

PHY 410 Homework 11 solutions

1. K. + K. Ch 14, #1 Mean speeds in Maxwellian distrib.

Speed distribution: $P(v) = 4\pi \left(\frac{m}{2\pi\tau} \right)^{3/2} v^2 e^{-mv^2/2\tau}$

a) $\langle v^2 \rangle = \int_0^\infty v^2 P(v) dv = 4\pi \left(\frac{m}{2\pi\tau} \right)^{3/2} \int_0^\infty v^4 e^{-mv^2/2\tau} dv$

put exponent in dimensionless form: $u = \sqrt{\frac{m}{2\tau}} v \quad du = \sqrt{\frac{m}{2\tau}} dv$

$$\langle v^2 \rangle = 4\pi \left(\frac{m}{2\pi\tau} \right)^{3/2} \left(\frac{2\tau}{m} \right)^{5/2} \int_0^\infty u^4 e^{-u^2} du$$

From Appendix A, equations (5), (6), (9):

$$\int_0^\infty u^4 e^{-u^2} du = \frac{1}{2} \Gamma\left(\frac{5}{2}\right) = \frac{1}{2} \cdot \frac{3}{2} \cdot \frac{1}{2} \cdot \sqrt{\pi} = \frac{3\sqrt{\pi}}{8}$$

$$\langle v^2 \rangle = 4\pi \cdot \frac{2\tau}{m} \cdot \frac{1}{\pi^{3/2}} \cdot \frac{3\sqrt{\pi}}{8} = \frac{3\tau}{m}$$

$$v_{rms} = \sqrt{\langle v^2 \rangle} = \sqrt{\frac{3\tau}{m}}$$

b) The most probable speed is where $P(v)$ has its maximum

$$0 = \frac{d}{dv} P(v) = 4\pi \left(\frac{m}{2\pi\tau} \right)^{3/2} \left[2v e^{-\frac{mv^2}{2\tau}} + v^2 \left(-\frac{mv}{\tau} \right) e^{-\frac{mv^2}{2\tau}} \right]$$

$$0 = 2v - \frac{mv^3}{\tau}$$

$$2 = \frac{mv^2}{\tau} \Rightarrow v_{mp} = \sqrt{\frac{2\tau}{m}}$$

c) Mean speed $\bar{c} = \int_0^{\infty} v P(v) dv$

$$\bar{c} = 4\pi \left(\frac{m}{2\pi\tau} \right)^{3/2} \cdot \int_0^{\infty} v^3 e^{-\frac{mv^2}{2\tau}} dv$$

$u = \sqrt{\frac{m}{2\tau}} v$
as before

$$= 4\pi \left(\frac{m}{2\pi\tau} \right)^{3/2} \cdot \left(\frac{2\tau}{m} \right)^2 \cdot \int_0^{\infty} u^3 e^{-u^2} du$$

Appendix A again: $\int_0^{\infty} u^3 e^{-u^2} du = \frac{1}{2} \Gamma(2) = \frac{1}{2}$

$$\bar{c} = 4\pi \left(\frac{2\tau}{m} \right)^{1/2} \frac{1}{\pi^{3/2}} \cdot \frac{1}{2} = 2 \left(\frac{2\tau}{\pi m} \right)^{1/2} = \underline{\underline{\left(\frac{8\tau}{\pi m} \right)^{1/2}}}$$

d) From class, $P(v_z) = \frac{1}{n} a(v_z) = \left(\frac{m}{2\pi\tau} \right)^{1/2} e^{-\frac{mv_z^2}{2\tau}}$

$$\langle |v_z| \rangle = \int_{-\infty}^{\infty} |v_z| P(v_z) dv_z = 2 \int_0^{\infty} v_z P(v_z) dv_z$$

$$= 2 \left(\frac{m}{2\pi\tau} \right)^{1/2} \int_0^{\infty} v_z e^{-\frac{mv_z^2}{2\tau}} dv_z$$

again $u = \sqrt{\frac{m}{2\tau}} v_z$

$$= 2 \left(\frac{m}{2\pi\tau} \right)^{1/2} \cdot \left(\frac{2\tau}{m} \right) \int_0^{\infty} u e^{-u^2} du$$

$$\int_0^{\infty} u e^{-u^2} du = \frac{1}{2}$$

$$= 2 \cdot \left(\frac{2\tau}{m} \right)^{1/2} \frac{1}{\pi^{1/2}} \cdot \frac{1}{2} = \underline{\underline{\left(\frac{2\tau}{\pi m} \right)^{1/2}}}$$

↑
Appendix A
again

2. If a molecule travels a distance L in a gas, the probability not to scatter is

$$P(L) = e^{-L/l}, \text{ where } l = \text{mean free path}$$

For this problem, $L = 20 \text{ cm} = 0.20 \text{ m}$, $P(L) = 0.9$

$$\ln P(L) = -\frac{L}{l}$$

$$l = \frac{-L}{\ln P(L)} = \frac{-0.20 \text{ m}}{\ln 0.9} = \frac{-0.20 \text{ m}}{-0.105} = 1.9 \text{ m}$$

In class we showed that $l = \frac{1}{\sqrt{2} n \sigma_c}$

$$n = \frac{1}{\sqrt{2} l \sigma_c} = \frac{1}{\sqrt{2} \cdot 1.9 \text{ m} \cdot 10^{-19} \text{ m}^2} = 3.7 \cdot 10^{18} \text{ m}^{-3}$$

Ideal gas: $pV = N\tau \rightarrow p = \frac{N}{V}\tau = n k_B T$

$$p = 3.7 \cdot 10^{18} \text{ m}^{-3} \cdot 1.38 \cdot 10^{-23} \text{ J/K} \cdot 300 \text{ K} = \underline{\underline{1.5 \cdot 10^{-2} \text{ Pa}}}$$

Many experimentalists still measure pressure in Torr

$$10^5 \text{ Pa} = 760 \text{ Torr}$$

$$p = 1.5 \cdot 10^{-2} \text{ Pa} \left(\frac{760 \text{ Torr}}{10^5 \text{ Pa}} \right) = 1.1 \cdot 10^{-4} \text{ Torr}$$

$$P_1 = N \frac{\tau_1}{V_1}$$

| | τ_2 | P_2 | ΔU | $\Delta \sigma$ | W | Q |
|------------|--------------------|-------------------|---|-----------------|---|------------------|
| Isothermal | τ_1 | $\frac{1}{3} P_1$ | 0 | $N \ln 3$ | $N \tau_1 \ln 3$ | $N \tau_1 \ln 3$ |
| Isentropic | $\tau_1 / 3^{2/3}$ | $P_1 / 3^{5/3}$ | $\frac{3}{2} N \tau_1 \left(\frac{1}{3^{2/3}} - 1 \right)$ | 0 | $\frac{3}{2} N \tau_1 \left(1 - \frac{1}{3^{2/3}} \right)$ | 0 |
| Sudden | τ_1 | $\frac{1}{3} P_1$ | 0 | $N \ln 3$ | 0 | 0 |

a) Isothermal: $\tau_2 = \tau_1$, $U = U(\tau) \Rightarrow \Delta U = 0$

Ideal gas $pV = N\tau$, $V_2 = 3V_1 \Rightarrow P_2 = \frac{1}{3} P_1$

$$W = \int_{V_1}^{V_2} p dV = \int_{V_1}^{V_2} \frac{N\tau}{V} dV = N\tau_1 \ln \frac{V_2}{V_1} = N\tau_1 \ln 3$$

1st Law: $\Delta U = Q - W$, $\Delta U = 0 \Rightarrow Q = W = N\tau_1 \ln 3$

$$\sigma = N \left[\ln \frac{n_Q}{n} + \frac{5}{2} \right] = N \left[\ln \left(\frac{V n_Q}{N} \right) + \frac{5}{2} \right]$$

$$\sigma_2 - \sigma_1 = N \ln \frac{V_2}{V_1} = N \ln 3 \quad \text{since } \tau \text{ is constant}$$

Alternative method: $d\sigma = \frac{dQ}{\tau} \Rightarrow \Delta\sigma = \frac{Q}{\tau} = N \ln 3$

no integral necessary since τ is constant

b) Isentropic: $\sigma_2 = \sigma_1 \Rightarrow \Delta\sigma = 0$

$$dQ = \tau d\sigma = 0 \Rightarrow Q = 0$$

continued

Monatomic gas has $\gamma = \frac{5}{3}$, $\tau V^{\gamma-1} = \tau V^{\frac{2}{3}} = \text{constant}$
 $\tau_1 V_1^{\frac{2}{3}} = \tau_2 V_2^{\frac{2}{3}} \rightarrow \tau_2 = \tau_1 \left(\frac{V_1}{V_2} \right)^{\frac{2}{3}} = \frac{\tau_1}{3^{\frac{2}{3}}}$

$$p V^{\gamma} = \text{constant} \rightarrow p_2 = p_1 \left(\frac{V_1}{V_2} \right)^{\frac{5}{3}} = \frac{p_1}{3^{\frac{5}{3}}}$$

$$\Delta U = U_f - U_i = \frac{3}{2} N \tau_2 - \frac{3}{2} N \tau_1 = \frac{3}{2} N \tau_1 \left(\frac{1}{3^{\frac{2}{3}}} - 1 \right)$$

$$\Delta U = Q - W, \quad Q = 0 \Rightarrow W = -\Delta U$$

c) Sudden: The gas does no work and absorbs no heat: $W = 0, \quad Q = 0$

$$\Delta U = Q - W = 0 \Rightarrow \tau \text{ is constant}$$

$$\text{Ideal gas law } p = \frac{N\tau}{V} \Rightarrow p_2 = \frac{1}{3} p_1$$

$$\sigma = N \left[\ln \left(\frac{V m_Q}{N} \right) + \frac{5}{2} \right] \quad \tau \text{ doesn't change}$$

$$\Delta \sigma = N \ln \frac{V_2}{V_1} = N \ln 3$$

Alternative method: Multiplicity increases by 3^N

$$\Delta \sigma = \ln \frac{g_2}{g_1} = \ln 3^N = N \ln 3$$

4. Initial state: 10 g water at $T_{1i} = 10^\circ\text{C}$

30 g water at $T_{2i} = 90^\circ\text{C}$

a) The two systems exchange heat, ending at a common temperature T_f .

$$Q_1 = -Q_2$$

$$m_1 c_p (T_f - T_{1i}) = -m_2 c_p (T_f - T_{2i})$$

$$T_f (m_1 + m_2) = m_2 T_{2i} + m_1 T_{1i}$$

$$T_f = \frac{m_1 T_{1i} + m_2 T_{2i}}{m_1 + m_2} = \frac{10\text{g} \cdot 10^\circ\text{C} + 30\text{g} \cdot 90^\circ\text{C}}{40\text{g}} = \underline{\underline{70^\circ\text{C}}}$$

b) Let the systems exchange heat slowly, so we can use $dS = \frac{dQ}{dT}$ with $dQ = m c_p dT$

$$\Delta S = \Delta S_1 + \Delta S_2 = \int_{T_{1i}}^{T_f} \frac{dQ_1}{T} + \int_{T_{2i}}^{T_f} \frac{dQ_2}{T}$$

$$= c_p \left[m_1 \int_{10^\circ\text{C}}^{70^\circ\text{C}} \frac{dT}{T} + m_2 \int_{90^\circ\text{C}}^{70^\circ\text{C}} \frac{dT}{T} \right]$$

Must convert temperatures to Kelvin
↓

$$= 4.19 \frac{\text{J}}{\text{gK}} \left[10\text{g} \cdot \ln \frac{343}{283} + 30\text{g} \cdot \ln \frac{343}{363} \right]$$

$$= 4.19 \frac{\text{J}}{\text{gK}} [1.923\text{g} \quad 1.700\text{g}] = \underline{\underline{0.93 \frac{\text{J}}{\text{K}}}}$$

5. Spins in a \vec{B} field; 2 states for each spin

$$\vec{m}_{\downarrow} = -m_0 \hat{e}_z$$

$$\vec{m}_{\uparrow} = m_0 \hat{e}_z$$

$$E_{\downarrow} = +m_0 B$$

$$E_{\uparrow} = -m_0 B$$

a) Probabilities: $P_{\uparrow} = \frac{e^{+m_0 B/\tau}}{e^{+m_0 B/\tau} + e^{-m_0 B/\tau}}$ $P_{\downarrow} = \frac{e^{-m_0 B/\tau}}{e^{+m_0 B/\tau} + e^{-m_0 B/\tau}}$

$$\langle \vec{m} \rangle = \sum_{j=1}^2 \vec{m}_j P_j = \frac{m_0 \hat{e}_z e^{+m_0 B/\tau} - m_0 \hat{e}_z e^{-m_0 B/\tau}}{e^{+m_0 B/\tau} + e^{-m_0 B/\tau}}$$

$$\langle \vec{m} \rangle = m_0 \hat{e}_z \tanh\left(\frac{m_0 B}{\tau}\right)$$

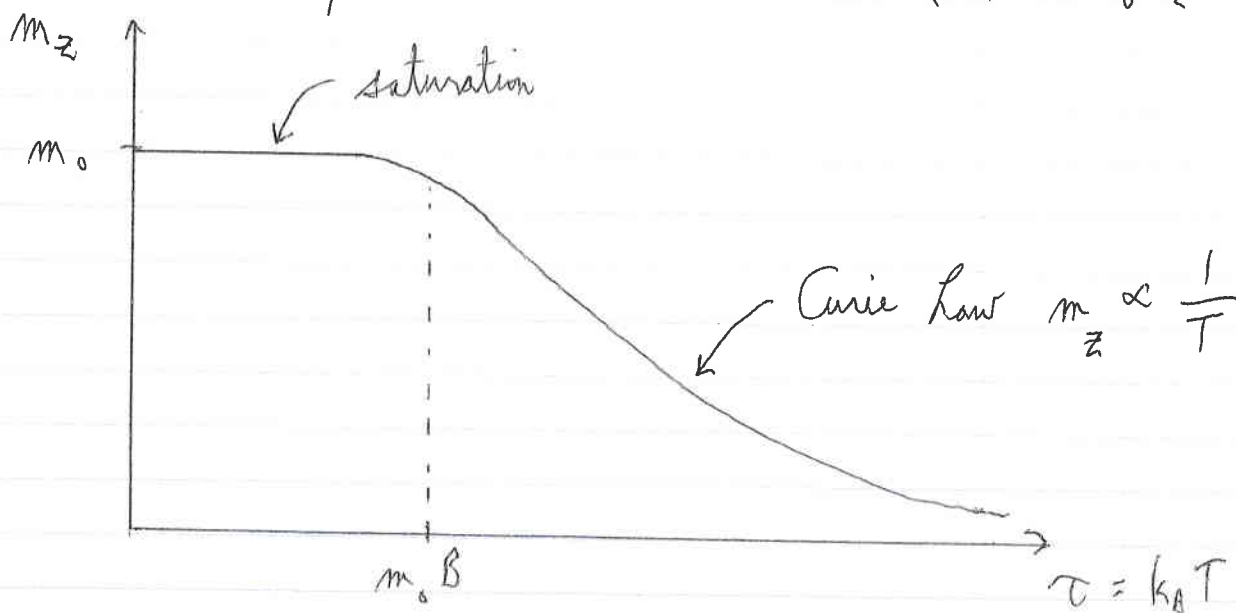
b) This part is tricky:

high temperature: $\tanh x \sim x$

$$\langle \vec{m} \rangle \approx m_0 \hat{e}_z \cdot \frac{m_0 B}{k_B T}$$

low temperature: $\tanh x \sim 1$

$$\langle \vec{m} \rangle = m_0 \hat{e}_z$$



$$6. \quad S = A V^{1/3} U^{1/2}$$

I am using U instead of E for "internal energy".

With N fixed, the Thermodynamic Identity is

$$dU = T dS - p dV$$

$$\text{or} \quad dS = \frac{1}{T} dU + \frac{p}{T} dV$$

The latter form is for $S = S(U, V)$, which is the situation for a Microcanonical Ensemble. You can see the derivatives directly:

$$(1): \frac{1}{T} = \left(\frac{\partial S}{\partial U} \right)_V \quad (2): \frac{p}{T} = \left(\frac{\partial S}{\partial V} \right)_U$$

$$(1): \frac{1}{T} = \frac{1}{2} A V^{1/3} U^{-1/2} \quad (2): \frac{p}{T} = \frac{1}{3} A V^{-2/3} U^{1/2}$$

$$\text{Solve (1) for } U^{1/2}: \quad U^{1/2} = \frac{1}{2} A V^{1/3} T$$

$$\text{Plug into (2):} \quad \frac{p}{T} = \frac{1}{3} A V^{-2/3} \cdot \frac{1}{2} A V^{1/3} T$$

$$p = \frac{1}{6} A^2 V^{-1/3} T^2 \quad \text{or} \quad \underline{\underline{p V^{1/3} = \frac{A^2}{6} T^2}}$$

7. Ideal Gas with $S = Nk_B \ln(VT^{5/2}) + \text{constant}$

Gas expands adiabatically from T_1, V_1 to T_2, V_2 .

a) Adiabatic = isentropic $\Rightarrow \Delta S = 0$

b) $S = \text{constant}$

$$\Rightarrow V_1 T_1^{5/2} = V_2 T_2^{5/2}$$

$$\left(\frac{T_2}{T_1}\right)^{5/2} = \frac{V_1}{V_2}$$

$$\frac{T_2}{T_1} = \left(\frac{V_1}{V_2}\right)^{2/5} \Rightarrow T_2 = T_1 \left(\frac{V_1}{V_2}\right)^{2/5}$$

Optional: $\Delta T = T_2 - T_1 = T_1 \left[\left(\frac{V_1}{V_2}\right)^{2/5} - 1 \right]$

c) Ideal gas $pV = Nk_B T$ $p = \frac{Nk_B T}{V} \propto \frac{T}{V}$

$$\frac{p_2}{p_1} = \frac{\left(\frac{T_2}{V_2}\right)}{\left(\frac{T_1}{V_1}\right)} = \frac{T_2}{T_1} \frac{V_1}{V_2} = \left(\frac{V_1}{V_2}\right)^{2/5} \cdot \left(\frac{V_1}{V_2}\right) = \left(\frac{V_1}{V_2}\right)^{7/5}$$

d) S is extensive: If $N \rightarrow 2N$, $V \rightarrow 2V$, $T \rightarrow T$, we must have $S \rightarrow 2S$. We can satisfy that by writing

$$\underline{S = Nk_B \ln\left(\frac{V}{N} T^{5/2}\right) + \text{constant}}$$

The missing term was $-Nk_B \ln N$