

# Calculating limits on Yukawa Interactions

Alexander Rider

May 6, 2016

## Potential Integral

A general Yukawa interaction leads to the interaction between two point masses given by

$$V(r) = \frac{-Gm_1m_2}{r}(1 + \alpha e^{-r/\lambda}) \quad (1)$$

where  $G$  is the gravitational constant,  $m_1$  and  $m_2$  are the respective point masses,  $\alpha$  is the dimensionless strength of the interaction, and  $\lambda$  is the length scale of the interaction. The potential due to the Yukawa interaction between a sphere of density  $\rho$  and point mass  $m$  can be written as the integral

$$V(s, r_b, \rho) = m \int_s^{s+2r_b} \int_0^\pi \int_0^{2\pi} \rho V(r) r^2 \sin(\theta) d\phi d\theta dr \quad (2)$$

The geometry of this setup is shown in figure 1. Where  $s$  is the separation between the point mass and the closest point on the surface of the sphere, and  $r_b$  is the radius of the sphere. Symmetry can be used to perform the integration over the  $\theta$  and  $\phi$  coordinates and sum up spherical caps weighted by  $V(r)$ . The

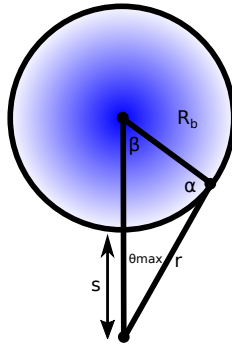


Figure 1: The geometry for integrating a function that is constant of radius between a point and a sphere.

area of the spherical cap intersecting the sphere at a distance  $r$  from the point mass is given by

$$A(r) = r^2 \int_0^{\theta_{max}} \int_0^{2\pi} \sin(\theta) d\phi d\theta = 2\pi r^2 (1 - \cos(\theta_{max})) \quad (3)$$

where  $\theta_{max}$  is the maximum azimuthal angle at which the spherical cap centered on the point mass intersects the sphere.  $\theta_{max}$  can be calculated from the law of cosines as

$$\theta_{max} = \text{Arccos}\left(\frac{(s + r_b)^2 + r^2 - r_b^2}{2r(s + r_b)}\right) \quad (4)$$

Inserting this into 3 we obtain

$$A(r) = \pi r \frac{(r - s)(s + 2r_b - r)}{s + r_b} \quad (5)$$

So we can write the total potential between a point mass and a sphere as

$$V(s, r_b, \rho) = \int_s^{s+r_b} V(r) A(r) dr \quad (6)$$

To integrate over the potential given in 1 we note that

$$\int r^n e^r dr = \int [\partial_\beta^n e^{r\beta}]_{\beta=1} dr = [\partial_\beta^n \int e^{\beta r} dr]_{\beta=1} \quad (7)$$

For the relevant cases of  $n = 1$  and  $n = 3$  we have that

$$\int r e^r dr = (r - 1) e^r \quad (8)$$

$$\int r^3 e^r dr = (r^3 - 3r^2 + 6r - 6) e^r \quad (9)$$

Now we can write

$$V(s, r_b, \rho, \lambda, \alpha, m) = -\frac{Gm\rho}{r_b + s} (v(r_b) + 3/2\alpha e^{-s/\lambda} v(\lambda) f(r_b/\lambda)) \quad (10)$$

where

$$v(r) = 4/3\pi r^3 \quad (11)$$

and

$$f(r_b/\lambda) = e^{-2r_b/\lambda} (1 + r_b/\lambda) + r_b/\lambda - 1 \quad (12)$$

We can now examine limiting cases to investigate the validity of this result. In the limit  $\alpha \rightarrow 0$  we see that  $V \rightarrow \frac{-Gm_1 m_2}{r}$  and standard Newtonian gravity is obtained. If we take the limit  $\alpha \gg \bar{G}$  and  $\frac{r_b}{\lambda} \ll 1$   $f(r/\lambda) \rightarrow 2/3(r_b/\lambda)^3$  so  $3/2v(\lambda)f(r_b/\lambda) \rightarrow v(r_b)$  and we get

$$V(s, r_b, \rho, \lambda, \alpha, m) \approx \alpha \frac{m\rho v(r_b) e^{-s/\lambda}}{r_b + s} \quad (13)$$

The sphere effectively interacts with the Yukawa potential as a point mass in this limit. The potential in (10) can be integrated over an attractor in 3 dimensions to efficiently calculate the interaction between a microsphere and an attractor.

## Nuclear Couplings

Yukawa violations of the inverse square law between two point masses of mass  $M_a$  and  $M_b$  are generally parameterized with the potential [1]

$$V_{ab}(r) = -\alpha G \frac{M_a M_b}{r} \exp(-r/\lambda) \quad (14)$$

For the special case of exchange of scalar or vector bosons  $\phi$  with mass  $m$  between two non-relativistic fermions generically produces a potential

$$V_{ab}(r) = \pm \frac{g_{S,V}^a g_{S,V}^b}{4\pi r} \exp(-r/\lambda) \quad (15)$$

where the  $-$  and  $+$  signs refer to scalar and vector interactions, respectively, and  $\lambda = \hbar/mc$ . For arbitrary vector interactions between electrically neutral atoms with proton number  $Z$ , and neutron number  $N$ ,

$$g_V = g_V^0 (Z \cos \tilde{\psi} + N \sin \tilde{\psi}) \quad (16)$$

where  $\tilde{\psi} = \arctan \frac{\tilde{q}_V^n}{\tilde{q}_V^p + \tilde{q}_V}$ . By comparing the generic Yukawa potential in 15 to the potential in 16 for scalar or vector boson exchange we can identify

$$\frac{g_a^0 g_b^0}{4\pi} = \alpha G u^2 \left( \left[ \frac{\tilde{q}}{\mu} \right]_a \left[ \frac{\tilde{q}}{\mu} \right]_b \right)^{-1} \quad (17)$$

where  $\mu = M/u$  with  $M$  being the atomic mass of the interacting nuclei and  $u$  being the atomic mass unit. Furthermore,

$$\left[ \frac{\tilde{q}}{\mu} \right] = \left[ \frac{Z}{\mu} \right] \cos \tilde{\psi} + \left[ \frac{N}{\mu} \right] \sin \tilde{\psi}. \quad (18)$$

For silicon,  $\left[ \frac{Z}{\mu} \right] = 0.504$  and  $\left[ \frac{N}{\mu} \right] = 0.496$  weighted by isotopic abundance. For oxygen,  $\left[ \frac{Z}{\mu} \right] = 0.500$  and  $\left[ \frac{N}{\mu} \right] = 0.500$  so that

$$\left[ \frac{\tilde{q}}{\mu} \right] \approx 0.50 \left( \cos \tilde{\psi} + \sin \tilde{\psi} \right). \quad (19)$$

Thus if we correct for the fact that 15 has been integrated over a mass distribution to correct for the fact that the nuclear potential is per nucleon we can write the correspondence between  $\alpha$  and  $\frac{g_a^0 g_b^0}{4\pi}$

$$\frac{g_a^0 g_b^0}{4\pi} = \alpha G u^2 \rho_{n,a} \rho_{n,b} 4.0 \left( \cos \tilde{\psi} + \sin \tilde{\psi} \right)^{-2} \quad (20)$$

where  $\rho_n$  is the density of nuclei per unit volume.

## References

- [1] E. G. Adelberger, B. R. Heckel, S. Hoedl, C. D. Hoyle, D. J. Kapner, and A. Upadhye. Particle-physics implications of a recent test of the gravitational inverse-square law. *Phys. Rev. Lett.*, 98:131104, Mar 2007.