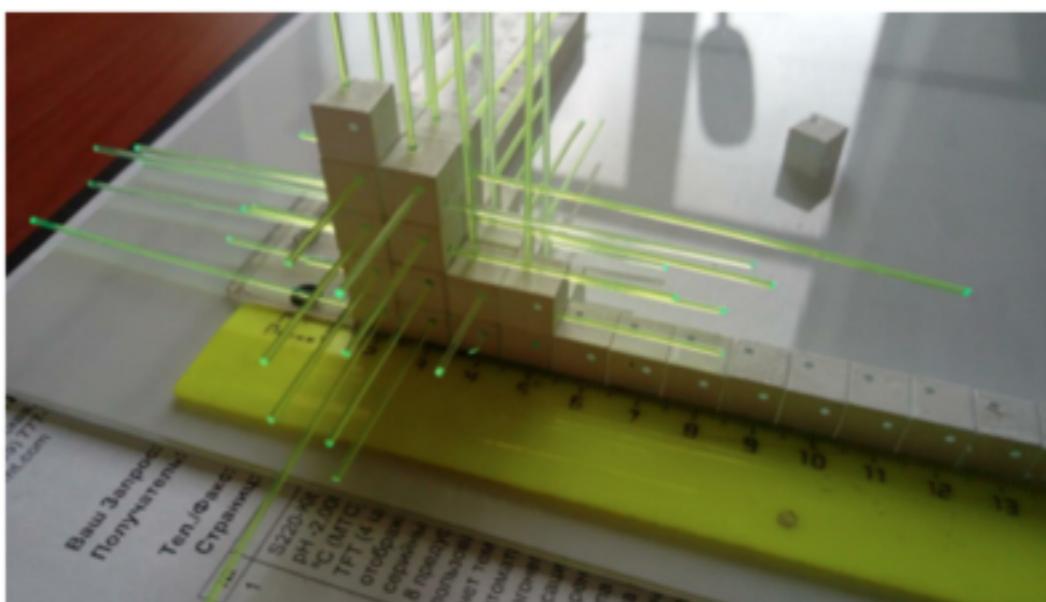
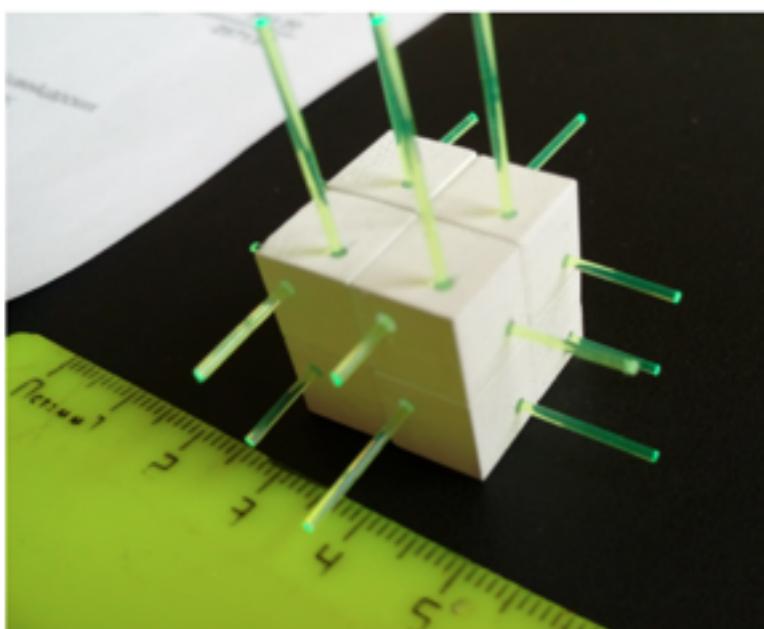
Planned ND280 Near Detector Upgrade



E61 Experiment

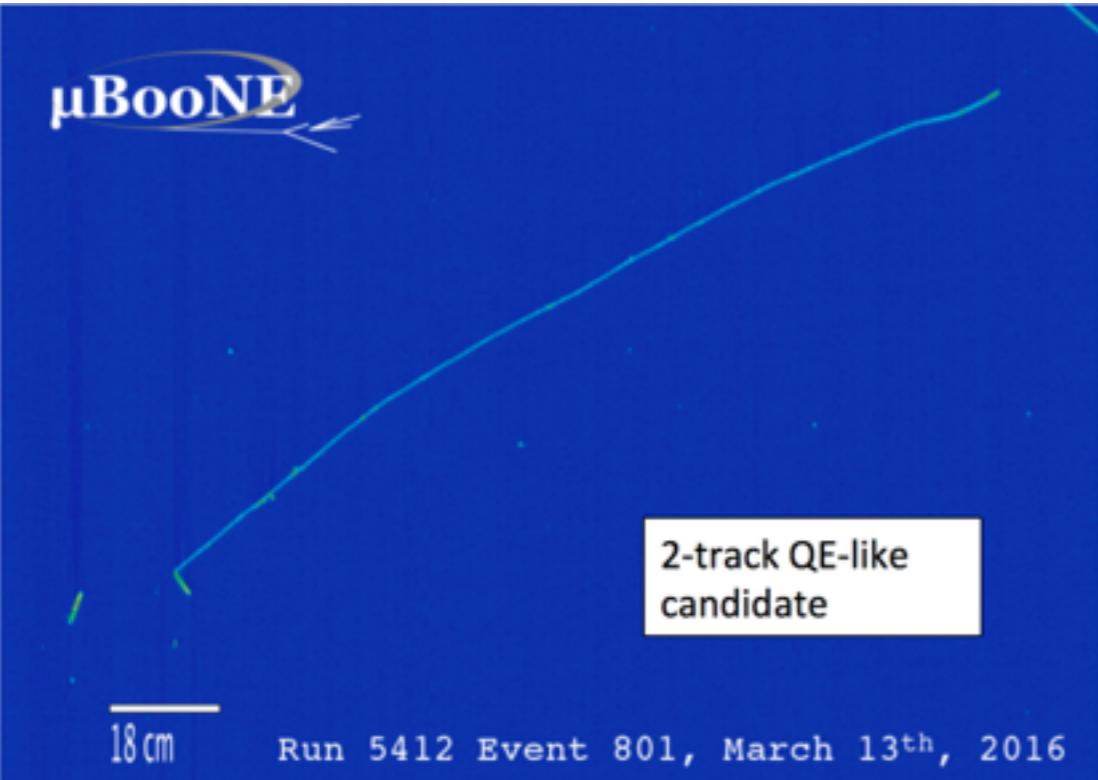


Near detector upgrades for T2K-II and T2HK era
New target with increased angular acceptance



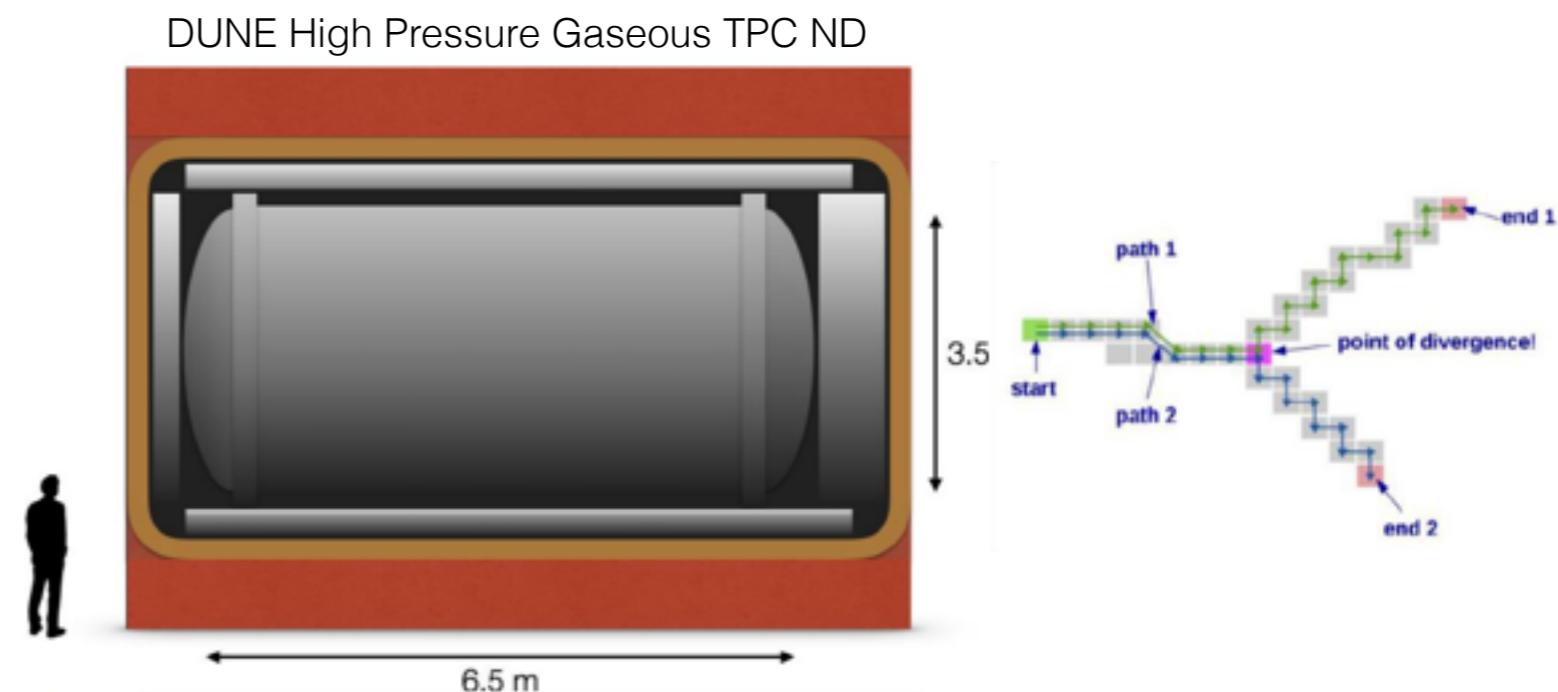
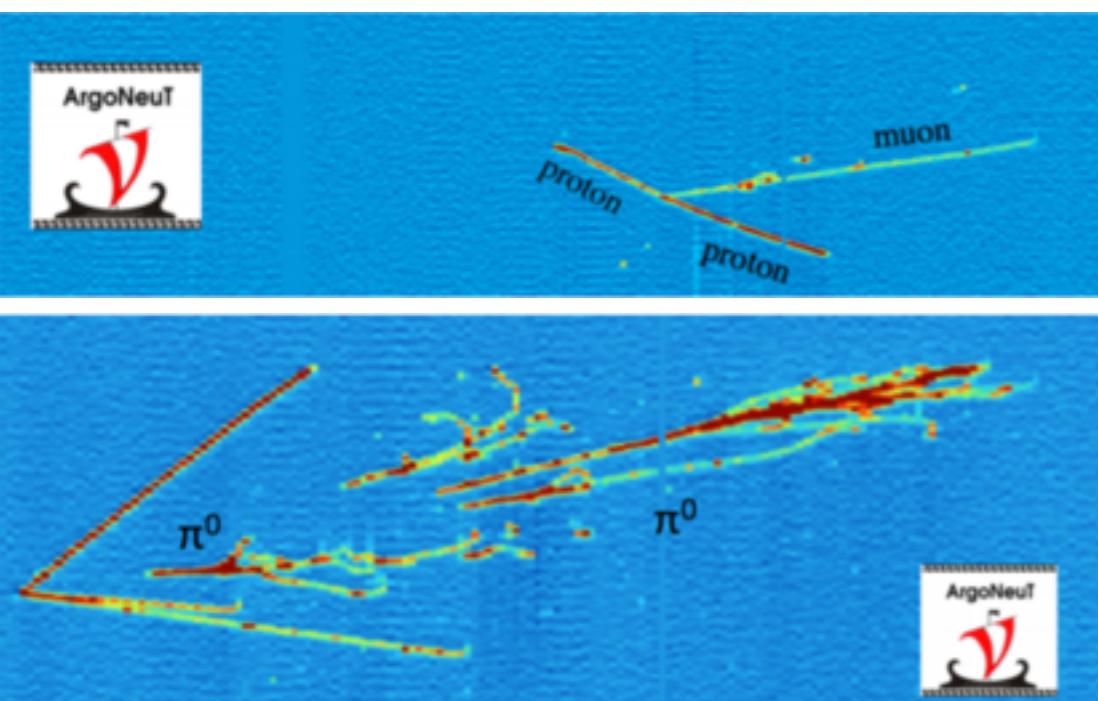
Intermediate Water-Cherenkov detector
Map detector response using multiple off-axis angles

Near Detector Development



Several Argon TPC experiments
Natural ND candidates for DUNE

Precisely image the neutrino interaction vertex
(better constraints on neutrino-nucleus
interaction models → better energy
measurement)



Ultra-low thresholds with gaseous TPC

Summary



Statistical precision promised by future high beam power and high mass experiments place high demands on the systematic uncertainties that experiments must reach

T2K and NOvA have reported systematics uncertainties in the range $\sim 3 - 10\%$ level

Reductions are needed today to make best use of the increasing statistical precision in the T2K and NOvA disappearance measurements

Improved flux determination, ν -nucleus interaction modelling and understanding of detector response will all play a role



Systematic Uncertainties in Long Baseline Neutrino Oscillation Measurements

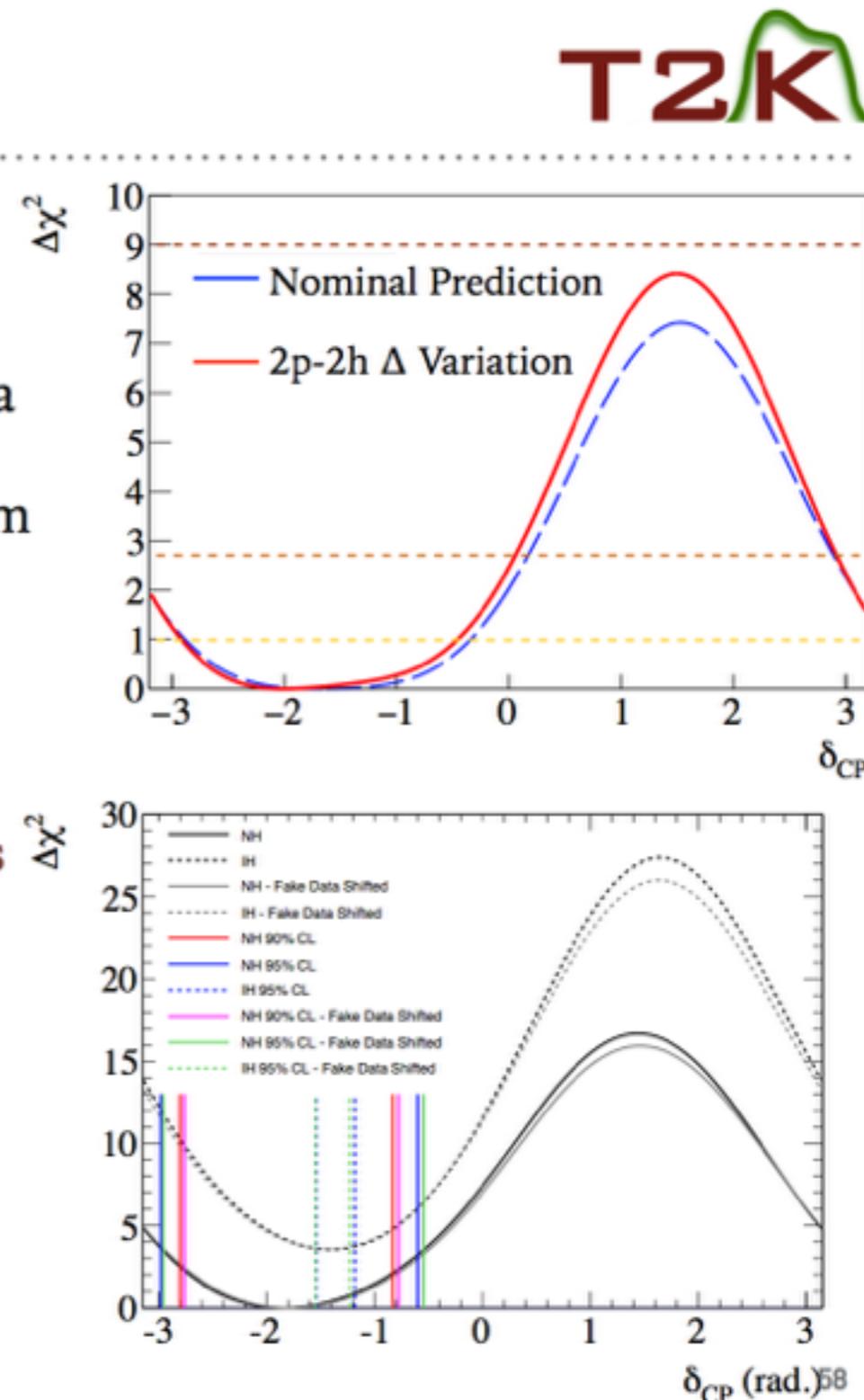


David Hadley
21st December 2017
Prospects in Neutrino Physics, NuPhys2017

Fake Data Studies

IMPACT ON CP PHASE

- ▶ Consider how changes to the $\Delta\chi^2$ impact intervals calculated from data
 - ▶ Shift $\Delta\chi^2$ observed in data (bottom plot) by difference observed in systematic study (top plot)
- ▶ Maximum shift in the NH 2σ confidence interval mid-point was 1.7%
- ▶ Maximum change to the NH 2σ confidence interval was 2.3%
- ▶ Impact on δ_{cp} intervals is small

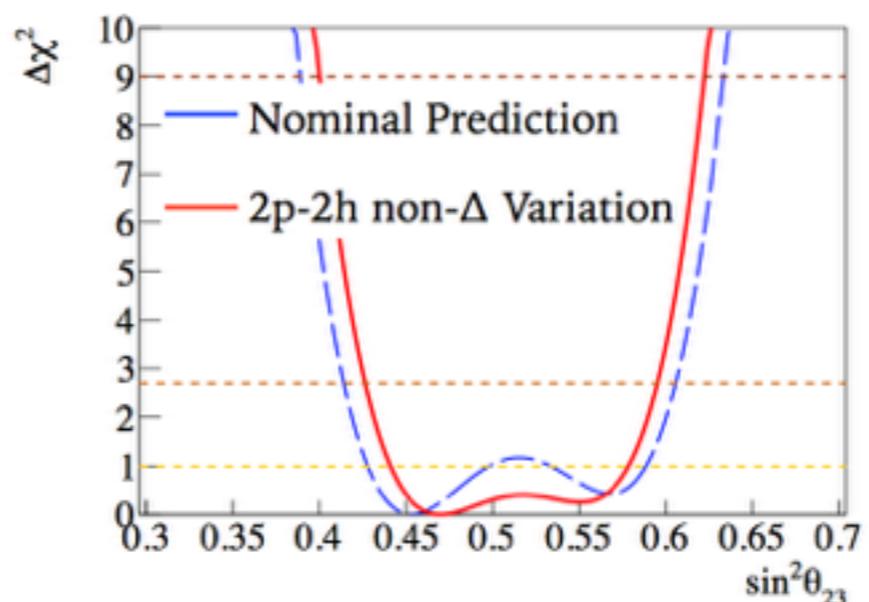
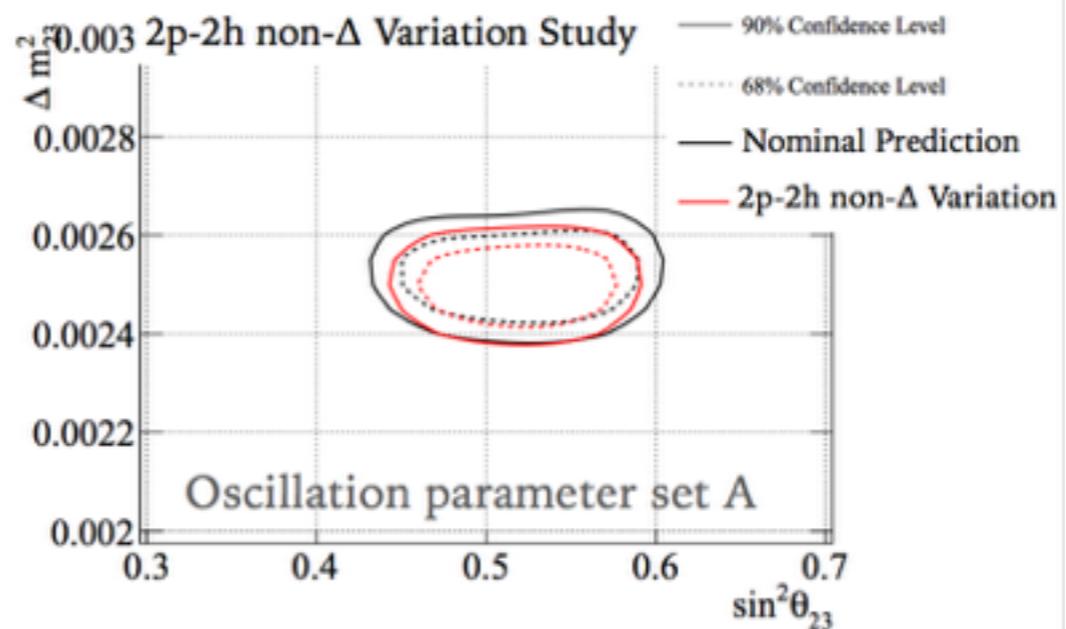


Fake Data Studies

IMPACT ON ATMOSPHERIC PARAMETERS



- In this study, Δm^2_{32} is biased to lower values
- $\sin^2\theta_{23}$ is biased towards maximal disappearance
 - Leads to narrower contour than fit to nominal prediction
- Shift towards maximal also seen in 1-D contour for oscillation parameter set B (bottom)



Fake Data Studies

ND280 DATA-DRIVEN VARIATION

- Take excess of data over prediction prior to ND280 fitting

- Assign excess to 1 of 3 types of interactions:

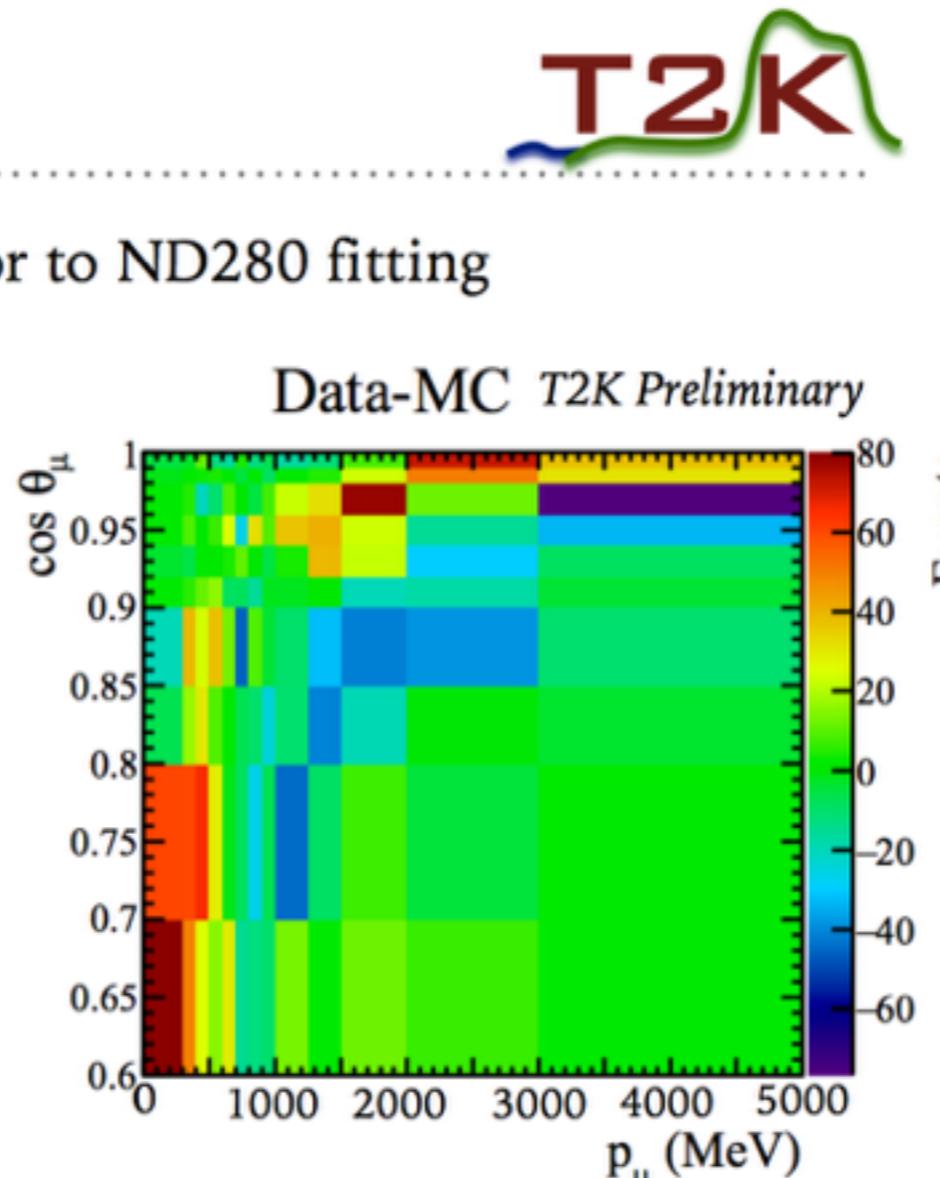
- CCQE
- 2p-2h Δ -enhanced
- 2p-2h non- Δ -enhanced

- Apply modeled excess to predict rates ND280 and SK

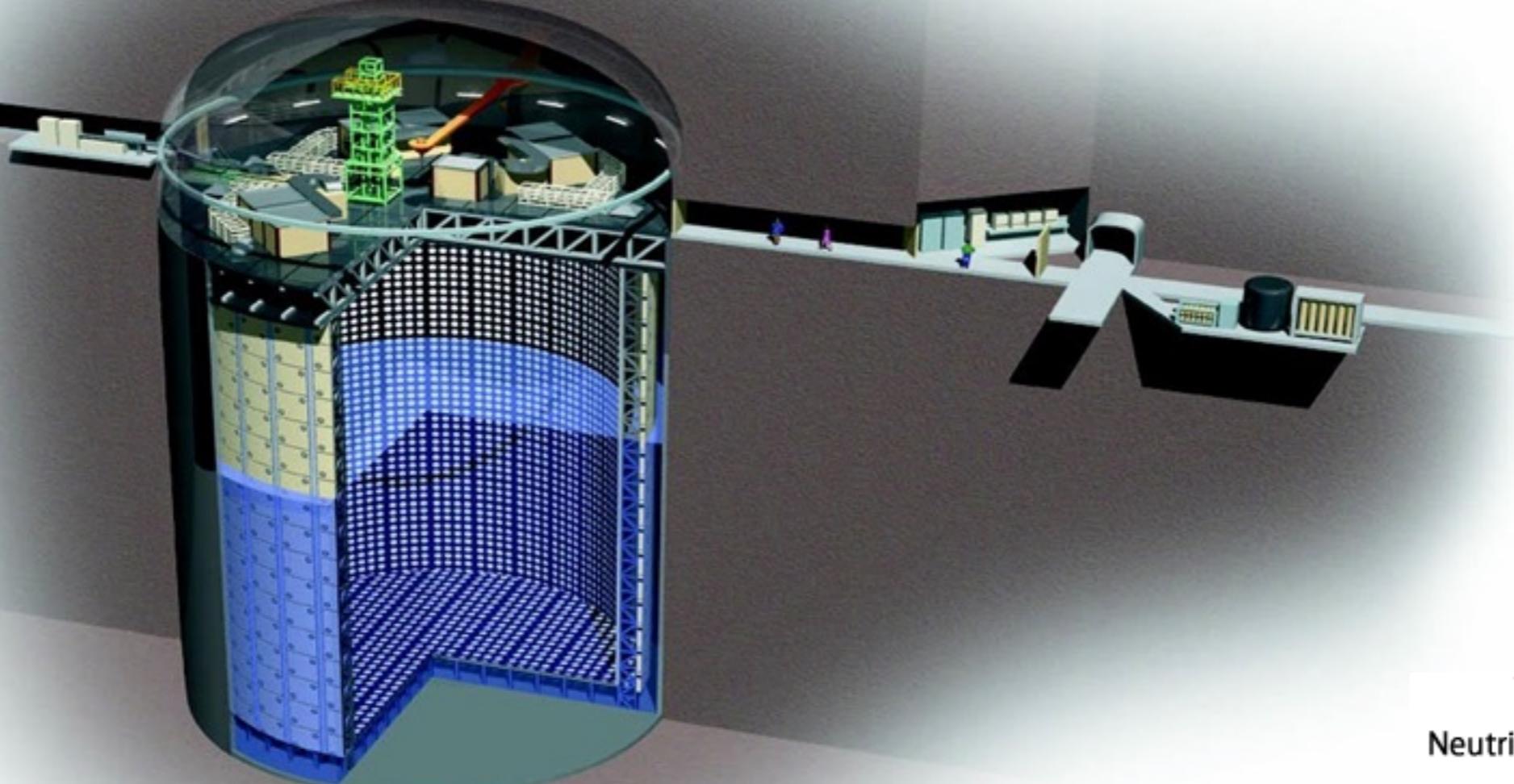
- Run fits

- Effect seen on $\sin^2\theta_{23}$ and Δm^2_{32}

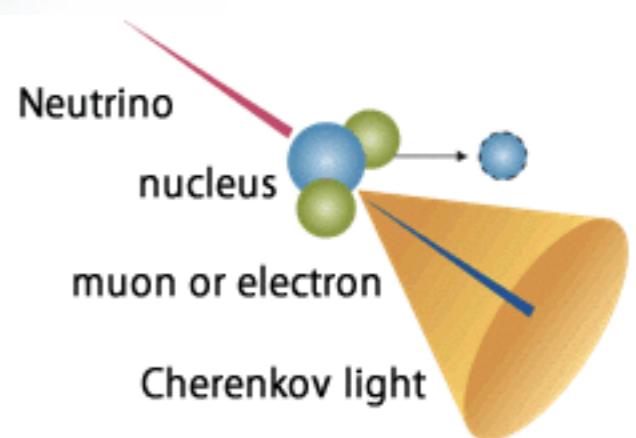
- No significant impact on the measured intervals for δ_{cp}



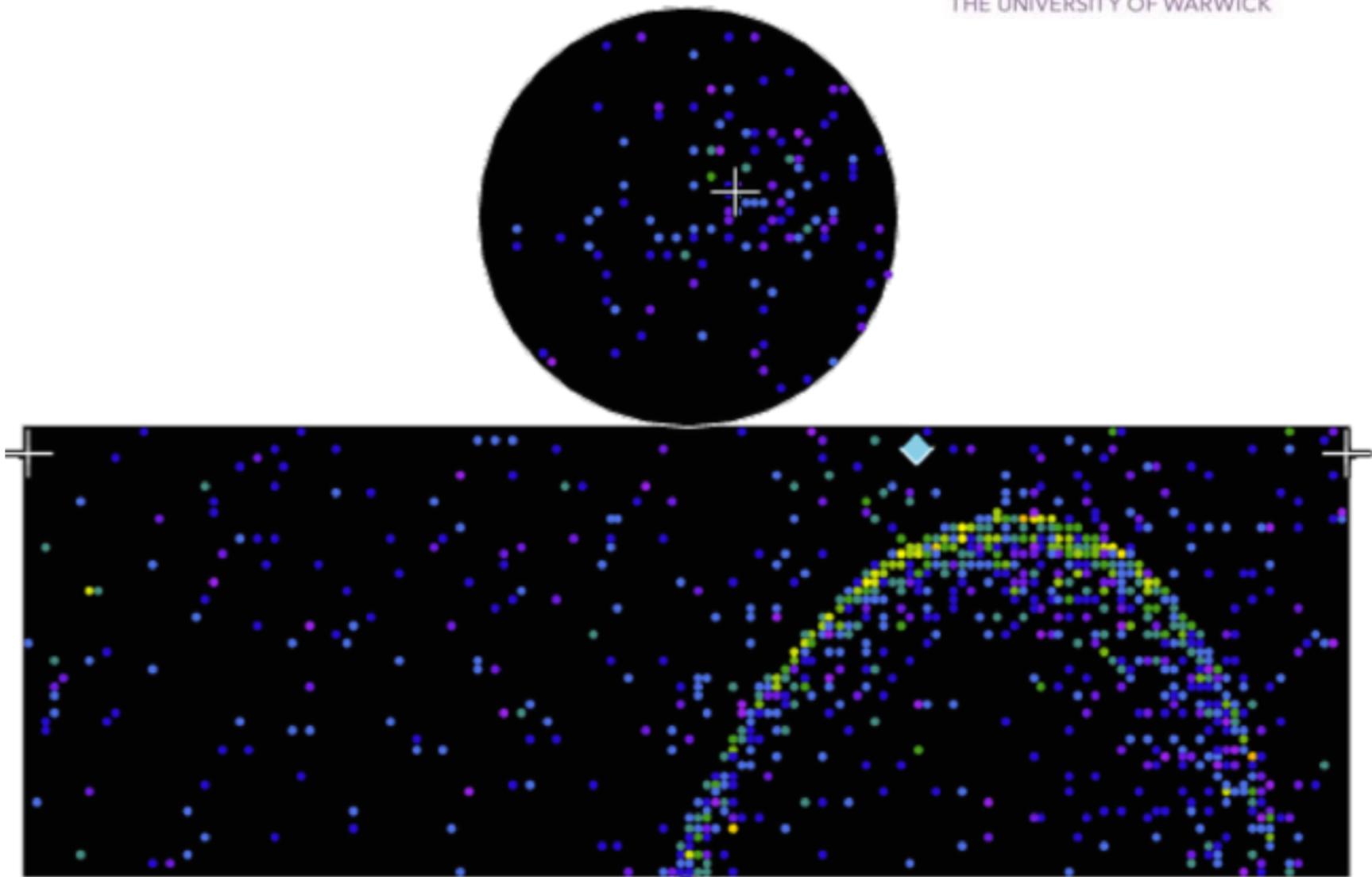
Super-Kamiokande



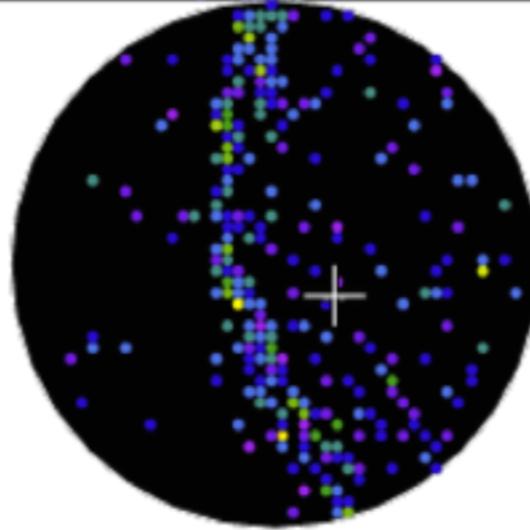
Water Cherenkov Detector
>22.5 kt fiducial mass



Water Cherenkov Technique

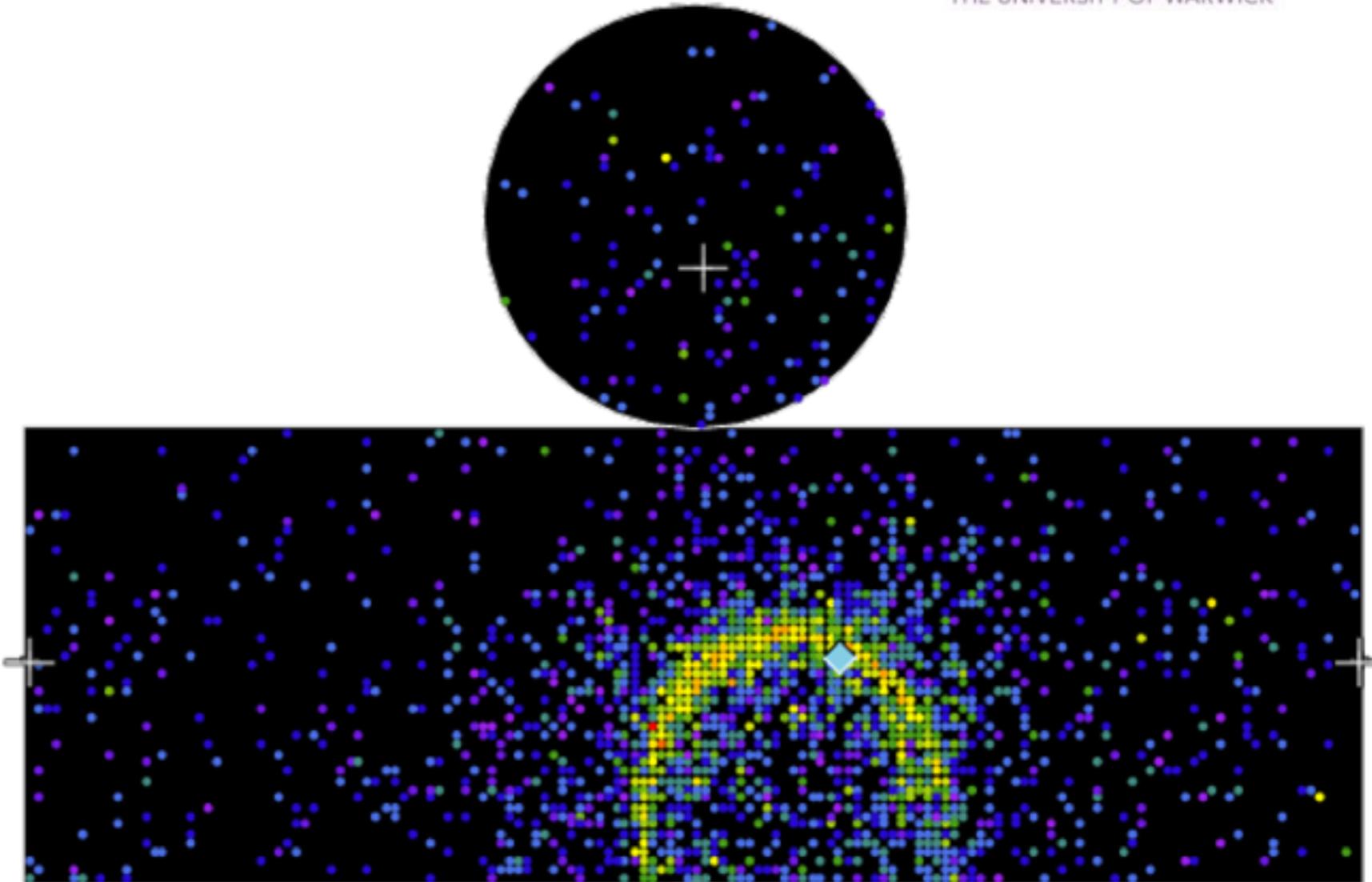
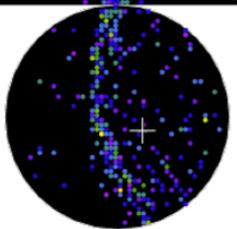
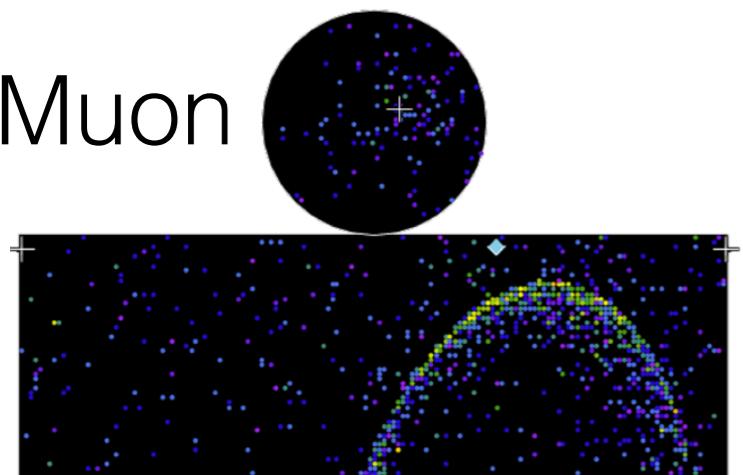


Muon

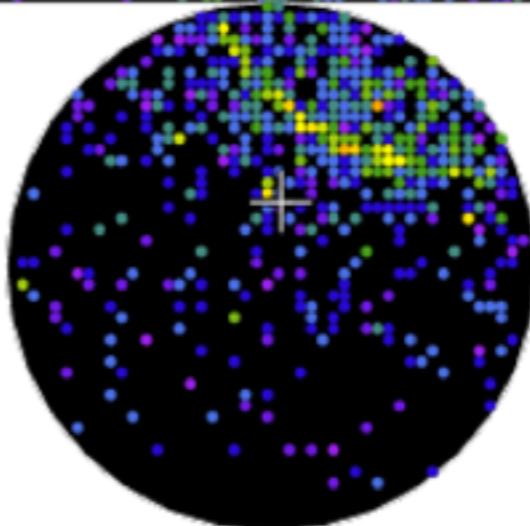


Water Cherenkov Technique

Muon



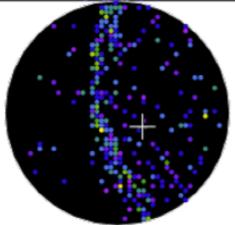
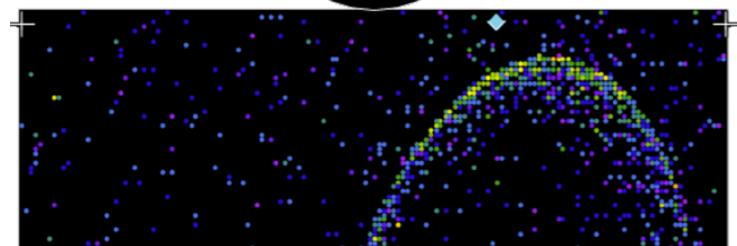
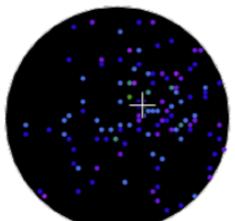
Electron



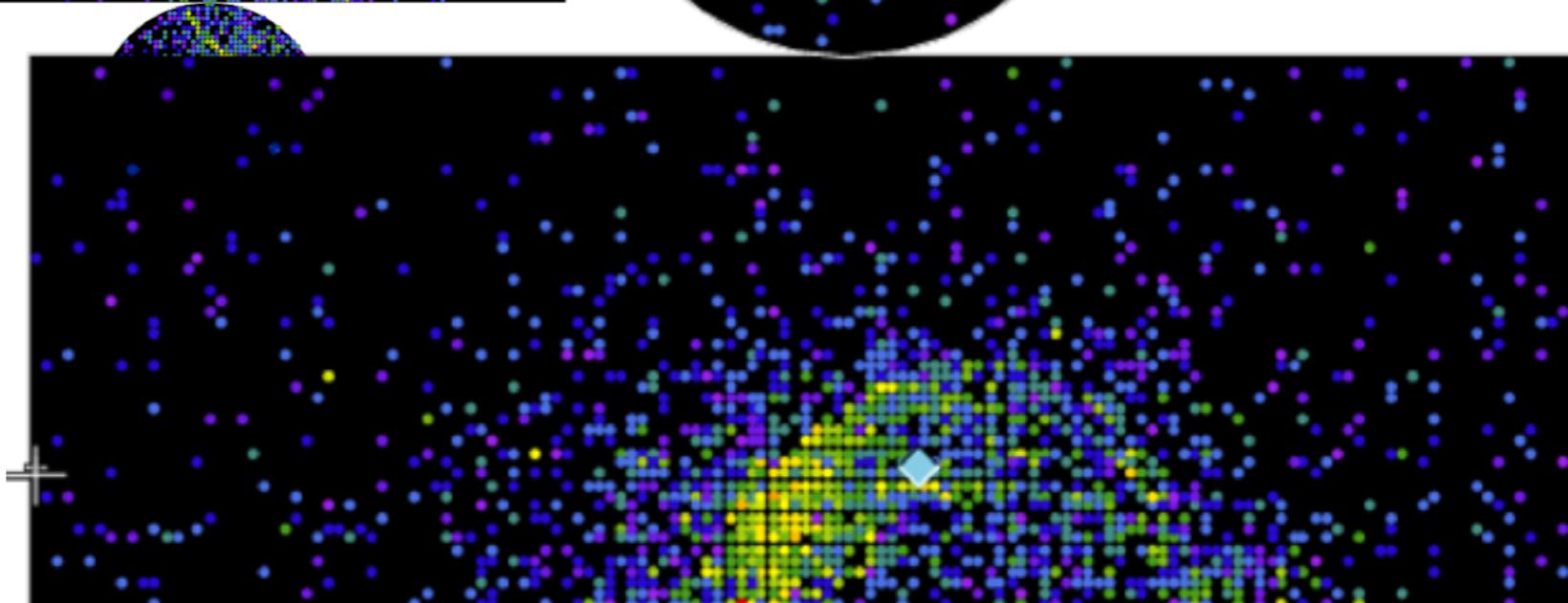
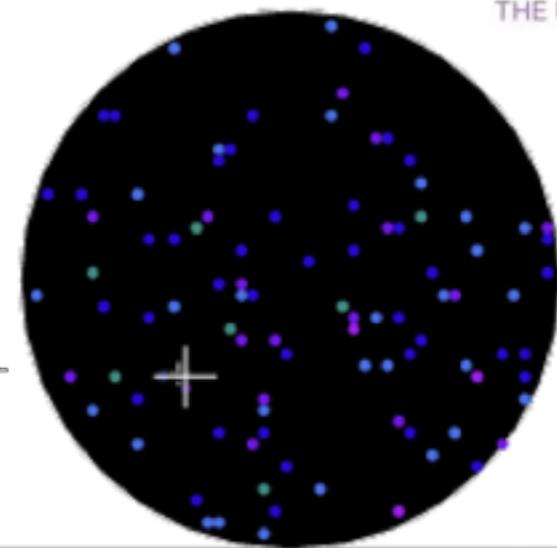
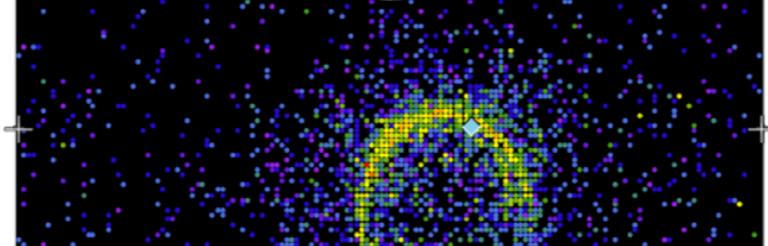
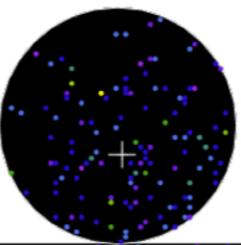
Water Cherenkov Technique



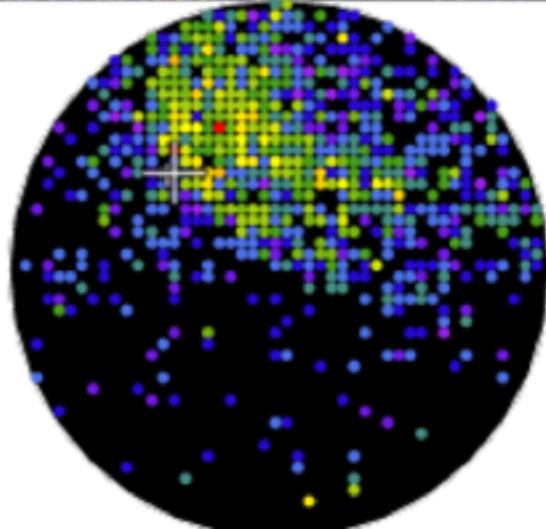
Muon



Electron

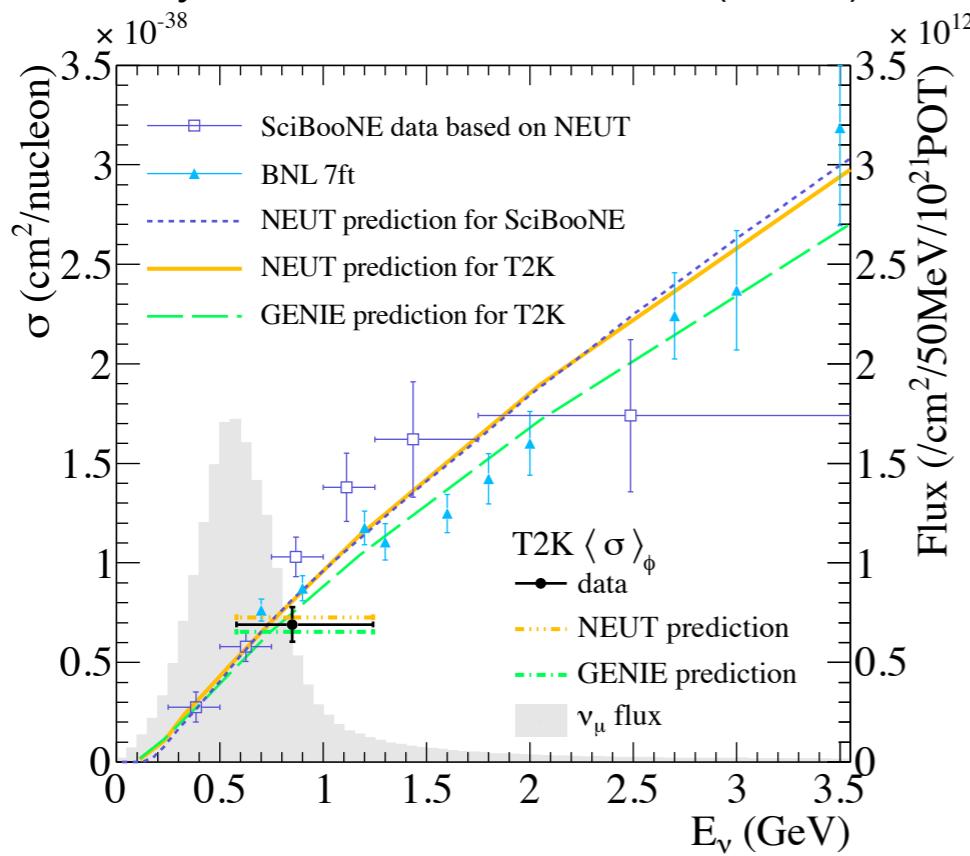


Neutral Pion

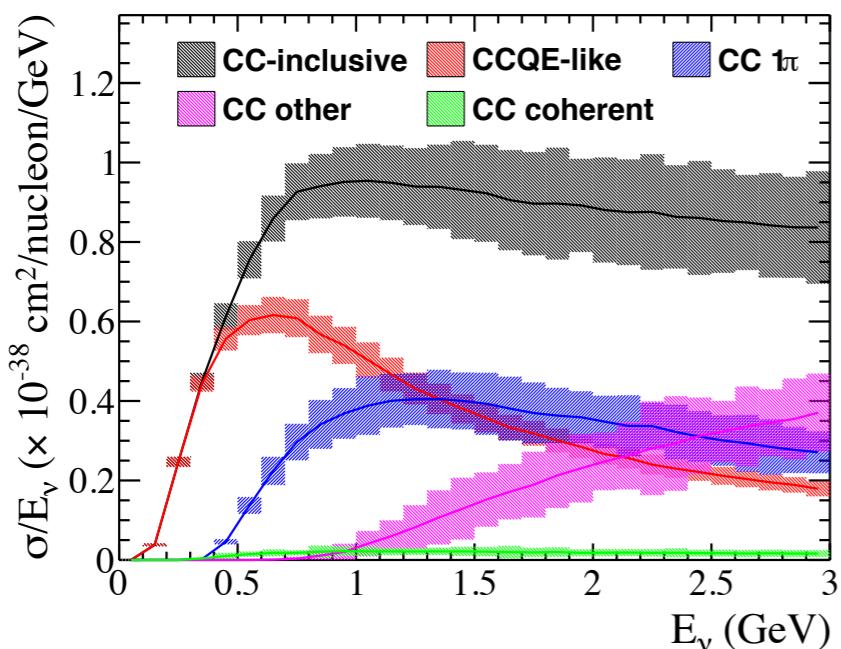
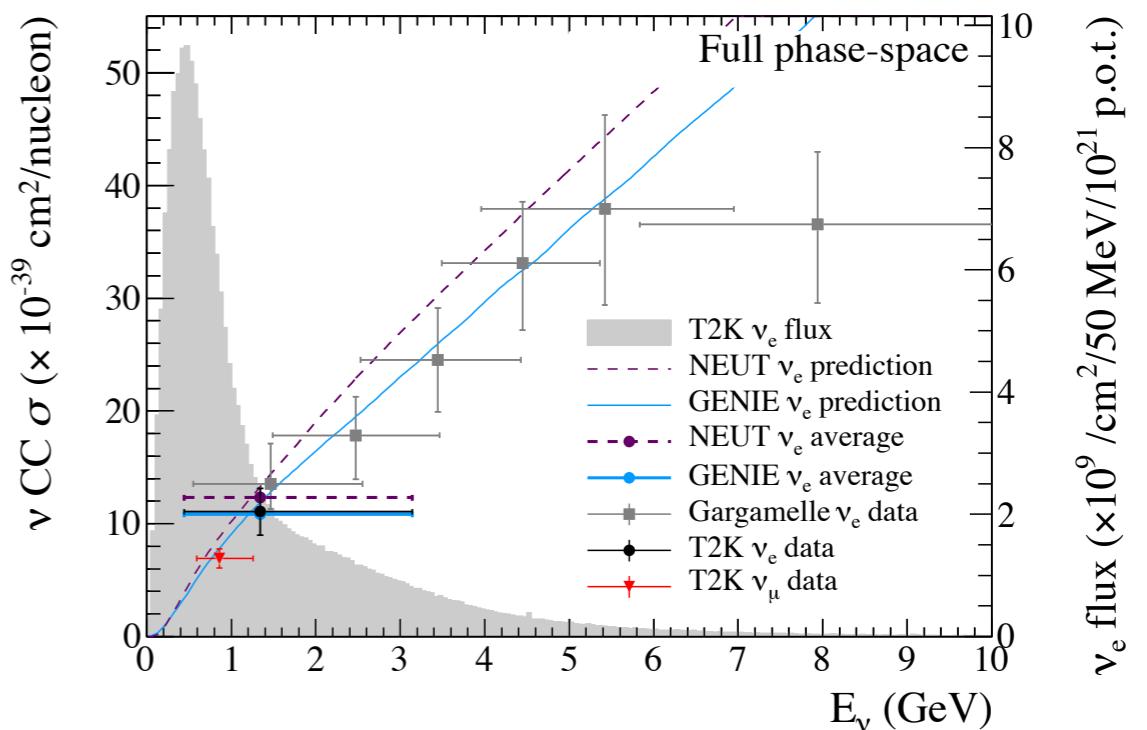


ND280 Flux

Phys. Rev. D 87, 092003 (2013)



Phys. Rev. Lett. 113, 241803 (2014)



In neutrino-mode

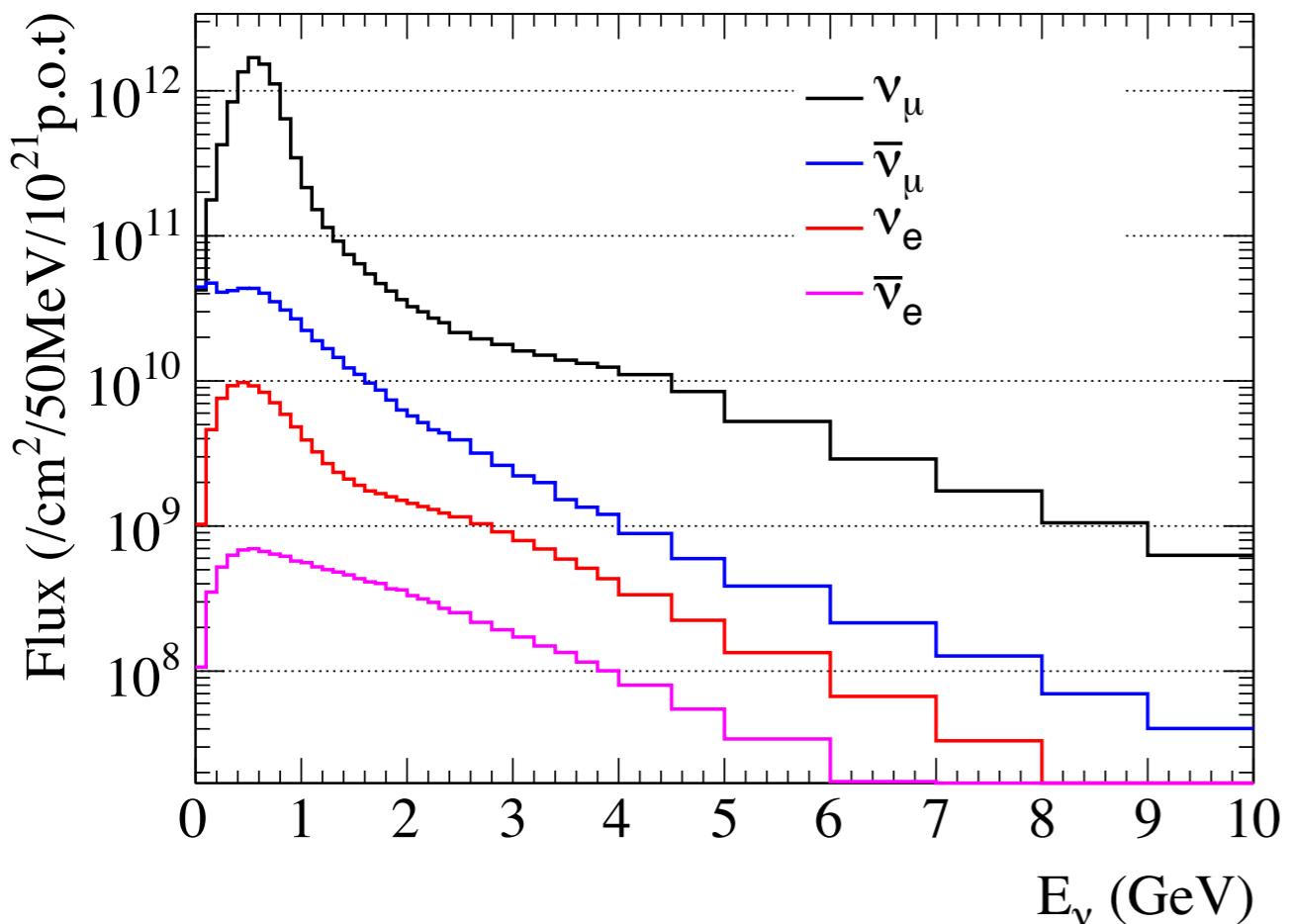
ν_μ : $\langle E \rangle = 0.85 \text{ GeV}, (\sim 90\%)$

ν_e : $\langle E \rangle = 1.3 \text{ GeV}, (\sim 1\%)$

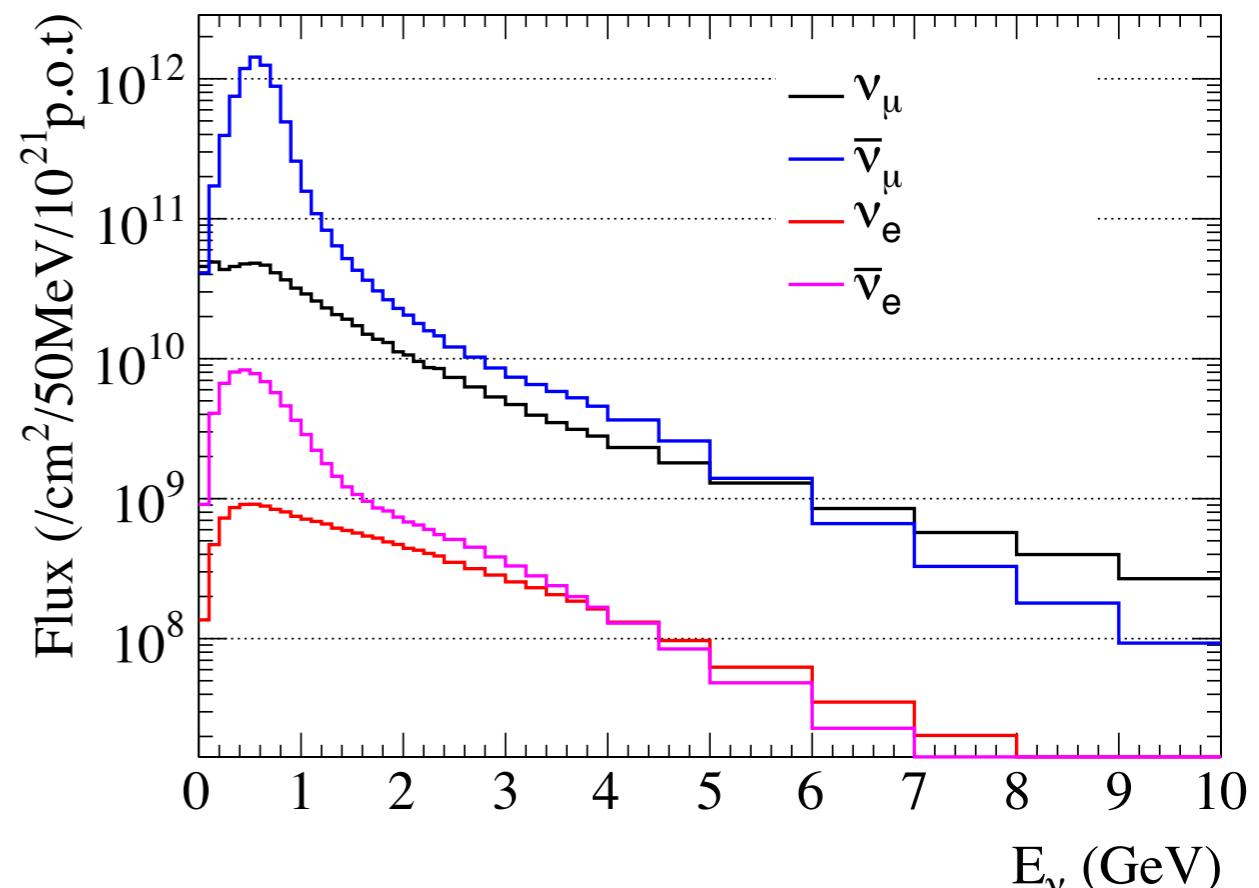
Dominant Reaction: CCQE
Single Pion Production

Flux at ND280

Neutrino Mode Flux at ND280



Antineutrino Mode Flux at ND280



In neutrino-mode

ν_μ : ⟨E⟩ = 0.85 GeV, (~90%)

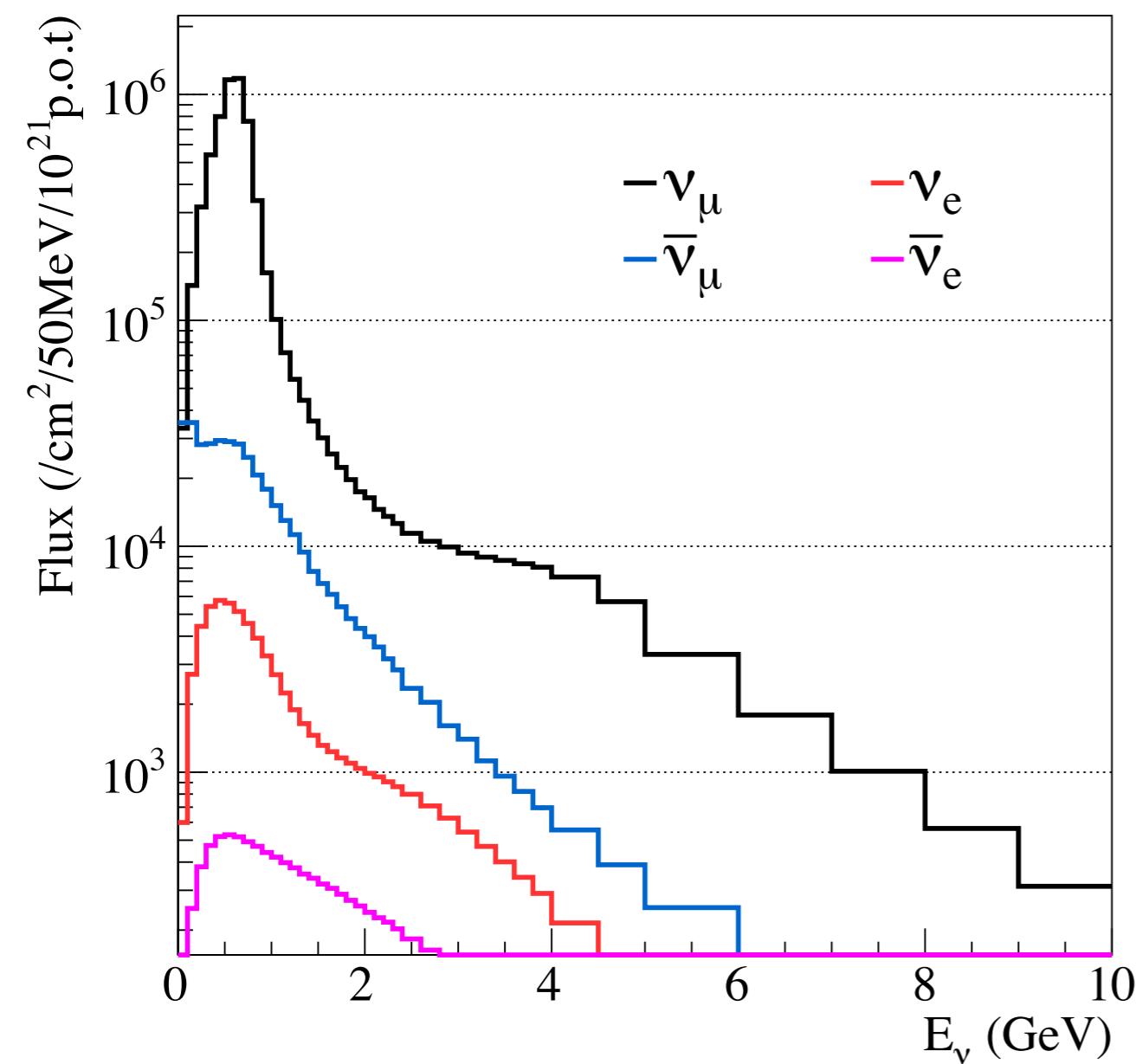
ν_e : ⟨E⟩ = 1.3 GeV, (~1%)

Dominant Reaction: CCQE
Single Pion Production

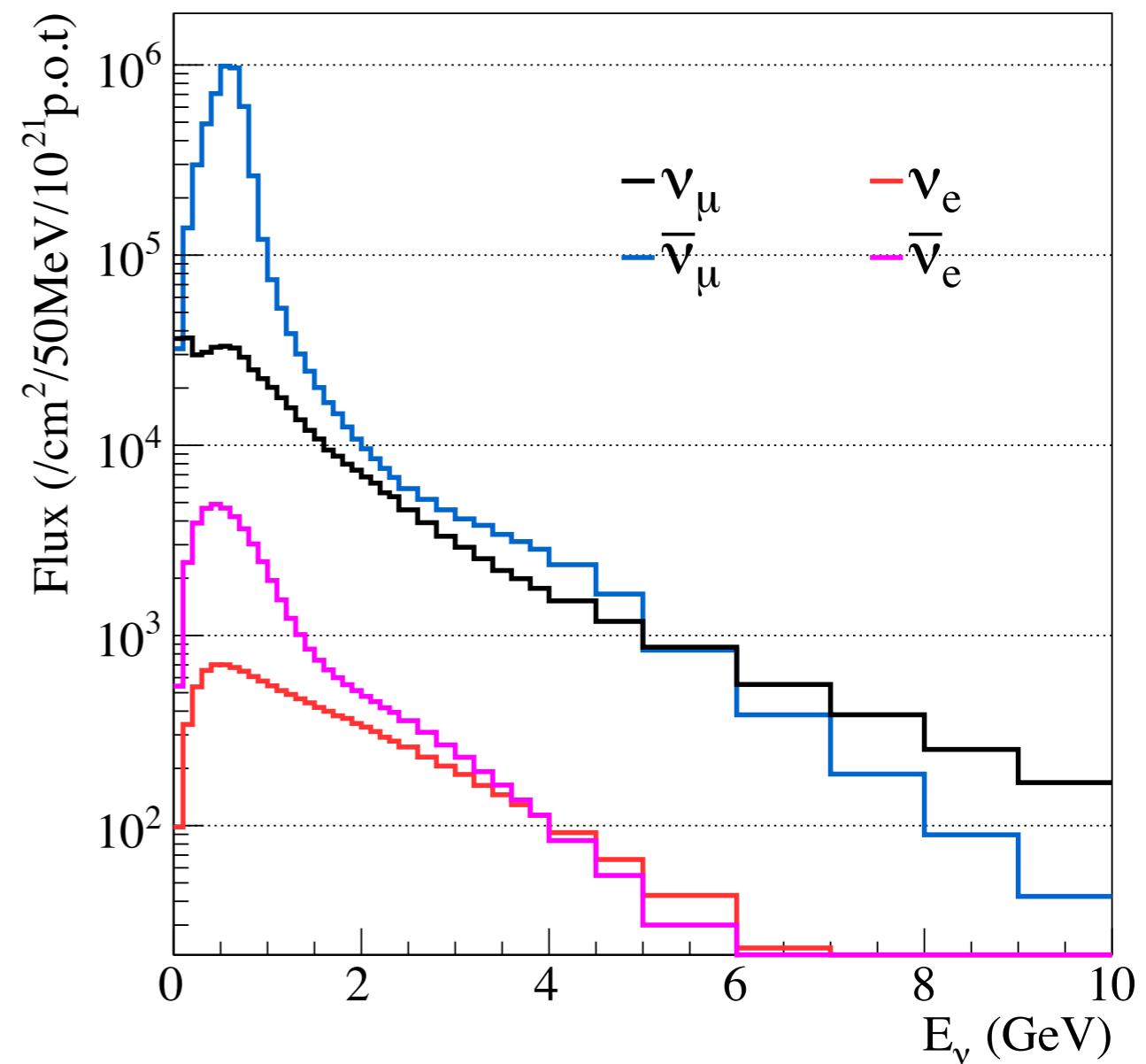
Flux at Super-K

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Neutrino Mode Flux at SK

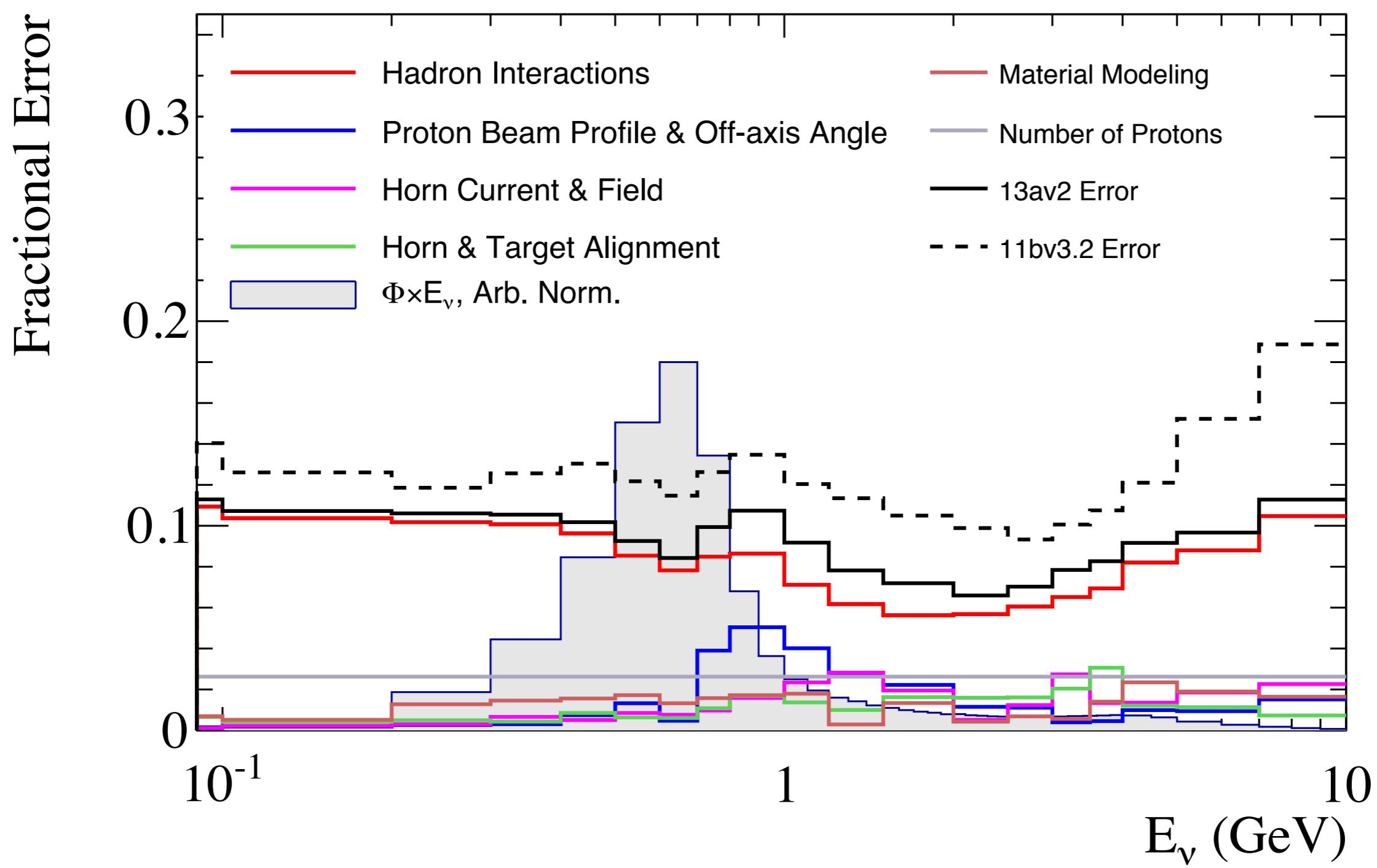


Antineutrino Mode Flux at SK

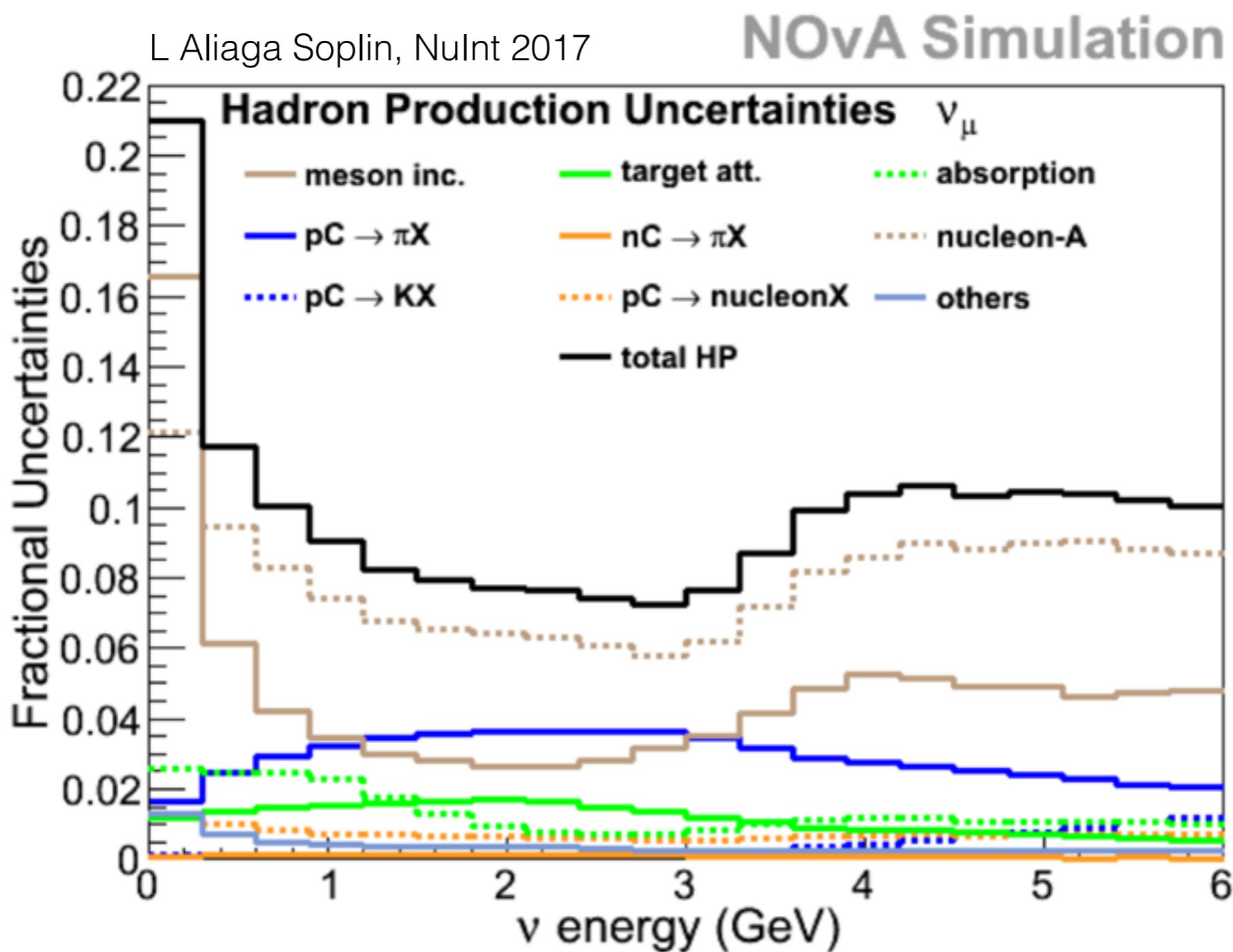


Flux Uncertainty

SK: Neutrino Mode, ν_μ

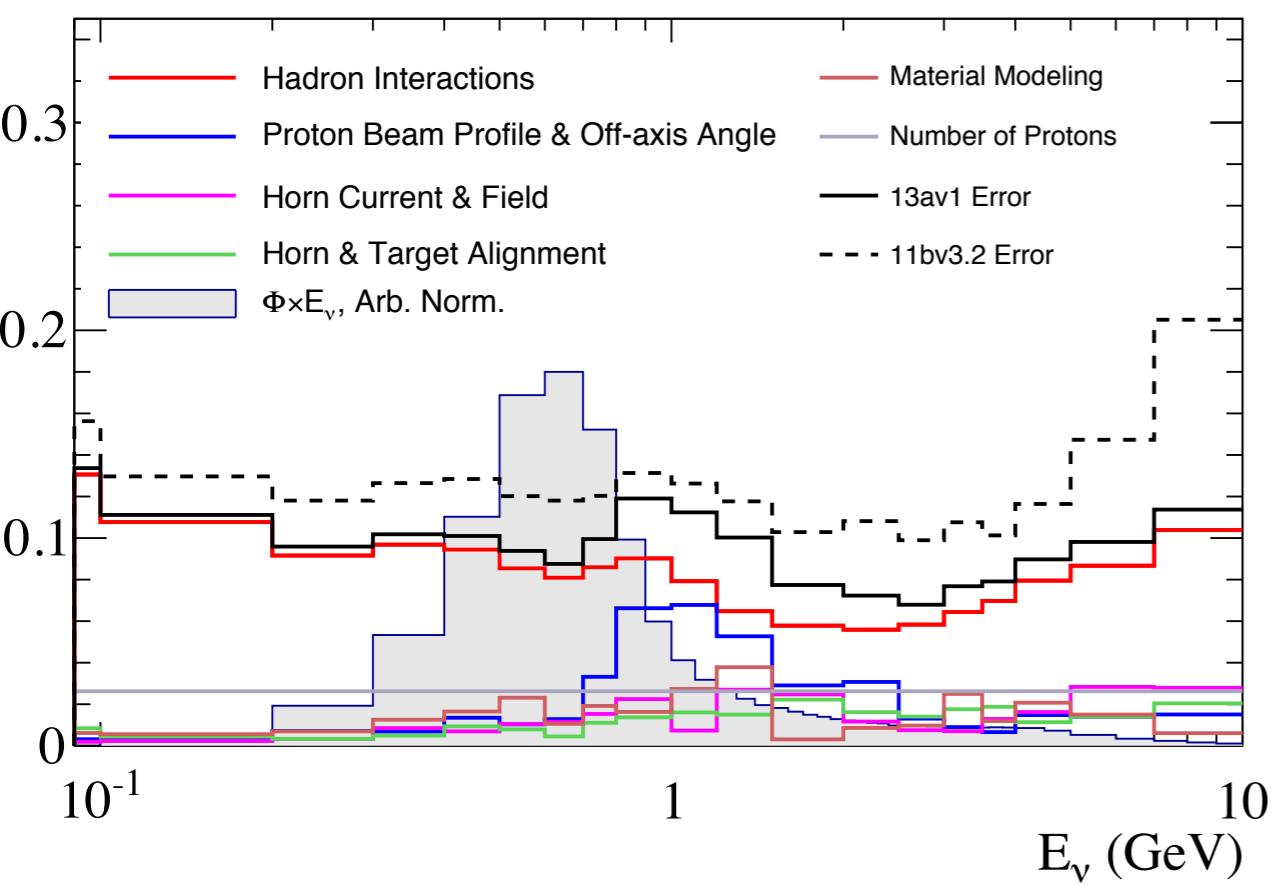


Flux Uncertainty

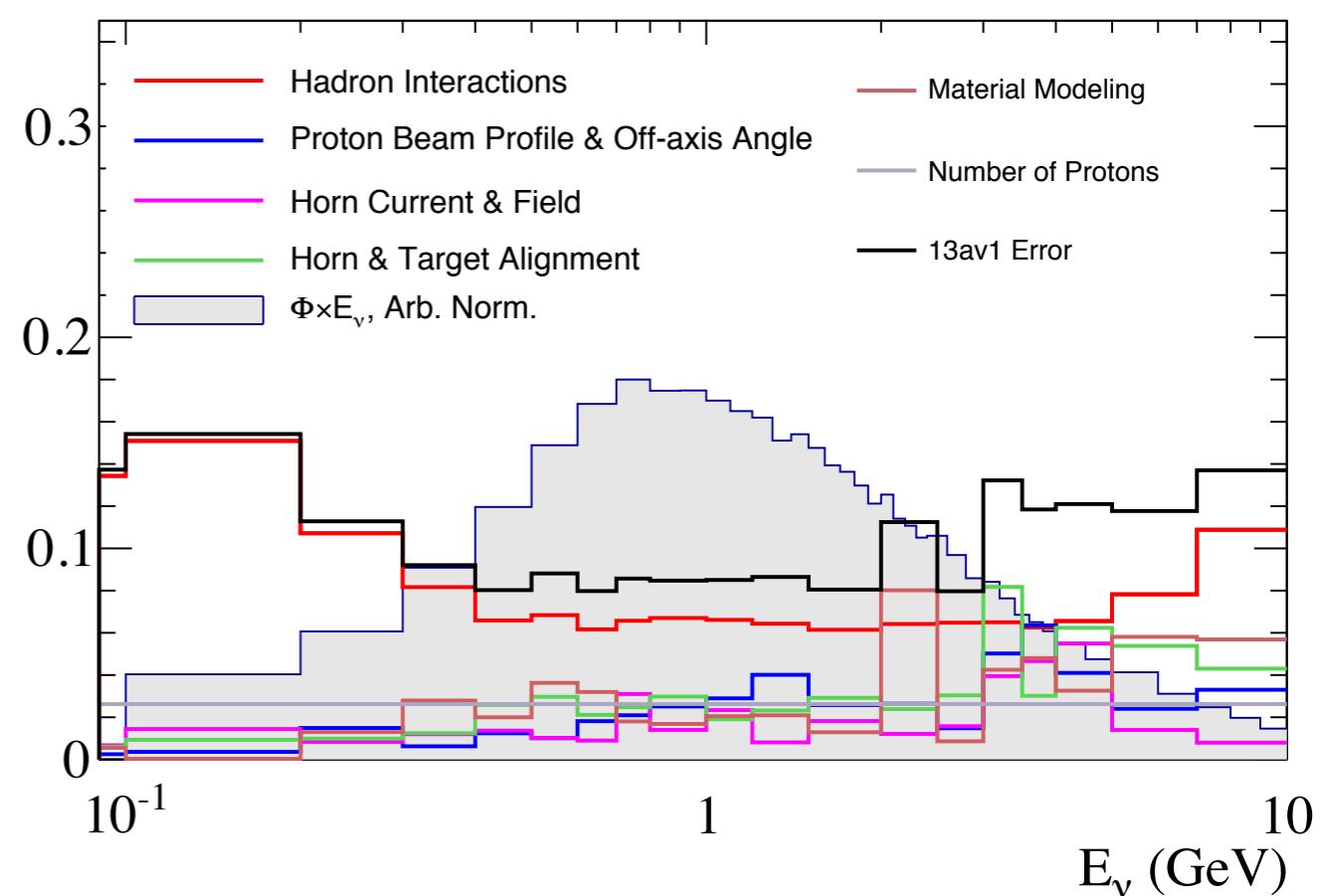


Flux at ND280

ND280: Neutrino Mode, ν_μ

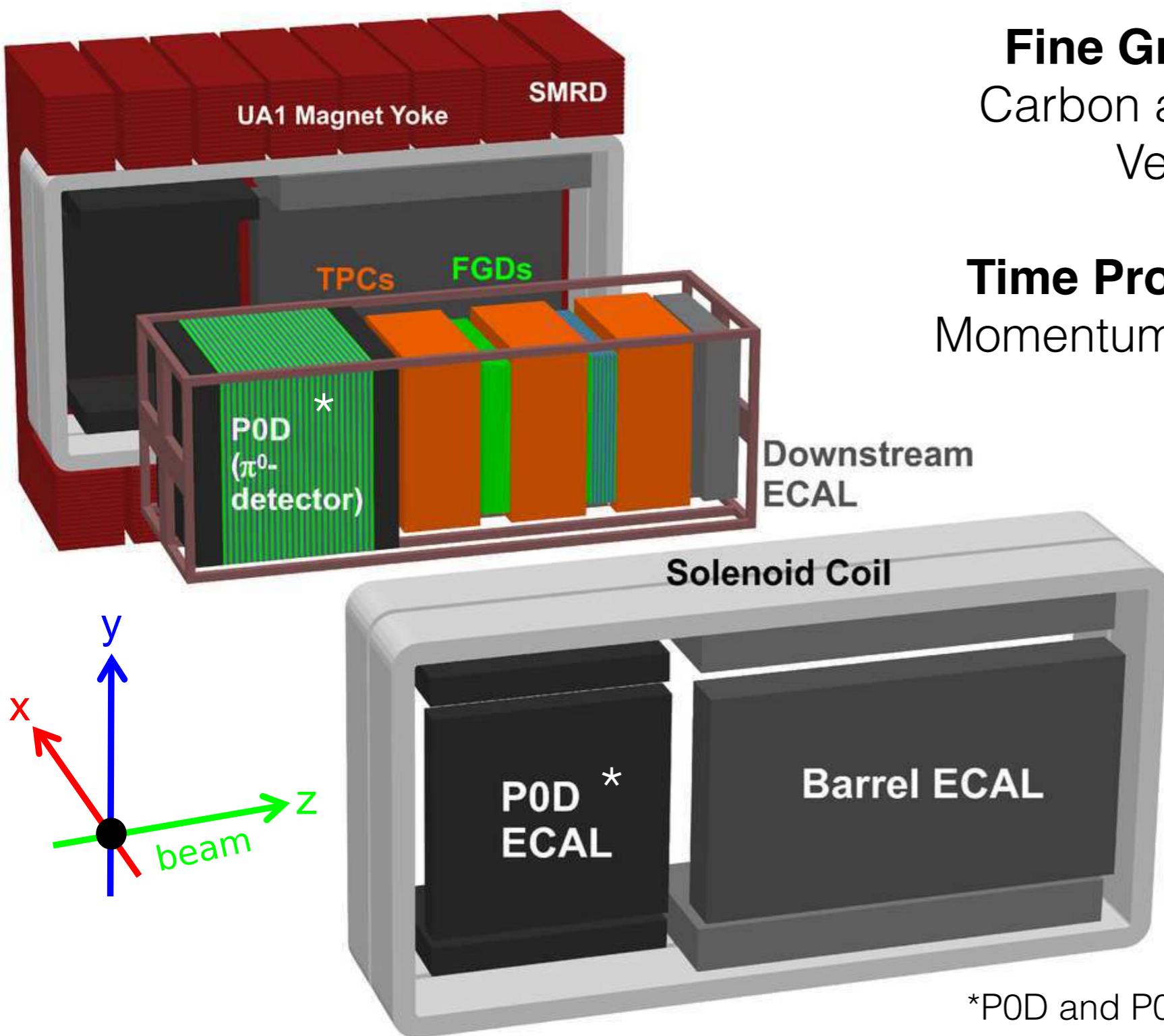


ND280: Antineutrino Mode, $\bar{\nu}_\mu$



ND280 Detector

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Fine Grained Detectors (FGD)

Carbon and Oxygen Target Mass,
Vertex reconstruction

Time Projection Chambers (TPC)

Momentum and Charge Measurement
Particle ID

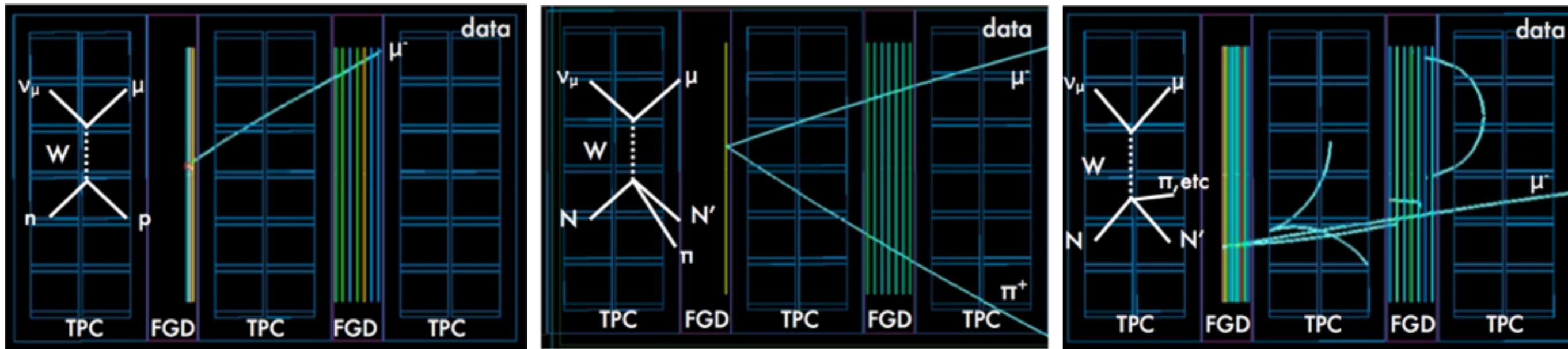
EM Calorimeters

Neutral Particle Reconstruction
Additional PID and
energy measurement
Tag entering backgrounds

*P0D and P0D ECal detectors not be discussed here.

See arXiv:1111.5030 and arXiv:1308.3445 for information on these detectors.

ND280 Input to T2K Oscillation Analysis



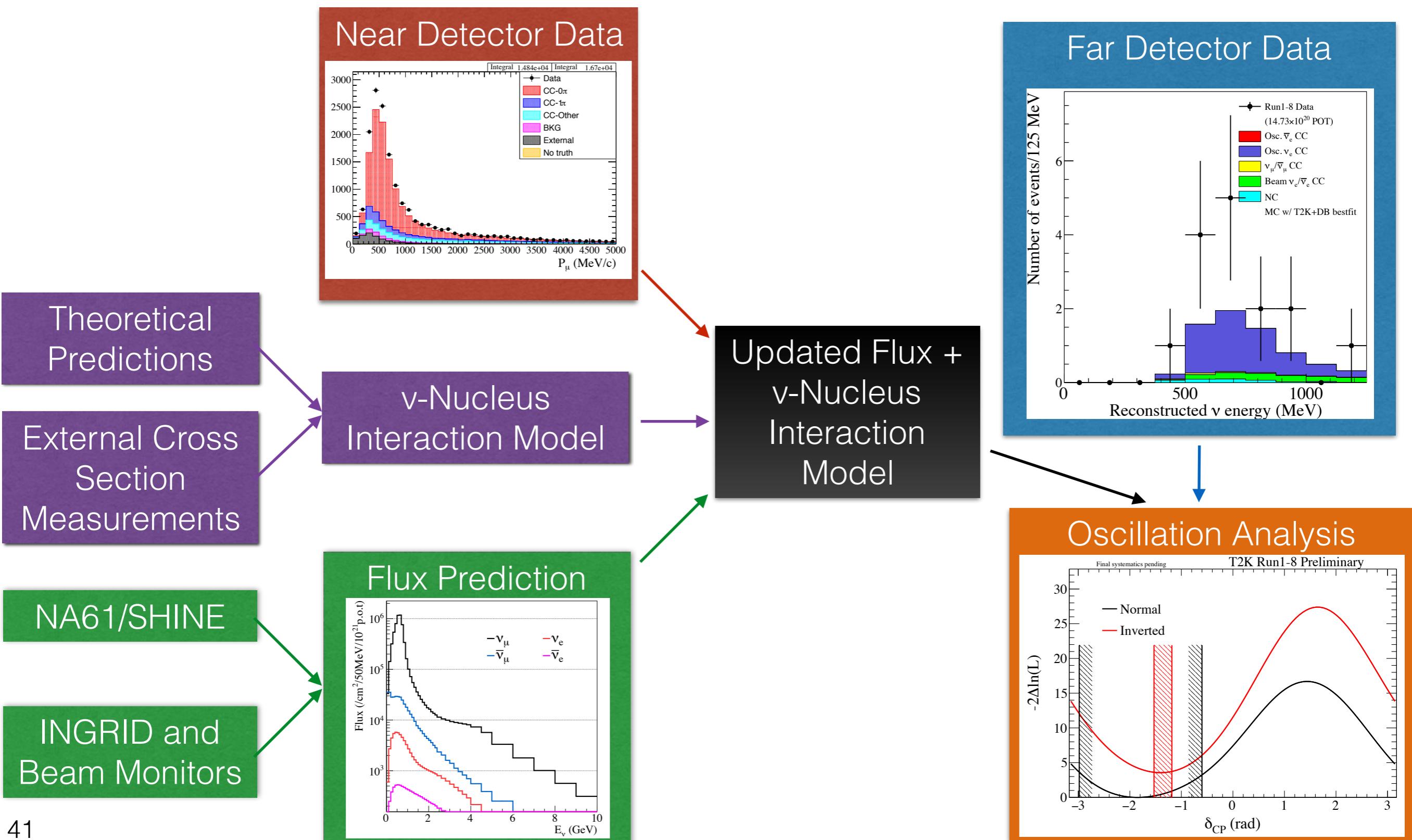
ND280 data split based on reconstructed topology enhanced in different interaction types

Fit flux + interaction model and propagate to far detector

As statistics increase and analysis becomes more sophisticated incorporate more channels

T2K Analysis Strategy

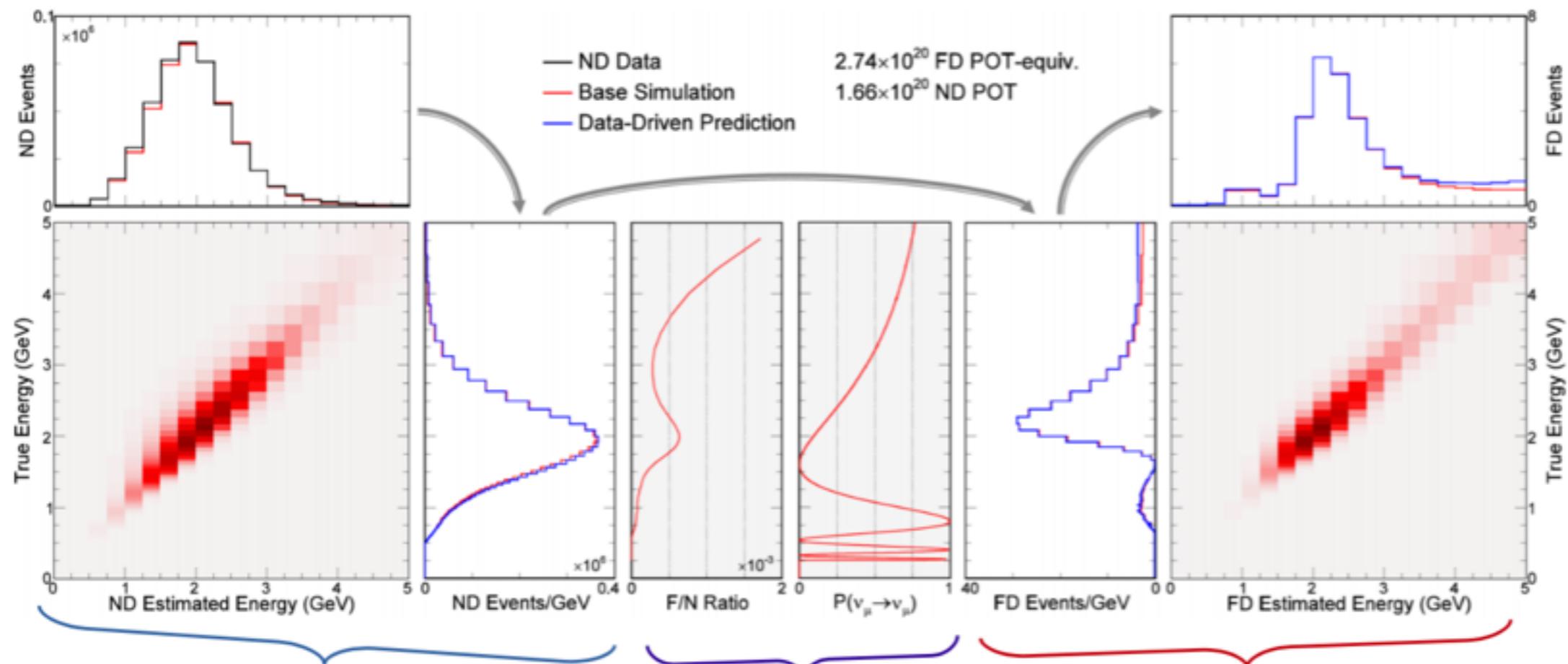
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NOvA Analysis Strategy

To produce a data-driven prediction at FD, based on ND:

J. Wolcot, Nulnt 2017



True energy distribution is corrected so that reconstructed data & MC agree at the ND...

...modified true energy distribution is propagated through predicted geometric beam dispersion & acceptance ratio, oscillations...

... and “extrapolated” reconstructed energy distribution computed to compare to data