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The Casimir effect within scattering theory

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New Journal of Physics **8** (2006) 243

Received 31 May 2006

Published 20 October 2006

Online at <http://www.njp.org/>

doi:10.1088/1367-2630/8/10/243

Abstract. We review the theory of the Casimir effect using scattering techniques. After years of theoretical effort, this formalism is now largely mastered so that the accuracy of theory–experiment comparisons is determined by the level of precision and pertinence of the description of experimental conditions. Due to an imperfect knowledge of the optical properties of real mirrors used in the experiment, the effect of imperfect reflection remains a source of uncertainty in theory–experiment comparisons. For the same reason, the temperature dependence of the Casimir force between dissipative mirrors remains a matter of debate. We also emphasize that real mirrors do not obey exactly the assumption of specular reflection, which is used in nearly all calculations of material and temperature corrections. This difficulty may be solved by using a more general scattering formalism accounting for non-specular reflection with wavevectors and field polarizations mixed. This general formalism has already been fruitfully used for evaluating the effect of roughness on the Casimir force as well as the lateral Casimir force appearing between corrugated surfaces. The commonly used ‘proximity force approximation (PFA)’ turns out to lead to inaccuracies in the description of these two effects.

Contents

1. Introduction	2
2. Specular scattering	5
2.1. Finite conductivity correction	7
2.2. Temperature correction	12
3. Non-specular scattering	14
3.1. Influence of surface roughness	15
3.2. Lateral Casimir force component	18
3.3. Comparison to experiments in a PS configuration	21
4. Conclusion	22
Acknowledgments	23
References	23

1. Introduction

After its prediction in 1948 [1], the Casimir force has been observed in a number of ‘historic’ experiments which confirmed its existence and main properties [2]–[5] (and references therein). With present-day technology, a new generation of Casimir force measurements has started since nearly a decade ago [6]–[12]. These experiments have reached a good enough accuracy to allow for a comparison between theoretical predictions and experimental observations which is of great interest for various reasons [13]–[15] (and references therein).

The Casimir force is the most accessible effect of vacuum fluctuations in the macroscopic world. As the existence of vacuum energy raises difficulties at the interface between the theories of quantum and gravitational phenomena, it is worth testing this effect with the greatest care and highest accuracy [16, 17]. A precise knowledge of the Casimir force is also a key point in many accurate force measurements for distances ranging from nanometre to millimetre. These experiments are motivated either by tests of Newtonian gravity at millimetric distances [18]–[21] or by searches for new weak forces predicted in theoretical unification models with nanometric to millimetric ranges [22]–[27]. Basically, they aim at putting limits on deviations of experimental results from present standard theory. As the Casimir force is the dominant force between two neutral non-magnetic objects in the range of interest, any new force would appear as a difference between experimental measurements and theoretical expectations of the Casimir force. On the technological side, the Casimir force has been shown to become important in the architecture of micro- and nano-oscillators (MEMS, NEMS) [28, 29]. In this context, it is extremely important to account for the conditions of real experiments.

The comparison between theory and experiment should take into account the important differences between the real experimental conditions and the ideal situation considered by Casimir. Casimir calculated the force between a pair of perfectly smooth, flat and parallel plates in the limit of zero temperature and perfect reflection (see figure 1). He found an expression for the force F_{Cas} and the corresponding energy E_{Cas} which only depend on the distance L , the area A and two fundamental constants, the speed of light c and Planck constant \hbar

$$F_{\text{Cas}} = \frac{\hbar c \pi^2 A}{240 L^4} = \frac{dE_{\text{Cas}}}{dL}, \quad E_{\text{Cas}} = -\frac{\hbar c \pi^2 A}{720 L^3}. \quad (1)$$

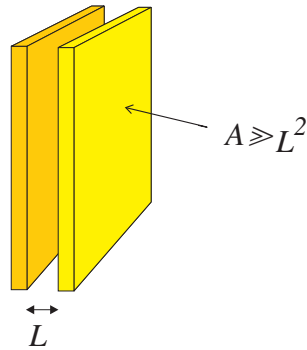


Figure 1. Original Casimir configuration of two plane parallel mirrors a distance L apart.

Each transverse dimension of the plates has been supposed to be much larger than L . Conventions of sign have been chosen so that F_{Cas} is positive while E_{Cas} is negative. They correspond to an attractive force ($\sim 0.1 \mu\text{N}$ for $A = 1 \text{ cm}^2$ and $L = 1 \mu\text{m}$) and a binding energy.

The fact that the Casimir force (1) only depends on fundamental constants and geometrical features is remarkable. In particular it is independent of the fine structure constant which appears in the expression of the atomic van der Waals forces. This universality property is related to the assumption of perfect reflection used by Casimir in his derivation. Perfect mirrors correspond to a saturated response to the fields since they reflect 100% of the incoming light. This explains why the Casimir effect, though it has its microscopic origin in the interaction of electrons with electromagnetic fields, does not depend on the fine structure constant.

However, no real mirror can be considered as a perfect reflector at all field frequencies. In particular, the most precise experiments are performed with metallic mirrors which show perfect reflection only at frequencies smaller than a characteristic plasma frequency ω_p which depends on the properties of conduction electrons in the metal. Hence the Casimir force between metal plates can fit the ideal Casimir formula (1) only at distances L much larger than the plasma wavelength

$$\lambda_p = \frac{2\pi c}{\omega_p}. \quad (2)$$

For metals used in the recent experiments, this wavelength lies in the $0.1 \mu\text{m}$ range (107 nm for Al and 137 nm for Cu and Au). At distances smaller than or of the order of the plasma wavelength, the finite conductivity of the metal has a significant effect on the force. The idea has been known since a long time [30]–[32] but a precise quantitative investigation of the effect of imperfect reflection has been systematically developed only recently [33]–[36]. As the effect of imperfect reflection is large in the most accurate experiments, a precise knowledge of its frequency dependence is essential for obtaining an accurate theoretical prediction of the Casimir force.

This is also true for other corrections to the ideal Casimir formula associated with the experimental configuration. For experiments at room temperature, the effect of thermal field fluctuations, superimposed to that of vacuum, affects the Casimir force at distances larger than a few micrometres. Again the idea has been known for a long time [37, 38] but a quantitative evaluation taking into account the correlation of this effect with that of imperfect reflection has been mastered only recently [39, 40]. A number of publications have given rise to contradictory

estimations of the Casimir force between dissipative mirrors at nonzero temperature [41]–[45]. Many attempts have been made to elucidate the problem by taking into account the low-frequency character of the force between metallic films [46], the spatial dispersion on electromagnetic surface modes [47] or the transverse momentum dependance of surface impedances [48]–[50]. Experimentally the effect of temperature of the Casimir force has not yet been conclusively measured [51]. For a recent analysis of this issue see [52].

Most experiments are performed between a plane and a sphere with the force estimation involving a geometry correction. Usually the Casimir force in the plane-sphere (PS) geometry is calculated using the proximity force approximation (PFA). This approximation amounts to the addition of force or energy contributions corresponding to different local inter-plate distances, assuming these contributions to be independent. But the Casimir force and energy are not additive, so the PFA cannot be exact, although it is often improperly called a theorem.

In this present paper, we consider both the original Casimir geometry with perfectly plane and parallel mirrors and the PS geometry when comparing to experiments. The PFA is expected to be valid in the PS geometry, when the sphere radius R is much larger than the separation L [53]–[55], which is the case for all present day experiments, and it will thus be used to connect the two geometries. In this case, the force F_{PS} between a sphere of radius R and a plane at a distance of closest approach L is given in terms of the energy E_{PP} for the plane-plane (PP) cavity as follows

$$F_{\text{PS}} = 2\pi R \frac{E_{\text{PP}}}{A}, \quad L \ll R. \quad (3)$$

Interesting attempts to go beyond this approximation concerning the PS geometry have been made recently [56, 57], also in the more general context of the connection between geometry and the Casimir effect [58]–[60].

Another important correction to the Casimir force is coming from surface roughness, which is intrinsic to any real mirror, with amplitude and spectrum varying depending on the surface preparation techniques. The departure from flatness of the metallic plates may also be designed, in particular under the form of sinusoidal corrugation of the plates which produce a measurable lateral component of the Casimir force [61]. For a long time, these roughness or corrugation corrections to the Casimir force have been calculated with methods valid only in limiting cases [62]–[67] or by using PFA [68]–[70]. Once again, it is only recently that emphasis has been put on the necessity of a more general method for evaluating the effect of roughness outside the region of validity of PFA with imperfect mirrors at arbitrary distances from each other [71]. While the condition $L \ll R$ is sufficient for applying the PFA in the PS geometry, more stringent conditions are needed for PFA to hold for rough or corrugated surfaces. The surfaces should indeed be nearly plane when looked at on a scale comparable with the separation L , and this condition is not always satisfied in experiments. When the PFA is no longer valid, the effect of roughness or corrugation can be evaluated by using the scattering theory extended to the case of non-specular reflection.

We review the Casimir effect within scattering theory and the theory of quantum optical networks. The main idea of this derivation is that the Casimir force has its origin in a difference of the radiation pressure of vacuum fields between the two mirrors and in the outer free field vacuum. This vacuum radiation pressure can be written as an integral over all modes, each mode being associated with reflection amplitudes on the two mirrors. We first present formulae written for specular reflection which are valid for lossless [72] as well as lossy mirrors [73]. We discuss the influence of the mirrors reflection coefficients at zero and nonzero temperature. We then

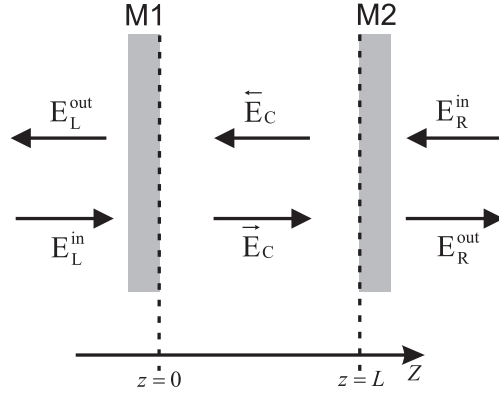


Figure 2. Schematic view of a Fabry-Pérot cavity of length L .

extend the approach to the case of non-specular reflection which mixes the field polarizations and transverse wavevectors. Finally, we apply the latter approach to the calculation of the roughness correction to the Casimir force between metallic mirrors [74, 75] and of the lateral component of the Casimir force between corrugated plates [76].

2. Specular scattering

Let us first consider the original Casimir geometry with perfectly plane and parallel mirrors aligned along the directions x and y . The two mirrors thus form a Fabry-Pérot cavity of length L as shown in figure 2. We analyse the cavity as a composed optical network, and calculate the fluctuations of the intracavity fields propagating along the positive and negative z -axis, \vec{E}_C and \vec{E}_C , in terms of the fluctuations of the incoming free-space fields E_L^{in} and E_R^{in} (the outgoing fields E_L^{out} and E_R^{out} are also shown).

The field modes are conveniently characterized by their frequency ω , transverse wavevector \mathbf{k} with components k_x, k_y in the plane of the mirrors and polarization p . As the configuration of figure 2 obeys a symmetry with respect to time translation as well as transverse space translations (along directions x and y), the frequency ω , transverse vector $\mathbf{k} \equiv (k_x, k_y)$ and polarization $p = \text{TE, TM}$ are preserved throughout the whole scattering processes on a mirror or a cavity. The scattering couples only the free vacuum modes which have the same values for the preserved quantum numbers and differ by the sign of the longitudinal component k_z of the wavevector. We denote by $(r_{\mathbf{k}}^p(\omega))_j$ the reflection amplitude of the mirror $j = 1, 2$ as seen from the inner side of the cavity. This scattering amplitude obeys general properties of causality, unitarity and high frequency transparency. The additional fluctuations accompanying losses inside the mirrors are deduced from the optical theorem applied to the scattering process which couples the modes of interest and the noise modes [77, 78].

The loop functions that characterize the optical response of the cavity to an input field play an important role in the following

$$f_{\mathbf{k}}^p[\omega] = \frac{\rho_{\mathbf{k}}^p[\omega]}{1 - \rho_{\mathbf{k}}^p[\omega]}, \quad \rho_{\mathbf{k}}^p[\omega] = (r_{\mathbf{k}}^p[\omega])_1 (r_{\mathbf{k}}^p[\omega])_2 e^{2ik_z L}, \quad (4)$$

$\rho_{\mathbf{k}}^p$ and $f_{\mathbf{k}}^p$ are respectively the open-loop and closed-loop functions corresponding to one round trip in the cavity. The system formed by the mirrors and fields is stable so that $f_{\mathbf{k}}^p$ is an analytic

function of frequency ω . Analyticity is defined with the following physical conditions in the complex plane

$$\omega \equiv i\xi, \quad \Re\xi > 0, \quad k_z \equiv i\kappa[\omega], \quad \kappa[\omega] \equiv \sqrt{\mathbf{k}^2 - \frac{\omega^2}{c^2}}, \quad \Re\kappa[\omega] > 0. \quad (5)$$

The quantum numbers p and \mathbf{k} remain spectator throughout the discussion of analyticity. The sum on transverse wavevectors may be represented as a sum over the eigenvectors $k_x = 2\pi q_x/L_x, k_y = 2\pi q_y/L_y$ associated with virtual quantization boxes along x, y or, at the continuum limit $L_x, L_y \rightarrow \infty$ with $A = L_x L_y$, as an integral

$$\sum_{\mathbf{k}} \equiv \sum_{q_x=-\infty}^{\infty} \sum_{q_y=-\infty}^{\infty} \rightarrow A \int_{-\infty}^{\infty} \frac{dk_x}{2\pi} \int_{-\infty}^{\infty} \frac{dk_y}{2\pi}. \quad (6)$$

We then introduce the Airy function defined in classical optics as the ratio of energy inside the cavity to energy outside the cavity for a given mode

$$g_{\mathbf{k}}^p[\omega] = 1 + \{f_{\mathbf{k}}^p[\omega] + \text{c.c.}\} = \frac{1 - |\rho_{\mathbf{k}}^p[\omega]|^2}{|1 - \rho_{\mathbf{k}}^p[\omega]|^2}. \quad (7)$$

$f_{\mathbf{k}}^p, g_{\mathbf{k}}^p$ depend only on the reflection amplitudes of mirrors as they are seen from the inner side. With these definitions, we write the Casimir force

$$F = -\hbar \sum_p \sum_{\mathbf{k}} \int_0^{\infty} \frac{d\omega}{2\pi} [i\kappa[\omega] f_{\mathbf{k}}^p[\omega] + \text{c.c.}], \quad (8)$$

or, equivalently, the Casimir energy

$$E = -\hbar \sum_p \sum_{\mathbf{k}} \int_0^{\infty} \frac{d\omega}{2\pi} \frac{1}{2i} \ln \left[\frac{1 - \rho_{\mathbf{k}}^p[\omega]}{1 - \rho_{\mathbf{k}}^p[\omega]^*} \right]. \quad (9)$$

Equations (8, 9) contain the contribution of ordinary modes freely propagating outside and inside the cavity with $\omega > c|\mathbf{k}|$ and k_z real. This contribution thus merely reflects the intuitive picture of a radiation pressure of fluctuations on the mirrors of the cavity [72] with the factor $g_{\mathbf{k}}^p - 1$ representing a difference between inner and outer sides. Equations (8, 9) also include the contribution of evanescent waves with $\omega < c|\mathbf{k}|$ and k_z imaginary. Those waves propagate inside the mirrors with an incidence angle larger than the limit angle and they also exert a radiation pressure on the mirrors, due to the frustrated reflection phenomenon [73]. Their properties are conveniently described through an analytical continuation of those of ordinary waves, using the well-defined analytic behaviour of κ and $f_{\mathbf{k}}^p$.

Using analyticity properties, we now transform (8) into an integral over imaginary frequencies by applying the Cauchy theorem on the contour enclosing the quadrant $\Re\omega > 0, \Im\omega > 0$. We use high frequency transparency to neglect the contribution of large frequencies. This leads to the following expression for the Casimir force

$$F = \hbar \sum_p \sum_{\mathbf{k}} \int_0^{\infty} \frac{d\xi}{2\pi} \{\kappa[i\xi] f_{\mathbf{k}}^p[i\xi] + \text{c.c.}\}, \quad (10)$$

which is now written as an integral over complex frequencies $\omega = i\xi$. In the same way, we obtain the Casimir energy as a function of imaginary frequencies

$$E = \frac{\hbar A}{2\pi} \sum_p \int \frac{d^2\mathbf{k}}{4\pi^2} \int_0^\infty d\xi \ln\{1 - \rho_{\mathbf{k}}^p(i\xi)\}. \quad (11)$$

Causality and passivity conditions assure that the integrand $\ln(1 - \rho_{\mathbf{k}}^p(i\xi))$ is analytical in the upper half space of the complex plane $\Re\xi > 0$. It is thus clear that both expressions for force and energy are equivalent.

2.1. Finite conductivity correction

Let us now review the correction to the Casimir force coming from the finite conductivity of any material. This correction is given by relations (8) or, equivalently, (10), as soon as the reflection amplitudes are known. These amplitudes are commonly deduced from models of mirrors, in particular bulk mirrors, slabs or layered mirrors, the optical response of metallic matter being described by some permittivity function. This function may be either a simple description of conduction electrons in terms of a plasma or Drude model or a more elaborate representation based upon tabulated optical data. At the end of the section, we will discuss the uncertainty in the theoretical evaluation of the Casimir force coming from the lack of knowledge of the specific material properties of a given mirror as was illustrated in [34].

Assuming that the metal plates have a large optical thickness, the reflection coefficients correspond to the ones of a simple vacuum–bulk interface [80]

$$r^{\text{TE}} = -\frac{\sqrt{\xi^2(\varepsilon(i\xi) - 1) + c^2\kappa^2} - c\kappa}{\sqrt{\xi^2(\varepsilon(i\xi) - 1) + c^2\kappa^2} + c\kappa}, \quad r^{\text{TM}} = \frac{\sqrt{\xi^2(\varepsilon(i\xi) - 1) + c^2\kappa^2} - c\kappa\varepsilon(i\xi)}{\sqrt{\xi^2(\varepsilon(i\xi) - 1) + c^2\kappa^2} + c\kappa\varepsilon(i\xi)}, \quad (12)$$

r^p stands for $r^p(i\xi, i\kappa)$ and $\varepsilon(i\xi)$ is the dielectric function of the metal evaluated for imaginary frequencies; the index \mathbf{k} has been dropped.

Taken together, the relations (10) and (12) reproduce the Lifshitz expression for the Casimir force [30]. Note that the expression was not written in this manner by Lifshitz. To our present knowledge, Kats [81] was the first to stress that Lifshitz expression could be written in terms of the reflection amplitudes (12). We then have to emphasize that (10) is much more general than Lifshitz expression since it still holds with mirrors characterized by reflection amplitudes differing from (12). As an illustration, we may consider metallic slabs having a finite thickness.

For a given polarization, we denote by r_{si} the reflection coefficient (12) corresponding to a single vacuum/metal interface and we write the reflection amplitude r for the slab of finite thickness through a Fabry–Pérot formula

$$r = r_{\text{si}} \frac{1 - e^{-2\delta}}{1 - r_{\text{si}}^2 e^{-2\delta}}, \quad \delta = \frac{D}{c} \sqrt{\xi^2(\varepsilon(i\xi) - 1) + c^2\kappa^2}. \quad (13)$$

This expression has been written directly for imaginary frequencies. The parameter δ represents the optical length in the metallic slab and D the physical thickness. The single interface expression (12) is recovered in the limit of a large optical thickness $\delta \gg 1$. With the plasma model, this condition just means that the thickness D is larger than the plasma wavelength λ_p .

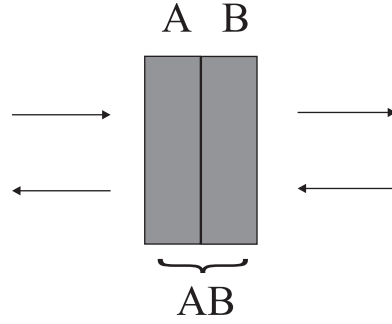


Figure 3. Composition of networks two networks labelled A and B are piled up to build up a network AB.

In order to discuss experiments, it may also be worth to write the reflection coefficients for multilayer mirrors. For example one may consider two-layer mirrors with a layer of thickness D of a metal A deposited on a large slab of metal B in the limit of large thickness as shown in figure 3. The reflection formulae are then obtained as in [82] but accounting for oblique incidence

$$r_{AB} = r_A + \frac{t_A^2 r_B}{1 - r_A r_B}, \quad t_{AB} = \frac{t_A t_B}{1 - r_A r_B}. \quad (14)$$

It reproduces the known results for the simple multilayer systems which have already been studied [13]. The combination of (13) and (14) allows to calculate most of the experimental situations precisely.

In order to assess quantitatively the effect of finite conductivity, we may in a first approach use the plasma model for the metallic dielectric function, with ω_p the plasma frequency,

$$\varepsilon[\omega] = 1 - \frac{\omega_p^2}{\omega^2}, \quad \varepsilon(i\xi) = 1 + \frac{\omega_p^2}{\xi^2}. \quad (15)$$

It is convenient to present the change in the Casimir force in terms of a factor η_F which measures the reduction of the force with respect to the case of perfect mirrors

$$F = \eta_F F_C \quad (16)$$

Using expressions (12) and (15), it is possible to obtain the reduction factor defined for the Casimir force through numerical integrations.

The result is plotted as the solid line on figure 4, as a function of the dimensionless parameter L/λ_p , that is the ratio between the distance L and the plasma wavelength λ_p . As expected the Casimir formula is reproduced at large distances ($\eta_F \rightarrow 1$ when $L \gg \lambda_p$). At distances smaller than λ_p in contrast, a significant reduction is obtained with the asymptotic law of variation read as [84, 85]

$$L \ll \lambda_p \rightarrow \eta_F = \alpha \frac{L}{\lambda_p} \quad \alpha \simeq 1.193. \quad (17)$$

This can be understood as the result of the Coulomb interaction of surface plasmons at the two vacuum/metal interfaces [83, 84]. The generalization of this idea at arbitrary distances is more

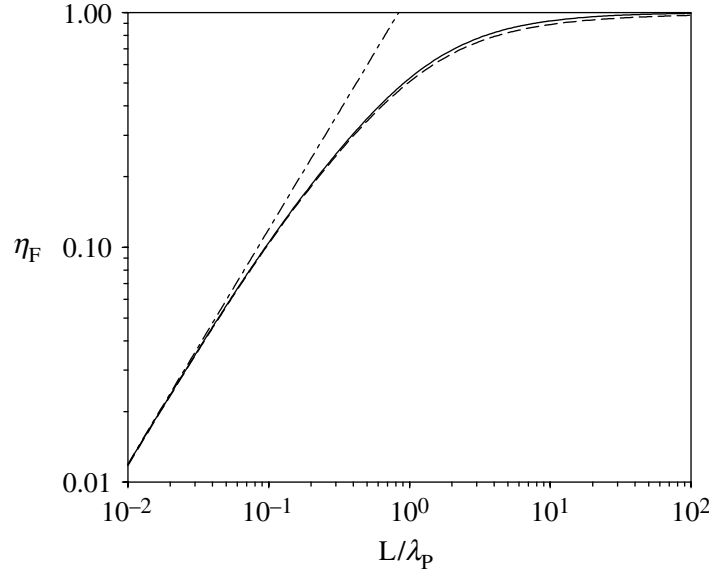


Figure 4. Reduction of the Casimir force compared to the force between perfect mirrors, when the finite conductivity is described by a plasma model (solid line) or a Drude model (dashed line) with a ratio $\frac{\gamma}{\omega_p}$ equal to 4×10^{-3} . The difference due to the relaxation parameter has only a small effect on the calculation of the Casimir force. The dotted-dashed line corresponds to the short distance asymptotic behaviour (17).

subtle since it involves a full electromagnetic treatment of the plasmon as well as ordinary photon modes [86].

The plasma model cannot provide a fully satisfactory description of the optical response of metals, in particular because it does not account for any dissipative mechanism. A more realistic representation is the Drude model [87]

$$\varepsilon[\omega] = 1 - \frac{\omega_p^2}{\omega(\omega + i\gamma)}, \quad \varepsilon(i\xi) = 1 + \frac{\omega_p^2}{\xi(\xi + \gamma)}. \quad (18)$$

This model describes not only the plasma response of conduction electrons with ω_p still interpreted as the plasma frequency but also their relaxation, γ being the inverse of the electronic relaxation time.

The relaxation parameter γ is much smaller than the plasma frequency. For Al, Au, Cu in particular, the ratio γ/ω_p is of the order of 4×10^{-3} . Hence relaxation affects the dielectric constant in a significant manner only at frequencies where the latter is much larger than unity. In this region, the metallic mirrors behave as a nearly perfect reflectors so that, finally, the relaxation does not have a large influence on the Casimir effect at zero temperature. This qualitative discussion is confirmed by the result of numerical integration reported as the dashed line on figure 4. With the typical value already given for γ/ω_p , the variation of η_F remains everywhere smaller than 2%.

For metals like Al, Au, Cu, the dielectric constant departs from the Drude model when interband transitions are reached, that is when the photon energy reaches a few eV. Hence, a more precise description of the dielectric constant should be used for evaluating the Casimir force in

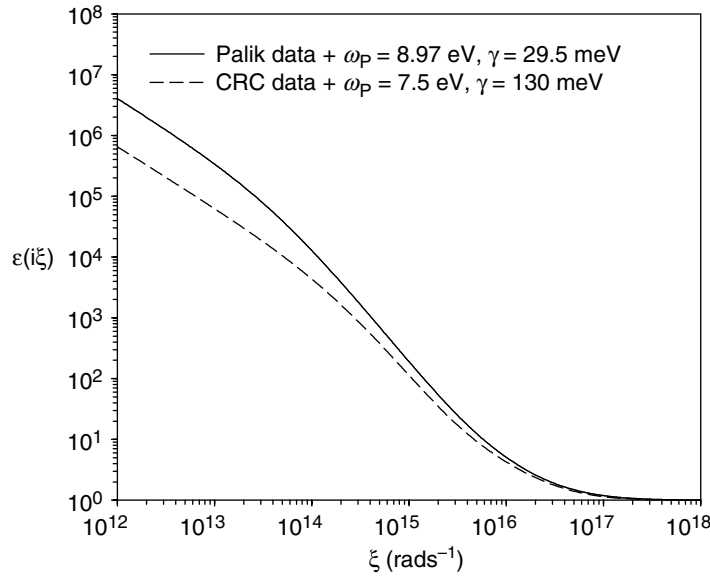


Figure 5. Dielectric function of Cu versus imaginary frequency. The solid line corresponds to the first set (optical data given by [89, 90], low frequencies extrapolated by a Drude model with $\omega_p = 8.97$ eV and $\gamma = 29.5$ meV), the dashed line to the second set (optical data given by [91], low frequencies extrapolated by a Drude model with $\omega_p = 7.5$ eV and $\gamma = 130$ meV).

the sub micrometre range. This description relies on one hand on the causality relations obeyed by the dielectric response function and on another hand on known optical data. The reader is referred to [34] for a detailed analysis, but we recall here the main argument and some important details. Let us first recall that frequencies are measured either in eV or in rad/s, using the equivalence $1 \text{ eV} = 1.519 \times 10^{15} \text{ rad s}^{-1}$. An erroneous conversion factor $1 \text{ eV} = 1.537 \times 10^{15} \text{ rad s}^{-1}$ was used in [34], which led to a difference in $\varepsilon(i\xi)$ of less than 1% over the relevant distance range. In the end of the calculation, this was corresponding to a negligible error in the Casimir force and energy [88].

The values of the complex index of refraction for different metals, measured through different optical techniques, are tabulated as a function of frequency in several handbooks [89]–[91]. Optical data may vary from one reference to another, not only because of experimental uncertainties but also because of the dispersion of material properties of the analysed samples. Moreover, the available data do not cover a broad enough frequency range so that they have to be extrapolated. These problems may cause variations of the results obtained for the dielectric function $\varepsilon(i\xi)$ and, therefore, for the Casimir force.

Figure 5 shows two different plots of $\varepsilon(i\xi)$ for Cu as a function of imaginary frequency ξ . The solid line corresponds to the first data set with data points taken from [89, 90] and extrapolation at low frequency with a Drude model with parameters $\omega_p = 8.97$ eV and $\gamma = 29.5$ meV in reasonable agreement with existing knowledge from solid state physics. However, as explained in [34], the optical data available for Cu do not permit an unambiguous estimation of the two parameters ω_p and γ separately. Other couples of values can be chosen which are also consistent with optical data. To make this point explicit, we have drawn a second plot on figure 5 (dashed line) with data taken from [91] and the low frequency interpolation given by a Drude model with

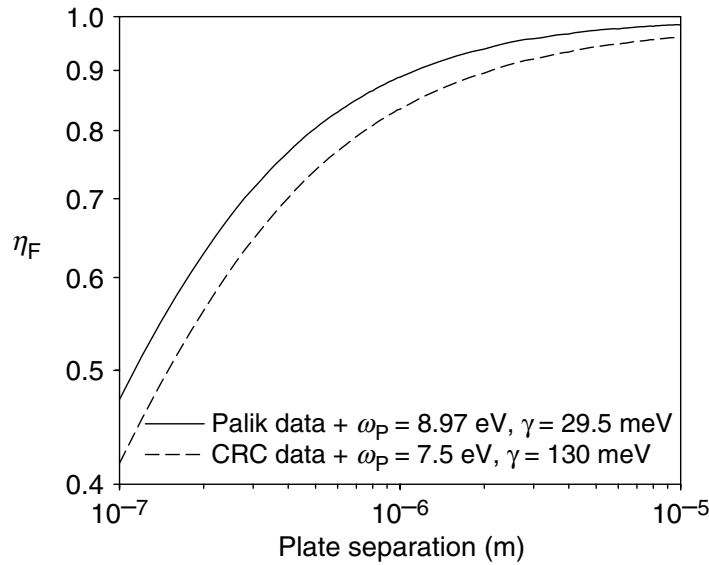


Figure 6. Reduction factor η_F for the Casimir force between two Cu plates as a function of the plate separation L . The solid line corresponds to the first set (optical data given by [89], low frequencies extrapolated by a Drude model with $\omega_P = 8.97$ eV and $\gamma = 29.5$ meV), the dashed line to the second set (optical data given by [91], low frequencies extrapolated by a Drude model with $\omega_P = 7.5$ eV and $\gamma = 130$ meV).

$\omega_P = 7.5$ eV and $\gamma = 130$ meV. These values lead to a dielectric function $\varepsilon(i\xi)$ smaller than in the first data set over the whole frequency range, but especially at low frequencies. An estimation of the uncertainties associated with this imperfect knowledge of optical data can be drawn from the computation of the Casimir force in these two cases.

Figure 6 shows the reduction factor η_F for the Casimir force between two Cu plates as a function of the plate separation L for the two sets of optical data. The two corresponding curves have similar dependences on the plate separation but the absolute values are shifted from one curve to the other. At a separation of 100 nm the difference can be as large as 5%. As the plasma frequency is basically the frequency above which the mirrors reflectivity diminishes considerably, the Drude parameters of the first set ($\omega_P = 8.97$ eV and $\gamma = 29.5$ meV) give a larger Casimir force than the second set, where the plasma frequency is lower ($\omega_P = 7.5$ eV and $\gamma = 130$ meV). A detailed analysis of this uncertainty has been recently reported [92].

Let us emphasize that the problem here is neither due to a lack of precision of the calculations nor to inaccuracies in experiments. The problem is that calculations and experiments may consider physical systems with different optical properties. Material properties of mirrors indeed vary considerably as a function of external parameters and preparation procedure [92]. This difficulty could be solved by measuring the reflection amplitudes of the mirrors used in the experiment and then inserting these informations in the formula giving the predicted Casimir force. In order to suppress the uncertainty associated with the extrapolation procedure, it would be necessary to measure the reflection amplitudes down to frequencies of the order of 1 meV, if the aim is to calculate the Casimir force in the distance range from 100 nm to a few micrometres.

2.2. Temperature correction

The Casimir force between dissipative metallic mirrors at nonzero temperature has given rise to contradictory claims which have raised doubts about the theoretical expression of the force. In order to contribute to the resolution of this difficulty, we now review briefly the derivation of the force from basic principles of the quantum theory of lossy optical cavities at nonzero temperature. We obtain an expression which is valid for arbitrary mirrors, including dissipative ones, characterized by frequency dependent reflection amplitudes. This expressions coincides with the usual Lifshitz expression when the plasma model is used to describe the mirrors material properties, but it differs when the Drude model is applied. The difference can be traced back to the validity of Poisson summation formula [40].

To discuss the effect of finite temperature, we use a theorem which gives the commutators of the intracavity fields as the product of those well known for fields outside the cavity by the Airy function. This theorem was demonstrated with an increasing range of validity in [72, 73, 79]. It is true regardless of whether the mirrors are lossy or not. Since it does not depend on the state of the field, it can be used for thermal as well as vacuum fluctuations. Assuming thermal equilibrium, the theorem leads to the expression of the field anticommutators, i.e. the field fluctuations. Note that thermal equilibrium has to be assumed for the whole system, which means that input fields as well as fluctuations associated with electrons, phonons and any loss mechanism inside the mirrors correspond to the same temperature T , whatever their microscopic origin may be. If parts of the system correspond to different temperatures, completely different results are obtained [94, 95].

The anticommutators of intracavity fields are given by those known for fields outside the cavity multiplied by the Airy function. Hence, the expression written in [73] for a null temperature is only modified through the appearance of a thermal factor in the integrand

$$F = -\hbar \sum_{p,\mathbf{k}} \int_0^\infty \frac{d\omega}{2\pi} \{i\kappa_{\mathbf{k}}[\omega] f_{\mathbf{k}}^p[\omega] c[\omega] + \text{c.c.}\} \quad c[\omega] \equiv \cot h \left(\frac{\pi\omega}{\omega_T} \right), \quad \omega_T \equiv \frac{2\pi k_B T}{\hbar}. \quad (19)$$

Using as before analyticity properties, we transform (19) into an integral over imaginary frequencies giving the following expression for the Casimir force

$$F = \hbar \sum_{p,\mathbf{k}} \int_0^\infty \frac{d\xi}{2\pi} \{ \kappa[i\xi + \eta] f_{\mathbf{k}}^p[i\xi + \eta] c[i\xi + \eta] + \text{c.c.} \}. \quad (20)$$

It is now written as an integral over complex frequencies $\omega = i\xi + \eta$ close to the imaginary axis, with the small positive real number $\eta \rightarrow 0^+$ maintaining the Matsubara poles $\omega_m = im\omega_T$ of $c[\omega]$ outside the contour used to apply the Cauchy theorem. Up to this point, the present derivation is similar to Lifshitz' demonstration [30], while being valid for arbitrary reflection amplitudes. The next steps in Lifshitz' derivation, scrutinized in [40], may raise difficulties for dissipative mirrors. Let us briefly recall the main arguments of [40].

We may first write a series expansion of the Casimir force (19) based upon the expansion of the function $\coth(\pi\omega/\omega_T)$ into a series of exponentials $\exp(-2n\pi i\xi/\omega_T)$ (see also [39] and references therein). This expansion obeys the mathematical criterion of uniform convergence so that, when it is inserted in (19), the order of the summation over n and integration over ξ may be

exchanged. It follows that the force (19) may also be read as

$$F = \frac{\hbar}{\pi} \sum_p \sum_{\mathbf{k}} \sum'_n \tilde{\phi}_{\mathbf{k}}^p \left(\frac{2n\pi}{\omega_T} \right), \quad \tilde{\phi}_{\mathbf{k}}^p(x) \equiv 2 \int_0^\infty d\xi \cos(\xi x) \phi_{\mathbf{k}}^p[\xi], \quad (21)$$

$$\phi_{\mathbf{k}}^p[\xi] \equiv \lim_{\eta \rightarrow 0^+} \kappa_{\mathbf{k}}[i\xi + \eta] f_{\mathbf{k}}^p[i\xi + \eta].$$

We have introduced the common summation convention

$$\sum'_n \varphi(n) \equiv \frac{1}{2} \varphi(0) + \sum_{n=1}^{\infty} \varphi(n). \quad (22)$$

The function $\phi_{\mathbf{k}}^p$ is well defined almost everywhere, the only possible exception being the point $\xi = 0$ where the limit $\eta \rightarrow 0^+$ may be ill defined for mirrors described by dissipative optical models [44]. Since this is a domain of null measure, the cosine Fourier transform $\tilde{\phi}_{\mathbf{k}}^p$ of $\phi_{\mathbf{k}}^p$ is well defined everywhere and the expression (21) of the Casimir force is valid for arbitrary mirrors, including dissipative ones. Note that the term $n = 0$ in (21) corresponds exactly to the contribution of vacuum fluctuations, or to the zero temperature limit, while the terms $n \geq 1$ give the corrections associated with thermal fields.

We come back to the derivation of the Lifshitz formula [30], often used as the standard expression of the Casimir force. This formula is directly related to the decomposition of the coth function into elementary fractions corresponding to the Matsubara poles $\Omega_m = im\omega_T$. If we assume furthermore that the function $\phi_{\mathbf{k}}^p$ is a sufficiently smooth test function, in the sense defined by the theory of distributions, we deduce that the expression (20) can also be read

$$F_{\text{Lif}} = \frac{\hbar\omega_T}{\pi} \sum_p \sum_{\mathbf{k}} \sum'_m \phi_{\mathbf{k}}^p[m\omega_T]. \quad (23)$$

This is the generalization of the Lifshitz' formula [30] to the case of arbitrary reflection amplitudes. It is a discrete sum over Matsubara poles with the primed summation symbol having the definition (22). This formula is known to lead to the correct result in the case of dielectric mirrors (for which it was derived in [30]), for perfect mirrors [37, 38] and also for metallic mirrors described by the lossless plasma model [39].

However, its applicability to arbitrary mirrors remains a matter of controversy [44]. The point is that the derivation of the Lifshitz' formula (23) requires that the function $\phi_{\mathbf{k}}^p$ be a sufficiently smooth test function, in the sense defined by the theory of distributions. Whether or not this is the case at $\xi = 0$ for $\phi_{\mathbf{k}}^p$ calculated from dissipative optical models constitutes the central question of the controversy on the value of the term $p = \text{TE}, m = 0$ in Lifshitz' sum [41]–[45]. Let us repeat that (21) is still a mathematically valid expression of the Casimir force even when $\phi_{\mathbf{k}}^p$ is ill defined for ξ in a domain of null measure. The question of validity of Lifshitz' formula (23) may also be phrased in terms of applicability of the Poisson summation formula [93]. This applicability depends on a smoothness condition which is met for dielectric mirrors, for perfect mirrors and for mirrors described by the plasma model and this explains why Lifshitz' formula (23) may be used as well as (21) in these cases [40].

In order to resolve this controversy, it is crucial to improve our knowledge of the reflection amplitudes at low frequencies. As already discussed, the best way to do that is to measure these

amplitudes on the mirrors used in the experiment at frequencies as low as possible. Although the theoretical question of a good modelling of mirrors at low frequencies is certainly of interest and needs to be answered, the crucial point for a reliable theory–experiment comparison is the necessity of assessing the real behaviour of the mirrors used in the experiments.

3. Non-specular scattering

We will now present a more general formalism to calculate the Casimir force and energy which takes into account non-specular reflection by the plates. Non-specular reflection is of course the generic reflection process on any mirror, while specular reflection is an idealization.

In order to introduce the more general formula, let us first rewrite expression (11) of the Casimir energy between two flat plates as a sum over modes labelled by the ξ and $m \equiv \mathbf{k}, p$

$$E_{\text{sp}} = \hbar \int_0^\infty \frac{d\xi}{2\pi} \text{Tr} \Delta_{\mathbf{k}}^p[i\xi], \quad \Delta_{\mathbf{k}}^p[i\xi] = \ln(1 - r_1 r_2 e^{-2\kappa L}). \quad (24)$$

This can be interpreted as the energy stored inside the cavity during the scattering process. It is expressed in terms of the phase shifts $\Delta_{\mathbf{k}}^p[i\xi]$ acquired by the field modes upon scattering on the cavity. These phase shifts are deduced from the S -matrix of the cavity [72] in such a manner that the Casimir energy is simply equal to the logarithm of the determinant of the S -matrix. Using the techniques of quantum field theory [59], this can also be written as the trace of matrix, here diagonal, defined on these modes

$$E_{\text{sp}} = \hbar \int_0^\infty \frac{d\xi}{2\pi} \sum_m \langle m | \ln(1 - r_1 r_2 e^{-2\kappa_i L}) | m \rangle. \quad (25)$$

Here, r_1 and r_2 are diagonal matrices which contain as their diagonal elements the specular reflection amplitudes, as they are seen from fields inside the cavity,

$$\langle m | r_i | m' \rangle \equiv \delta_{m,m'} r_i(\xi, m), \quad \delta_{m,m'} \equiv 4\pi^2 \delta(\mathbf{k} - \mathbf{k}') \delta_{p,p'}, \quad (26)$$

while κ is a matrix diagonal over the same modes

$$\langle m | \kappa | m' \rangle \equiv \sqrt{\mathbf{k}_m^2 + \xi^2} \delta_{m,m'}. \quad (27)$$

It is now easy to write down a more general formula of the Casimir energy for the case of stationary but non-specular scattering

$$E_{\text{nsp}} = \hbar \int_0^\infty \frac{d\xi}{2\pi} \text{Tr} \ln(1 - \mathcal{R}_1 e^{-\kappa L} \mathcal{R}_2 e^{-\kappa L}). \quad (28)$$

The two matrices \mathcal{R}_1 and \mathcal{R}_2 are no longer diagonal on plane waves since they describe non-specular reflection on the two mirrors. The propagation factors remain diagonal on plane waves. Note that the matrices appearing in (28) no longer commute with each other. As a consequence, the two propagation matrices in (28) can be moved through circular permutations in the product but not adjoined to each other.

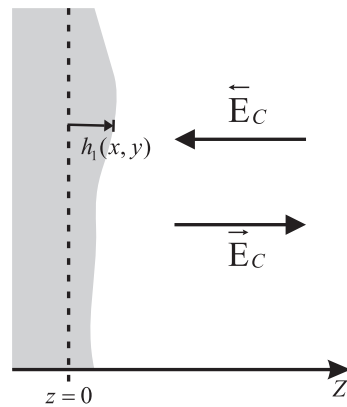


Figure 7. Magnified detail of the internal surface of mirror M1.

Formula (28) has already been used to evaluate the effect of roughness [74, 75] or corrugation [76] of the mirrors on the Casimir force. To this aim, it was expanded at second order in the profiles of the mirrors, with the optical response of the bulk metals described by the plasma model. The non-specular reflection amplitudes were then deduced from techniques developed for treating rough plates [96, 97]. The condition of validity of this expansion is that the roughness or corrugation amplitude is the smallest of length scales involved in the problem. In this regime, it was nevertheless possible to investigate various domains for the roughness or corrugation wavelength and thus to investigate the effect of roughness or corrugation outside as well as inside the range of validity of the PFA.

We may again emphasize at this point that the formula (28) has a wider range of validity than used in those applications. It can in principle describe mirrors with nanostructured surfaces corresponding to large amplitudes which cannot be treated as a small perturbation. It can as well deal with more complicated optical responses which are described neither by a plasma nor by a Drude model. As was extensively discussed above for the case of specular reflection, the formula (28) remains valid for arbitrary mirrors, the only problem being to obtain the precise form of the reflection matrices to be inserted into it.

3.1. Influence of surface roughness

Let us now recall how the non-specular scattering formula (28) can be used to calculate the effect of roughness on the Casimir force. Taking this effect into account simultaneously with that of finite conductivity is essential, because both of them are important at short distances. In order to analyse the roughness effect between two metallic plates, we will describe the optical properties of the mirrors by the plasma model. The values for the plasma wavelength, the mirror separation and the roughness correlation length will be arbitrary with respect to each other, the roughness amplitude remaining the smallest length scale for perturbation theory to hold. We will review some simple analytical expressions for several limiting cases, as well as numerical results allowing one for a reliable calculation of the roughness correction in real experiments [74, 75].

In a PP geometry, the surface profiles are defined by the functions $h_i(x, y)$ ($i = 1, 2$) giving the local heights with respect to the mean separation L along the z -direction as shown in figure 7. These functions are defined so that they have zero averages. We consider the case of stochastic

roughness characterized by spectra

$$\sigma_{ij}(\mathbf{k}) = \int d^2\mathbf{r} e^{-i\mathbf{k}\cdot\mathbf{r}} \langle h_i(\mathbf{r}) h_j(\mathbf{0}) \rangle, \quad i = 1, 2. \quad (29)$$

We suppose the surface A of the plates to contain many correlation areas, which allows us to take ensemble or surface averages interchangeably. The two plates are considered to be made of the same metal and the crossed correlation between their profiles is neglected ($\sigma_{12}(\mathbf{k}) = 0$).

We obtain the following variation of the Casimir energy E_{PP} up to second order in the perturbations h_i [71]

$$\delta E_{\text{PP}} = \int \frac{d^2\mathbf{k}}{4\pi^2} G_r(\mathbf{k}) \sigma(\mathbf{k}), \quad \sigma(\mathbf{k}) = \sigma_{11}(\mathbf{k}) + \sigma_{22}(\mathbf{k}). \quad (30)$$

With our assumptions, the spectrum $\sigma(\mathbf{k})$ fully characterizes the roughness of the two plates. The correlation length ℓ_C is defined as the inverse of its width. The response function $G_r(\mathbf{k})$ then describes the spectral sensitivity to roughness of the Casimir effect. Symmetry requires that it only depends on $k = |\mathbf{k}|$. The dependance of G_r on k reflects that not only the roughness amplitude but also its spectrum plays a role in diffraction on rough surfaces [96, 97]. The formula (30) has been obtained for the energy in the PP configuration but it also determines the force correction δF_{PS} in the PS configuration since the PFA is still used for describing the weak curvature of the sphere (see below).

We now focus our attention on the validity of PFA for treating the effect of roughness and notice that this validity only holds at the limit of smooth surface profiles $k \rightarrow 0$. In fact, the following identity is obeyed by our result [74], for arbitrary values of L and λ_P ,

$$G_r(k \rightarrow 0) = \frac{E''_{\text{PP}}(L)}{2}, \quad (31)$$

where the derivative is taken with respect to the plate separation L . If we now suppose that the roughness spectrum $\sigma(k)$ is included inside the PFA sector where $G_r(k) \simeq G_r(0)$, G_r may be replaced by $G_r(0)$ and factored out of the integral (30) thus leading to the PFA expression [71]

$$\delta E_{\text{PP}} = \frac{E''_{\text{PP}}(L)}{2} a^2, \quad a^2 = \int \frac{d^2\mathbf{k}}{4\pi^2} \sigma(\mathbf{k}) \equiv \langle h_1^2 + h_2^2 \rangle. \quad (32)$$

In this PFA limit, the correction depends only on the variance a^2 of the roughness profiles, that is also the integral of the roughness spectrum.

In the general case in contrast, the sensitivity to roughness depends on the wavevector k . This key point is emphasized by introducing a new function $\rho_r(k)$ which measures the deviation from the PFA [71]

$$\rho_r(k) = \frac{G_r(k)}{G_r(0)}. \quad (33)$$

This function is plotted on figure 8 for several values of L . As for all numerical examples considered below, we take $\lambda_P = 137$ nm which corresponds to gold covered plates.

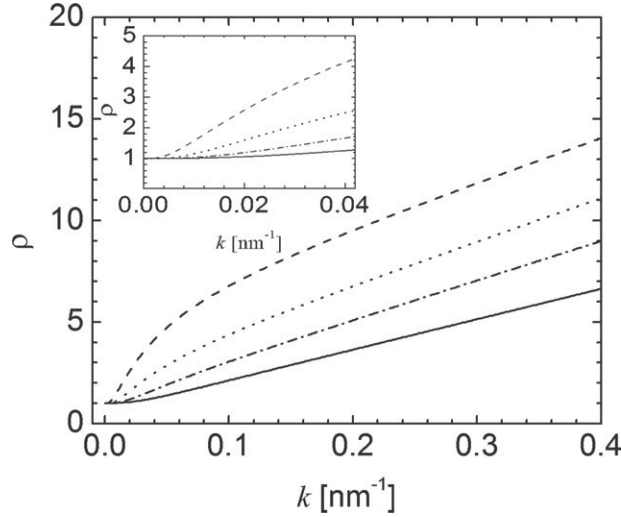


Figure 8. Variation of ρ_r versus k for $L = 50, 100, 200, 400$ nm (from bottom to top curve).

The ratio $\rho_r(k)$ is almost everywhere larger than unity, which means that the PFA systematically *underestimates* the roughness correction. The inset shows $\rho_r(k)$ for small values of k where the PFA is a good approximation. To give a number illustrating the deviation from the PFA, we find $\rho_r \simeq 1.6$ for $L = 200$ nm and $k = 0.02$ nm⁻¹, which means that the exact correction is 60% larger than the PFA result for this intermediate separation and a typical roughness wavelength $2\pi/k \simeq 300$ nm.

Figure 8 indicates that $\rho_r(k)$ grows linearly for large values of k . This is a general prediction of our full calculations [74], for arbitrary values of L and λ_p ,

$$\rho_r(k) = \alpha_r k \quad \text{for } k \gg \frac{2\pi}{\lambda_p}, \frac{1}{L}. \quad (34)$$

The dimensionless parameter α_r/L depends on $2\pi L/\lambda_p$ only, and its expression is given by equation (8) in [74]. In figure 9, we plot the coefficient α_r as a function of L with $\lambda_p = 137$ nm. In the limit of short distances, we recover the expression which was drawn in [71] from older calculations [64] (after a correction by a global factor 2)

$$\alpha_r = 0.4492L \quad \text{for } k^{-1} \ll L \ll \frac{\lambda_p}{2\pi}. \quad (35)$$

In the opposite limit of large distances, the coefficient α_r is found to saturate [74]

$$\alpha_r = \frac{14}{15} \frac{\lambda_p}{2\pi} \quad \text{for } k^{-1} \ll \frac{\lambda_p}{2\pi} \ll L. \quad (36)$$

It is interesting to note that this result differs from the long distance behaviour which was drawn in [71] from the reanalysis of calculations of the effect of sinusoidal corrugations on perfectly reflecting plates [66]. Perfect reflectors indeed correspond to the limiting case where

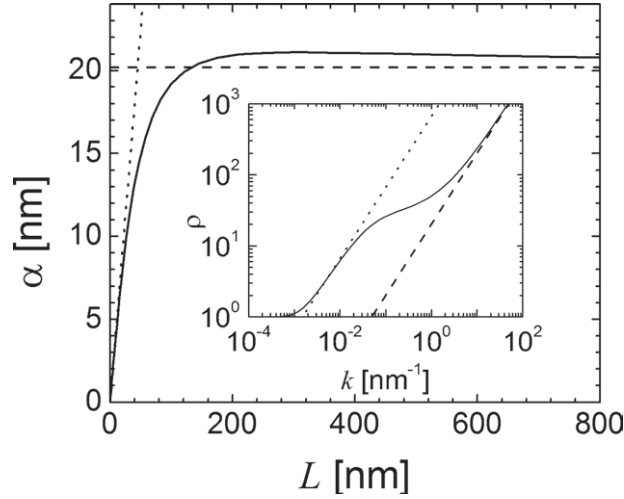


Figure 9. Variation of the coefficient α_r versus L . The analytical result for $k^{-1} \ll L \ll \lambda_P$ is shown as the dotted line and for $k^{-1} \ll \lambda_P \ll L$ as the dashed line. A comparison between this second result (dashed straight line) and the exact $\rho_r(k)$ (solid line) is shown in the inset for $L = 2 \mu\text{m}$. The analytical result $\rho_r = Lk/3$ predicted by the model of perfect reflectors (dotted line) is valid only in the intermediate range $\lambda_P \ll k^{-1} \ll L$.

λ_P rather than $1/k$ is the shortest length scale. The following result is obtained in this case [74], which effectively fits that of [66],

$$\rho_r = \frac{1}{3}Lk \quad \text{for } \lambda_P \ll k^{-1} \ll L. \quad (37)$$

The long-distance behaviour is thus given by (36) when $1/k \ll \lambda_P \ll L$ but by (37) when $\lambda_P \ll 1/k \ll L$. The cross-over between these two regimes is shown in the inset of figure 9, where we plot ρ_r as a function of k for $L = 2 \mu\text{m}$. The failure of the perfect reflection model for $1/k \ll \lambda_P$ has been given an interpretation in [74]: it results from the fact that not only the incoming field mode but also the outgoing one have to see the mirror as perfectly reflecting for formula (37) to be valid.

These numerical results can be used to assess the accuracy of the PFA applied to the problem of roughness. PFA is indeed recovered at the limit of very smooth surface profiles and the deviation from PFA given by our results as soon as the roughness wavevector goes out of this limit. The mirrors used in a given experiment have a specific roughness spectrum which can, and in our opinion must be, measured when the experiments are performed. The integral (30) then leads to a reliable prediction for the roughness correction, as soon as the spectral sensitivity $G_r(k)$ and the real spectrum $\sigma(k)$ are inserted into it.

3.2. Lateral Casimir force component

The spectral sensitivity $G_r(k)$ involved in the calculation of the roughness correction can be considered as a further prediction of quantum electrodynamics, besides the more commonly studied mean Casimir force, so that the comparison of its theoretical expectation with experiments

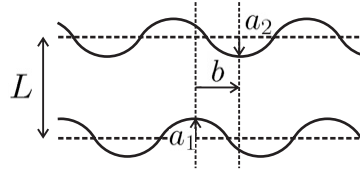


Figure 10. Surface profiles considered for the lateral component of the Casimir force. Both surfaces have a sinusoidal corrugation with a_1 and a_2 being the corrugation amplitudes, b the mismatch between the two sinusoidal functions.

is an interesting prospect. But this comparison can hardly rely on the roughness correction (30) which remains in any case a small variation of the longitudinal Casimir effect. A more stringent test can be performed by studying the lateral component of the Casimir force which arises between corrugated surfaces. This lateral Casimir force would indeed vanish in the absence of surface corrugation so that the expression of the spectral sensitivity will thus appear directly as a factor in front of the lateral Casimir force. For reasons which will become clear below, the spectral sensitivity involved in the calculation of corrugation effect is a different function $G_C(k)$.

Nice experiments have shown the lateral Casimir force to be measurable at separations of a few hundred nanometres [61], that is of the same order of magnitude as the plasma wavelength λ_P . It follows that these experiments can neither be analysed by assuming the mirrors to be perfect reflectors [67], nor by using the opposite limit of plasmon interaction [64]. It is no more possible to use the PFA if we want to be able to treat arbitrary values of the ratio of the corrugation wavelength λ_C to the interplate distance L . This is why we emphasize the results drawn from the non-specular scattering formula (28) which can be used for calculating the lateral Casimir force for arbitrary relative values of λ_P , λ_C and L . The only drawback of this calculation is that it is restricted to small enough corrugation amplitudes, since the latter have to remain the smallest length scale for perturbation theory to hold. But the lateral force is known to be experimentally accessible in this regime. Again we model the optical response of the metallic plates by the plasma model.

The surface profiles of the corrugated plates are defined by two functions $h_i(\mathbf{r})$, with $\mathbf{r} = (x, y)$ is the lateral position along the surfaces of the plates, while $i = 1, 2$ labels the two plates. As in experiments [61], we consider the simple case of uniaxial sinusoidal corrugations imprinted on the two plates (see figure 10) along the same direction, say the y -direction, and with the same wavevector $k \equiv 2\pi/\lambda_C$

$$h_1 = a_1 \cos(kx), \quad h_2 = a_2 \cos[k(x + b)]. \quad (38)$$

Both profiles h_1 and h_2 have zero spatial averages and they are counted as positive when they correspond to local length decreases below the mean value L .

For the purpose of the calculation of the lateral Casimir force, the non-specular reflection matrix \mathcal{R}_j have to be developed up to the first order in the deviations h_j from flatness of the two plates. They are thus written as the sum of a zeroth-order contribution identifying with the specular reflection amplitude and of a first-order contribution proportional to the Fourier component at wavevector $(\mathbf{k} - \mathbf{k}')$ of the surface profiles, this Fourier component being able to induce a scattering of the field modes from the wavevector \mathbf{k} to \mathbf{k}' [76]. The correction of the Casimir energy δE_{PP} induced by the corrugations arises at second order in the corrugations, with crossed terms of the form $a_1 a_2$ which have the ability to induce lateral forces. In other words,

the corrugation sensitivity function $G_C(k)$ obtained below depends on the crossed correlation between the profiles of the two plates, in contrast to the function $G_r(k)$ calculated above for describing the roughness spectral sensitivity. The latter were depending on terms quadratic in h_1 or h_2 , and their evaluation required that second-order non-specular scattering be properly taken into account. Here, first-order non-specular amplitudes evaluated on both plates are sufficient.

The result of the calculation is read as a second-order correction induced by corrugations

$$\delta E_{PP} = 2 \int \frac{d^2 \mathbf{k}}{(2\pi)^2 A} G_C(\mathbf{k}) H_1(\mathbf{k}) H_2(-\mathbf{k}), \quad (39)$$

with the function $G_C(\mathbf{k})$ given by equation (3) in [76]. For isotropic media, symmetry requires $G_C(\mathbf{k})$ to depend only on the modulus of the wavevector $k = |\mathbf{k}|$. We may also assume for simplicity that the two plates are made of the same metallic medium. The energy correction thus depends on the lateral mismatch b between the corrugations of the two plates, which is the cause for the lateral force to arise. Replacing the ill-defined $(2\pi)^2 \delta^{(2)}(0)$ by the area A of the plates, (compare with (31))

$$\delta E_{PP} = a_1 a_2 \cos(kb) G_C(k). \quad (40)$$

Once again, the result of the PFA is recovered from equation (40) as the limiting case $k \rightarrow 0$, that is also for long corrugation wavelengths. This corresponds to nearly plane surfaces where the Casimir energy can be obtained from the energy E_{PP} calculated between perfectly plane plates by averaging the ‘local’ distance $\mathcal{L} = L - h_1 - h_2$ over the surface of the plates. Expanding at second order in the corrugation amplitudes and disregarding squared terms in a_1^2 and a_2^2 because they cannot produce a lateral dependence, we thus recover expression (40) with $G_C(k)$ replaced, for small values of k or equivalently large values of λ_C , by $G_C(0)$ given by (compare with 31)

$$G_C(k \rightarrow 0) = \frac{E''_{PP}}{2}. \quad (41)$$

This property is ensured, for any specific model of the material medium, by the fact that G_C is given for $k \rightarrow 0$ by the specular limit of non-specular reflection amplitudes [76]. For arbitrary values of k , the deviation from the PFA is then described by the ratio

$$\rho_C(k) = \frac{G_C(k)}{G_C(0)}. \quad (42)$$

In the following, we discuss explicit expressions of this ratio ρ_C given by its general expression (equation (3) in [76]). For the numerical examples, we take $\lambda_P = 137$ nm, corresponding to gold covered plates. The result ρ_C is plotted in figure 11 as a function of k , for different values of the distance L .

For example for a distance $L = 50$ nm, the PFA is correct in the range $k \leq 0.01$ nm⁻¹ (i.e. $\lambda_C \geq 628$ nm) covered by the plot in figure 11. However, for typical separations of 100 nm or larger, ρ_C drops significantly below its PFA value of unity. A more detailed discussion can be found in [76].

For still larger values of kL , the functions $G_C(k)$ and $\rho_C(k)$ decay exponentially to zero. If we also assume that $k\lambda_P \gg 1$, we find $G_C(k) = \alpha_C k \exp(-kL)$ where the parameter α_C now depends on λ_P and L only. This is in striking contrast with the behaviour of the response function for

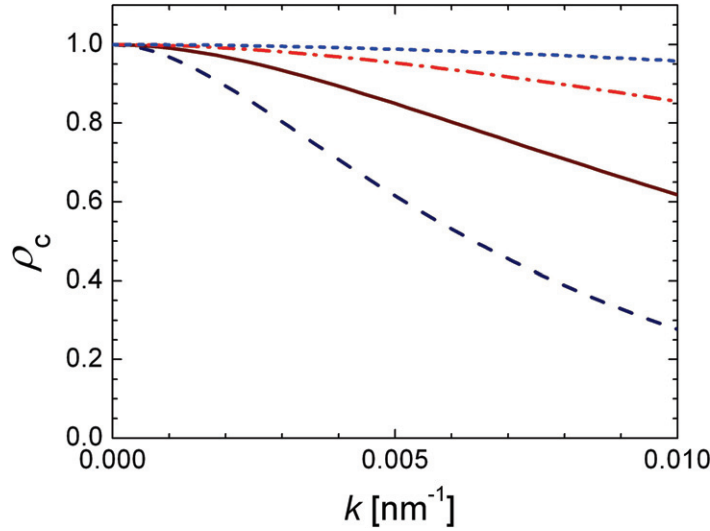


Figure 11. Variation of ρ_C versus k with $\lambda_P = 137$ nm and for $L = 50$ nm (dotted line), 100 nm (dash-dotted line), 200 nm (solid line) and 400 nm (dashed line).

stochastic roughness, which *grows* linearly with k for large k due to the contribution of the second-order reflection coefficients [75]. These coefficients do not contribute to the second-order lateral effect, which is related to two first-order non-specular reflections at different plates, separated by a one-way propagation with a modified momentum of the order of k . The resulting propagation factor is, in the large- k limit, $\exp(-\kappa L) \approx \exp(-kL)$, thus explaining the exponential behaviour.

3.3. Comparison to experiments in a PS configuration

In order to compare the theoretical expression of the lateral Casimir force to experiments, we have to consider the PS geometry [61] rather than the PP one. As $R \gg L$, we use the PFA to connect the two geometries. Any interplay between curvature and corrugation is avoided provided that $RL \gg \lambda_C^2$. These two conditions are met in the experiment reported in [61], where $R = 100 \mu\text{m}$, $\lambda_C = 1.2 \mu\text{m}$ and $L \sim 200$ nm.

We thus obtain the energy correction δE_{PS} between the sphere and a plane at a distance of closest approach L as an integral of the energy correction δE_{PP} in the PP geometry

$$\delta E_{PS}(L, b) = \int_{\infty}^L \frac{2\pi R dL'}{A} \delta E_{PP}(L', b). \quad (43)$$

Then the lateral force is deduced by varying the energy correction (43) with respect to the lateral mismatch b between the two corrugations. Simple manipulations then lead to the lateral Casimir force in the PS geometry

$$F_{PS}^{\text{lat}} = \frac{2\pi a_1 a_2}{A} k R \sin(kb) \int_{\infty}^L dL' G_C(k, L') \quad (44)$$

The force attains a maximal amplitude for $\sin(kb) = \pm 1$, which is easily evaluated in the PFA regime $k \rightarrow 0$ where $G_C(k)$ does not depend on k , so that F_{PS}^{lat} scales as k . As k increases, the

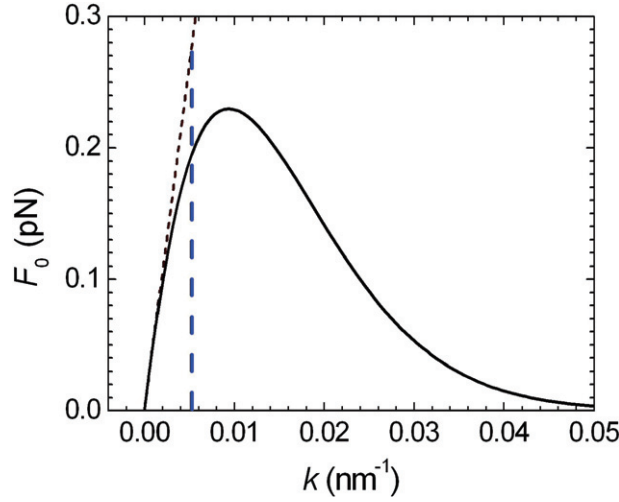


Figure 12. Lateral force amplitude for the PS setup, as a function of k , with figures taken from [61]. The experimental value $k = 0.0052 \text{ nm}^{-1}$ is indicated by the vertical dashed line.

amplitude increases at a slower rate and then starts to decrease due to the exponential decay of $G_C(k)$. For a given value of the separation L , the lateral force reaches an optimum for a corrugation wavelength such that kL is of order of unity, which generalizes the result obtained for perfect reflectors in [66]. In figure 12, we plot the force $F_{\text{PS}}^{\text{lat}}$ (for $\sin(kb) = 1$) as a function of k , with figures taken from the experiment of [61]. We also use the values $a_1 = 59 \text{ nm}$ and $a_2 = 8 \text{ nm}$ of the amplitudes for measuring the force as in [61], recalling however that our calculations are valid in the perturbative limit $a_1, a_2 \rightarrow 0$.

The plot clearly shows the linear growth for small k as well as the exponential decay for large k . The maximum force is at $k = 0.009 \text{ nm}^{-1}$ so that $kL \simeq 2$. The experimental value $k = 0.0052 \text{ nm}^{-1}$ is indicated by the dashed line in figure 12, and the force obtained as 0.20 pN , well below the PFA result, indicated by the straight line and corresponding to a force of 0.28 pN .

Such a variation in the lateral Casimir force should in principle be measurable in an experiment. This could lead to the first unambiguous evidence of the limited validity of the PFA, that is also to the first observation of a non trivial effect of geometry on the Casimir force.

4. Conclusion

In this paper, we have described the theory of the Casimir effect using the techniques of scattering theory. We have recalled how this formalism allows one to take into account the real conditions under which Casimir force measurements are performed.

In particular, the finite conductivity effect can be treated in a very precise manner, which is a necessity for a reliable theory-experiment comparison. There however, remain inaccuracies in this comparison if the reflection amplitudes are drawn from optical models, because of the intrinsic dispersion of optical properties of samples fabricated by different techniques. We have emphasized that these inaccuracies could be circumvented by measuring these reflection amplitudes rather than modelling them.

We have then presented the scattering formulation of the Casimir force at nonzero temperature. This formulation clears out the doubt on the expression of the force, while again requiring to have at one's disposal reflection amplitudes representing the real properties of the mirrors used in the experiments. Let us at this point emphasize that the effect of temperature has not been unambiguously proven in experiments, and that its observation is one of the most urgent challenges of experimental research in the domain. An interesting possibility would be to perform accurate measurements of the force at distances larger than a few micrometres, for example by using torsional balances [98].

In the second part of the paper, we have presented a more general scattering formalism which takes into account non-specular reflection. We have also discussed the application of this formalism for the calculation of the roughness correction to the longitudinal Casimir force as well as of the lateral component of the Casimir force arising between corrugated surfaces. We have argued that the spectral sensitivity functions appearing in these expressions have to be considered as a new prediction of quantum electrodynamics, which differ from the more commonly studied mean Casimir force as soon as one goes out of the domain of validity of the PFA. This new test seems to be experimentally feasible and constitutes another challenge of great interest to be faced in the near future.

Acknowledgments

PAMN thanks R Rodrigues for discussions and CNPq and Instituto do Milenio de Informação Quântica for partial financial support. AL and SR acknowledge fruitful discussions with M T Jaekel and C Genet. AL acknowledges partial financial support by the European Contract STRP 12142 NANOCASE.

References

- [1] Casimir H B G 1948 *Proc. K. Ned. Akad. Wet.* **51** 793
- [2] Sparnaay M J 1989 *Physics in the Making* ed A Sarlemijn and M J Sparnaay (Amsterdam: North-Holland) 235
- [3] Milonni P W 1994 *The Quantum Vacuum* (New York: Academic)
- [4] Mostepanenko V M and Trunov N N 1997 *The Casimir Effect and its Applications* (Oxford: Clarendon)
- [5] Lamoreaux S K 1999 *Resource Letter Am. J. Phys.* **67** 850
- [6] Lamoreaux S K L 1997 *Phys. Rev. Lett.* **78** 5
- [7] Mohideen U and Roy A 1998 *Phys. Rev. Lett.* **81** 4549
- [8] Harris B W, Chen F and Mohideen U 2000 *Phys. Rev. A* **62** 052109
- [9] Ederth Th 2000 *Phys. Rev. A* **62** 062104
- [10] Bressi G, Carugno G, Onofrio R and Ruoso G 2002 *Phys. Rev. Lett.* **88** 041804
- [11] Decca R S, López D, Fischbach E and Krause D E 2003 *Phys. Rev. Lett.* **91** 050402
- [12] Decca R S, López D, Fischbach E, Klimchitskaya G L, Krause D E and Mostepanenko V M 2005 *Ann. Phys.* **318** 37
- [13] Bordag M, Mohideen U and Mostepanenko V M 2001 *Phys. Rep.* **353** 1
- [14] Lambrecht A and Reynaud S 2002 *Poincaré Semin. Vacuum Energy Renormalization* **1** 107 (Preprint quant-ph/0302073)
- [15] Milton K A 2005 *J. Phys. A* **20** 4628
- [16] Reynaud S, Lambrecht A, Genet C and Jaekel M T 2001 *C. R. Acad. Sci. Paris* **IV-2** 1287 (Preprint quant-ph/0105053)

- [17] Genet C, Lambrecht A and Reynaud S 2002 *On the Nature of Dark Energy* ed U Brax, J Martin and J P Uzan, 121 (Frontier Group) (*Preprint* [quant-ph/0210173](#))
- [18] Fischbach E and Talmadge C 1998 *The Search for Non Newtonian Gravity* (New York: Springer/AIP Press)
- [19] Hoyle C D, Schmidt U, Heckel B R, Adelberger E G, Grundlach J H, Kapner D J and Swanson H E 2001 *Phys. Rev. Lett.* **86** 1418
- [20] Adelberger E G 2002 *Proc. Second Meeting on CPT and Lorentz Symmetry* ed V A Kostelecky (Singapore: World Scientific) also in (*Preprint* [hep-ex/0202008](#))
- [21] Long J C *et al* 2003 *Nature* **421** 922
- [22] Carugno G, Fontana Z, Onofrio R and Rizzo C 1997 *Phys. Rev. D* **55** 6591
- [23] Bordag M, Geyer B, Klimchitskaya G L and Mostepanenko V M 1999 *Phys. Rev. D* **60** 055004
- [24] Fischbach E and Krause D E 1999 *Phys. Rev. Lett.* **82** 4753
- [25] Long J C, Chan H W and Price J C 1999 *Nucl. Phys. B* **539** 23
- [26] Fischbach E, Krause D E, Mostepanenko V M and Novello M 2001 *Phys. Rev. D* **64** 075010
- [27] Decca R S, Fischbach E, Klimchitskaya G L, Krause D E, Lopez D L and Mostepanenko V M 2003 *Phys. Rev. D* **68** 116003
- [28] Buks E and Roukes M L 2001 *Phys. Rev. B* **63** 033402
- [29] Chan H B, Aksyuk V A, Kleiman R N, Bishop D J and Capasso F 2001 *Science* **291** 1941
Chan H B, Aksyuk V A, Kleiman R N, Bishop D J and Capasso F 2001 *Phys. Rev. Lett.* **87** 211801
- [30] Lifshitz E M 1956 *Sov. Phys.-JETP* **2** 73
- [31] Heinrichs J 1975 *Phys. Rev. B* **11** 3625
- [32] Schwinger J, de Raad L L and Milton K A 1978 *Ann. Phys.* **115** 1
- [33] Lamoreaux S K 1999 *Phys. Rev. A* **59** R3149
- [34] Lambrecht A and Reynaud S 2000 *Euro. Phys. J. D* **8** 309
- [35] Klimchitskaya G L, Mohideen U and Mostepanenko V M 2000 *Phys. Rev. A* **61** 062107
- [36] Bezerra V B, Klimchitskaya G L and Mostepanenko V M 2000 *Phys. Rev. A* **62** 014102
- [37] Mehra J 1967 *Physica* **57** 147
- [38] Brown L S and Maclay G J 1969 *Phys. Rev.* **184** 1272
- [39] Genet C, Lambrecht A and Reynaud S 2000 *Phys. Rev. A* **62** 012110
- [40] Reynaud S, Lambrecht A and Genet C 2004 *Quantum Field Theory Under the Influence of External Conditions* ed K A Milton (Princeton, NJ: Rinton) p 36 (*Preprint* [quant-ph/0312224](#))
- [41] Boström M and Sernelius Bo E 2000 *Phys. Rev. Lett.* **84** 4757
- [42] Svetovoy V B and Lokhanin M V 2000 *Mod. Phys. Lett. A* **15** 1013, 1437
- [43] Bordag M, Geyer B, Klimchitskaya G L and Mostepanenko V M 2000 *Phys. Rev. Lett.* **85** 503
- [44] Klimchitskaya G L and Mostepanenko V M 2001 *Phys. Rev. A* **63** 062108
Klimchitskaya G L 2002 *Int. J. Mod. Phys. A* **17** 751
- [45] Hoyer J S, Brevik I, Aarseth J B and Milton K A 2003 *Phys. Rev. E* **67** 056116
- [46] Torgerson J R and Lamoreaux S K 2004 *Phys. Rev. E* **70** 047102
- [47] Sernelius B E 2005 *Phys. Rev. B* **71** 235114
- [48] Esquivel R and Svetovoy V B 2004 *Phys. Rev. A* **69** 062102
- [49] Brevik I, Aarseth J B, Hoyer J S and Milton K A 2005 *Phys. Rev. E* **71** 056101
- [50] Hoyer J S, Brevik I, Aarseth J B and Milton K A 2005 *J. Phys. A* (*Preprint* [quant-ph/0506025](#))
- [51] Mostepanenko V M, Bezerra V B, Decca R S, Geyer B, Fischbach E, Klimchitskaya G L, Krause D E, Lopez D and Romero C 2005 *J. Phys. A* to appear (*Preprint* [quant-ph/0512134](#))
- [52] Brevik I, Ellingsen S A and Milton K 2006 *Preprint* [quant-ph/0605005](#)
- [53] Deriagin B V, Abrikosova I I and Lifshitz E M 1968 *Q. Rev.* **10** 295
- [54] Langbein D 1971 *J. Phys. Chem. Solids* **32** 1657
- [55] Kiefer J E *et al* 1978 *J. Colloid Interface Sci.* **67** 140
- [56] Schröder O, Sardicchio A and Jaffe R L 2005 *Phys. Rev. A* **72** 012105
- [57] Gies H and Klingmüller K 2006 *Preprint* [quant-ph/0601094](#)

- [58] Balian R and Duplantier B 1978 *Ann. Phys.* **112** 165
- [59] Plunien G, Muller B and Greiner W 1986 *Phys. Rep.* **134** 87
- [60] Balian R and Duplantier B 2004 *Preprint* [quant-ph/0408124](#)
- [61] Chen F *et al* 2002 *Phys. Rev. Lett.* **88** 101801
Chen F *et al* 2002 *Phys. Rev. A* **66** 032113
- [62] van Bree J, Poulis J, Verhaar B and Schram K 1974 *Physica* **78** 187
- [63] Maradudin A A and Mills D L 1975 *Phys. Rev. B* **11** 1392
- [64] Maradudin A A and Mazur P 1980 *Phys. Rev. B* **22** 1677
Mazur P and Maradudin A A 1981 *Phys. Rev. B* **23** 695
- [65] Nieto-Vesperinas M 1982 *J. Opt. Soc. Am. A* **72** 538
- [66] Emig T, Hanke A, Golestanian R and Kardar M 2001 *Phys. Rev. Lett.* **87** 260402
Emig T, Hanke A, Golestanian R and Kardar M 2003 *Phys. Rev. A* **67** 022114
- [67] Emig T 2003 *Europhys. Lett.* **62** 466
- [68] Bordag M, Klimchitskaya G L and Mostepanenko V M 1995 *Phys. Lett. A* **200** 95
- [69] Klimchitskaya G M, Roy A, Mohideen U and Mostepanenko V M 1999 *Phys. Rev. A* **60** 3487
- [70] Blagov E V *et al* 2004 *Phys. Rev. A* **69** 044103
- [71] Genet C, Lambrecht A, Maia Neto P A and Reynaud S 2003 *Europhys. Lett.* **62** 484
- [72] Jaekel M T and Reynaud S 1991 *J. Physique* **I-1** 1395 (*Preprint* [quant-ph/0101067](#))
- [73] Genet C, Lambrecht A and Reynaud S 2003 *Phys. Rev. A* **67** 043811
- [74] Maia Neto P A, Lambrecht A and Reynaud S 2005 *Europhys. Lett.* **69** 924
- [75] Maia Neto P A, Lambrecht A and Reynaud S 2005 *Phys. Rev. A* **72** 012115
- [76] Rodrigues R B, Maia Neto P A, Lambrecht A and Reynaud S 2006 *Phys. Rev. Lett.* **96** 100402
- [77] Barnett S M, Gilson C R, Huttner B and Imoto N 1996 *Phys. Rev. Lett.* **77** 1739
- [78] Courty J M, Grassia F and Reynaud S 2000 *Noise, Oscillators and Algebraic Randomness* ed M Planat, 71 (Berlin: Springer) (*Preprint* [quant-ph/0110021](#))
- [79] Landau L and Lifshitz E M 1980 *Landau and Lifshitz Course of Theoretical Physics: Electrodynamics in Continuous Media* ch X (Oxford: Butterworth-Heinemann)
- [80] Kats E I 1977 *JETP* **46** 109
- [81] Lambrecht A, Jaekel M T and Reynaud S 1997 *Phys. Lett. A* **225** 188
- [82] Genet C, Intravaia F, Lambrecht A and Reynaud S 2004 *Ann. Fond. L. de Broglie* **29** 311 (*Preprint* [quant-ph/0302072](#))
- [83] Henkel C, Joulain K, Mulet J Ph and Greffet J J 2004 *Phys. Rev. A* **69** 023808
- [84] Van Kampen N G, Nijboer B R A and Schram K 1968 *Phys. Lett. A* **26** 307
- [85] Intravaia F and Lambrecht A 2005 *Phys. Rev. Lett.* **94** 110404
- [86] Ashcroft N W and Mermin N D 1976 *Solid State Physics* (Philadelphia, PA: HRW International)
- [87] Further details on optical data are available; please contact Online at astrid.lambrecht@spectro.jussieu.fr
- [88] *Handbook of Optical Constants of Solids* Palik E D (ed) 1995 (New York: Academic)
- [89] *Handbook of Optics II* 1995 (New York: McGraw-Hill)
- [90] Lide D R (ed) 1998 *CRC Handbook of Chemistry and Physics* 79th edn. (Boca Raton, FL: CRC Press)
- [91] Pirozhenko I, Lambrecht A and Svetovoy V B 2006 *New J. Phys.* **8** 238
- [92] Barnett S M, Jeffers J, Gatti A and Loudon R 1998 *Phys. Rev. A* **57** 2134
- [93] Antezza M, Pitaevskii L P and Stringari S 2005 *Phys. Rev. Lett.* **95** 113202
- [94] Henkel C, Joulain K, Mulet J P and Greffet J-J 2002 *J. Opt. A* **4** S109
- [95] Morse P M and Feshbach H 1953 *Methods of Theoretical Physics* (New York: McGraw-Hill) part I ch. 4.8
- [96] Agarwal G S 1977 *Phys. Rev. B* **15** 2371
- [97] Greffet J-J 1988 *Phys. Rev. B* **37** 6436
- [98] Lambrecht A, Nesvizhevsky V, Onofrio R and Reynaud S 2005 *Class. Quantum Grav.* **22** 5397