

2001 Physics GRE Solutions

1 Pendulum bob

This is *total* acceleration, so there will be two components acting on the bob: gravity and tension. Tension acts in the direction of the string, and gravity always points straight down. For this reason, we can immediately rule out (A) and (E), as those are clearly ignoring any force other than what is along the string. Furthermore, at point *c*, the string (and therefore the tension) is pointing in the same direction as gravity, so the arrow should be pointing totally along the string. This rules out (B) and (D).

Answer: (C).

2 Friction on a turntable

There are two forces which must cancel out: Centripetal and frictional.
Therefore:

$$m\omega^2 r = mg\mu_s, \quad (1)$$

$$\Rightarrow r = \frac{g\mu_s}{\omega^2}. \quad (2)$$

However, the units need to be taken care of. We are given a frequency $\nu = 33.3$ revolutions per minute, while $[g] = [\text{m}][\text{s}]^{-2}$. We want to convert $[\nu]$ to an angular frequency $[\omega] = [\text{rad}][\text{s}]^{-1}$.

$$\omega = 2\pi \frac{\text{rad}}{\text{rev}} \frac{1\text{min}}{60\text{s}} \nu, \quad (3)$$

$$\Rightarrow r = \frac{9.8 \times 0.3}{\left(33.3 \times \frac{2\pi}{60}\right)^2} = 0.242. \quad (4)$$

Answer: (D).

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3 Satellite orbit

I can't think of many clever ways to narrow down the answers. It is probably just good to have Kepler's Third Law memorized:

$$T = 2\pi\sqrt{\frac{a^3}{GM}}, \quad (5)$$

where

T is the orbital period,

a is the semi-major axis, and

GM is the planet's mass times the gravitational constant.

Answer: **(D)**.

4 One-dimensional inelastic collision

As usual, start with conservation of momentum:

$$m_1v_1 + m_2v_2 = (m_1 + m_2)v_{12}. \quad (6)$$

With $m_1 = 2m$, $m_2 = m$ and $v_2 = 0$ (the problem states that the particle is at rest), we have

$$2mv_1 = 3mv_{12} \quad (7)$$

$$\Rightarrow v_{12} = \frac{2}{3}v_1. \quad (8)$$

Then, with conservation of energy,

$$\frac{1}{2}m_1v_1^2 + \frac{1}{2}m_2v_2^2 = \frac{1}{2}(m_1 + m_2)(v_{12})^2 \quad (9)$$

$$\Rightarrow \frac{1}{2}2mv_1^2 = \left(\frac{1}{2}3m\right)\left(\frac{2}{3}v_1\right)^2 \quad (10)$$

$$\Rightarrow mv_1^2 = \left(\frac{3}{2}m\right)\frac{4}{9}v_1^2 \quad (11)$$

$$\Rightarrow KE_i = \frac{2}{3}KE_f \quad (12)$$

So the final energy is $\frac{2}{3}$ of the initial, meaning the system has lost a third of its kinetic energy.

Answer: **(C)**.

5 Average kinetic energy of a 3D harmonic oscillator

The equipartition theorem says that there is $\frac{1}{2}k_B T$ of energy for each degree of freedom *in the Hamiltonian*. The Hamiltonian for a harmonic oscillator is

$$H = \frac{1}{2m} (p_x^2 + p_y^2 + p_z^2) + \frac{1}{2}k (q_x^2 + q_y^2 + q_z^2), \quad (13)$$

which has 6 degrees of freedom (3 positions and 3 momenta), so the average kinetic energy is

$$\langle KE \rangle = 6 \times \frac{1}{2}k_B T = 3k_B T. \quad (14)$$

Answer: (D).

6 Work done on a gas

Adiabatic means that heat is conserved. **Adiabatic heating occurs when the pressure of a gas is increased from work done on it by its surroundings.** For an adiabatic process, $PV^\gamma = \text{const.}$, and the work done by applying pressure to a volume is $W = \int P dV$. Using these, one can derive the worthy-of-memorizing equation for the work done on a gas in an adiabatic process:

$$W_a = -\frac{1}{1-\gamma} (P_f V_f - P_i V_i). \quad (15)$$

Isothermal means that the temperature is conserved. **An isothermal process occurs slowly enough to allow the system to continually adjust to the temperature of the reservoir through heat exchange.** Therefore, $P_i V_i = nRT_i = P_f V_f = nRT_f$, and using the same definition of work above, one can derive the worthy-of-memorizing equation for the work done on a gas in an isothermal process:

$$W_i = nRT_i \ln \left(\frac{V_f}{V_i} \right) = P_i V_i \ln \left(\frac{V_f}{V_i} \right). \quad (16)$$

Right off the bat, we know that $V_f = 2V_i$, so neither W_i nor W_a will be equal to 0. This eliminates (B) and (D). Furthermore, we can see from the formulae that $W_i \neq W_a$. This makes the problem as deciding whether W_a is greater than W_i or vice-versa. One way to solve this would be by noting that for a monatomic gas, $\gamma = \frac{5}{3}$ and solving.

Another method is this: For an adiabat, $P \propto \frac{1}{V^\gamma}$, and $\gamma > 1$ for an ideal gas. For an isotherm, $P \propto \frac{1}{V}$. Therefore, for the same change in volume, the pressure of an adiabat changes more than the pressure of an isotherm. Therefore, the adiabat does more work.

Answer: (E).

7 Two bar magnets

Going through the options:

(A) cannot be true because both the magnets are standing on their north ends. Matching polarities always repel.

(C) seems to imply that there are magnetic monopoles, which don't exist. In this picture, $\nabla \cdot \mathbf{B} \neq 0$.

(D) doesn't work because you'd expect cylindrical symmetry.

(E) would be good if there were longitudinal current through the cylinders, but there's not.

Answer: (B).

8 Charge induced on an infinite conductor

By the conservation of charge, there should be an equal and opposite charge induced on the top face of the conductor, giving an answer of $-Q$.

Answer: (D).

9 Charges on a circle

All the charges are symmetrically placed around the circle, so at the center, there is an equal contribution from every direction pointing towards the charges, giving a net magnitude of 0. One suggestion from the Internet is to take the limit as infinite charges are equally placed around the circle: This is a conducting ring, which has no field on the interior.

Answer: (A).

10 Energy across a capacitor

The energy stored in a capacitor can be derived by finding the work required to establish an electric field in the capacitor:

$$W = \int_0^Q V(q) dq = \int_0^Q \frac{q}{C} dq = \frac{1}{2} \frac{Q^2}{C} = \frac{1}{2} CV^2 = \frac{1}{2} VQ, \quad (17)$$

where

V is the voltage difference across the capacitor,

C is the effective capacitance, and

Q is the total charge across one of the plates.

Furthermore, it is helpful to remember the additive properties of capacitors. In parallel, n capacitors add normally:

$$C_{\text{eff}} = C_1 + C_2 + \cdots + C_n. \quad (18)$$

In series, the reciprocals add:

$$\frac{1}{C_{\text{eff}}} = \frac{1}{C_1} + \frac{1}{C_2} + \cdots + \frac{1}{C_n}. \quad (19)$$

So, for this problem we need the effective capacitance of two capacitors in series:

$$\frac{1}{C_{\text{eff}}} = \frac{1}{3\mu\text{F}} + \frac{1}{6\mu\text{F}}, \quad (20)$$

$$= \frac{9}{18\mu\text{F}}, \quad (21)$$

$$\Rightarrow C_{\text{eff}} = \frac{18}{9}\mu\text{F}. \quad (22)$$

Finally, using $\frac{1}{2}CV^2$ for the energy stored, we get

$$\frac{1}{2} \times \frac{18}{9} \times 300^2 = 300^2 = 900,000. \quad (23)$$

Not knowing the units is fine: There is only one answer which is 9 times some number of tens.

Answer: **(A)**.

11 Two-lens image

This requires two applications of the thin-lens limit of the lensmaker's formula,

$$\frac{1}{d_o} + \frac{1}{d_i} = \frac{1}{f}, \quad (24)$$

where d_o is the distance from the object to the lens, d_i is the distance from the lens to the image, and f is the focal length of the lens.

For this problem, we apply the formula once for the object and the first lens, and then using that image as the “object” for the second lens. Solving for d_i for the first lens:

$$\frac{1}{d_i} = \frac{1}{20} - \frac{1}{40}, \quad (25)$$

$$= \frac{1}{40}, \quad (26)$$

$$\Rightarrow d_i = 40. \quad (27)$$

So the first image is 40 cm behind the first lens. This places the new “object” 10 cm behind the second lens. Because this is now on the opposite side of the second lens as the side which we established was positive for the first lens, we say that d_o for the second lens is -10 .

Then, following through with the second lensmaker’s equation,

$$\frac{1}{d_i} = \frac{1}{10} + \frac{1}{10}, \quad (28)$$

$$= \frac{1}{5}, \quad (29)$$

$$\Rightarrow d_i = 5. \quad (30)$$

So the final image is 5 cm behind the second lens.

Answer: **(A)**.

12 Spherical concave mirror

We can start with the lensmaker’s equation as used above,

$$\frac{1}{d_i} = \frac{1}{f} - \frac{1}{d_o}. \quad (31)$$

Looking at the picture in the problem, we can see that $d_o < f$. Therefore, $\frac{1}{d_o} > \frac{1}{f}$, and so d_i will be negative, placing the image behind the mirror.

Answer: **(E)**.

13 Resolving two stars

I can’t think of any way to do this problem but to memorize the Rayleigh criterion:

$$\theta = 1.22 \frac{\lambda}{D}, \quad (32)$$

which describes the requirement on the diameter of a lens, D , in order to resolve two objects separated by an angle θ , at some wavelength of light, λ .

In this particular problem, we have $\theta = 3 \times 10^{-5}$ rad, and $\lambda = 600$ nm. Plugging it all in, we have

$$D = \frac{1.22 \times 600 \times 10^{-9} \text{ m}}{3 \times 10^{-5} \text{ rad}} = 2.4 \text{ cm} \quad (33)$$

Answer: **(B)**.

14 Cylindrical gamma-ray detector

The “8-centimeter-long” fact is not useful. All that matters is the flux through the face of the detector. When the detector is *right at* the source, it’s collecting all of the radiation that leaves from that half of the source – hence it’s collecting 50% of emitted radiation. At a distance of 1 meter, the radiation is now spread out over a sphere of surface area $4\pi r^2 = 4\pi \text{ m}^2$. The face of the detector is only occupying $16\pi \times 10^{-4} \text{ m}^2$. Therefore, the fraction of radiation which is detected is a ratio of these two areas:

$$\frac{16\pi \times 10^{-4} \text{ m}^2}{4\pi \text{ m}^2} = 4 \times 10^{-4}. \quad (34)$$

Answer: (C).

15 Measurement precision

Precision is a measure of *random* error. If the precision is high, there is low random error. With low random error, the statistical variability is low, so the measurements are clustered closely regardless of how well they agree with the actual value.

Answer: (A).

16 Uncertainty

In addition to generally knowing about random error estimation, you need to know two things specifically for this problem: 1) that radioactive decay follows a Poisson distribution, and 2) that the standard deviation of a Poisson distribution is the square root of the average.

Then, you can estimate the expected uncertainty (usually taken as the error of the standard deviation, σ_x) from the fact that the error in the standard deviation, $\Delta\sigma_x$ goes as

$$\Delta\sigma_x = \frac{\sigma_x}{\sqrt{N}}, \quad (35)$$

for N trials.

So, calculating the average, we have

$$\langle x \rangle = \frac{3 + 0 + 2 + 1 + 2 + 4 + 0 + 1 + 2 + 5}{10} = 2, \quad (36)$$

so the standard deviation is

$$\sigma_x = \sqrt{\langle x \rangle} = \sqrt{2}. \quad (37)$$

Assuming that by “uncertainty of 1 percent” they mean that the error should be the average plus or minus 1%, we are looking for an uncertainty of $\Delta\sigma_x = 0.002$. So, using the fact

mentioned above, we can solve for the number of trials required for this uncertainty.

$$0.02 = \frac{\sqrt{2}}{\sqrt{N}}, \quad (38)$$

$$\Rightarrow N = \frac{2}{0.0004} = 5,000. \quad (39)$$

Answer: (D).

17 Filling electron shells

For problems like these, I think it's best to just memorize how electron shells are filled up. The rules I keep in mind are these:

1. *s*-orbitals can take 2 electrons, *p*-orbitals can take 6 and *d*-orbitals can take 10.
2. The pattern is *s-s-p-s-p-s-d*.
3. If the orbitals are in the right order, and they each have as many electrons as they can take (except, perhaps the last one), then that's the right answer.

Answer: (B).

18 Energy to remove a helium electron

The Bohr model is applicable here, because helium isn't too different from hydrogen. The energy for an electron in the n^{th} state of a hydrogenic atom with Z protons is given by

$$E_n = Z^2 \frac{R_E}{n^2}, \quad (40)$$

where R_E is a Rydberg, the energy of the ground state of hydrogen, 13.6 eV.

Because Helium has two protons ($Z = 2$), it can accept two electrons into its ground state, so $n = 1$ for this problem. If the helium were singly ionized and only had a single electron, as He^+ , then the energy to remove that electron would be, from the Bohr model above,

$$E_1 = 4 \times 13.6 \text{ eV} = 54.4 \text{ eV}. \quad (41)$$

It's tempting to say that the energy required to remove the two electrons of neutral helium would just be $2 \times 54.4 \text{ eV} = 108.8 \text{ eV}$, but this is not the case. Considering one electron at a time, one can see that the first one is not *just* experiencing the tug of the nucleus, but is *also* being pushed away by a Coulomb repulsion from the *other electron*. Therefore, we expect it to be slightly easier to remove the first of two electrons than to remove one solitary electron.

In lieu of perturbation theory, we can subtract the theoretical value of 54.4 eV for a single ground state energy from the experimental value of 79 eV for complete ionization. What's left over will be the energy that was required to remove an electron whose exist was being pushed along by a Coulomb repulsion. So,

$$79 \text{ eV} - 54.4 \text{ eV} = 24.6 \text{ eV}. \quad (42)$$

Answer: (A).

19 Source of the Sun's energy

It is not necessary to know the whole proton-proton chain to answer this problem. Just note that a helium has 4 things of roughly proton mass (2 protons and 2 neutrons), and that $n + e^- \rightarrow p$. So in order to make helium, you need 4 hydrogens and enough energy to collide half of the electrons and protons.

Answer: (B).

20 Definition of bremsstrahlung

The way I remember this is that bremsstrahlung is German for “breaking radiation” which reminds me that it's analogous to breaking the sound barrier (it doesn't matter that they're different meanings of “break”). The only answer that sounds like it has anything to do with breaking a sound barrier is the one that mentions rapid deceleration in a medium.

Answer: (E).

21 Comparing Lyman- α and Balmer- α wavelengths

This requires knowing the Rydberg formula for the energy, E , of a photon released in the transition from one energy level, n_i to another, n_f :

$$E = E_f - E_i = R_E \left(\frac{1}{n_f^2} - \frac{1}{n_i^2} \right), \quad (43)$$

where R_E is the Rydberg energy, 13.6 eV. The energy of a photon emitted by a Lyman- α transition (from $n = 2$ to $n = 1$) is then

$$E_{L\alpha} = R_E \left(1 - \frac{1}{4} \right) = \frac{3}{4} R_E. \quad (44)$$

While Balmer- α transitions (from $n = 3$ to $n = 2$) release photons with energies of

$$E_{B\alpha} = R_E \left(\frac{1}{4} - \frac{1}{9} \right) = \frac{5}{36} R_E, \quad (45)$$

and, taking the ratio,

$$\frac{E_{L\alpha}}{E_{B\alpha}} = \frac{3 \times 36}{4 \times 5}, \quad (46)$$

$$= \frac{27}{5}. \quad (47)$$

Finally, the problem is asking for the wavelength of the light. $E = hc/\lambda$, and taking the ratio has dealt with the implicit hc , so we only need to take the reciprocal of $27/5$.

Answer: (B).

22 Orbital system parameters

The hint is in the phrase “very small moon”. The question is implying that the mass of the moon is negligible compared to the mass of the planet. This means that the center of gravity will be very close to the center of the planet, and so you can treat the planet as stationary. Then, think about the equations of orbital motion. Kepler’s Third Law (problem 3) describes the period and semi major axes of the orbit, but makes no mention of mass. If you’re solving for a circular orbit with forces (or any orbit, but circular is easy), the smaller mass will cancel:

$$\frac{GMm}{r^2} = \frac{mv^2}{r} \quad (48)$$

There are probably other examples, but these should be sufficient to convince you the answer is the moon’s mass.

Answer: (A).

23 Particle accelerating around a circle

I think this problem is worded confusingly. There’s a particle going around a circle at a rate of 10 ms^{-1} , and so it’s experiencing the standard centripetal acceleration of

$$\mathbf{a}_c = -\frac{v^2}{r} \hat{\mathbf{r}} = -\frac{100}{10} \hat{\mathbf{r}} = -10 \hat{\mathbf{r}}. \quad (49)$$

But this particle is *also* undergoing a tangential acceleration of 10 ms^{-2} , in the direction of the particle’s velocity. That means that the particle is experiencing two orthogonal

accelerations of equal magnitude. The only way that these can combine is to form a vector at 45° to both. To answer the question, note that one of the accelerations is in the same direction as the velocity, so the angle is the same.

Answer: (C).

24 Throwing a stone

There is no air resistance, so v_x should be constant, so we should only consider solutions with plot II in the first column. The ball starts by going up, so v_y is at positive at first, and should undergo constant deceleration. As the ball begins to fall, v_y is negative. The only plot which fits this description is III.

Answer: (C).

25 Moment of inertia of packed pennies

To do this problem, refer to the front of the test for the moment of inertia of a solid disk about an axis going through its center: $I = \frac{1}{2}mr^2$. A very quick way I did this was to say that the whole collection of pennies was a solid disk of mass $7m$ and radius $3r$. Then,

$$I = \frac{1}{2}mr^2, \quad (50)$$

$$m \rightarrow 7m, r \rightarrow 3r, \quad (51)$$

$$= \frac{1}{2} \times 7m \times (3r)^2, \quad (52)$$

$$= \frac{63}{2}mr^2. \quad (53)$$

The available answer this is closest to is $55/2$, which is the correct answer. However, if the correct answer weren't also the largest, this estimate is not nearly close enough to be sure. A rigorous way to do this is to use the parallel axis theorem.

If you know the moment of inertia of an object around one axis, the parallel axis theorem provides a way to calculate the moment of inertia around any other axis which is parallel to the original. In this case, we know the moment of inertia of one penny around the axis that goes through its center of mass,

$$I_{\text{CM}} = \frac{1}{2}mr^2 \quad (54)$$

If we now rotate that penny around an axis a distance d from its center, the parallel axis theorem says that the new moment of inertia is given by

$$I = I_{\text{CM}} + md^2 \quad (55)$$

In calculating the contribution to the total angular momentum of each extremal penny in the arrangement given in the problem, we are rotating each of the exterior pennies around an axis $2r$ from their centers. Then, we can add these moments to the moment of the middle penny being rotated around *its* center to get the total moment of inertia. For the moment of inertia of each exterior penny,

$$I_{\text{ext.}} = I_{\text{CM}} + m(2r)^2, \quad (56)$$

$$= \frac{1}{2}mr^2 + m \times 4r^2, \quad (57)$$

$$= \frac{9}{2}mr^2. \quad (58)$$

So each exterior penny contributes $(9/2)mr^2$ to the total moment of inertia which we can now calculate by summing the moments of all the constituents.

$$I_{\text{tot.}} = I_{\text{CM}} + 6 \times \frac{9}{2}mr^2, \quad (59)$$

$$= \frac{1}{2}mr^2 + \frac{54}{2}mr^2, \quad (60)$$

$$= \frac{55}{2}mr^2. \quad (61)$$

Answer: (E).

26 Speed of the end of a falling rod

This problem also requires looking at the front of the test for (or memorizing) the moment of inertia of a rod of length L and mass M about its center of mass (so that it's spinning like a baton): $I = \frac{1}{12}ML^2$. However, the rod in the problem is not rotating around its center of mass. It's pivoting on one end, a distance $L/2$ from its center of mass. We can get the moment of inertia of this rotation with the parallel axis theorem.

$$I = I_{\text{CM}} + Md^2, \quad (62)$$

$$d \rightarrow \frac{L}{2}, \quad (63)$$

$$I = M\frac{L^2}{12} + M\frac{L^2}{4}, \quad (64)$$

$$= \frac{1}{3}ML^2. \quad (65)$$

Now we can use conservation of energy to solve for the speed of the end of the rod. Setting the rotational energy equal to the gravitational potential energy from the rod's center of

mass (which is located at $L/2$),

$$\frac{1}{2}I\omega^2 = Mg\frac{L}{2}, \quad (66)$$

$$\Rightarrow \omega^2 = \frac{MgL}{I}, \quad (67)$$

$$= \frac{3g}{L}. \quad (68)$$

To convert the radial speed, ω to the linear speed of the end of the rod, we note that $\omega = v/r$, and in the case of the end of the rod, $r = L$. So,

$$\frac{v^2}{L^2} = \frac{3g}{L}, \quad (69)$$

$$\Rightarrow v = \sqrt{3gL}. \quad (70)$$

Answer: (C).

27 Eigenvalues of the Hamiltonian operator

An axiom of quantum mechanics is that the Hamiltonian operator is hermitian, and a requirement of hermitian operators is that their eigenvalues are always real. If the Hamiltonian operator weren't hermitian, then observations of a particle's energy could give imaginary results, which would have no physical meaning.

Answer: (A).

28 Orthogonal states

That two vectors are orthogonal means their inner product is equal to zero. Taking the inner product of ψ_1 and ψ_2 ,

$$\langle \psi_1 | \psi_2 \rangle = (5 \times 1) + ((-3) \times (-5)) + (2 \times x) = 20 + 2x \quad (71)$$

The problem asks for the value of x which ensures ψ_1 and ψ_2 are orthogonal, which, from demanding the above inner product be equal to zero, is -10 .

Answer: (E).

29 Expectation value of an operator

The whole point of eigenvalues is that you can replace an operator acting on an eigenvector with the corresponding eigenvalue, a scalar. In this case, we take the inner product of the

wavefunction with the wavefunction after it has been acted upon by the operator \hat{O} ,

$$\langle \psi | \hat{O} \psi \rangle. \quad (72)$$

To calculate $\langle \hat{O} \psi \rangle$, we can just push the operator through ψ ,

$$\langle \hat{O} \psi \rangle = \frac{1}{\sqrt{6}} \hat{O} \psi_{-1} + \frac{1}{\sqrt{2}} \hat{O} \psi_1 + \frac{1}{\sqrt{3}} \hat{O} \psi_2, \quad (73)$$

$$\text{replacing the operator with the eigenvalues,} \quad (74)$$

$$= \frac{1}{\sqrt{6}} (-1) \psi_{-1} + \frac{1}{\sqrt{2}} (1) \psi_1 + \frac{1}{\sqrt{3}} (2) \psi_2, \quad (75)$$

$$= -\frac{1}{\sqrt{6}} \psi_{-1} + \frac{1}{\sqrt{2}} \psi_1 + \frac{2}{\sqrt{3}} \psi_2. \quad (76)$$

Then, calculating the inner product with the original wavefunction,

$$\langle \psi | \hat{O} \psi \rangle = \left(-\frac{1}{\sqrt{6}} \times \frac{1}{\sqrt{6}} \right) + \left(\frac{1}{\sqrt{2}} \times \frac{1}{\sqrt{2}} \right) + \left(\frac{1}{\sqrt{3}} \times \frac{2}{\sqrt{3}} \right), \quad (77)$$

$$= \frac{1}{6} + \frac{1}{2} + \frac{2}{3} = 1. \quad (78)$$

Answer: (C).

30 Possible Hydrogen wavefunctions

Any wavefunction must be normalizable, which is to say

$$\int_{-\infty}^{\infty} \psi^* \psi dr = 1, \quad (79)$$

because between the limits of $\pm\infty$, the particle must be *somewhere* (that is, if you're in position space; the analogous requirement in momentum space is that the particle must have *some* finite momentum). The only function which is normalizable is option I.

Answer: (A).