## Physical mechanics

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#### 1 Cinematics

Having:

- $\mathcal{R}$  as rotation matrix of S' seen from S
- $\mathbf{R}$  the relative position of S' seen from

the space coordinates of a point are:

$$r = \mathcal{R}^T r' + \mathbf{R}$$
  
 $r' = \mathcal{R}(r - \mathbf{R})$ 

We can then define the antisymmetric angular momentum matrices as:

$$\hat{w} = \dot{\mathcal{R}}^T \mathcal{R} = \begin{pmatrix} 0 & -w_z & w_y \\ w_z & 0 & -w_x \\ -w_y & w_x & 0 \end{pmatrix}$$

$$\hat{w}' = \mathcal{R} \dot{\mathcal{R}}^T = \begin{pmatrix} 0 & -w_z' & w_y' \\ w_z' & 0 & -w_x' \\ -w_y' & w_x' & 0 \end{pmatrix}$$

And the a can then define the dual vectors too:

$$w = (w_x, w_y, w_z)$$
$$w' = (w'_x, w'_y, w'_z)$$

We get the relationships:

$$\hat{w}q = w \times q$$

$$\hat{w}'q = w' \times q$$

$$\hat{w} = \mathcal{R}^T \hat{w}' \mathcal{R}$$

$$\hat{w}' = \mathcal{R} \hat{w}$$

Defining **V** as the speed of S' as seen from S, we have the velocities (deriving the position one):

$$v = \mathcal{R}^{T}(v' + w' \times r') + \mathbf{V}$$
$$v' = \mathcal{R}[v - \mathbf{V} - w \times (r - \mathbf{R})]$$

Then with  $\alpha$  as the angular acceleration, A as the acceleration of S' as seen by S, and deriving this once more we get the acceleration:

$$\begin{split} a' &= -\alpha' \times r'; \text{acimutal} \\ &-2w' \times v'; \text{coriolis} \\ &-w' \times (w' \times r'); \text{centrifugal} \\ &+ \mathcal{R}a \\ &- \mathcal{R}\mathbf{A}; \text{drag} \end{split}$$

$$a = \mathbf{A} + \mathcal{R}^{T}[a'$$

$$+\alpha' \times r'$$

$$+1w' \times v'$$

$$+w' \times (w' \times r')]$$

## 2 Dynamics

# **2.1 Newton dynamics**The three newton laws:

• In absence of forces, a body remains it's constant lineal movement  $\mathbf{p} = \mathbf{m}\mathbf{v}$ :

$$\mathbf{0} = rac{\mathrm{d}}{\mathrm{d}\mathbf{t}}\mathbf{p} \longrightarrow \mathbf{p} = \mathrm{ct}.$$

• The change of lineal momentum is given by the force that acts on a body:

$$\mathbf{F} = \frac{\mathrm{d}}{\mathrm{d}\mathbf{t}}\mathbf{p}; \quad \mathbf{m} = \mathrm{ct.} \to \mathbf{F} = \mathbf{ma}$$

• Given two particles that interact, the force on particle 1 is the same strength and opposite sign than the force on particle 2.

$$\mathbf{F_1} = -\mathbf{F_2}$$

## 2.2 Common differential equations and solutions

Common types of problems and their differential equations and solutions:

• Uniform rectilinear movement:

$$\ddot{q} = 0 \longrightarrow q(t) = q_0 + \dot{q}_0 t$$

• Uniformly accelerated movement:

$$\ddot{q} = a \longrightarrow q(t) = q_0 + \dot{q}_0 t + \frac{1}{2}at^2$$

• Uniformly accelerated with friction:

$$\ddot{q} = a - b\dot{q} \longrightarrow \dot{q}(t) = \frac{a}{b} + \left(\dot{q}_0 - \frac{a}{b}\right)e^{-bt}$$

Where a/b is the limit speed.

Harmonic oscillator:

$$\ddot{q} + \omega^2 q = 0 \to$$

$$q(t) = \frac{\dot{q}_0}{\omega} \sin(\omega t) + q_0 \cos(\omega t)$$

$$= A \sin(\omega t + \phi)$$

Where

$$A = \sqrt{q_0^2 + \frac{\dot{q}_0^2}{\omega^2}}$$
$$\phi = \arccos \frac{\dot{q}_0}{\sqrt{q_0^2 \omega^2 + \dot{q}_0^2}}$$

 Charged particle in an electric and magnetic field (Lorentz force):

$$\mathbf{F} = \mathbf{q}\mathbf{E} + \mathbf{q}\mathbf{v} \times \mathbf{B}$$

Ends up creating a spiral with the axes perpendicular to the magnetic field, that turns with a frequency  $\omega_c$  called cyclotronic frequency, and a constant modulo of the speed.

#### 2.3 Non-inertial reference frames

When applying newton dynamics from the point of view of a non-inertial reference

frame, we get fictional forces:

$$\begin{split} m\mathbf{a}' &= \mathbf{F}' \quad ; \mathcal{R}\mathbf{F} \\ &- m\mathcal{R}\mathbf{A} \quad ; \mathbf{F}'_{\text{drag}} \\ &- m\dot{\boldsymbol{\omega}}' \times \mathbf{r}' \quad ; \mathbf{F}'_{\text{acimutal}} \\ &- 2m\boldsymbol{\omega}' \times \mathbf{v}' \quad ; \mathbf{F}'_{\text{Coriolis}} \\ &- m\boldsymbol{\omega}' \times (\boldsymbol{\omega}' \times \mathbf{r}') \quad ; \mathbf{F}'_{\text{centrifugal}} \end{split}$$

Then, for a particle on the surface of earth we have:

$$\mathbf{a}' = rac{1}{\mathbf{m}} \mathcal{R} \mathbf{F}_{ ext{ng}} + \mathbf{g}'_{ ext{effective}} - \mathbf{2} oldsymbol{\omega}' imes \mathbf{v}'$$

where  $\mathbf{F}_{ng}$  is the non-gravitational force and the effective gravitational acceleration is:

$$\mathbf{g}'_{\text{effective}} = (-\omega_{\mathbf{T}}^2 \mathbf{R}_{\mathbf{T}} \cos \lambda \sin \lambda) \mathbf{e}'_{\mathbf{y}}$$
$$+(-g + \omega_T^2 R_T \cos^2 \lambda) \mathbf{e}'_{\mathbf{z}}$$

Where  $\lambda$  is the latitude.

## 3 Geometry of particle systems

#### 3.1 Center of mass

$$egin{aligned} \mathbf{R}_{CM} &\equiv rac{1}{M_{ ext{total}}} \sum_{i=1}^{N} m_i \mathbf{r}_i \ \mathbf{R}_{CM} &\equiv rac{1}{M_{ ext{total}}} \int_{\mathcal{V}} 
ho(\mathbf{r}) \mathbf{r} d\mathcal{V} \end{aligned}$$

Note that for non-euclidean coordinates we have different  $d\mathcal{V}$ , so the volume integrals have extra members (generic volume integrals):

• Cylindrical (for cylinder of radius R and height Z):

$$\int_{0}^{R} \int_{0}^{Z} \int_{0}^{2\pi} r \ d\Theta dz dr$$

• Spherical, for sphere of radius R:

$$\int_0^R \int_0^{2\pi} \int_0^{\pi} r^2 \sin \varphi \ d\varphi d\Theta dr$$

#### 3.2 Inertia tensor

Symmetric tensor, defined as:

$$\mathbf{I} \equiv \sum_{i=1}^{N} m_i \mathbf{r}_i^2 \mathcal{I} - \sum_{i}^{N} m_i \mathbf{r}_i \mathbf{r}_i^T$$

In matrix form (if integrating, replace sum by volume/surface integral):

$$\sum_{i=1}^{N} \begin{pmatrix} y_i^2 + z_i^2 & -x_i y_i & -x_i z_i \\ -x_i y_i & x_i^2 + z_i^2 & -y_i z_i \\ -x_i z_i & -y_i z_i & x_i^2 + y_i^2 \end{pmatrix}$$

And coordinate integral form (as before, keep in mind the coordinates for the integral):

$$I_{lphaeta} = \int_{\mathcal{V}} 
ho(\mathbf{r}) [\delta_{lphaeta}r^2 - r^{lpha}r^{eta}] d\mathcal{V}$$

 $I_{xx}$ ,  $I_{yy}$  and  $I_{zz}$  are called **moments of** inertia with respect to the axis. While the

non-diagonal terms are called **products of inertia**.

For flat surfaces in the XY plane, we have the theorem of perpendicular axes:

$$I_{zz} = I_{xx} + I_{yy}$$

## 3.3 Inertia tensor in different reference systems

We have the general formula for the inertia tensor in the reference system S':

$$\mathbf{I}' = \mathcal{R}[$$

$$\mathbf{I} - \mathbf{I}_{M}$$

$$+ M(\mathbf{R} - \mathbf{R}_{CM})^{2} \mathcal{I}$$

$$- M(\mathbf{R} - \mathbf{R}_{CM})(\mathbf{R} - \mathbf{R}_{CM})^{T}$$

$$] \mathcal{R}^{T}$$

Where we defined the **inertia tensor of the center of mass**:

$$\mathbf{I}_{M} \equiv M\mathbf{R}_{CM}^{2} \mathcal{I} - M\mathbf{R}_{CM}\mathbf{R}_{CM}^{T}$$

#### 3.4 Particular cases

■ When  $S' \equiv S''_{CM}$  has the origin in the center of mass ( $\mathbf{R} = \mathbf{R}_{CM}$ ) and parallel axes to S, you get the Steiner theorem:

$$\mathbf{I} = \mathbf{I}_{CM}'' + \mathbf{I}_{M}$$

The inertia tensor on S is equal to the inertia tensor from  $S_{CM}^{\prime\prime}$  plus the inertia tensor of a single particle with the same mass in the center of mass.

• When only the origin is the same as the center of mass ( $\mathbf{R} = \mathbf{R}_{CM}$ ), then we have the general expression:

$$\mathbf{I}'_{CM} = \mathcal{R}\mathbf{I}''_{CM}\mathcal{R}^T$$

Where  $S_{CM}^{\prime\prime}$  is the reference system from the previous point.

Given the **spectral decomposition theorem**, it's always possible to find a rotation matrix  $\mathcal{R}$  that  $\mathbf{I}'_{CM} = \mathbf{I}'_{D}$  gets

diagonalized. That new system is called **principal axes reference system**.

### 3.A1 Diagonalizing the inertia tensor

- Build the characteristic polynomial:  $det(\mathbf{I} \lambda \mathcal{I})$
- Make it equal to 0 and resolve (3rd degree polynomial in worst case), that gives you the **eigenvalues** ( $\lambda$ ).
- Get each of the eigenvectors (v) with
   Iv = λv for each found eigenvalue λ.
- Now you get the rotation matrix R that diagonalizes the inertia tensor by doing:

$$\mathcal{R} = \begin{pmatrix} v_{1x} & v_{2x} & v_{3x} \\ v_{1y} & v_{2y} & v_{3y} \\ v_{1z} & v_{2z} & v_{3z} \end{pmatrix}$$

• And finally get the inertia tensor by applying  $I_D = \mathcal{R}I\mathcal{R}^T$