

## Spatial Detectors

- ◆ “Count ‘hits’”
- ◆ Spatial (or temporal) distribution
- ◆ “0”, “1”, “2” dimensional detectors



Detection system →

Particle to be detected

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## Quantum Efficiency

<p>Baseball: Batting Average = hits / at bats</p> 	<p>Particle detecton: Quantum efficiency= detected / incident quanta</p> 
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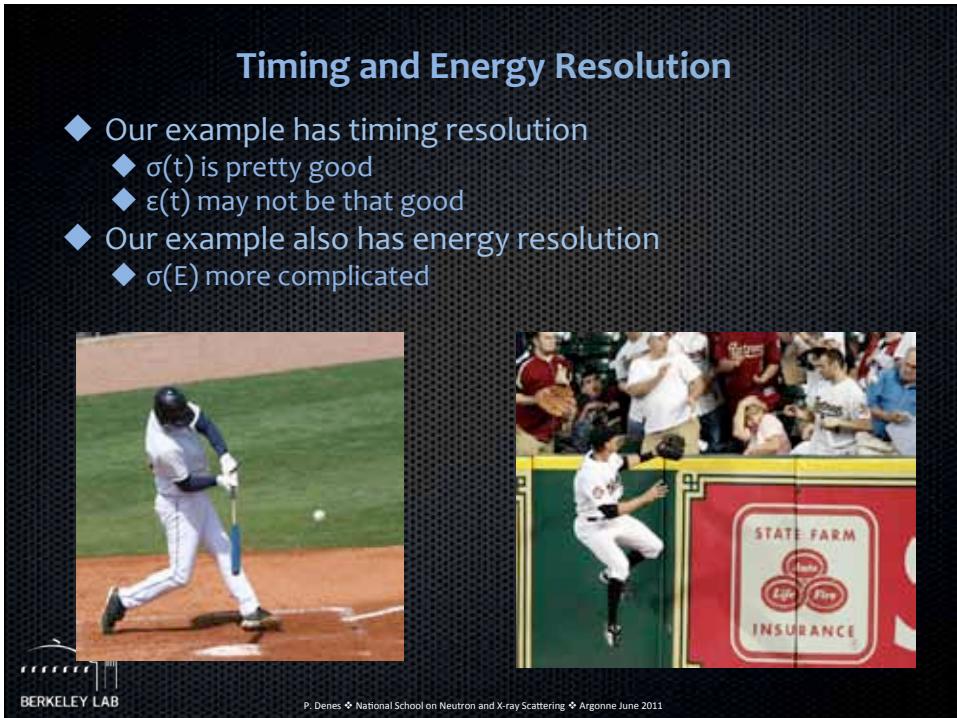
Note that the Q.E. may  
 depend on the energy  
 of the incident quanta  
 (we'll come back to this)

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## Timing and Energy Resolution

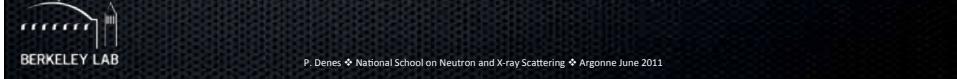
- ◆ Our example has timing resolution
  - ◆  $\sigma(t)$  is pretty good
  - ◆  $\epsilon(t)$  may not be that good
- ◆ Our example also has energy resolution
  - ◆  $\sigma(E)$  more complicated

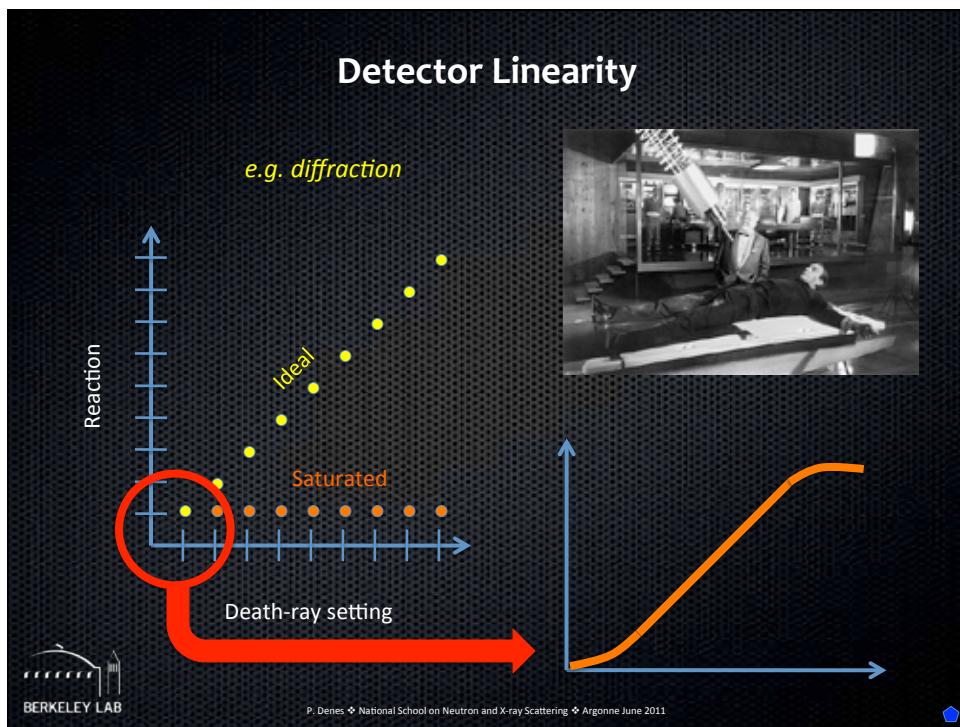
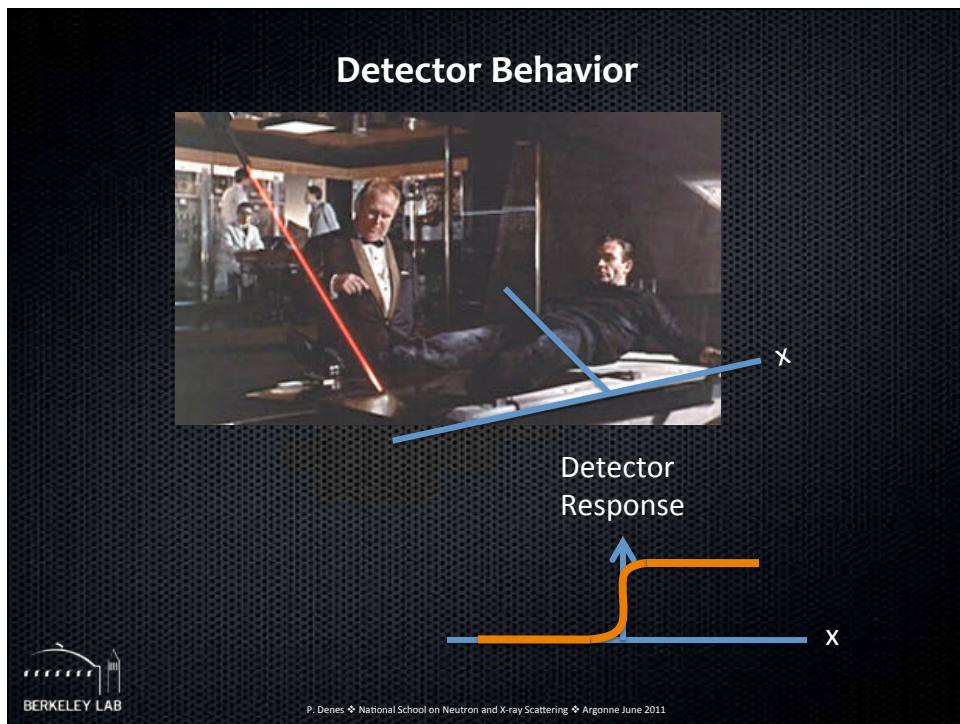


## Calorimetric Photon Detector



Calorimetric detector: absorbed energy measured by change of temperature  
(more generally, “calorimeters” measure total absorbed energy)





## Spatial Detector Properties

A “point” detector (“0D”)  
Responds to hits in sensitive area



No way to know where in the sensitive area the hit occurred

There may be additional information



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## Day-to-day oD Detector Example

Airport (pulsed induction) metal detector



“yes / no” – along with additional information




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### Day-to-day 1D Example



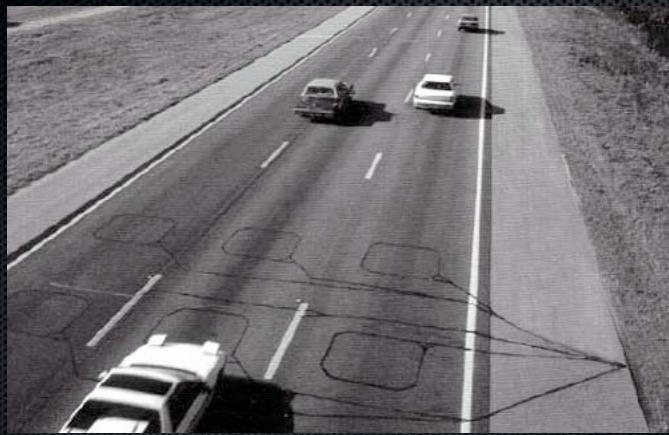
Theory

Experiment



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### Day-to-day 2D Detector Example

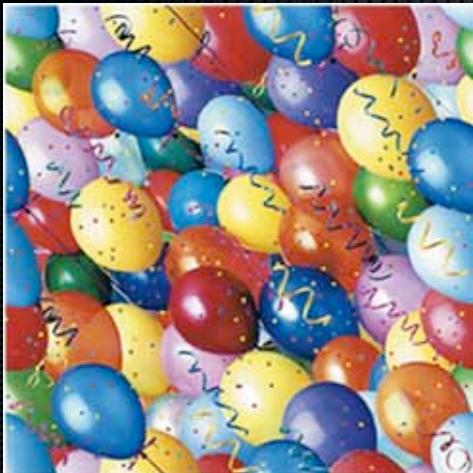


$$v = \Delta x / \Delta t$$



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## An Example 2D Detector



- ◆ 2D arrangement of our oD detector elements
- ◆ Which are quite non-linear
- ◆ Arranged in random sizes and orientations
- ◆ But with each element very small



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## Early X-ray Detection

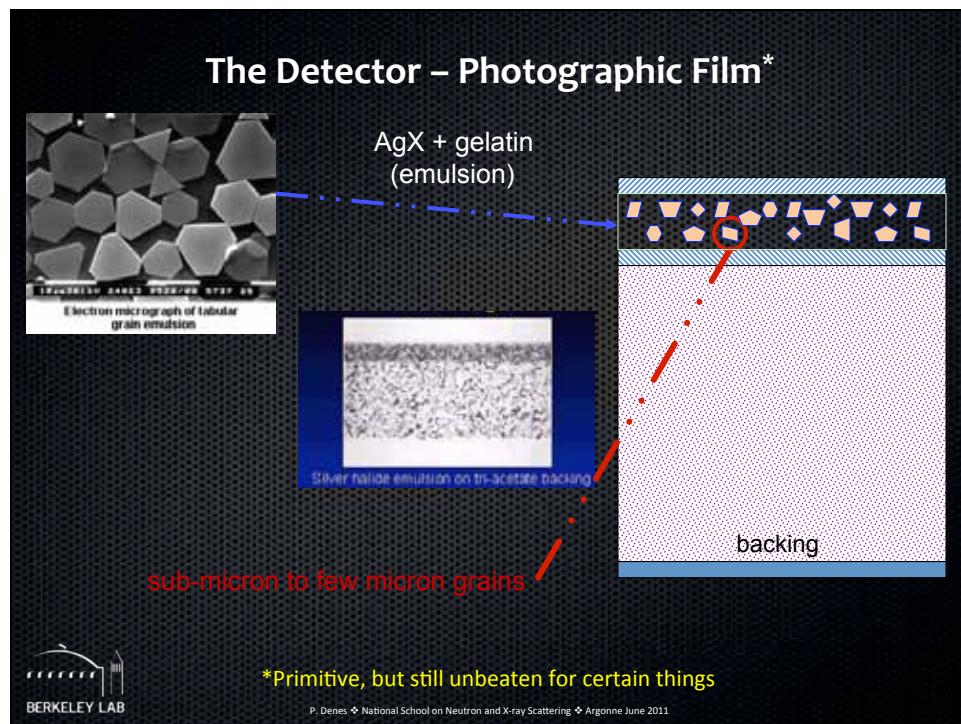
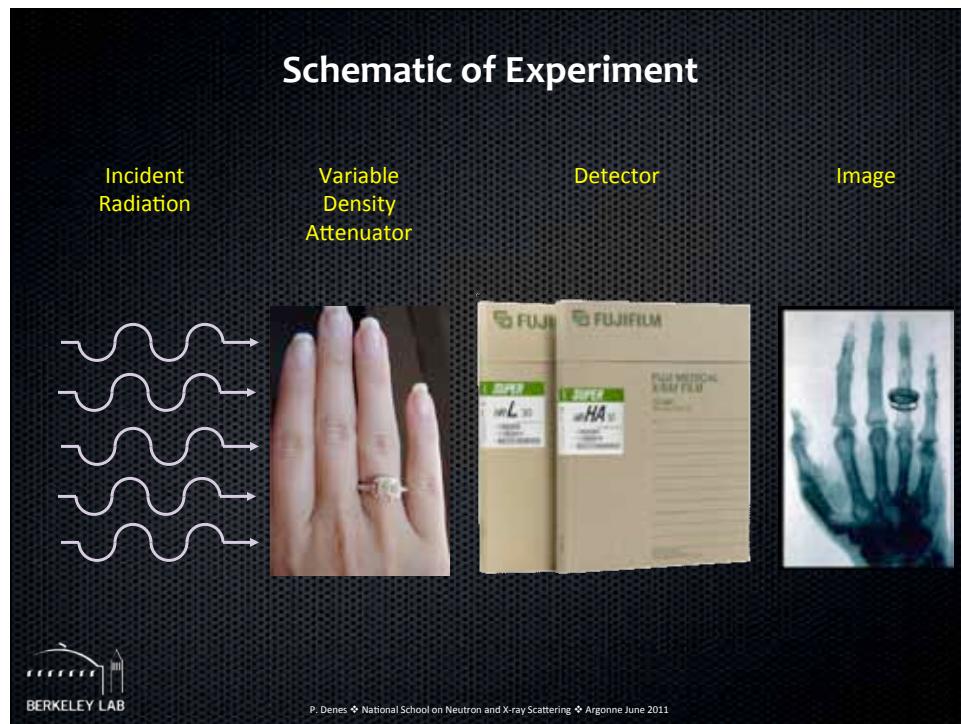
Herr Röntgen

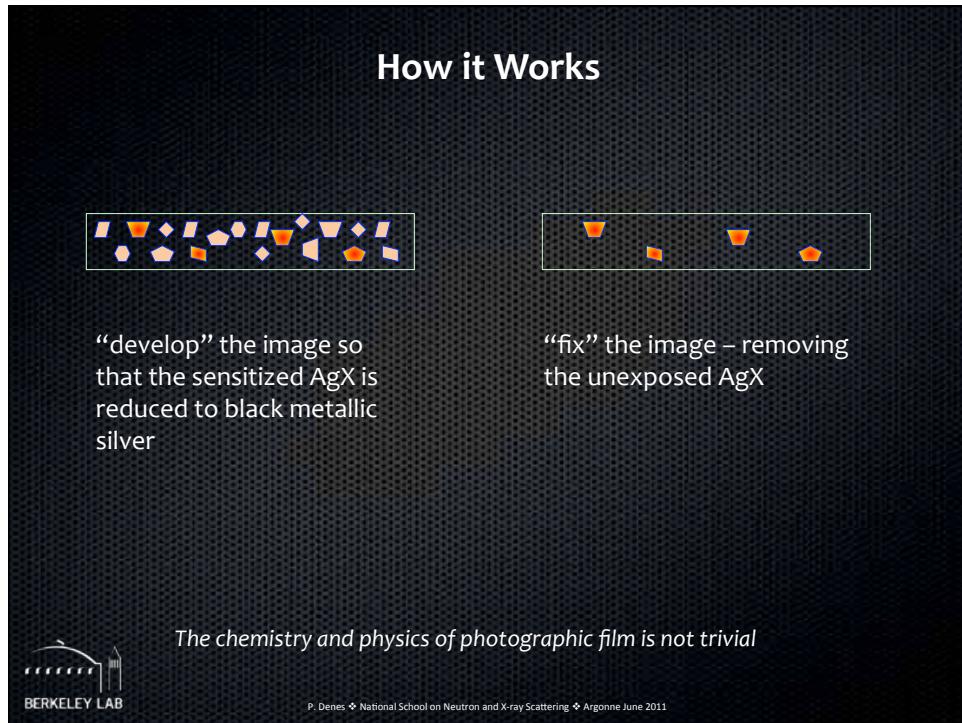
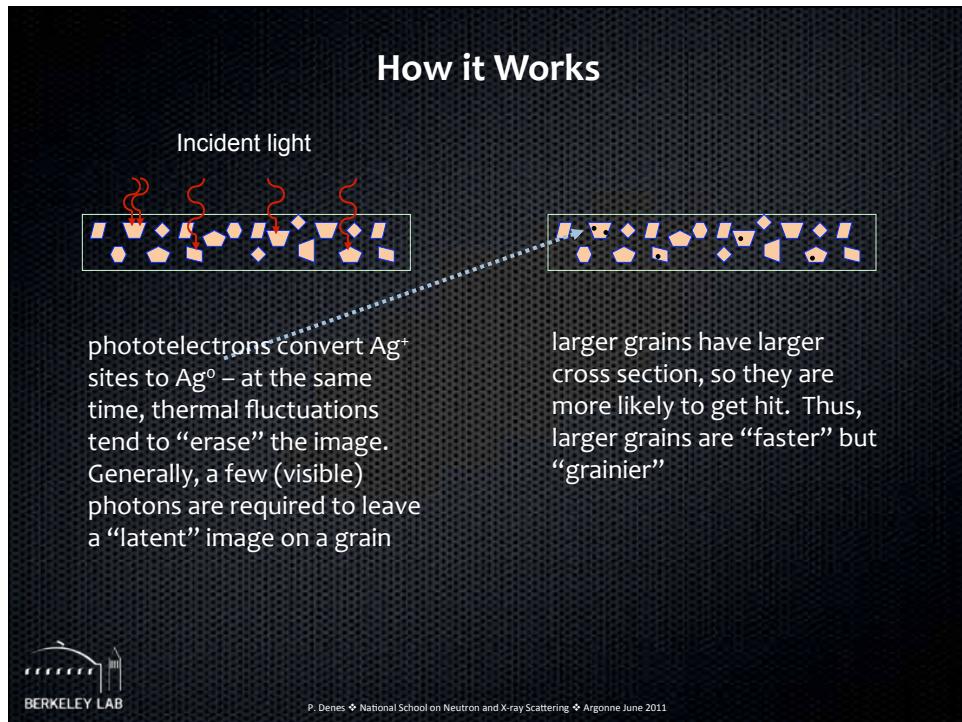


Frau Röntgen



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## Spatial Imaging Characteristics – PSF



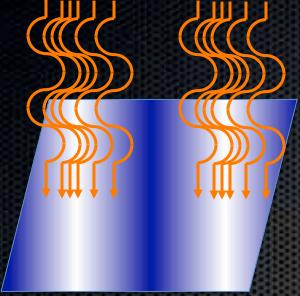
- ◆ Point Spread Function
- ◆  $\delta$ -function input
- ◆  $PSF(x_0, y_0, x, y)$
- ◆ Image is convolution of input at PSF
- ◆ “Black box” PSF includes all effects that might broaden or scatter the input



  
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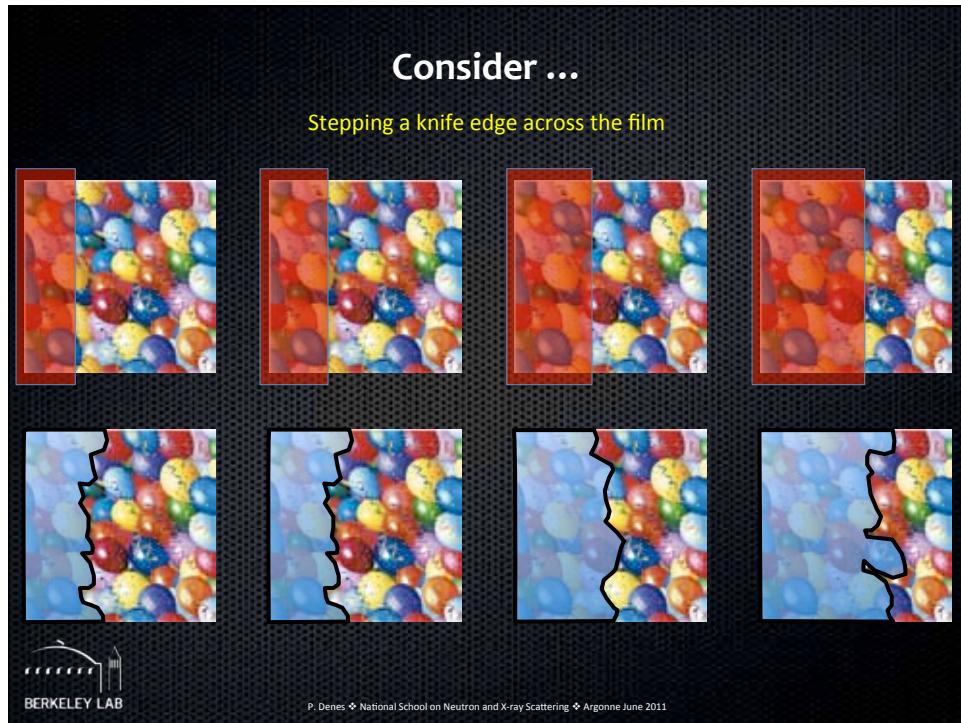
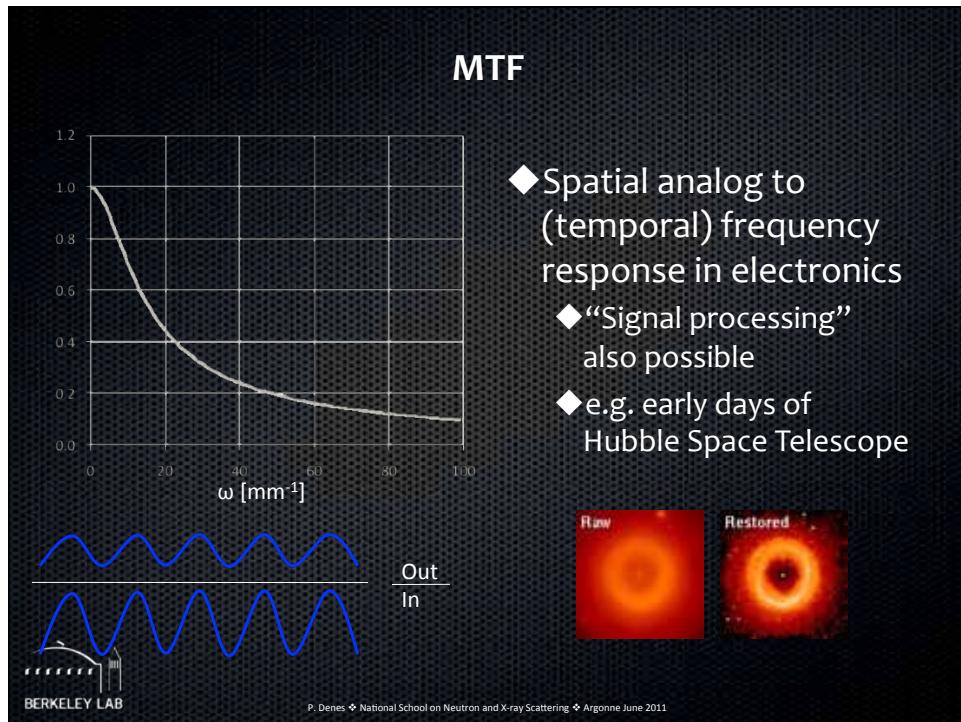
## Spatial Imaging Characteristics – MTF



- ◆ Modulation Transfer Function
- ◆  $\sin \omega x$  input
- ◆  $MTF(\omega)$
- ◆  $MTF(\omega_x, \omega_y)$
- ◆  $MTF = | FT(PSF) |$
- ◆ Related to **contrast**

  
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## Spatial Detector Concepts

- ◆ Quantum Efficiency
- ◆ Active area
- ◆ Contrast (PSF, MTF)
- ◆ Spatial (frequency dependence)

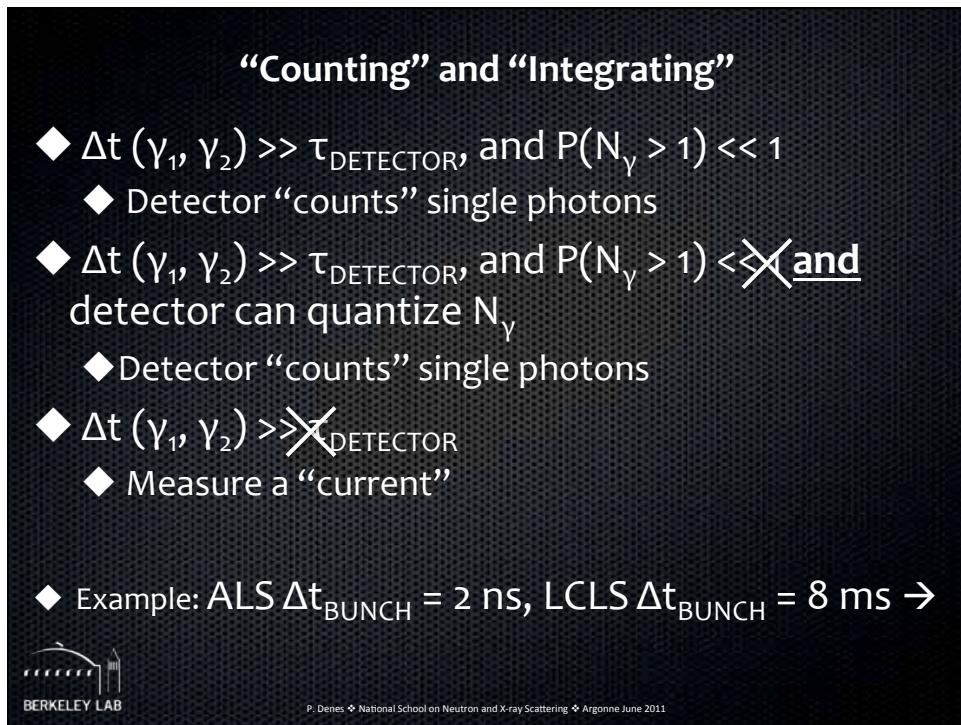
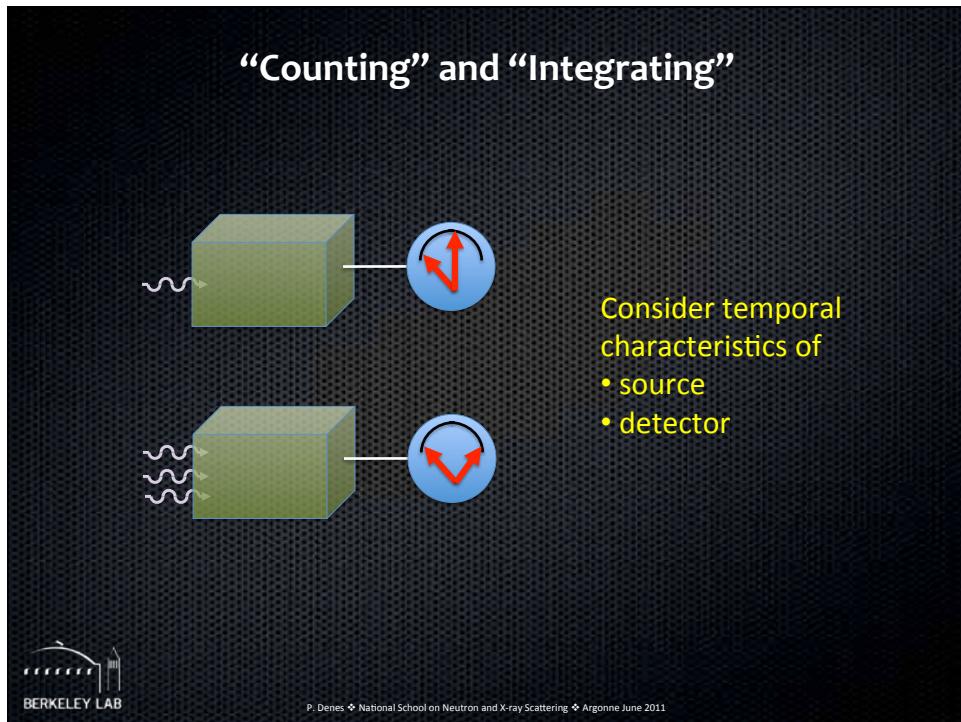
PSF = 0                    PSF = 1%                    PSF = 5%

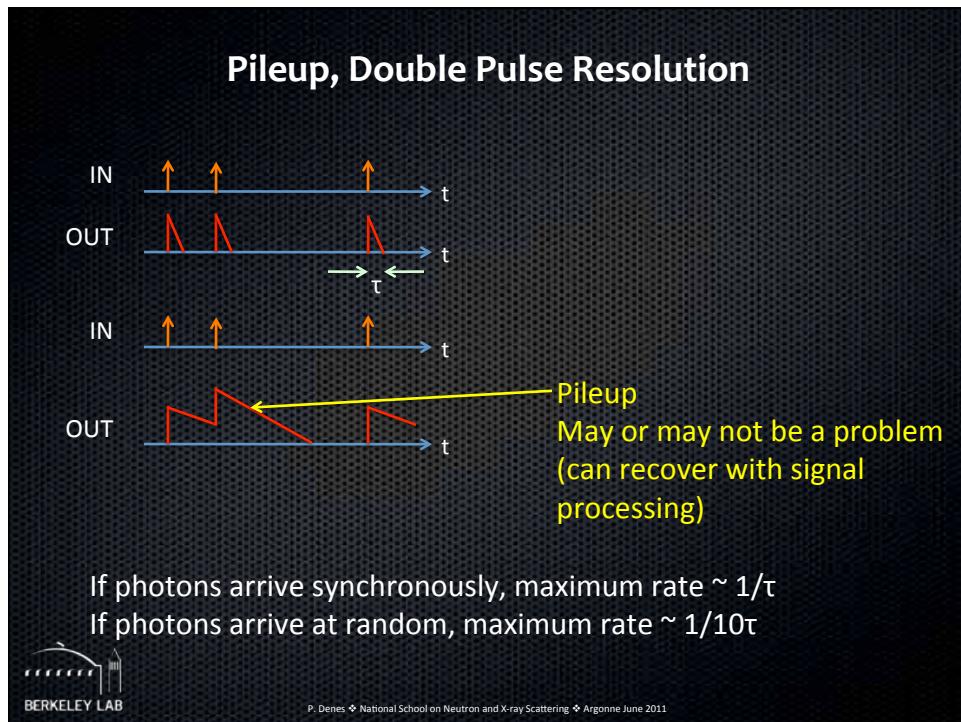
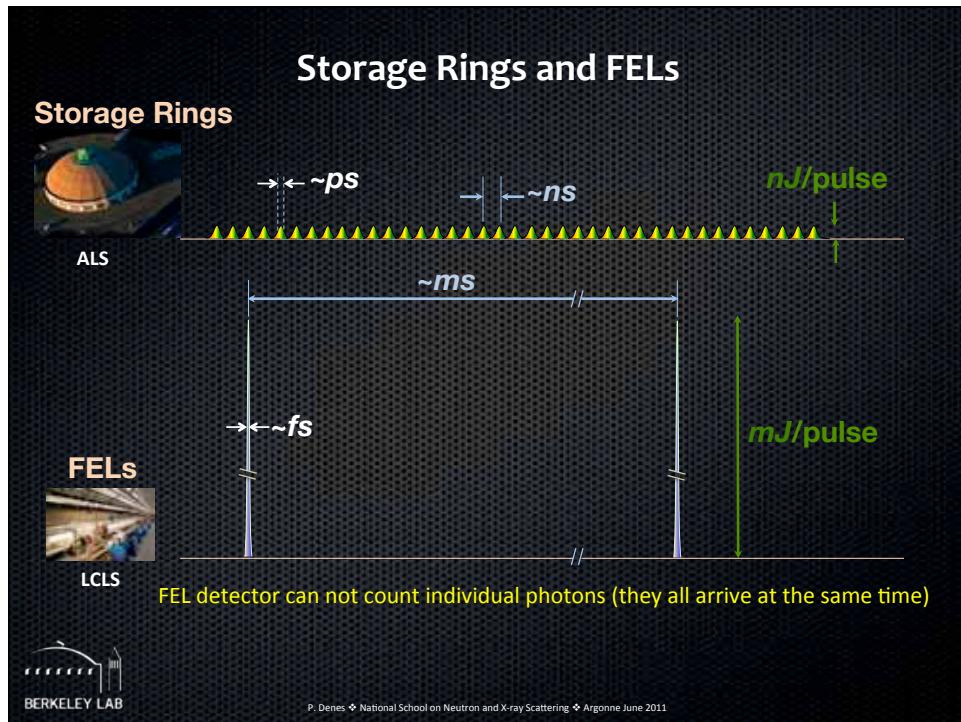
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## Detector Temporal Response

Pulsed Operation

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## Look Further Into “Detector”

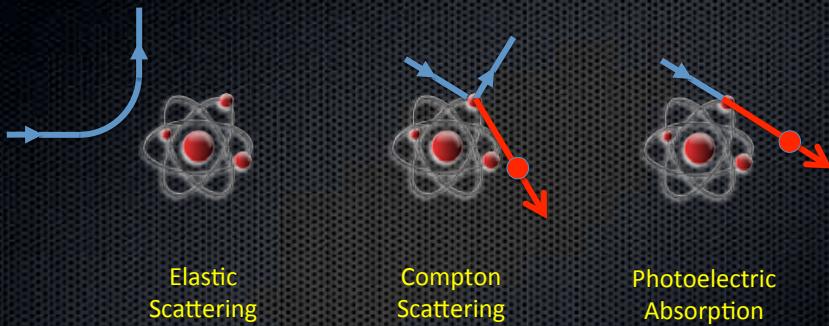
- ◆ Rarely does a (practical) photon detector actually detect photons
- ◆ Generally the photon is converted into one (or more) secondary particles
- ◆ Those secondary particles (usually electrons) are then detected, or create tertiary particles which are detected
- ◆  $\gamma \rightarrow e^-$



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## X-ray Interaction in Detector

Practically speaking, 3 possibilities:



Elastic  
Scattering

Compton  
Scattering

Photoelectric  
Absorption

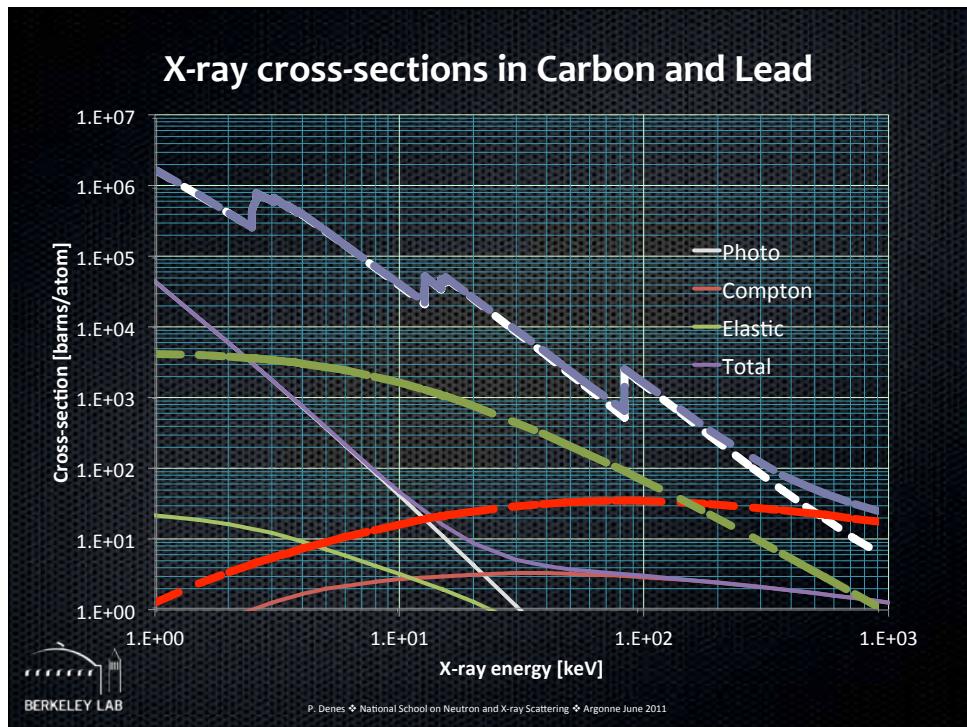
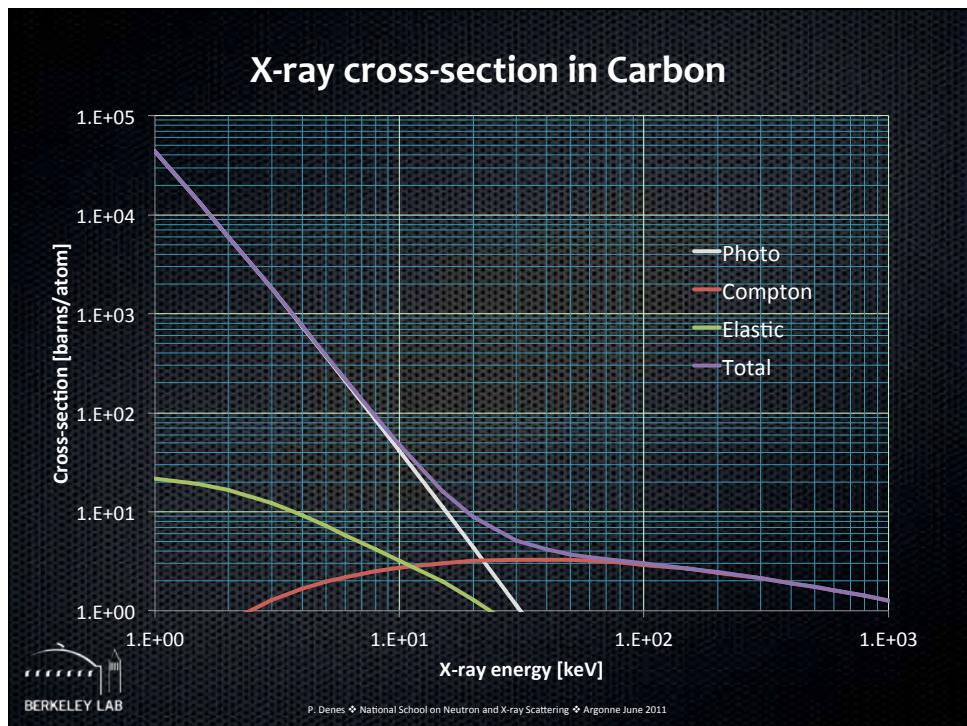
$$E_e \neq E_\gamma$$

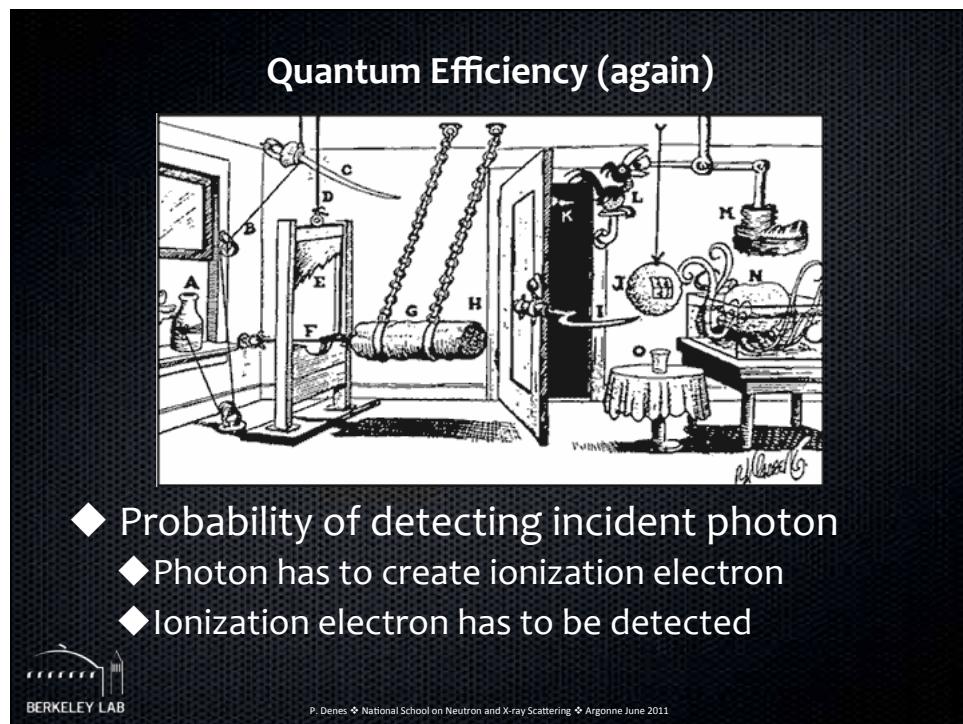
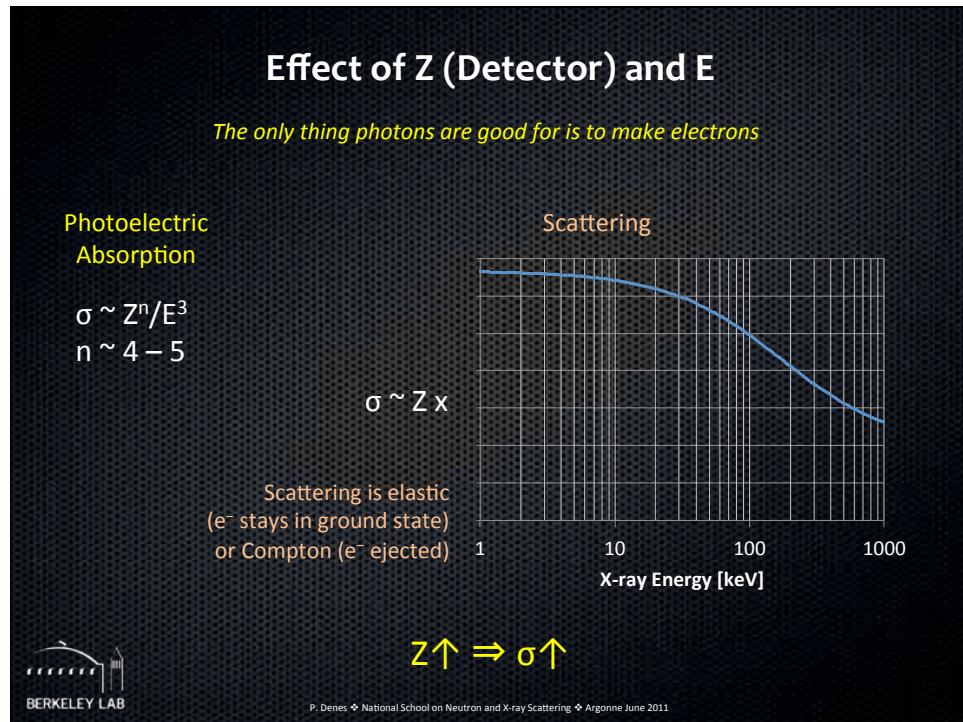
$$E_e = E_\gamma$$

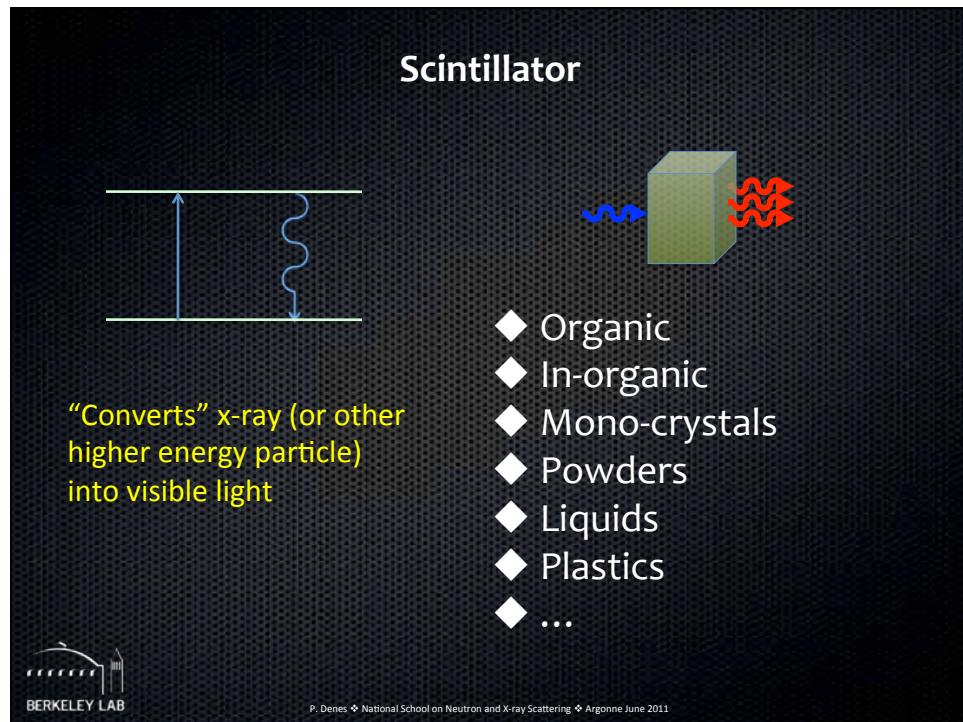
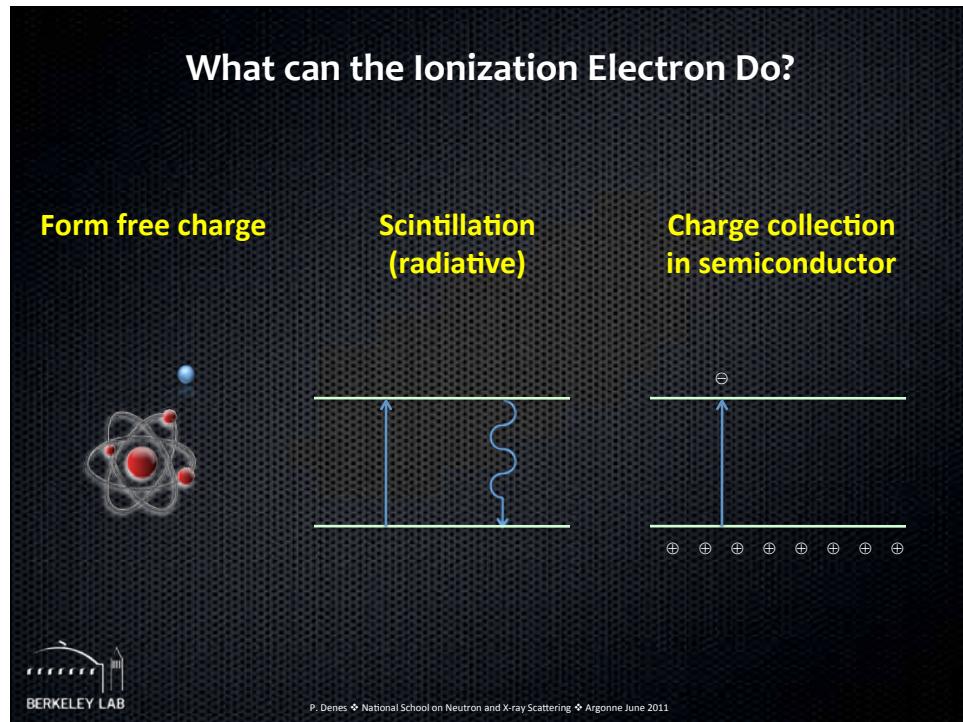
*Electron range (very crudely)  $R [\mu\text{m}] \approx E [\text{keV}]$*

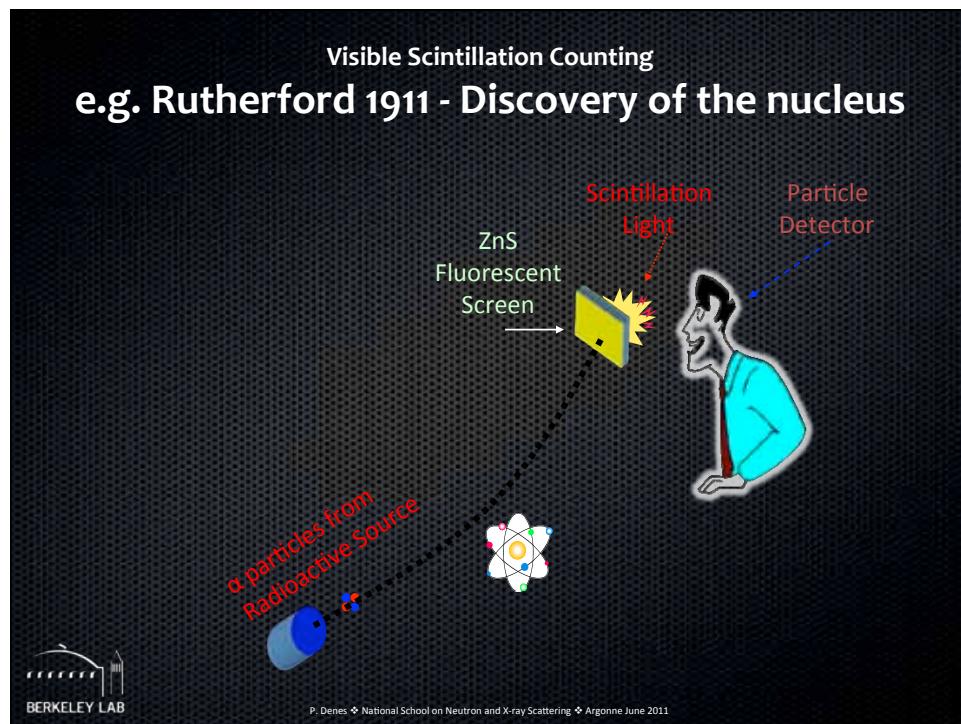
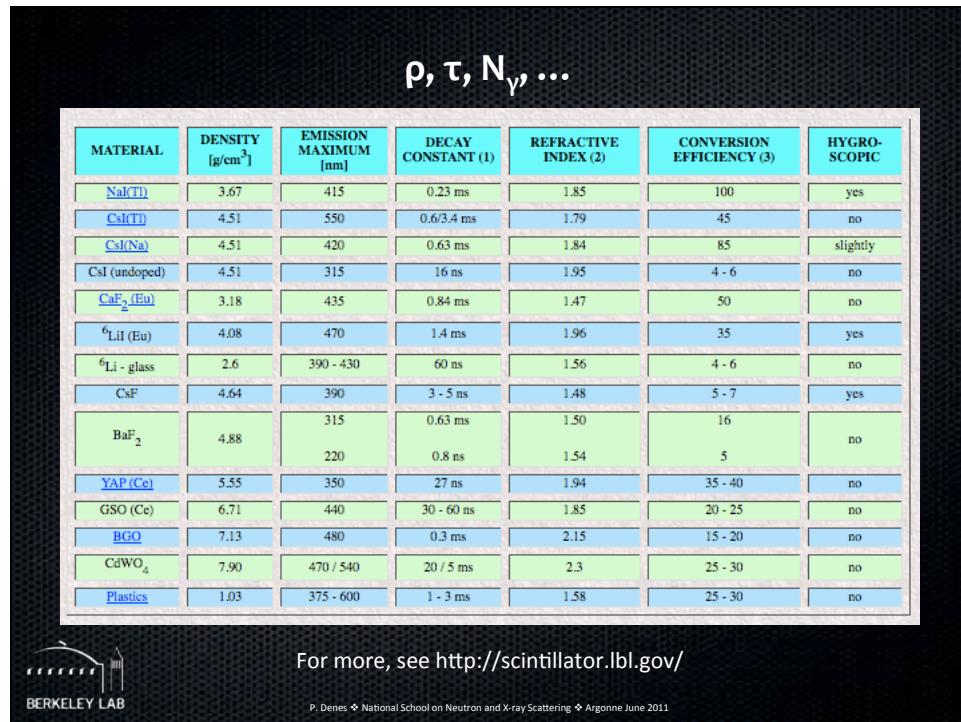


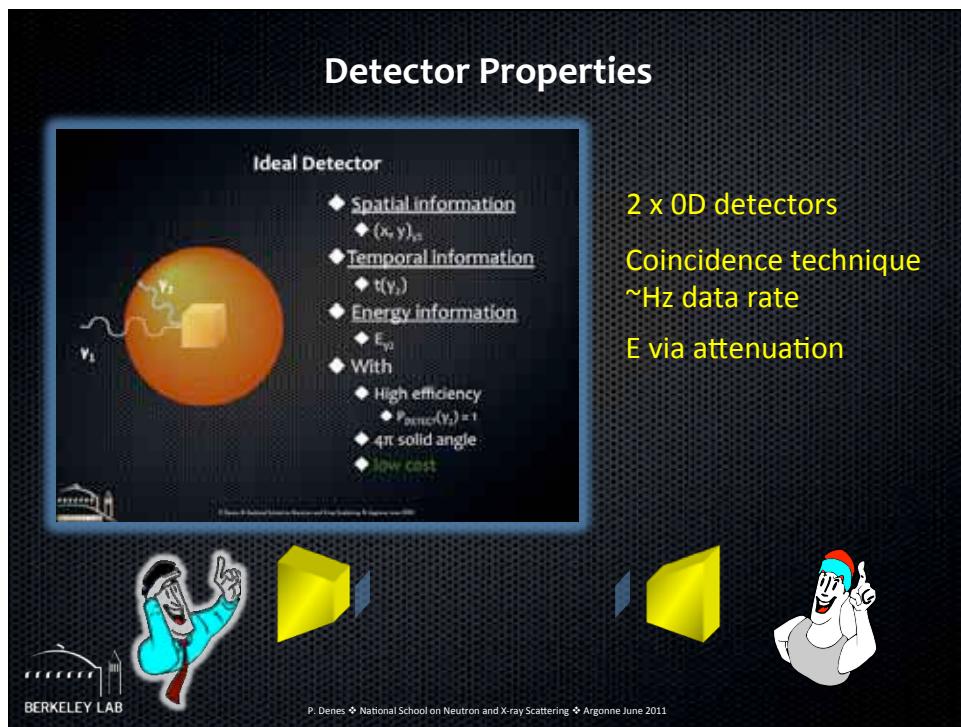
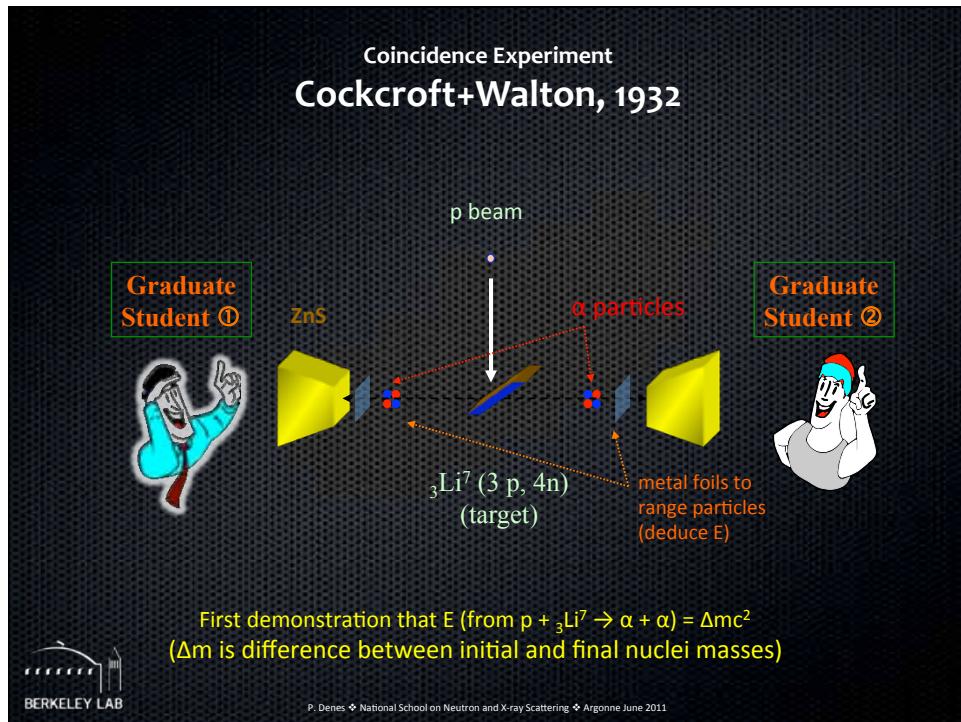
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$\oplus$  and  $\ominus$  of this technique

- Low Power (graduate students don't need much food)
- Low Speed - counting rate limitations  $\sim 1$  Hz
- Threshold sensitivity

(although Marsden could distinguish  $\alpha$  and  $p$  by brightness)

At  $\lambda \sim 500$  nm, Threshold<sub>TRAINED OBSERVERS</sub>  $\sim 17 \gamma$  for  $t_{FLASH} > 40 \mu\text{s}$

- Yield: "...at one famous laboratory during this period all intending research students were tested in the dark room for their ability to count scintillations accurately. Only those whose eyesight measured up to the standards required were accepted for nuclear research; the others were advised to take up alternative, less exacting, fields of study"

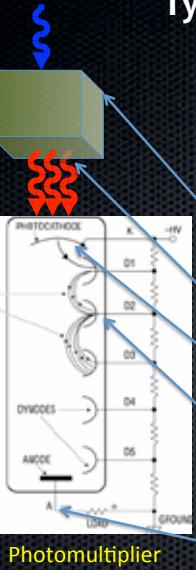
(from Birks)



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### Typical Scintillation Detector



**Photomultiplier**

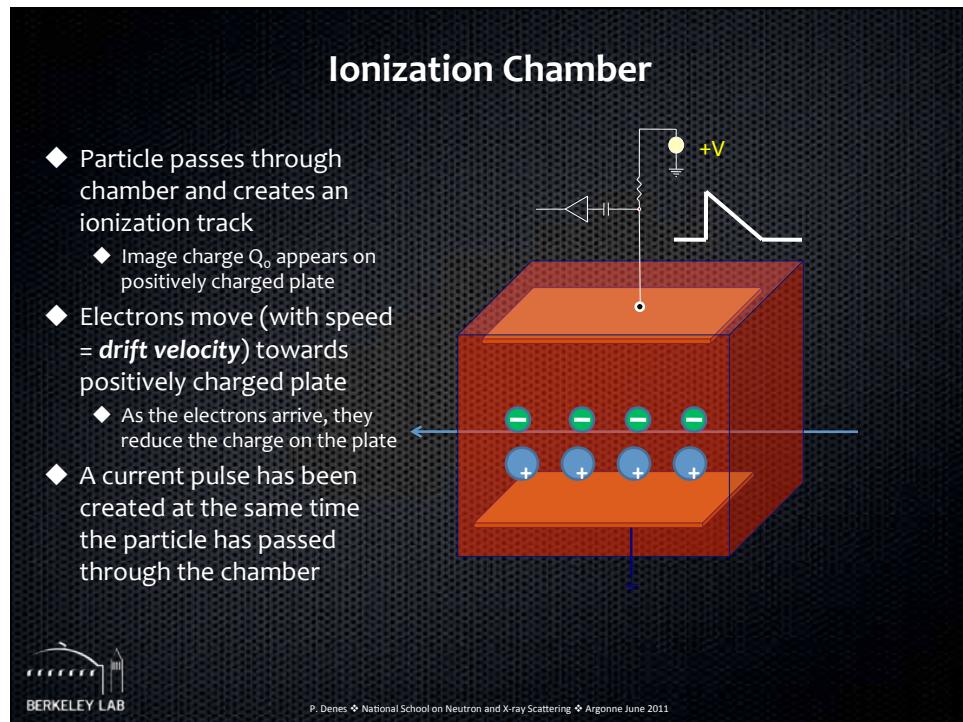
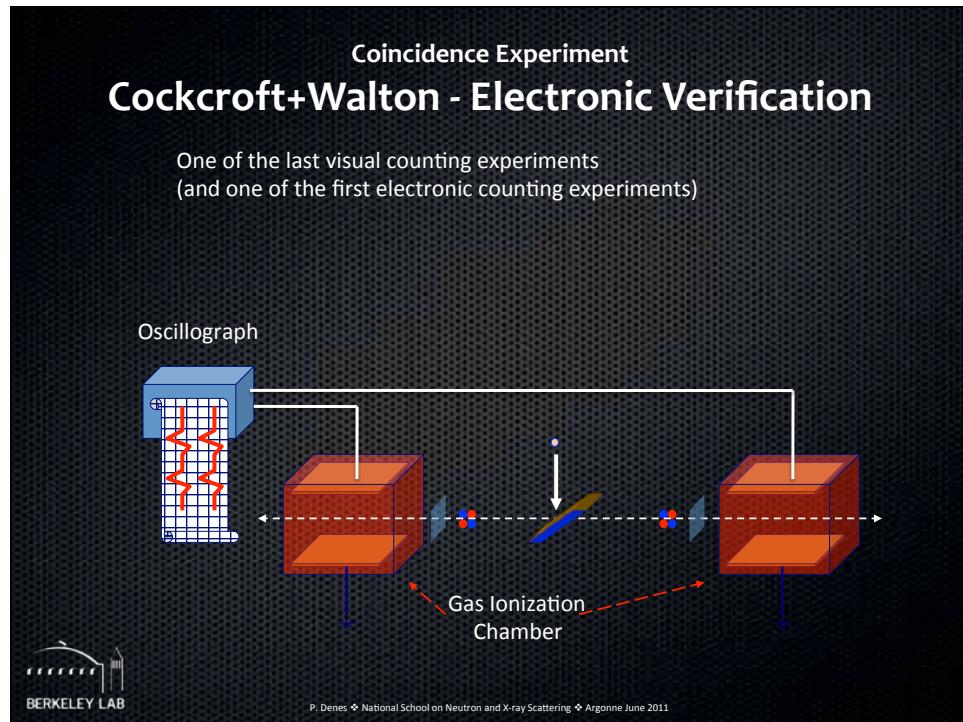
The diagram illustrates a typical scintillation detector. An incident photon (represented by a blue wavy arrow) strikes a green rectangular scintillator crystal. The crystal emits red wavy arrows representing visible photons. These photons enter a yellow cylindrical photomultiplier tube. Inside the tube, the photons are converted into electrons (indicated by small black dots). These electrons are then amplified by secondary emission, represented by a series of curved lines. Finally, an output current is detected at the anode.

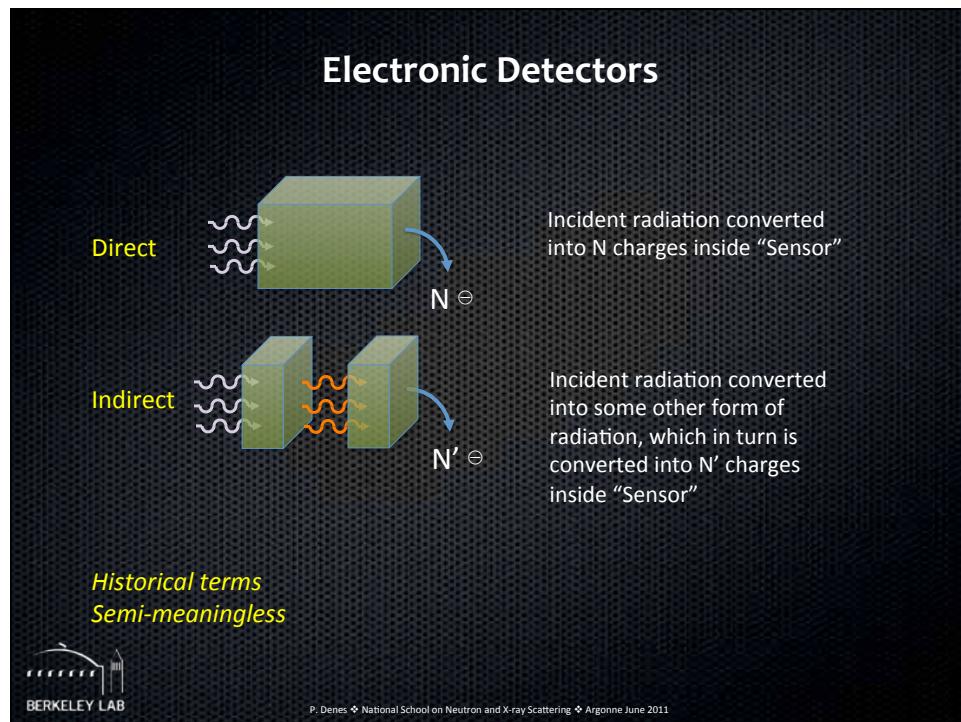
- ◆ Incident photon creates (ionization) electron
- ◆ Ionization generates visible photons
- ◆ Visible photons converted into electrons
- ◆ Electrons “amplified” by secondary emission
- ◆ Output current detected



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## Energy Needed for Detection

"Sensor"	$\eta = E$ per secondary quanta	Mechanism
Gas	30 eV	$e^-/ion$ pairs
Scintillator	10 – 1000 eV	optical excitation
Semiconductor	1 – 5 eV	$e^-/hole$ pairs
Superconductor	~meV	breakup of Cooper pairs
Superconducting calorimeters	~meV	phonons

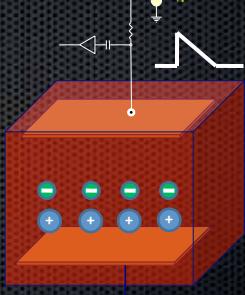
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Stolen from H. Spieler

## Statistics – Fano Factor

"Sensor"	$\eta = E$ per secondary quanta	Mechanism
Gas	30 eV	$e^-/ion$ pairs
Scintillator	10 – 1000 eV	optical excitation
Semiconductor	1 – 5 eV	$e^-/hole$ pairs
Superconductor	$\sim meV$	breakup of Cooper pairs
Superconducting calorimeters	$\sim meV$	phonons

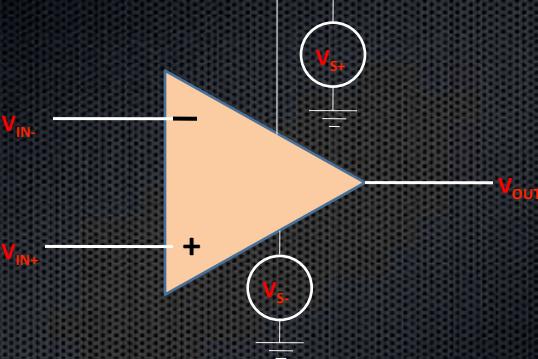


$N_{\pm} = \frac{E}{\eta}$   
 $\sigma_N = \sqrt{FN}$

- ◆ Intrinsic resolution is *Fano-limited*
- ◆  $\sigma_N/N \downarrow$  as  $\eta \downarrow$ 
  - ◆ Hence interest in superconducting calorimeters
- ◆ There are additional ways to have fluctuations on N

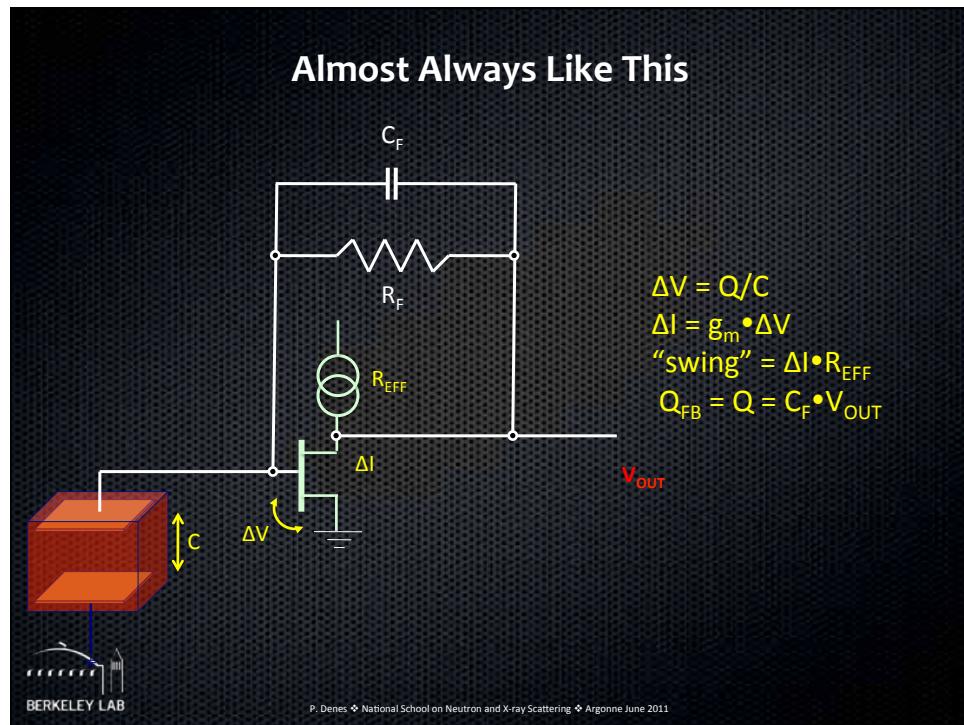
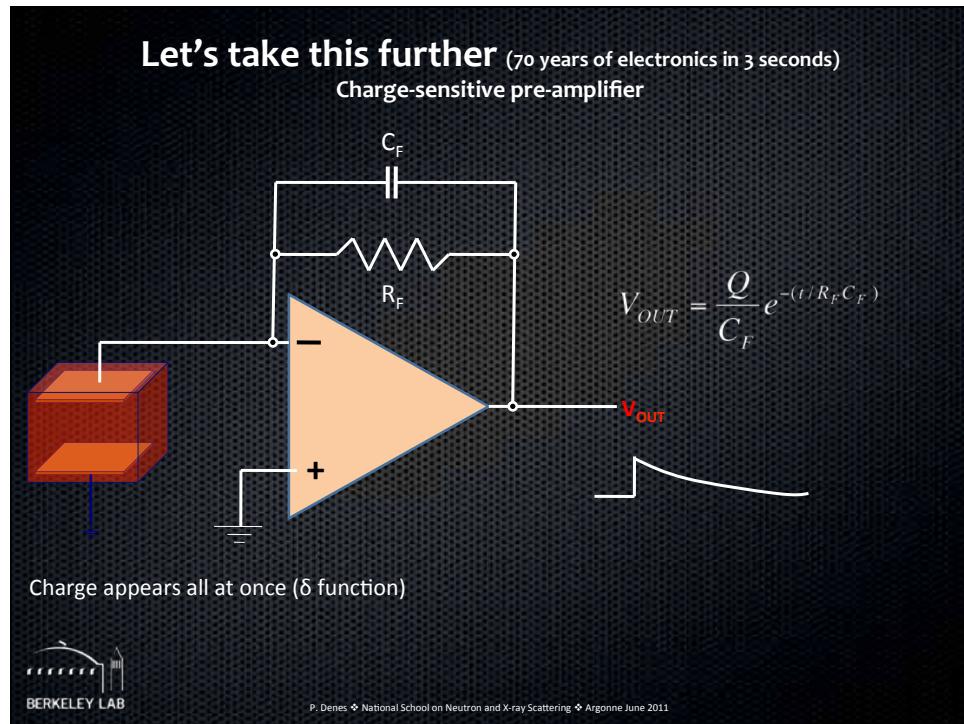
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### Next Problem – the current pulse is usually very small It must be amplified



$$V_{OUT} = \begin{cases} V_{S+} & \text{if } V_{IN+} > V_{IN-} \\ V_{S-} & \text{if } V_{IN+} < V_{IN-} \end{cases}$$

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## Noise and Statistics

Some terms to get started

- ◆ Incident photon creates electron of energy  $E_\gamma$  (photoelectric) or  $< E_\gamma$  (Compton) (with probability “QE”)
- ◆ Electron creates **on average**  $N = E_e/\eta$  e/h pairs
- ◆ Output pulse height = Gain x N Volts
- ◆ Output electronic noise  $V_N$  Volts

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## Semiconductor Detector

p-i-n diode

Reverse-biased diode  
Similar to gas ionization chamber

$$N = E_\gamma/\epsilon$$

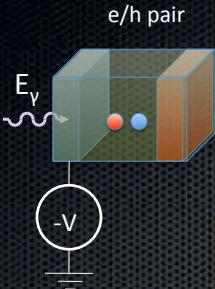
$$\sigma^2_N = F \cdot E_\gamma/\epsilon, F = \text{Fano factor}$$

Material	Si	Ge	GaAs	Diamond
$\eta$ [eV]	3.6	3.0	4.4	13.1
F	0.12	0.13	0.10	0.08
$\rho$ [g/cm³]	2.3	5.3	5.3	3.5

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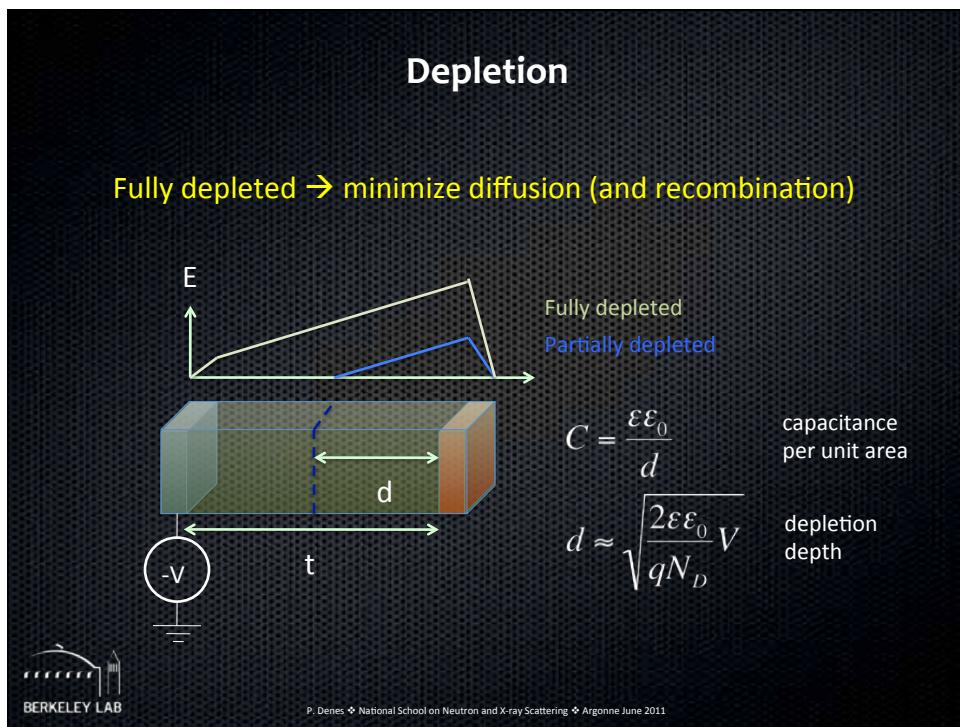
## How it Works

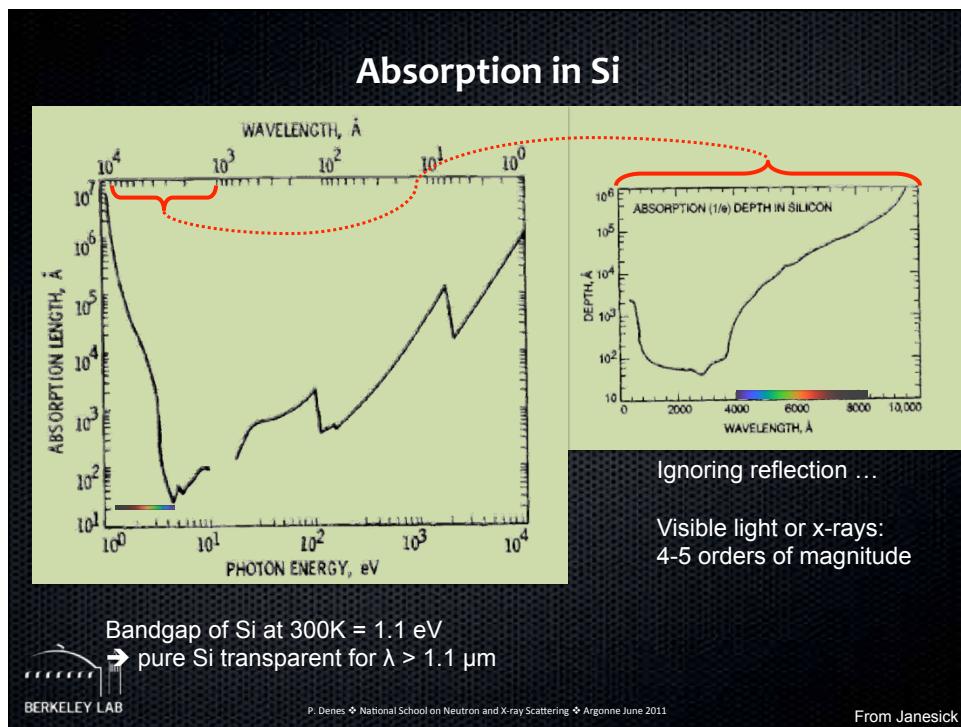
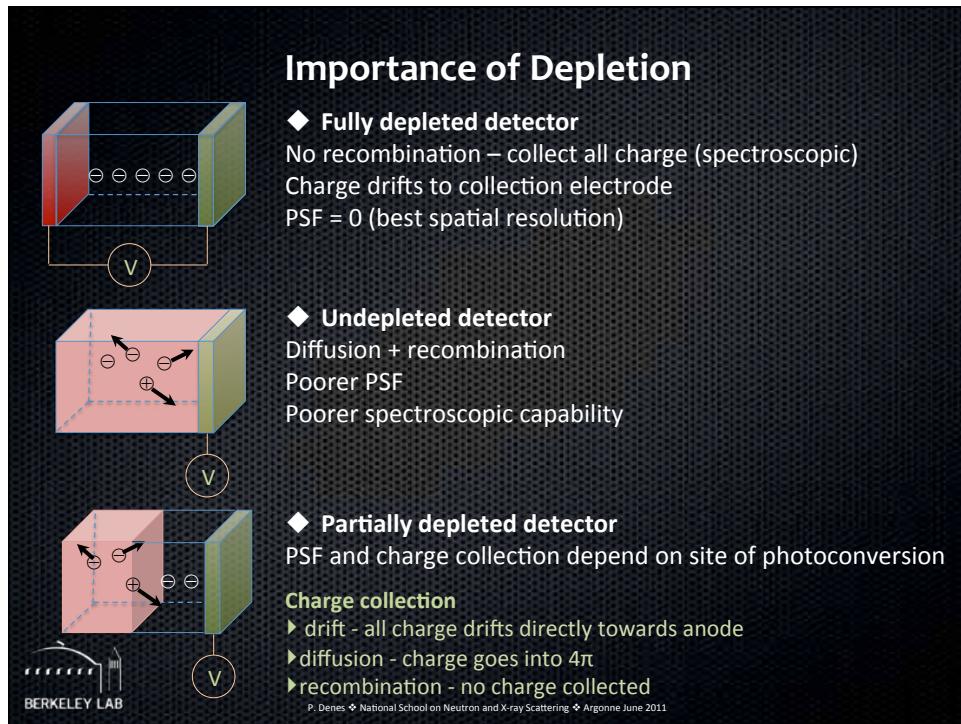


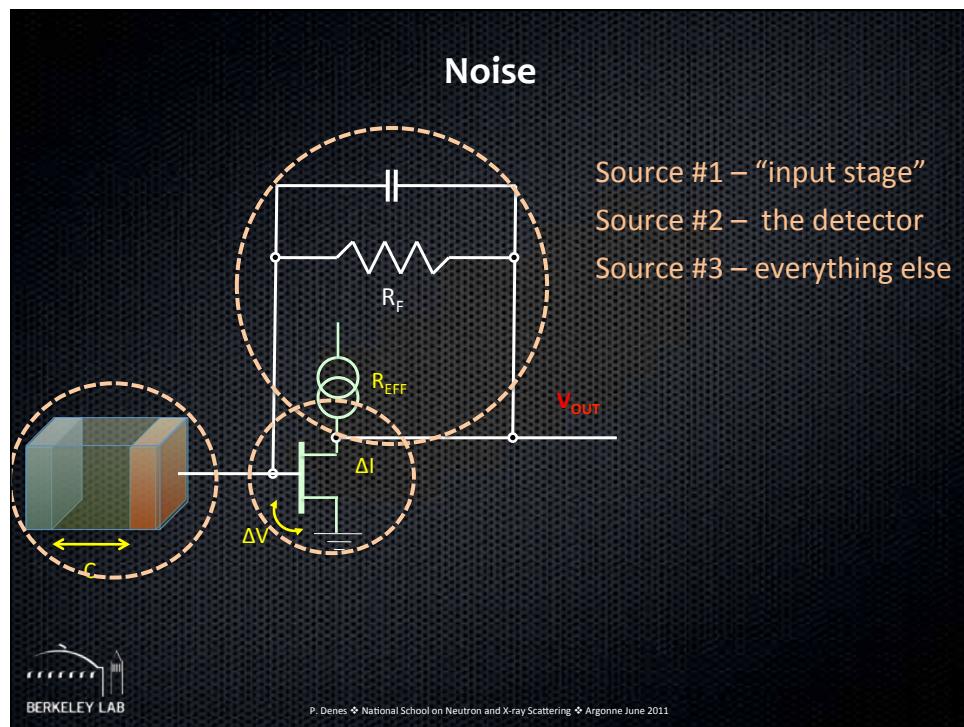
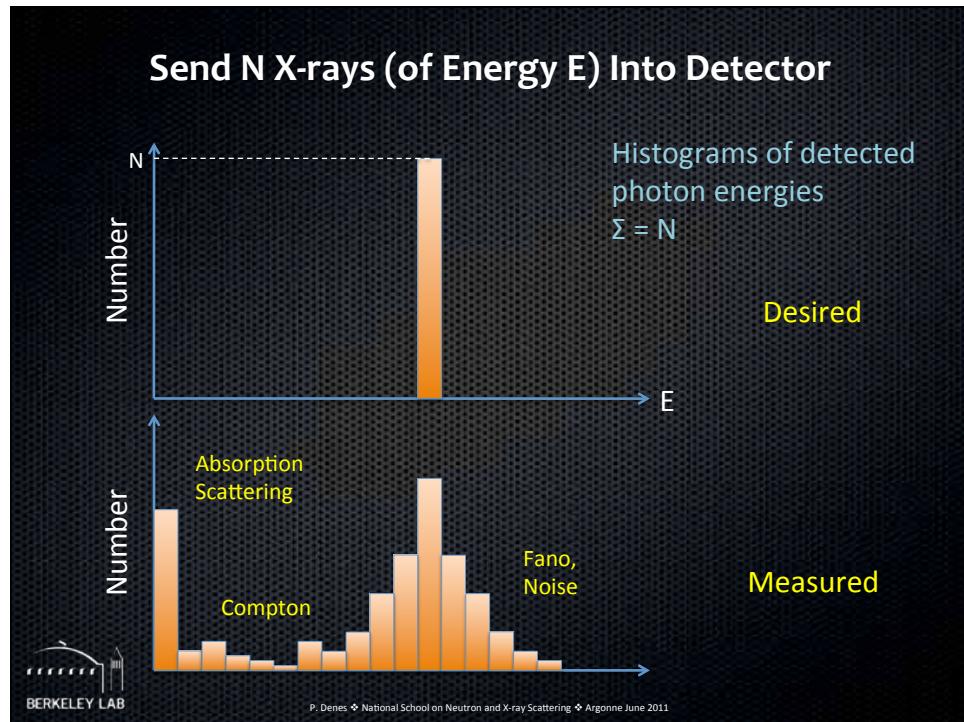
- ◆ Recombination
  - ◆  $e^-$  recombination time  $\propto 1 / \text{hole concentration}$
- ◆ Diffusion
  - ◆ In field-free region,  $e^-$  diffuses (into  $4\pi$ )  
 $D = (kT/q)\mu$  ( $\mu$  = mobility)
- ◆ Drift
  - ◆ In non-zero field region  $e^-$  moves towards positive plate with velocity  $\mu \cdot E$

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## The Detector Makes Noise?



**Semiconductor** detector  
i.e. valence band  $\sim$ full,  
conduction band  $\sim$ empty

eV band gaps  $\rightarrow$  thermal  
excitation of carriers

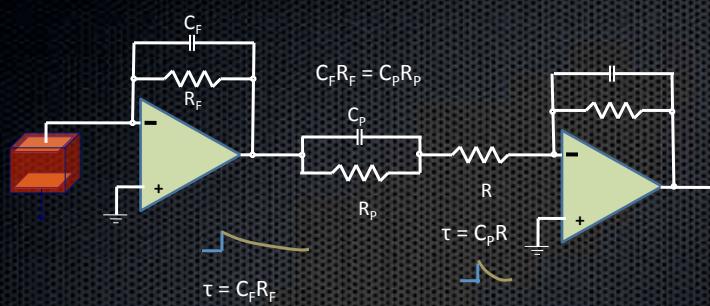
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- ◆ Thermal excitation
- ◆ “leakage” or “dark” current ( $I_{LEAK}$  e $^-$ /s)
- ◆ “looks like” signal
- ◆ (“shot noise”)
- ◆ Reduced by cooling
- ◆ Noise,  $\propto \sqrt{I_{LEAK}}$ , because leakage is not orderly



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## Some More Electronics



$\tau = C_F R_F$

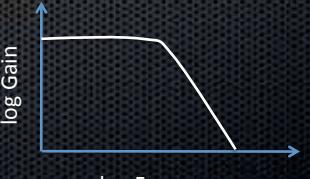
$C_F R_F = C_P R_P$

$\tau = C_P R$

“shaping time”  $\tau$

Bandwidth =  $1 / 2\pi \tau$

In frequency domain:



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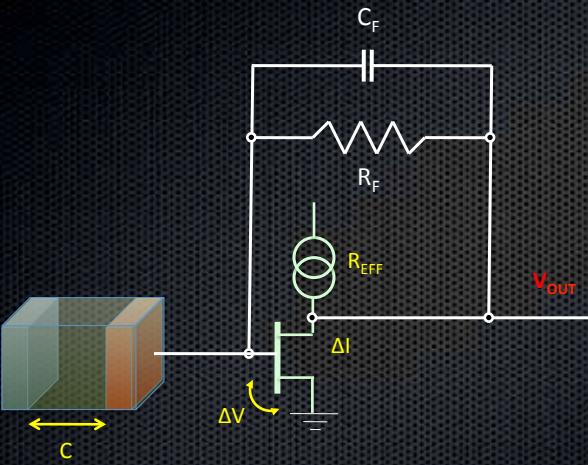
**Things  $\propto \tau$**

- ◆ Double pulse resolution  $\propto \tau$
- ◆ Noise due to leakage current  $\propto$ 
  - ◆  $\sqrt{I}$  – random arrival of leakage charge
  - ◆  $I \sim e^{-T/T^2}$
  - ◆  $\sqrt{\tau}$  – i.e.  $\sqrt{[e^-/s] \cdot [s]}$
- ◆ Longer integration time ( $\tau$ ) increases noise due to leakage current
- ◆ Must want short integration time

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### Electronic Noise



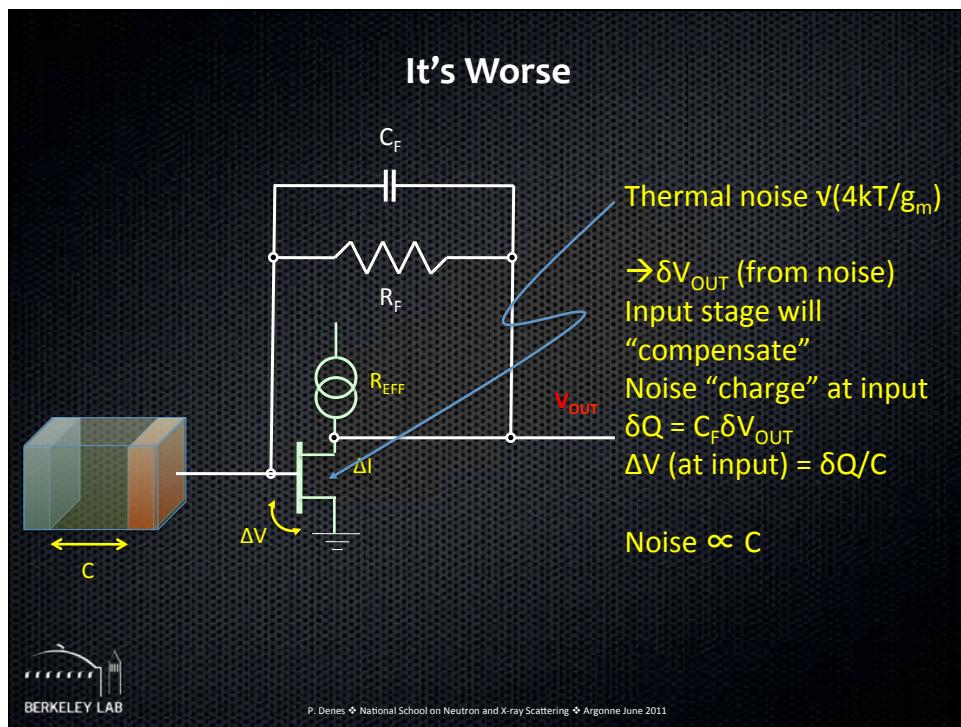
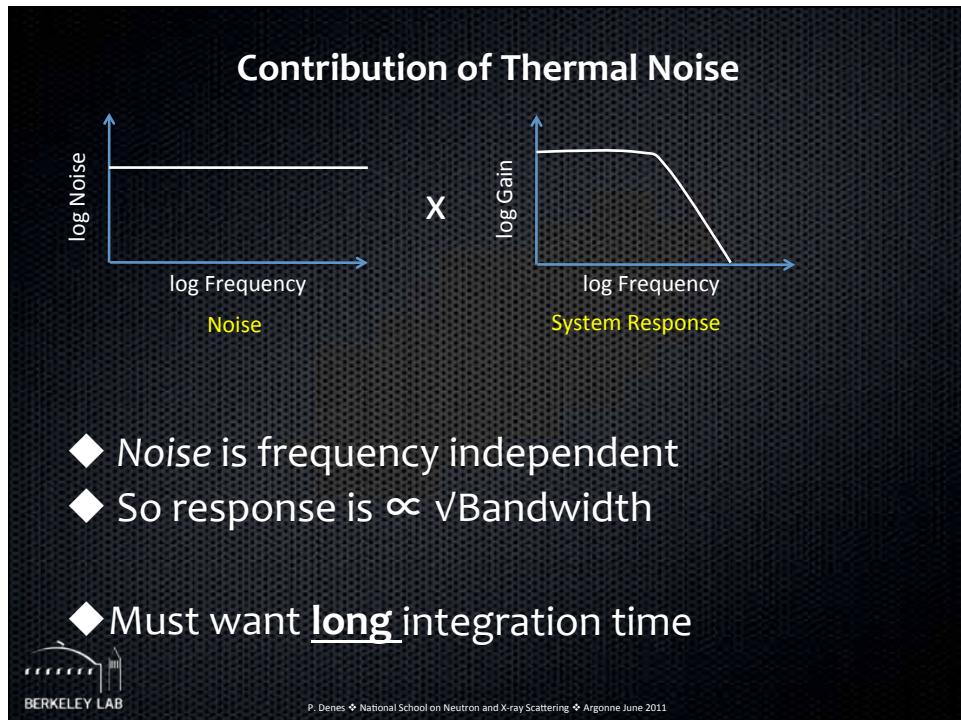
Resistors make noise  
(Thermal excitation of carriers in resistor means  $I \times R = V_{NOISE}$ )

Thermal noise is truly random,  $V_N \sim \sqrt{4kTR}$



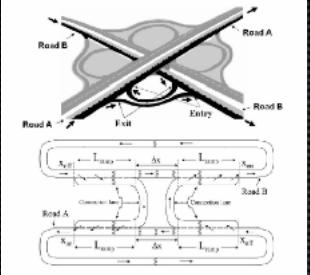
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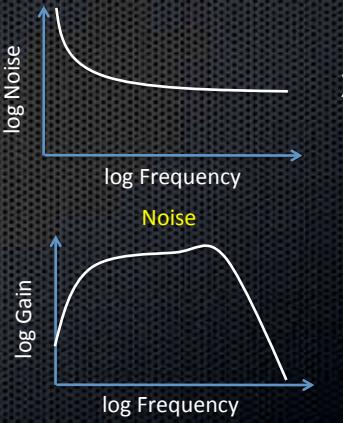
**It's Even Worse**

Many physical systems are subject to fluctuations  $\sim 1/f^\alpha$   
You know this from driving:



RMS of time you wait getting onto the freeway  $\sim 1/f$

Same with electronics. So there is an optimum



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**Not so Simple**

1. Fluctuations in number of photons “absorbed”
2. Fluctuations in number of secondary particles created
3. (Fluctuations in number of tertiary particles created)
4. Electronic noise
  - ◆ Energy resolution: 2, 3 and 4
  - ◆ Quantum efficiency: 1 (but maybe 2, 3 and 4)

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## Detective Quantum Efficiency

- ◆ Combine notion of Quantum Efficiency (probability of detecting a particle) with spatial response (probability of detecting/quantifying N (x,y) particles → DQE
- ◆ How faithfully does the detector transfer the (spatially varying) fluctuations of the input signal
- ◆  $DQE(\omega_x, \omega_y)$
- ◆ Many definitions – most common is  $DQE = \frac{(S/N)_{OUT}^2}{(S/N)_{IN}^2}$
- ◆ Example, flat field illumination (flux  $\phi$ ) of detector with certain QE
- ◆  $(S/N)_{IN} = \frac{\phi A \tau}{\sqrt{\phi A \tau}}$  (Poisson)
- ◆  $(S/N)_{OUT} = \frac{QE \times \phi A \tau}{\sqrt{QE \times \phi A \tau + \sigma_N^2}}$  for electronic noise  $\sigma_N$

  
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## S/N, Dynamic Range, Number of Bits

Usually mis-stated!

- ◆ Si:  $\eta = 3.6$  eV. Inject 3.6 keV  $\gamma$ s (generates on average 1,000 e/h pairs) and measure the output pulse height → “conversion gain” = Volts / e<sup>-</sup> = V<sub>e</sub>
- ◆ RMS noise at output = V<sub>N</sub>
  - ◆ ENC (Equivalent Noise Charge) = V<sub>N</sub>/V<sub>e</sub>
- ◆ If the maximum voltage that the system can measure is V<sub>MAX</sub>, then the dynamic range is V<sub>MAX</sub> / V<sub>N</sub>
  - ◆ Example: V<sub>e</sub> = 1 μV / e<sup>-</sup>, V<sub>N</sub> = 100 μV
    - ◆ ENC = 100 e<sup>-</sup> = 360 eV [RMS]
    - ◆ V<sub>MAX</sub> = 1V → DR = 1V / 100 μV = 10<sup>4</sup>
  - ◆  $N_{BITS} = \ln(DR) / \ln(2)$ 
    - ◆  $\ln(10^4) / \ln(2) = 13$  bits (i.e.  $2^{13} \approx 10^4$ )
- ◆ S/N has specific meanings, that are not any of these!

  
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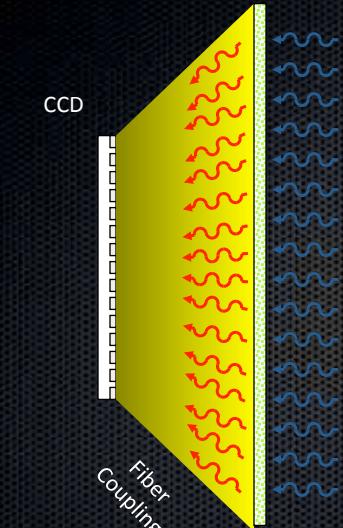
**A tale of 3 different  $\sqrt{N}$**

- ◆ Uniform flux  $\phi$  [ $\gamma/\text{cm}^2/\text{s}$ ] on area A yields  
 $N = \phi A [\gamma/\text{s}] \pm \sqrt{N}$  incident photons/s
  - ◆ photostatistics
- ◆ Each one (that is converted) produces  $N_{\pm} = NE/\eta \pm \sqrt{(FN_{\pm})}$  e/h pairs/s
  - ◆ intrinsic resolution
- ◆ Which, as a current sampled in time  $\tau$  has fluctuations  $\sim \sqrt{(N_{\pm}\tau)}$ 
  - ◆ shot noise



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**“Classical” X-ray Detector**



- ◆ Phosphor (powdered scintillator)
- ◆ Fiber-optically coupled to a CCD (2D solid-state detector) camera
- ◆ + and -

  - ◆ “general purpose”
  - ◆ radiation damage
  - ◆ area
  - ◆ phosphor
  - ◆ fiber-optic



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## Scientific CCDs (Charge-Coupled Devices)

Dumbbell nebula - LBNL CCD  
Blue: H<sub>α</sub> at 656 nm  
Green: SII at 955 nm  
Red: R II at 1.02 mm

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- ◆ CCD invented in 1969 by Boyle and Smith (Bell Labs) as alternative to magnetic bubble memory storage
- ◆ LST (“Large Space Telescope” – later Hubble) 1965 – how to image?
  - ◆ Film was obvious choice, but - It would “cloud” due to radiation damage in space Changing the film in the camera not so trivial
- ◆ 1972 CCD proposed

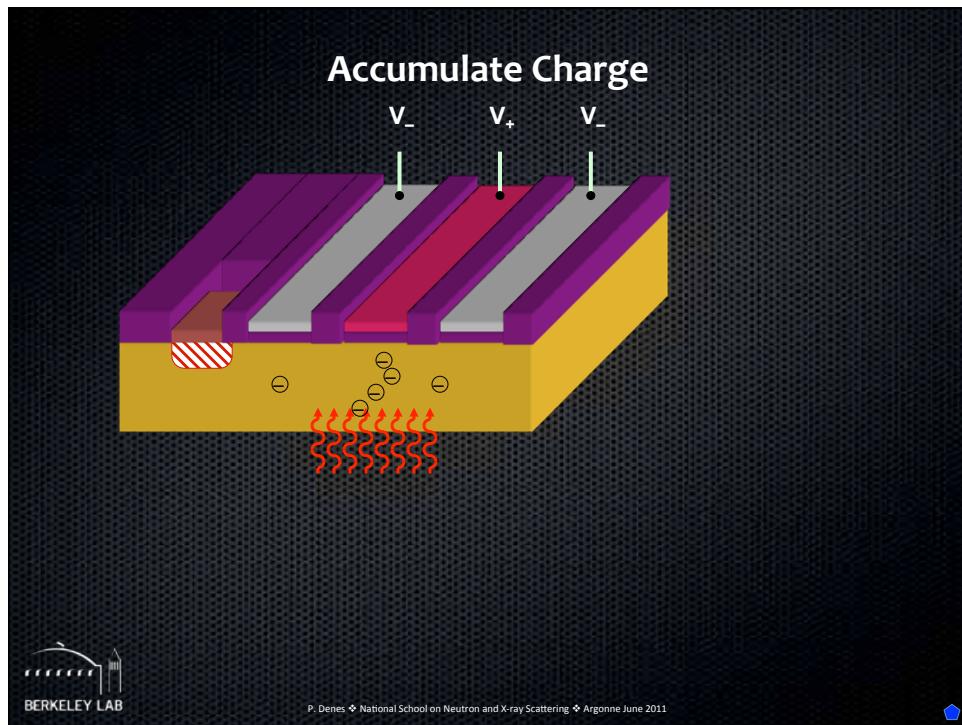
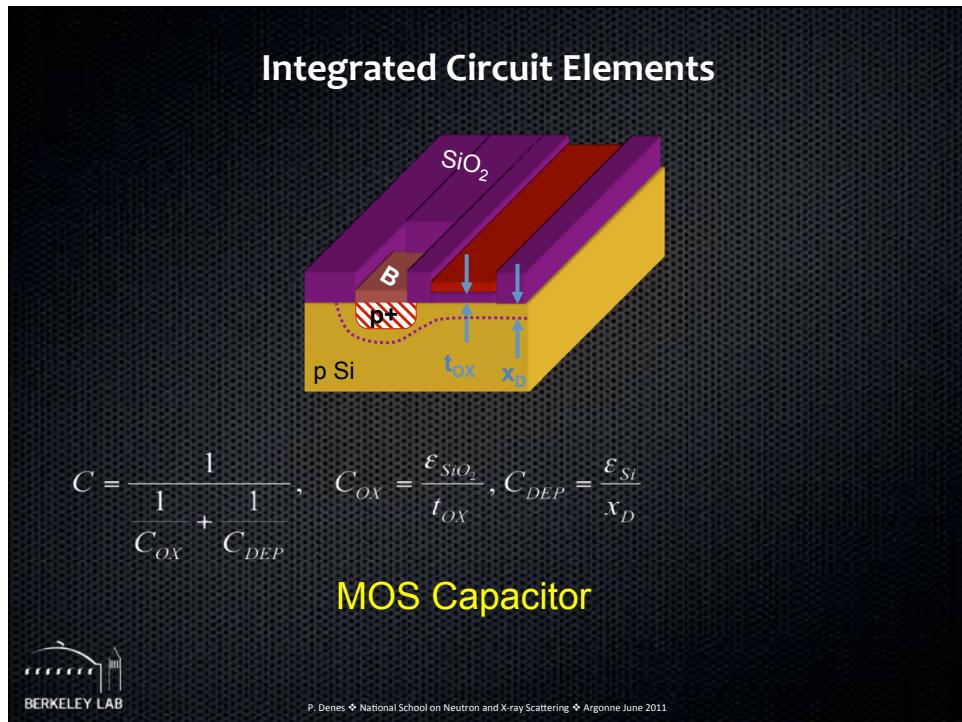
## Si Processing: Integrated Circuit Elements

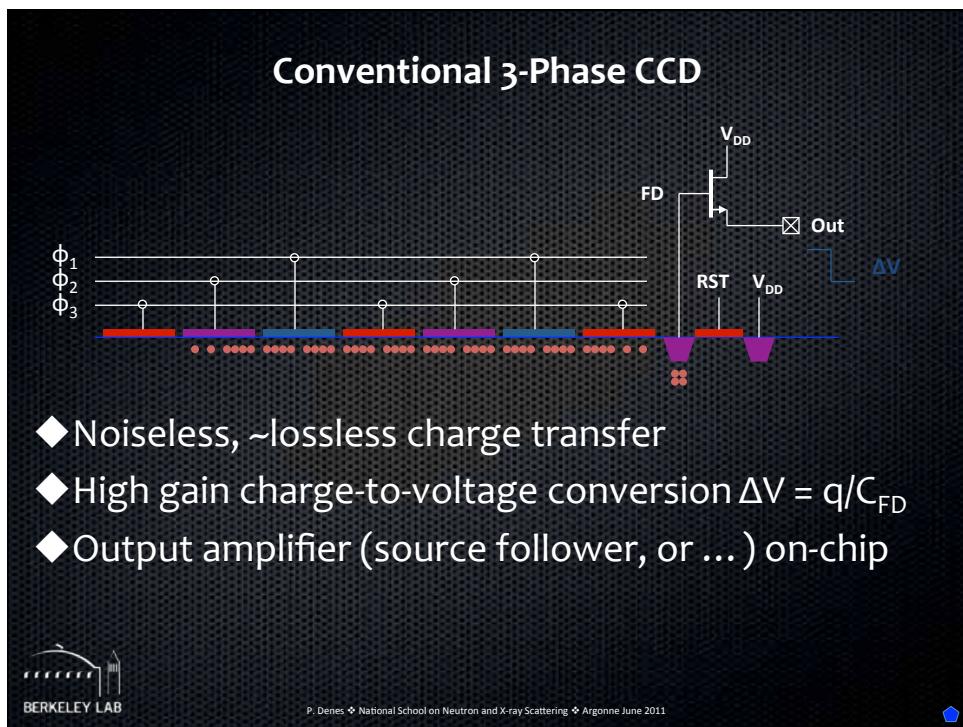
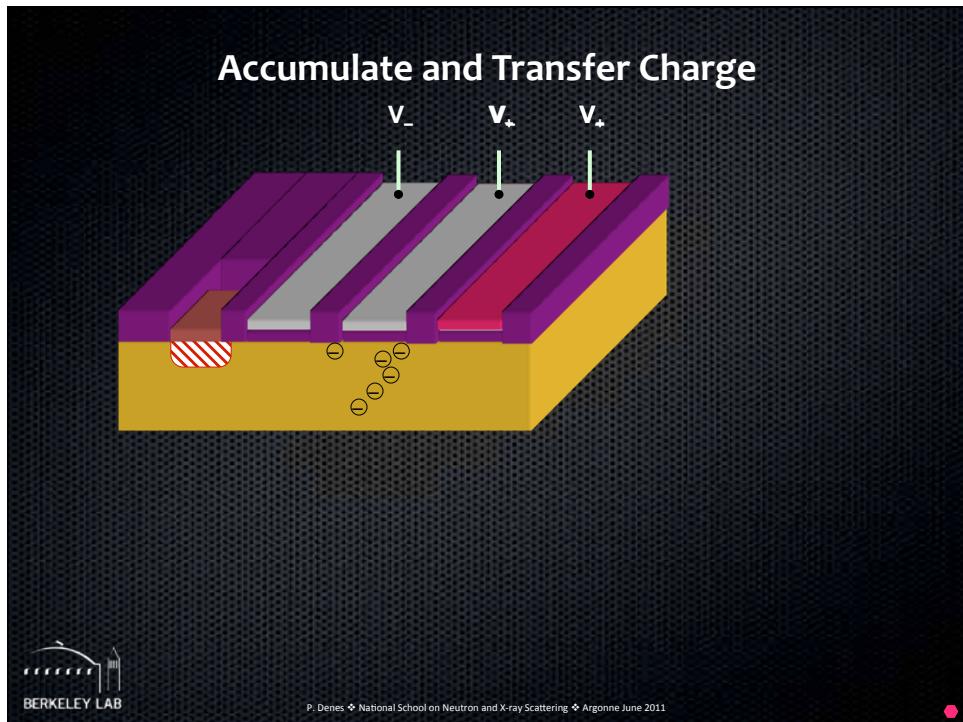
MOS Transistor

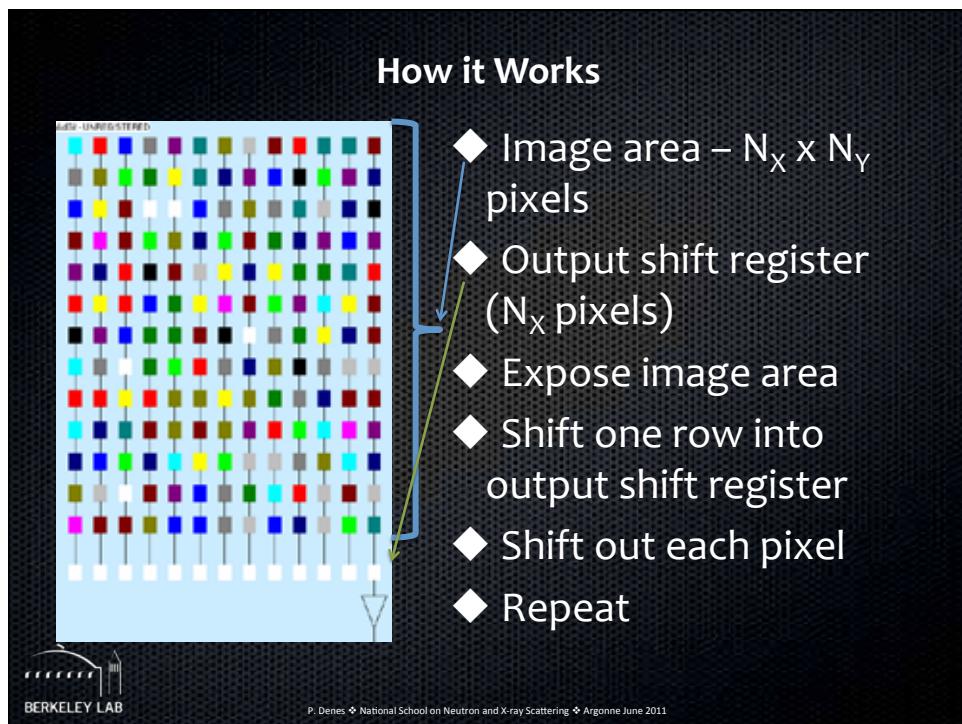
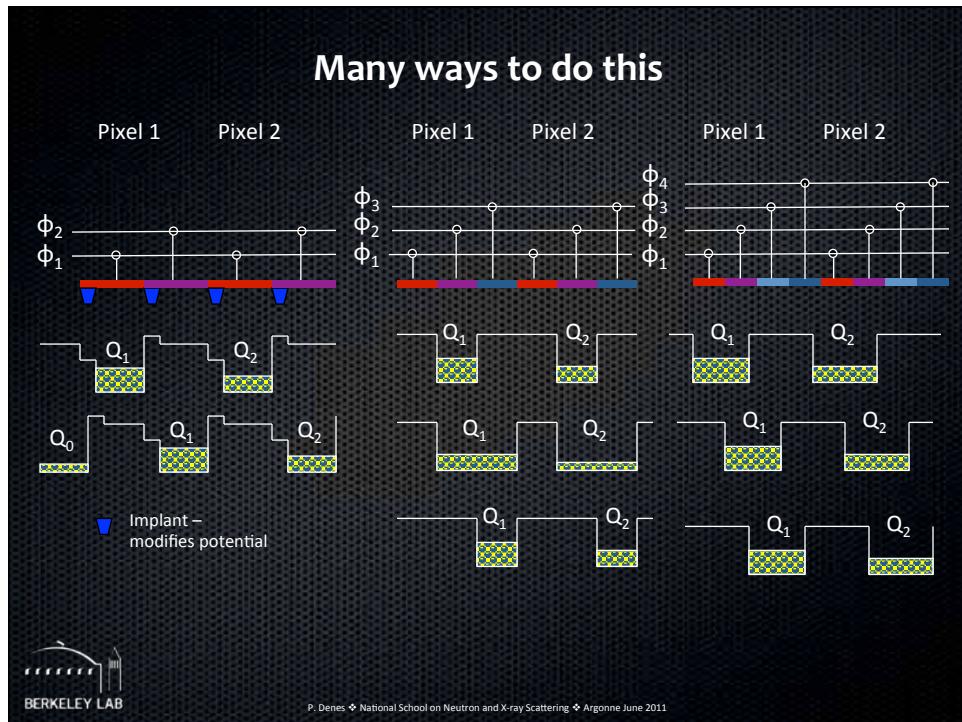
pn Diode

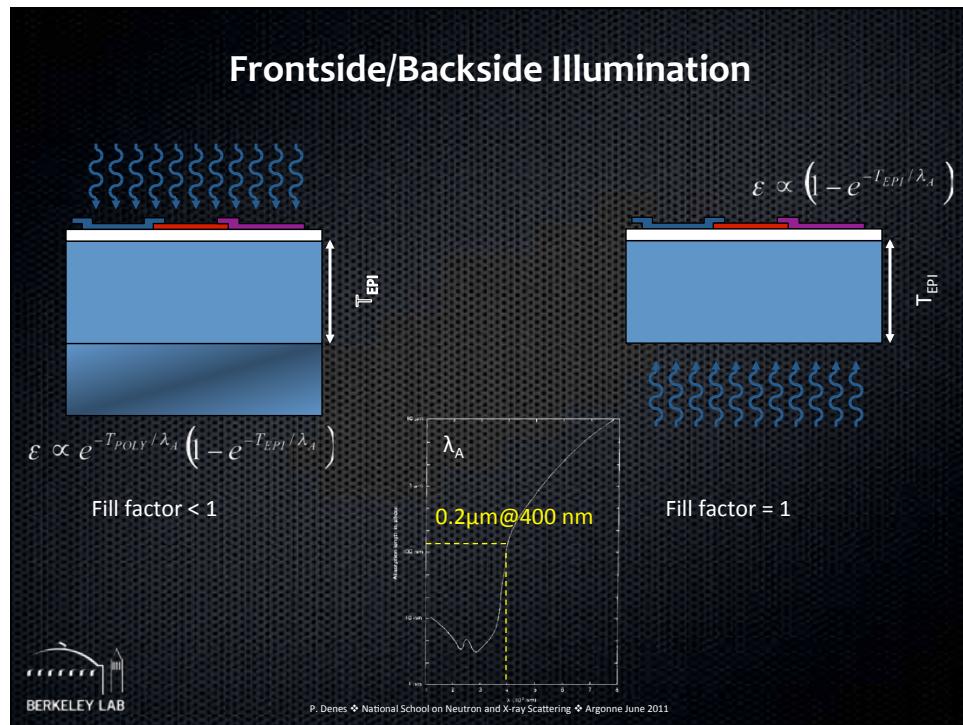
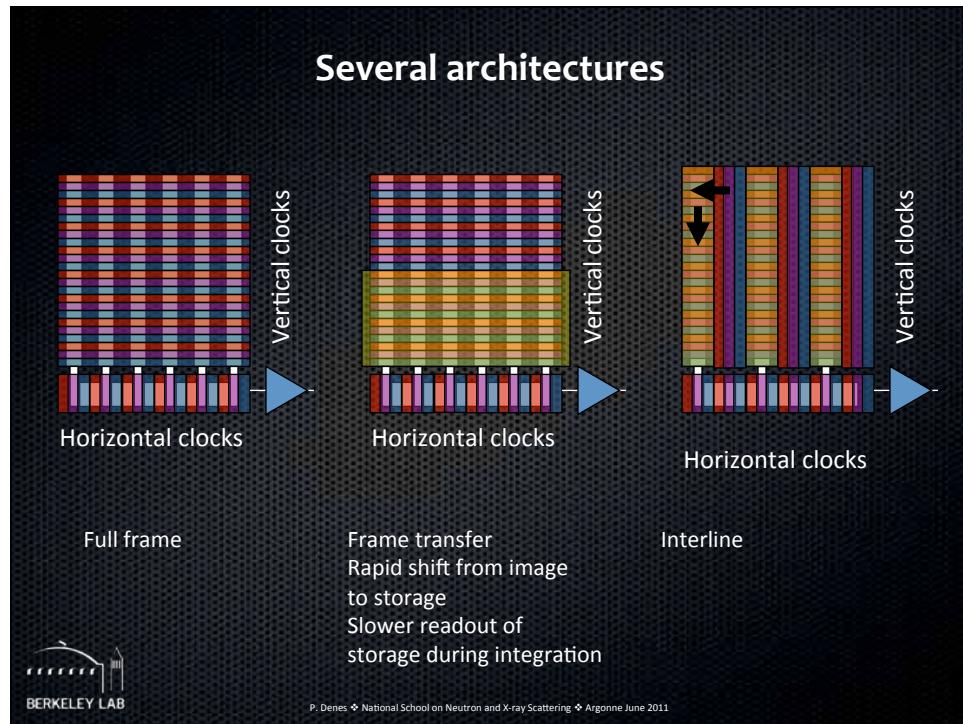
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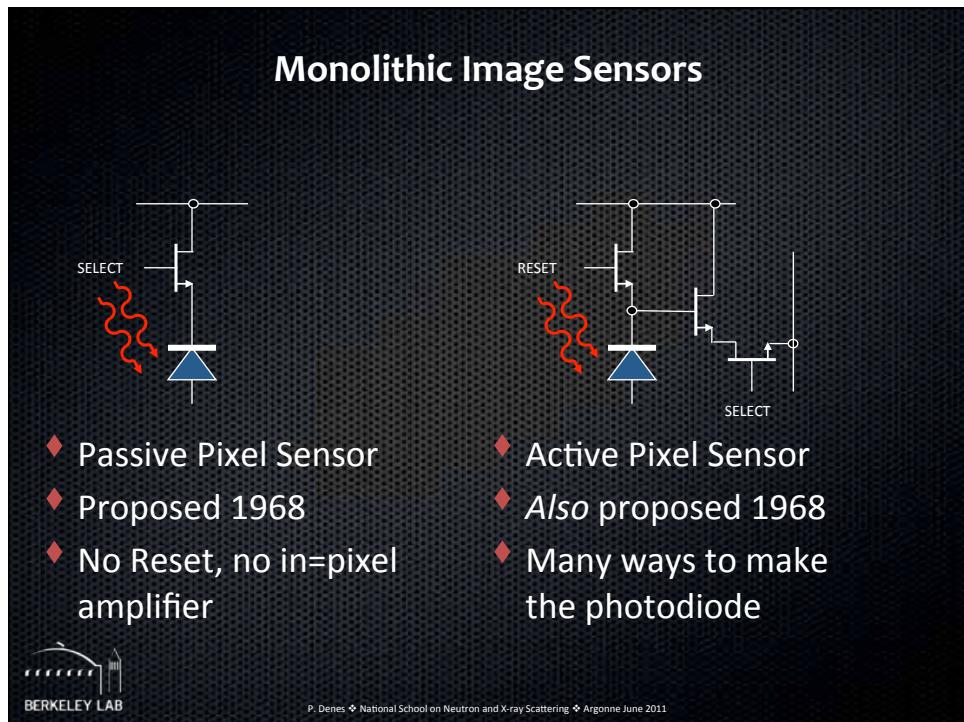
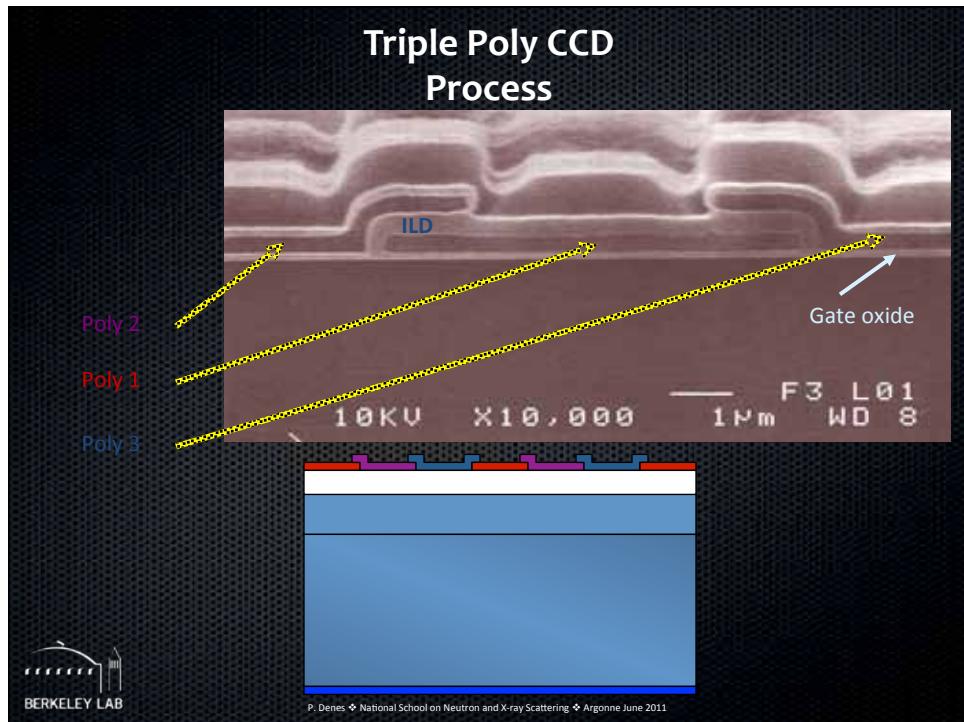
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## CCD vs APS

- ◆ APS – transfers a voltage down the column
- ◆ CCD – (noiselessly) transfers a charge down the column
- ◆ APS – can be more sensitive (source follower does not have to drive off-chip)
- ◆ APS – fill factor < 1 in general
- ◆ Photogate APS – like a matrix of individual CCDs
- ◆ Backside illumination – attempted for APS, work-in-progress

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## CMOS, CMOS “opto” and CCD processes

**CMOS driven by constant field scaling**

$t_{ox} \rightarrow t_{ox}/\kappa$

$V \rightarrow V/\kappa$

$n+S$

$n+D$

$p \text{ substrate}$

$Doping - N_A \rightarrow \kappa N_A$

$\rightarrow \leftarrow$

Channel Length  $L \rightarrow L/\kappa$

	CCD	CMOS
$t_{ox}$ ( $\text{\AA}$ )	500 - 1000	5-20
Well depth ( $\mu\text{m}$ )	2.5	0.5 deeper for RF
Implant ( $\mu\text{m}$ )	$\sim 1$ channel stop	0.1 S/D implants
$V$	$\geq 10$	<3.3 <2.5 <1.x ...
Poly layers	3 (2)	1 2 for analog
Subst. quality	Low leakage	Don't care Except opto

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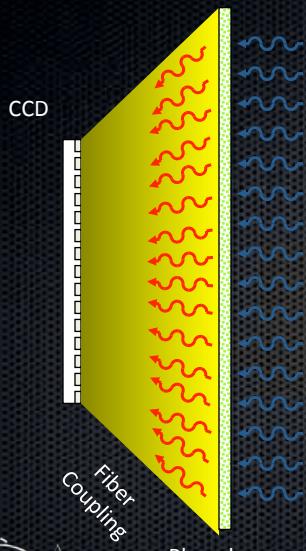
## Why CCDs?

- ◆ Low noise (noiseless charge transfer, do everything to make  $C_{FD}$  small in order to get large conversion gain)
- ◆ Fill-factor = 1 (for backside illumination)
- ◆ Linear and easy to calibrate
- ◆ **Long history of scientific use**
- ◆ Large area devices easier (cheaper) to develop as CCDs than as state of the art CMOS devices
  - ◆ Readily wafer scale
- ◆ Commercially produced



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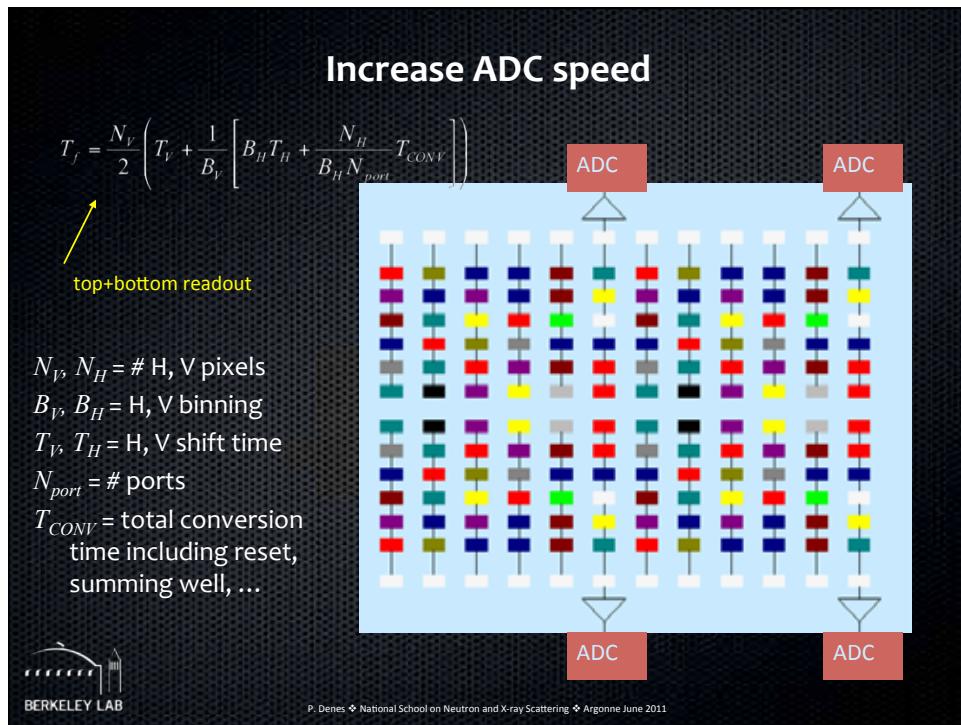
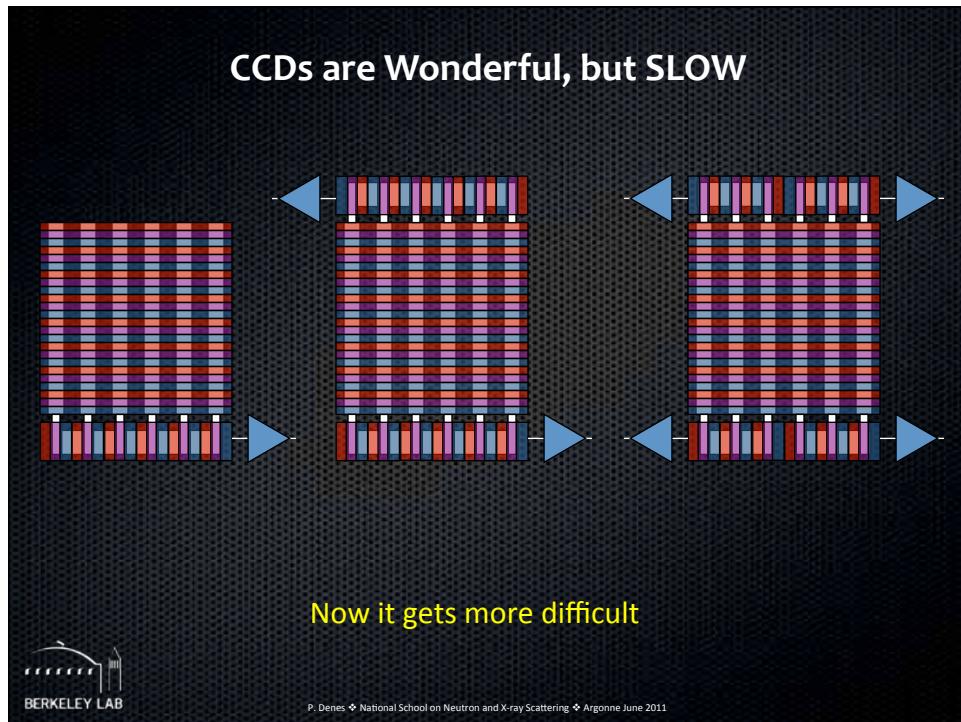
## “Classical” X-ray Detector



- ◆ Phosphor (powdered scintillator)
- ◆ Fiber-optically coupled to a CCD (2D solid-state detector) camera
- ◆ + and –
  - ◆ “general purpose”
  - ◆ radiation damage
  - ◆ area
  - ◆ phosphor
  - ◆ fiber-optic



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**For example**

**Increase readout/ADC speed**

- ◆ Dalsa – FT50M
- ◆  $1024 \times 1024 \times 5.6 \mu\text{m}$  pixel
- ◆ Frame transfer / 2 ports
- ◆  $100 \text{ fps} = 100 \text{ MPix/s}$
- ◆ 11.1 bits [67 dB] at 30/60 fps
- ◆ 10.1 bits [61 dB] at 50/100 fps

**S/F Limitations**

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**Limitations**

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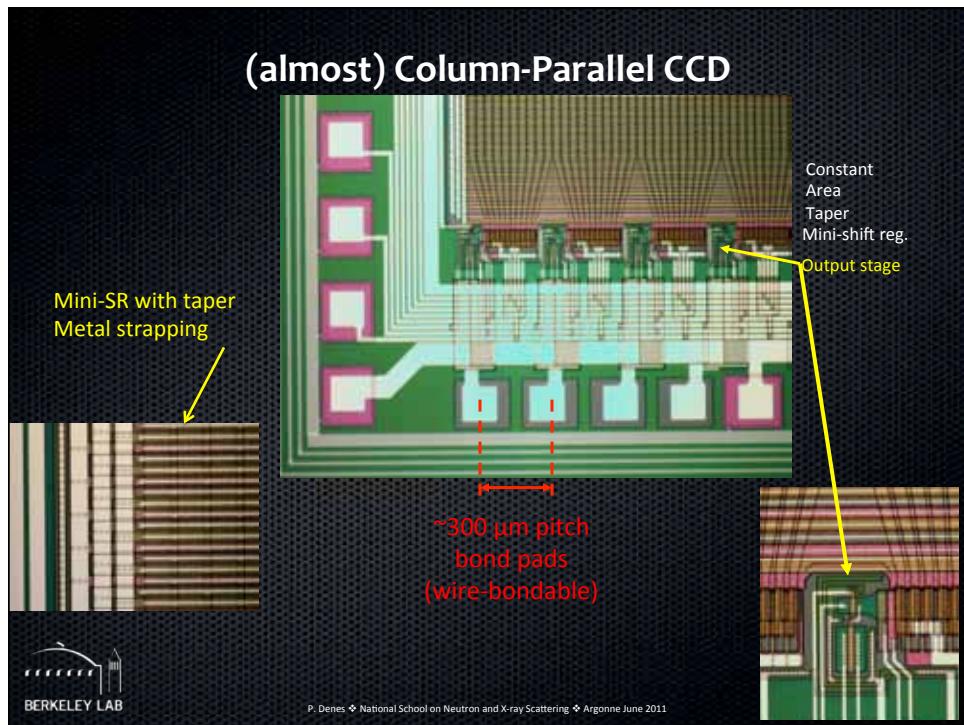
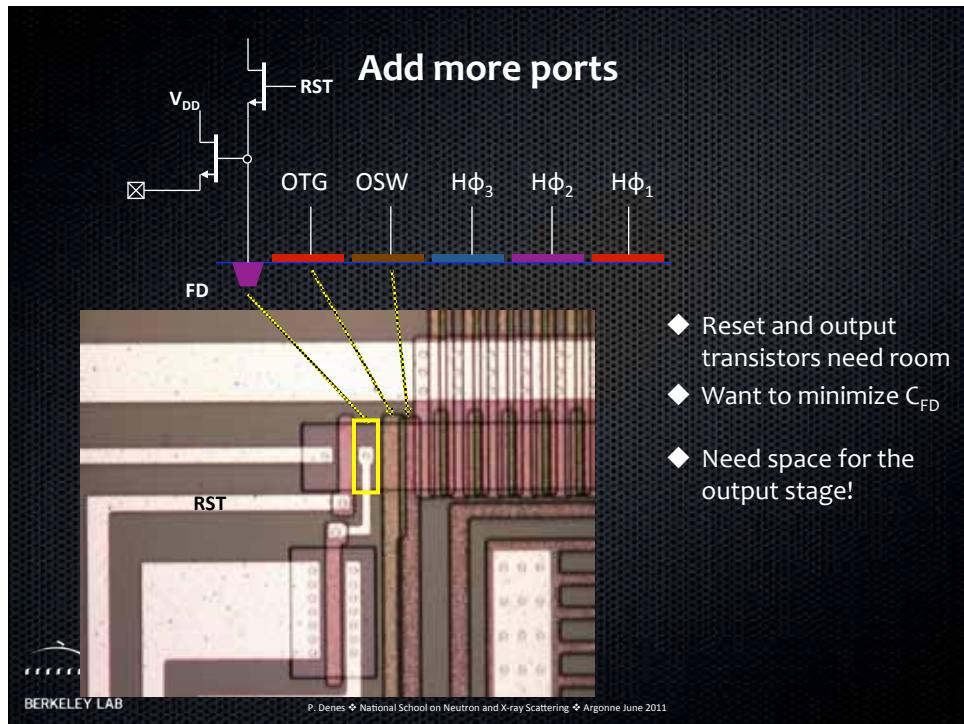
Noise contribution from  $M_R$  (reset switch) removed by CDS (correlated double sampling – measure  $V_R$  and  $V_R + V_S$ )

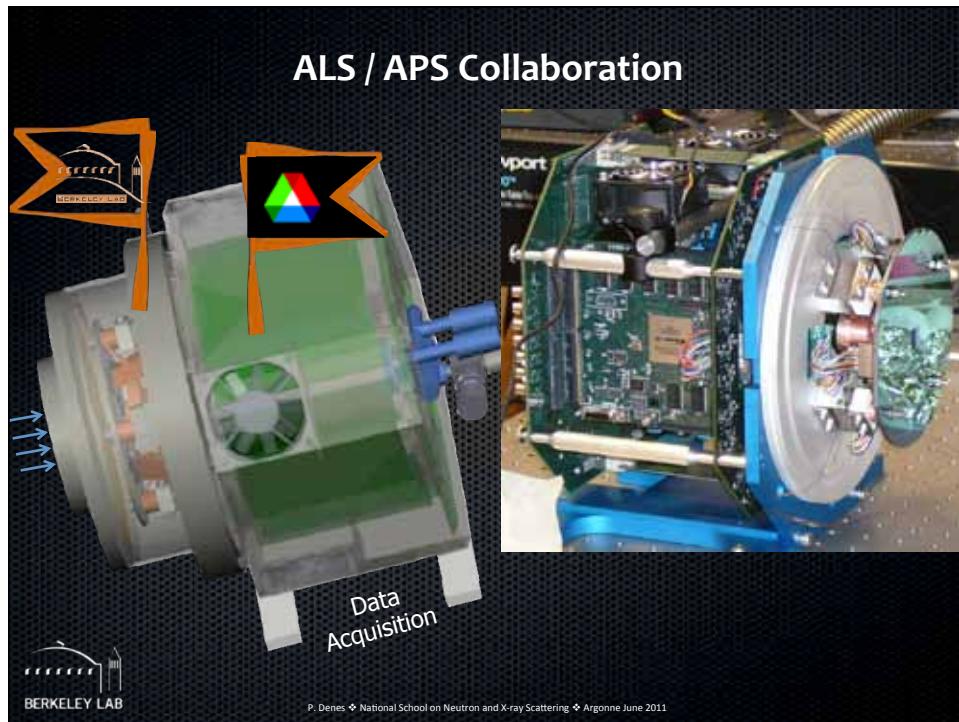
Noise contributions from  $M_S$  (source follower)

- ◆ Thermal noise  $V_n^2 \sim 4kTg_m \int H^2(f) df$
- ◆ 1/f noise  $V_n^2 \sim \frac{K}{C_{ox}WL} \int H^2(f) \frac{1}{f} df$
- ◆ Noise from current source

$\uparrow \sim \sqrt{\text{rate}}$

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## Direct Detection

Previous example of CCD usage was for optical photons. What about x-rays?

This should be depleted – generally thin with conventional processes  
→ add a layer which can be used as an electrode

**PROPOSAL:**  
*Make a thick CCD on a high-resistivity n-type substrate, operate fully depleted with rear illumination.*

**3-phase CCD structure**  
Poly gate electrodes  
buried p channel  
n<sup>-</sup> (10 kΩ·cm)  
photo-sensitive volume (300 μm)  
Transparent rear window

**Advantages:**

- 1) Conventional MOS processes with no thinning => "inexpensive"
- 2) Full quantum efficiency to > 1 μm => no fringing
- 3) Good blue response with suitably designed rear contact
- 4) No field-free regions for charge diffusion, good PSF

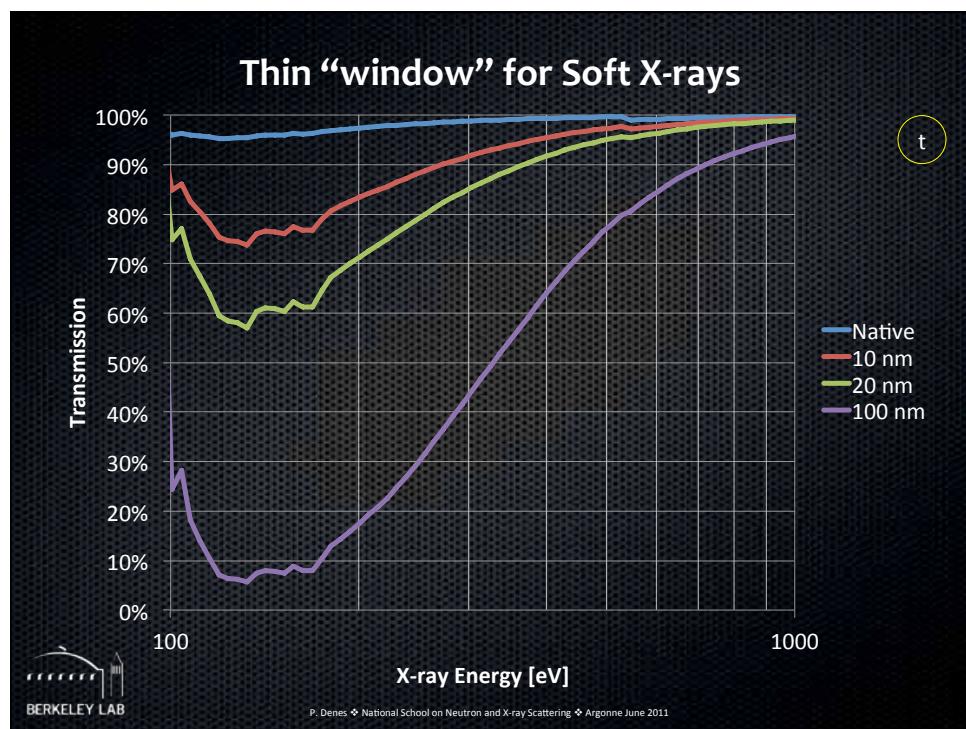
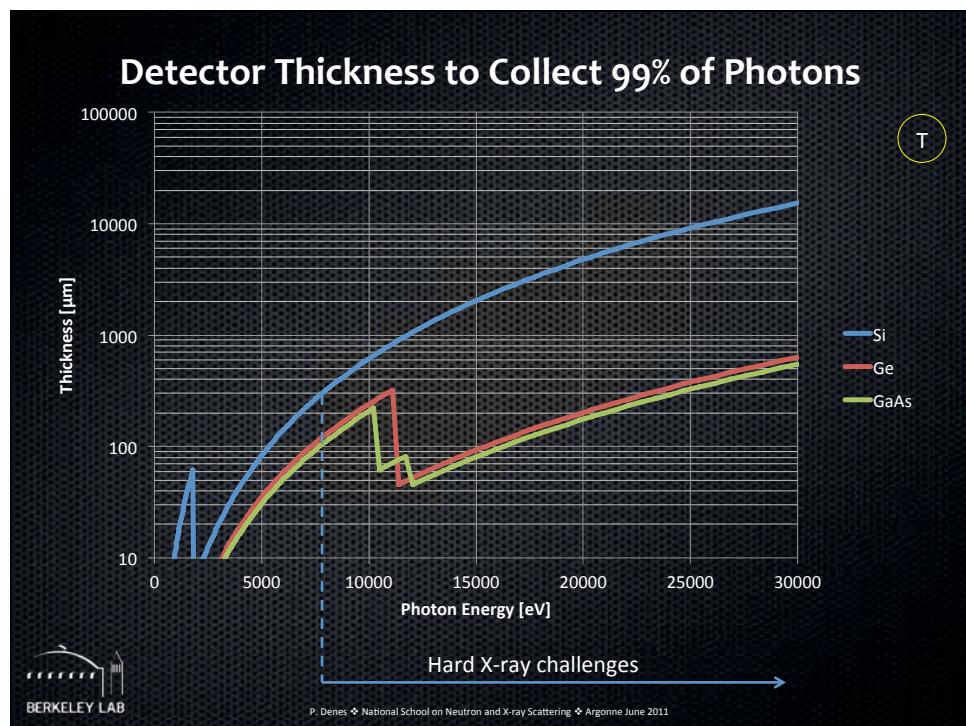
**Disadvantages:**

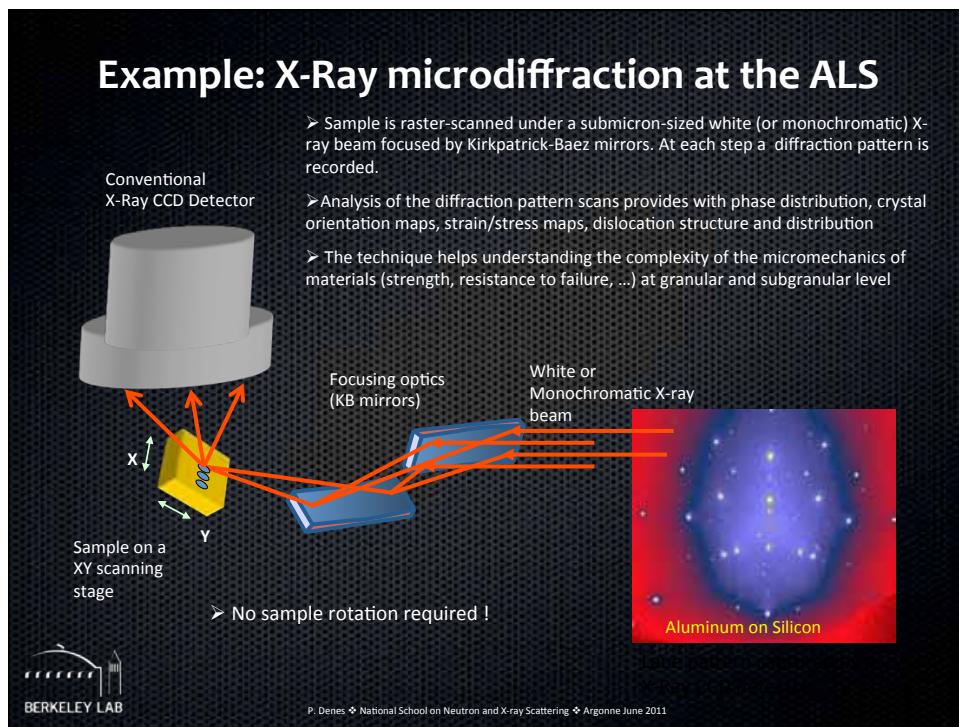
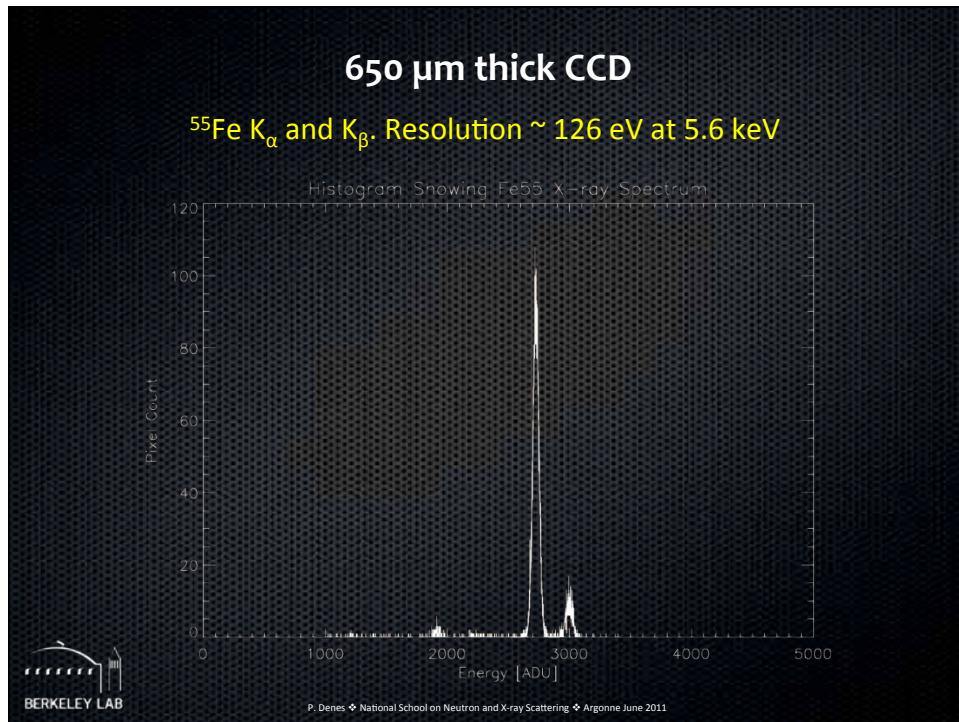
- 1) Enhanced sensitivity to radiation (x-rays, cosmic rays, radioactive decay)
- 2) More volume for dark current generation
- 3) Dislocation generation

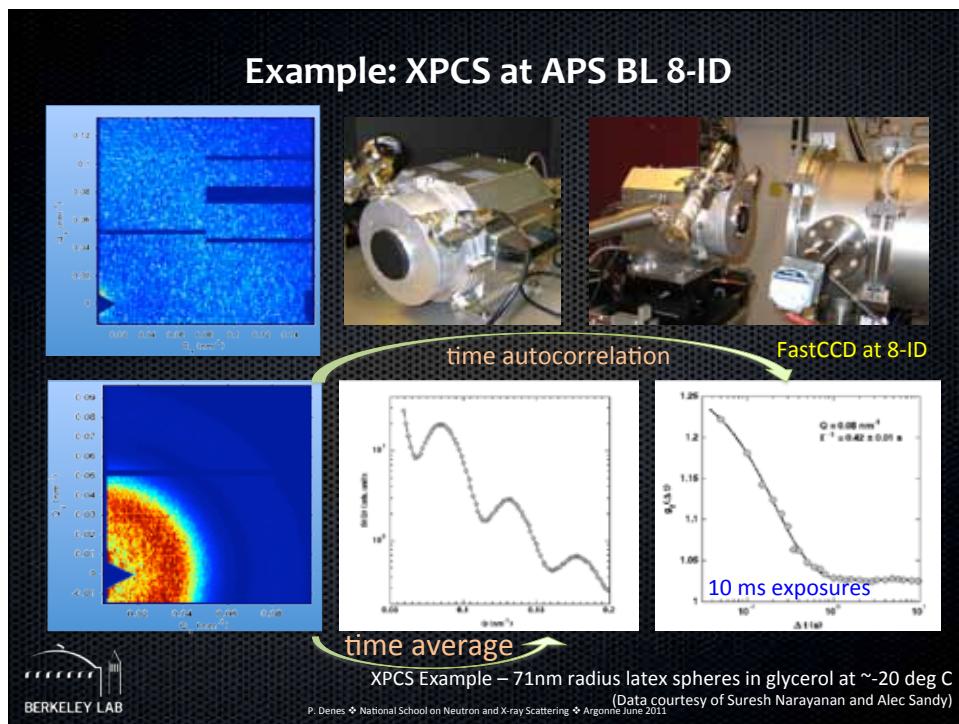
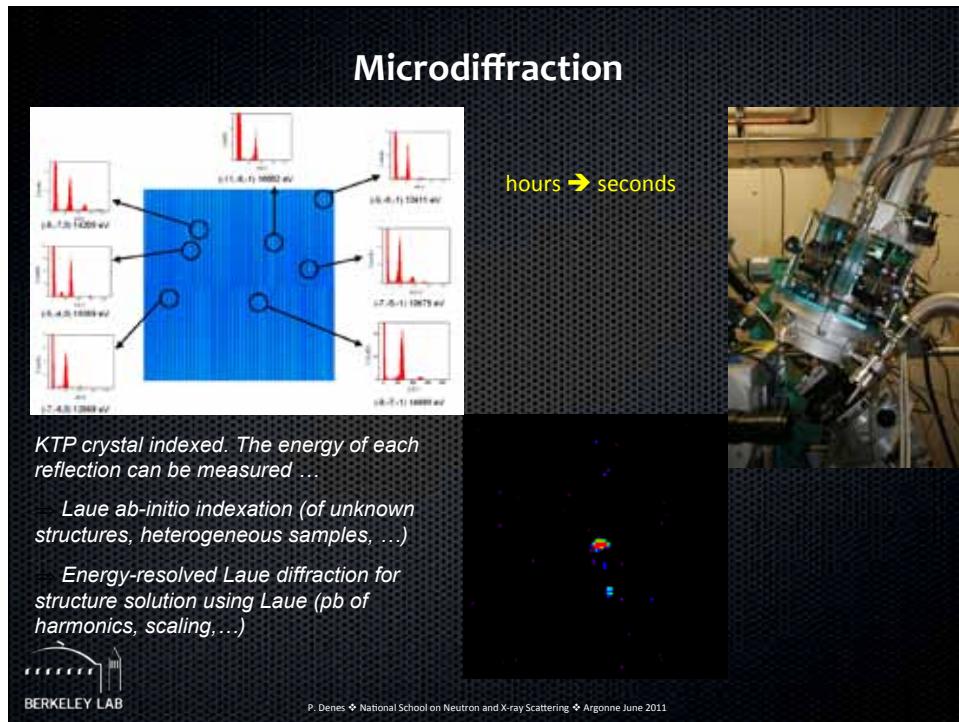
LBL CCD – S. Holland et al.

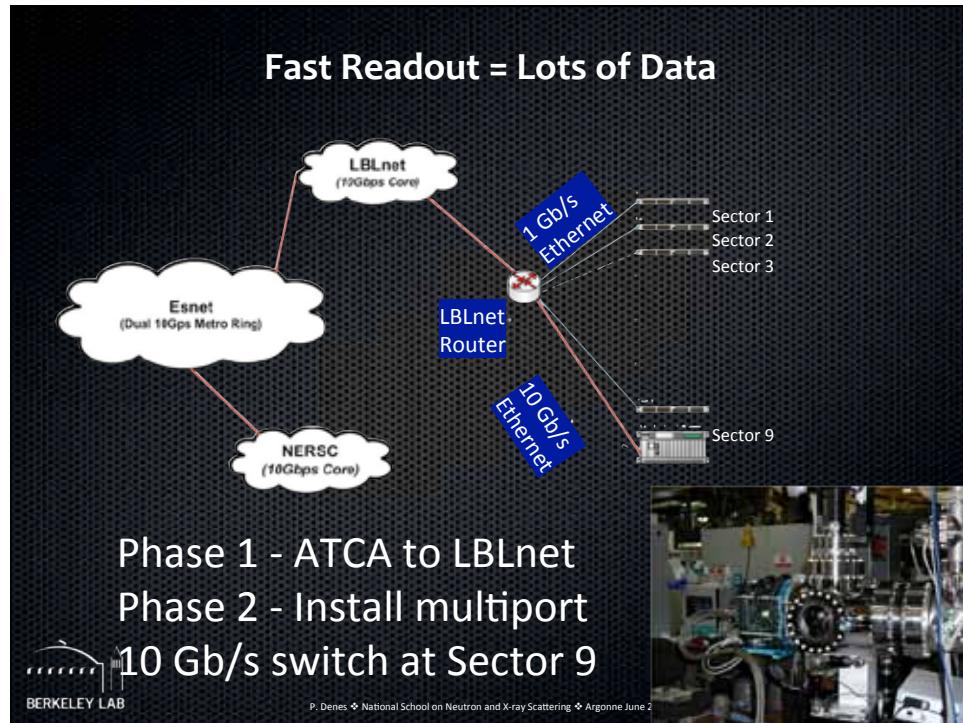
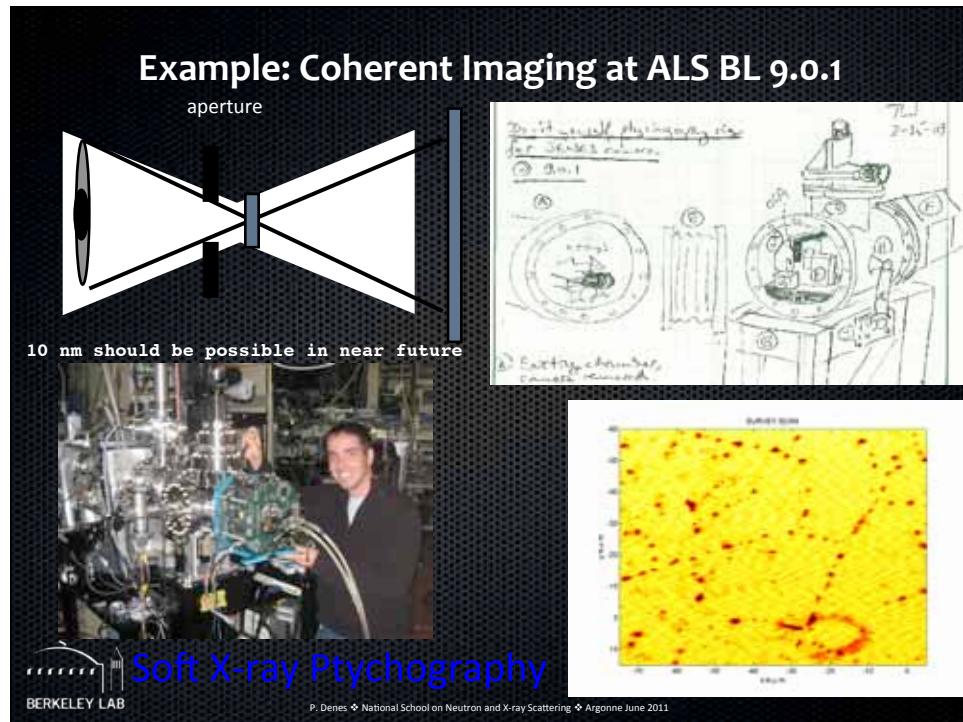
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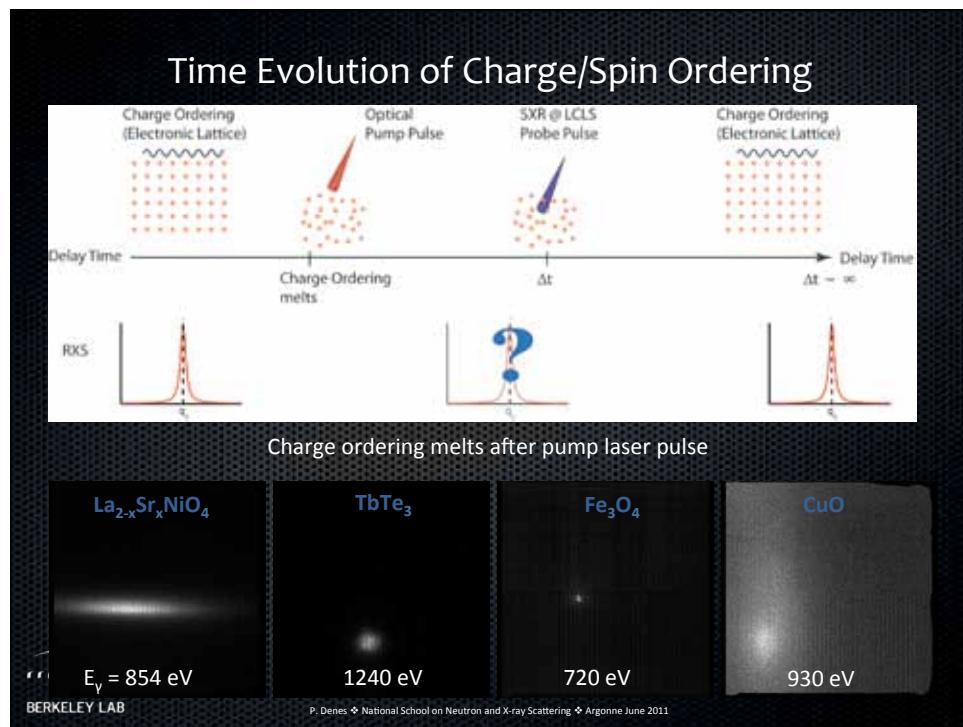
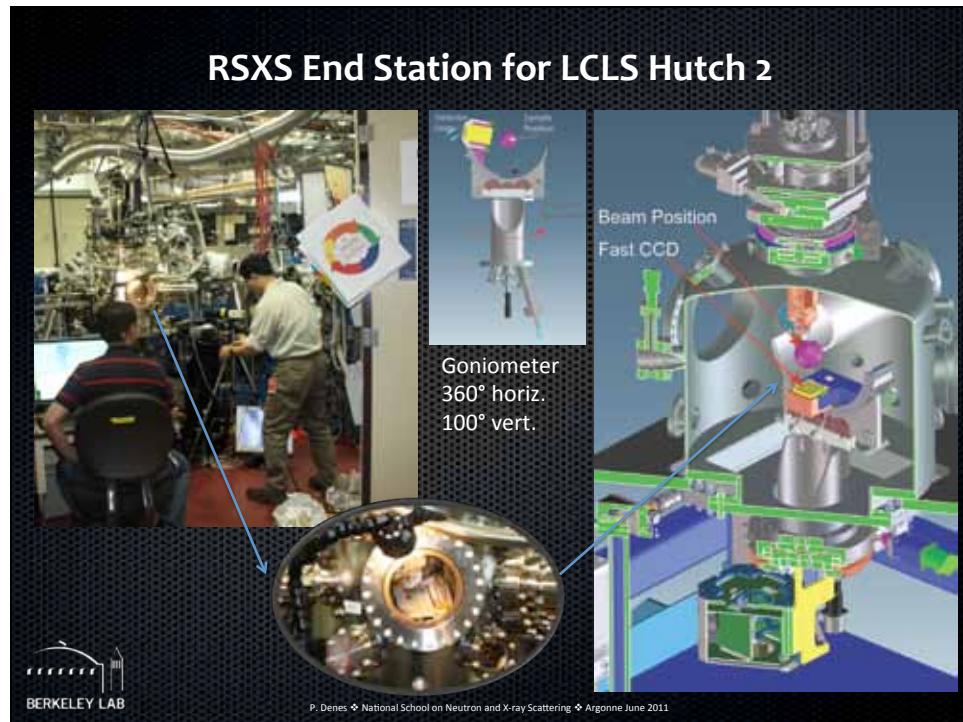
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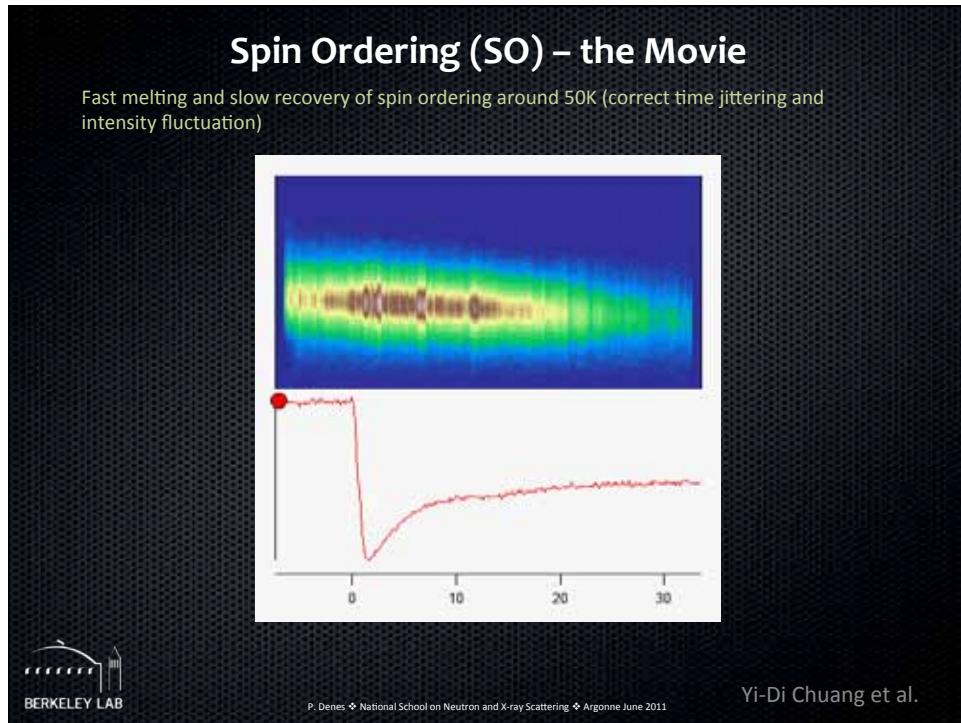
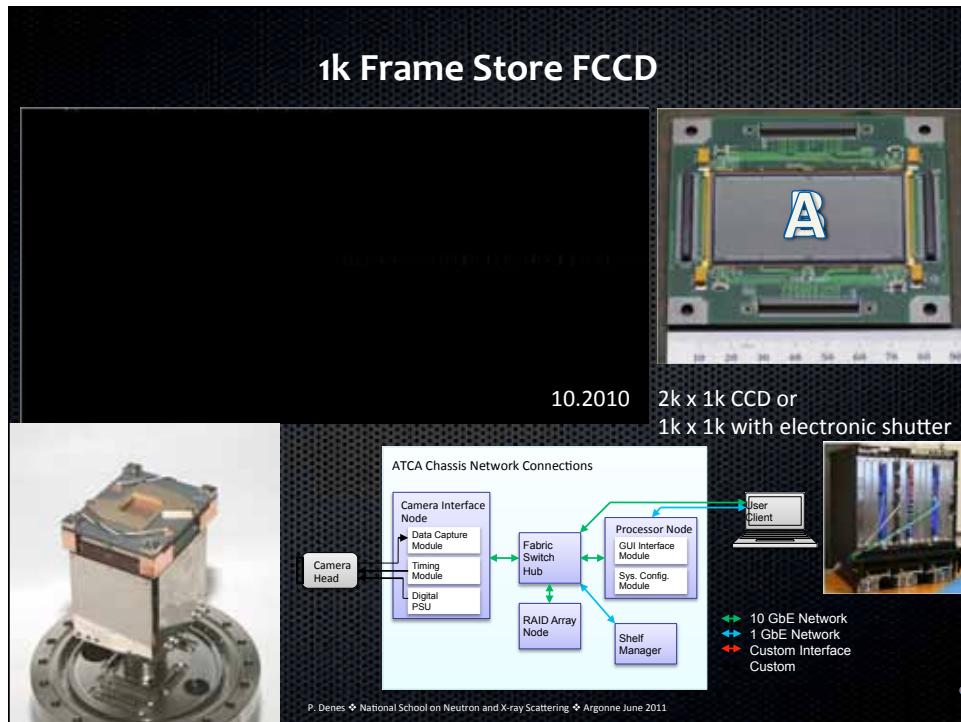


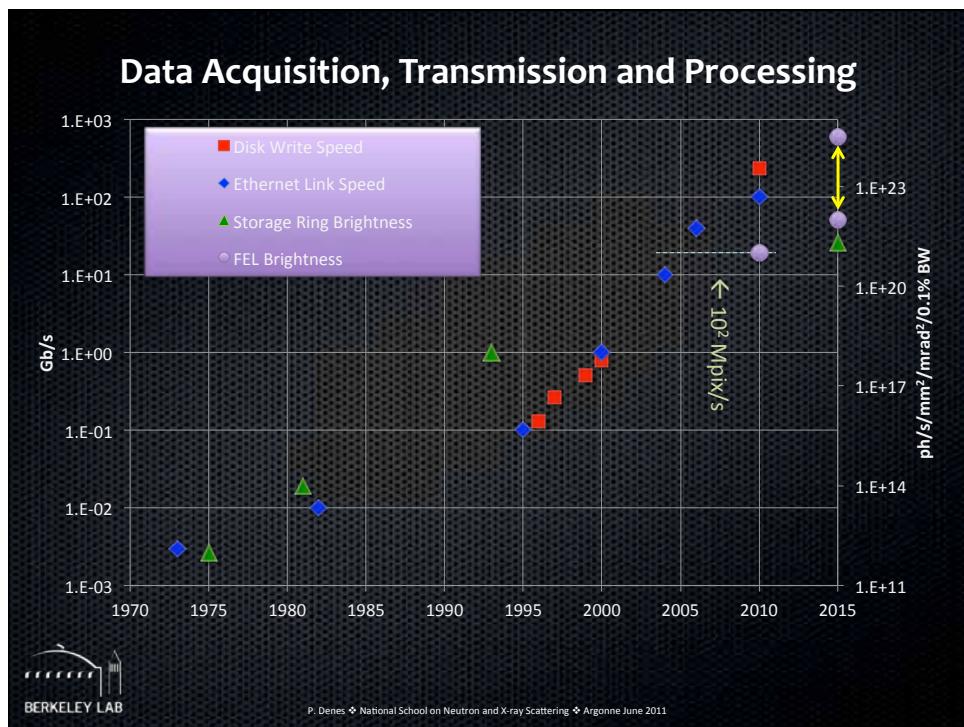
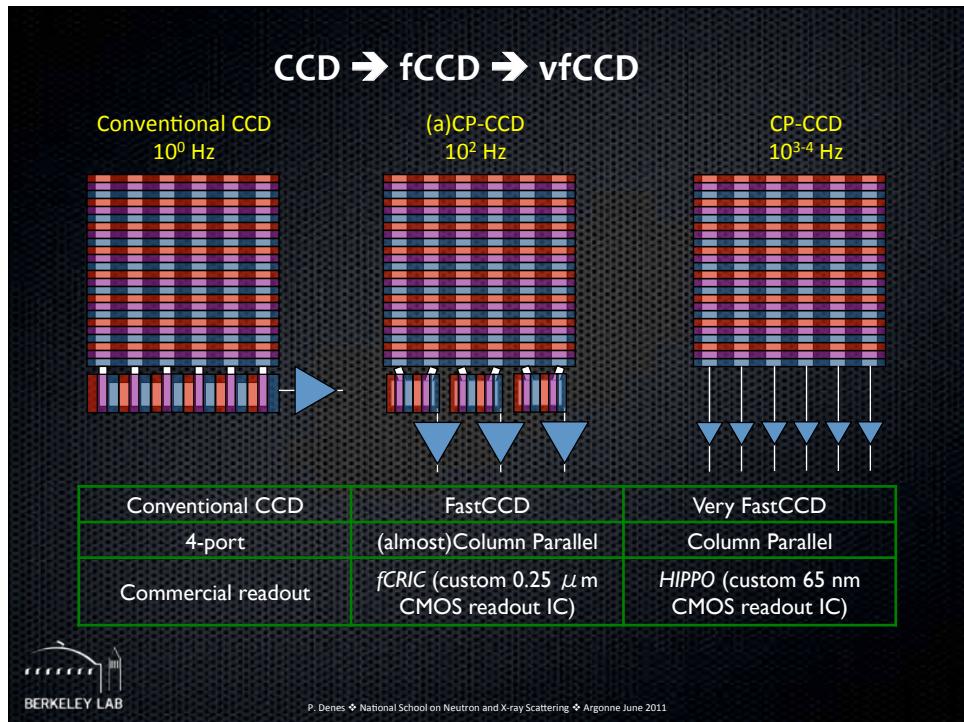


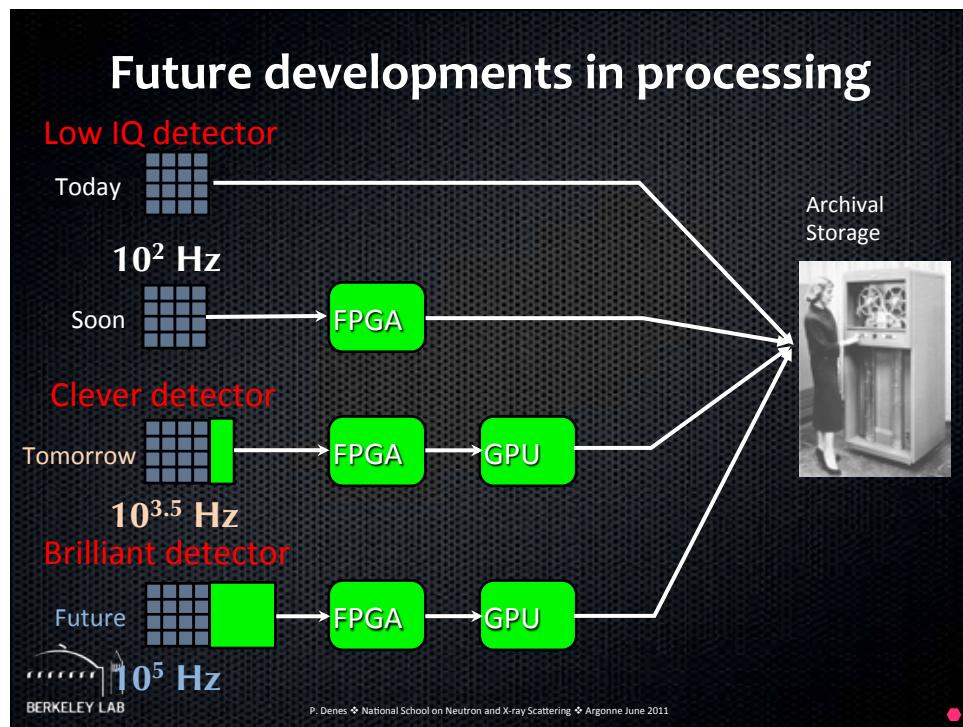
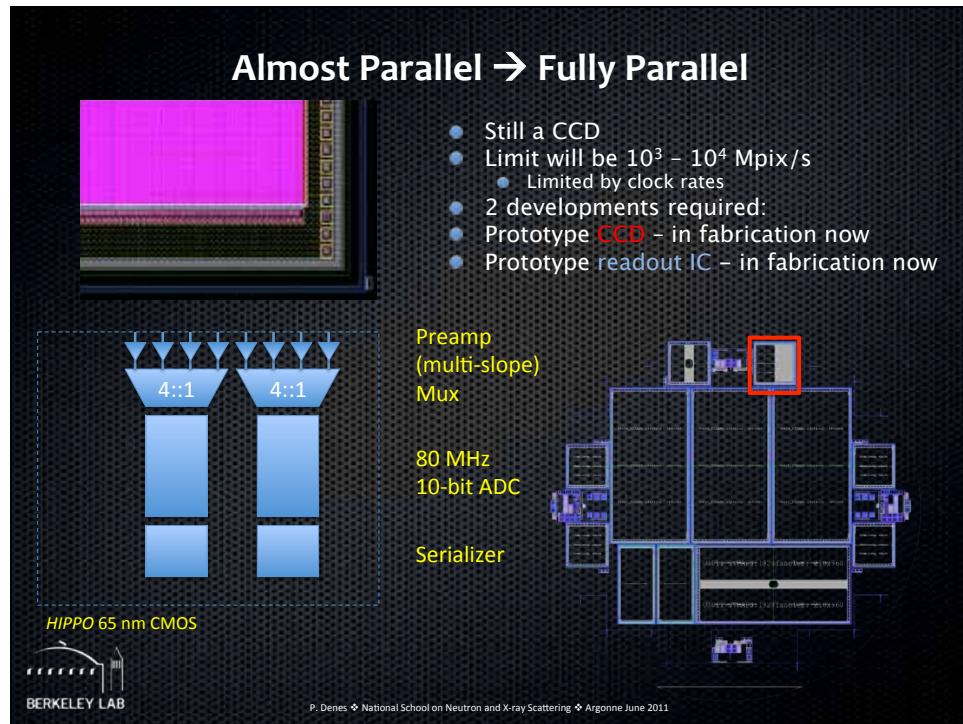


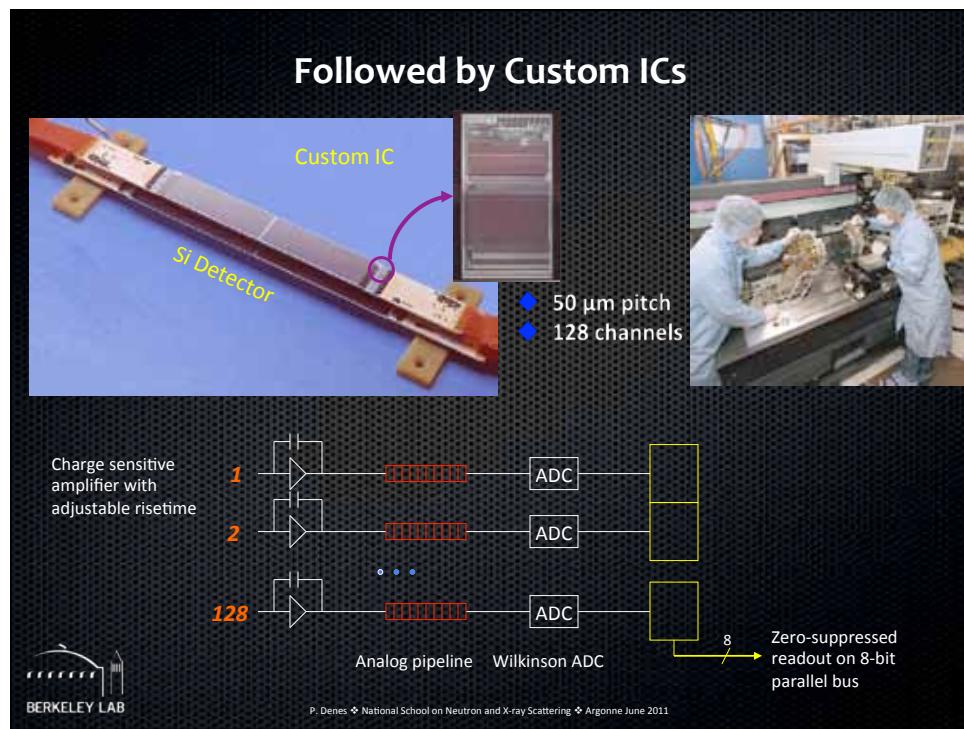
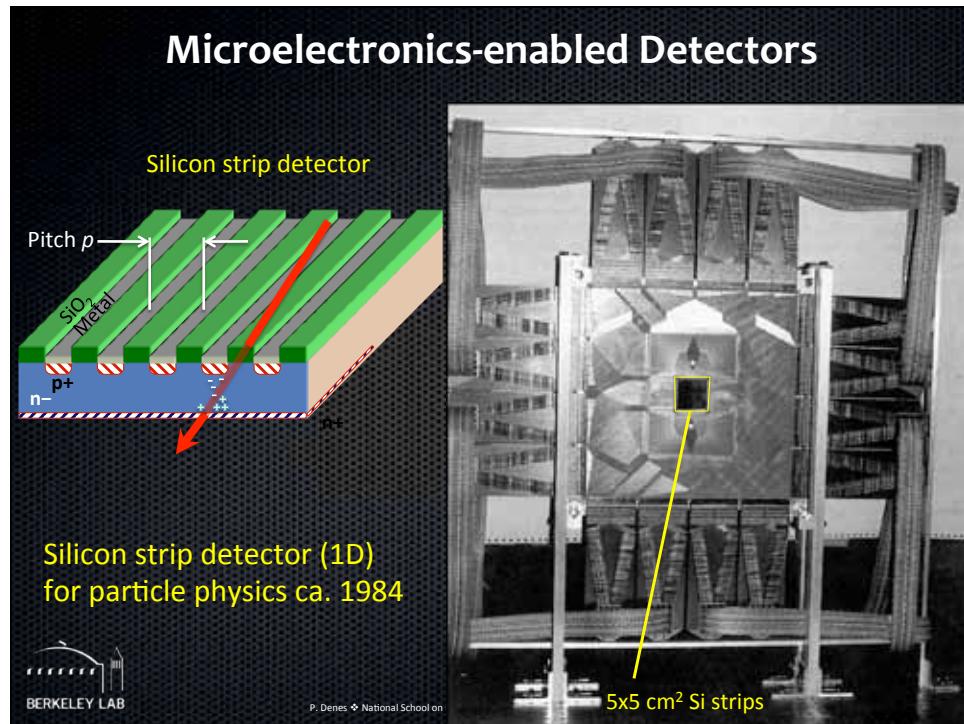


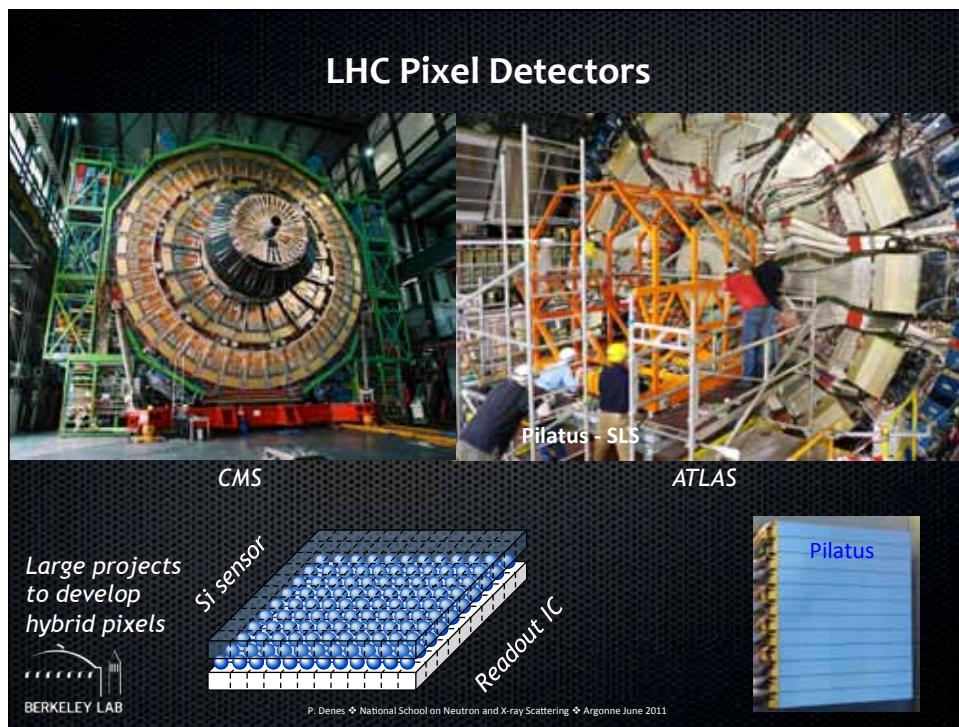
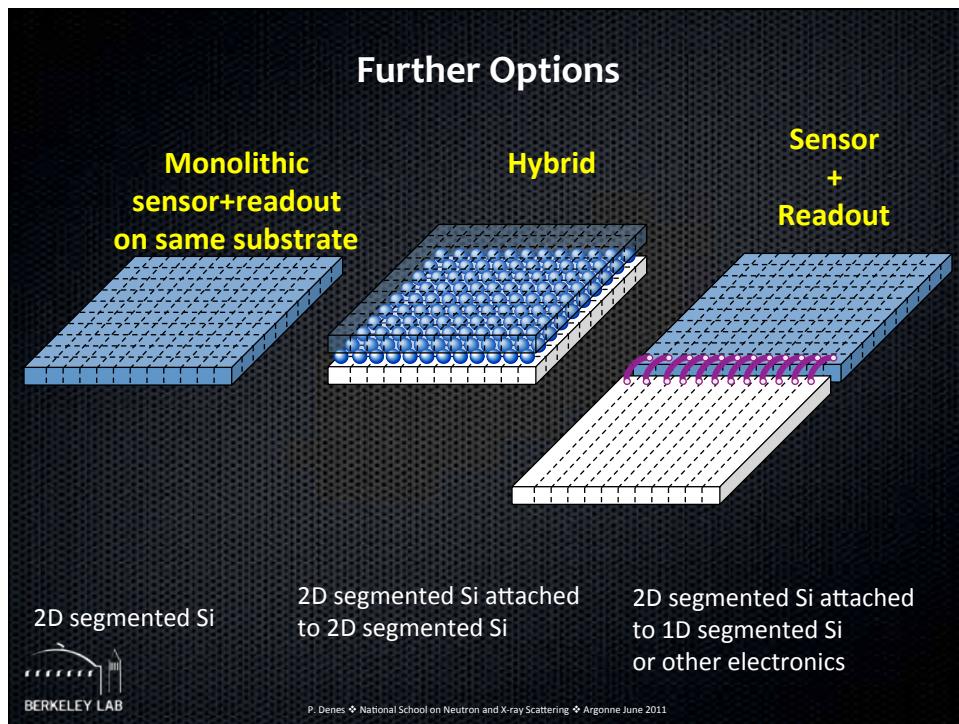


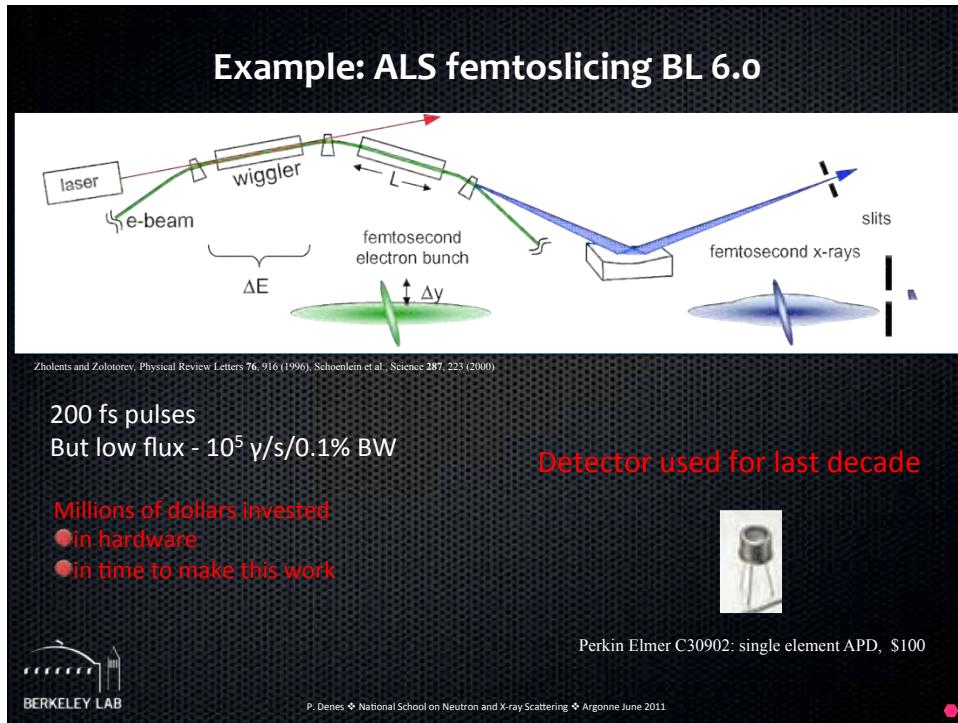
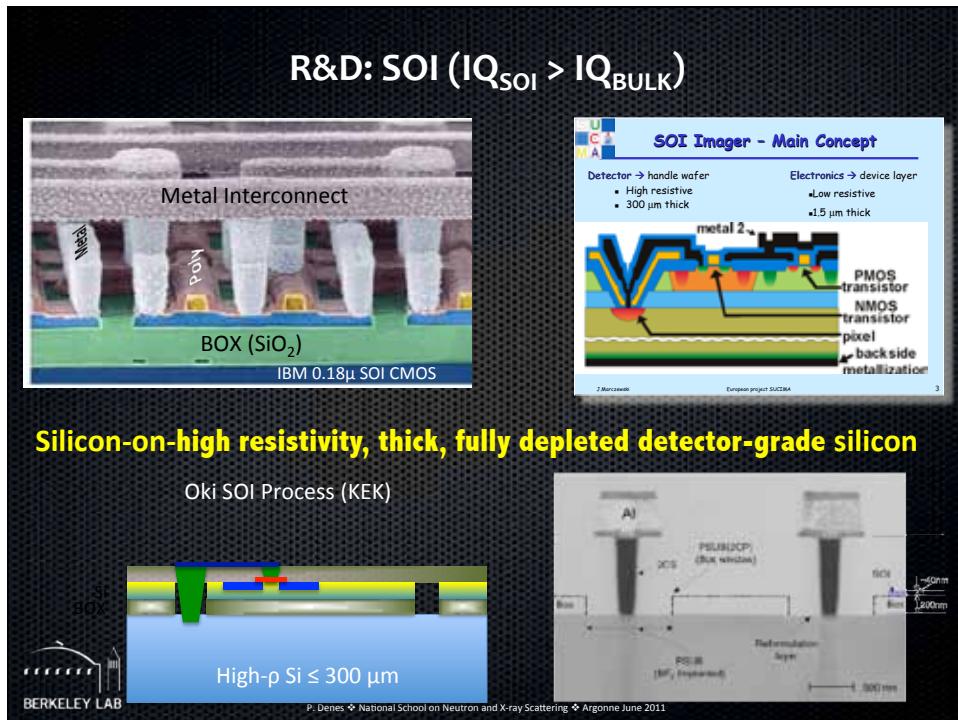


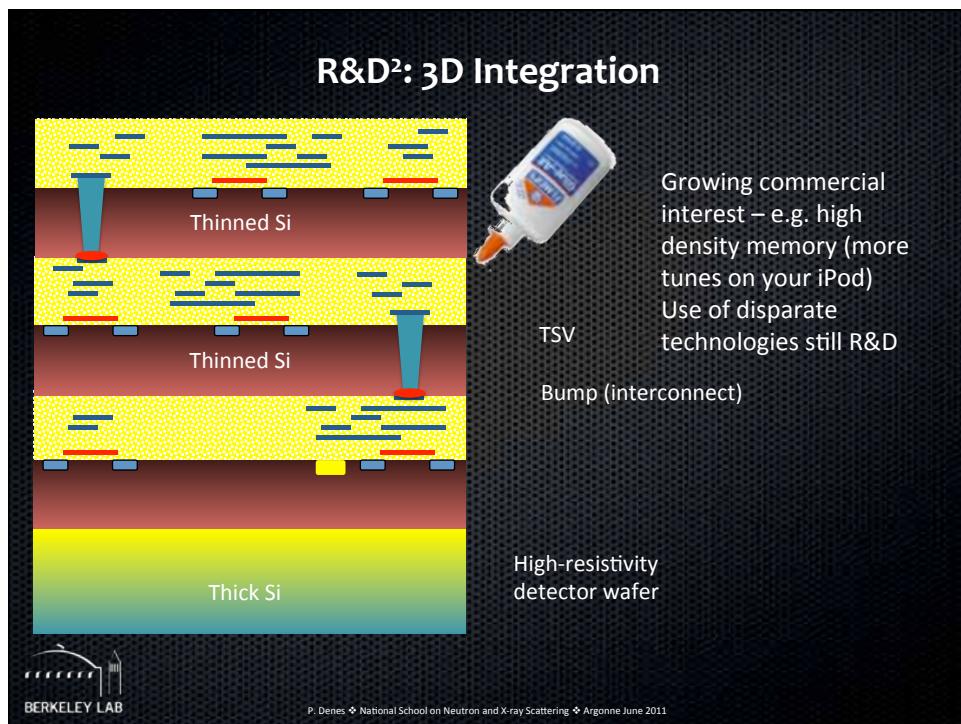
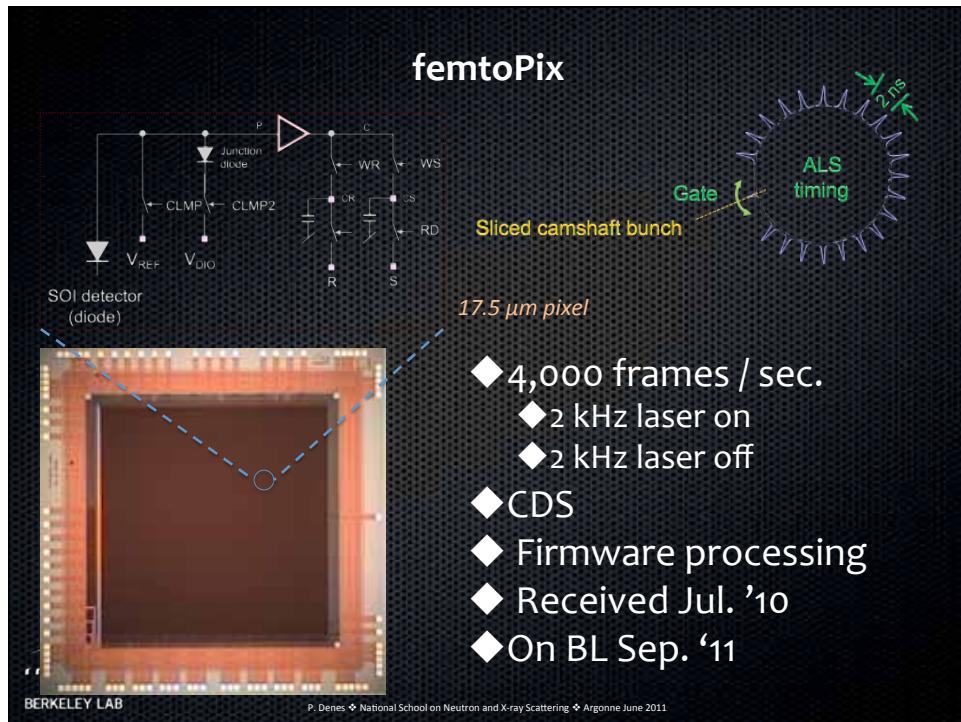


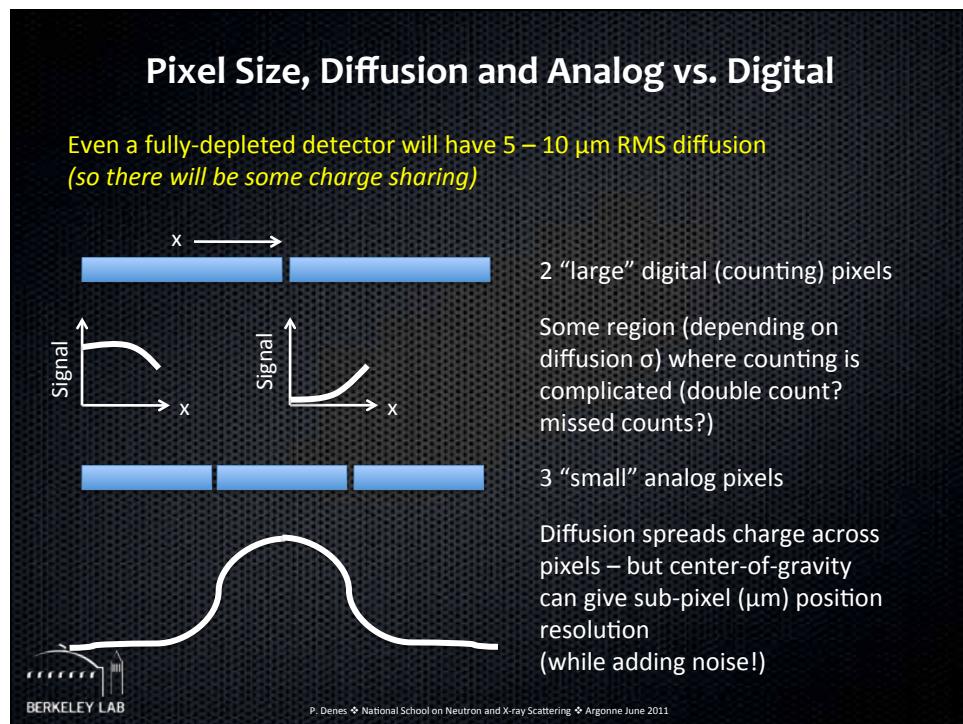
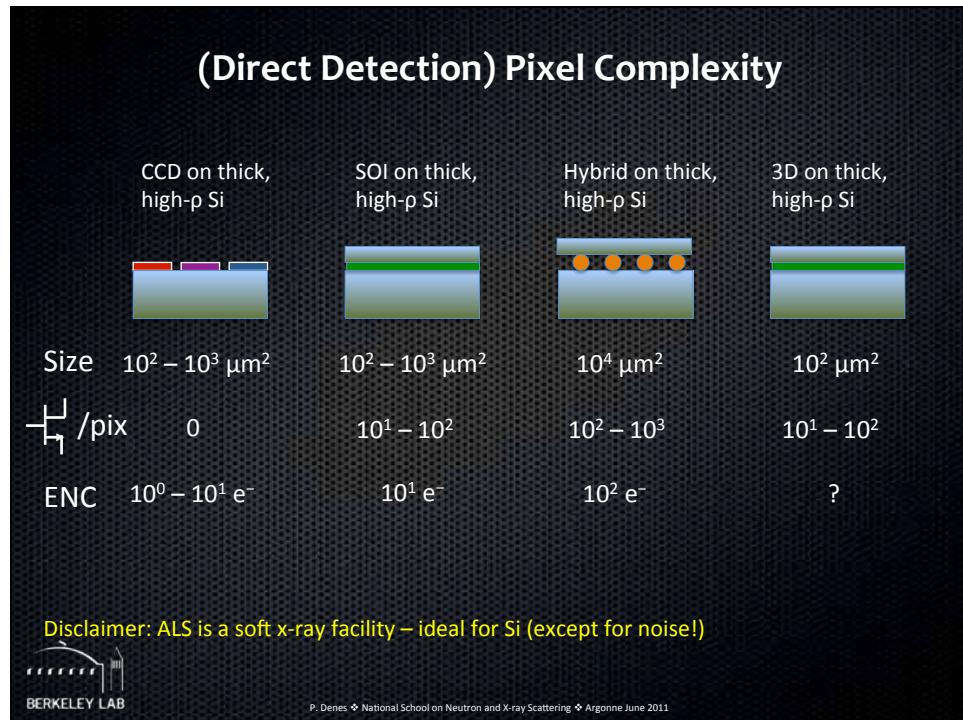












## Other Examples of 2D Detectors

**Large-area, flat-panel x-ray detector**

**Scintillator [e.g. CsI(Tl)]**

**aSi + TFT Passive Pixel readout**

**MCP**

**Photocathode**

**Hybrid pixel IC (or CCD)**

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## Summary (1)

- ◆ For a detector, the only useful thing a photon can do is create an electron
  - ◆ Note to accelerator people: the only useful thing an electron can do is create a photon
- ◆ Detection mechanisms
  - ◆ “Direct” (includes film, image plates, ... )
  - ◆ “Indirect” – usually via scintillator
- ◆ Sensor “properties” critical
  - ◆ Density (stopping power,  $\sigma_{PE}$ , ... )
  - ◆ Band gap, light yield, ...

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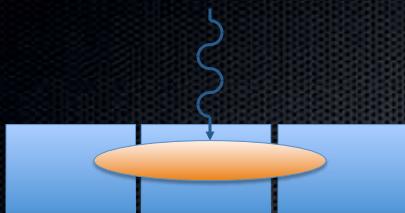
## Summary (2)

- ◆ Fluctuations
  - ◆  $0 \leq E_e \leq E_\gamma$  in “detector”
  - ◆ Number ( $N \propto E_e$ ) of secondary (tertiary) particles
  - ◆ Electronic noise
    - ◆ Thermal
    - ◆ Faster is (generally) noisier
- ◆ Spatial resolution (PSF, MTF) (diffusion)
- ◆ Temporal resolution (noise is important)
- ◆ DQE
- ◆ Radiation damage (not discussed, but important)



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## In other words



- ◆ \*Detectable = f(Electronics)
  - ◆ DQE  $\sim 1 / [\text{Electronic}] \text{ Noise}$
  - ◆ Many ways to say  $5\sigma$  (c.f. Rose criterion)
- ◆  $\sigma(E) \sim F \oplus \text{Noise}$
- ◆  $\sigma(t) \sim \text{Noise}$

- ◆ Photon incident at  $(0,0)$ 
  - ◆ Probability Q.E. of creating a detectable\* signal
  - ◆ Signal  $\propto 1/\eta$ 
    - ◆ Photostatistics
    - ◆ Fano factor
- ◆ Spatial resolution (PSF, MTF) (diffusion)



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## Summary (3)

- ◆ Like parking spaces, “no lack of detectors, only lack of imagination”
- ◆ Microelectronics-enabled detector development in particle physics starting to spill over into synchrotron radiation research
- ◆ Semiconductor detectors!
- ◆ DAQ, computing and processing!
- ◆ Si excellent for  $E < 10$  keV (and benefits from commercial processing)
  - ◆ Other developments, e.g. involving avalanche multiplication, that there was no time to discuss
  - ◆ For higher energies, have candidate materials (GaAs, Ge, CdTe, ...) but need R&D
- ◆ Future will be detectors designed for experiments (not experiments designed for detectors)



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## Questions?

**Grateful acknowledgements to:**  
 ALS Experimental Systems Group  
 ALS Scientific Systems Group  
 APS Beamline Technical Support Group  
 Electronic Systems Group  
 Integrated Circuit Design Group  
 MicroSystems Laboratory  
 National Center for Electron Microscopy  
 Physics Division  
 Engineering Division

