

David Fernández

## **Development of calibration and limit checking modules for a satellite's ground control software**

**Aalto University School of Electrical Engineering**

**Department of Radio Science and Engineering**

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**Thesis supervisor and instructor:** D.Sc. (Tech.) Jaan Praks



**Aalto University**  
**School of Electrical**  
**Engineering**

**Author:** David Fernández

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*All our dreams can come true, if we have the courage to pursue them.*

Walt Disney

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## Preface

# Acknowledgements

Otaniemi, Espoo, August 2013.

David.

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# Chapter 1

## Introduction

The relationship between human beings and space has existed since the beginning of time. It has always been a source of mystery, something we want to understand. More than 4000 years ago, the Egyptians and the Babylonians were influenced by the movements of the sun and the planets and, based on those movements, developed calendars for their crops. Later, the ancient Greeks developed the concept of *astronomy*, the science of the heavens. Subsequently, we can find philosophers such as Nicolaus Copernicus, Johannes Kepler, who explained the motion of the planets, and Galileo Galilei. In the 17th century, Sir Isaac Newton invented calculus, developed his law of gravitation and performed important experiments in optics.

The technological advancements of the 20th century, specially accelerated by the World War II, made physical exploration of space become possible. This thesis is oriented towards one of those advancements, artificial satellites.

A *satellite* is a natural or artificial object moving around a bigger body. This motion is defined as an orbit, enabled by the dominant force of gravity from the bigger body. In early 1945, the United States started the Vanguard Rocket development to launch a satellite. However, after several failed launches, it was the Soviet Union who took the advantage by launching the Sputnik-1 (Figure 1.1) on October 4, 1957. During the Cold War, the space race between the Soviet Union and the United States was a hard fight which made space technology advance quite rapidly.

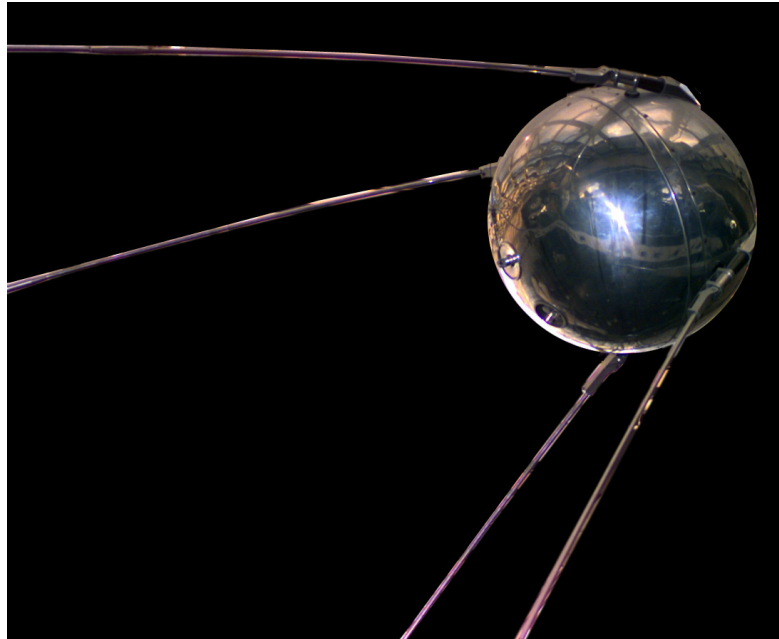


Figure 1.1: Sputnik 1 (NASA Public Domain)

Finally, on January 31, 1958 the United States managed to launch their first artificial satellite, the Explorer-1 (Figure 1.2).

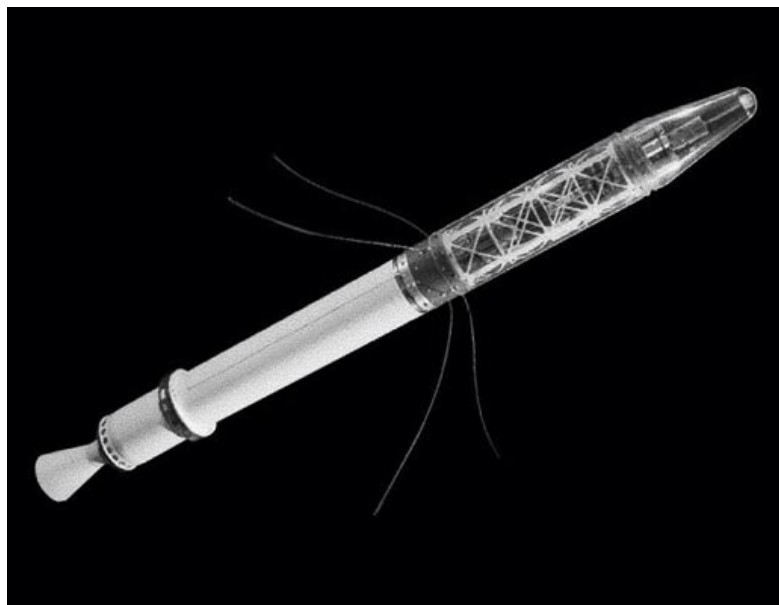


Figure 1.2: Explorer 1 (NASA Public Domain)

As of October 1, 2011 there were 966 operating satellites in orbit. About two-thirds of these were owned by the United States, Russia and China[1]. Their major types are:

- Communications: used for television, radio, Internet and telephone services.
- Navigation: using radio time signals, this satellites allow mobile receivers on the ground to determine their exact position. They are also used to determine the location of satellites situated in lower orbits.
- Exploration: used to observe distant planets, galaxies and other outer space objects by using telescopes and other sensors.
- Remote sensing: Remote sensing satellites are used to gather information about the nature and condition of Earth. The sensors in this kind of satellites receive electromagnetic emissions in several spectral bands and can detect the object's composition and temperature. Also, environmental conditions and so on. These satellites have also been widely used as military "spy satellites".

The constant evolution of technology and the growth of human needs have made the mission requirements rise throughout the last decades, thus, satellite mass has grown from Sputnik's 84 kg. and Explorer-1's 14k kg. to over 6,000 kg in 2007[2]. The main consequence of this, amongst others, has been an increment in mission costs.

To counter this trend, the small satellite movement was created by the academic community and it has shown how mission costs can be cut dramatically to a point in which a university can build a launch their own satellite. (Figure 1.3). Due to its success, it has become vigorous industry. Aalto-1 and ESTCube-1, projects with which this thesis is related, are nanosatellites, and the perfect example of this new concept.

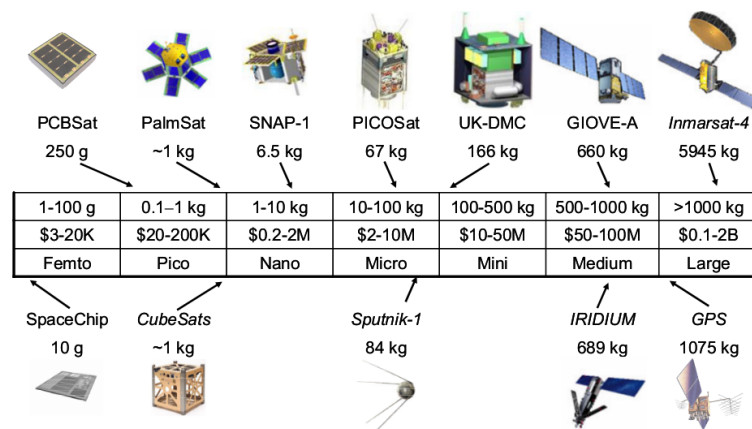


Figure 1.3: Satellite Mass and Cost Classification [2]

## 1.1 Background

As it was stated before, this thesis is closely related to two different projects. Aalto-1 and ESTCube-1. These two projects are based on the most common standard use by universities, CubeSat[3]. An open standard developed by the California Polytechnic State University and Stanford University.

### 1.1.1 Aalto-1

Led by Aalto University, Aalto-1 project aims to build a multi-payload remote sensing nanosatellite (Figure 1.4). The size of the satellite is approximately 34 cm x 10 cm x 10 cm with a mass of less than 4 kg[4]. Aalto-1 will also be the first Finnish satellite.



Figure 1.4: Aalto-1

There are different institutions cooperating to make this possible. The main payload, the imaging spectrometer, has been designed and built by VTT Technical Research Centre of Finland. The Radiation Monitor (RADMON) has been designed by the Universities of Helsinki and Turku in cooperation with the Finnish Meteorological Institute (FMI). The Plasma Brake has been designed by a consortium including the FMI, the Department of Physics of the University of Helsinki, the Departments of Physics and Astronomy and Information Technology of the University of Turku, the Accelerator laboratory of the University of Jyväskylä, Aboa Space Research Oy, Oxford Instruments Oy and other Finnish companies. Meanwhile, Aalto University is responsible for designing and building the satellite platform and the day-to-day operation of the project.[5]

Aalto-1's mission is to validate the technologies used by the payloads in space environment and measure their performance. In addition, it is also an educational project. Students are the main workforce towards its success. Being the the first Finnish student satellite mission, is a good tool to improve Finnish space teaching and also allow students to be in touch with prominent partners, both domestic and international, in the space technology field.

### 1.1.2 ESTCube-1

ESTCube-1 is a single-unit CubeSat (Figure 1.5). The size of the satellite is approximately 10 cm x 10 cm x 10 cm with a maximum mass of 1.33 kg. It has been build by students of the universities of Tartu and Tallinn, in Estonia, and it is the first Estonian satellite [6]. It was launched from the Guaiana Space Centre on May 7, 2013 as one of the three payloads of the Vega VV02 rocket [7].

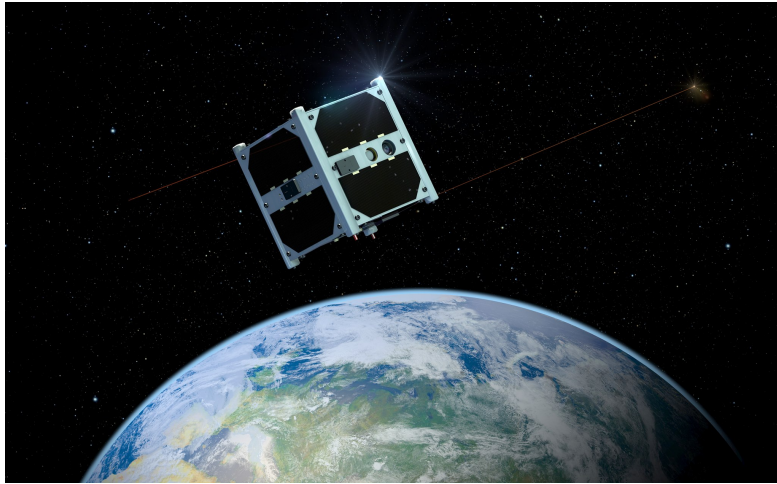


Figure 1.5: ESTCube-1

ESTCube-1's main goal is to observe and measure the E-sail effect for the first time. It has been placed into a polar low Earth orbit (LEO) and will deploy a single 10 m long and 9 mm wide tether[6]. This has relation with Aalto-1's Plasma Brake, which will deploy a 100 m long tether. The require duration of ESTCube-1's mission is a few weeks and it can be extended to about a year.

## 1.2 Problem statement

## 1.3 Research objectives and scope

The purpose of this thesis is to develop a calibration system for a ground station which will solve the problem mentioned above. In addition, it is necessary to check if the received values are within a expected range, so a different system will be developed for that purpose.

Using ESTCube-1 project as a baseline, the author aims to develop such systems which are generic enough to be used by any number of satellite missions. The modules developed will be integrated into *Hummingbird*, an open source platform for monitoring and control which will be explained later in this work.

## 1.4 Motivations

## 1.5 Outline of the thesis



# Chapter 2

## Satellite Communications

This chapter covers the topic of how the satellite communicates with Earth and back. One of the most important constraints that needs to be approached when working with ground station is the fact that the satellite is orbiting around Earth. To understand this, the first section of this chapter will cover what orbits are, what elements describe them and how that information can be delivered by a computer. Following, the different types of data exchanged by the satellite and the ground station will be covered. The last section of the chapter goes through what a ground station is, its hardware and software components and how it communicates with a satellite in space.

### 2.1 Orbits

An orbit is the **gravitationally curved path of an object around a center of mass**. Examples of orbits can be the Earth around the Sun or artificial satellites around the Earth.

#### 2.1.1 Kepler's Laws

Planetary movements were first mathematically defined by the German mathematician, astronomer and astrologer Johannes Kepler in the 17th century. He concentrated his observations into three simple laws[8]:

- The orbit of each planet is an ellipse with the Sun occupying one focus.
- The line joining the Sun to a planet sweeps out equal areas in equal intervals of time.
- A planet's orbital period is proportional to the mean distance between the Sun and the planet, raised to the power of  $3/2$ .

These laws generally apply to every celestial body. When analysing two bodies, if one is much bigger than the other, it conforms the "two-body problem". It assumes that both bodies are spherical and they are modelled as if they were point particles. This means that influences from any third body are discarded. The analysis of this problem has resulted in six elements that completely define an orbit, which will be explained in the next section.

### 2.1.2 Classical Orbital Elements

The Classical Orbital Elements are six parameters which uniquely identify an orbit. They also can be used to predict future positions of the satellite.[9]

The first two elements, the orbit's size and shape are defined based on a 2D representation on an ellipse (Figure 2.1).

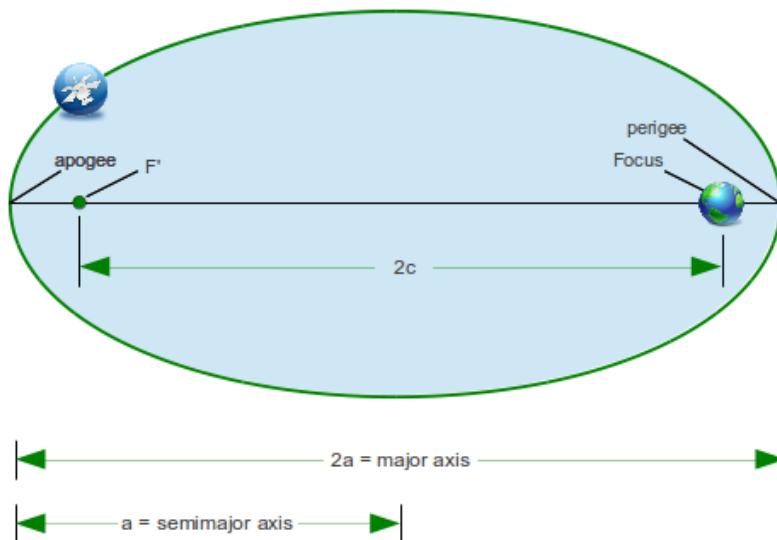


Figure 2.1: Semimajor Axis

- The *semimajor axis* ( $a$ ) is one half the distance across the long axis of the orbit, and it represents the orbit's size.
- The *eccentricity* represents the shape of the orbit. It describes how much the ellipse is elongated compared to a circle. Based on the latter, the orbit can have the following shapes, as shown in Figure 2.2

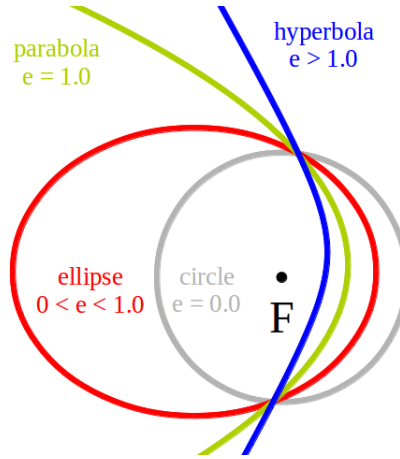


Figure 2.2: Eccentricity

Before jumping onto the next Orbital Elements it is necessary to point out that the Geocentric-equatorial Coordinate System will be used. It is now a 3D representation, where the fundamental plane is Earth's equatorial plane and the principal direction is in the vernal equinox direction (see Figure 2.3).

The following orbital elements define the orientation of the orbital plane:

- The inclination ( $i$ ) describes the tilt of the orbital plane with respect to the reference plane. It is measured at the ascending node. This is, where the orbit crosses with the reference plain when moving upwards.
- The right ascension of the ascending node ( $\Omega$ ) represents the angle between the principal direction and the point where the orbital plane crosses the reference plane from south to north measured eastward.

Based on this two elements, the orbits can be classified as shown in Table 2.1

| Inclination ( $i$ )           | Orbital Type   | Diagram |
|-------------------------------|--|---------|
| $0^\circ$ or $180^\circ$      | Equatorial   |         |
| $90^\circ$                    | Polar  |         |
| $0^\circ \leq i < 90^\circ$   | Direct or Prograde (moves in the direction of Earth's rotation)          |         |
| $90^\circ < i \leq 180^\circ$ | Indirect or Retrograde (moves against the direction of Earth's rotation) |         |

Table 2.1: Types of Orbits and Their Inclination [9].

Although it is not part of the COEs, orbits can also be sorted by their altitude. NASA's classification divides orbits in three groups (Table 2.2).

| Orbit  | Altitude ( $a$ )          | Uses                                       |
|--|---------------------------|--|
| Low Earth Orbit (LEO)                          | $a < 2000Km$              | Scientific and weather satellites          |
| Medium Earth Orbit (MEO)                       | $2000Km \leq a < 36000Km$ | GPS  |
| High Earth Orbit (HEO) or Geosynchronous (GSO) | $36000Km$                 | Communications (phones, television, radio) |

Table 2.2: NASA's classification of orbits. [10]

It is now time to go through the last two COEs:

- The argument of perigee ( $\omega$ ) is the angle between the ascending node and the perigee, measured in the direction of the satellite's motion.
- The true anomaly ( $v$ ) specifies the location of the satellite within the orbit. Amongst all the CEOs, this is the only one which changes over time. It is the angle between the perigee and the satellite's position vector measured in the direction of its motion.

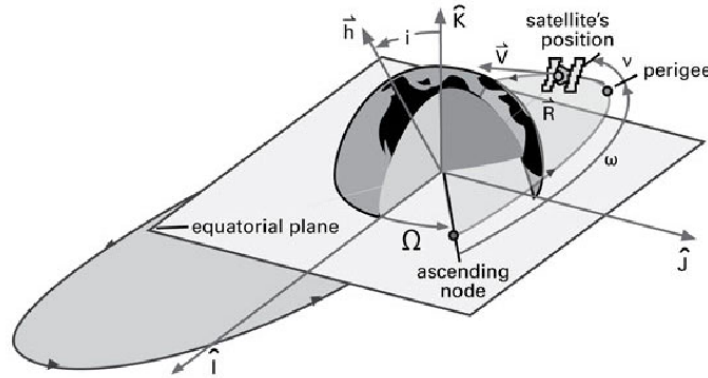


Figure 2.3: Classical Orbital Elements [9]

### 2.1.3 Ground Tracks

The satellite ground tracks are the projection of its orbit onto Earth. An example of this can be Figure 2.4.

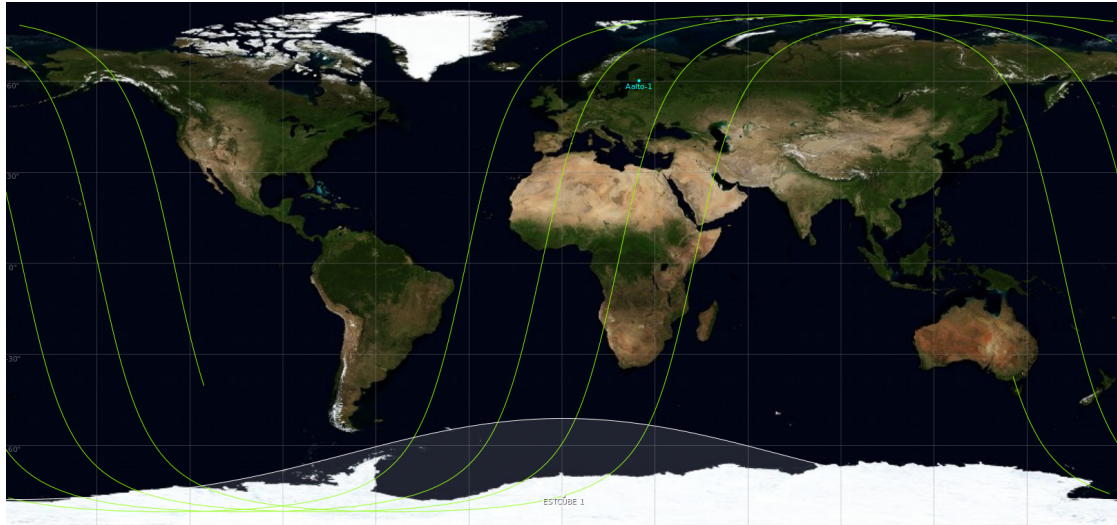


Figure 2.4: Ground Tracks of ESTCube-1

Since the orbital plane does not move in inertial space, the satellite's orbit will always be the same. If Earth did not move the representation of the orbit would be a single line, as the ground track would continuously repeat. However, Earth rotates at 1600 km/hr. Thus, even if the orbit does not change, from the Earth-based observer's point of view it appears to shift to the west.

### 2.1.4 Two Line Elements

A two line element set (TLE) is a data format created by the North American Aerospace Defense Command (NORAD) and NASA to transport sets of orbital elements describing satellite orbits around Earth. These TLEs can be later processed by a computer to calculate the position of a satellite at a particular time and are usually used by ground stations in order to track them.

The following snippet shows an example of a TLE for the International Space Station.

```
ISS (ZARYA)
1 25544U 98067A 13166.62319444 .00005748 00000-0 10556-3 0 120
2 25544 51.6483 116.0964 0010829 73.3727 265.7013 15.50799671834453
```

## CHAPTER 2. SATELLITE COMMUNICATIONS

| Field | Columns | Content  | Example      |
|-------|---------|--|--------------|
| 1     | 01      | Line number  | 1            |
| 2     | 03-07   | Satellite number   | 25544        |
| 3     | 08      | Classification (U=Unclassified)  | U            |
| 4     | 10-11   | International Designator<br>(Last two digits of launch year)                           | 98           |
| 5     | 12-14   | International Designator<br>(Launch number of the year)                                | 067          |
| 6     | 15-17   | International Designator (Piece of the launch)   | A            |
| 7     | 19-20   | Epoch Year (Last two digits of year)   | 08           |
| 8     | 21-32   | Epoch (Day of the year and fractional<br>portion of the day)                           | 264.51782528 |
| 9     | 34-43   | First Time Derivative of the Mean Motion   | -0.00002182  |
| 10    | 45-52   | Second Time Derivative of Mean Motion<br>(decimal point assumed)                       | 00000-0      |
| 11    | 54-61   | BSTAR drag term (decimal point assumed)  | -11606-4     |
| 12    | 63      | Ephemeris type   | 0            |
| 13    | 65-68   | Element number   | 292          |
| 14    | 69      | Checksum (Modulo 10)<br>(Letters, blanks, periods, plus signs = 0;<br>minus signs = 1) | 7            |

Table 2.3: Two-Line Element Set Format Definition, Line 1

| Field | Columns | Content   | Example     |
|-------|---------|---|-------------|
| 1     | 01      | Line number   | 1           |
| 2     | 03-07   | Satellite number  | 25544       |
| 3     | 09-16   | Inclination [Degrees]   | 51.6416     |
| 4     | 18-25   | Right Ascension of the Ascending<br>Node [Degrees]                  | 247.4627    |
| 5     | 27-33   | Eccentricity (decimal point assumed)<br>(Launch number of the year) | 0006703     |
| 6     | 35-42   | Argument of Perigee [Degrees]                                       | 130.5360    |
| 7     | 44-51   | Mean Anomaly [Degrees]  | 325.0288    |
| 8     | 53-63   | Mean Motion [Revs per day]  | 15.72125391 |
| 9     | 64-68   | Revolution number at epoch [Revs]                                   | 56353       |
| 10    | 69      | Checksum (Modulo 10)  | 00000-0     |

Table 2.4: Two-Line Element Set Format Definition, Line 2

## 2.2 Data

The data exchanged between a satellite and the ground stations on Earth can be divided into three different categories: the beacon, the telemetry and the telecommands.

### 2.2.1 Beacon

A *beacon* is a radio signal transmitted continuously or periodically over a specified radio frequency. It provides a small amount of information such as identification or location, but it can have more applications. Examples of these are: adjust the power of the ground station signal based on the beacon's strength or tune the ground station to compensate the doppler shift.

### 2.2.2 Telemetry

*Telemetry* data is sent from the satellite to the ground station and can also be divided in three sub-categories.

The *housekeeping data* provides information about the health and operating status of the satellite. Examples of this data can be pressure, voltages and currents, or also bits representing the operational status of all the components as it is shown in Figure 2.5. The size of this data is usually quite small, so a bit rate on only a few hundreds of bits per second is enough to complete the transmission successfully.

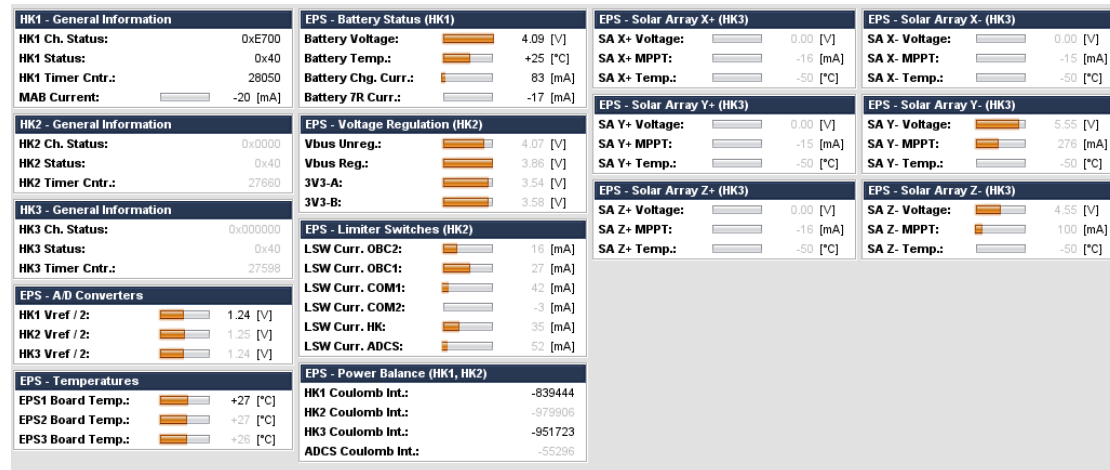


Figure 2.5: Housekeeping data of the satellite Masat-1

*Attitude data* is generated by different sensors, such as magnetometers, gyroscopes, accelerometers and Sun, Earth and star sensors.

*Payload data* changes with every mission and needs to be considered individually. Scientific or Earth-observing mission normally generate very large data volumes, specially in the form of images. An example of this can be Figure 2.6, the first picture taken by the Hungarian nanosatellite Masat-1[13].

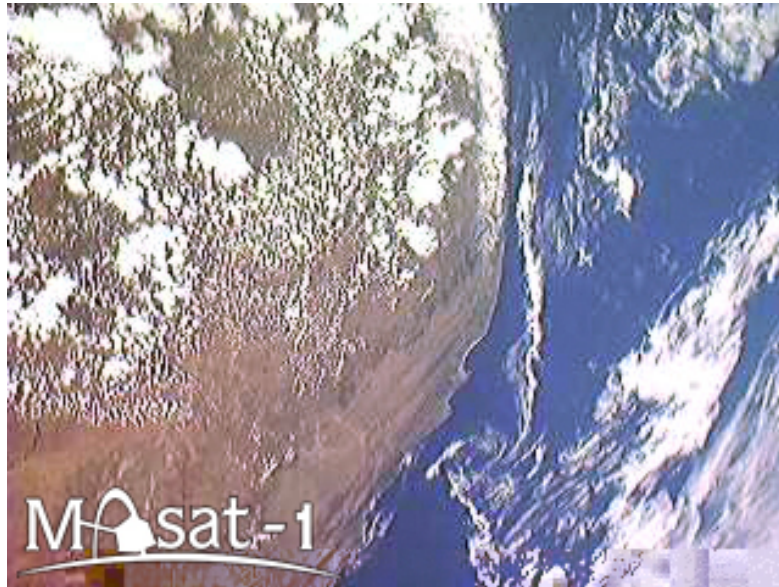


Figure 2.6: Picture of South Africa taken from the nanosatellite Masat-1

### 2.2.3 Telecomands

The telecommands are sent from the ground station to the satellite. They are used to remotely control its functions and are divided in three basic types [8]:

- *Low-level on-off commands.* These are logic-level pulses used to set or reset log flip-flops.
- *High-level on-off commands.* Higher-powered pulses, capable of operating a latching relay or RF waveguide switch directly.
- *Proportional commands.* Digital words. Used for purposes such as reprogramming memory locations on the on-board computer or setting up registers in the attitude control subsystem.



## 2.3 Ground Station

One integral part of every satellite mission is the ground station. It works as the first and final piece of the communication link. Its main functions are the following:

- Tracking the satellite to determine its position in orbit.
- Gather data to keep track of the satellite's data and status.
- Command operations to control the different functions of the satellite.
- Process the received engineering and scientific data to present it in the required formats.

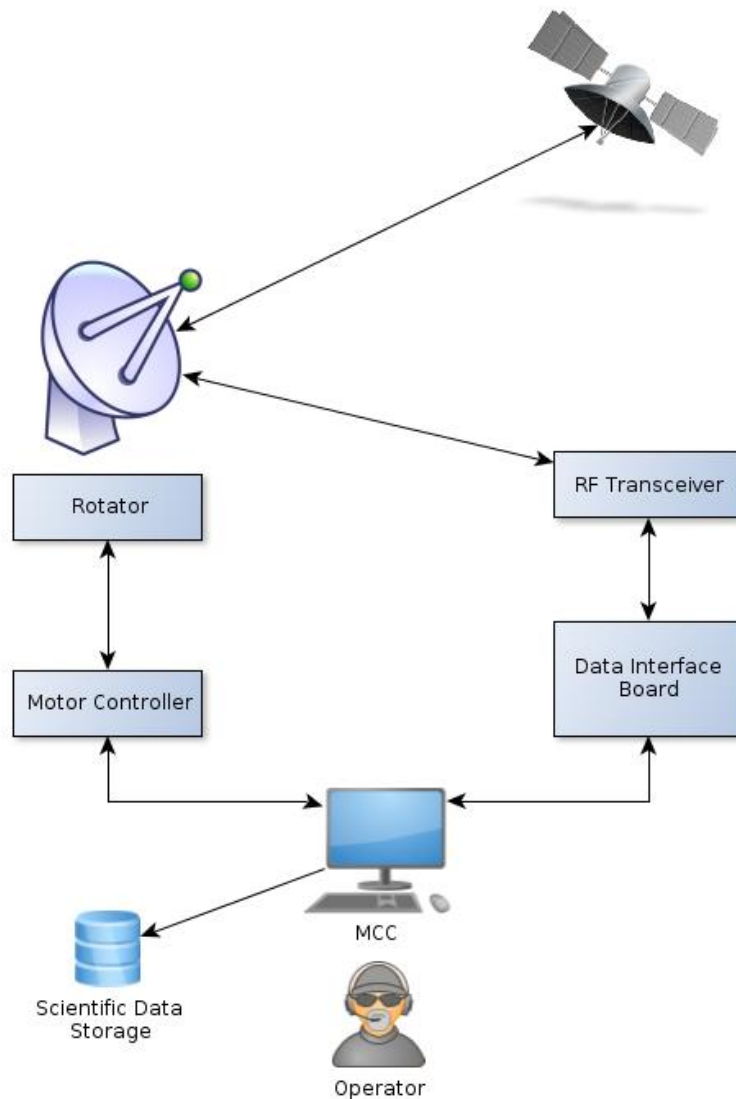


Figure 2.7: Diagram of a ground station

It is important to remember that university satellites are usually classified as amateur satellites. This means that they use amateur radio frequencies and the usage of the ground station is bound to each country's amateur radio regulations.

### 2.3.1 Hardware

The main components of a ground station are the antenna, the transceiver, the data recorders and the computers and their peripherals.

#### Antennas

The main hardware component of a ground station is the antenna. Its functions may include tracking, receiving telemetry, sending telecommands, etc.

The frequencies most commonly used for amateur satellites are shown in Table 2.5.

Table 2.5: TODO: Table with Frequencies.

#### Transceiver

A transceiver is a hardware unit containing both a transmitter and a receiver. It acts as an intermediary between the antenna and a computer, changing the radio frequency into bytes and viceversa.

### 2.3.2 Software

The activity in the ground station does not start when the satellite is passing over it and does not end once it is gone. There are certain tasks that need to be done before, during and after the pass.

Before the satellite arrives it is necessary to determine and predict its orbit. Based on this prediction the software will schedule future passes and generate the command list which will be sent during the pass.

The real-time software comes into operation when the satellite is visible from the ground station. It is in charge of controlling the antenna rotor to follow it across the sky; it will also send telecommands to the satellite and verify their correct reception. In addition, it will receive the data being transmitted from the satellite, which will be processed later.

Once the satellite is not visible any more the post-pass software comes into play. The data received during the pass is now processed and stored so the specialists can analyse it.

There are many different kinds of software oriented towards its use by amateur satellite missions. Examples of this are **GPredict**[14] and **Orbitron**[15], which are used for tracking and prediction, or **Carpcomm**[16], which amongst other functionalities aims to build a network of ground stations.

There is also professional software in use aiming to build ground station networks. One example is **GENSO**, a project of the European Space Agency (ESA) coordinated by its Education Office. The University of Vigo in Spain hosts the European Operations Node and coordinates the access to the network[17]. At the same time, and as a cooperation with the ESTCube-1 project, CGI is supervising the development of similar solution, **Hummingbird**[18], which will be explained more deeply in the following chapters, as this thesis is part of the mention project.

### 2.3.3 Protocols

A protocol is an agreement between the communicating parties on how communication is to proceed[19]. This section will be focused on the OSI Reference model as well as on some of the most popular protocols for amateur radio communications.

#### The OSI Reference Model

The Open Systems Interconnection (OSI) Reference Model was developed in 1983 and revised in 1995. This model deals with connecting systems that are open for communication with other systems. It consists on seven layers which are explained in Table 2.6.

| OSI Model    |                 |                 |  |
|--------------|-----------------|-----------------|--|
|              | Data Unit       | Layer           | Function   |
| Host Layers  | Data            | 7. Application  | Network process to application.  |
|              |                 | 6. Presentation | Data representation, encryption and decryption, convert machine dependent data to machine independent data |
|              |                 | 5. Session      | Interhost communication, managing sessions between applications  |
|              | Segments        | 4. Transport    | End-to-end connection, reliability and flow control  |
| Media Layers | Packet/Datagram | 3. Network      | Path determination and logical addressing  |
|              | Frame           | 2. Data link    | Physical addressing  |
|              | Bit             | 1. Physical     | Media, signal and binary transmission  |

Table 2.6: OSI Model[19].

## AX.25

AX.25 is a data link layer protocol designed for use by amateur radio operators. It occupies the first, second and third layers of the OSI model. However, AX.25 was developed before the model came into action, so its specification was not written to separate into OSI layers.

The link-layer packet radio transmission takes place in small blocks of data called frames. Those frames are represented in the Figures 2.8 and 2.9.

| Flag     | Address      | Control   | Info     | FCS     | Flag     |
|----------|--------------|-----------|----------|---------|----------|
| 01111110 | 112/224 Bits | 8/16 Bits | N*8 Bits | 16 Bits | 01111110 |

Figure 2.8: Supervisory and Unnumbered frames [20]

| Flag     | Address         | Control   | PID    | Info     | FCS     | Flag     |
|----------|-----------------|-----------|--------|----------|---------|----------|
| 01111110 | 112/224<br>Bits | 8/16 Bits | 8 Bits | N*8 Bits | 16 Bits | 01111110 |

Figure 2.9: AX.25 Information frame [20]

## FX.25

FX.25 is an extension to the AX.25 protocol. It has been created to complement the AX.25 protocol, providing an encapsulation mechanism that does not alter the AX.25 data or functionalities. AX.25 packets are easily damaged, and this extension intends to remedy the situation by providing a Forward Error Correction (FEC) capability at the bottom of Layer 2.

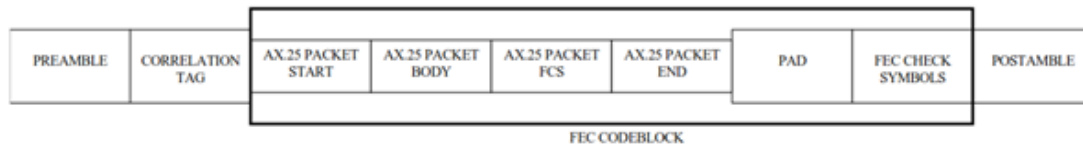


Figure 2.10: FX.25 frame structure [21]

## Chapter 3

# Hummingbird: the open source platform for monitoring and controlling small satellites

The continuous growth of the popularity of small satellites has been accompanied by an increased interest in creating more a better solutions for the ground segments of the missions. These missions are more complex every time and the use of software oriented for amateurs (FIXME!!!) is starting to prove insufficient. For this reason there are nowadays several projects to build professional-like software and ground station networks to control small satellites. This chapter introduces one of them, *Hummingbird*; the project into which the work carried on in this thesis will be integrated.

## Chapter 4

# Requirements

# Chapter 5

## Implementation

### 5.1 Design

### 5.2 Technologies used

As it has been explained previously, the two modules developed as part of the work for this thesis have been designed to be integrated with the Hummingbird project. To do so, Java[22] has been chosen as the main programming language, as it is the language used for the development of Hummingbird. In the same way, Apache Camel[23] and ActiveMQ [24] are used for the communication with the rest of the modules. XStream[25] has been chosen as the library used to parse the XML files used to specify the scientific information.

The final piece of technology used is BeanShell[26], a Java-like scripting language and interpreter which runs in the Java Runtime Environment. The calibration module has been designed to be generic, adaptable to every mission. Also, the goal was to make the calibration information input easy for the scientists, meaning this no need to any difficult Java programming, compilation and so on. BeanShell integrates with the Java code and allows to run those scripts **in (under, on????) runtime**.

### 5.3 User manual

#### 5.3.1 Calibration module

This short user manual covers the use of the calibration module. The process is fully automated, so the user only needs to configure the pertinent XML file containing the information related to all the parameters which need to be calibrated.



There can be as many files as needed, although it is recommended to have one file per subsystem. This way, the person who is making the changes will not have to be worried about modifying some other parts they do not understand. It is advisable that each file is called as the corresponding subsystem.

The location of the folder where the XML files are stored is fully configurable by a system property. It can be set like this: **-Dpath="/path/to/the/folder"**.

The XML file has the following format:

```
<calibration>
  <entry>
    <id></id>
    <description></description>
    <outputId></outputId>
    <unit></unit>
    <scriptInfo>
      <isVector></isVector>
      <resultVariable></resultVariable>
      <auxParameters></auxParameters>
      <script></script>
    </scriptInfo>
  </entry>
</calibration>
```

Table 5.1: blablabla

- **id**: name of the parameter to calibrate.
- **Description**: description of the parameter.
- **outputId**: name of the parameter generated after the calibration. If left blank, it will be the same as **id**. Please note that all calibrated parameters names end in **\_cal**.
- **unit**: Units in which the value is represented.
- **isVector**: **true** if the result of the calibration is a vector with several values (which generates several new parameters) or **false** if the calibration returns a single value.
- **resultVariable**: variable in the script in which the result will be stored.
- **auxParameters**: if there are extra parameters needed for the calibration process it is necessary to list them here separated by commas (','),. Please note that if the extra parameters also need to be calibrated the parameters needed for its calibration also must be included here instead of the original one (See Figure ??).

- **script**: script to generate the calibrated value. Please note that if extra parameters are needed, their calibration script must be included here, not the parameter name (See Figure ??).

```

<calibration>
  <entry>
    <id>parameterA</id>
    <description>Example of parameter for
      simple calibration</description>
    <outputId>generatedA</outputId>
    <unit>E</unit>
    <scriptInfo>
      <isVector>>false</isVector>
      <resultVariable>result</resultVariable>
      <auxParameters></auxParameters>
      <script>result = (parameterA*779.09823)/3145.2839</script>
    </scriptInfo>
  </entry>
</calibration>

```

Table 5.2: Example of simple calibration

Figure ?? represents the simplest example of calibration information. **parameterA** is the parameter to be calibrated and the user has chosen that the name of the calibrated parameter will be **generatedA**. The software will automatically append **\_cal**, so the final output name will be **generatedA\_cal**. The information about the calibration script states that the result will not be a vector and the value after the calculations will be stored in a variable called **result**. There are no extra parameters needed and the calibration script is  $(parameterA * 779.09823) / 3145.2839$ .

```

<calibration>
  <entry>
    <id>parameterA</id>
    <description>Example of parameter which depends
      on others to be calibrated</description>
    <outputId></outputId>
    <unit>E</unit>
    <scriptInfo>
      <isVector>>false</isVector>
      <resultVariable>result</resultVariable>
      <auxParameters>parameterB,parameterC</auxParameters>
      <script>result = (parameterA*(parameterB*2345/37))
        /3145.2839 + (parameterC*2)</script>
    </scriptInfo>
  </entry>
</calibration>

```

Table 5.3: Example of calibration depending on other parameters

Figure ?? shows an example of a parameter which depends on others for calibration. Again, **parameterA** is the name of the parameter to be calibrated. In this case the user has not selected an output ID, so it will by default be **parameterA\_cal**. The result of the calibration will not be a vector and it needs **parameterB** and **parameterC** to be calibrated. The calibration script can be explained as follows:

- Calibration script for **parameterA**:  $result = ((parameterA * (parameterB\_cal)) / 3145.2839) + (parameterC\_cal)$
- Instead of just stating that **parameterB\_cal** is needed to carry on the calibration, the user must specify its calibration script:  $parameterB * 2345/37$
- Same thing with **parameterC\_cal**:  $parameterC * 2$
- The final result is what can be seen in the example:  $result = (parameterA * (parameterB * 2345/37)) / 3145.2839 + (parameterC * 2)$

```
<calibration>
  <entry>
    <id>parameterA</id>
    <description>Example of parameter for simple calibration</description>
    <outputId>generatedA</outputId>
    <unit>E</unit>
    <scriptInfo>
      <isVector>>false</isVector>
      <resultVariable>result</resultVariable>
      <auxParameters></auxParameters>
      <script>result = (parameterA*779.09823)/3145.2839</script>
    </scriptInfo>
  </entry>
</calibration>
```

Table 5.4: Example of simple calibration

### 5.3.2 Limit checking module

This subsection covers the user manual for the limit checking module. The process is fully automated, so the user only needs to configure the pertinent XML file containing the information related to the limits of every parameter.

There can be as many files as needed, although it is recommended to have one file per subsystem. This way, the person who is making the changes will not have to be worried about modifying some other parts they do not understand. It is advisable that each file is called as the corresponding subsystem.

The location of the folder where the XML files are stored is fully configurable by a system property. It can be set like this: **-Dpath="/path/to/the/folder"**.

The XML file has the following format:

```
<limitChecking>
  <entry>
    <id></id>
    <limits>
      <sanityLower></sanityLower>
      <hardLower></hardLower>
      <softLower></softLower>
      <softUpper></softUpper>
      <hardUpper></hardUpper>
      <sanityUpper></sanityUpper>
    </limits>
  </entry>
</limitChecking>
```

Table 5.5: blablabla

- **id:** name of the parameter to calibrate which limits are to be checked.
- **Sanity limits:** Optional. If the value is below the lower limit or above the upper limit it is discarded.
- **Hard limits:**
  - If the sanity limits are available anything between these limits and the sanity limits is considered an error.
  - If the sanity limits are disabled anything below the lower limit or above the upper limit is considered an error.
- **Soft limits:**
  - Anything between the lower and upper soft limits is considered an OK value.
  - Anything between the soft limits and the hard limits is considered OK, but with a warning.

The following two examples show the two ways in which the limit checking module can be configured:

```
<limitChecking>
  <entry>
    <id>parameterA</id>
    <limits>
      <sanityLower>-100</sanityLower>
      <hardLower>-75</hardLower>
      <softLower>-20</softLower>
      <softUpper>20</softUpper>
      <hardUpper>75</hardUpper>
      <sanityUpper>100</sanityUpper>
    </limits>
  </entry>
</limitChecking>
```

Table 5.6: Limit checking with sanity limits available

```
<limitChecking>
  <entry>
    <id>parameterA</id>
    <limits>
      <hardLower>-75</hardLower>
      <softLower>-20</softLower>
      <softUpper>20</softUpper>
      <hardUpper>75</hardUpper>
    </limits>
  </entry>
</limitChecking>
```

Table 5.7: Limit checking without sanity limits available

## Chapter 6

### Future work

## Chapter 7

## Conclusions



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