

# Numerical solution of the 1D grey radiation hydrodynamics equations with an entropy-based artificial viscosity.

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## Abstract

The entropy viscosity method is extended to the 1-D grey radiation-hydrodynamic equations. The method employs a viscous regularization to stabilize the numerical scheme with a viscosity coefficient modulated by the entropy production that is known to be peaks in shocks. The dissipative terms are consistent with the entropy minimum principle which requires a new functional form of the entropy residual to be derived, suitable for radiation-hydrodynamic equations. The equations are discretized with a standard Continuous Galerkin Finite Element Method (CGFEM) and implicit temporal integrator within the Moose multiphysics framework. The method of manufactured solution (MMS) is employed to demonstrate second-order accuracy in both the equilibrium diffusion and streaming limits. Several typical 1-D radiation-hydrodynamic test cases with shocks (from Mach 1.05 to Mach 50) are computed to establish the ability of the technique at capturing and resolving shocks.

**Keywords:** radiation-hydrodynamics, artificial viscosity, entropy viscosity method.

# 1 Introduction

Solving the radiation hydrodynamic equations is a challenging task for multiple reasons. First, the characteristic time scales between the radiation and hydrodynamics are different by several orders of magnitude which often requires the radiation part to be solved implicitly to ensure stability. Second, as with any wave-dominated problems, high resolution schemes are needed to accurately resolve shocks. Third, achieving high-order accuracy is challenging but some recent developments provided high-order accuracy results both in time and space when discretized either Euler equations [1, 2, 3, 4] or the radiation equation independently from each other.

Significant effort has been put into developing Riemann solvers for both the radiation and hydrodynamic equations. Balsara [5] developed a Riemann solver for the radiation-hydrodynamic equations by considering the frozen approximation that decouple the two physics components. However, such an approach can reveal itself questionable in the equilibrium diffusion limit. In this case, the coupling terms drive the physics and have to be accounted for. A *generalized Riemann solver* which accounts exactly for the relaxation terms was developed in [6]. Another approach, applicable in the equilibrium limit, consists of recasting the system of equations in a conservative form. Then, a numerical solution can be obtained by modifying the equation of state, yielding the radiation-modified equation of state (REOS). This is straightforward to implement in any solver for Euler equations by replacing the equation of state by the REOS [7]. Edwards and al. proposed a two stage semi-implicit scheme IMEX to solve the coupled radiation-hydrodynamic equations [10]. They applied a Trapezoidal/BDF2 temporal discretization scheme to the nonlinear grey radiative diffusion. The radiation and hydrodynamic equations are solved implicitly and explicitly, respectively. A Riemann solver along with a flux limiter is used to resolve shocks and other waves. Their results show good agreement with semi-analytical solutions. In this article we propose to solve the 1-D radiation-hydrodynamics equations by using *the entropy viscosity method*. This technique, developed by Guermond et al. for hyperbolic system of equations [2, 3], consists in adding appropriate dissipative terms to the governing equations whose viscosity coefficient is modulated by the local entropy production. These dissipative terms are devised so as to stabilize the numerical scheme and to remove the non-physical oscillations appearing at the shock locations. Generally speaking, entropy is produced at shocks [11]. Thus, by

setting the viscosity coefficient proportional to the entropy production, shocks can be detected and tracked, and an adequate amount of viscosity is added locally to stabilize the numerical scheme. The entropy production is computed on the fly, by analyzing the entropy residual; this residual is strongly peaked in shocks and small elsewhere. The entropy viscosity method has been shown to achieve high-order accuracy away from the shock regions and was successfully applied to non-linear hyperbolic equation using various discretization methods (finite volume, continuous and discontinuous finite elements, spectral method) and yielded high-order accuracy on non-uniform meshes and complex geometries [3, 12]. Because of the similarity between Euler equations and the radiation hydrodynamic equations, it is conjectured that the entropy viscosity method may be a good candidate for resolving shocks occurring in radiation-hydrodynamic phenomena.

The 1-D grey radiation-hydrodynamic (GRH) equations are recalled in Eq. (1) and the corresponding variables are defined for clarity purposes:

$$\left\{ \begin{array}{l} \partial_t (\rho) + \partial_x (\rho u) = 0 \\ \partial_t (\rho u) + \partial_x (\rho u^2 + P + \frac{\epsilon}{3}) = 0 \\ \partial_t (\rho E) + \partial_x [u (\rho E + P)] = -\frac{u}{3} \partial_x \epsilon - \sigma_a c (aT^4 - \epsilon) \\ \partial_t \epsilon + \frac{4}{3} \partial_x (u \epsilon) = \frac{u}{3} \partial_x \epsilon + \partial_x \left( \frac{c}{3\sigma_t} \partial_x \epsilon \right) + \sigma_a c (aT^4 - \epsilon) \end{array} \right. , \quad (1)$$

where  $\rho$ ,  $u$ ,  $E$ ,  $\epsilon$ ,  $P$  and  $T$  are the material density, material velocity, material specific total energy, the radiation energy density, material pressure and temperature, respectively. The total and absorption cross-sections,  $\sigma_t$  and  $\sigma_a$  are either constant or density- and temperature-dependent. The variables  $a$  and  $c$  are the Boltzman constant and the speed of light, respectively. Lastly, the symbols  $\partial_t$  and  $\partial_x$  denote the temporal and spatial partial derivatives, respectively. The material temperature and pressure are computed with the Ideal Gas equation of state (IGEOS):

$$\left\{ \begin{array}{l} P = (\gamma - 1) C_v \rho T \\ e = C_v T \end{array} \right. \quad (2)$$

where  $e$  is the specific internal energy and is obtained from the expression  $e = E - 0.5u^2$ . The heat capacity  $C_v$  and the heat ratio coefficient  $\gamma$  are assumed constant.

This paper is organized as follows. In Section 2, the entropy viscosity method is extended to the grey radiation-hydrodynamic equations; details regarding the derivation of the adequate dissipative terms and definitions for the new viscosity coefficients are provided. Spatial and temporal discretization schemes are discussed in Section 3 along with the solution algorithm employed to solve the discretized equations. Numerical results are presented in Section 4 where second-order accuracy of the scheme is demonstrated in both the equilibrium diffusion and streaming limits, using the method of manufactured solutions applied to the GRH equations. Then, several numerical test cases, taken from the published literature, are provided; in these simulations, the Mach number varies from 1.05 to 50 [13]. Conclusions are presented in Section 5.

## 2 The entropy-based viscosity method applied to the 1-D Radiation-Hydrodynamic equations

In this section, we extend the entropy viscosity method [2, 3, 12] to the 1-D radiation-hydrodynamic equations in a staged process. First, the reader is guided through the main steps that lead to the derivation of the dissipative terms, using the entropy minimum principle [14]. Then, a definition for the entropy viscosity coefficient based upon the entropy production is given.

We recall that the entropy viscosity method was developed for hyperbolic system of equations. However, the radiation hydrodynamic equations are not strictly hyperbolic but several numerical techniques are based on the study of their hyperbolic parts [5, 15]. Thus, following the same rationale, the system of equations given in Eq. (1) is made hyperbolic by assuming an infinite opacity (the frozen approximation) and by ignoring the relaxation terms. These two assumptions yield the following system of equations:

$$\left\{ \begin{array}{l} \partial_t (\rho) + \partial_x (\rho u) = 0 \\ \partial_t (\rho u) + \partial_x (\rho u^2 + P + \frac{\epsilon}{3}) = 0 \\ \partial_t (\rho E) + \partial_x [u (\rho E + P)] = -\frac{u}{3} \partial_x \epsilon \\ \partial_t \epsilon + \frac{4}{3} \partial_x (u \epsilon) = \frac{u}{3} \partial_x \epsilon \end{array} \right. \quad (3)$$

The jacobian matrix of the hyperbolic terms can be computed to derive the eigenvalues:

$$\lambda_1 = u - c_m, \lambda_{2,3} = u \text{ and } \lambda_4 = u + c_m, \quad (4)$$

where  $c_m$  is the material speed of sound and is defined as follows:

$$c_m^2 = \underbrace{P_\rho + \frac{P}{\rho^2} P_e}_{c_{Euler}^2} + \frac{4\epsilon}{9\rho} \quad (5)$$

with  $P_x$  the standard shorthand notation for  $\partial_x P$ , and  $c_{Euler}^2$  denotes the definition of the speed of sound when only considering the 1-D Euler equations. The above hyperbolic system of equations can be recast in a conservative form. This allows us to assume the existence of an entropy function  $s$  [16] that depends upon the internal energy  $e$ , the density  $\rho$ , and the radiation energy density  $\epsilon$ . Then, following some algebra given in Appendix A, an equation satisfied by the entropy  $s$  is obtained:

$$\rho \frac{ds}{dt} = \rho (\partial_t s + u \partial_x s) = 0, \quad (6)$$

where  $\frac{d}{dt}$  denotes the total or material derivative. Eq. (6) is often referred as the entropy residual and is used to prove the entropy minimum principle,  $\frac{ds}{dt} \geq 0$  [14].

When adding dissipative terms to each equation of Eq. (3) as required in the entropy viscosity method, the entropy residual equation is modified and some additional terms will appear in the right hand side of Eq. (6). The sign of these extra terms needs to be studied for the entropy minimum principle to hold. As such, the entropy minimum principle is invoked to guide in the derivation of appropriate expressions for each of the dissipative terms. Obtaining the final expression of the dissipative terms is a lengthy process and only the final result along with the key assumptions are stated here. The reader is referred to Appendix A for the details of the

derivation. The system of equations with the dissipative terms is the following:

$$\begin{cases} \partial_t (\rho) + \partial_x (\rho u) = \partial_x (\kappa \partial_x \rho) \\ \partial_t (\rho u) + \partial_x (\rho u^2 + P + \frac{\epsilon}{3}) = \partial_x (\kappa \partial_x \rho u) \\ \partial_t (\rho E) + \partial_x [u (\rho E + P)] + \frac{u}{3} \partial_x \epsilon = \partial_x (\kappa \partial_x (\rho E)) \\ \partial_t \epsilon + \frac{4}{3} \partial_x (u \epsilon) - \frac{u}{3} \partial_x \epsilon = \partial_x (\kappa \partial_x \epsilon) \end{cases}, \quad (7)$$

where  $\kappa$  is a locally defined positive viscosity coefficient. It was assumed that the following conditions hold:

$$\begin{cases} P \frac{\partial s}{\partial e} + \rho^2 \frac{\partial s}{\partial \rho} + \frac{4}{3} \rho \epsilon \frac{\partial s}{\partial \epsilon} = 0 \\ s(\rho, e, \epsilon) = \hat{s}(\rho, e) + \frac{\rho_0}{\rho} \tilde{s}(\epsilon) \end{cases} \quad (8)$$

where  $-\tilde{s}$  is convex with respect to the radiation energy density  $\epsilon$  and  $-\hat{s}$  is convex with respect to the internal energy  $e$  and the specific volume  $\frac{1}{\rho}$ . The constant  $\rho_0$  is of order one and appears only for dimensionality purposes.

Once the dissipative terms are obtained, it remains to define the local viscosity coefficient  $\kappa(x, t)$ .

We require the following to hold in the prescription for  $\kappa$ :

- An upper bound on  $\kappa$  is required for stability of the numerical scheme: when considering explicit temporal integrators, the maximum value of the viscosity coefficient is related to the Courant-Friedrichs-Lewy number (CFL). This upper bound is defined by analogy to the standard upwind (Godunov) scheme that is known to efficiently smooth out oscillations (but is only first-order accurate). With implicit temporal integrators, the same reasoning is used even if the CFL number is, in theory, no longer required. This upper bound will be referred to as the *first-order viscosity*, denoted by  $\kappa_{max}(x, t)$ .
- Since the entropy residual is a measure of the entropy production that occurs in shock regions, it is natural to define a viscosity coefficient proportional to the entropy residual. This will enable shock detection and tracking and will also provide a measure of the viscosity required to stabilize the scheme. This viscosity coefficient is referred to as the *entropy viscosity coefficient* or *second-order viscosity coefficient* and is denoted by  $\kappa_e(x, t)$ .
- The viscosity coefficient that is actually used in the dissipative terms,  $\kappa$ , of Eq. (3) is defined as follows:  $\kappa(x, t) = \min(\kappa_e(x, t), \kappa_{max}(x, t))$ . With such a definition, the viscosity added

to the system of equation will saturate to the first order viscosity in the shock regions.

Elsewhere, the entropy production and the viscosity coefficient  $\kappa$  are expected to be small.

Next, we define the first- and second-order viscosity coefficients. Following the work of Zingan et al. [12], the first-order viscosity definition is based on the local largest eigenvalue that is known to be  $|u| + c_m$  in 1-D:

$$\kappa_{max} = \frac{h}{2} (|u| + c_m) \quad (9)$$

where  $h$  is the local grid size. This definition is derived based on the upwind scheme and a simple derivation can be found in [2] in the case of a scalar hyperbolic equation. Through the definition of the speed of sound  $c_m$ , both the material and radiation properties are accounted for in the definition of the first-order viscosity coefficient.

The definition of the second order viscosity coefficient,  $\kappa_e(x, t)$ , is based upon the entropy residual (Eq. (6)) that can be recast as a function of pressure  $P$ , density  $\rho$  and radiation energy density  $\epsilon$ :

$$\tilde{D}_e(x, t) = \frac{s_e}{P_e} \underbrace{\left( \frac{dP}{dt} - c_{Euler}^2 \frac{d\rho}{dt} \right)}_{\hat{D}_e(x, t)} \quad (10)$$

The term  $s_e$  is the inverse of the material temperature (Appendix A) and  $P_e$  is computed from the IGEOS. These two terms are positive so that the sign of the entropy residual  $\tilde{D}_e(x, t)$  can be determined by simply inspecting the terms inside the parentheses, denoted by  $\hat{D}_e(x, t)$ . Such an expression is easier to compute than the one given in Eq. (6) which required an analytical expression for the entropy function. In addition to the entropy residual, inter-element jumps in the pressure and density gradients,  $J$ , are also accounted for. The objective is to be able to also detect discontinuities that are not shocks, such as contact waves (there is no entropy production in a contact wave), in order to stabilize them as well.

Thus, the entropy viscosity coefficient  $\kappa_e(x, t)$  is set to be proportional to  $\hat{D}_e(x, t)$  and  $J$  with the following form:

$$\kappa_e(x, t) = h^2 \frac{\max(|\hat{D}_e(x, t)|, J)}{n_P} \quad (11)$$

where  $J = \max_i(J(x_i, t))$ , and  $J(x_i, t)$  is the jump of a given quantity at cell interface  $x_i$ , and  $n_P$  is a normalization function (of the same unit as pressure) that has to be chosen so that



the viscosity coefficient  $\kappa$  is of unit  $m^2/s$ . It is chosen to take the following definition for the normalization function:  $n_P = \rho c_m^2$ . Thus, the final definition for the viscosity coefficient  $\kappa$  is the following:

$$\kappa_e(x, t) = h^2 \frac{\max(|\hat{D}_e(x, t)|, J)}{\rho c_m^2} \quad (12)$$

The jump  $J$  in the definition of  $\kappa(x, t)$  is piecewise-constant. Its definition is discretization-dependent and defined as follows for Continuous Galerkin FEM:

$$\begin{cases} J_P(x_i, t) = |u|[[\partial_x P]] \\ J_\rho(x_i, t) = c_m^2 |u|[[\partial_x \rho]] \\ J(x_i, t) = \max(J_\rho(x_i, t), J_P(x_i, t)) \end{cases} \quad (13)$$

The symbol  $[[\cdot]]$  denotes the jump at the interface.

The entropy viscosity method is now well defined for the hyperbolic system given in Eq. (3) and will be used to solve for the grey radiation-hydrodynamic equations given in Eq. (1). However, one may question how the relaxation source terms,  $\sigma_a c(aT^4 - \epsilon)$  and the physical diffusion term,  $\partial_x(D\partial_x \epsilon)$ , may affect the entropy viscosity method. When applying the entropy viscosity method, the radiation energy density equation will now contain a diffusive term and a numerical dissipative term with a vanishing viscosity coefficient  $\kappa$ . As long as the diffusive coefficient  $D = \frac{c}{3\sigma_t}$  is larger than the viscosity coefficient  $\kappa$ , the numerical dissipative term should not be required. A way to ensure consistency and prevent the formation of oscillations in the frozen limit is to merge the two second-order derivative terms into one as follows:

$$\partial_x \left( \frac{c}{3\sigma_t} \partial_x \epsilon \right) + \partial_x (\kappa \partial_x \epsilon) = \partial_x \left[ \max \left( \frac{c}{3\sigma_t}, \kappa \right) \partial_x \epsilon \right] \quad (14)$$

Thus, as long as the artificial viscosity coefficient  $\kappa$  is locally smaller than the physical diffusive coefficient  $D = \frac{c}{3\sigma_t}$ , no artificial viscosity is required to ensure stability of the numerical scheme. As the diffusive coefficient  $D$  goes to zero, shocks can form in the radiation energy density profile and will require a certain amount of viscosity in order to prevent oscillations from appearing. The effect of the relaxation source terms onto the entropy viscosity method can become problematic in the equilibrium diffusion limit ( $\sigma_a c \rightarrow \infty$ ): the relaxation source terms behave as

dissipative terms and make the system parabolic [4]. In [19], a study on the impact of various artificial viscosity methods onto hyperbolic systems with relaxation terms was carried out. It was shown that high-order viscosity coefficients are better suitable since they do not alter the physical solution as much as first-order viscosity such as the upwind scheme. A manufactured method solution is designed in Section 4.1 to test the convergence of the numerical solution in the equilibrium-diffusion limit.

**Remark 2.1.** *The normalization factor has to be larger than  $h$  in order to conserve the high-order accuracy.*

**Remark 2.2.** *The reader will notice that, except for the definition of the jumps, the whole method is independent of the spatial discretization employed. The technique could be used with discontinuous Galerkin finite element or finite volume methods. In both cases, an adequate definition of the jump terms can be found in [12].*

### 3 Numerical scheme and solution technique

The 1-D radiation-hydrodynamic equations Eq. (1) are discretized with the continuous Galerkin finite element method (CGFEM) under the Moose framework [21].

To obtain a weak form, the following generic form of Eq. (1) is considered:

$$\partial_t U + \partial_x F(U) = S + \partial_x H(U) \quad (15)$$

where  $U$  is the vector solution,  $F$  is a conservative vector flux,  $S$  is a vector containing the relaxation source terms and non-conservative terms, and  $H$  is the artificial viscosity dissipative

flux:

$$\begin{aligned}
U &= \begin{bmatrix} \rho \\ \rho u \\ \rho E \\ \epsilon \end{bmatrix}, \quad F(U) = \begin{bmatrix} \rho u \\ \rho u^2 + P + \frac{\epsilon}{3} \\ u(\rho E + P) \\ \frac{4}{3}u\epsilon \end{bmatrix}, \\
S &= \begin{bmatrix} 0 \\ 0 \\ -\frac{u}{3}\partial_x \epsilon - \sigma_a c (aT^4 - \epsilon) \\ \frac{u}{3}\partial_x \epsilon + \partial_x \left( \frac{c}{3\sigma_t} \partial_x \epsilon \right) + \sigma_a c (aT^4 - \epsilon) \end{bmatrix} \quad \text{and} \\
H(U) &= \begin{bmatrix} \kappa \partial_x \rho \\ \kappa \partial_x (\rho u) \\ \kappa \partial_x (\rho e) + \frac{u^2}{2} \kappa \partial_x \rho + \rho u \mu \partial_x u \end{bmatrix}
\end{aligned}$$

In order to apply the continuous finite element method, Eq. (15) is multiplied by a smooth test function  $\phi$ , integrated by parts over each finite element  $e$  of the discrete mesh  $\Omega$  bounded by  $\partial\Omega$ , to obtain a weak solution:

$$\begin{aligned}
&\sum_e \int_e \partial_t U \phi - \sum_e \int_e F(U) \partial_x \phi + \int_{\partial\Omega} F(U) \mathbf{n} \phi - \\
&\sum_e \int_e S \phi + \sum_e \int_e H(U) \partial_x \phi - \int_{\partial\Omega} H(U) \mathbf{n} \phi = 0
\end{aligned} \tag{16}$$

where  $\mathbf{n}$  is the outward normal vector to the boundary of the computational domain.

The integrals over the elements  $e$  are evaluated using either trapezoidal or Gauss quadrature rules. The time-dependent term will be evaluated using the implicit second-order temporal integrator BDF2 [4]. Only linear polynomials test functions are considered in this paper.

The integral on  $\partial\Omega$  requires computation of the flux  $F(U)$  and  $H(U)$  on the boundary. It is chosen to treat the boundary for each physics component independently and details will be given in Section 4.2 on how to compute  $F(U)$ . The viscosity dissipative flux  $H(U)$  is zero on the boundaries since the viscosity coefficients  $\kappa$  and  $\mu$  are set to zero [2, 3, 12].

The Jacobian-Free Newton Krylov method [22] is employed to solve the system for each time

step: Newton’s and Krylov methods are used for the outer nonlinear and inner linear solves, respectively. The full Jacobian matrix is used as a preconditioner and is computed by finite difference: this options is available in Moose and remains efficient for 1-D simulations.

**Remark 3.1.** *The entropy residual expression is not integrated over the cell volume as it is usually done in the Galerkin finite element method. The variable values and their gradients are available at quadrature points and at different times, and, thus, can be used to evaluate the entropy residual.*

## 4 Numerical results

In this section, numerical results of the dimensional 1D grey radiation-hydrodynamic equations are presented using the entropy viscosity method. First, second-order accuracy of the method is demonstrated using the method of manufactured solution (MMS). Then, some typical test cases for radiation-hydrodynamic equations are given.

### 4.1 Space/time accuracy

The same manufactured solution as in [18] is used in order to test both the diffusive and streaming limit solutions in a slab of thickness  $L = 2\pi$  cm. The manufactured solutions are composed of trigonometric functions. Periodic boundary conditions are used for all of the variables. The  $L_2$  norm errors between the numerical and exact solutions are computed for density, momentum, total material energy, and radiation energy density. For each new simulation, the time step is divided by two and the number of spatial degrees of freedom is doubled. With such settings, the error is expected to decrease by a factor 4 if second-order convergence is achieved.

The first manufactured solution is designed to test the equilibrium-diffusion limit. In that case, the radiation energy is in equilibrium with the material temperature and the opacity is large which means that the radiation mean-free path is not resolved. The following exact solution was

used:

$$\begin{cases} \rho = \sin(x - t) + 2 \\ u = \cos(x - t) + 2 \\ T = \frac{0.5\gamma(\cos(x-t)+2)}{\sin(x-t)+2} \\ \epsilon = aT^4 \end{cases} \quad (17)$$

The cross sections  $\sigma_a$  and  $\sigma_t$  are assumed constant and set to the same value  $1000 \text{ cm}^{-1}$ . The simulation is run until  $t = 3 \text{ sh}$ . The  $L_2$  error norm along with the ratio between consecutive simulations are given in Table 1 for the equilibrium diffusion limit:

Table 1: Convergence study for the first manufactured solution.

DoF	$\Delta t$ (sh)	$\rho$	ratio	$\rho E$	ratio
20	$10^{-1}$	0.590766	NA	1.333774	NA
40	$5 \cdot 10^{-1}$	0.290626	2.03	0.478819	2.79
80	$2.5 \cdot 10^{-2}$	0.0959801	3.021	0.154119	3.11
160	$1.25 \cdot 10^{-2}$	0.02593738	3.70	0.0405175	3.80
320	$6.25 \cdot 10^{-3}$	$6.471444 \cdot 10^{-3}$	4.00	$9.90446 \cdot 10^{-3}$	4.09
640	$3.125 \cdot 10^{-3}$	$1.584158 \cdot 10^{-3}$	4.01	$2.44727 \cdot 10^{-3}$	4.04
DoF	$\Delta t$ (sh)	$\epsilon$	ratio	$\rho u$	ratio
20	$10^{-1}$	0.00650085	NA	0.910998	NA
40	$5 \cdot 10^{-1}$	0.00124983	5.20	0.4090946	2.23
80	$2.5 \cdot 10^{-2}$	0.000262797	4.76	0.125943	3.25
160	$1.25 \cdot 10^{-2}$	$6.17726 \cdot 10^{-5}$	4.25	$3.381042 \cdot 10^{-3}$	3.72
320	$6.25 \cdot 10^{-3}$	$1.509184 \cdot 10^{-5}$	4.09	$8.373657 \cdot 10^{-3}$	4.04
640	$3.125 \cdot 10^{-3}$	$3.72548 \cdot 10^{-6}$	4.05	$2.070538 \cdot 10^{-3}$	4.04

The second manufactured solution is used to test the streaming limit: the radiation streaming dominates the absorption/re-emission term and evolves at a fast time scale. The exact solutions

are the following:

$$\begin{cases} \rho = \sin(x - t) + 2 \\ u = (\sin(x - t) + 2)^{-1} \\ T = 0.5\gamma \\ \epsilon = \sin(x - 1000t) + 2 \end{cases} \quad (18)$$

For this manufactured solution, the cross sections are still assumed constant and set to the same value  $1 \text{ cm}^{-1}$ . The final time is  $t_{final} = 3. \text{ sh}$ . Once again, the  $L_2$  error is given in Table 2 for the density, momentum, material total energy and radiation energy density.

Table 2: Convergence study for the second manufactured solution.

DoF	$\Delta t$ (sh)	$\rho$	ratio	$\rho E$	ratio
20	$10^{-1}$	$1.4373 \cdot 10^{-2}$	NA	$5.88521 \cdot 10^{-1}$	NA
40	$5 \cdot 10^{-2}$	$3.760208 \cdot 10^{-3}$	3.82	$1.4244 \cdot 10^{-1}$	4.13
80	$2.5 \cdot 10^{-2}$	$9.91724 \cdot 10^{-4}$	3.79	$3.2047 \cdot 10^{-2}$	4.44
160	$1.25 \cdot 10^{-2}$	$2.4455 \cdot 10^{-4}$	4.06	$7.4886 \cdot 10^{-3}$	4.28
320	$6.25 \cdot 10^{-3}$	$6.280715 \cdot 10^{-5}$	3.89	$1.82327 \cdot 10^{-3}$	4.11
640	$3.125 \cdot 10^{-3}$	$1.57920 \cdot 10^{-5}$	3.98	$4.50463 \cdot 10^{-4}$	4.05
1280	$1.5625 \cdot 10^{-4}$	$3.96096 \cdot 10^{-6}$	3.99	$1.12061 \cdot 10^{-4}$	4.02
DoF	$\Delta t$ (sh)	$\epsilon$	ratio	$\rho u$	ratio
20	$10^{-1}$	$3.82001 \cdot 10^{-1}$	NA	$2.354671 \cdot 10^{-3}$	NA
40	$5 \cdot 10^{-2}$	$1.21500 \cdot 10^{-1}$	3.14	$6.138814 \cdot 10^{-4}$	3.84
80	$2.5 \cdot 10^{-2}$	$3.27966 \cdot 10^{-2}$	3.70	$1.74974 \cdot 10^{-4}$	3.51
160	$1.25 \cdot 10^{-2}$	$8.38153 \cdot 10^{-3}$	3.91	$3.61297 \cdot 10^{-5}$	4.84
320	$6.25 \cdot 10^{-3}$	$2.10925 \cdot 10^{-3}$	3.97	$9.03866 \cdot 10^{-6}$	3.99
640	$3.125 \cdot 10^{-3}$	$5.28472 \cdot 10^{-4}$	3.99	$2.25649 \cdot 10^{-6}$	4.01
1280	$1.5625 \cdot 10^{-4}$	$1.322268 \cdot 10^{-4}$	3.99	$5.69984 \cdot 10^{-7}$	3.95

For both manufactured solutions the error is divided by four as the time step  $\Delta t$  and the spatial mesh  $\Delta x$  are reduced by a factor two. Thus, we conclude that the entropy viscosity

method allows the scheme to be second-order accurate when the numerical solution is smooth.

## 4.2 Radiation shock simulations

The purpose of this section is to show that the entropy-based viscosity method (Section 2) can resolve shocks occurring in radiation-hydrodynamic simulations. Multiple test cases are considered with Mach numbers of 1.05, 1.2, 2, 5 and 50 [13]. All of the simulations are run with 500 cells and with a Courant-Friedrichs-Lewy (CFL number) of 10 until steady-state (even if the scheme employed here is fully implicit, a CFL number can still be computed and is a good reference for comparison against semi-implicit or fully explicit codes). Linear Lagrange polynomials and the second-order temporal integrator BDF2 are once again used. For clarity, the initial conditions of each test case will be given in a table and plots of the density  $\rho(x)$ , the radiation  $\theta(x)$  and material  $T(x)$  temperatures at steady-state will be shown as well as the local viscosity coefficients  $\kappa(x)$  and  $\kappa_{max}(x)$ . The computational domain consists of a 1D slab of thickness  $L$ . The initial step is located at  $x_0$  and will be specified for all of test cases. For all of the test cases presented in this paper the cross sections  $\sigma_a$  and  $\sigma_t$  are assumed constant and set to  $853.144 \text{ cm}^{-1}$  and  $390.711 \text{ cm}^{-1}$ , respectively, if not otherwise specified. The heat capacity at constant specific volume is set to  $C_v = 0.12348 \text{ jerks}/(g - keV)$ .

For the Mach 2 simulation, results will be shown when employing the first-order viscosity ( $\kappa(x, t) = \kappa_{max}(x, t)$ ). The rationale for this is to show the benefits of using a high-order viscosity coefficients.

The inlet and outlet BCs are given next. The Euler equations and radiation equation are considered independently since the later one is parabolic. At the inlet, the flow is supersonic and, therefore, no physical information exits the system. Thus, Dirichlet boundary condition can be used. At the outlet, the flow become subsonic which requires a particular treatment. Following the work from [23], a static boundary condition is implemented. Only the back pressure is provided and the other variables are computed using the characteristic equations. For the radiation equation, vacuum boundary conditions are used at both inlet and outlet.

#### 4.2.1 An equilibrium diffusion test:

For this test, the inlet Mach number is set to 1.05. The radiation field and material are in equilibrium. The initial conditions are given in Table 3. The computational domain is of size

Table 3: Initial conditions for Mach 1.05.

	left	right
$\rho$ ( $g/cm^3$ )	1.	1.0749588
$u$ ( $cm/sh$ )	0.1228902	0.1144127
$T$ ( $keV$ )	0.1	0.1049454
$\epsilon$ ( $jerk s/cm^3$ )	$1.372 \cdot 10^{-6}$	$1.6642117 \cdot 10^{-6}$

$L = 0.030$  cm and the initial step was located in the middle at  $x_0 = 0.015$  cm. The numerical solutions at steady-state are given in Fig. 1, Fig. 2 and Fig. 3. The solution was run with  $CFL = 10$ .

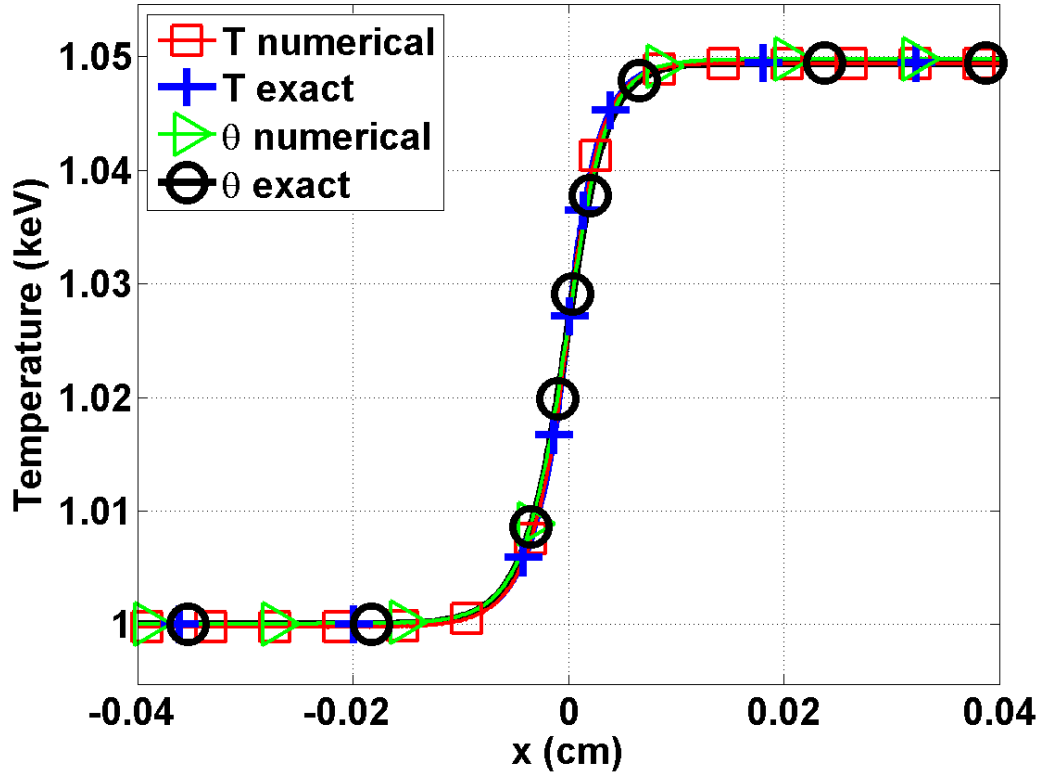


Figure 1: Material and radiation temperature profiles at steady-state for Mach 1.05 test.



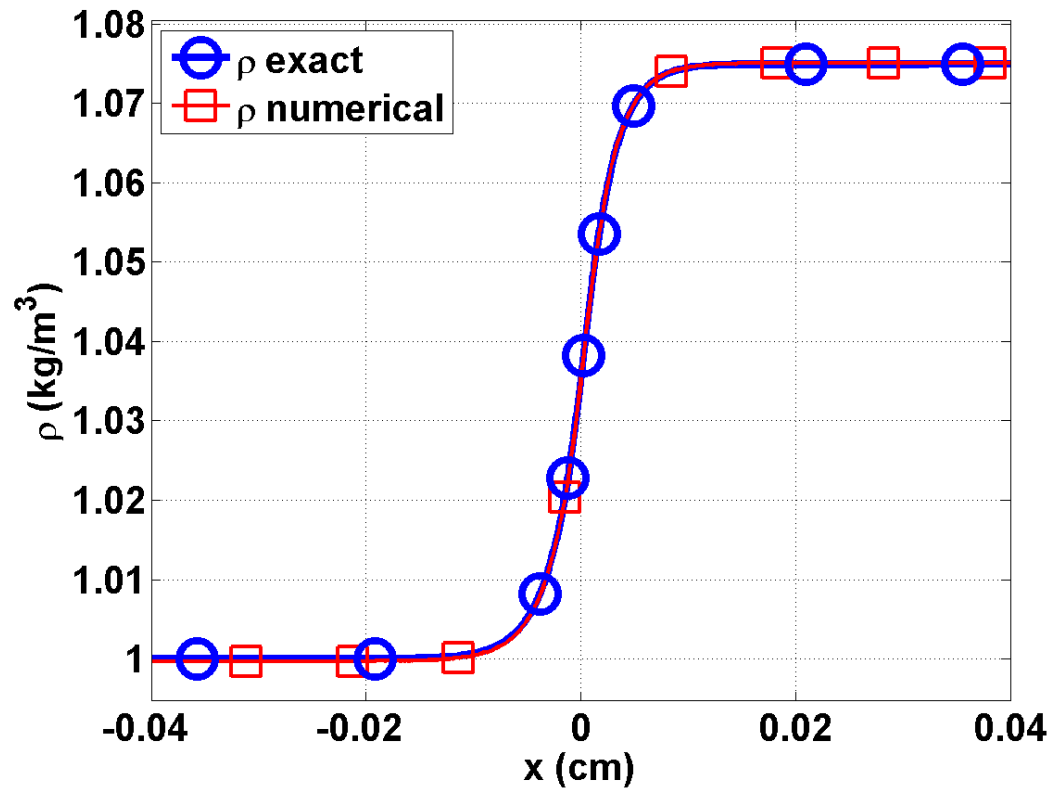


Figure 2: Material density profile at steady-state for Mach 1.05 test.

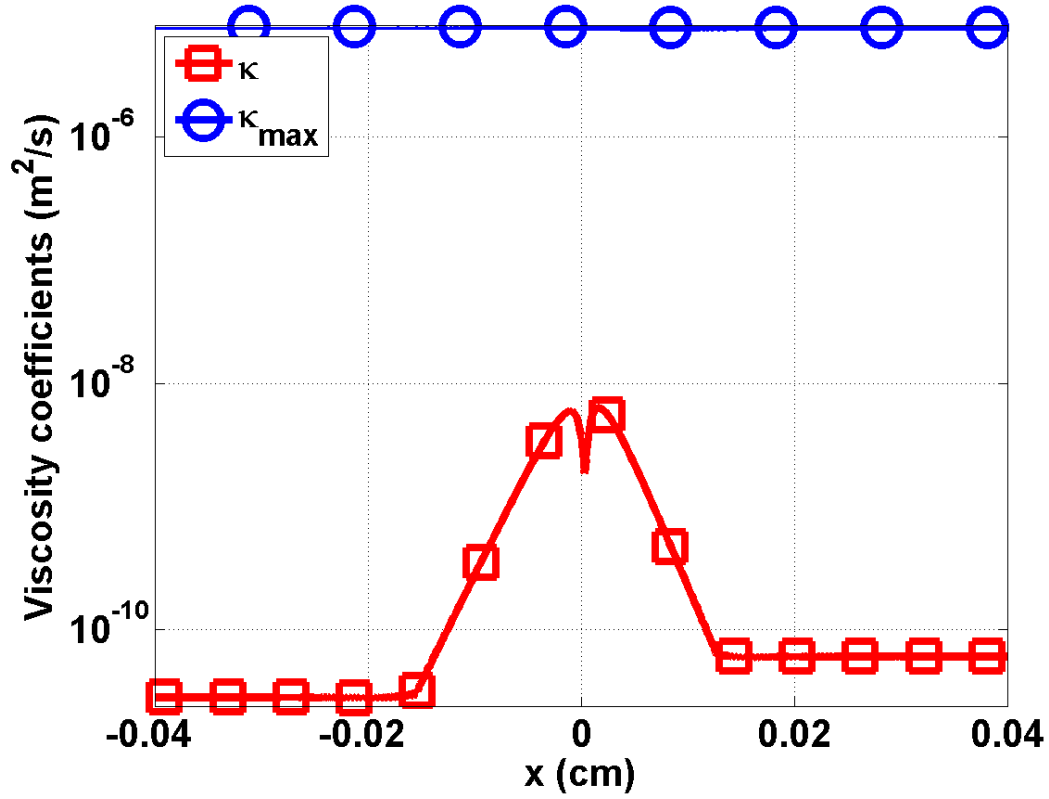


Figure 3: First-order viscosity  $\kappa_{max}$  and second-order viscosity  $\kappa$  profiles at steady-state for Mach 1.05 test (logarithm scale).

The energy transfer between the material and radiation fields is not large enough to form a shock in the material. Thus, all of the material variables are smooth (Fig. 1 and Fig. 2) as well as the radiation temperature  $\theta$ . Because of the smoothness of the solution, the viscosity coefficient  $\kappa$  is three order of magnitude smaller than the first-order viscosity coefficient  $\kappa_{max}$  (Fig. 3): in other words, the viscosity added to the system is large enough to stabilize it, but sufficiently small so that the physical solution is not disturbed.

#### 4.2.2 A 1.2 Mach hydrodynamic shock:

In this test, the material experiences a shock and the radiation energy density remains smooth. The initial conditions, corresponding to a Mach number of 1.2 at the inlet, are as follows:

Table 4: Initial conditions for Mach 1.2.

	left	right
$\rho$ ( $g/cm^3$ )	1.	1.0749588
$u$ ( $cm/sh$ )	0.1405588	0.1083456
$T$ ( $keV$ )	0.1	0.1194751
$\epsilon$ ( $jerk/cm^3$ )	$1.372 \cdot 10^{-6}$	$2.7955320 \cdot 10^{-6}$

The slab thickness is set to  $L = 310^{-2}$  cm and the initial step was located at  $x_0 = 0$  cm.

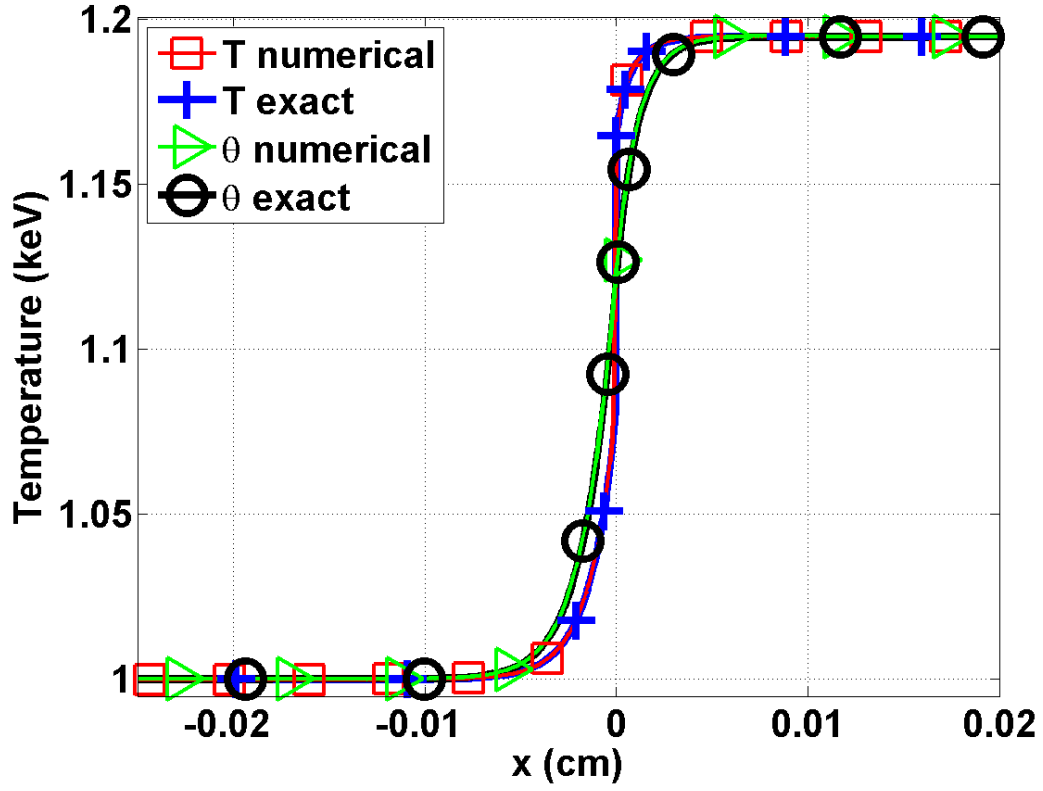


Figure 4: Material and radiation temperature profiles at steady-state for Mach 1.2 test.

The radiation and material temperatures have two different behaviors (Fig. 4): the later experiences an embedded hydrodynamic shock, whereas the radiation temperature is smooth

because of the diffusion term. The material temperature profile does not show any pre- and post-shock oscillations. In Fig. 5, the material density profile has a shock as well. The viscosity coefficient (Fig. 6) is peaked in the shock as expected but does not saturate to the first-order viscosity. It is conjectured that the diffusion term in the radiation equation brings extra stability to the system.

Overall, the numerical solution behaves as expected in the shock and the entropy-based viscosity method seems to efficiently smooth oscillations out.

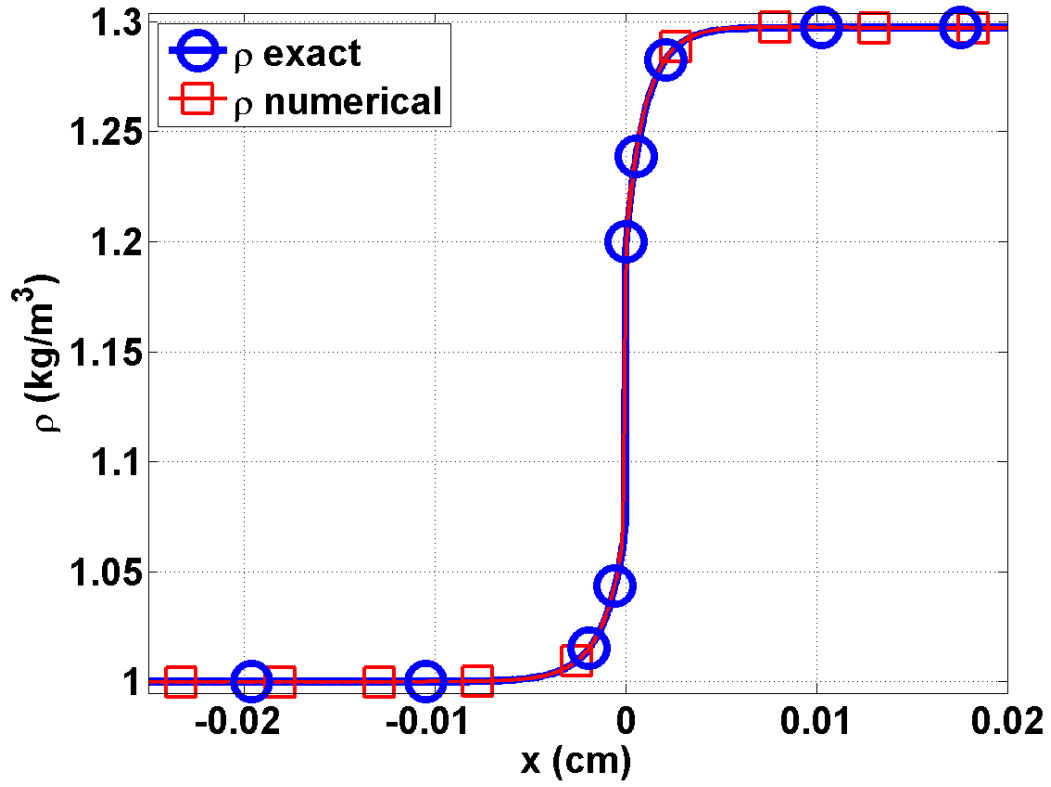


Figure 5: Material density profile at steady-state for Mach 1.2 test.

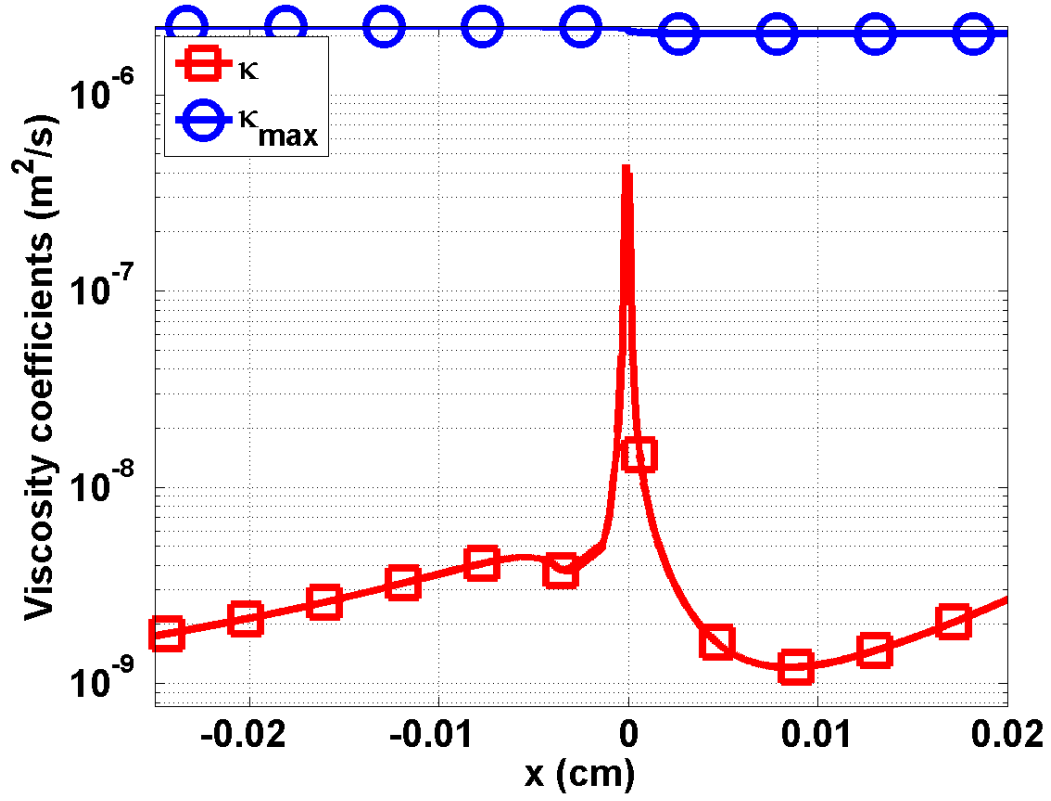


Figure 6: First-order viscosity  $\kappa_{\max}$  and second-order viscosity  $\kappa$  profiles at steady-state for Mach 1.2 test (logarithm scale).

#### 4.2.3 A Mach 2 shock:

The Mach 2 shock test has two features: a hydrodynamic shock and a Zeldovich spike, which make it interesting for testing the robustness of the entropy-based viscosity method. The initial conditions are specified in Table 5 for a slab of length  $L = 310^{-2}$  cm with  $x_0 = 0$  cm.

Table 5: Initial conditions for Mach 2.

	left	right
$\rho$ ( $g/cm^3$ )	1.	1.0749588
$u$ ( $cm/sh$ )	0.1405588	0.1083456
$T$ ( $keV$ )	0.1	0.1194751
$\epsilon$ ( $jerk s/cm^3$ )	$1.372 \cdot 10^{-6}$	$2.7955320 \cdot 10^{-6}$

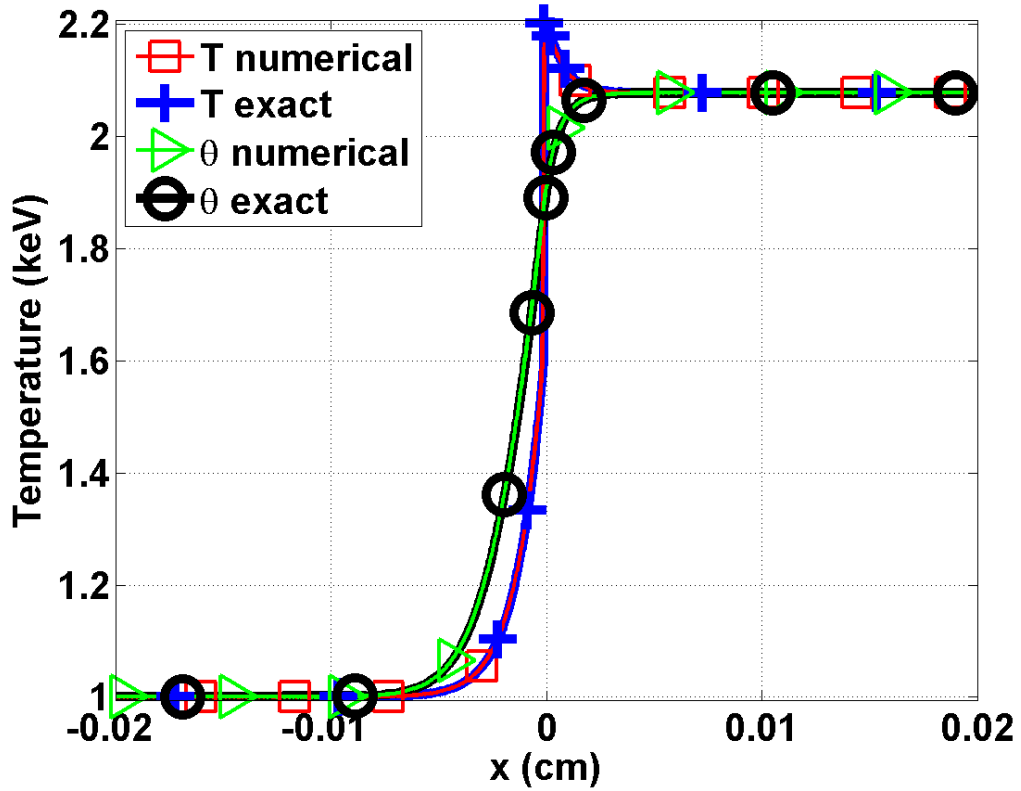


Figure 7: Material and radiation temperature profiles at steady-state for Mach 2 test.

Once again, the radiation temperature profile is smooth and the material temperature experiences an embedded hydrodynamic shock and a peak as shown in Fig. 7. In Fig. 8, the shock is well resolved. The viscosity coefficient profile is given in Fig. 9 and is peaked, once again, in the shock region.

For academic purpose, the same simulation was run with the first-order viscosity i.e.,  $\kappa$  was set equal to  $\kappa_{max}$  in the whole domain, in order to see the advantage of using a second-order viscosity coefficient. The results are given in Fig. 10 for the material density and temperature. Numerical solutions with first- and second-order viscosity coefficients are graphed. The radiation temperature profile (not shown here) is not affected by the first-order viscosity and curves are confounded: this is expected because of the way the artificial viscosity term is treated in the radiation equation (Section 2). However, on the same figure, the shock and peak in the material temperature profile are smoothed out: the shock is not as sharp and the peak amplitude is reduced because of the large amount of viscosity added to the system. This test shows the benefits of using a high-order viscosity coefficient in order to avoid over dissipation.

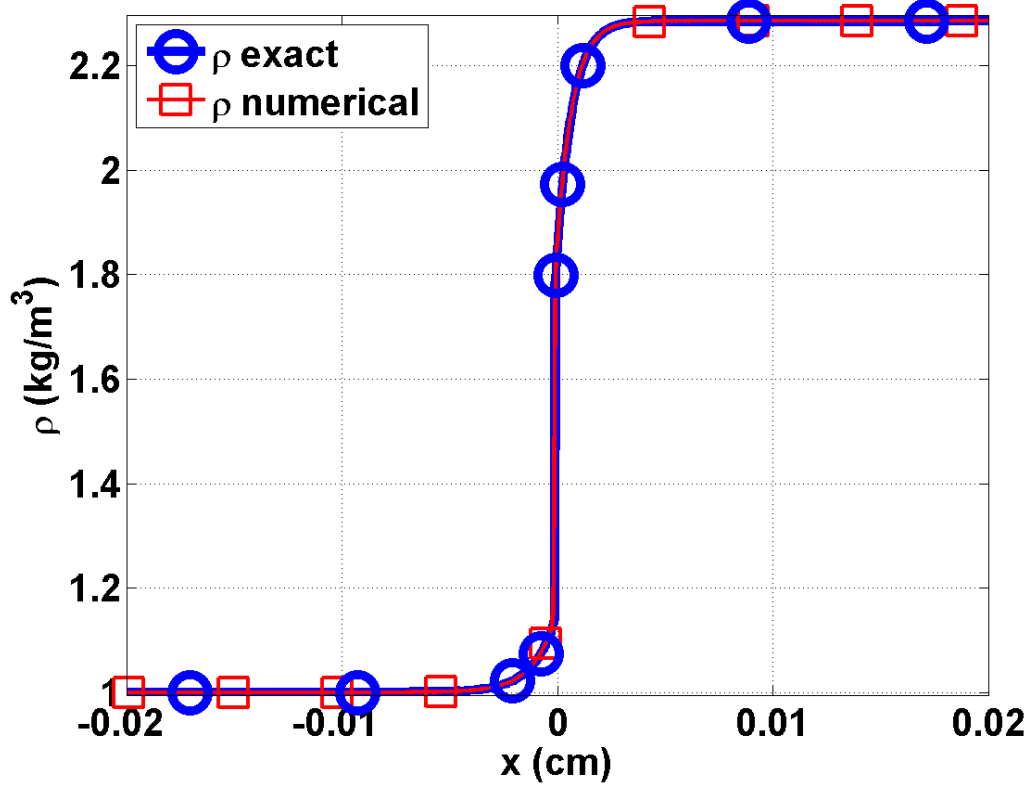


Figure 8: Material density profile at steady-state for Mach 2 test.

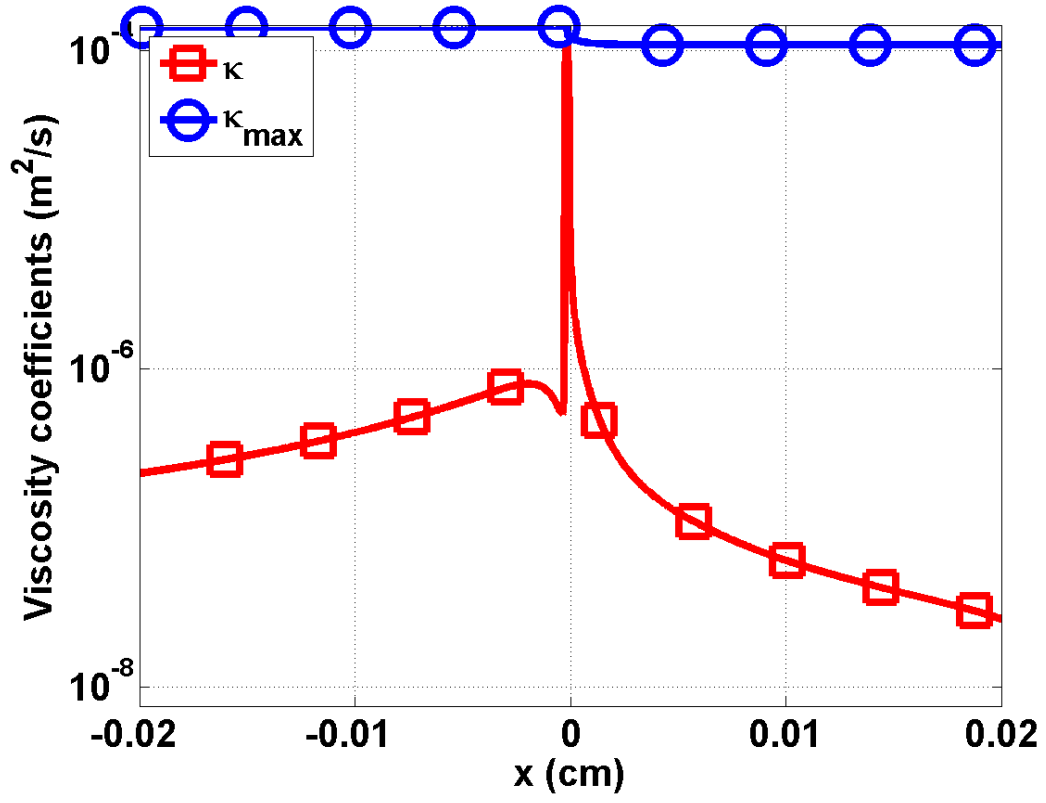


Figure 9: First-order viscosity  $\kappa_{\text{max}}$  and second-order viscosity  $\kappa$  profiles at steady-state for Mach 2 test.



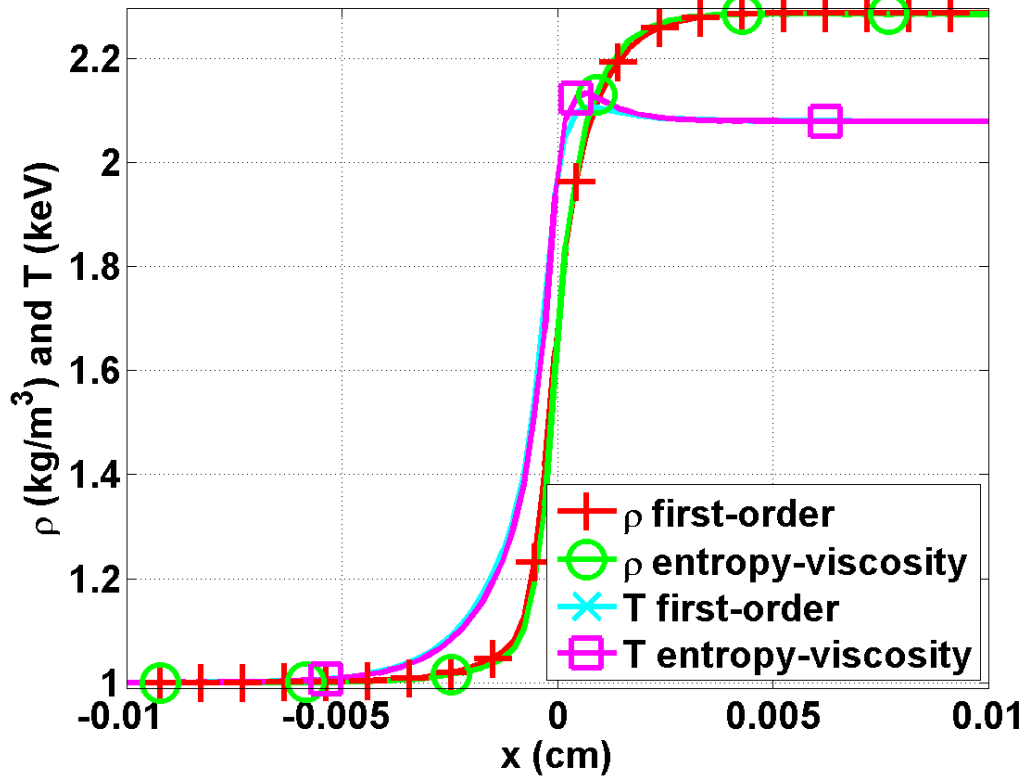


Figure 10: Comparison between the material density and temperature profiles run with the high-order and first-order viscosity coefficients.

#### 4.2.4 Mach 5 shock:

A Mach 5 test is run with the initial conditions of Table 6 on a computational domain of length  $L = 0.05 \text{ cm}$  and with a  $CFL$  of 10 ( $x_0 = 0 \text{ cm}$ ). Steady-state results are shown in Fig. 11, Fig. 13 and Fig. 14 for the material and radiation temperatures, the density and the viscosity coefficients, respectively.

Table 6: Initial conditions for Mach 5.

	left	right
$\rho$ ( $g/cm^3$ )	1.	1.0749588
$u$ ( $cm/sh$ )	0.1405588	0.1083456
$T$ ( $keV$ )	0.1	0.1194751
$\epsilon$ ( $jerk/cm^3$ )	$1.372 \cdot 10^{-6}$	$2.7955320 \cdot 10^{-6}$

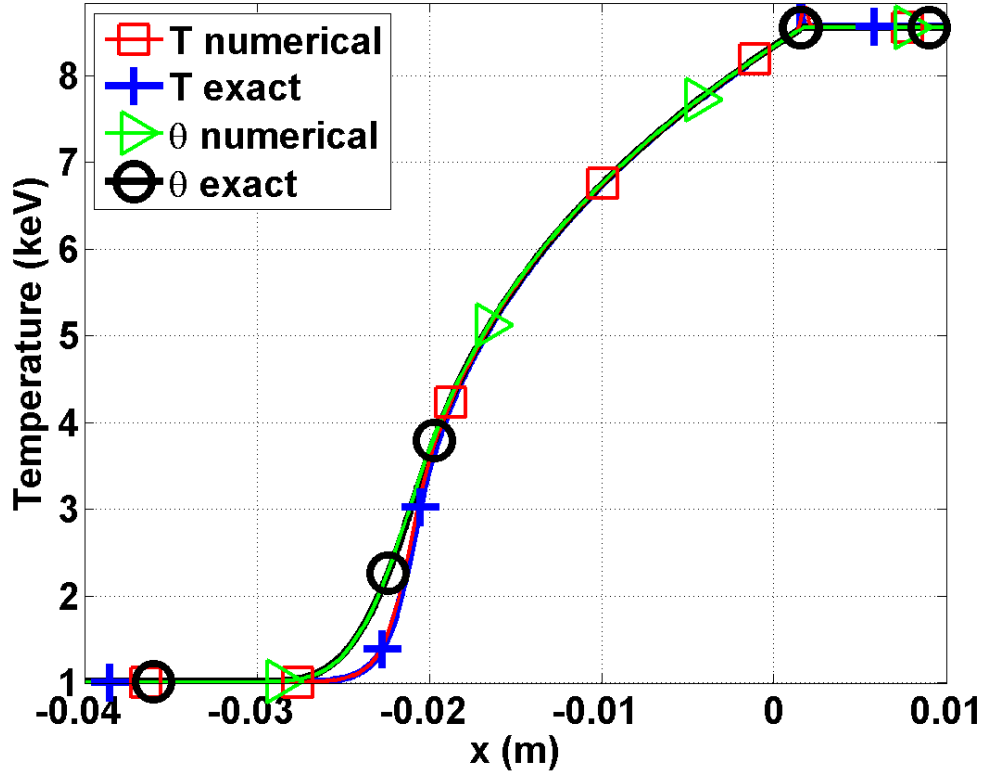


Figure 11: Material and radiation temperature profiles at steady-state for Mach 5 test. Zoom at the location go the peak using different mesh resolutions.

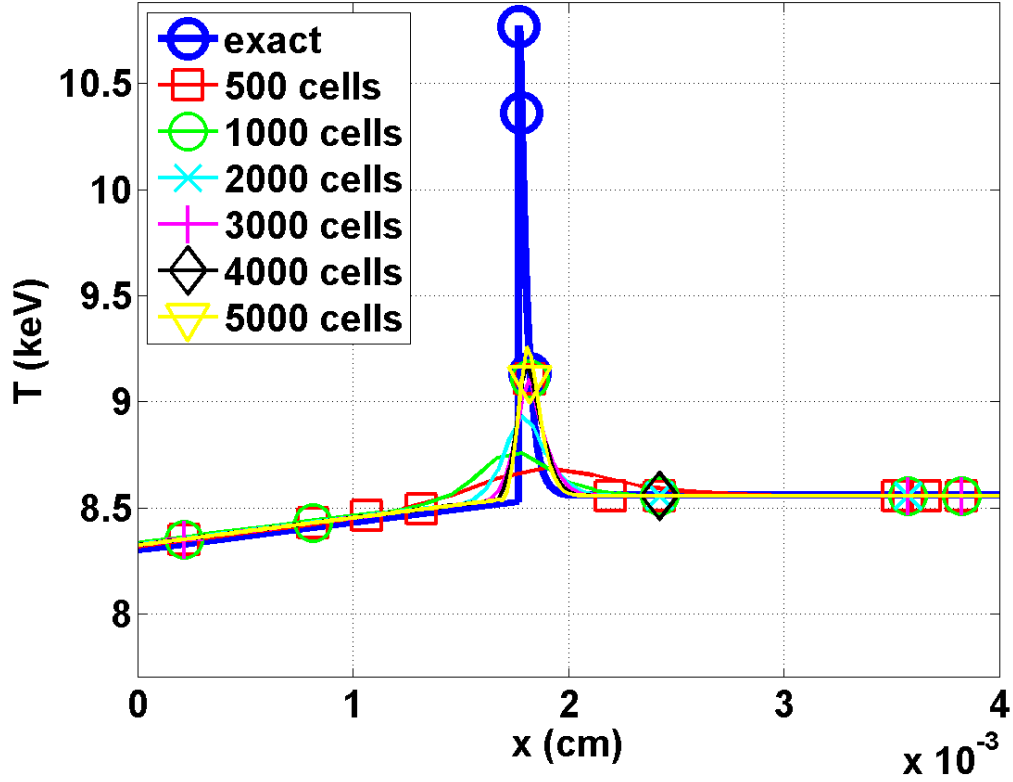


Figure 12: Material and radiation temperature profiles at steady-state for Mach 5 test. Zoom at the location go the peak using different mesh resolutions.

In Fig. 11, the radiation temperature profile is smooth. The material temperature no longer displayed an embedded hydrodynamic shock but has a Zeldovich spike. The mesh with 500 elements is not fine enough to correctly resolve the Zeldovich spike. In Fig. 12, the Zeldovich spike region is plotted for different mesh from 500 to 5000 elements: the peak is better resolved when using large numbers of elements and its position seems to be independent of the mesh size when correctly refined. The density profile, Fig. 13, has a shock located at the same position as the Zeldovich spike of the material temperature profile. The viscosity coefficient  $\kappa$  is also peaked in the shock region, as expected. The material and radiation variables do not present any oscillations.

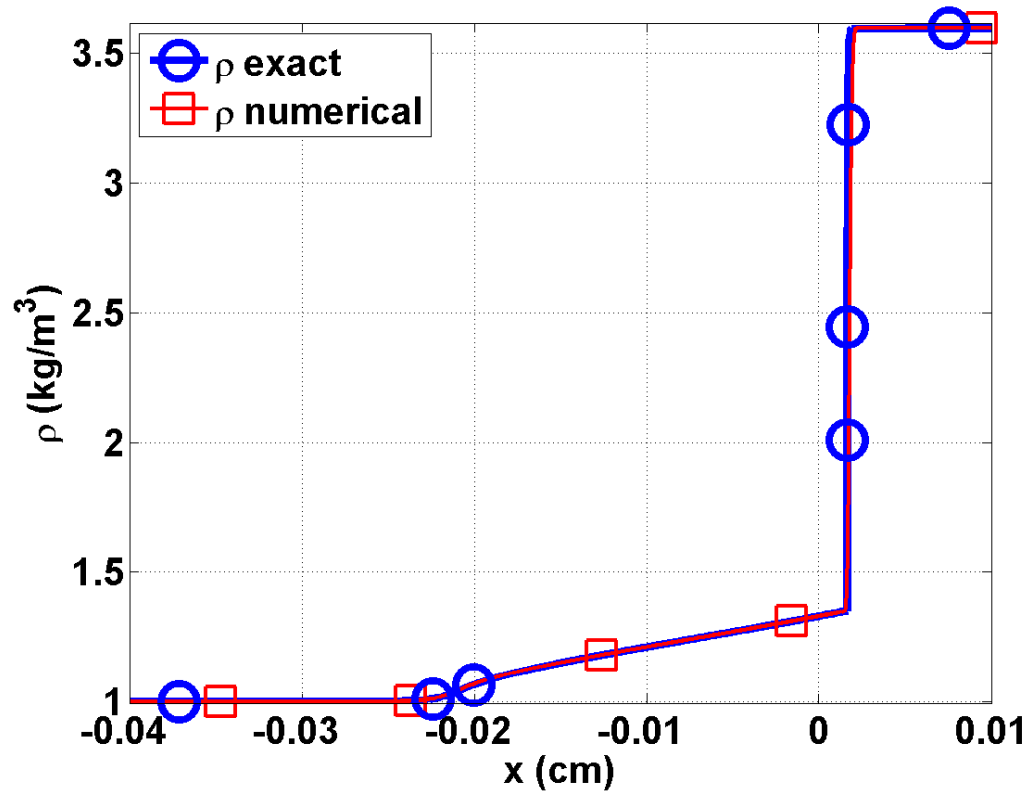


Figure 13: Material density profile at steady-state for Mach 5 test.

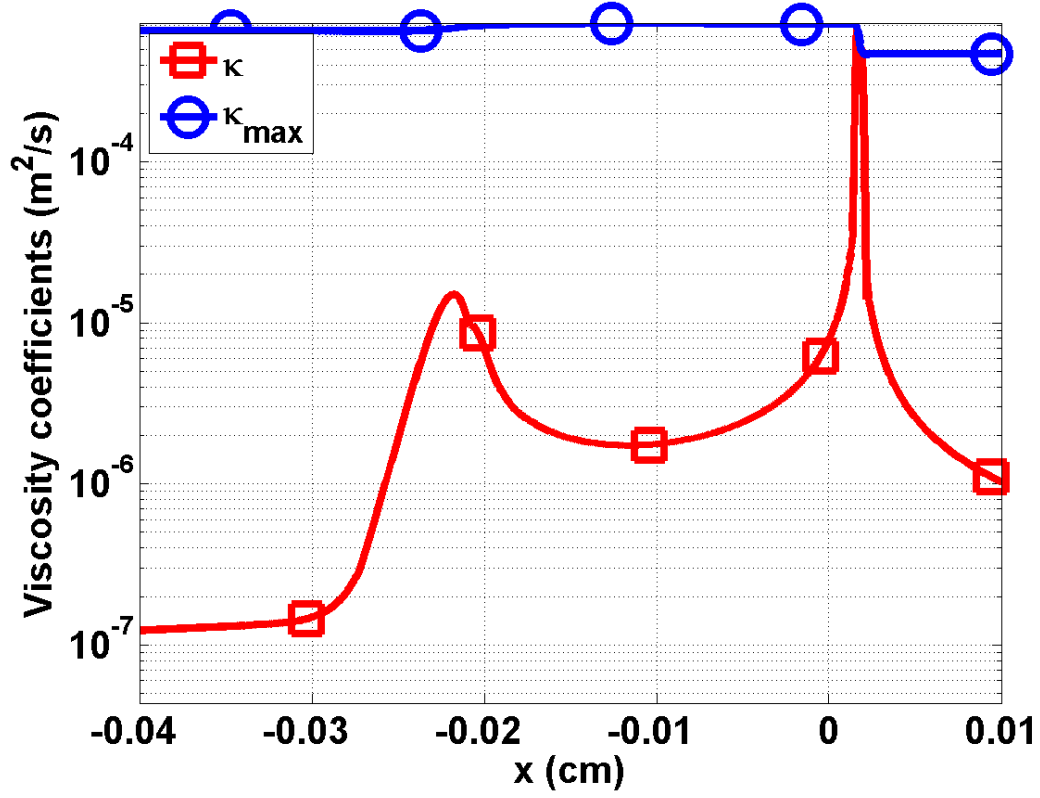


Figure 14: First-order viscosity  $\kappa_{max}$  and second-order viscosity  $\kappa$  profiles at steady-state for Mach 5 test.

#### 4.2.5 Mach 50 shock:

Mach 50 test is known to be challenging. The initial conditions are given in Table 7. The computational domain is of length  $L = 0.2 \text{ cm}$ . Results are once again given at steady-state.

Table 7: Initial conditions for Mach 50.

	left	right
$\rho \text{ (g/cm}^3\text{)}$	1.	6.5189217
$u \text{ (cm/sh)}$	585.6620	89.84031
$T \text{ (keV)}$	1.0	85.51552
$\epsilon \text{ (jerk/cm}^3\text{)}$	$1.372 \cdot 10^{-2}$	$7.33726 \cdot 10^5$

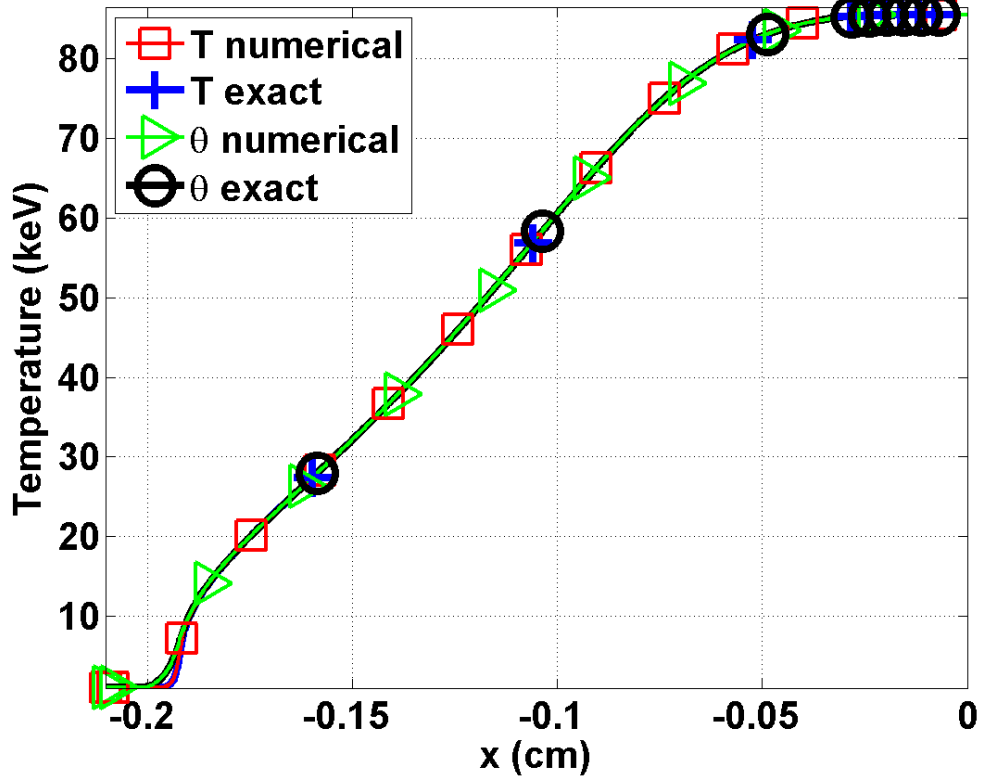


Figure 15: Material and radiation temperature profiles at steady-state for Mach 50 test.

At Mach 50, there is no embedded hydrodynamic shock forming. The density profile is smooth as shown in Fig. 16. In Fig. 15, the material and radiation temperatures overlap on all of the computational domain except for a small region located between  $x = -0.2$  and  $x = -0.18$  cm. In this particular region, the viscosity coefficient saturates to the first-order viscosity (Fig. 17) because of the inflection point in the material temperature profile. The artificial dissipative terms correctly stabilize the material temperature profile without altering the physical solution: the radiation temperature is expected to increase ahead of the material temperature.

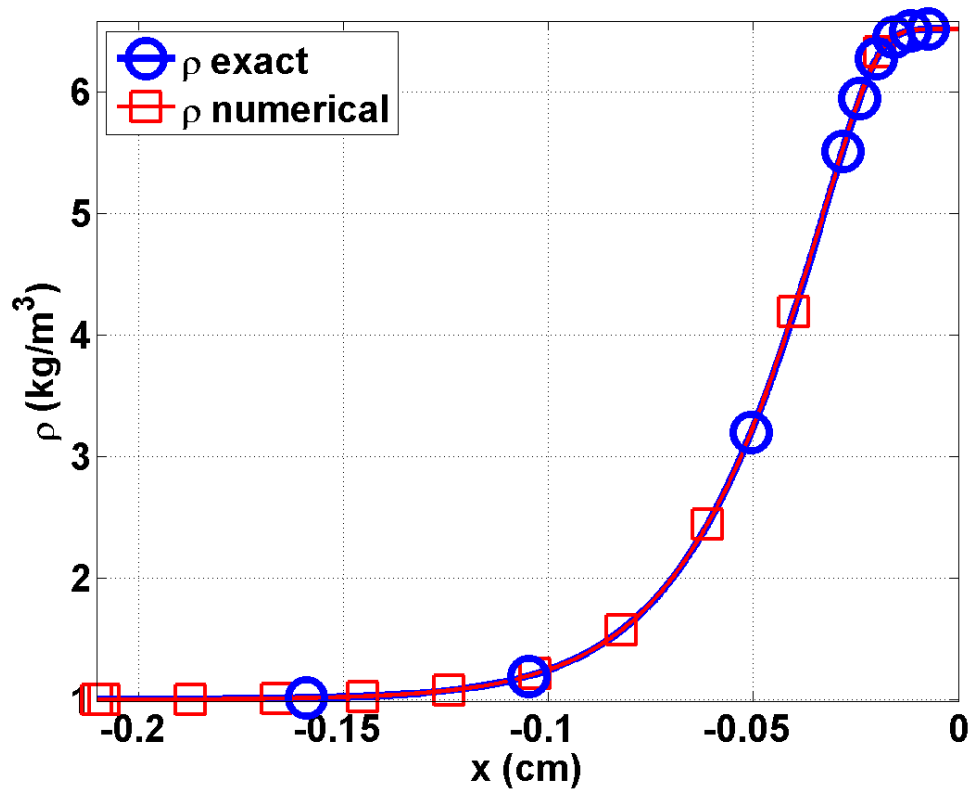


Figure 16: Material density profile at steady-state for Mach 50 test.

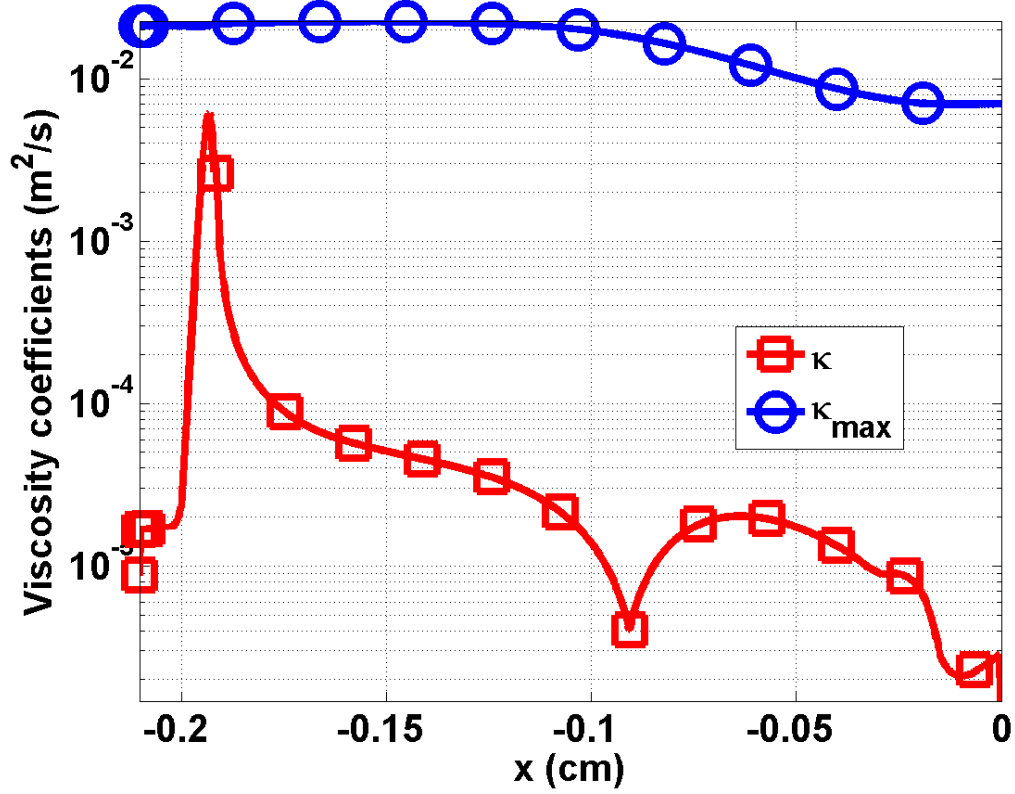


Figure 17: First-order viscosity  $\kappa_{max}$  and second-order viscosity  $\kappa$  profiles at steady-state for Mach 50 test.

## 5 Conclusions

In this paper, we have shown that the entropy-based viscosity method is a valid candidate for solving the 1-D radiation-hydrodynamic equations. A theoretical derivation is given for the derivation of the dissipative terms that are consistent with the entropy minimum principle. The viscosity coefficient  $\kappa$  is defined proportional to the entropy residual that measures the local entropy production allowing detection of shocks. Through the manufactured solution method, it is demonstrated, firstly, that second-order accuracy is achieved when the solution is smooth, and secondly, that the artificial dissipative terms do not affect the physical solution in the equilibrium-diffusion limit.

The method behaves well in the tests performed for Mach numbers ranging from 1.05 to 50. The



main features such as the embedded hydrodynamic shock and the Zeldovich peak are resolved accurately without spurious oscillations. The viscosity coefficient is peaked in the shock region only and behaves as expected. All of these results were obtained by using an unique definition of the viscosity coefficient, that is computed on the fly. The addition of dissipative terms to the set of equations, requires more computational work but is rather simple to implement.

As future work, extension to the multi-D equations could be considered: all of the derivations presented in this paper hold. The definition of the viscosity coefficients  $\kappa$  and  $\kappa_{max}$  do not need to be modified.

## A Proof of the entropy minimum principle for the radiation-hydrodynamic equations with dissipative terms:

In this appendix, a demonstration of the entropy minimum principle for the system of equations Eq. (7) is given. This proof, inspired by [20], details the steps that lead to the derivation of the dissipative terms for the multi-D Euler equations by using the entropy minimum principle.

We start with the hyperbolic system given in Eq. (3) and add dissipative terms to each equation as follows:

$$\begin{cases} \frac{d\rho}{dt} + \rho\partial_x u = \partial_x f \\ \partial_t(\rho u) + \partial_x(\rho u^2 + P + \frac{\epsilon}{3}) = \partial_x g \\ \partial_t(\rho E) + \partial_x[u(\rho E + P)] = \partial_x(h + ug) \\ \partial_t\epsilon + u\partial_x\epsilon + \frac{4}{3}\epsilon\partial_x u = \partial_x l \end{cases} \quad (19)$$

where  $f$ ,  $g$ ,  $h$  and  $l$  are dissipative terms to be determined. Eq. (19) is then recast as a function of the primitive variables  $(\rho, u, e, \epsilon)$  to yield:

$$\begin{cases} \frac{d\rho}{dt} + \rho\partial_x u = \partial_x f \\ \rho\frac{du}{dt} + \partial_x(P + \frac{\epsilon}{3}) = \partial_x g - u\partial_x f \\ \rho\frac{de}{dt} + P\partial_x u = \partial_x h + g\partial_x u + (0.5u^2 - e)\partial_x f \\ \frac{d\epsilon}{dt} + \frac{4}{3}\epsilon\partial_x u = \partial_x l \end{cases} \quad (20)$$

The right-hand side of the internal energy equation can be simplified by choosing the dissipative terms  $g$  and  $h$  as follows:  $h = \tilde{h} - 0.5u^2 f$  and  $g = \rho\mu\partial_x u + uf$  where  $\mu \geq 0$  is a dissipative coefficient. Using these definitions, the system of equation given in Eq. (20) becomes:

$$\begin{cases} \frac{d\rho}{dt} + \rho\partial_x u = \partial_x f \\ \rho\frac{du}{dt} + \partial_x(P + \frac{\epsilon}{3}) = \partial_x g - u\partial_x f \\ \rho\frac{de}{dt} + P\partial_x u = \rho\mu(\partial_x u)^2 + \partial_x \tilde{h} - e\partial_x f \\ \frac{d\epsilon}{dt} + \frac{4}{3}\epsilon\partial_x u = \partial_x l \end{cases} \quad (21)$$

This system of equation admits an entropy function  $s$  that depends on density  $\rho$ , internal energy  $e$  and radiation energy density  $\epsilon$ . In order to prove the entropy minimum principle, a conservation

statement satisfied by the entropy is needed. This equation which is referred to as an entropy residual  $D_e(x, t)$ , can be obtained by a combination of the equations given in Eq. (21). This process is motivated by the following (chain rule)

$$\partial_\alpha s = \partial_\rho s \partial_\alpha \rho + \partial_e s \partial_\alpha e + \partial_\epsilon s \partial_\alpha \epsilon, \quad (22)$$

which holds for any independent variable  $\alpha = x, t$ . It is also required to define the dissipative terms  $\tilde{h}$ ,  $f$  and  $l$ . The following definitions are chosen:

$$\begin{cases} f &= \kappa \partial_x \rho \\ \tilde{h} &= \kappa \partial_x (\rho e) \\ l &= \kappa \partial_x \epsilon \end{cases} \quad (23)$$

where  $\kappa$  is another positive dissipative coefficient.

Thus, using the continuity, the internal energy and the radiation equations of Eq. (21) and using Eq. (22) along with the definition of the dissipative terms, a conservation statement satisfied by the entropy  $s$  is obtained:

$$\frac{ds}{dt} + \underbrace{\left( P \partial_e s + \rho^2 \partial_\rho s + \frac{4}{3} \rho \epsilon \partial_\epsilon s \right)}_{(a)} \partial_x u = \partial_x (\rho \kappa \partial_x s) + \kappa \partial_e s \partial_x s - \rho \kappa \underbrace{X A X^t}_{(b)} + \underbrace{s_e \rho \mu (\partial_x u)^2}_{(c)} \quad (24)$$

where  $X$  is a row vector defined as  $X = (\rho, e, \epsilon)$  and  $A$  is the 3x3 symmetric matrix:

$$A = \begin{bmatrix} \partial_\rho (\rho^2 \partial_\rho s) & \partial_{\rho, e} s & \partial_\rho (\rho \partial_\epsilon s) \\ \partial_{\rho, e} s & \partial_{e, e} s & \partial_{e, \epsilon} s \\ \partial_\rho (\rho \partial_\epsilon s) & \partial_{e, \epsilon} s & \partial_{\epsilon, \epsilon} s \end{bmatrix} \quad (25)$$

In order to show that an entropy minimum principle holds, the signs of the terms (a), (b) and (c) in Eq. (24) need to be studied.

Regarding (a), it is assumed that  $P \partial_e s + \rho^2 \partial_\rho s + \frac{4}{3} \rho \epsilon \partial_\epsilon s = 0$ . The motivation for this is two-fold: First, in order to have a negative sign for the term (a), it would require  $P \partial_e s + \rho^2 \partial_\rho s + \frac{4}{3} \rho \epsilon \partial_\epsilon s$  to have a sign of opposite to that  $\partial_x u$ . The thermodynamic variables cannot be a function of the

material velocity or its derivative under a non-relativistic assumption. Such a statement would not be true when dealing with relativistic equations of state. Second, a similar assumption was made in [20] for multi-D Euler equations (without the radiation energy):  $P\partial_e s + \rho^2\partial_\rho s = 0$ . The term (b),  $XAX^t$ , is a quadratic form and its sign is determined by simply looking at the positiveness of the matrix  $A$  [17]. Here we need to prove that  $A$  is negative-definite which is equivalent to showing the three following inequalities:

$$\begin{cases} A_1 \geq 0 \\ A_2 \leq 0 \\ A_3 = A \geq 0 \end{cases} \quad (26)$$

where  $A_k$  is the  $k^{th}$  order leading principle minor. Determining the sign of the last inequality that corresponds to the determinant of the 3 by 3 matrix  $A$  can be difficult and needs to be simplified. Zeroing out the off-diagonal entries of the last row or column would simplify the expression for the determinant of  $A$ . This can be achieved by assuming  $\partial_\rho(\rho\partial_\epsilon s)$  and  $\partial_{e,\epsilon}s$  are zero, which requires the following form for the entropy function:

$$s(\rho, e, \epsilon) = \tilde{s}(\rho, e) + \frac{\rho_0}{\rho} \hat{s}(\epsilon). \quad (27)$$

where  $\tilde{s}$  and  $\hat{s}$  are two functions whose properties will be provided later. The constant  $\rho_0$  is used for a dimensionality purpose. Next, using the expression of the entropy given in Eq. (27), matrix  $A$  becomes:

$$A = \begin{bmatrix} \partial_\rho (\rho^2 \partial_\rho \tilde{s}) & \partial_{\rho,e} \tilde{s} & 0 \\ \partial_{\rho,e} \tilde{s} & \partial_{e,e} \tilde{s} & 0 \\ 0 & 0 & \rho^{-1} \partial_{\epsilon,\epsilon} \hat{s} \end{bmatrix}$$

Proving that the matrix  $A$  is negative-definite is now straightforward by inspecting the sign of the leading principal minors:

$$\begin{cases} A_1 = \partial_\rho (\rho^2 \partial_\rho \tilde{s}) \leq 0 \\ A_2 = \partial_\rho (\rho^2 \partial_\rho \tilde{s}) \partial_{e,e} \tilde{s} - (\partial_{\rho,e} \tilde{s})^2 \geq 0 \\ A_3 = \rho^{-1} \partial_{\epsilon,\epsilon} \hat{s} A_2 \leq 0 \end{cases} \quad (28)$$

This is easily achieved when assuming that the functions  $-\tilde{s}$  and  $-\hat{s}$  are convex. Thus, the sign of (b) is now determined.

Finally, it remains to determine the sign of the term (c)  $= \partial_e s \rho \mu (\partial_x u)^2$ . The density  $\rho$  and the viscosity coefficient  $\mu$  are both positive: the latest proof for positivity of the density can be found in [20]. Then, only the sign of  $\partial_e s$  remains unknown but it can be determined by studying (a). It was assumed earlier in this appendix that  $P \partial_e s + \rho^2 \partial_\rho s + \frac{4}{3} \rho \epsilon \partial_e s = 0$ . This equation is now recast and split into two equations using Eq. (27). Separation of variables yields:

$$P \partial_e \tilde{s} + \rho^2 \partial_\rho \tilde{s} = \alpha \text{ and } \hat{s} - \frac{4\epsilon}{3} \partial_e \hat{s} = \alpha$$

where  $\alpha$  is a constant to determine. If one sets  $\alpha = 0$ , then the two physics are decoupled, which allows us to reconnect to the result derived in [20] for the multi-D Euler equations:  $P \partial_e \tilde{s} + \rho^2 \partial_\rho \tilde{s} = 0$ . Then, following [20], definitions for  $\partial_e \tilde{s}$  and  $\partial_\rho \tilde{s}$  are obtained:

$$\begin{cases} \partial_e s = \partial_e \tilde{s} = T^{-1} \\ \partial_\rho \tilde{s} = -\frac{P}{\rho^2} \partial_e \tilde{s} \end{cases}$$

where  $T$  is the material temperature which ensures positivity of  $\partial_e s$ . Thus, (c) is positive.

From the above results, the entropy minimum principle follows, so that the sign of the entropy residual is known:

$$\boxed{\partial_t s + u \partial_x s \geq 0} \tag{29}$$

**Remark A.1.** By assuming  $\alpha = 0$ , an expression for the  $\hat{s}$  can be derived by solving the ODE,  $\hat{s} - \frac{4\epsilon}{3} \partial_e \hat{s} = 0$ , which yields:  $\hat{s}(\epsilon) = \beta \exp\left(\frac{4\epsilon^2}{3}\right)$ , where  $\beta$  is a constant. The sign of  $\beta$  is determined by using the condition,  $\partial_{\epsilon,\epsilon} \hat{s} \leq 0$ , derived above, so that  $\beta \leq 0$ .

**Remark A.2.** The viscous regularization derived in this appendix, has two viscosity coefficients:  $\mu$  and  $\kappa$ . For the purpose of this paper, these coefficients are set equal. Under this assumption, the above viscous regularization is equivalent to the parabolic regularization of [24].

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