

Homework #7

- (Kittel 6.4) **Energy of gas of extreme relativistic particles.** Extreme relativistic particles have momenta p such that $pc \gg Mc^2$, where M is the rest mass of the particle. The de Broglie relation $\lambda = h/p$ for the quantum wavelength continues to apply. Show that the mean energy per particle of an extreme relativistic ideal gas is 3τ if $\epsilon \cong pc$ in contrast to $\frac{3}{2}\tau$ for the nonrelativistic problem. (An interesting variety of relativistic problems are discussed by E. Fermi in Notes on Thermodynamics and Statistics, University of Chicago Press, 1966, paperback.)

Solution.

$$\langle \epsilon \rangle = \frac{\int_0^\infty \epsilon D(\epsilon) e^{-\beta \epsilon} d\epsilon}{\int_0^\infty D(\epsilon) e^{-\beta \epsilon} d\epsilon}$$

For $\epsilon \cong pc$, $D(\epsilon) \propto \epsilon^2$. Hence

$$\langle \epsilon \rangle = \frac{\int_0^\infty \epsilon \epsilon^2 e^{-\beta \epsilon} d\epsilon}{\int_0^\infty \epsilon^2 e^{-\beta \epsilon} d\epsilon}$$

Let $x = \beta \epsilon = \epsilon / k_B T$. Then

$$\langle \epsilon \rangle = k_B T \frac{\int_0^\infty x^3 e^{-x} dx}{\int_0^\infty x^2 e^{-x} dx} = k_B T \frac{3 \times 2}{2 \times 1} = 3k_B T$$

- (Kittel 6.5) **Integration of the thermodynamic identity for an ideal gas.** From the thermodynamic identity at a constant number of particles we have

$$d\sigma = \frac{dU}{\tau} + \frac{pdV}{\tau} = \frac{1}{\tau} \left(\frac{\partial U}{\partial \tau} \right)_V d\tau + \frac{1}{\tau} \left(\frac{\partial U}{\partial V} \right)_\tau dV + \frac{pdV}{\tau}$$

Show by integration that for an ideal gas the entropy is

$$\sigma = C_V \log \tau + N \log V + \sigma_1$$

where σ_1 is a constant independent of τ and V .

Solution.

$$d\sigma = \frac{1}{\tau} \left(\frac{\partial U}{\partial \tau} \right)_V d\tau + \frac{1}{\tau} \left(\frac{\partial U}{\partial V} \right)_\tau dV + \frac{pdV}{\tau}$$

For an ideal gas, $pV = Nk_B T = N\tau$.

$$\left(\frac{\partial U}{\partial V} \right)_\tau = T \left(\frac{\partial S}{\partial V} \right)_\tau - p = T \left(\frac{\partial p}{\partial T} \right)_V - p = T \frac{Nk_B}{V} - p = 0$$

$$d\sigma = \frac{1}{\tau} \left(\frac{\partial U}{\partial \tau} \right)_V d\tau + \frac{pdV}{\tau} = \frac{C_V}{\tau} d\tau + \frac{N}{V} dV$$

$$\sigma = \int \frac{C_V}{\tau} d\tau + \frac{N}{V} dV = C_V \ln \tau + N \ln V + \sigma_1$$

3. (Kittel 6.7) **Relation of pressure and energy density.**

(a) Show that the average pressure in a system in thermal contact with a heat reservoir is given by

$$p = - \frac{\sum_s (\partial \mathcal{E}_s / \partial V)_N \exp(-\mathcal{E}_s / \tau)}{Z}$$

where the sum is over all states of the system.

(b) Show for a gas of free particles that

$$\left(\frac{\partial \mathcal{E}_s}{\partial V} \right)_N = - \frac{2}{3} \frac{\mathcal{E}_s}{V}$$

as a result of the boundary conditions of the problem. The result holds equally well whether \mathcal{E}_s refers to a state of N noninteracting particles or to an orbital.

(c) Show that for a gas of free non-relativistic particles

$$p = 2U/3V$$

where U is the thermal average energy of the system. The result is not limited to the classical regime; it holds equally well for fermion and boson particles, as long as they are nonrelativistic.

Solution.

$$p = -\frac{1}{Z} \sum_s \left(\frac{\partial \epsilon_s}{\partial V} \right) e^{-\epsilon_s / \tau}$$

$$\epsilon_s = \frac{\hbar^2}{2m} k_s^2 \quad \text{with } k_s \approx \frac{1}{L} = \frac{1}{V^{1/3}}$$

$$\epsilon_s \propto \frac{1}{V^{2/3}}$$

Let $\epsilon_s \propto \frac{A_s}{V^{2/3}}$. Then

$$\frac{\partial \epsilon_s}{\partial V} = -\frac{2}{3} \frac{A_s}{V^{5/3}} = -\frac{2}{3} \frac{\epsilon_s}{V}$$

$$p = \frac{2}{3} \frac{1}{Z} \sum_s \frac{\epsilon_s}{V} e^{-\epsilon_s / \tau} = \frac{2}{3V} \frac{1}{Z} \sum_s \epsilon_s e^{-\epsilon_s / \tau} = \frac{2U}{3V}$$

4. (Kittel 6.9) **Gas of atoms with internal degrees of freedom.** Consider an ideal monatomic gas, but one for which the atom has two internal energy states, one an energy Δ above the other. There are N atoms in volume V at temperature τ . Find the (a) chemical potential; (b) free energy; (c) entropy; (d) pressure; (e) heat capacity at constant pressure.

Solution.

$$N = \lambda \sum_{\epsilon} e^{-\beta \epsilon}$$

$$\epsilon = \frac{p^2}{2m} + \epsilon_{\text{int}}$$

$$\epsilon_{\text{int}} = 0 \Rightarrow \sum_{\epsilon} e^{-\beta \epsilon} = Z = V n_Q$$

$$\epsilon_{\text{int}} = \Delta \Rightarrow \sum_{\epsilon} e^{-\beta \epsilon} = e^{-\beta \Delta} Z = V n_Q e^{-\beta \Delta}$$

$$N = \lambda \sum_{\epsilon} e^{-\beta \epsilon} = \lambda V n_Q (1 + e^{-\beta \Delta})$$

$$\lambda = \frac{n}{n_Q} \frac{1}{1 + e^{-\beta \Delta}}$$

$$\lambda = e^{\mu / \tau}$$

$$\mu = \tau \ln \frac{n}{n_Q} - \tau \ln (1 + e^{-\beta \Delta})$$

(b)

$$Z_1 = (1 + e^{-\beta \Delta}) Z_{ideal} = (1 + e^{-\beta \Delta}) V n_Q$$

$$Z = \frac{Z_1^N}{N!} = \frac{(V n_Q)^N}{N!} (1 + e^{-\beta \Delta})^N$$

$$F = -\tau \ln Z = -\tau \ln \left(\frac{V n_Q}{N!} \right)^N - N \tau \ln (1 + e^{-\beta \Delta}) = N \tau \left(\ln \frac{n}{n_Q} - 1 \right) - N \tau \ln (1 + e^{-\beta \Delta})$$

(c)

$$\sigma = - \left(\frac{\partial F}{\partial \tau} \right)_V = N \left[\ln \frac{n_Q}{n} + \frac{5}{2} \right] + N \ln (1 + e^{-\beta \Delta}) + N \tau \frac{e^{-\beta \Delta}}{1 + e^{-\beta \Delta}} \frac{\Delta}{\tau^2}$$

$$\text{ie. } \sigma = N \left[\ln \frac{n_Q}{n} + \frac{5}{2} \right] + N \ln (1 + e^{-\beta \Delta}) + \frac{N \Delta}{\tau} \frac{1}{e^{\beta \Delta} + 1}$$

(d)

$$p = - \left(\frac{\partial F}{\partial V} \right)_\tau = \frac{N \tau}{V}$$

(e)

$$C_p = \tau \left(\frac{\partial \sigma}{\partial T} \right)_p = k_B \tau \left(\frac{\partial \sigma}{\partial \tau} \right)_p = k_B \tau \frac{\partial}{\partial \tau} N \left(\ln \frac{n_Q}{n} + \frac{5}{2} \right) + k_B \tau \frac{\partial}{\partial \tau} \left(N \ln (1 + e^{-\beta \Delta}) + \frac{N \Delta}{\tau} \frac{1}{e^{\beta \Delta} + 1} \right)$$

Hence,

$$C_p = \frac{5}{2} Nk_B + Nk_B \tau \left(\frac{e^{-\beta\Delta}}{1 + e^{-\beta\Delta}} \frac{\Delta}{\tau^2} - \frac{1}{e^{\beta\Delta} + 1} \frac{\Delta}{\tau^2} + \frac{\Delta}{\tau} \frac{e^{\beta\Delta}}{(e^{\beta\Delta} + 1)^2} \frac{\Delta}{\tau^2} \right) = \frac{5Nk_B}{2} + Nk_B \frac{\Delta^2}{\tau^2} \frac{e^{\beta\Delta}}{(e^{\beta\Delta} + 1)^2}$$

5. (Kittel 6.11) **Convective isentropic equilibrium of the atmosphere.** The lower 10-15 km of the atmosphere - the troposphere - is often in a convective steady state at constant entropy, not constant temperature. In such equilibrium pV^γ is independent of altitude, where $\gamma = C_p/C_v$. Use the condition of mechanical equilibrium in a uniform gravitational field to: (a) Show that $dT/dz = \text{constant}$, where z is the altitude. This quantity, important in meteorology, is called the dry adiabatic lapse rate. (Do not use the barometric pressure relation that was derived in Chapter 5 for an isothermal atmosphere.) (b) Estimate dT/dz , in degrees Celsius per km. Take $\gamma = 7/5$. (c) Show that $p \propto \rho^\gamma$, where ρ is the mass density. If the actual temperature gradient is greater than the isentropic gradient, the atmosphere may be unstable with respect to convection.

Solution.

(a)

$$[p(z) - p(z + dz)]A = nmAdzg \Rightarrow -\frac{dp}{dz} = nmg$$

$$\frac{dT}{dz} = \frac{dT}{dp} \frac{dp}{dz} = -\frac{dT}{dp} nmg \quad (\text{Eqn 1})$$

$$pV^\gamma = \text{constant, with } pV = Nk_B T \text{ gives } p \left(\frac{T}{p} \right)^\gamma = \text{constant and so } \frac{T^\gamma}{p^{\gamma-1}} = \text{constant}$$

$$\ln \frac{T^\gamma}{p^{\gamma-1}} = \text{constant} \Rightarrow \gamma \ln T - (\gamma - 1) \ln p = \text{constant}$$

$$\frac{\gamma}{T} dT - (\gamma - 1) \frac{dp}{p} = 0$$

$$\frac{dT}{dp} = \frac{(\gamma - 1)T}{\gamma p} \quad (\text{Eqn 2})$$

Eqn 2 can also be obtained from

$$Td\sigma = dU + pdV = c_v dT + pdV$$

For isentropic processes $d\sigma = 0$ and so

$$C_v dT + pdV = 0$$

$$pdV + Vdp = Nk_B dT$$

$$C_v dT + Nk_B dT - Vdp = 0$$

$$C_p dT - \frac{Nk_B T}{p} dp = 0$$

(since $\gamma = \frac{C_p}{C_v}$ and $Nk_B = C_p - C_v$).

$$\gamma dT - (\gamma - 1) \frac{T dp}{p} = 0 \Rightarrow \frac{dT}{dp} = \frac{\gamma - 1}{\gamma} \frac{T}{p}$$

Substitute Eqn 2 into Eqn 1 to get

$$\frac{dT}{dz} = -\frac{\gamma - 1}{\gamma} \frac{T p}{p} mg$$

Now $p = nk_B T$ and so

$$\frac{dT}{dz} = -\frac{\gamma - 1}{\gamma} \frac{mg}{k_B} = \text{constant}$$

(b)

Take m the mass of N_2

$$m = 28 \times 1.67 \times 10^{-27} \text{ kg} = 4.68 \times 10^{-26} \text{ kg}$$

Now $\gamma = \frac{7}{5}$ and so

$$\frac{dT}{dz} = -\frac{7/5 - 1}{7/5} \frac{4.68 \times 10^{-26}}{1.38 \times 10^{-23}} 9.8 K/m = 9.5 \times 10^{-3} K/m$$

T drops about 10 degrees Celsius every km.

(c)

$$pV^\gamma = \text{constant}, \text{ and } \rho \propto V^{-1} \Rightarrow p\rho^{-\gamma} = \text{constant}, \text{ and so } p \propto \rho^\gamma$$

6. (Kittel 6.12) ***Ideal gas in two dimensions.*** (a) Find the chemical potential of an ideal monatomic gas in two dimensions, with N atoms confined to a square of area $A = L^2$. The spin is zero. (b) Find an expression for the energy U of the gas. (c) find an expression for the entropy σ . The temperature is τ .

Solution.

$$(a) \quad N = \lambda Z_1 \text{ with } Z_1 = \sum e^{-\beta\epsilon} = \int D(\epsilon) e^{-\beta\epsilon} d\epsilon.$$

$$D(\epsilon) d\epsilon = \frac{A 2\pi p dp}{h^2} = \frac{2\pi A}{h^2} \frac{1}{2} dp^2 = \frac{2\pi A m}{h^2} d\epsilon$$

$$Z_1 = \frac{2\pi A m}{h^2} \int_0^\infty e^{-\beta\epsilon} d\epsilon = \frac{2\pi A m}{h^2} \tau$$

$$N = e^{\mu/\tau} \frac{2\pi A m}{h^2} \tau \Rightarrow \mu = \tau \ln \frac{N h^2}{2\pi A m \tau}$$

(b)

$$\bar{\epsilon} = \frac{1}{2} k_B T + \frac{1}{2} k_B T = k_B T$$

$$U = N\bar{\epsilon} = N k_B T$$

(c)

$$Z = \frac{Z_1^N}{N!} \Rightarrow F = -\tau \ln Z = -\tau \left(N \ln \left(\frac{2\pi A m}{h^2} \tau \right) - N \ln N + N \right)$$

$$F=-\tau N\left(\ln\frac{2\pi Am\tau}{h^2N}+1\right)$$

$$\sigma=-\left(\frac{\partial F}{\partial \tau}\right)_A=N\left(\ln\frac{2\pi Am\tau}{h^2N}+1\right)+N\tau\frac{1}{\tau}=N\left(\ln\frac{2\pi Am\tau}{h^2N}+2\right)$$

$$U=F+\tau\sigma=N\tau$$