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LECTURE 4: CONSERVED CHARGE OF ROTATIONAL INVARIANCE  
Wednesday, January 22, 2020

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Recall Noether's theorem from the previous lecture:

$$\frac{d}{dt} \left[ \frac{\delta L}{\delta \dot{\vec{x}}} \delta \vec{x} \right] = 0$$

If the action is rotationally invariant,

$$\vec{x} \rightarrow R(\hat{\mathbf{n}}, \theta) \vec{x}$$

where

$$R(\hat{\mathbf{n}}, \theta) = e^{i \vec{L} \cdot \hat{\mathbf{n}} \theta}$$

In the previous lecture, we found that

$$i(L^a)_{ij} = \epsilon_{ij}^a \equiv \epsilon_{aij}$$

If we expand the exponential to a few terms, we find

$$e^{i \vec{L} \cdot \hat{\mathbf{n}} \theta} \rightarrow 1 + i \vec{L} \cdot \hat{\mathbf{n}} \theta + \mathcal{O}(\theta^2)$$

as  $\theta \rightarrow 0$ . We find  $\delta \vec{x}$  to be

$$\begin{aligned} \delta \vec{x} &= \left[ (i \vec{L} \cdot \hat{\mathbf{n}})_{ij} \theta \right] x_j \\ &= (i(L^a)_{ij} n^a x_j) \theta \\ &= (\epsilon_{aij} n_a x_j) \theta \end{aligned}$$

If our Lagrangian has the form

$$L = \frac{1}{2} m \dot{x}^2 - V(x)$$

we find that Noether's theorem gives us

$$\begin{aligned} \frac{d}{dt} [m \dot{x}_i \delta x_i] &= 0 \\ &= \frac{d}{dt} m [\dot{x}_i (n_a \theta \epsilon_{aij} x_j)] \\ &= \frac{d}{dt} [m \dot{x}_i \epsilon_{aij} x_j n_a \theta] \end{aligned}$$

Because  $\hat{\mathbf{n}}$  and  $\theta$  are arbitrary and this equation must be true for all  $\hat{\mathbf{n}}$  and  $\theta$ ,

$$\frac{d}{dt} \left[ \underbrace{m \dot{x}_i}_{p_i} \epsilon_{aij} x_j \right] = 0$$

so

$$p_i x_j \epsilon_{aij} = \vec{x} \times \vec{p} = \vec{L} \quad \longrightarrow \quad \text{invariant}$$

## 0.1 Conservation Laws in Quantum Mechanics

The fundamental time-evolution equation in QM is the Schrödinger equation:

$$i\hbar \frac{\partial}{\partial t} |\psi\rangle = H |\psi\rangle$$

The Schrödinger picture is a formulation where we make the operators independent of time, but allow the wave functions to be time-dependent.

The Heisenberg picture is a formulation where all of the operators are time-dependent whereas the wave functions are time-independent.

There is a simple way to transform between the two using a time-evolution operator:

$$U(t', t) = e^{-iH(t'-t)/\hbar}$$

If we work in the Schrödinger picture, we know that  $|\psi(t')\rangle = U(t', t) |\psi(t)\rangle$ . If we consider the expectation value of some operator:

$$\langle \psi(t') |_S O_S |\psi(t)\rangle_S = \langle \psi(t') | U^\dagger(t', t) O_S U(t', t) |\psi(t)\rangle$$

We could equivalently define

$$O_H(t') = U^\dagger(t', t) O_S U(t', t)$$

such that

$$\langle \psi(t') |_S O_S |\psi(t)\rangle_S = \langle \psi |_H O_H(t') |\psi\rangle_H$$

where

$$|\psi\rangle_H \equiv U(t', t) |\psi(t)\rangle_S$$

We can use the Schrödinger equation on the Heisenberg picture operator:

$$\begin{aligned} i\hbar \frac{\partial}{\partial t} O_H(t) &= \\ i\hbar \frac{\partial}{\partial t} \langle \psi |_S O_S |\psi\rangle_S &= \left[ i\hbar \frac{\partial}{\partial t} \langle \psi |_S \right] O_S |\psi\rangle_S + \langle \psi |_S O_S \left[ i\hbar \frac{\partial}{\partial t} |\psi\rangle_S \right] \\ &= -\langle HO |\psi| HO \rangle + \langle OH |\psi| OH \rangle \\ &= \langle [O, H] |\psi| [O, H] \rangle \\ &= i\hbar \langle \psi |_H \frac{d}{dt} O_H(t) |\psi\rangle_H \end{aligned}$$

so

$$i\hbar \frac{d}{dt} O_H(t) = [O, H]$$

What does this have to do with conserved quantities? If  $[O, H] = 0$ ,  $O$  is time independent. In quantum mechanics, a symmetry is always expressible in terms of a unitary transformation

$$U = e^{i\vec{X} \cdot \vec{\lambda}}$$

where  $\vec{X}^\dagger = \vec{X}$  are the generators of the symmetry which obey a Lie algebra  $[X_a, X_b] = if_{abc} X_c$ .

The difference between classical and quantum mechanics is that everything is an operator, and operators transform under symmetries:

$$O \rightarrow U^\dagger(\vec{\lambda}) O U(\vec{\lambda})$$

where  $\vec{\lambda}$  is the set of parameters which determine the group element. Now consider some of the typical operators and how they transform. Under rotations, the position operator transforms as

$$\vec{x} \rightarrow U(\hat{n}, \theta) \vec{x} U(\hat{n}, \theta)$$

or

$$\vec{x}' = e^{-i\vec{L} \cdot \hat{n} \theta / \hbar} \vec{x} e^{i\vec{L} \cdot \hat{n} \theta / \hbar}$$

Consider an infinitesimal rotation ( $\theta \rightarrow 0$ ):

$$\vec{x}' = (1 - i\frac{\vec{L} \cdot \hat{n} \theta}{\hbar}) \vec{x} (1 + i\frac{\vec{L} \cdot \hat{n} \theta}{\hbar}) + \mathcal{O}(\theta^2)$$

or

$$\vec{x}' = \vec{x} - (\imath \frac{\vec{L}}{\hbar} \cdot \hat{n} \theta) \vec{x} + \vec{x} (\imath \frac{\vec{L}}{\hbar} \cdot \hat{n} \theta)$$

so

$$\delta \vec{x} = \vec{x} (\imath \frac{\vec{L}}{\hbar} \cdot \hat{n} \theta) - (\imath \frac{\vec{L}}{\hbar} \cdot \hat{n} \theta) \vec{x}$$

so

$$\delta x_a = [(x_a (\imath L_b) / \hbar) - (\imath L_b X_a) / \hbar] \hat{n}_b \theta$$

We define the angular momentum operator as

$$\vec{L} \equiv \vec{x} \times \vec{p} = \vec{x} \times \left( \imath \hbar \frac{\partial}{\partial \vec{x}} \right)$$

We can also write this in index notation:

$$L_b = -\imath \hbar x_i \partial_j \epsilon_{ijb}$$

Let's now apply this to our  $\delta x_a$  formula:

$$\delta x_a = (\imath \hat{n}_b \theta) \left[ x_a, \imath \frac{L_b}{\hbar} \right]$$

Now we just need to figure out what the commutator is.

$$\left[ x_a, \imath \frac{L_b}{\hbar} \right] = [x_a, x_c p_d \epsilon_{cdb}]$$

$\epsilon_{cdb}$  is just a constant, we can take it out, and we are left with

$$[x_a, \imath x_c p_d] = \imath [x_a, x_c] p_d + x_c \imath [x_a, p_d]$$

since

$$[A, BC] = [A, B]C + B[A, C]$$

Position commutes with itself and  $[x_a, p_d] = \imath \hbar \delta_{ad}$  so

$$[x_a, x_c p_d] = -x_c (\delta_{ad})$$

Finally

$$\begin{aligned} \delta x_a &= \imath \hat{n}_b \theta \epsilon_{cdb} (\imath \hbar x_c \delta_{ad}) \\ &= -\hat{n}_b \theta x_c \epsilon_{cab} \\ &= -\hat{n}_b \theta x_c \epsilon_{abc} \end{aligned}$$

This is very similar to the classical case where

$$\delta x_a = \hat{n}_b \theta x_c \epsilon_{abc}$$

Whereas in the quantum case we have

$$\delta x_a = -\hat{n}_b \theta x_c \epsilon_{abc}$$

We have shown that the operator  $\vec{x}$  transforms just like a vector under rotation. If you did the same thing with  $\vec{p}$ , you would find the exact same result (and a similar result with any vector operator).