The amplitudes as measured by 2 observers are

$$f(\vec{x},t) = e^{i(\vec{k}\cdot\vec{x}-\omega t)}, \quad f(\vec{x},t) \neq e^{i(\vec{k}\cdot\vec{x}-\omega t)}$$

we lemand the amplies be equal
$$e^{i(\vec{k}\cdot\vec{x}-\omega t)} = e^{i(\vec{k}\cdot\vec{x}-\omega t')}$$
This suggests  $x^{M} = (t, \vec{x}), \quad k'' = (\frac{\omega}{c}, \vec{k})$ 

This suggests  $x^{M} k'' g = x^{M} k$  is Lorentz invariant,

consequently,  $x^{M}$  is approximate.

By covariance of  $k'''$ , we know it follows Loventz transformate

Recall the transformation on spacetime coordinates:
$$f' = f(t - \vec{p} \cdot \vec{r})$$

$$f'' = f' (t - \vec{p} \cdot \vec{r})$$
This suggests the transformation for  $k'''$ :
$$k'' = f'(k' - \vec{p} \cdot \vec{k})$$

$$= f'' = f'(k' - \vec{p} \cdot \vec{k})$$

$$= f'' = f''$$

This is not the familiar form of relativistic doppler; to get such, do assume  $\vec{\beta} \cdot \vec{k} = \beta k$ , (busts allign with  $\vec{k}$ ) w= r (w - cpk) Then Thuse dispersion  $\frac{w}{k} = c$ ,  $k = \frac{w}{c}$ 0/= X(W-BW) When we dead of the state of th = w (1-B)(1-B)  $w' = w \int CI - \beta ds$   $\int (1+\beta ds)$ DAMP TO GEO GIVES SHACKELY FRINGE TO 의 10 = ZWpGPR May 2 전환

To w > 1 = 12

pandson Changes 2. 8.2024